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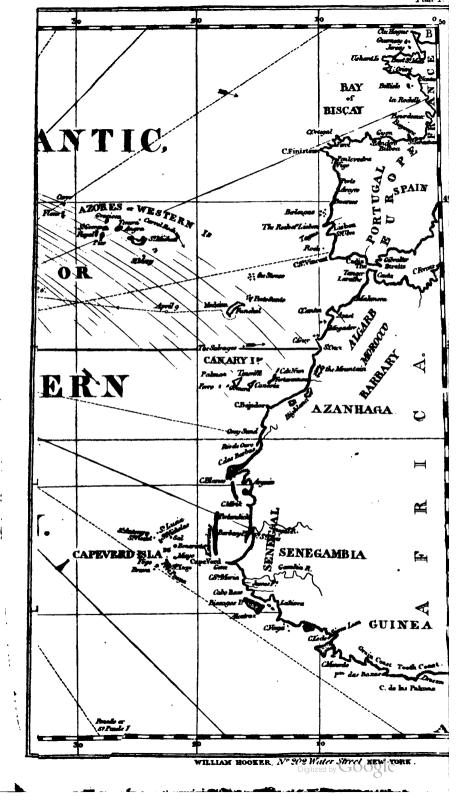




Favis Goodbreak

Many Charles Court





THE

NEW AMERICAN PRACTICAL NAVIGATOR:

BRING AN

EPITOME OF NAVIGATION:

CONTAINING ALL THE TABLES NECESSARY TO BE SED WITH THE

NAUTICAL ALMANAC.

IN DETERMINING

THE LATITUDE. AND THE LONGITUDE LUNAR OBSERVATIONS:

KEEPING A COMPLETE RECKONING AT SEA:

ILLUSTRATED BY

PROPER RULES AND EXAMPLES:

WHOLE EXEMPLIFIED IN A JOURNAL. THEKEPT FROM BOSTON TO MADEIRA.

IN WHICH ALL THE BULES OF NAVIGATION ARE INTRODUCED."

ALSO.

THE DEMONSTRATION OF THE USUAL RULES OF TRIGONOMETRY: PROBLEMS

IN MENSURATION, SURVEYING AND GAUGING: DICTIONARY OF SEA-TERMS:

AND THE MANNER OF

PERFORMING THE MOST USEFUL EVOLUTIONS AT SEA.

WITH AN

APPENDIX.

METHODS OF CALCULATING ECLIPSES OF THE SUN AND MOON, AND OCCULTATIONS OF THE FIXED STARS: RULES FOR FINDING THE LONGITUDE OF A PLACE BY OBSERVATIONS OF ECLIPSES OR OCCULTATIONS: AND A NEW METHOD FOR FINDING THE LATITUDE BY TWO ALTITUDES.

BY NATHANIEL BOWDITCH, LL. D.

Fellow of the Royal Societies of London, Edinburgh, and Dublin; of the American Philosophical Society,
held at Philodelphia: of the American Londomy of Arts and Sciences; of the Connecticut Academy
of Arts and Sciences; of the Literary and Philosophical Society of New-York, 4c.

SIXTH STEREOTYPE EDITION.

----**NEW-YORK:**

PUBLISHED BY EDMUND M. BLUNT, PROPRIETOR, AND AUTHOR OF THE AMERICAN COAST PILOT, No. 202, WATER-STREET.

John Gray & Co. Print.

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1826.

STANDARD WORKS.

PUBLISHED BY EDMUND M. BLUNT,

Anthor of the American Coast Pilot, &c.

202

WATER, CORNER OF FULTON STREET—NEW-YORK, [OLD ESTABLISHED STAND.] BOOKS.

BOWDITCH'S PRACTICAL NAVIGATOR, 6th edition, stereotyped. This work has been re-published

BOWDITCH'S PRACTICAL NAVIGATOR, the cition, stereotyped. This work has been re-published in London, and has a decided preference to empessant.

BLUNT'S AMERICAN COAST PILOT, 10th edition, greatly improved.

THE MERCHANT'S AND SHIP MASTERS ASSISTANT, comprehending all the necessary mercantile information for Merchants and Shipmasters. In this work all recent commercial regulations are introduced, and the most experienced will find something new.]

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edition.

NAUTICAL ALMANACS, from the year 1811 to 1828, both inclusive—to be continued annually. Explanation stereotyped, and English copy corrected.

Errors in the English copies of the Nautical Almanac,

(1826, Over Fifty errors Corrected in BLUNT'S EDITIONS ONLY.

[1825, Over Twenty errors SEAMANSHIP and NAYAL TACTIOS. Second Edition, with Places.

WARD'S LUNAR TABLES. 2d. edition.

WARD'S LUNAR TABLES, 2d. edition.

CHARTS:

A NEW CHART,	extending from New-York to Havana, including Bahama Banks and Channels, improved by actual Surveys of the Chesapeake Bay, by order of the Navy Department, and of Cape Hatteras, Look Out and Fear Shoats, in conformity to an act of Congress of the United States, and conducted under the direction of J. D. KLLIOT, Esq. Capt, in the U. S. Navy, and by permission of Hon. SMITH THOMPSON, Secretary of the Navy, copied, and contains all the Surveys made on the Coast of North Carolina to the present time (1626.) This Chart has since been improved by a survey from Sandy Hook to Cape May, under direction of Capt. JONATHAN COLESWORTHY, and EDMUND BLUNT, hydrographer, in the sloop New Packet. It has also several
	PLANS of HARBOURS, from actual surveys. of the Mississippi River, extending to New Orleans, including Mobile, &c. with sailing directions, and Plan of Mobile, on a large scale, from actual survey, published in 1825.
,	of Bahama Bank, from actual survey mode in sloop Orbit, in 1820, with salling directious, by the direction and at the expense of E. M. BLUNT, by E. C. WARD, U. S. Navy, and EDMUND BLUNT, hydrographer. [It is worthy of remark, that nine-teen vessels were lost on the Bahama Bank the year previous to this survey, since which, accidents have rarely occurred, and the correctness of the Chart admits not a question, but has received the approbation of thousands, as being the most correct Chart of Bahama Bank extant.]
-	from New-York to Nova protfa, extending from latitude 39° N. to latitude 47° N. longitude 68° W. to longitude 74° W. including the whole of St. George's Bank, improved to August 1821, by government and other surveys.
,	of the Atlantic or Western Ocean, improved to 1820, with an Analysis of the authorities upon which the dangers have been inserted on the Chart. The Tracks extend to the Equator, and are continued on the Chart of the South Atlantic Ocean, of the South Atlantic Ocean, containing more authentic information than any extant,
······································	part of which describes dangers lately discovered, with original Plans of Harbours and Views. of the North Coast of Brazil, showing the entrances and courses of the Rivers Para and Amazon.
	of the West Indies, on six sheets, which may be had separate. of River La Plare, according to late Surveys, with Salling Directions on the Chart. of the Coast of Guayana.—of the Coast of Brazil.—of the Island of Bermudes, with Sailing Directions.—of Long-Island Sound, improved to 1825.—of the Coast of Labra- dor.—of Newfoundland.—of the Coast of Brazil, from Maranbam to the river La

Plate, &c. &c. the sailing Directions for which are in Blunt's American Coast Pilot.

PLAN of New-London Harbour, surveyed by CHARLES MORRIS, Esq. of the United States Mavy, by order of Commodore RODGERS, and to him respectfully dedicated.

SOUTHERN DISTRICT OF NEW-YORK, SS.

BE IT REMEMBERED, That on the twenty-second day of June, A. D. 1826, in the 'fiftieth year of the Independence of the United States of America. Edmund M. Biunt, of the said district, hath deposited in this office the title of a book, the right whereof he claims as proprietor, in the words following, to wit:

"The New American Practical Navigator: being an Epitome of Navigation: containing all the Ta"thes necessary to be used with the Nautical Almanac, in determining the Latitude, and the Longitude
to Lunar Observations; and keeping a complete Reckoning at Sea; illustrated by proper Rules
and Examples: the whole exemplified in a Journal, kept from Boston to Madeira, in which all the
rules of Navigation are introduced. Also, the Demonstration of the usual Rules of Trigonometry;
Problems in Mermuration, Surveying, and Gauging; Dictionary of Sea-Terms; and the manner of performing the most useful Evolutions at Sea. With an Appendix, containing Methods of calculating Eclipses of the Sun and Moon and Occultations of the Fixed Stars; Rules for finding the latitude of a
place by Observations of Eclipses or Occultations; and a new method for finding the latitude by
"two altitudes. By Nathaniel Bowditch, LL. D. Fellow of the Royal Societies of London, Edinburgh,
and Dablin; of the American Philosophical Society, held at Philadelphia; of the American Academy
of Arts and Sciences; of the Connecticut Academy of Arts and Sciences; of the Literary and Philospophical Society of New-York, &c. Sixth stereotype edition."

In conformity to the act of congress of the United Stares, entitled "An Act for the encouragement
of learning, by securing the copies of Map, Charts, and Books, to the Authors and Proprietors
of such copies, during the time therein mentioned." And also to an act, entitled "an act supplementary"
to an act, entitled an act for the encouragement of learning, by securing the copies of Map, Charts, and
Books, to the authors and proprietors of such copies, during the times thereon and extendland the prints."

Clerk of the Southern District of New-York.

Clerk of the Southern District of New-York.

VK555 B7 1826

REPORT

Of the Committee, appointed by the EAST INDIA MARINE SOCIETY of Salem, at their meeting on the 6th of May, 1801, to examine a work called "The New American Practical Navigator, by Nathaniel Bouditch, F. A. A."

AFTER a full examination of the system of Navigation presented to the society by one of its members (Mr. Nathaniel Bowditch) they find, that he has corrected many thousand errors existing in the best European works of the kind; especially those in the Tables for determining the latitude by two altitudes, in those of difference of latitude and departure, of the sun's right ascension, of amplitudes, and many others necessary to the Navigator. Mr. Bowditch has likewise, in many instances, greatly improved the old methods of calculation, and added new ones of his own. That of clearing the apparent distance of the moon, and sun or stars, from the effect of parallax and refraction, is peculiarly adapted to the use of seamen in general, and is much facilitated (as all other methods are) in the present work, by the introduction of a proportional table into that of the correction of the moon's altitude. His Table nineteenth, [the twentieth of the present edition] of corrections to be applied in the lunar calculations, has the merit of being the only accurate one the Committee are acquainted with. He has much improved the table of latitudes and longitudes of places, and has added those of a number on the American coast, hitherto very inaccurately ascertained.

This work, therefore, is, in the opinion of the Committee, highly deserving of the approbation and encouragement of the society, not only as being the most correct and ample now extant, but as being a genuine American production; and as such they hesitate not to recommend it to the attention of Navigators, and to the public at large.

JONATHAN LAMBERT, BENJAMIN CARPENTER. JOHN OSGOOD, JOHN GIBAUT, JACOB CROWNINSHIELD,

Committee.

APPROVED. BENJAMIN HODGES, President.

A TRUE COPY. MOSES TOWNSEND, Sec'ry.

Salem, May 13, 1801.

PREFACE.

IN the preface to the first edition of this work, it was observed, that the object of the publication was to collect into one volume all the rules, examples and tables necessary for forming a complete system of practical navigation. To do this, those authors were consulted whose writings afforded the best materials for the purpose,* and such additions and improvements were introduced as were suggested by a close attention to the subject; and the accuracy of the tables accompanying the work was ensured by actually going through all the calculations necessary to a complete examination of them, making the last figure exact to the nearest unit. In performing this, above eight thousand errors were discovered and corrected in Moore's Practical Navigator, and above two thousand in the second edition of Maskelyne's Requisite Tables.† Almost all the errors in Maskelyne's collection were in the last decimal place, and in most cases would but little affect the result of any nautical calculation; but when it is considered that most of those tables are useful in other calculations, where great accuracy is required, it will not be deemed an unnecessary improvement to have corrected so great a number of small errors.

Several articles were added in the second edition, particularly the description and use of the circular instrument of reflection, methods of surveying harbours, new tables, &c. In the third and subsequent editions, several improvements have been made, particularly in the method of correcting the dead reckoning, and in the articles of surveying. An Appendix is given, containing methods of projecting and calculating eclipses of the moon and sun, and occultations of the fixed stars or planets by the moon; rules for deducing the longitude of a place from observations of eclipses of the sun or occultations; a new and short method of calculating the altitude and longitude of the nonagesimal degree of the ecliptic; solutions of several useful problems of Nautical Astronomy, and an improvement of Napier's rules for the solution of spheric triangles. Several new tables were added. The

table of latitudes and longitudes is much increased and corrected.

Also an entirely new article is given in this edition on the method of finding the latitudes by two altitudes of the same, or of different objects: the solutions being direct and simple, embracing all the cases of the problem: a point which has not been attended to in some works of celebrity. This is an important addition to the present work, and it is recommended to the considera-

tion of navigators.

The tables published separately in the Appendix of the first edition are introduced into the body of this work, and are extended so as to render the use of them more simple. The short and easy method of working a lunar observation, published in that Appendix, which has one great advantage ever all other approximate methods, in the manner of applying the corrections (all' them being additive) is here explained and illustrated by several examples. Two other methods of correcting the apparent distance are given; one being that invented by the author of this work in the year 1795;

^{*} The works chiefly consulted were those published by Maskelyne, Roberston, Patoun, Rios, &c. and a treatise on "Seamanship," published at London in 1785. In this new edition, the work of the Chavaller de Borda, entitled "Description et Usage du Cercle de Reflection," &c. has also been

The third edition of that work the errors of the table of proportional logarithms are corrected.

This method was communicated to Mr. Be Lambre, who published an account of it in the Commissance des temps pour Pannee, 1802."

the other an improvement of Witchell's method, in which, without altering materially the calculation, the number of cases is considerably reduced.

Any person who wishes to examine the tables, may do it by the methods used for that purpose, which will here be explained with some additional remarks.

Tables I. and H. were calculated by the natural sines taken from the fourth edition of Sherwin's logarithms, which were previously examined, by differences; when the proof-sheets of the first edition were examined, the numbers were again calculated by the natural sines in the second edition of Hutton's logarithms; and if any difference was found, the numbers were calculated a third time by Taylor's logarithms.

TABLE III. contains the meridional parts for every degree and minute of

the quadrant, calculated by the following rule, viz.

M=T×0.0007915704468. in which T is the log-tangent less radius of half the latitude increased by 45° taken to seven places of figures, reckoned as integers, and M is the meridional parts of that latitude in miles.

Table IV. contains the declination of the sun, which was compared with the Nautical Almanacs for the years 1824, 1825, 1826, and 1827, and mark-

ed to the nearest minute.

Table IV. A. The Equation of Time, for the years 1824, 1825, 1826,

and 1827.

Table V. contains the correction of the sun's declination, as published by Dr. Maskelyne. The correction taken from this table will rarely differ more than 16 or 17 seconds from the truth.

TABLE VI. contains the mean of the sun's right ascension, taken from

the Nautical Almanacs for the years 1824, 1825, 1826, and 1827.

Table VI. A. contains the correction for the daily variation of the Equation of Time.

TABLE VII. contains the amplitudes of the sun for various latitudes and

declinations calculated by Taylor's logarithms by this rule:

Log. sec. lat. + Log. sine declination—10.0000000 = Log. sine amplitude. Table VIII. contains the right ascensions and declinations of 76 stars of the first and second magnitudes, with their annual variations, adapted to the beginning of the year 1820. This table was formed from that published by the Astronomer Royal at Greenwich (Mr. Pond) in the Nautical Almanac for 1823, with the addition of a number of stars from the Catalogue of Baron Von Zach.

TABLE IX. contains the time of the sun's rising and setting, calculated by

Taylor's logarithms by this rule :

Log. cos. hour=Log. tang. declin. + Log. tang. latitude-10.0000000,

Table X. contains the distances at which any object is visible at sea, calculated by the rule given in § 195 of Vince's Astronomy, in which the terrestrial refraction was noticed: this circumstance was neglected by Robertson, Moore, and others, and of course their tables are erroneous. The rule given by Mr. Vince, expressed in logarithms, is this:

0.12155+Half log. of height in feet=Log. of dist. in statute miles.

In reducing the rule to logarithms, the radius of the earth was called 20911790 feet, which agrees nearly with the mean value given in De La Lande's Astronomy.

TABLE XI. is a common table of proportional parts, the construction of

which does not need any explanation.

TABLE XII. contains the refraction of the heavenly bodies, calculated by Dr. Bradley's rule, supposing the refraction to be as the tangent of the apparent zenith distance of the object decreased by three times the refraction, the horizontal refraction being supposed equal to 35'.

The rule expressed in logarithms is this:

Log. tang. (app. zen. dist.—3. refraction)—8.2438534—Log. of ref. in sec.

The numbers calculated by this rule agree nearly with those published

in Table I. of Maskelyne's Requisite Tables.

TABLE XIII. contains the dip of the horizon for various heights, calculated by the rule in § 197 of Vince's Astronomy, in which the terrestrial refraction is allowed for. All the numbers of this table differ a little from those published by Dr. Maskelyne, who had made a different allowance for that refraction. The rule given by Mr. Vince, expressed in logarithms, is: 1.7712711 + half the log. of the height in feet = Log. dip in seconds.

1.7712711 + half the log. of the height in feet = Log. dip in seconds.

TABLE XIV. contains the sun's parallax in altitude, calculated by multiplying the natural sine of the apparent zenith distance by the sun's horizontal parallax 83". The numbers in this table agree with those published by

Dr. Maskelyne.

Table XV. contains the augmentation of the moon's semi-diameter == 15". 626 × sine) 's altitude. This table agrees nearly with that published by Maskelyne.

TABLE XVI. contains the dip for various distances and heights, calcula-

ted by this rule.

$$D = \frac{5}{7}d + 0.56514 \times \frac{h}{d}$$

in which D represents the dip in miles or minutes, d the distance of the land

in sea miles, and k the height of the eye of the observer in feet.

Tables XVII. XVIII. and XIX. were first calculated by the author of this work, and published in the Appendix to the first edition. The correction in the first of these tables is equal to the difference between the star's refraction and 60′. The correction of Table XVIII. is equal to the difference between 60′ and the correction of the sun's altitude for parallax and refraction. The correction of Table XIX. is equal to the difference between 59′ 42″ and the correction of the moon's altitude for parallax and refraction. The logarithms in each of these tables may be found by adding together the constant log. 9.6990, the log. co-sine of the apparent altitude of the object, the proportional logarithm of the correction of the altitude of the object for parallax and refraction, and rejecting 20 from the index. The method of performing these calculations are so obvious, that it is unnecessary to enter into any farther explanation. Most of the numbers in these tables were calculated three different times.

Table XX. There are two columns in this table corresponding to each degree, the numbers in one column exceed those of the other by 18", the numbers in the least column express the difference b, between the base B and the hypotenuse $B\pm b$ of a right angled spheric triangle, whose third side P, never exceeds 60'; the argument at the top of the table being B, and at the side $60'\pm P$. The value of b being found by this rule by Taylor's logarithms:

Log. b in seconds=Log. co-tang. B+Log. vers. sine P-14.6855749-Diff.

log. sines of B and $B \pm \frac{1}{4}b$

in which the last term may in most cases be neglected.

Table XX. (New Form.) corrections in seconds additive. See appendix, page 618.

Table XXI for turning time into degrees, is the same as in other works of this kind.

TABLE XXII. contains the proportional logarithms for three hours. The numbers of this table may be found by subtracting the logarithm of the time in seconds from the log. of 10800"; or, which is the same thing, by the following rule:

Prop. log. T=4.0864758—log. of T in seconds, neglecting the three right

hand figures of the remainder.

TABLE XXIII. was first constructed by Mr. Douwes of Amsterdam, about the year 1740, for which he received £.50 of the Commissioners of Longitude in England. This table was published in the first and second editions of the Requisite Tables; in the former of which it was carried as far as six hours; in the latter the table of Log. Rising was extended to 9 hours; in the present edition of this work it is extended to 12 hours. The numbers in this table are easily deduced from the log. sines, log. co-secants, and log. versed sines of the hour to which they correspond. Thus, if the time, opposite to any number of these tables turned into degrees, is H, we shall have

> Log. 4 elapsed time of H=log. co-secant H-10.0000000 Log. middle time =Log. sine H-4.6989700 Log. rising H $\begin{cases} = \text{Log. versed sine H} - 5.00000000 \\ = 2 \times \log. \text{ sine } \frac{1}{2} \text{ H} - 14.6989700 \end{cases}$

By means of these formulæ, the numbers of Table XXIII. were calculated by Sherwin's, Hutton's and Taylor's logarithms, and above a thousand errors were discovered in the second edition of the Requisite Tables, most of which were in the additional three hours (from six to nine hours) not published in About two thirds of these additional numbers differ from the first edition. their true values by one or two units.

TABLE XXIV. was compared with Sherwin's and Hutton's Tables, and a

few errors corrected.

TABLE XXV. contains the log. sines, log. tangents, &c. corresponding to points and quarter points of the compass. This was compared with Sherwin's, Hutton's, and Taylor's logarithms.

TABLE XXVI. contains the common logarithms of numbers, which was

compared with Sherwin's, Hutton's and Taylor's logarithms.

TABLE XXVII. contains the common log. sines, tangents, secants, &c. This was compared with Sherwin's, Hutton's and Taylor's tables. Two additional columns are given in this table, which are very convenient in finding the time from an altitude of the sun. The degrees are marked to 180°, which saves the trouble of subtracting the given angle from 1800 when it exceeds 90°.

TABLE XXVIII. was calculated by proportioning the daily variation of

the time of the moon's passing the meridian.

Table XXIX. contains the correction of the moon's altitude for parallax

and refraction, corresponding to the parallax 57' 80".

TABLES XXX. and XXXI. are tables of proportional parts, taken from the

Requisite Tables, with a few corrections.

Table XXXII. contains the variation of the altitude of any heavenly body for one minute of time from noon, for various degrees of latitude and The following method was used in constructing the table:-A and B were calculated for each degree of declination by these

Log. A=Log. 1".96349+2 log. cos. declination-20.00000.

Log. B=Log. A. + log. tang. declination-10.00000. and then the correction of the table corresponding to the zenith distance Z (=Lat. +Dec.) was found by this formulæ. $A \times co$ -tang. Z + B. To fa-

cilitate the computation of these numbers, a table of the products of A by

the whole numbers from 1 to 9 was calculated.

TABLE XXXIII. contains the squares of the minutes and parts of a minute corresponding to every second from 0" to 12' 59". This requires no ex-

planation.

TABLE XXXIV. contains the error of an observed angle arising from a deviation of 1' in the parallelism of the surfaces of the central mirror, those surfaces being supposed to be perpendicular to the plane of the instrument. The correction in the fifth column of this table corresponding to any angle

A in the first column may be found nearly by Hutton's logarithms, as forlows: to the constant logarithm 0.07340 add the log, secant of A A, find this in the column of log-tangents and take out the corresponding material secunt B. then the correction will be 2' (B-1,55.) The numbers in the second tolumn are nearly equal to those in the fifth corresponding to the angle A + 20°, decreased by 1". 63. The numbers in the third column are equal to the difference between 1". 63, and the numbers in the fifth corresponding to A 100 20°. The numbers in the fourth column are equal to the half difference between 1". 63. ence of the numbers on the same horizontal line in columns second and third when it exceeds 40°, otherwise their half sum.

TABLE XXXV. contains the correction to be applied to an observation taken in a direction inclined to the plane of the instrument. The following rule was used in calculating this table: Find an arch A such that

Log. sine A=Log. sine 4 observed angle+log. co-sine of error of inculation. Then the difference between 2 A and the observed angle will be the

tabular correction.

TABLE XXXVI. contains the variation of the mean refraction (given in Table XII.) for various temperatures and densities of the air. The correct tion given in this Table is nearly the same as that deduced from Dr. Brade v rule, which is as follows :- As the mean height of the barometer 29,0 inches is to the true height, so is the mean retraction to the corrected refragalism; and as 350 increased by the height of Fahrenheit's thermometer is to ann

so is the corrected to the true refraction.

TABLE XXXVII, contains the latitudes and longitudes of the fixon alves of the 1st. 2d. and 3d. magnitudes. The nine stars, from which the distances are marked in the Nautical Almanac, are given from the table pall-lished in the Nautical Almanac for 1820. The rest were deduced from the table published in the second edition of Doctor Mackag's treatise on longitude, supposing the annual procession 50". Sh and the secular appatien as in his table. In the third edition of his work, the procession was ablowed on the latitude as well as the longitude, which causes an error of about w in the latitudes of the stars in that edition.

TABLE XXXVIII. was calculated by this rule. Suppose L to be the latitude, R the reduction of latitude, then log, co-tang, (L-R) = 0.0020001 + log, co-tang, L. The reduction of parallels corresponding to 55', 57' and a log, co-tang, L. was found by these formulæs respectively \$ 3-5% 3 cos. : Let x', t-3"

cos. 2 L; 6". 1-6" 1 cos. 2 L.

TABLE XXXIX, was calculated by the rule in vol. 1., page 854 of Vmce's Astronomy, supposing S to be the place of the sun, P-that of the planet, and T that of the earth.

Aberration = -20'.cos. STP -20" V BT. coy. SPT. Making use of the di-

tances, &c. given by La Place in vol. III: of his Mecanique Celoute. small alteration was made in the rule in calculating the aberration of Mer.

TABLE XL. was calculated by - 17".9. sine long. D's unde.

TABLE XLL was calculated by - 20", cos. argument.

Table XLill. Part I. =-19", 175 cos. arg. Part II. =0",357 cos. arg. Part III. =-3".9314 cos. seg.

Part II: =-6" 50 cas, arg. Part II. =-1", 42 cos, arg. Part III. =-16",582 cinc arg. TABLE XLIII. Part I.

TABLE XLIV. Part I. = 8".1845 sine arg. Part II. = (arg. in seconds)

Part III. = 960" × sine h's par, in lat × tang- h's true lat. -950. Yersed sine par, in lat.

If we suppose the sum of these three parts to be S seconds, and the moon's horizontal semi-diameter to be D minutes. Part IV. corresponding to S and D will be S × (D + 16)(D-16)

TABLE XLV. The arguments at the side being B and 12—B hours, and the second difference at the top A, the correction of this table will be $A \times B(12-B)$

288

Table XLVI. contains the Latitudes and Longitudes of the most remarkable ports, harbours, &c. in the world. Great alterations were made in this table in the fifth edition, particularly by the insertion of more than thirteen hundred additional places in the India Seas and in the Pacific Ocean, besides various corrections in other parts of the world, consulting the latest and best authorities. Several corrections and additions have also been made in the present edition.

TABLE XLVII. contains the times of high water on the full and change of the moon, with the vertical rise of the tide, at many ports, harbours, &c. in the world. This table (like the preceding) depending wholly on observations, is therefore liable to be erroneous, though great pains have been taken

to make it as correct as possible.

Table XLVIII. (Appendix, pages 616 and 617) contains the variation of the altitude of an object arising from a change of 100 seconds in the declination.

Most of the tables of this collection have been republished in London in several editions of a work having the following title: "The Improved Practical Navigator, originally written and calculated by NATHANIEL BOWDITCH; revised, recalculated, and newly arranged, by Thomas Kirry." But a number of mistakes have been made in printing the Tables of Mr. Kirby's first edition, some of which have been taken notice of by Dr. Mackay, in the preface of his "Complete Navigator;" and as the manner in which those mistakes are mentioned might lead the reader to suppose that the same errors existed in the American Tables, it is thought proper explicitly to state, that not one of the "many errors and contradictions," Doctor Mackay has mentioned, is to be found therein.

It may be observed that the first method of working double altitudes, given in page 133, is an improvement of a method published by Mr. Ivory.

CONTENTS.

Ph.	me.l		agu.
SIGNS and Abbreviations used in	-	to a degree of longitude at every	
this work	X	degree of latitude	164
Decimal Arithmetic		Questions to exercise the learner	
The state of the s	100	in parallel sailing	465
Demonstration of the most useful	800		Gir
Demonstration of the most useful	92	Middle latitude sailing	COL.
propositions of geometry		Theorems in middle latitude sail-	-
Demonstration of theorems in	100	ing	67
plane trigonometry		Table of solutions of the several	
Geometrical problems	17	cases of middle latitude sailing	08
Construction of the plane scale	24	Questions to exercise the learner	-
Description of Gunter's scale	21	in middle latitude sailing	79,7
Description and use of the sliding	100	Mercator's sailing	77
rule	24	To find the meridianal parts cor-	
Description and use of the sector	26	responding to any degree and	- 300
To find the logarithm of any num-	GG.	minute	77
ber and the contrary	08	Table of solutions of the various	
Multiplication by logarithms	31	cases of Mercator's sailing	75.
Division by logarithms	81		
Involution by logarithms	52		
Evolution by logarithms	82		85
The rule of three by logarithms	52	Construction and use of Merca-	
To calculate compound inferest		tor's chart	TO
by logarithms	33	Of the log-line and half-minute	
To find the log. sine, tangent, &c.	64.3	glass	
corresponding to any number		Description and use of a quadrant	
of degrees and minutes	88		EIX
To find the degrees, minutes, and	-	THE RESERVE OF THE PERSON NAMED IN COLUMN TWO IS NOT THE PERSON NAMED IN COLUMN TWO IS NAMED IN COLUMN TW	00
The state of the s		To adjust a quadrant	-
seconds corresponding to any	34	To take an altitude by a fore ob-	
log. sine, tangent, &c.	2.4	servation	
To find the arithmetical comple-	01	To take the sun's altitude by a	
ment of any logarithm	34	The country of the co	92
Plane trigonometry	35	THE REAL PROPERTY AND ADDRESS OF THE PARTY O	
Table of solutions of the various		a quadrant	
cases of trigonometry		Description and use of a sextant	
Right-angled plane trigonometry	.37		0.0
Questions to exercise the learner in		To adjust a sextant	105
right-angled plane trigonometry		To measure the angular distance	
Oblique trigonometry	40	Control of the Contro	107
A short introduction to astronomy	м	To measure the angular distance	
and geography	144		87
Explanations of the ferms used in	100	Verification of the mirrors and co-	
astronomy and geography	97	Joured glasses	97
Examples in geography	50	parameter and annual parameter and annual	
Plane Sailing	23	of reflection	90
A table of the angles which every		Adjustments of the circle of reflec-	
point of the compass makes		tion	100
with the meridian	53	To observe the meridian altitude	
A table of solutions of the several		of an object by a circle	102
cases of plane sailing	58	To measure the angular distance	
Questions to exercise the learner		of the sun from the moon by a	
in plane sailing	-59	cirele-	103
Traverse sailing	59	To measure the angular distance of	
Parallel sailing	63		
Theorems for solving the several		Verification of the mirrors and	
cases of parallel miling	64		106
A table showing bow many miles	-	On parallax, refraction, and dir	1
of meridian distance correspond		of the horizon	3.07

			•	•
	p	age.	1 . ا	Page:
	To find the distance of the land in		Second method of finding the ap-	
-	order to calculate the dip	109		156
40.5				
T			Third method of finding the appa-	
	Variation of the compass	111	_	158
	To observe an amplitude or azi-		To find the apparent time by an	
	muth by the compass	112	altitude of a fixed star	158
	 To calculate the true amplitude 	112	To regulate a watch by equal alti-	-
	To calculate the true azimuth	119		160
	. Questions to exercise the learner		To find the longitude at sea by lu-	
	in calculating an azimuth	114	nar observations	162
	Having the true and magnetic am-			
			Method of finding the stars used	
•	plitude or azimuth to find the		in lunar observations	163
34)	variation	114	General observations on the tak-	
	To calculate the variation by azi-		ing a lunar observation	164
	muths observed at equal alti-		To work a lunar observation	166
	tudes before and after passing		Examples of lunar observations	168
•	the meridian	115	Second method of working a lu-	
	Variation observed	117	nar observation	174
lui V			Witchell's improved method of	
	To find the letitude has a mortiline	110		
	To find the latitude by a meridian	100	finding the true distance	175
	annuae of the sun of fixed star	120	Method of taking a lunar observa-	
-CM	To find the time of the moon's		tion when you have only one	
	passing the meridian	123	observer	176
	To find the moon's declination;	124	To calculate the sun's altitude at	;
189	To find the latitude by the moon's	1	any time	178
	meridian altitude	125	To calculate the altitude of any	
	To find the latitude by the meri-		star	178
1.55	dian altitude of a planet	107	To calculate the altitude of the	
		121		
	To find the latitude by double al-		moon'	180
1000	titudes		To find the longitude by the eclip-	
1	of the sun	128	ses of Jupiter's satellites	181
	— of a star	129	To find the longitude by an eclipse	
1000	of a planet	129	of the moon	181
175	of the moon	130	To find the longitude by a time-	
-	- of two different objects,	- 1	keeper or chronometer	182
	taken within a few minutes of	!	To regulate a chronometer by lu-	
15 mil	The state of the s	130	nar observations	185
				100
400	of two different objects.		To find the longitude by a varia-	
		131	tion chart	185
			Problems useful in navigation	186
	Second method	138	To find the difference between the	
	Third method	142	true and apparent directions of	•
•	Questions to exercise the learner		the wind	191
28		148	To determine the height of a	
24	To find the latitude by one alti-		mountain by barometers	192
100	tude of the sun, having your	- 1	Mensuration	193
				197
-	The state of the s	,	Gauging	-
2.75	To find the latitude by the mean	1.		·199
	of several altitudes of the sun	· [To find the content of a field by	•
	taken near noon by a sextant or	- 1	the table of diff. lat. and depart-	٠.
	circle	150	ure	201
130	To find the latitude on shore by	ŀ	To survey a coast in sailing along	
	means of an artificial horizon		shore	203
15	To find the latitude by the polar		To survey a harbour by observa-	7.7
1		158	tions on shore	205
.3	The state of the s			#UV
13	To find the time at sea and regu-		Methods of surveying a small	,
-3	THE RESERVE OF THE PARTY OF THE	154	bank or shoal where great accu-	000
w	Examples to exercise the learner		racy is required	206

P	age.	: Р	age.
To reduce soundings taken at any.	U	Problem VII. To calculate the	. 0 -
time of the tide to low water	209	longitude of a place from the	
To reduce a draught to a smaller		observed beginning and end of	
scale	210		586
Of winds	210	Problem VIII. To find the longi-	
Directions for sailing from Ameri-		tude of a place from the begin-	
ca to India	212	ning or end of a solar eclipse	589
Tides	213	Problem IX. To find the longi-	
To find the time of high water by	٠,	tude of a place from the begin-	
a Nautical Almanac	214	ning or end of an occultation	590
To find the time of high water by		Problem X. To project an eclipse	
the tables C and D	216	of the moon	591
Tables for calculating the time of		Problem XI. To project an eclipse	
high water	217	of the sun .	595
Currents	218	Problem XII. To project an oc-	
Gulf Stream	218	cultation of a fixed star	598
Method of keeping a reckoning at		Problem XIII. To calculate the	
sea	219	beginning or end of an eclipse	
To find the lee way and allow for it		or occultation	601
To correct the dead reckoning	228	Problem XIV. To find the appa-	
Rules for working a day's work	225	rent time at Greenwich from	
Examples for working a day's		the moon's longitude	602
work	227	Problem XV. To find the longi-	
Journal from Boston to Madeira	231	tude of a place by measuring	
Explanations of sea terms	249	the distance of the moon from	
Evolutions at sea	264	a fixed star not marked in the	, .
Catalogue of the Tables, with ex-		Nautical Almanac	603
amples of the uses of those not	•	Problem XVI. To find the longi-	000
explained in other parts of the		tude of a place by the moon's	
work	279	passage over the meridian	60 4
Tables from I. to XLVIII.		Problem XVII. Given the lati-	001
10000 1000 1000 1000 1000		tude of the moon and longi-	
APPENDIX.		tudes of the moon and sun to	
Addition and subtraction, using		find their angular distance	605
the signs as in algebra	571	Problem XVIII. Given the longi-	•••
Problem I. To find the longitude,		tudes and latitudes of the moon	
latitude, &c. of the moon	571	and a star to find their angular	
Problem II. To find the horary		distance	600
motion of the moon	574	Problem XIX. Given the right	0.70
Problem III. To find the eclip-		ascension and declination to	
tic conjunction or opposition of		find the longitude and latitude	606
the moon and sun, or a star	575	Problem XX. Given the longi-	•••
Problem IV. To find the altitude	-	tude and latitude to find the	
and longitude of the nonagesi-		right ascension and declination	607
mal	577	Spheric Trigonometry	607
Table to facilitate the calcula-		Improvement of Napier's rules	•••
tion	578		608
Abridged rule for calculating the		Theorems in Spherics	609
altitude and longitude of the		On finding the latitudes by two	5 5.,
nonagesimal	579		611
Problem V. To calculate the		Examples for exercise	614
moon's parallax in latitude and		Table shewing the variation of the	-
longitude	580		
Problem VI. To calculate the lon-		from a change of 100 seconds	٠.
gitude of a place from the ob-		in its declination	616
served beginning and end of a	_	Table XX. (New Form) correc	
solor actions	585		613

ARRANGEMENT OF THE TABLES.

Table.	Table.
DIFFERENCE of latitude	To find the time of the moon's
and departure for points I	passing the meridian XXVIII
Ditto, for degrees II	Correction of the moon's alti-
Meridional parts III	tude for parallax and refrac-
Sun's declination IV	tion XXIX
23 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4 4	To find the moon's declina-
For reducing the Sun's declina-	
	To find the sun's right ascen-
Sun's right-ascension VI	sion XXXI
Correction for the daily varia-	Variation of the sun's altitude
tion of the Equation of Time VI	in one minute from noon XXXII
Amplitudes VII	To reduce the numbers of Table
Right ascensions and declina-	XXXII. to other given inter-
tions of the fixed stars VIII	vals from noon XXXIII
Sun's rising and setting IX	Errors arising from a deviation
For finding the distance of ter-	of one minute in the parallel-
restrial objects at sea X	ism of the surfaces of the cen-
Proportional parts XI	tral mirror XXXIV
Refraction of the heavenly bo-	Error arising from a deviation .
dies XII	of the telescope from a plane
Dip of the horizon XIII	parallel to the plane of the in-
Sun's parallax in altitude XIV	strument XXXV
Augmentation of the moon's	Correction of the mean refrac-
semidiameter XV	tion for various heights of the
Dip for different heights and	thermometer and barome-
	ter . XXXVI
	Longitudes and Latitudes of
rithm of a lunar observation when a star is used XVII	
	Reduction of latitude and hori-
To find the correction and loga-	zontal parallax XXXVIII
rithm of a lunar observation	Aberration of the planets in lon-
when the sun is used XVIII	gitude XXXIX
To find the correction and loga-	Equation of the equinoxes in
rithm of a lunar observation	longitude
depending on the moon's al-	Aberration of the fixed stars in
titude XIX	latitude and longitude XLI
For finding the third correction	Aberration of the fixed stars in
of a lunar observation XX	right ascension and declina-
(New Form) Corrections in se-	tion' XLII
conds additive. See Append-	Nutation in right ascension and
ix, page 818 XX	declination XLIII
For turning degrees and min-	Augmentation of the moon's
utes into time, and the con-	semidiameter, found by nona-
trary XXI	gesimal XLIV
Proportional logarithms XXII	Equation of second differences XLV
For finding the latitude by two	Table of latitudes and longi-
altitudes of the sun XXIII	tudes • XLVI
Natural sines and co-sines XXIV	Tide Table XLVII
Log. sines, tangents, &c. to	Table, shewing the variation
points and quarter points XXV	of the altitude of an object
Logarithms of numbers XXVI	arising from a change of 100
Logarithmic sines, tangents, .	seconds in its declination,
and seconds is VVIII	(A

SIGNS AND ABBREVIATIONS USED IN THIS WORK.

+ Is the sign of addition, and denotes that whatever number or quantity follows the sign, must be added to those that go before it, thus 9+8 signifies that 8 is to be added to 9. Or A+B implies that the quantities represented by A and B are to be added. The sign + is called the positive sign.

The sign of subtraction; and denotes that the number following it must, be subtracted from those going before it, thus 7—5, signifies that 5 must

be subtracted from 7. The sign — is called the negative sign.

X Is the sign of multiplication, and shows that the numbers placed before and after it are to be multiplied, thus, 7×9 signifies 7 multiplied by 9, which makes 63; and 7×6×2 signifies the continued product of 7 by 8 and by 2, which makes 112. Multiplication is also denoted by placing a point between the quantities to be multiplied; thus A.B signifies that A is to be multiplied by B.

is the sign of division, and signifies that the number that stands before it is to be divided by the number following it, as 72 + 12 shows that 72 is to be divided by 12. Division may also be denoted by placing two points between the numbers, thus, 72:12 represents 72 divided by 12

points between the numbers, thus, 72: 12 represents 72 divided by 12

or by placing the numbers thus, — which signifies 72 divided by 12.

() or _____. Either of these marks is used for connecting numbers to-

gether, thus, $3+4\times 6$, or $(3+4)\times 6$, signifies that the sum of 3 and 4 is to be multiplied by 6.

= Is the sign of equality, and shows that the numbers or quantities placed before it are equal to those following it: thus 8 × 12=96. Or 8 mul-

tiplied by 12 are equal to 96, and $7+2\times4=98$.

:::: Is the sign of proportion, and is marked thus, 7:14::10:20, that is, as 7 is to 14, so is 10 to 20. Or A:B:: C:D, that is, as A is to B, so is C to D.

Signifies degrees; thus, 45° represents 45 degrees.

' Signifies minutes; thus, 24' or 24 minutes,

" Signifies seconds; thus, 44", or 44 seconds.

Signifies thirds or sixtieth parts of seconds; thus, 44", or 44 thirds.

S. Signifies sine. N. S. Signifies Natural sine.

Sec. Signifies Secant.

Tan Signifies Tangent.

Co-sine, Co-tangent, or Co-secant of an arch signifies the sine, tangent or secant of the complement of that arch respectively.

Signifies Angle; with an s at top Angles, < ...</p>

 \wedge d Angled.

△ Signifies Triangle. △'s Triangles.

Signifies a square.

or The Sun. O or I the Moon. * a Star. L. L. Lower Limb.
U. L. Upper Limb. N. L. Nearest Limb. S. D. Semi-diameter.
P. L. Proportional Logarithm. N. A. Nautical Almanac. Z. D.
Zenith Distance. D. R. Dead Reckoning.

DIRECTIONS FOR THE BINDER.

		•
PLATE	I.	TO FRONT THE TITLE PAGE
	II.	TO FRONT PAGE 20.
	III.	TO FRONT PAGE 44.
	IV.	TO FRONT PAGE 46.
	v.	TO FRONT PAGE 50.
	VI.	TO FRONT PAGE 52.
	VII.	TO FRONT PAGE 96.
	VIII.	TO FRONT PAGE 106.
	IX.	TO FRONT PAGE 110.
	Χ.	TO FRONT PAGE 203.
	XI.	TO FRONT PAGE 205.
	XII	TO FRONT PAGE 582.

DECIMAL ARITHMETIC.

MANY persons who have acquired considerable skill in common Arithmetic, are unacquainted with the method of calculating by decimals, which is of great use in Navigation; for which reason it was thought proper to

prefix the following brief explanation.

Fractions or Vulgar Fractions are expressions for any assignable part of an unit; they are usually denoted by two numbers, placed the one above the other, with a line between them: thus, 4 denotes the fraction one-fourth, or one part out of four of some whole quantity, considered as divisible into four equal parts. The lower number 4 is called the denominator of the fraction, showing into how many parts the whole or integer is divided; and the upper number 1, is called the numerator, and shows how many of those equal parts are contained in the fraction. And it is evident that if the numerator and denominator be varied in the same ratio, the value of the fraction will remain unaltered: thus if the numerator and denominator of the fraction 4 be multiplied by 2, 3, or 4, &c. the fractions arising will be $\frac{2}{4}$, $\frac{1}{12}$, $\frac{1}{13}$, &c, which are evidently equal to $\frac{1}{4}$.

Decimal Fraction is a fraction whose denominator is always an unit with some number of ciphers annexed, the numerators of which may be any numbers whatever; as $\frac{7}{45}$, $\frac{7}{16}$, $\frac{7}{16}$, $\frac{7}{16}$, $\frac{7}{16}$. And as the denominator of a decimal is always one of the numbers 10, 1000, &c. the inconvenience of writing these denominators may be avoided, by placing a point between the integral and the fractional part of the number; thus $\frac{7}{16}$ is written .5; and $\frac{7}{16}$ is written .14; the mixed number $\frac{7}{16}$, consisting of whole num-

bers and fractional ones is written 5.14.

In setting down a decimal fraction, the numerator must consist of as many places as there are ciphers in the denominator; and if it has not so many figures the defect must be supplied by placing ciphers before them; thus, $\frac{16}{160}$, =.16, $\frac{1}{1000}$ =.016, $\frac{1}{1000}$ =.0016, &c. And as ciphers on the right hand side of integers increase their value in a tenfold proportion, as 2, 20, 200, &c. so when set on the left hand of decimal fractions, they decrease their value in a tenfold proportion, as 2, .02, .002, &c. but ciphers set on the right hand of these fractions make no alteration in their value, neither of increase or decrease; thus, .2 is the same as .20 or .200. The common arithmetical operations are performed the same way in decimals, as they are integers; regard being had only to the particular notation, to distinguish the integral from the fractional part of a sum.

ADDITION OF DECIMALS.

Addition of decimals is performed exactly like that of whole numbers, placing the numbers of the same denomination under each other, in which case the decimal separating points will range straight in one column.

		EXAMPLES.	
	Miles.	Feet.	Inches.
	26.7	1.26	272.3267
	32.15	2.31	.0134
	143.206	1.785	2.1576
	.003	2.0	31.4
	-	-	
Sum	202.059	7.85 5	305.8977

SUBTRACTION OF DECIMALS.

Subtraction of decimals is performed in the same manner as in whole numbers, by observing to set the figures of the same denomination and the ser parating points directly under each other.

	 EXAMPLE 	S.	
From 31.267	36.75	1.254	1364.2
Take 2.63	.026	.316	25.163
	-		
Diff. 28.637	36.724	.938	. 1339.037

MULTIPLICATION OF DECIMALS.

Multiply the numbers together the same as if they were whole numbers, and point of as many decimals from the right hand as there are decimals in both factors together; and when it happens that there are not so many figures in the product as there must be decimals, then prefix as many ciphers to the left hand as will supply the defect.

EXAMPLE I.	EXAMPLE IV, Multiply .17 by .06	
Multiply 3.25 by 4.5	.17	
3.25	.06	
4.5		
	Answer .0102	
1.625	In each of the factors are two decimals	
13.00	the product ought therefore to contain 4	
	and there being only three figures in the	
Answer 14.625	product I prefix a cipher.	
In one of the factors is one decimal and	EXAMPLE V. Multiply .18 by 24.	
in the other two, their sum 3 is the number	.18	
of decimals of the product.	24	
EXAMPLE II.		
Multiply 0.5 by 0.7	72	
0.5	36	
0.7	Answer 4.32	
	EXAMPLE VI. Multiply 36.1 by 2.5	
0.35 A nswer.	36.1	
EXAMPLE III.	2.5	
Multiply 3.25 by .05		
3.25	18.05	
.05	72.2	

DIVISION OF DECIMALS.

Division of decimals is performed in the same manner as in whole numbers; only observing that the number of decimals in the quotient must be equal to the excess of the number of decimals of the dividend above those of the divisor.—When the divisor contains more decimals than the dividend, ciphers must be affixed to the right hand of the latter to make the number equal or exceed that of the divisor.

Divide 14.625 by 3.25 3.25)14.625(4.5 1300
1625 1625

In this example there are 2 decimals in the divisor, and 3 in the dividend, hence there is one decimal in the quotient.

.1625 Product.

EXAMPLE II. Divide 0.35 by 0.7 .7).35(.5 .35

EXAMPLE III. Divide 3.1 by .0062

Previous to the division I affix a number of ciphers to the right hand of 3.1, which does not alter its value.

90.25

.0062)3.100000(500.00 310

0000 Therefore the answer is 500.00 or 500.

Digitized by

EXAMPLE IV. Divide 9.6 by .06

160 Answer.

Here by affixing a cipher to 9.6 it becomes 9.60, and has then 2 decimals in it, which is the same number as is in the divisor, therefore the quotient is an integer number.

EXAMPLE V. Divide 17.256 by 1.16 1.16)17.25600(14.875

116 565

> 464 1016 928

REDUCTION OF DECIMALS.

If you wish to reduce a vulgar fraction to a decimal, you may add any number of ciphers to the numerator, and divide it by the denominator, the quotient will be the decimal fraction; the decimal point must be so placed that there may be as many figures to the right hand of it as you added ciphers to the numerator; if there are not as many figures in the quotient, you must place ciphers to the left hand to make up the number.

EXAMPLE I. Reduce 1 to a decimal.

2 Answer

EXAMPLE II. Reduce ? to a decimal.

.375 Answer.

EXAMPLE III. Reduce 3 inches to the decimal of a foot.

Since 12 inches=1 foot, this fraction is 3.

12)3.00

.25 Answer.

EXAMPLE IV. Reduce 34 inches to the decimal of a foot.

 $3\frac{1}{2} = \frac{7}{4}$: this divided by 12 is $\frac{7}{4}$. 24)7.000(.291 Answer.

48 220

216

24

EXAMPLE V. Reduce 1 foot and 6 inches to the decimal of a yard. Here 1 foot 6 inches=18 inches. And 1 yard=36 inches, therefore this fraction is \(\frac{1}{3} \frac{8}{6} \).
36) 180(5 Answer.

If you have any decimal fraction, it is easy to find its value in the lower denominations of the same quantity; thus if the fraction was the decimal of a yard, by multiplying it by 3 we have its value in feet and parts; if we multiply this by 12, the product is its value in inches and parts; and in the same manner the values may be obtained in other cases.

EXAMPLE VI.

Required the value of 3.25 yards.

3.25 12

9.00

Answer S yards, 0 feet, 9 inches.

EXAMPLE VII.

Required the value of 7.231 days.

7.231 24 924

38.400 Answer 7 days, 5 hours, 52 minutes, . and 38 seconds.

GEOMETRY.

GEOMETRY is the Science which treats of the description, properties. and relations of magnitudes in general, of which there are three kinds or species, viz. a line which has only length without either breadth or thickness; a superfices, comprehended by length and breadth, and a solid, which has length, breadth, and thickness.

A POINT considered mathematically has no length, breadth, or thickness.

A STRAIGHT LINE OR RIGHT LINE is the shortest distance between the two

points which limits its length, as

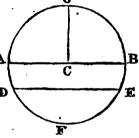
A PLANE SUPERFICES is that in which any two points being taken, the straight line between them lies wholly in that surface.

PARALLEL LINES are such as are in the same plane and which extended infinitely do never meet, as AB, DC.

A CIRCLE is a plane figure, bounded by an uniform curve line; it is commonly described with a pair of compasses; one point of which is fixed, whilst the other is turned round to the place where the motion first began; the fixed point is called the CENTRE, and the line described by the other point is called the circumference.

The RADIUS of a circle, or SEMIDIAMETER, is a right line drawn from the centre to the circumference, as AC; or it is that line which is taken between the points of the compasses to describe the circle.

A DIAMETER of a circle is a right line drawn through the centre and terminated at both ends p by the circumference, as ACB, and is the double of the radius AC. A diameter divides the circle, and its circumference into two equal VII.



An ARCH of a circle is any part or portion of the circumference, as DFE. VIII.

The CHORD of an arch is a straight line joining the ends of the arch; it divides the circle into two unequal parts, called SEGMENTS, and is a chord to them both, as DE is the chord of the arches DFE and DGE. IX.

A SEMICIRCLE, or half circle, is a figure contained under a diameter and the arch terminated by that diameter, as AGB or AFB. Any part of a circle contained between two radii and an arch, is called a SECTOR.

A QUADRANT is half-a semicircle, or one-fourth part of a whole circle, as the figure CAG.

NOTE. All circles, whether great or small, are supposed to have their circumference divided into 360 equal parts, called degrees, and each degree into 60 equal parts, called minutes; and each minute into 60 equal parts, called seconds, and so on into thirds, fourths,* &c. and an arch is said to be of as . many degrees as it contains parts of the 360, into which the circumference is divided.

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^{*} A new division of the circumférence of the circle has lately been adopted by several eminent French mathematicians, in which the quadrant is divided into 100°, each degree into 100°, each minute into 100°, dec and tables of logarithms have been published conformable thereto. The general adoption of this division would tend greatly to facilitate most of the calculations of navigation and

XI.

An ANGLE is the inclination of two lines which meet, but not in the same direction.

An angle is usually expressed by the letter placed at the angular point, as the angle A. But when two or more angles are at the same point, it is then necessary to express each by three letters, and the letter at the angular point is placed between the two. Thus, the angle formed by the lines AB, AC, is call-A ed the angle BAC or CAB, and that formed by AB, AD, is called the angle BAD, or DAB.

An angle is measured by the arch of a circle comprehended between the two

legs that form the angle, the centre of the circle being the angular point.

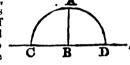
Thus the angle A is measured by the arch BC described round the point A as a centre, and the angle is said to be of as many degrees as the arch is, that is, if the arch BC is 50°, then the angle BAC, is said to be an angle of 50 A degrees.



XII.

If a right line AB, fall upon another DC, so as to incline neither to the one

side nor the other, but makes the angles ABC, ABD, equal to each other; then the line AB is said to be perpendicular to the line DC, and each of these angles is called a right angle, being each equal to a quadrant or 90°; because the sum of the two angles ABC, ABD, is measured by the semicircle DAC, described on the diameter DBC, and centre B.



B

XII

An Acute Angle is less than a right angle, as ABC.

XIV.

An OBTUSE ANGLE is greater than a right

angle, as GEH.

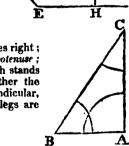
The least number of right lines that can include a space, are three which form a finelude a space, are three which some six and the space of the

gure called a *Triangle*, consisting of six parts, viz. three sides and three angles: it is distinguished into three sorts, viz. a right angled triangle, an obtuse-angled triangle,

and an acute angled triangle.

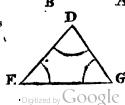
YV

A RIGHT RELED TRIANGLE has one of its angles right; the side opposite the right angle is called the hypotenuse; and the other two sides are called legs; that which stands upright, is called the perpendicular, and the other the base; thus BC is the hypotenuse, AC the perpendicular, and AB the base; the angles opposite the two legs are both acute.



XVI.

An Acute-angles Triangle has each of its angles acute, as DEG.



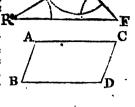
XVII.

An Obtuse-Angled Triangle has one of its angles obtuse, or greater than a right angle, as RAF; the other two angles are acute.

RAF; the other two angles are acute.

NOTE. All triangles that are not right angled, whether they are acute or obtuse, are in general terms called oblique-angled triangles, without any other distinction.

A QUADRILATERAL figure is one bounded by four sides, as ACDB. If the opposite sides are parallel they are called Parallelograms. Thus if AC be parallel to BD, and AB parallel to CD, the figure ACDB is a parallelogram. A parallelogram having all its sides equal, and its angles right, is called a SQUARE, as B. When the angles are right, and the opposite sides only equal, it is called a RECTANGLE, as A.



 ${f B}$

XIX.

The SINE of an arch is a line drawn from one end of the arch perpendicular to a diameter drawn through the other end of the same arch; thus RS is the sine of the arch AS, RS being a line drawn from one end of that arch, perpendicular to DA which is the diameter passing through the other end A of the arch.

XX.

The Co-SINE of an arch is the sine of the complement of that arch, or of what that arch wants of a quadrant; D thus AH being a quadrant, the arch SH is the complement of the arch AS; SZ is the sine of the arch SH, or the co-sine of the arch AS.

H Cotangent of the diameter contained be-

XXI.

The Versed Sine of an arch is that part of the diameter contained between the sine and the arch; thus RA is the versed sine of the arch AS, and DCR is the versed sine of the arch DHS.

XXII.

The Tangent of an arch is a right line drawn perpendicular to the diameter passing through one end of the arch, and terminated by a line drawn from the centre through the other end of the arch; thus AT is the tangent of the arch AS.

XXIII.

The Co-tangent of an arch is the tangent of the complement of that arch to a quadrant; thus HG is the tangent of the arch HS or the co-tangent of the arch AS.

XXIV.

The SECANT of an arch is a right line drawn from the centre through one end of the arch to meet the tangent drawn from the other end; thus CT is the secant of the arch AS.

XXV.

The Co-secant of an arch is the secant of the complement of that arch to a quadrant, thus CG is the secant of the arch SH, or co-secant of the arch AS. XXVI.

What an arch wants of a semicircle is called the Supplement of the arch, thus, the arch DHS is the supplement of the arch AS. The sine, tangent, or secant of an arch, is the same as the sine, tangent, or secant of its supplement; thus, the sine of 80°=sine of 100°, and the sine of 70°=sine of 110°, &c.

XXVII. If one line AB fall any way upon another CD, the sum of the two angles ABD, ABC is always equal to

two right angles.

For on the point B as a centre, describe the circular arch CAD, cutting the line CD in C and D; then (by art. 6) this arch is equal to a semicircle, but it is also equal to the sum of the arches CA and AD, the measures of the two angles ABC, ABD; therefore the sum of the two angles is equal to a semicircle, or two right angles. Hence it is evident that all the angles which can be made from a point in any line, towards one side of the line, are equal to two right angles, and that all the angles which can be made about a point, are equal to four right angles.

XXVIII.

If a line AC cross another BD in the point E, the opposite angles will be equal, viz. BEA=CED, and BEC=AED.

Upon the point E as a centre, describe the circle ABCD; then it is evident that ABC is a semicircle, as also BCD (by the 6th) therefore the arch ABC = $\operatorname{arch} \mathbf{B}$ BCD, taking from both the common arch BC, there remains AB=CD, that is, the angle BEA equal to the angle CED. After the same manner we may prove that the angle BEC is equal to the angle AED.

XXIX. If a line GH cross two parallel lines, AB, CD, it makes the external opposite angles equal to each other, viz. GEB=CFH and AEG=HFD.

For since AB and CD are parallel to each other, they may be considered as one broad line, and GH crossing it; then the A vertical or opposite angles GEB, CFHare equal (by art. 28) as also AEG=HFD.

 \mathbf{B} ${f n}$

If a line GH cross two parallel lines AB, CD (see the figure) the alternate angles AEF and EFD, or CFE and FEB are equal.

For GEB=AEF (art. 28) as also CFH=EFD (by the same art.) but GEB=CFH by the last. Therefore AEF is equal to EFD; in the same way may we prove FEB=CFE.

XXXI. If a line GH cross two parallel lines AB, CD (see the preceding figure) the external angle GEB is equal to the internal opposite one EFD, or AEG equal to CFE.

For the angle AEF is equal to the angle EFD by the last, and AEF=GEB (by art. 28) therefore GEB=EFD; in the same way we may prove AEG=CFE. XXXII.

If a line GH cross two parallel lines AB, CD (see the preceding figure) the sum of the two internal angles BEF and DFE, or AEF and CFE is equal to two right angles.

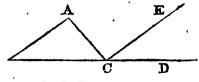
For since the angle GEB is equal to the angle EFD (by the last) to both

add the angle BEF, and we have GEB+BEF=BEF+EFD, but GEB+BEF=two right angles. (art. 27.) Hence BEF+EFD=two right angles; and in the same manner we may prove AEF+CFE=two right angles.

XXXIII.

In any triangle ABC, one of its legs, as BC being produced towards D, the external angle ACD is equal to the sum of the internal and opposite angles ABC, BAC.

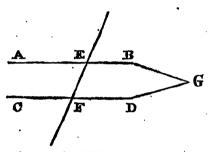
To prove this, through C draw
CE parallel to AB; then since
CE is parallel to AB and the
lines AC, BD cross them, the angle ECD=ABC (by article 31)
and ACE=BAC (by article 30) B
adding these together we have
ECD+ACE=ABC+BAC; but
ECD+ACE=ACD, therefore AC



ECD+ACE=ACD, therefore ACD=ABC+BAC.
XXXIV.

Hence it may be proved that if any two lines AB and CD, be crossed by a third line EF, and the alternate angles AEF and EFD be equal, the lines AB and CD will be parallel.

For if they are not parallel, they must meet each other on one side of the line EF (suppose at G) and so form the triangle EGF, one of whose sides, GE being produced to A, the exterior angle AEF must (by the preceding article) be equal to the sum of the two angles EFG and EGF; but by supposition it is equal to the angle EFG alone; therefore the angle AEF must be equal to the sum of the two angles are the sum of the two angles are supposition.



gles EFG and EGF, and at the same time equal to EFG alone, which is absurd; therefore the lines AB, CD cannot meet, and must be parallel.

XXXV.

In any right lined triangle ABC, the sum of the three angles is equal to two

right angles.

To prove this, you must produce BC (in the fig. art. 33,) towards D, then (by art. 35) the external angle ACD=ABC+BAC, to both add the angle ACB and we have ACD+ACB=ABC+BAC+ACB, but ACD+ACB=two right angles (by art. 27.) Hence ABC+BAC+ACB=two right angles; therefore the sum of the three angles of any plain triangle ACB is equal to two right angles.

XXXVI.

Hence ist any plain triangle, if one of its angles be known, the sum of the other two will be also known.

For by the last article the sum of all three angles is equal to two right angles or 180°, hence, by subtracting the given angle from 180°, the remainder

will be the sum of the other two.

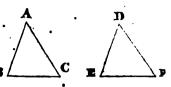
In any right angled triangle, the two acute angles taken together are just equal to a right angle: for all three angles being equal to two right angles, and one angle being right by supposition, the sum of the other two must be equal to a right angle, consequently any one of the acute angles being given, the other one may be found, by subtracting the given one from 90 degrees.

XXXVII.

If in any two triangles ABC, DEF, two legs of the one, AB, AC, be equal to two legs of the other DE, DF, each to each respectively, that is AB=DE and

AC=DF, and the angles BAC, EDF included between the equal legs be equal; then the remaining leg of the one will be equal to the remaining leg of the other, and the angles opposite to the equal legs will be equal, that is, BC=EF, ABC=DEF, and ACB=DFE.

For if the triangle ABC be supposed to be lifted up and put upon the triangle DEF, with the point A on the point D and the line AB upon DE; it is plain, since AB=DE, that the point B will fall upon E, and since the angles BAC, EDF are equal, the line AC will B



fall upon DF, and these lines being of equal length, the point C will fall upon F, consequently the line BC will fall exactly upon the line EF, and the triangle ABC will in all respects be exactly equal to the triangle DEF, and the angle ABC will be equal to the angle DEF, also the angle ACB will be equal to the angle DEF.

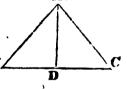
DEF, also the angle ACB will be equal to the a

After the same manner it may be proved that if in any two triangles ABC; DEF (see the preceding figure) two angles ABC and ACB of the one, be equal to two angles DEF, DFE of the other, and the included side BC be equal to EF; the remaining sides and included angles will also be equal to each other respectively, that is, AB=DE, AC=DF, and the angle BAC=angle EDF.

For if the triangle ABC be supposed to be lifted up and laid upon the triangle DEF, the point B being upon the point E, and the line BC upon the line EF; then since BC=EF the point C will fall upon the point F, and since the angle ACB=the angle DFE, the line CA will fall upon the line FD, and by the same way of reasoning, the line BA will fall upon the line ED, therefore the point of intersection A of the two lines BA, CA, will fall upon D, the point of intersection of the lines ED, FD, consequently AB=DE, AC=DF, and the angle BAC=the angle EDF.

If two sides of a triangle are equal, the angles opposite these sides will also be equal; that is, if AB=AC, the angles ABC, ACB will also be equal. For draw the line AD bisecting the angle BAC, and meeting the line BC in D, dividing the triangle

and meeting the line BC in D, dividing the triangle BAC into two triangles ABD, ACD, in which the B side AB = AC, the side AD is common to both triangles, and the angle BAD = the angle DAC; con-



sequently (by art. 37) the augle ABD must be equal to the angle ACD.

The converse of this proposition is also true; that is, if two angles of a triangle are equal, the opposite sides are also equal.

This is demonstrated nearly in the same manner by means of art. 38.

Any angle at the circumference of a circle is half the angle at the centre

standing upon the same arch.

Thus, the angle BAD is half the angle BCD standing upon the same arch BD of the circle BEDA, whose centre is C. To demonstrate this, draw through A and the centre C the right line ACE, then (by art. 85) the angle CAD+angle CDA=angle ECD, but AC=CD (being two radii of the same circle) therefore (by art. 89) the angle CAD=the angle CDA, and the sum of these two angles is the double of either of them, that is, CAD+CDA=twice

CAD, therefore ECD=twice CAD; in the same manner it may be proved that BCR=twice BAC, and by adding these together, we have ECD+BCE=twice CAD+twice BAC, that is, BCD=twice BAD, or BAD equal to half of BCD. The demonstration is similar when B, D, fall on the same

side of E.

XLI.

An angle at the circumference is measured by half the arch it subtends.

For an angle at the centre standing on the same arch is measured by the whole arch (by art. 11); but since an angle at the centre is double that at the circumference, (art. 40) it is evident that an angle at the circumference must be measured by half the arch it stands upon. Hence all angles ACB, ADB, AEB, &c. at the circumference of a circle standing on the same chord AB are equal to each other; for they are all measured by the same arch, viz. half the arch AB.



XLII.

An angle in a segment greater than a semicircle is less

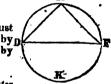
than a right angle.

Thus if ABC be a segment greater than a semicircle, the arch AC on which it stands must be less than a semicircle, and the half of it less than a quadrant or a right angle; but the angle ABC in the segment is measured by the half of the arch AC; therefore it is less than a right angle.



An angle in a semicircle is a right angle.

For since DEF is a semicircle, the arch DKF must also be a semicircle; but the angle DEF is measured by nhalf the arch DKF, that is, by half a semicircle or by a quadrant; therefore the angle DEF is a right one.

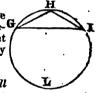


An angle in a segment less than a semicircle is greater

than a right angle.

Thus if GHI be a segment less than a semicircle, the arch GLI on which it stands must be greater than a semicircle, and its half greater than a quadrant or right angle: therefore the angle GHI which is measured by half the arch GLI is greater than a right angle.

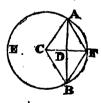
XLIII.



If from the centre C of the circle ABE, there be let fall the perpendicular CD on the chord AB; it will bisect the

chord in the point D.

Draw the radii CA, CB; then (by art. 59) the angle CBA=the angle CAB, and as the angles at D are right, the angle ACD must be equal to the angle BCD (by art. 56.) Hence in the triangles ACD, BCD, we have the angle ACD equal to the angle BCD, CA=CB and CD common to both triangles, consequently (by art. 57) AD=DB, that is, AB is bisected at D.



XLIV

If from the centre C of the circle ABE there be drawn a perpendicular CD, to the chord AB, and it be continued to meet the circle in F, it will bisect the arch AFB in F. (See the preceding figures)

For in the last article it was proved that the angle ACD = the angle BCD,

hence (by art. 11) the arch AF = the arch FB.

XLV.

Any line bisecting a chord at right angles is a diameter.

For since (by art. 43) a line drawn from the centre perpendicular to a chord, bisects that chord at right angles, therefore conversely a line bisecting a chord at right angles, must pass through the centre, and consequently be a diameter.

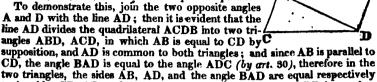
XLVI.

The sine of any arch is equal to half the chord of twice that arch. For (in the last scheme) AD is the sine of the arch AF, and AF is equal to

half the arch AFB and AD half the chord AB, whence the proposition is manifest.

XLVII.

If two equal and parallel lines AB, CD, be joined by A two others, AC, BD, these will be also equal and parallel.



to the sides CD, AD, and the angle ADC; hence (by art. 37) BD is equal to

AC, and the angle DAC equal to the angle ADB; therefore (by art. 34) the lines BD, AC, must be parallel.

Cor. Hence it follows that the quadrilateral ABDC is a parallelogram, since the opposite sides are parallel. It is also evident that in any parallelogram, the line joining the opposite angles (called the diagonal) as AD, divides the figure into two equal parts, since it has been proved that the triangles ABD,

ACD, are equal to each other.

XLVIII.

It follows also from the preceding article, that a triangle ACD (see the preceding figure) on the same base, and between the same parallels with a parallelogram ABDC, is the half of that parallelogram.

From the same article it also follows, that the opposite sides of a parallelogram are equal. For it has been proved, that ABDC being a parallelogram, AB is equal to CD, and AC equal to BD.

All parallelograms on the same or equal bases, and between the same parallels, are equal to each other; that is, if BD and GH be equal, and the lines BH, AF be parallel, the parallelograms ABDC, BDFE and EFHG will be equal to each other.

For AC is equal to EF each being equal to BD (by art. 49) to both add CE and we have AE equal to CF; therefore in the two triangles ABE, CDF; AB is equal to CD, and AE is equal to CF, and the angle BAE is equal to DCF (by art. 51,) therefore B



the two triangles ABE, CDF are equal (by art. 37) and taking the triangle CKE from both, the figure ABKC is equal to the figure KDFE, to both which add the little triangle KBD, and we have the parallelogram ABDC equal to the parallelogram BDFE. In the same way it may be proved that the parallelogram EFHG is equal to the parallelogram BDFE; therefore the three parallelograms ABDC, BDFE, and EFHG are equal to each other.

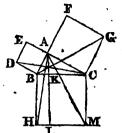
Cor. Hence it follows, that triangles on the same base and between the same parallels are equal, since they are the half of the parallelograms on the same

hase and between the same parallels (by art. 48.)

In any right angled triangle, the square of the hypotenuse is equal to the sum of the squares of the two sides. Thus if BAC be a right angled triangle the square of the hypotenuse BC, viz. BCMH, is equal to the sum of the squares made on the two sides AB and AC, viz. to ABDE and ACGF.

To demonstrate this, through the point A draw AKL perpendicular to the hypotenuse BC. Join AH, AM, DC, and BG; then it is evident, that DB

is equal to BA (by art. 18) and BH equal to BC, therefore in the triangles DBC, ABH, the two legs DB, BC of the one are equal to the two legs AB, BH, of the other; and the included angles DBC and ABH are also equal. (for DBA is equal to CBH being both right, to each add ABC and we have DBC equal to ABH) there-fore the triangles DBC, ABH are equal (by art. 37) but the triangle DBC is half of the square ABDE (by art. 48) and the triangle ABH is half the parallelogram BKLH (by the same art.) consequently the square ABDE is equal to the parallelegram BKLH. In the same way it may be proved that the square ACGF is equal to the parallelogram KCML. Therefore the sum of the squares ABDE and



ACGF is equal to the sum of the parallelograms BKLH and KCML; but the sum of these parallelograms is equal to the square BCMH, therefore the

sum of the squares on AB and AC is equal to the square on BC.

Cor. Hence in any right angled triangle, if we have the hypotenuse and one of the legs, we may easily find the other leg, by taking the square of the given leg from the square of the hypotenuse, the square root of the remainder will be the sought leg. Thus if the hypotenuse was 13, and one leg was 5. the other leg would be 12, for the square of 5 is 25, and the square of 13 is 169, subtracting 25 from 169 leaves 144, the square root of which is 12. both legs are given, the hypotenuse may also be found by extracting the square root of the sum of the squares of the legs; thus if one leg was 6, and the other 8, the square of the first is 36, the square of the second is 64, adding 36 and 64 together gives 100, whose square root is 10, which is the sought hypotenuse.

t.II.

Four quantities are said to be proportional, when the magnitude of the first compared with the second is the same as the magnitude of the third compared

with the fourth.

Thus 4, 8, 12 and 24, are proportional; because 4 is half of 8, and 12 is half of 24; and if we take equi-multiples $A \times a$, $A \times b$, of the quantities a and b, and other equi-multiples $B \times a$, $B \times b$, of the same quantities a and b, the four quantities $A \times a$, $A \times b$, $B \times a$, $B \times b$ will be proportional, for $A \times a$ compared with $A \times b$ is of the same magnitude as a compared with b, and $B \times a$ compared with b. pared with $B \times b$ is also of the same magnitude as a compared with b.

LIII. In any triangle AGg if a line Ee be drawn parallel to either of the sides as

Gg, the side Ag will be to AE, as Ag to Ae, or as Gg to Ee.

To demonstrate this, upon the line AG take the line AB so that a certain multiple of it may be equal to AE, and another multiple of it may be equal to AG; this may be always done accurately when AE and AG are commensurable; if they are not commensurable, accurately quantity AB may be taken so small that certain multiples of it may differ from AE and AG respectively by

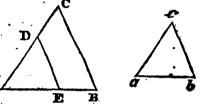


quantities less than any assignable. On the line AG, take BC, CD, DE, EF, FG, &c. each equal to AB, and through these points draw the lines Bb, Ce, &c. parallel to Gg, cutting the line Ag in the points b, c, d, e, &c. draw also the lines BM, CL, DK, &c. parallel to Ag, cutting the former parallels in the points N, O, P, &c. and the line Gg in the points M, L, K, &c. Then the triangles ABb, BCN, CDO, &c. are similar and equal to each other: for the . lines Bb, CN are parallel, therefore the angle ABh=BCN (by art. 31) and

by the same article the angle BAb is equal to CBN (because BN is parallel to Ab) and by construction AB=BC, therefore (by art. 38) the triangles ABb and BCN are equal to each other; and in the same manner we may prove that the others CDO, DEP, EFQ, &c. are equal to ABb. Therefore Ab=BN=CO=DP, &c. and Bb=CN=DO=EP, &c. but (by art. 49) BN=Bc, CO=cd, DP=de; therefore Ab=be=cd=de, &c. and since (by construction) AB=BC=CD, &c. any line AE is the same multiple of AB as the corresponding line Ae is of Ab; and AG is the same multiple of AB as Ag is of Ab; therefore the lines AG, AE, Ag, Ae, are proportional (by art. 52;) that is, AG is to AE as Ag is to Ae; and in a similar manner we may prove that AG is to AE as Gg is to Ee.

If any two triangles, ABC, abc, are similar, or have all the angles of the one, equal to all the angles of the other, each to each respectively, that is, CAB=cab, ACB=acb, ABC=abc; the legs opposite to the equal angles will be proportional, viz. AB: ab:: AC: ac; AB: ab:: BC: bc; and AC: ac:: BC: bc.

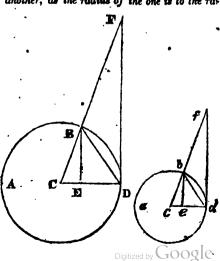
To prove this, set off upon a side AB of the largest triangle AE = ab, and through E draw ED parallel to BC, to meet AC in D, then since DE, BC are parallel; the angle AED is equal to ABC (by art. 31) and this (by supposition) is equal to the angle abc; also the angle DAE is (by A supposition) equal to eab; there-



fore in the triangles ADE, abc, the two angles DAE, AED of the one, are equal to the two angles cab, abc of the other, each to each respectively, and the included side AE is (by construction) equal to the included side ab; therefore (by art. S8)* AD is equal to ac, and DE equal to ic: but since in the triangle ABC there is drawn DE parallel to BC one of its sides, to meet the other two sides in the points DE; therefore (by the preceding art.) AB: AE:: AC: AD, and AB: AE:: BC: DE, and AC: AD:: BC: DE; if in these three proportions for DE we put its equal bc, for AE put ab, and for AD put ac; they will become AB: ab:: AC: ac, and AB: ab:: BC: bc, and AC: ac:: BC: bc.

The chord, sine, tangent, &c. of any arch in one circle, is to the chord, sine, tangent, &c. of the same arch in another, as the radius of the one is to the radius of the other.

Let ABD, abd, be two circles; BD, bd, two arches of these circles, equal to one another, or consisting of the same number of degrees ;— FD, fd, the tangents; Bd, bd, the chords; BE, be, the sines, &c. of these two arches BD, bd, and CD, cd, the radii of the circles; then CD: cd :: FD : fd, and CD : cd ::* BD: bd, and CD: cd:: BE: be, &c. For since the arches BD, bd, are equal, the angles BCD, bcd, are also equal, and PD, fd, being tangents to the points D and d, the angles CDF, cdf are each equal to a right angle (by art. 22;)



therefore since in the two triangles CDF, cdf, the two angles FCD, CDF of the one, are equal to the two angles fcd, cdf, of the other, each to each, the remaining angle CFD is also equal to the remaining angle cfd, (by art. 36;) consequently the triangles CFD, cfd, are similar. The triangles BCD, bed are also similar, for the angle CBD is equal to the angle CDB, being each subtended by the radius; therefore (by art. 36) each of these angles is equal to half the supplement of the angle BCD; and in the same manner the angle cbd or cdb is equal to half the supplement of the angle bcd, and since the angle BCD is equal to bcd, the angles of these two triangles must be equal, consequently they are similar. The triangles BCE, bce are also similar, because BE is parallel to FD, and be parallel to fd. Hence we obtain (by art. 54) the following analogies. CD: cd::FD: fd; CD: cd::BD: bd; CB: cb:: BE: be, &c.

LVI.

Let ABD be a quadrant of a circle, described by the radius CD, BD any arch of it, BA its complement, BG or CF the sine, CG or BF the co-sine, DE the tangent, AH the co-tangent, CE the secant, and CH the co-secant of that arch BD. Then since the triangles CDE, CGB, are similar or equi-angular we shall have (by art. 54) DE: CE::

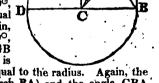
BG: CB, that is, the tangent of an arch, is to secant of the same, as the sine of it is to radius. Also, CE: CD:: CB:

CG; that is, the secant is to radius as the radius to the cosine of an arch. Also, CF: CA:: CB: CH, that is, the C: GD: DE, that is, the similar to the triangle CDE, we have AH: CA:: CD: DE, that is, the co-tangent is to the radius as the radius to the tangent of an arch.

In all circles, the sine of 90°, the tangent of 45°, and the chord of 60°, are

each equal to the radius.

For in the circle DFAEB, let the arch BE be 45°, the arch BA 60°, and BF 90°. Draw through the centre C the diameter DCB and perpendicular thereto the tangent BG meeting CE produced in G; draw the chord BA, and join CF, CA.—Then since the arch BF is 90°, DF must be 90°, whence (by art. 12 & 19) the radius CF is equal to the sine of the arch BF, or sine of 90°. Again, in the triangle CBG, since the angle CBG is 90°, and BCG is 45° by supposition, the angle CGB is also 45° (by art. 36) therefore (by art. 39) BG is



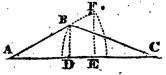
equal to CB, that is, the tangent of 45° is equal to the radius. Again, the angle ACB is 60° (being measured by the arch BA) and the angle CBA is also 60° (being measured by half the arch AD=120° by art. 40) therefore (by art. 39) CA=AB, that is, the chord of 60° is equal to the radius.

The four following propositions contain the demonstration of the rules by which all the calculations of trigonometry may be made; they were inserted here in order to prevent any embarrassment of the young calculator, from the introduction of the demonstrations among the precepts for calculation.

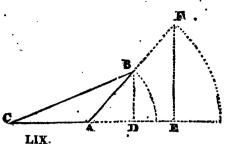
LVIII.

In any plane triangle, the sides are proportional to the sines of the opposite angles.

Let ABC be the triangle; produce the lesser side AB to F, making AF equal to BC; from B and F let fall the perpendiculars BD, FE, upon AC (produced if necessary;) then FE is the sine of the angle A, and BD is the sine of the angle C, the



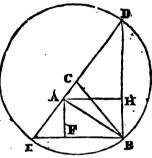
tadius being BC equal to AF; now the triangles ABD, AFE, having the angle A common to both, and the angle D equal to the angle E (being each equal to a right angle) are similar; hence (by art. 54) as AF (or its equal BC) is to AB, so is FE to BD; that is, BC is to AB as the sine of the angle A is to the sine of the angle C.



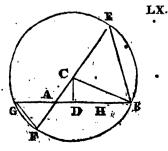
In any triangle (supposing any side to be the base, and calling the other two the sides) the sum of the sides is to their difference, as the tangent of half the sum of the angles at the base is to the tangent of half the difference of the same angles.

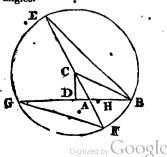
Thus, in the triangle ABC, if we call AB the base, it will be as the sum of AC and CB is to their difference, so is the tangent of half the sum of the angles ABC, BAC, to the tangent of half their difference.

Dem. With the longest leg CB as radius, describe a circle about the centre C, meeting the shorter side AC (produced on each side) in the points D and E, join EB, DB; draw AH perpendicular to DB, and AF perpendicular to EB; then (by art. 42) the angle EBD, being in a semi-circle, is a right angle; and the triangles AHD, AFE, are similar, and AF is equal to HB. Moreover,



since CB is equal to CD or CE, AD is the sum and AE is the difference of the legs AC, CB; likewise (by art. S3) the angle BCD is equal to the sum of the angles BAC, ABC, and therefore (by art. 40) the angle DEB, or its equal DAH, is equal to half the sum of the angles at the base ABC, BAC. Again (by art. S3) the angle BAC is equal to the sum of the angles CEB (or CBE) and ABE, and therefore is equal to the sum of the angle ABC, and twice the angle ABE; hence the angle ABE or its equal BAH, is equal to half the difference of the angles at the base. But in the right angled triangles AHD, AHB, making AH radius, the legs DH, HB are the tangents of the angles DAH, BAH, or the tangents of half the sum and half the difference of the angles at the base; but by reason of the similar triangles AHD, AFE, we have AD: AE: DH: AF or HB; that is, AD, the sum of the legs AC and CB, is to AE their difference, as DH the tangent of half the sum of the angles at the base (the radius being AH) is to MB the tangent of half the difference of the same angles, (to the same radius,) and therefore (by art. 55) as the tabular tangent of half the sum of the same angles at the base is to the tabular tangent of half the difference of the same angles.





In any plane triangle ABC, if the line CD be drawn perpendicular to the base AB, dividing it into two segments, AD, DB, and the base AB be bisected in the point H, we shall have,

As the base AB is to the sum of the sides, AC, BC, so is the difference of the sides to twice the distance DH of the perpendicular from the middle of the base.

Dem. With the greater side CB as radius, describe about the centre C the circle BFGE, meeting the other side produced in the points E and F, and the base AB produced in G; join GF and BE. Then AE is the sum, and AF the difference of the sides AC, CB; and since CD is perpendicular to GB, the line GB is bisected in D (by art. 45) and as AB is bisected in H, the line AG is equal to twice DH. Now in the triangles BAE, GAF, the angles ABE, GFA are equal (by art. 41) and the angle BAE is equal to GAF (by art. 28) therefore the remaining angles AEB, AGF, are equal, and the triangles BAE, GAF, are similar; consequently (by art. 54) AB: AE::AF:AG, or twice HD, which is the proposition to be demonstrated. Having thus obtained HD, we may find the segments AD, DB, by adding HD to the half base HA or HB and by taking their difference.

In any plane triangle, the square of radius is to the square of the co-sine of half of either of the angles, as the rectangle contained by the two sides including that angle is to the rectangle contained by the A C H D E half sum of the sides, and that half sum decreased by the side opposite to that angle.

CB+CE|*-BE* CB+CE+BE×CB+CE-BE
Again, AD=AC+CD=CB+CD;

we have

CB,

adding

this

CA

OT

hence AD|2=CB2+2CB·CD+CD2; also, BD2=CB2-CD2; hence AB2= AD2+BD2=-2CB2+2CB.CD=2CB×CB+CD=2CB.AD; hence AB2: AD2::

*CB : AD= ; but AB being radius, AD is the co-sine

The other cases of this proposition may be demonstrated in the same manner.

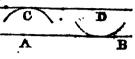
S.CE .

GEOMETRICAL PROBLEMŚ.

PROBLEM I.

To draw a Right Line CD parallel to a given Right Line AB, at any given distance, as at the point D.

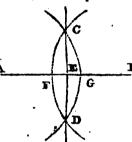
WITH a pair of compasses take the nearest distance between the point D and the given right line AB; with that distance set one foot of the compasses any where on the line AB, as at A. and draw the arch C on the same side of the line AB as the point D, from the point D draw a line so as just to touch the arch C, and it is done; for the line CD will be partile! to the line AB, and at the distance of the point given D, as was required.



PROBLEM II.

To bisect or divide a given line AB into two equal parts.

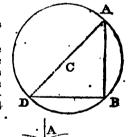
Take any distance in your compasses great- A er than half the line AB, then with one foot in B, describe the arch CFD; with the same distance, and one foot in A, describe the arch CGD. cutting the former arch in C and D; draw the line CD, and it will bisect AB in the point E.



PROBLEM III.

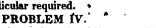
To erect a perpendicular BA on the end of a given Right Line DB.

Take any extent in your compasses, and with one foot in B fix the other in any point C without the given line; then with one point of the compasses in C, describe with the other the circle ABD; through D and C draw the diameter DCA meeting the circle in A; join B and A and it is done; for BA will be the required line (by art. 42 Geom.)

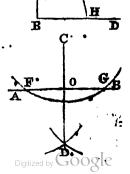


Or thus,

Take any convenient distance as BH in your compasses, and with one foot in B describe the erch HFG, upon which set off the same distance as a chord from H to F, and from F to G, upon F and G describe two arches intersecting each other in A: draw a line from B to A and it is done; for BA will be the perpendicular required.



From a given point as C, to let fall a perpendicular CO, on a given Right Line AB. Take any extent in your compasses greater than the least distance between C and the given line AB; with one foot in C, describe an arch to cut the given line AB in F and G;—with one foot in G describe an arch, and with the same distance, and one foot in F, describe another arch cutting the former in D; from C to D draw the line COD, cutting AB in O; then CO will be the perpendicular required.



PROBLEM V.

From a given point C to let fall a perpendicular CB on a given Line AB, when the perpendicular is to fall so near the end of the given line that it cannot be done as above.

Upon any point A of the line AB as a centre, and with the distance AC describe an arch E; choose any other point in the line AB, as D, and with the distance DC describe another arch intersecting the former in E, join CE cutting AB in B, and it is done, for CB will be the perpendicular required.

PROBLEM VI. To make an angle that shall contain any proposed number of degrees, from a given point in a given line. H CASE 1. When the given angle is right, or contains 900 let CA be the given line, and C the given point.

On C erect a perpendicular CD, and it is done; for the angle DCA is an angle of 90°. Or thus, on the point C as a centre, with the chord of 600* describe an arch GH, and set off thereon from G to H the distance of the chord of 900 and from C through H draw CHD, which will form the angle DCA of C

90° required.
CASE 2. When the angle is acute, as for example 36° 50' let CB be the given line and C the point

at which the angle is to be made.

With the chord of 600 in your compasses, and one foot on C, as a centre, draw the arch FB, on which set off from B to F, the given angle 5640 taken from the line of chords; through F and the centre C draw the right line A.C. and it is done; for the angle ACB will be an angle of 36° 30' as was required

.Case S. When the given angle is obtuse, as for example 1270 20' let CB be the given line and C the angular point.

Take the chord of 60° in your compasses, and with one foot on C as a centre, describe an arch BGHE, upon which set off the chord of 60° (which you already have in) your compasses) from B to G, and from G to H; then set off from G to E, the excess of the given angle above 60°, which is 671° taken from the line of chords, or you may set off from H to E, the excess of the given angle above 1200, which. is 710; draw the line CE, and it is done, for the angle ECB will be an angle of 127° 20'.

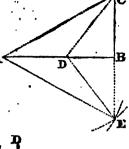
Were it required to measure a given angle, the process would have been nearly the same, by sweeping an arch as BE, and measuring it on the line of

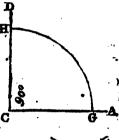
chords, as is evident.

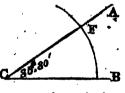
PROBLEM VII.

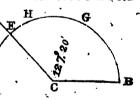
To bisect a given arch of a circle AB, whose centre is C.

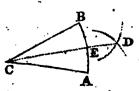
Take in your compasses any extent greater than the half of AB, and with one foot in A, describe an arch; with the same extent and one foot in B, describe another arch cutting the former in D; join CD and it is done, for this line will bisect the arch AB in the point E. also evident that the line CD bisects the angle BCA, or divides it into two equal parts.











^{*} For a description of the line of Chords see page 20.

PROBLEM VIII.

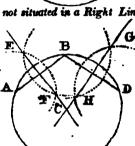
To find the centre of a given Circle.

With any radius, and one foot in the circumference as at A, describe an arch of a circle, as CBD, cutting the given circle in B; with the same extent, and one foot in B, describe. another arch CAD, cutting the former in C and D; through C and D draw the line CD, which will pass through the centre of the circle; in like manner may another right line be drawn, as EFG, which shall cross the first right line at the centre required. This construction depends upon article 43 of Geometry.

PROBLEM IX.

To draw a Circle through any three given points not situated in a Right Line.

Let A, B and D be the given points. Take in your compasses any distance greater than half AB, and with one foot in A describe an arch EF; with the same extent, and one foot in B, describe another arch cutting the former in the . points E, F, through which draw the indefinite right line EFC; then take in your compasses A any extent greater than half BD, and with one foot in B, describe an arch GH; with the same extent, and one foot in D, describe another arch cutting the former in the points G, H, through which draw the right line GHC, cutting the former right line EFC, in the point C; upon the point C as a centre, with an extent equal to CA, CB, or CD, as radius, describe the sought circle.

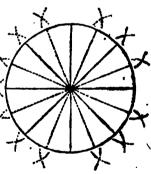


PROBLEM X.

To divide a Circle into 2, 4, 8, 16, or 32,

equal parts.

Draw a diameter through the centre, diriding the circle into two equal parts; bisect this diameter by another drawn perpendicular thereto, and the circle will be divided, into four equal parts or quadrants; bisect each of these quadrants again by right lines drawn through the centre, and the circle will be divided into eight equal parts; and so you may continue the bisections any number of times. This problem is useful in constructing the mariner's compass.



PROBLEM XI.

To divide a given Line into any number of equal parts.

Let it be required to divide the line AB into five equal parts.—From the point A draw any line AD, making an angle with the line AB; then through the point B draw a line BC parallel to AD; and from A, withany small opening in your compasses, set off A a number of equal parts on the line AD, less by one than the proposed number (which C number of equal parts in this example is 4:) 4 then from B set off the same number of the same parts on the line BC, then join 4 and 1, 3 and 2, 2 and 3, 1 and 4, and these lines will cut the given line as required.

CONSTRUCTION OF THE PLANE SCALE.

1st. WITH the radius you intend for your scale, describe a semicircle ADB, (Plate II. fig. 1.) and from the centre C draw CD perpendicular to AB, which will divide the semicircle into two quadrants, AD, BD; continue CD towards S, draw BT perpendicular to CB, and join BD and AD.

2dly. Divide the quadrant BD into 9 equal parts, then will each of these be 10 degrees; subdivide each of these parts into single degrees, and if your radius will admit of it, into minutes or some aliquot parts of a degree greater

than minutes.

sdly. Set one foot of the compasses in B and transfer each of the divisions of the quadrant BD to the right line BD, then will BD be a line of chords.

4thly. From the points 10, 20, 30, &c. in the quadrant BD draw right lines parallel to CD, to cut the radius CB, and they will divide that line into a line of sines which must be numbered from C towards B.

5thly. If the same line of sines be numbered from B towards C, it will become a line of versed sines, which may be continued to 180°, if the same divisions be transferred on the same line on the other side of the centre C.

6thly. From the centre C, through the several divisions of the quadrant BD, draw right lines till they cut the tangent BT, so will the line BT become a line of tangents.

Whly. Setting one foot of the compasses in C, extend the other to the several divisions 10, 20, 80, &c. in the tangent line BT, and transfer these extents severally to the right line CS, then will that line be a line of secants.

8thly. Right lines drawn from A to the several divisions 10, 20, 30, &c. in the quadrant BD, will divide the radius CD into a line of semi-tangents.

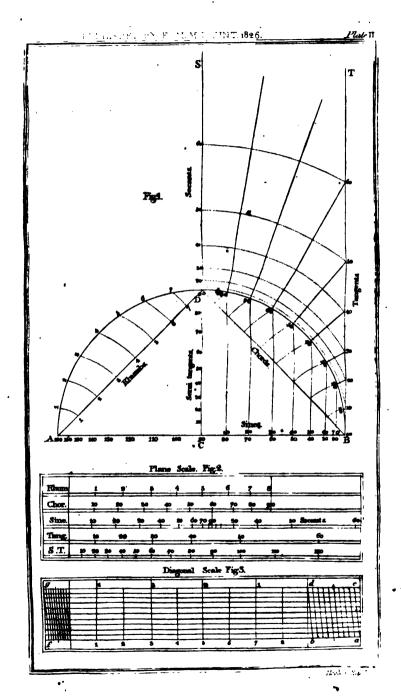
9thly. Divide the quadrant AD into eight equal parts, and from A as a centre transfer these divisions severally into the line AD, then will AD be a line of Rhumbs, each division answering to 110 15 upon the line of chords. use of this line is for protracting and measuring angles, according to the common division of the mariner's compass. If the radius AC be divided into 100 or 1000, &c. equal parts, and the lengths of the several sines, tangents, and secants, corresponding to the several arches of the quadrant, be measured thereby, and these numbers be set down in a table, * each in its proper column, you will by these means have a collection of numbers by which the several cases in trigonometry may be solved. Right lines graduated as above. being placed severally upon a ruler, form the instrument called the Plane Scale, (see Plate II. fig. 2.) by which the lines and angles of all triangles may be measured. All right lines, as the sides of plain triangles, &c. when they are considered simply as such without having any relation to a circle, are measured by scales of equal parts, one of which is subdivided equally into 10, and this serves as a common division to all the rest. In most scales an inch is taken for a common measure, and what an inch is divided into is generally set at the end of the scale. By any common scale of equal parts, divided in this manner, any number less than 100 may be readily taken; but if the number should consist of three places of figures, the value of the third figure cannot be exactly ascertained, and in this case it is better to use a diagonal scale, by which any number consisting of three places of figures, may be exactly found. The figure of this scale is given in Plate II. fig. 3: its construction is as follows.

Having prepared a ruler of convenient breadth for your scale, draw near the edges thereof two right lines, af, cg, parallel to each other; divide one of these lines as af, into equal parts, according to the size of your scale; † and

^{*} In table XXIV. is given the sine and co-sine to every minute of the quadrant, to five places of decimals.

The leggth of one of these equal parts at the end of the scale to which this description refers to ob.

The length of one of the equal parts of the scale of the other end being the half of ob.



through each of these divisions draw right lines perpendicular to af, to meet cg, then divide the breadth into 10 equal parts, and through each of these divisions draw-right lines parallel to af and cg; divide the lines ab, cd, into 10 equal parts, and from the point a to the first division in the line cd, draw a diagonal line; then parallel to that line, draw diagonal lines through all the other divisions, and the scale is complete. Then, if any number, consisting of three places of figures, as 256, be required from the larger scale gd, you must place one foot of the compasses on the figure 2 on the line gd, then the extent from 2 to the point d will represent 200. The second figure being 5, count live of the smaller divisions from d towards e, and the extent from 2 to that point will be 250. Move both points of the compasses downwards till they are on the sixth parallel line below gd, and open them a little till the one point rests on the vertical line drawn through 2, and the other on the diagonal line drawn through 5; the extent then in the compasses will represent 256. In the same way the quantities 25,6; 2,56; 0,256, &c. are measured. Besides the lines already mentioned, there is another on the Plane Scale

marked ML, which is joined to a line of chords, and shows how many miles of easting or westing correspond to a degree of longitude in every latitude.* These several lines are generally put on one side of a ruler, two feet long; and on the other side is laid down a scale of the logarithms of the sines, tangents, and numbers, which is commonly called Gunter's Scale, and as it is of general use, it requires a particular description.

GUNTER'S SCALE.

ON GUNTER'S SCALE are eight lines, viz.

Sine Rhumbs, marked (SR) corresponding to the logarithms of the natural sines of every point of the mariner's compass, numbered from the left hand towards the right, with 1, 2, 3, 4, 5, 6, 7, to 8, where is a brass pin. This line is also divided, where it can be done, into halves and quarters.

Tangent rhumbs, marked (TR) correspond to the logarithms of the tangents of every point of the compass, and are numbered 1, 2, 3, to 4, at the right hand where there is a pin, and thence towards the left hand with 5, 6,

7; it is also divided, where it can be done, into halves and quarters.

The line of numbers, marked (Num.) corresponds to the logarithms of numbers, and is marked thus; near the left hand it begins at 1, and towards the right hand are 2, 3, 4, 5, 6, 7, 8, 9; and 1 in the middle, at which is a brass pin, then 2, 3, 4, 5, 6, 7, 8, 9, and 10 at the end, where there is another pip. The values of these numbers and their intermediate divisions depend on the estimated values of the extreme numbers 1 and 10; and as this line is of great importance, a particular description of it will be given. 1 may be counted for 1, 10, 100, or 1000, &c. and then the next 2, will be 2, 20, 200, or 2000, &c. respectively. Again, the first 1 may be reckoned 1 tenth, 1 hundredth, or 1 thousandth part, &c. then the next will be 2 tenths, or 2 hundredths, or 2 thousandths parts, &c. so that if the first 1 be esteemed 1, the middle 1 will be 10; 2 to its right 20; 3, 30; 4, 40; and 10 at the end 100; again, if the first 1 is 10, the next 2 is 20, 3 is 30, and so on, making the middle 1, 100, the next 2 is 200, 3 is 300, 4 is 400, and 10 at the end is 100% In like manner, if the first 1 be esteemed 1 tenth part, the next 2 will be 2

The description and use of logarithms are given in page 29, et seq. The logsines, tangents, &c. are marked on these scales by means of a line of equal parts,

corresponding to the size of the scale:

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^{*} As it would confuse the adjoined figure to describe on it the line of longitudes, it is neglected, but the construction is as follows: divide the line CB into 60 equal parts (if it can be done) and through each point draw lines parallel to CD to interset the arch BD: about B, as a centre, transfer the several points of intersection to the line BD, and then number it from D, towards B, from 0 to 60, and it will be the line of longitudes.

tenth parts, and the middle 1 will be 1; the next 2, k; and 10 at the end will be 10. Again, if the first 1 be counted 1 hundredth part, the next 2 hundredth parts, the middle 1 will be 10 hundredth parts, or one tenth part, and the next 2 two tenth parts, and 10 at the end will be but one whole number or integer.

As the figures are increased or diminished in their value, so in like manner must all the intermediate strokes or subdivisions be increased or diminished: that is, if the first 1 at the left hand be counted 1, then 2 (next following it) will be 2, and each subdivision between them will be 1 tenth part, and so all the way to the middle 1, which will be 10, the next 2, 20, and the longer strokes between 1 and 2 are to be counted from 1 thus, 11, 12, (where is a brass pin) then 15, 14, 15, sometimes a longer stroke than the rest, then 16, 17, 18, 19, 20, at the figure 2; and in the same manner the short strokes between the figures 2 and 3, 3 and 4, 4 and 5, &c. are to be reckoned as units. Again, if 1 at the left hand be 10, the figures between it and the middle 1 will be common tens; and the subdivisions between each figure will be units; from the middle 1 to 10 at the end, each figure will be so many hundreds; and between these figures each longer division will be 10. From this description it will be easy to find the divisions representing any given number, thus: Suppose the point representing the number 12, were required; take the division at the figure 1 in the middle, for the first figure of 12; then for the second figure count two tenths, or longer strokes to the right hand, and this will be the point representing 12, where the brass pin is.

Again, suppose the number 22 were required; the first figure 2 is to be found on the scale, and for the second figure 2, count 2 tenths onwards, and

that is the point representing 22.

Again, suppose 1728 were required; for the first figure 1, I take the middle 1, for the second figure 7, count onwards as before, and that will be 1700. And as the remaining figures are 28 or nearly 30, I note the point which is nearly 12, of the distance between the marks 7 and 8, and this will be the point representing 1728.

If the point representing 435 was required; from the 4 in the second interval count towards 5 on the right, three of the larger divisions and one of the smaller (this smaller division being midway between the marks 3 and 4) and that will be the division expressing 435. In a similar manner other

numbers may be found.

All fractions found in this line must be decimals; and if they are not; they must be reduced into decimals, which is easily done by extending the compasses from the denominator to the numerator; that extent laid the same way, from 1 in the middle or right hand, will reach to the decimal required.

EXAMPLE. Required the decimal fraction equal to \$: Extend from 4 to 5; that extent will reach from 1 on the middle to .75 towards the left hand.

The like may be observed of any other vulgar fraction.

Multiplication is performed on this line, by extending from 1 to the multi-

plier: that extent will reach from the multiplicand to the product.

Suppose, for example, it were required to find the product of 16 multiplied by 4, extend from 1 to 4; that extent will reach from 16 to 64, the product required.

Division being the reverse of multiplication, therefore extend from the divisor to unity; that extent will reach from the dividend to the quotient.

Suppose 64 to be divided by 4; extend from 4 to 1, that extent will reach

from 64 to 18, the quotient.

Questions in the Rule of Three are solved by this line as follows: Extend from the first term to the second, that extent will reach from the third term to the fourth. And it ought to be particularly noted, that if you extend to the left, from the first number to the second, you must also extend to the left, from the third number to the fourth; and the contrary.

^{*} Or you may extend from the first to the third, for that extent will reach from the second to the fourth. 'This method must be adopted when using the lines of sizes, tangents, &c. if the first and third terms are of the same name, and different from the second and fourth.

EXAMPLE. If the diameter of a circle be 7 inches, and the circumference 22, what is the circumference of another eircle, the diameter of which is 14 inches? Extend from 7 to 22, that extent will reach from 14 to 44 the same way.

.The superficial content of any parallelogram is found by extending from 1 to the breadth; that extent will reach from the length to the superficial content.

EXAMPLE. Suppose a plank or board, 15 inches broad and 27 feet long, 'the content of which is required: Extend from 1 to 1 foot 8 inches (or 1,25;) that extent will reach from 27 feet to 38,75 feet, the superficial content. Or extend from 12 inches to 15, &c.

The solid content of any bale, box, chest, &c. is found by extending from 1 to the breadth; that extent will reach from the depth to a fourth number, and the extent from 1 to that fourth number will reach from the length to the

solid content.

EXAMPLE 1st. What is the content of a square pillar, whose length is 21 feet 9 inches, and breadth 1 foot 3 inches? The extent from 1 to 1,25 will reach from 1,25 to 1,56, the content of one foot in length; again, the extent from 1 to 1,56 will reach from the length 21,75 to 33,9, or 34, the solid content in feet.

EXAMPLE 2d. Suppose a square piece of timber, 1,25 feet broad, ,56 deep, and 36 long, be given to find the content: extend from 1 to 1,25; that extent will reach from ,56 to ,7; then extend from 1 to ,7; that extent will reach from 36 to 25,2 the solid content. In like manner may the contents of bales, &c. be found, which divided by 40 will give the tonnage.

4thly. The line of sines marked (Sin.) corresponding to the log-sines of the degrees of the quadrant, begins at the left hand, and is numbered to the right thus; 1, 2, 3, 4, 5, &c. to 10; then 20, 30, 40, &c. ending at 90 degrees,

where is a brass centre pin, as there is at the right end of all the lines.

5thly. The line of versed sines, marked (V. S.) corresponding to the log versed sines of the degrees of the quadrant, begins at the right hand against 90° on the sines, and from thence is numbered towards the left hand thus: 10, 29, 50, 40, &c. ending at the left hand at about 169°; each of the subdivisions, from 10 to 30, is in general two degrees, from thence to 90 is single degrees, from thence to the end, each degree is divided into 15 minutes.

Sthly. The line of tangents, marked (Tang.) corresponding to the log-tangents of the degrees of the quadrant, begins at the left hand, and is numbered towards the right thus: 1, 2, 3, &c. to 10, and so on 20, 30, 40, and 45, where is a brass pin under 90° on the sines; from thence it is numbered backwards, 50, 60, 70, 80, &c. to 89, ending at the left hand where it began at 1 degree. The subdivisions are nearly similar to those of the sines. When you have any extent in your compasses, to be set off from any number less than 45° on the line of tangents, towards the right, and it is found to reach beyond the mark of 45°, you must see how far it extends beyond that mark, and set it off from 45° towards the left, and see what degree it falls upon, which will be the number sought, which must exceed 45°; if, on the contrary, you are to set off such a distance to the right from a number greater than 45°, you must proceed as before, only remembering, that the answer must be less than 45°, and you must always consider the degrees above 45° as if they were marked on the continuation of the line to the right hand of 45°.

7thly. The line of the meridional parts, marked (Mer.) begins at the right hand, and is numbered thus; 10, 20, 50, &c. to the left hand, where it ends at 87 degrees. This line, with the line of equal parts, marked (E. P.) under it, are used together, and only in Mercator's Sailing. The upper line contains the degrees of the meridian, or latitude in a Mercator's chart, corresponding

to the degrees of longitude on the lower line.

The use of this Scale in solving the usual problems of Trigonometry, Plane Sailing, Middle Latitude Sailing, and Mercator's Sailing, will be given in the course of this work; but it will be unnecessary to enter into an explanation of its use in calculating the common problems of Nautical Astronomy, as it is much more accurate to perform those calculations by logarithms.

ON THE SLIDING RULE.

THE Sliding Rule consists of a fixed part and a stider, and is of the same dimensions, and has the same lines marked on it as on a common Gunter or Plane Scale, which may be used with a pair of compasses in the same manner as those scales; and as a description of those lines has already been given, it will be unnecessary to repeat it here, it being sufficient to observe, that there are two lines of numbers, a line of log-sines and a line of log-tangents on the slider, and that it may be shifted so as to fix any face of it on either side of the fixed part of the scale, according to the nature of the question to be solved.

In solving any problem in Arithmetic, Trigonometry, Plane Sailing, &c. let the proposition be so stated that the first and third terms may be alike, and of course the second and fourth terms alike; then bring the first term of the analogy on the fixed part, against the second term on the slider, and against the third term on the fixed part will be found the fourth term on the slider; or if necessary the first and third terms may be found on the slider, and the second and fourth on the fixed part. Multiplication and Division are performed by this rule, in considering unity as one of the terms of the analogy.

Thus, to perform multiplication, set 1 on the line of numbers of the fixed part against one of the factors on the line of numbers of the slider, then against the other factor on the fixed part will be found the product on the

slider.

EXAMPLE. To find the product of 4 by 12 draw out the slider till 1 on the fixed part coincides with 4 on the slider, then opposite 12 on the fixed part will be found 48 on the slider.

To perform Division, set the divisor on the line of numbers of the fixed part against 1 on the slider, then against the dividend, on the fixed part, will

be found the quotient on the slider.

EXAMPLE. To divide 48, by 4—set 4 on the fixed part against 1 on the slider, then against 48 on the fixed part will be found 12 on the slider.

EXAMPLES IN THE RULE OF THREE.

If a ship sail 25 miles in 4 hours, how many miles will she sail in 12 hours at the same rate?

Bring 4 on the line of numbers of the fixed part against 25 on the line of numbers of the slider, then against 12 on the fixed part will be found 75 on the

slider, which is the answer required.

EXAMPLE. If S pounds of sugar cost 21 cents, what will 27 pounds cost? Bring S on the line of numbers of the fixed part, against 21 on the line of numbers of the slider, then against 27 on the fixed part, will be found 189 on the slider.

EXAMPLE IN TRIGONOMETRY.

In the oblique-angled triangle ABC, let there be given AB=56, AC=64, angle ABC= 46° 30' to find the other angles and the side BC.

In this case we have (by art. 58 Geometry) the following canons. AC (64): sine < B (46° 30'):: AB (56): sine < C, and sine < B: AC :: sine < A: BC. Therefore, to work the first proposition by the sliding rule,

^{*} If the first and second terms are alike, instead of the first and third, you must bring the first term on the slider against the third on the fixed part, and against the second term on the slider, will be found the fourth term on the fixed part. Or, if necessary, the first and second terms may be found on the fixed part, and the third and fourth on the slider.

we must bring 64 on the line of numbers of the fixed part, against 46° 30' on the line of sines of the slider, then against 56 on the former will be 390 24' on the latter, which will be the angle C. The sum of the angles B and C being subtracted from 1800 leaves the angle A=940 6'. Then, by the second canon, bring the angle B=46° 30' on the line of sines of the slider against AC=64 on the line of numbers of the fixed part, then against the angle A=94° 6' (or its supplement 85° 54') on the slider will be found the side BC=88 on the fixed part.

In a similar manner may the other propositions in trigonometry be solved. From what has been said, it will be easy to work all the problems in Plane, Middle Latitude, and Mercator's Sailing, as in the three following examples, which the learner may pass over until he can solve the same problems by the Scale. If any one wishes to know the use of the sliding rule in problems of Spherical Trigonometry, he may consult the treatises written expressly on that subject: but it may be observed, that in such calculations the aliding rule is rather an object of curiosity than of real use, as it is much more accurate to make use of logarithms.

Example 1. Given the course sailed 1 point, and the distance 85 miles-

required the difference of Latitude and Departure?

By Case 1st of Plane Sailing, we have these canons:

Radius (8 points): Distance (85):: Sine. Co. Course (7 points): Diff. Lat.;

and Radius (8 points): Distance (85):: Sine Course (1 point): Departure.

Hence we must bring the radius 8 points on the fixed part of the Sine
Bhumbs against 85 on the line of numbers on the slider, then against 7 points on the sine rhumbs will be found the diff. of lat. 834 on the slider, and against one point will be found the departure 161 miles.

If the course is given in degrees, you must use the line marked Sin.

Example 2. Given the diff. of lat. 40 miles, and departure 30 miles—re-quired the course and distance?

By case 6, of Plane Sailing, we have this canon for the course:-Diff. Lat. (40): Radius 45°:: Departure (30): Tang. Course.

Hence we must bring 40 on the line of numbers of the slider against 450 on the line of tangents on the fixed part, then against 30 on the slider will be found the course 370 nearly.

Again, the canon for the distance gives :

Sine Course (87°): Departure (50):: Radius (90°): Distance.

Hence we must bring 370 on the line of sines of the fixed part against 30 on the line of numbers on the slider, then against 900 on the line of sines of the fixed part will be found the distance 50 on the slider.

EXAMPLE 3. Given the Middle Lat. 400 and the departure 30 miles-re-

quired the Diff. of Long.?

By case 6, of Middle Latitude Sailing, we have this canon :-

Sine Comp. Mid. Lat. (50°): Departure (30):: Radius (90°): Diff. Long. Hence by bringing 500 on the line of sines of the fixed part against 50 on the line of numbers on the slider, then against 90° on the fixed part, we shall find 39 on the slider, which will be the difference of longitude required.

DESCRIPTION AND USE OF THE SECTOR.

THIS instrument consists of two rules or legs, moveable round an axis or joint, as a centre, having several scales drawn on the faces, some single, others double; the single scales are like those upon a common Gunter's Scale; the double scales are those which proceed from the centre, each being laid twice on the same face of the instrument, viz. once on each leg. From these scales, dimensions or distances are to be taken, when the legs of the

instrument are set in an angular position.

The single scales being used exactly like those on the common Gunter's Scale, it is unnecessary to notice them particularly; we shall therefore only enumerate a few of the uses of the double scale, the number of which is seven, viz. the scale of Lines, marked Lin. or L. the scale of chords, marked Cho. or C. the scale of Sines, marked Sin. or S. the scale of Tangents to 45°, and another scale of tangents from 45° to about 76°, both of which are marked Tan. or T. the scale of Secants, marked Sec. or S. and the scale

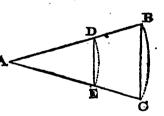
of Polygons, marked Pol.

The scale of lines, chords, sines, and tangents, under 45°, are all of the same radius, beginning at the centre of the instrument, and terminating near the other extremity of each leg, viz. the lines at the division 10, the chords at 60°, the sines at 90°, and the tangents at 45°; the remainder of the tangents, or those above 45°, are on other scales, beginning at a quarter of the length of the former, counted from the centre, where they are marked with 45°, and extend to about 76 degrees. The secants also begin at the same distance from the centre, where they are marked with 0, and are from thence continued to 750: The scales of polygons are set near the inner edge of the legs, and where these scales begin, they are marked with 4, and from thence are numbered backward or towards the centre, to 12.

In describing the use of the sector, the terms lateral distance and transverse distance often occur. By the former is meant the distance taken with the compasses on one of the scales only, beginning at the centre of the sect. Ir; and by the latter, the distance taken between any two corresponding divisio us of the scales of the same name, the legs of the sector being in an angula w

position.

The use of the sector depends upon the porportionability of the corresponding sides of similar triangles, (demonstrated in art. 58, Geometry.) For if in the triangle ABC we take AB=AC and AD= AE, and draw DE, BC, it is evident that DE and BC will be parallel; therefore by above-mentioned proposition AB:A BC :: AD : DE; so that whatever part AD is of AB, the same part DE will be of BC; hence, if DE be the chord, sine, or tangent of any arch to the radius AD, BC will be the same to the radius AB.



Use of the line of Lines.

The line of lines is useful to divide a given line into any number of equal parts, or in any proportion, or to find 3d and 4th proportionals, or mean pro-

portionals, or to increase a given line in any proportion.

Example 1. To divide a given line into any number of equal parts, as suppose 9: make the length of the given line a transverse distance to 9 and 9, the number of parts proposed; then will the transverse distance of 1 and 1 be one of the parts, or the ninth part of the whole; and the transverse distance of 2 and 2 will be 2 of the equal parts or } of the whole line, &c.

Example 2. If a ship sails 52 miles in 8 hours, how much would she sail

in 3 hours at the same rate?

Take 52 in your compasses as a transverse distance, and set it off from 8 to 8, then the transverse distance 3 and 3 being measured laterally, will be found equal to 19 and a half, which is the number of miles required.

Example 3. Having a chart constructed upon a scale of 6 miles to an inch, it is required to open the sector, so that a corresponding scale may be

taken from the line of lines?

Make the transverse distance 6 and 6, equal to 1 inch, and this position of the sector will produce the given scale.

EXAMPLE 4. It is required to reduce a stale of 6 inches to a degree, to

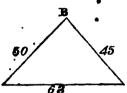
another of S inches to a degree ?

Make the transverse distance 6 and 6, equal to the lateral distance 3 and 5: then set off any distance from the chart laterally, and the corresponding transverse distance will be the reduced distance required.

EXAMPLE 5. One side of any triangle. being given, of any length, to measure the other two sides on the same scale.

Suppose the side AB of the triangle ABC. measures 50, what are the measures of the

other two sides?



Take AB in your compasses, and apply it transversely to 50 and 50; to this opening of the sector apply the distance AC in your compasses to the same number on both sides of the rule transversely; and where the two points fall will be the measure on the line of lines of the distance required; the distance AC will fall against 63, 63, and BC against 45, 45, on the line of lines.

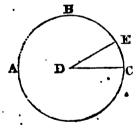
Use of the line of Chords on the Sector.

The line of chords upon the sector is very useful for protracting any angle, when the paper is so small that an arch cannot be drawn upon it with the radius of a common line of chords.

Suppose it was required to set off an arch of 30°, from the point C of the

small circle ABC.

Take the radius DC in your compasses, and set it off transversely from 60° to 60° on the chords, then take the transverse extent from 30° to 30° on the chords; and place one foot of the compasses in C, the other will reach to E, and CE will be the arch required. And by the converse operation any angle or arch may be measured, viz. with any radius describe an arch about the angular point; set that radius transversely from 60° to 60°; then take the distance of the arch, intercepted between the two



legs, and apply it transversely to the chords, which will show the degrees of

the given angle.

Note. When the angle to be protracted exceeds 60°, you must lay off and twice. 80°, and then the remaining part; or if it be above 120°, lay off 60° twice. and then the remaining part. And in a similar manner any arch above 600 may be measured.

. Uses of the lines of Sines, Tangents, and Secants.

By the several lines disposed on the sector, we have scales of several radii, so that,

1st. . Having a length or radius given, not exceeding the length of the sector when opened, we can find the chord, sine, &c. of an arch to that radius; thus, suppose the chord, sine, or tangent of 20 degrees to a radius of 2 inches be required. Make 2 inches the transverse opening to 60° and 60° on the chords; then will the same extent reach from 45° to 45° on the tangents, and

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from 90° to 90° on the sines; so that to whatever radius the line of chords is set, to the same are all the others set also. In this disposition, therefore, if the transverse distance between 20° and 20° on the chords be taken with the compass, it will give the chord of 20 degrees; and if the transverse of 20° and 20° be in like manner taken on the sines, it will be the sine of 20 degrees; and lastly, if the transverse distance of 20° and 20° be taken on the tangents, it will be the tangent of 20 degrees to the same radius of two inches.

2dly. If the chord or tangent of 70° were required. For the chord you must first set off the chord of 60° (or the radius) upon the arch, and then set off the chord of 10°. To find the tangent of 70 degrees, to the same radius, the scale of upper tangents must be used, the under one only reaching to 45°; making therefore 2 inches the transverse distance to 45° and 45° at the beginning of that scale, the extent between 70° and 70° on the same will be the tangent of 70 degrees to 2 inches radius.

tangent of 70 degrees to 2 inches radius.

3dly. To find the secant of any arch; make the given radius the transverse distance between 0 and 0 on the secants; then will the transverse distance of 20° and 20°, or 70° and 70°, give the secant of 20° or 70° respec-

tively.

4thly. If the radius and any line representing a sine, tangent, or secant, be given, the degrees corresponding to that line may be found by setting the sector to the given radius, according as a sine, tangent, or secant is concerned; then taking the given line between the compasses, and applying the two feet transversely to the proper scale, and sliding the feet along till they both rest on like divisions on both legs; then the divisions will show the degrees and parts corresponding to the given line.

Use of the line of Polygons. .

The use of this line is to inscribe a regular polygon in a circle. For example, let it be required to inscribe an octagon in a circle. Open the sector till the transverse distance 6 and 6 be equal to the radius of the circle; then will the transverse distance of 8 and 8 be the side of the inscribed polygon.

Use of the sector in Trigonometry.

All proportions in trigonometry are easily worked by the double lines on the sector; observing that the sides of triangles are taken off the line of lines, and the angles are taken off the sines, tangents, or secants, according to the nature of the proportion. Thus, if in the triangle ABC we have given AB=56, AC=64, and the angle $ABC=46^{\circ}$ 30' to find the rest. In this case we have (by art. 58, Geometry) the following proportions, as AC (64): sine < B (46°) 30°):: AB (56): sine < C, and as sine B: AC :: sine A: BC. Therefore to work these proportions by the sector, take the lateral distance 64=AC from the B2 line of lines, and open the sector to make this a transverse distance of 46° 30'=< B on the sines; then take the lateral distance 56=AB on the lines, and apply it transversely on the sines, which will give 39° 24' ≤ C. Hence the sum of the angles B and C is 85° 54', which taken from 180°, leave the angle $A = 94^{\circ} 6'$. Then to work this second proportion, the sector being set at the same opening as before, take the transverse distance of 94° 6' = the angle A on the sines, or, which is the same thing, the transverse distance of its supplement 85° 54'; then this applied laterally to the lines, gives the sought side BC=88. In the same manner we might solve any problem in trigonometry, where the tangents and secants occur, by only measuring the transverse distances on the tangents or secants, instead of measuring them on the sines, as in the preceding example. All the problems that occur in Nautical Astronomy may be solved by the sector, but as the calculation by logarithms is much more accurate, it will be useless to enter into a further detail on this subject.

LOGARITHMS.

IN order to abbreviate the tedious operations of multiplication and division with large numbers, a series of numbers, called logarithms, were invented by Lord Napier, Baron of Marchinston in Scotland, and published in Edinburgh in 1614; by means of which the operation of multiplication may be performed by addition, and division by subtraction; numbers may be involved to any power by simple multiplication, and the root of any power extracted by

simple division.

In Table XXVI. are given the logarithms of all numbers from 1 to 9999; to each one bught to be prefixed an index, with a period or dot to separate it from the other part, as in decimal fractions; the numbers from 1 to 100 are published in that table with their indices; but from 100 to 9999 the index is left out for the sake of brevity, but it may be supplied by this general rule, viz. the index of the logarithm of any integer, or mixed number, is always one less than the number of integer places in the natural number. Thus the index of the logarithm of any number (integer or mixed) between 10 and 100 is 1, from 100 to 10000 is 3, &c. the method of finding the logarithms from this table will be evident from the following examples.

To find the logarithms of any number less than 100. Enter the first page of the table, and opposite the given number

will be found the logarithm with its index prefixed.

Thus, opposite 71 is 1.85126, which is its logarithm.

To find the logarithm of any number between 100 and 1000.

RULE. Find the given number in the left hand column of the table of logarithms, and immediately under 0 in the next column, is a number, to which must be prefixed the number 2 as an index (because the number consists of three places of figures) and you will have the sought logarithm.

Thus, if the logarithm of 649 was required; this number being found in the left hand column, against it in the column marked 0 at the top (or bottom) is found \$1224, to which prefixing the index 2, we have the logarithm

of 649=2.81224.

To find the logarithm of any number between 1000 and 10000.

RULE. Find the three left hand figures of the given number, in the left hand column of the table of logarithms, opposite to which, in the column that is marked at the top (or bottom) with the fourth figure, is to be found the sought logarithm; to which must be prefixed the index 3, because the number contains 4 places of figures.

Thus, if the logarithm of 6495 was required; opposite to 649, and in the column marked 5 at the top (or bottom) is \$1258, to which prefix the index

3 and we have the sought logarithm 3.81258.

To find the logarithm of any number above 10000.

RULE. Find the three first figures of the given number, in the left hand column of the table, and the fourth figure at the top or bottom, and take out the corresponding number as in the preceding rule; take also the difference between this logarithm and the next greater, and multiply it by the given number exclusive of the four first figures, cross off at the right hand of the product as many figures as you had figures of the given number to multiply by; then add the remaining left hand figures of this product to the logarithm taken from the table, and to the sum prefix an index equal to one less than the number of integer figures in the given number, and you will have the sought logarithm.

Thus, if the logarithm of 64957 was required: opposite to 649 and under 5 is 81258, the difference between this and the next greater number 81265 is 7, this multiplied by 7 (the last figure of the given number) gives 49, crossing off the right hand figure leaves 4.9 or 5 to be added to 81258, which makes 81265, to this prefixing the index 4, we have the sought logarithm 4.81265.

Again, if the logarithm of 6495738 was required; the logarithm corresponding to 649 at the left, and 5 at the top, is as in the last example 81258,

the difference between this and the next greater is 7, which multiplied by 758 (which is equal to the given number excluding the four first figures) gives 5166, crossing off the three right hand figures of this product (because the number 738 consists of three figures) we have the correction 5 to be added to 81258; and the index to be prefixed is 6 because the given number consists of 7 places of figures, therefore the sought logarithm is 6.81263.

To find the logarithm of any mixed decimal number.

Find the logarithm of the number as if it was an integer by the last rule, to which profix the index of the integer part of the given number.

Thus, if the logarithm of the mixed decimal 649.5738 was required; find the logarithm of 6495788 without noticing the decimal point; this, in the last example, was found to be 81263, to this we must prefix the index 2, corresponding to the integer part 649; the logarithm sought will therefore be 2.81265.

To find the logarithm of any decimal fraction less than unity.

The index of the logarithm of any number less than unity is negative, but to avoid the mixture of positive and negative quantities, it is common to borrow 10 or 100 in the index, which must afterwards be neglected in summing them with other indices; thus instead of writing the index - 1, it is generally written + 9 or + 99; but in general it is sufficient to borrow 10 in the index, and it is what we shall do in the rest of this work. In this way we may find the logarithm of any decimal fraction by the following rules.

RULE. Find the logarithm of a fraction as if it was a whole number;—see how many ciphers precede the first figure of the decimal fraction, subtract that number from 9 and the remainder will be the index of the given fraction.

Thus the log. of 0,0391 is 8.59218; the log. of 0,25 is 9.59794; the log. of

0,0000025 is 4.59794, &c.

To find the logarithm of a vulgar fraction.

Subtract the logarithm of the denominator from the logarithm of RULE. the numerator (borrowing 10 in the index when the denominator is the greatest) the remainder will be the logarithm of the fraction sought.

EXAMPLE I. Required the log. of ?? From log. of 3 Take log. of 8

EXAMPLE II.

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Required the log. of 81 or 13? 0.47719 Prom log. of 13 1.11394 0.90309 Take log. of 4 0.60206

Rem. log. # or ,375

9.57403 Rem. log. of 31 or 3,25

0.51189

To find the number corresponding to any logarithm.

RULE. In the column marked 0 at the top (and bottom) of the table, seek for the next less logarithm, neglecting the index; note the number against it, and carry your eye along that line until you find the nearest less logarithm to the given one, and you will have the fourth figure of the given number at the top, which is to be placed to the right of the three other figures; if you wish for greater accuracy, you must take the difference between this tabular logarithm and the next greater, also the difference between that least tabular logarithm and the given one; to the latter difference annex 2 or more civilized to the latter difference annex 2 or more civilization. phers at the right hand, and divide it by the former difference, and place the quotient* to the right hand of the four figures already found; and you will have the number sought expressed in a mixed decimal, the integer part of which will consist of a number of figures (at the left hand) equal to the index of the logarithm increased by unity.†

Thus, if the number corresponding to the logarithm 1.52634 was required; I look for 52634 in the column marked 0 at the top or bottom, and find it standing opposite to 336; now the index being 1, the sought number must

consist of two integer places, therefore it is 33,6.

^{*} This quotient must consist of as many places of figures as there were ciphers senexed, conformable to the rules of the division of decimals. Thus, if the divisor was 40, and the number to which two ciphers were annexed was 2, making 2,00, the quotient must not be estimated as 5, but as 05, and then two figures must be placed to the right of the four figures before found.

[If, the index corresponds to a fraction less than unity, you must place as many ciphers to the left of that fittinger as are equal to the index subtracted from 9, the decimal point being placed to the left of these ciphers; in this manner you will obtain the sought number.

If the given logarithm was 2.52858; I find that 52858 stands in the column marked 0 at the top or bottom, directly opposite to 215 which is the number sought, because the index being 2, the number must consist of 8

places of figures.

If the number corresponding to the logarithm 2.57345 was required; I look in the column 0, and find in it, against the number 374, the logarithm 57287, and guiding my eye along that line, I find the given logarithm 57345 in the column marked 5; therefore the mixed number sought is 3745, and since the index is 2, the integer part must consist of 3 places, therefore the number sought is 374,5. If the index had been 1, the number would have been 37,45; and if the index had been 0, the number would have been 3,745. If the index had been 8 corresponding to a number less than unity, the answer would have been 0,03745, &c.

Again, if the number corresponding to the logarithm 5,57811 was required; I look in the column 0, and find in it against 378, and under 5, the logarithm 57807, the difference between this and the next greater logarithm 57818 being 11, and the difference between 57807 and the given number 57811 being 4, to this 4 I affix two ciphers, which make 400, and divide it by 11, the quotient is 36 nearly; this number connected with the former four agures make 878536, which is the number required, since the index being 5

the number must consist of six places of figures.

MULTIPLICATION BY LOGARITHMS.

RULE. Add the logarithms of the two numbers to be multiplied, and the sum will be the logarithm of their product.

EXAMPLE I.	• 1	EXAMPLE II.	
Multiply 25 by 35.	- 1	Multiply 22,4 by 1,8.	
25 log.	1.39794	22,4 log.	1.35025
35 log.	1.54407	1,8 log.	0.25527
T OFF T	0.04004	T - 1 4 40 00 1	
Product 875 log.	2,94201	Product 40,32 log.	1. 60 55 %
EXAMPLE III.	. 1	EXAMPLE IV.	
Multiply 3,26 by 0,0025.	i i	Multiply 0,25 by 0,003.	
3,26 log.	0.51322	0,25 log.	9.39794
0,9025 log.	7.39794	0,003 log.	7.47719
Product 0,00815 log.	7.91116	Product 0,00075 log.	6.87506

In the last example the sum of the two indices is 16, but since 10 was borrowed in each number, I have neglected 10 in the sum, and the remainder 6 being less than the other 10, is evidently the index of the logarithm of a fraction less than unity:

·· DIVISION BY LOGARITHMS.

RULE. From the logarithm of the dividend subtract the logarithm of the divisor, the remainder will be the logarithm of the quotient.

EXAMPLE I.	i	EXAMPLE II.	
Divide 875 by 25.	I	Divide 40,32 by 22,4.	
875 log.	2.94201	40,32 log.	1.6055%
25 log.	1.39794	22,4 log.	1.35025
Quotient 35 log.	1.54407	Quotient 1,8 log.	0.25527
EXAMPLE, III.	1	EXAMPLE IV.	
. Divide 0,00815 by 0,0025.	ł	Divide 0,00075 by 0,025.	
0.00815 log.	7.91116	0,00075 log. · ·	6.875 06
0,0025 log.	7.39794	0,025 log.	8.39794
•			
A	0 61 200	O	0 47710

Quotient 3,26 log. 0,51322 Quotient 0,03 log. 8.47712 In Example III. both the divisor and dividend are fractions less than unity, and the divisor is the least, consequently the quotient is greater than unity. In Example IV. both fractions are less than unity, and since the divisor is the greatest, its logarithm is greater than that of the dividend; for that reason it was necessary to borrow 10 in the index previous to making the subtraction, hence the quotient is less than unity.

INVOLUTION BY LOGARITHMS.

RULE. Multiply the logarithm of the number given, by the index of the power to which the quantity is to be raised, the product will be the logarithm of the power sought. But in raising the powers of any decimal fraction it must be observed, that the first significant figure of the power must be put as many places below the place of units as the index of its logarithm wants of 10 multiplied by the index of the power.

EXAMPLE I.		EXAMPLE II.	•
Required the square of 18?		Required the cube of 13?	
18 log. ' . '	1.25527	13 log.	1.11394
, , ,	2		3
Answer 324 log.	2.51054	Answer 2197 log.	3.34182
EXAMPLE III.	.4	EXAMPLE IV.	
Required the square of 6,4?	l l	Required the cube of 0,25?	
6.4 log.	0.80618	0,25 log.	9.39794
, 0	. 2		3
Answer 40,96 log.	1.61236	Answer 0,015625.	28.19382
In the last example the ind	AT 98 TOUR	is 9 of 90 (the product of	10 hv the

In the last example the index 28 wants 2 of 30 (the product of 10 by the power 3) therefore the first significant figure of the answer, viz. 1, is placed two figures distant from the place of units.

· EVOLUTION BY LOGARITHMS.

RULE. Divide the logarithm of the number by the index of the power, the quotient will be the logarithm of the root sought. But if the power whose root is to be extracted is a decimal fraction less than unity, prefix to the index of its logarithm a figure less by one than the index of the power, and divide the whole by the index of the power, the quotient will be the logarithm of the root sought.

	-			
EXAMPLE I	. 1	EXAMPLE	II.	
What is the square root of	7 394 ?	Required the cube root of 2197?		
324 log.	2)2.51055	• 2197 log.	3)3.34183	
Answer 18 log.	1.25527	Answer 13	log. 1.11394	
EXAMPLE II	I.	EXAMPLE	IV.	
· Required the square root	of 40,96?	Required the cube root	of 0,0156 25 ?	
40,96 log.	2)1.61236	0,015695 log.	8.19382	
Answer 6,4	log. 0.80618	Prefix 2 to the index	3)28.19392	
•		Answer 0,25	log. 9.39794	
MO WORK MITT	TITE			

TO WORK THE RULE OF THREE BY LOGARITHMS.

When three numbers are given to find a fourth proportional in arithmetic, we make a statement and say, as the first number is to the second so is the third to the fourth; and by multiplying the second and third together, and dividing the product by the first, we obtain the fourth number sought. To obtain the same result by logarithms, we must add the logarithms of the second and third numbers together, and from the sum subtract the logarithm of the first number, the remainder will be the logarithm of the sought fourth number.

•			J J		-
EXAMPLE I.		•	EXAMPLE	II.	
If 6 yards of cloth cost 5	dolla	rs. what	If a ship sails 20 miles	in 7 hot	ırs. how
will 20 yards cost?		,	much will she sail in 21	hours at 1	the same
As 6	log.	0.77815	rate?		
Is to 5	log.	0.69897	As 7	log.	0.84510
So is 20	log.	1.30103	. Is to 20	log.	1.30103
Sum of 2d. and 3d.		8.00000	. So is 21	log.	1.32222
Subtract first		0.77815	Sum of 2d. and 3d.		2.62325
To 16.67	log.	1.22135	Subtract the first		0.84510
The answer therefore is 1				log.	1.7781.5
67-100ths or 16 dollars and 67	7 cent	s.	.The answer is	1	60 miles.

^{*} In this rule it is supposed that 10 was borrowed in finding the index of the decimal according to the rule page 30.

. To calculate COMPOUND INTEREST by Logarithms.

To 100 dollars add its interest for one year; find the logarithm of this sum and reject 2 in the index, then multiply it by the number of years and parts of a year, for which the interest is to be calculated; to the product add the logarithm of the sum put at interest; the sum of these two logarithms will be the logarithm of the amount of the given sum for the given time.

EXAMPLE.

Required the amount of the principal and interest of 355 dollars, let at 6 per cent. compound interest for 7 years?

Adding 6 to 100 gives 106, whose logarithm, rejecting 2 in the index, is

0.02531

tum, rejecting z in the index Multiplied		7
Product Principal 355 dollars	log.	0.17717 2.55023
the log. of 533,83	log.	2.72740

Therefore the amount of principal and interest is 533 dollars and 83 cents.

Sum gives

To find the Logarithm-sine, Tangent, or Secant, corresponding to any number of Degrees and Minutes, by Table XXVII.

The given number of degrees must be found at the bottom of the page when between 45° and 135°, otherwise at the top, the minutes being found in the column marked M, which stands on the side of the page on which the degrees are marked; thus, if the degrees are less than 45, the minutes are to be found in the left hand column, &c. and it must be noted that if the degrees are found at the top, the names of hour, sine, co-sine, tangent, &c. must also be found at the bottom, the names sine, co-sine, &c. must also be found at the bottom. Then opposite to the number of the minutes will be found the log-sine, log-secant, &c. in the column marked sine, secant, &c. respectively.

EXAMPLE I.

EXAMPLE II.

Required the log. sine of '28° 37'?

Required the log. secant of 126° 20'?

Find 38° at the top of the page, directly below which, in the left hand column, find 37'; against which in the column marked find 20'; against which, in the column marked is 9.68029, the log, sine of the given column of the given column of the given column of the page, discount of the

To find the Logarithm-sine, Co-sine, &c. for Degrees, Minutes, and Seconds, by Table XXVII.

Find the numbers corresponding to the even minutes next above and below the given degrees and minutes, and take their difference; then say, as 60" is to the number of seconds in the proposed number, so is that difference to a correction to be applied to the number corresponding to the least number of degrees and minutes; additive if it is the least of the two numbers taken from the table, otherwise subtractive.

EXAMPLE I.

EKAMPLE II.

Required the log. sine of 24° 16′ 48″?
Sine of 24° 16′ 9.61382 Secant of 105.20 log. 10.67768
Sine of 24 17 9.61411 106.21 105.2722

Diff. 29 Diff. 46

Then, as 60": 48":: 29: 23, which, added Then as 60": 16":: 46: 12, which, subtracted to the number corresponding to 24° 16', from the number corresponding to 105° 20', gives 2.61405 the log. sine of 24° 16' 48". gives 10.57756, the log. sec. of 105° 20' 16".

If the given seconds be $\frac{1}{4}$, $\frac{1}{4}$, $\frac{1}{4}$, or $\frac{1}{6}$, or any other even parts of a minute, the like parts may be taken of the difference of the logarithms, and added or subtracted as above, which may be frequently done by inspection.

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To find the Degrees, Minutes, and Seconds, corresponding to any gives Logarithm-sine, Co-sine, &c. by Table XXVII.

Find the two nearest numbers to the given logarithm-sine, co-sine, &c. in the column marked sine, co-sine, &c. respectively, one being greater and the other less, and take their difference; take also the difference between the given logarithm and the logarithm corresponding to the least number of degrees and minutes: then say, as the first found difference is to the second difference, so is 60" to a number of seconds to be annexed to the smallest number of degrees and minutes before found.

EXAMPLE L

Find the degrees, minutes, and seconds (less than 90°) corresponding to the log. sine 9.61405.

 Next less log. 94° 16′
 9.61382
 Log. of least numb. 24° 16′ is
 9.61382

 Greater
 24° 17°
 9.61411
 Given log.
 9.61405

29
Then say, as 29: 23:: 60": 48" which annexed to 24° 16' give 24° 16' 48" answering to log. sine 9.61405. Subtracting 24° 16' 48" from 180°, there remains 155° 43' 12', the log_sine of which is also 9.61405.

EXAMPLE II.

Find the degrees, minutes, and seconds (above 90°) corresponding to the log secant 10.56703.

Secant 105° 43' log. 10.56722 Log. of the least numb. 105° 43' 10.56722 Secant 105 44 10.56677 Given log. 10.56703

Then as 45 is to 19, so is 60" to 25", which annexed to 105° 43', gives 105° 43' 25", the degrees, minutes, and seconds required.

To find the Arithmetical Complement of any Logarithm.

The arithmetical complement of any logarithm, is what it wants of 10,00000 and is used to avoid subtraction. For when working any proportion by logarithms, you may add the arithmetical complement of the logarithm of the first term, instead of subtracting the logarithm itself, observing to neglect 10 in the index of the sum of the logarithms. The arithmetical complement of any logarithm is thus found:—Begin at the index, and write down what each figure wants of 9, except the last significant figure, which take from 10*.

EXAMPLE.

Required the arithmetical complement of 9.62595?

For the first figure 9, write 0; for 6, 5; for 2, 7; for 5, 4; for 9, 0; and for the last figure 5, write 5; thus the arithmetical complement is 0,57405.

In the same manner the arithmetical complement of 1.86563 is 8.18437, the ar. co. of 10.33133 is 9.66867, and the ar. co. of 1.22800 is 8.77200. To illustrate the method of using the arithmetical complement of any logarithm, I shall here calculate the examples as given in page 32.

As 6 Is to 5 So is 20	EXAMPLI log. ar. co. log. log.		9.22185 As 0.69897 Is 1.30103 So	to 20	EXAMPLE II. log. ar. co. log. log.	9.15490 1.30103 1.32222
To 16,67	log.	•	1.22185 To	60	log.	1.77915

[•] When the index of the given logarithm is greater than 10, as in some of the numbers of Table XXVII. the left hand figure of it must be neglected; and when there are any ciphers to the right hand of the last significant figure, you may place the same number of ciphers to the right hand of the other figures of the arithmetical complement.

PLANE TRIGONOMETRY.

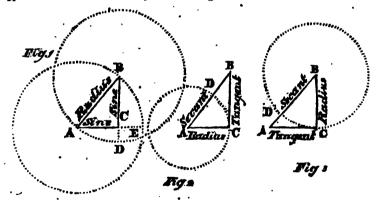
PLANE TRIGONOMETRY is the science which shows how to find the measures of the sides and angles of plane triangles, some of them being already known. It is divided into two parts, right-angled and oblique-angled: in the former case, one of the angles is a right angle or 90°; in the latter

they are all oblique.

In every plane triangle there are six parts, viz. three sides and three angles; any three of which being given (except the three angles) the other three may be found by various methods, viz. by Gunter's Scale, by the sliding rule, by the sector, by geometrical construction, or by arithmetical calculation. We shall explain each of these methods; but the latter is by far the most accurate; it is performed by the help of a few theorems, and a trigonometrical canon, exhibiting the natural or the logarithmic sines, tangents, and secants, to every degree and minute of the quadrant. The theorems alluded to are the following:

THEOREM I.

In any right angled triangle, if the hypotenuse be made radius, one side will be the sine of the opposite angle, and the other its co-sine; but if either of the legs be made radius, the other leg will be the tangent of the apposite angle, and the hypotenuse will be the secant of the same angle.



1st. If in the right-angled plane triangle ACB (fig. 1) we make the hypotenuse AB radius, and upon the centre A describe the arch BE, to meet AC produced in E; then it is evident that BC is the sine of the arch BE (or the sine of the angle BAC) and that AC is the co-sine of the same angle: and if the arch AD be described about the centre B, AC will be the sine of the angle ABC, and BC its co-sine.

2dly. 'If the leg AC (fig. 2) be made radius, and the arch CD be described about the centre A; CB will be the tangent of that arch, or the tangent of the

angle CAB; and AB will be its secant.

Sdly. If the leg BC (fig. 3.) be made radius, and the arch CD be described about the centre B; CA will be the tangent of that arch, or the tangent of

the angle B; and AB will be its secant.

Now it has been already demonstrated (in art. 55. Geom.) that the sine, tang. sec. &c. of any arch in one circle, is to the sine, tang. sec. &c. of a similar arch in another circle as the radius of the former circle to the radius of the latter. And since in any right-angled triangle there are given either two sides, or the angles and one side, to find the rest; we may, if we wish to find

[•] It will not be necessary to add any further description of the uses of the sector or sliding rule, for what we have already given will be sufficient for any one, tolerably well versed in the use of Gunter's Scale.

† See Twoles XXIV. and XXVII.

a side, make any side radius; then say, as the tabular number of the same name as the given side is to the given side of the triangle, so is the tabular number of the same name as the required side, to the required side of the triangle. If we wish to find an angle, one of the given sides must be made radius; then say, as the side of the triangle made radius, is to the tabular radius, so is the other given side to the tabular sine, tangent, secant, &c. by it represented; which being sought for in the table of sines, &c. will correspond to the degrees and minutes of the required angle.

THEOREM II.

In all plane triangles, the sides are in direct proportion to the sines of their opposite angles (by art. 58, Geom.)

Hence, to find a gide, we must say, as the sine of an angle is to its opposite. side, so is the sine of either of the other angles to the side opposite thereto. But if we wish to find an angle, we must say, as any given side is to the sine of its opposite angle, so is either of the other sides to the sine of its opposite angle.

THEOREM III.

In every plane triangle, it will be, as the sum of any two sides is to their difference, so is the tangent of half the sum of the two apposite angles to the tangent of half their difference (by art. 59, Geom.)

THEOREM IV.

As the base of any plane triangle is to the sum of the two sides, so is the difference of the two sides to twice the distance of a perpendicular (let fall upon the base from the opposite angle) from the middle of the base (by art. 60, Geom.)

THEOREM V.

In any plane triangle, as the rectangle contained by any two sides including a sought angle, is to the rectangle contained by the half sum of the sides and the half sum decreased by the other side, so is the square of radius to the square of the co-sine of half the sought angle (by art. 61, Geom.)

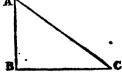
[In addition to these theorems. it will not be amiss for the learner to recall to mind the following articles, must of which have been already demonstrated.]

In every triangle, the greatest side is opposite to the greatest angle; and the greatest angle opposite to the greatest side. 2. In every triangle equal sides subtend equal angles. (Art. 39, Geom.)

The three angles of any plane triangle are equal to 1800. (Art. 35,

Geom.)

- If one angle of a triangle be obtuse, the rest are acute; and if one angle be right, the other two together make a right angle, or 900; therefore, if one of the acute angles of a right-angled triangle be known, the other is found by subtracting the known angle from 90°. If one angle of any triangle be known, the sum of the other two is found by subtracting the given angle from 1800; and if two of the angles be known, the third is found by subtracting their sum from 180°.
 - 5. The complement of an angle is what it wants of 90°, and the supplement of an angle is what it wants of 1800.



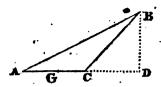
In the two following tables we have collected all the rules necessary for solving the various cases of Right-Angled and Oblique-Angled Trigonometry.

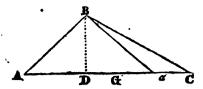
RIGHT-ANGLED TRIGONOMETRY.

Case.	Given.	Sought.	Solutions.
1	Hyp. AC. Angles.	Leg BO. Leg AB.	Rad.: hyp. AC.:: sine A.: leg RC. Rad.: hyp. AC.:: sine C.: leg AB.
2 & 3	Leg BO. Angles.	Log AB. Hyp. AC.	Rad. : leg BO. : : tang. C. : leg AB. { Rad. : leg BC. : 2 acc. C. : hys. AO. { Or Sine A : leg BC. : c rad. : hyp. AO.
44.5	Hyp. AC. Leg AB.	Angles. Leg BC.	Hyp. AC.: rad.: : leg AB.: sine C. whose comp. is A. Rad.: hyp. AC.: sine A.: leg. BC.
6	Both legs. AB & BC.	Angles. Hyp. AC.	Leg BC. : rad. : leg AB. : tang. C. whose comp. is A. Sine C. : leg AB. : : rad. : hyp. AC

E

B





Case	Given.	Sought.	. Solutions.
1	The angles and side A B.	Side BC. Side AC.	Sine C : side AB : : sine A : side BC. Sine C : side AB : : sine B : side AC.
243	Two sides A B, BC, and angle C opposite to one of them.	Angle A. Angle B. Side AC.	Side AB : sine C : side BC : sine A, which added to C, and the sum subtracted from 180° gives B. Sine C : side AB : sine B : side AC.
445	Two sides A C, A B and the in- cluded angle A	and B.	Subtract half the given angle A from 800, the remainder is half the sum of the other angles. Then say, as the sum of the sides AO, AB is to their difference, so is the tangent of the half sam of the other angles to the tangent of half their difference; which added to and subtracted from the half sum, will give the two angles B and C the greatest angle being opposite to the greatest side. Sine B: side AC: : sine A: side BC.
6	All three sides	gles.	Let fall a perpendicular BD opposite to the required angle; then as AC: sum of AB, BC: their difference: twice DG, the distance of the perpendicular from the middle of the base; hence, AD, CD are known, and the triangle ABC is divided into two rights angled triangles BCD, BAD, then by cases 4 and 5 of right-angled trigonometry, we may find the hogle A or C.
		}	Either of the angles, as A, may also be found by the following rule. From half the sum of the three-sides subtract the side BC opposite to the sought angle; take the logarithms of the half sum and remainder, to which add the arithmetical complements of the logarithms of the sides AB, AC (including the sought angle;) half the sum of these four logarithms will be the logarithmic co-sine of balf the sought angle.

In working by logarithms with any of the preceding rules, you must remember, that the legarithm of the first term of the enclose is to be subtracted from the man of the logarithms of the second and third terms, the remainder will be the logarithm of the second fourth term. When the first term is radius (whose logarithm is 10,0000) you need only reject an unit in the second left hand figure of the index of the sum of the second and third terms. But when the radius occurs in the second or third term, you must suppose a lumit to be added to the second left hand figure of the index of the other term, and subtract therefrom the logarithm of the first term.

RIGHT-ANGLED TRIGONOMETRY.

Solution of the six cases in right-angled Trigonometry.

CASE I.

The angles and hypotenuse given, to find the legs.

· Given the hypotenuse AC 250 leagues, and the angle C opposite to the side AB=35° 30' to find the base CB and perpendicular AB.

BY PROJECTION.

Draw the base CB of any length; with an extent equal to the chord of 60° and on C as a C centre, describe the arch DE; from E to D lay off 35° 30' taken from the line of chords,*

through C and D draw the line AC, which make equal to 250; from A let fall the perpendicular AB, to cut CB in B, and it is done; for CB will be 203.5, and AB=145.2.

⁴ In all projections of this kind the angles are measured from the line of chords, the radius used for describing arches by which the angles are to be measured, being equal to the chord of 66°, the sides of the triangles are measured by scales of equal parts as was before observed. Digitized by

BY LOGARITHMS. •

By making the hypotenuse CA radius, it will be,

i radius, it was so, To find the perpendicular Alla. 10.00000 To find the base BC. 10.00000 As radius As radius Is to the hypot. AC 250 2.39794 is to the hypot. AC 250 So is the sine ang. A 540 30' 9.91069 So is sine ang. C 35° 30'

2.30963 To the per. AB 145.2 To the base BC 203,5 2.16189

BY GUNTER'S SCALE.

In all proportions wrought by Gunter's Scale, when the first and second terms are of the same kind, the extent from the first term to the second, will reach from the third sthe fourth.

Or when the first and third terms are of the same kind.

The extent from the first term to the third will reach from the second fo the fourth; that is, set one point of the compasses on the division expressing the first term, and extend the other point to the division expressing the third term; then, without altering the opening of the compasses, set one point on the division representing the second term, and the other point will fall on the division showing the fourth term or answer.

In the present example the work will be as follows:

Extend from radius or 90°, to 54° 30° on the line of sines; that extent will reach from 250, the hypotenuse, to 203.5, the base on the line of numbers; and the extent from radius, or 90°, to 35° 30' on the line of sines, will reach from 250 to 145.2 on the line of numbers.

Observe the like in all that follows, except in those proportions where the word secant is mentioned, which cases must be wrought by considering the hypotenuse radius,* there being no line of secants on the common Gunter's Scale, although it might easily be marked on the line of sines.

Note. The radius, according to the nature of the proportion, may be either of the following quantities.

8 points on the line of rhumbs. 4 points on the line of tan-rhumbs. 90° on the line of sines. 450 on the line of tangents.

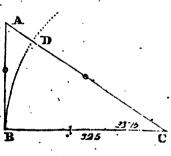
CASES II. AND III.

The angles and one leg given, to find the hypotenuse and other leg.

The angle ACB 55° 15', the leg BC 525 miles, given, to find the hypotenuse and the other leg.

BY PROJECTION.

Draw the line BC, which make equal to 325 miles; on B erect the perpendicular BA; on C, as a centre, with the chord of 600 sweep the arch BD, which make equal to 53° 15'; draw CD, and continue it to cut AB in A, and it is done; for AB being B measured on the same scale that BC was, will be 213.1, and AC 388.6 miles.



2.39794

9.76395

BY LOGARITHMS.

By making the base BC radius, it will be,

To find the perpendicular AB. To find the hypotenuse AC. 10.00000 As radius 900 As radius 45º 10.00000 Is to the base BC 325 2.51188 is to the base BC 325 2.51188 So is tang. eng: C 33° 15' 9.81666 So is sec. ang. C 33° 15' 10.07765 To the perpen. AB 213.1 2.32854 To the hypot. AC 388.6 2.58953 BY GUNTER.

Extend from 45° to 85° 15' on the line of tangents; that extent will reach

from the base 325 to the perpendicular 213,1 on the line of numbers.

2dly. Extend from 56° 45' to radius on the line of sines; that extent will reach from the base 325 to the hypotenuse 388,6 on the line of numbers.

Or by using in the analogy radius: cos. angle instead of secant angle: Radius and radius: sine angle instead of co-secant angle: Radius. Digitized by GOOGIC

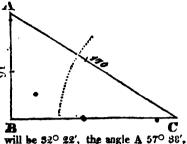
CASES IV. AND V.

The hypotenuse and one leg given, to **S**ad the angles and other leg.

The leg AB 91, the hypotenuse AC 170, given, to find the angle ACB, or BAC, and the leg BC.

BY PROJECTION.

Draw BC at pleasure; on B erect the perpendicular BA, which make equal to 91; take 170 in your compasses, and with one foot on A, describe an arch to cut BC in C; join A 🗷 and C, and it is done; for the angle C will be 52° 22', the angle A 57° 88', and BC 143,6.



BY LOGARITHMS.

By making the hypotenuse radius it will be,

To find the angle C. As the hypotenuse 170 Is to radius So is the perpendicular 91

To sine angle C 32° 22'

To find the base BC.* 2.23045 As radius 10.00000 2,23045 10.00000 is to the hypotenuse 170 9.92667 1.95904 So is sine ang. A 570 38'

9.72859 To the base BC 143.6

2.15712

BY GUNTER.

Extend from the hypotenuse 170 to the perpendicular 91 on the line of numbers; that extent will reach from radius to the angle C, or the complement of angle A=320 22' on the line of sines.

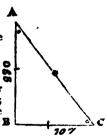
2dly. Extend from radius to the angle A 57° 38' on the line of sines; that extent will reach from the hypotenuse 170 to the base 143.6, on the line of numbers.

CASE VI.

The legs given, to find the angles and hypotenuse. The legs AB 890, BC 707, given, to find the angle BAC or ACB, and the hypotenuse AC.

BY PROJECTION.

Make BC=707, and on B erect the perpendicular BA, which make equal to 890; join AC, and it is done; for the angle C will be 510 32'; consequently the angle A 38° 28', and the hypotenuse 1137.



To find the hypoteruse AC.

BY LOGARITHMS.

By making the base radius, it will be, 2.84942 As radius

To find the angle C. As the base 707

Is to radius So is the perpendicular 890

2.94939 So is the sec. ang. C 51° 32 10.09997 To the hyp. AC=1137

10.00000 is to the base 707

2.84942 10.20617

10.00000

To tan. ang. C=51° 32'

3.05559

BY GUNTER.

The extent from 707 to 890 on the line of numbers will reach from radius, or 45 degrees, to the angle C 51° 32' on the line of tangents.

2dly. The extent from the angle C 51° 32' to radius, or 90°, on the line of sines, will reach from the base 890 to the hypotenuse 1137, on the line of numbers.

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When you take the log, sines, or tangents to the nearest minute only, it is best to use this canon for finding BC, which is more correct than the one found by making the perpendicular radius; because the variation of the log, sine of an arch is less than the corresponding variation of the log, tangent.

† When you are finding AC it is best to make the greatest side radius, for the reason mentioned in

QUESTIONS

To exercise the learner in Right-Angled Plane Trigonometry. ...

Quest. 1. The hypotenuse 498 miles, and the angle opposite to the base 56° 15', given, to find the base and perpendicular.

Ans. Base 412.4, and the perpendicular 275.6 miles.

Quest. 2. The perpendicular 275 leagues, and the angle opposite to the base 56° 15', given, to find the hypotenuse and base.

Ans. The hypotenuse 495, and base 411.6 leagues.

Quest. 3. The base 33 yards, and the angle opposite to the perpendicular 53° 26', given, to find the hypotenuse and perpendicular.

Ans. Hypotenuse 55.39, and the perpendicular 44.49 yards.

Quest. 4. The hypotenuse 575, and perpendicular 50 miles, given, to find the base.

Ans. Base 572.8 miles.

Quest. 5. The hypotenuse 59, and the base 35 miles, given, to find the perpendicular.

Ans. Perpendicular 48.9 miles.

Quest. 6. The base 33, and perpendicular 52 leagues given, to find the hypotenuse.

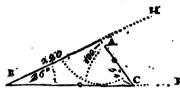
Ans. Hypotenuse 61.59 leagues.

OBLIQUE TRIGONOMETRY.

CASE I.

Two angles and one side given, to find either of the legs.

Given the angle BAC=100° the angle ACB=540 and the leg AB=220 to find the sides.



BY PROJECTION.

Subtract the sum of the angles A and C from 1800, the remainder will be the angle B=260. Draw the indefinite line BE, also the line BH, making the angle EBH=250, on BH set off BA 220. On A make the angle BAC 1000: then AC will intersect the line BE in the point C, which completes the triangle, and BC will measure (on the same scale from which BA was laid down) 268 nearly, and AC 119.

BY LOGARITHMS by Theorem II

DI LIUU	Titilities of theorem if.		
. To find BC.	To find DC.		
As the sine of the angle C 54°	9.90796 As sine ang; C 54°		9.90796
Is to the side Al 220	2.34242 Is to the side AB 220		2.34242
So is the sine of the ang. A 100°	9.99335 So is the sine ang. B 26°	•	9.64184
	12.33577		11.98426
	9.90796		9.90796
To the side BC 267.8	2.42781 To side AC 119.2		2.07630
•	BY GUNTER.		,•

The extent from the angle C=540 to the angle A or its supplement 800 on the sines, will reach from AB=220 to BC=268 on the line of numbers.

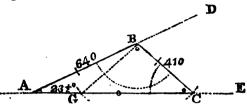
2dly. The extent from the angle $C=54^{\circ}$ to the angle $B=26^{\circ}$, on the sines, will reach from AB=220 to AC=119 on the line of numbers.

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CASES II. AND III.

Two sides and an angle opposite to one of them being given, to find the other angles and the third side.

NOTE. It may be proper to observe, that if the given angle be obtuse, the angle sought will be acute: but when the given angle is acute and opposite to a K lesser given side, then it is doubtful



whether the required angle is acute or obtuse; it ought therefore to be given

by the conditions of the problem.

EXAMPLE. Let there be given the side BC 410, the side AB 640, and the angle A 2340, to find the other side AC, and the angles ABC, BCA.

BY PROJECTION.

Draw the indefinite line FE, make the angle DAE = 234°, on AD set off AB=640, then on B, with 410 in your compasses taken from the same scale, describe an arch cutting FE in the points C and G, join BC, BG, and it is done; for the triangle may be either ACB, or AGB, according as the angle C, or G, is acute or obtuse; if that angle be acute, the triangle will be ABC; the side AC will measure 908, the angle ACB will measure 5810, and the angle ABC will measure 1180 nearly; but if the angle at the base be obtuse, the triangle will be AGB; the side AG will measure 266, the angle AGB will measure 14140 and the angle ABG 150.

If the side BC had been given greater than AB, there could have been only one answer to this problem; for in that case, the point G would have fallen on the continuation of the line CA towards F, in which case the angle A of the triangle would become equal to FAB, instead of being equal to its supple-

ment, as is required by the conditions of the problem.

		BY	LOG	3 A	RITHM	IS, by Theorem II.	
To f	nd t	he angle	C or	G.		To find AC.	
	As the side BC 410					As sine angle C 38° 30'	9.79410
Is to the sine	of a	ngle A	2310		9.60070	Is to AB 640	2.80615
So is the side			•		2.80618	So is sine angle ABC 1185	9.94593
			•		12.40688		12.75211
					2.61278		9.79410
To sine ang.			141 23		9,79410	To the side AC 907,8	2.95801
				_		To find AG.	
Subtract	62	0 or	165	0		As sine angle G 141° 30'	9.79410
From	180	0	180	0		Is to AB 640	2.80618
				_		So is sine angle ABG 15°	9.41300
Ang. ABC	113	O AB	G 15	()	•	l	
						į.	12.21919
							9.79410
						Fo the side AG 266,1	2.42503
					RV GI	INTER.	

1st. The extent from BC=410 to AB=640, on the line of numbers, will reach from A=2340, to 3840, on the line of sines, which is equal to the angle. C; its supplement 141° 30' being equal to the angle G.

2dly. The extent from the angle C=38° 30' to 62° 0' (the supplement of the angle ABC, 118° 0') on the sines, will reach from AB=640 to 908, on the line of numbers; therefore the side AC=908.

Or, the extent from 38° 30' (the supplement of the angle G) to the angle ABG=150 0' on the sines, will reach from AB=640 to 266, on the line of numbers: hence AG = 266.

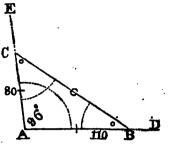
CASES IV. AND V.

Two sides and their contained angle being given, to find either of the other angles and the third side.

Given the side AB 110 m. AC 80 m. and angle BAC 96° 0′ to find the angles BCA and CBA.

BY PROJECTION.

Draw the indefinite right line AD, on which set off AB=110; make the angle EAB=96°; and on AE set off AC=80; join BC, and it is done; for BC will measure on the former scale 143, and the angles B and C will measure 55° 55′ and 50° 5′ respectively on the line of chords.



BY LOGARITHMS.

To find the angles B an As sum of sides AC and A Is to their difference 30 So is tang. ½ sum op. <'s or comp. ½ ang. Å.	18 190	2.27875 1.47712	To find the si As sine ang. B 33 Is to AC 80 So is sine ang. A or its sup.	° 55′	by Theor	rem II. 9.74662 1.90309 9.99761
		11.43156 2.27875		•		11.90070 9.74662
To tang. half diff.	8º 5'	=9.15281	To side BC 142.6			2.15408
Sum is angle C Diff. is angle B	50 5 33 55	•	•			

BY GUNTER.

1st. The extent from the sum of the sides 190 to their difference 30 on the line of numbers, will reach from the half sum of the angles B and C 42° to their half difference 8° 5′ on the line of tangents. The sum of which half sum and half difference gives the angle C 50° 5′ and their difference the angle B 33° 55′; the greatest angle being opposite to the greatest side.

2dly. The extent from the angle B 38° 55', to the angle A 96° (or its supplement 84°) on the line of sines, will reach from the side AC 80 to the

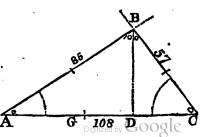
side BC 142.6 on the line of numbers.

CASE VI.

The three sides of a plane triangle given, to find the angles.

The sides AB 85, BC 57, AC 108 given, to find the angles ABC, BAC, BCA.
BY PROJECTION.

Draw the line AC, and make it equal to 108; take 85 in your compasses, and with one foot on the point A, describe an arch; then take the distance 57 in your compasses, and with one foot on C, describe another arch intersecting the former arch in the point B; join AB, CB, and it is done. For the angle A being measured will be found \$3140, B=970, and the angle C=5140 nearly.



BY LOGARITHMS, by Theorem IV.

Suppose BD to be drawn perpendicular to AC, and that AG = GC.

	Side AB=85 Side BC=57	As the base AC 108 Is to the sum of the sides AB and BG 142 So is diff. of the sides AB and BC 28		2.03342 2.15229 1.44716
Sum of the sides Diff. of the sides	143 28		•	3.59945 2.03342
Half base AC DG	54 18.4	To twice DG Its half=DG	36,9 18,4	1.56603
Sum is greater seg. Diff. is least seg. DO			,-	

Having divided the triangle into two right-angled triangles, the hypotenuses and bases of which are given, we may find the angles by Theorem I.

		To find the angle BCD.		
		As the hypotenuse BC 57	1.75587	
		Is to radius 90°	10.00000	
To find the angle BAD.		So is the lesser seg. DC 35,6	1.55145	
As the hypotenuse AB 85	1.92942	,		
Is to radius 90°	10.00000	To co-sine of BCD=51° 21'	9.79558	
So is the greater seg. AD 72,4	1.85974	BAD=31 36		
				
To co-sine BAD=31° 36'	9.93032	Sum 82 57		
		· Subtract from 180		
		Remains angle ABC 97 3		

BY GUNTER.

1st. The extent from the base AC=108, to the sum of the sides 142, on the line of numbers, will reach from the difference of the sides 28, to twice DG 36.8 on the same line of numbers.

· 2dly. The extent from the hypotenuse AB=85, to the greater segment AD 72.4, on the line of numbers, will reach on the sines from the radius 90° to 58° 24' which is the complement of the angle BAD.

to 58° 24' which is the complement of the angle BAD.

3dly. The extent from the hypotenuse CB 57, to the lesser segment DC 55.6 on the line of numbers, will reach on the sines from the radius 90° to 58° 59', which is the complement of the angle BCA.

This case may be solved without dividing the triangle into two right-angled triangles by Theorem V.

To find the angle A. $BC = 57$		Having the angle A, we may find C by Theorem II.	the angle
AB= 85 log. co. ar.	8.07058	As BC 57	1.75587
AC=108 log. co. ar.	7.96658	Is to sine angle A 31° 36'	9.71938
•		So is AB 85	1.92942
Sum 250			
Half sum 125 log.	2.09691	1	11.64874.
Half sum less BC 68 log.	1.83251		1.75587
Sum	19.96658	To the sine of angle C 51° 23'	9,99287
Half sum 15° 48′ Co-sine 2	9.98329		•
Doubled is 31 36=angle A.		1	•_

ASTRONOMY AND GEOGRAPHY.

ASTRONOMY is the science which treats of the motions and distances of the heavenly bodies, and of the appearances thence arising.

Geography is the science which treats of the situations and distances of

the various parts of the surface of the earth.

The common opinion of astronomers of the present day is, that the universe is composed of an infinite number of systems or worlds; that in every system there are certain bodies moving in free space, and revolving at different distances round a sun, placed in or near the centre of the system; and that these suns and other bodies are the stars which are seen in the heavens.

The SOLAR SYSTEM, so called, is that in which our earth is placed, and in which the sun is supposed to be fixed near the centre, with several bodies, similar to our earth, revolving round at different distances. This hypothesis, which is fully confirmed by observation, is called the Copernican System, from Nicholas Copernicus, a Polish Philosopher, who revived it about the

year 1500, after it had been buried in oblivion many ages.

Stars are distinguished into two kinds, fixed and wandering. The former are supposed to be suns in the centres of their systems, shining with their own light, and preserving nearly the same situation with respect to each other. They are usually distinguished by their brightness, the largest being called of the first magnitude, and the smallest visible to the naked eye being of the sixth or seventh magnitude. A Constellation is a number of stars lying in the neigbourhood of one another on the surface of a celestial sphere, which astronomers, for the sake of remembering with greater ease, suppose to be circumscribed by the outlines of some animal or other figure. The wandering stars are those bodies within our system, or celestial sphere, which revolve round the sun; they appear luminous by reflecting the light of the sun, and are of three kinds, namely primary planets, secondary planets, and comets.

The Primary Planets are bodies which revolve round the sun, as the centre of their courses, the motions being regularly performed in tracks or paths (called orbits) that are nearly circular and concentric with each other. A Secondary Planet, Satellite, or Moon, is a body which, while it is carried round the sun, revolves also round a primary planet. Comets are a sort of planets moying round the sun in very eccentric orbits, with vast atmospheres about

them, and tails derived from the same.

There are eleven primary planets, which, reckoned in order from the sun, are as follows: Mercury, Venus, the Earth, Mars, Vesta, Juno, Pallas, Ceres,

Jupiter, Saturn, and Uranus.

Mercury and Venus are called inferior planets, because their orbits are within the earth's; the others are called superior planets, as their orbits include the coath

clude that of the earth.

The Sun, the first and greatest object of astronomical knowledge, is placed near the centre of the orbits of all the planets, and turns round its axis in 254 days; its diameter is 885,000 English miles, and its mean distance from the earth 95 millions of miles.

MERCURY is the least of all the planets, known before the discovery of Vesta, Juno, Pallas, and Ceres, and is the nearest to the sun, his mean distance from that luminary being 37 millions of miles. His periodic revolution in his orbit round the sun is performed in 87 days 23 hours, and his diameter is about 3200

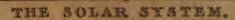
miles.

VENUS is the brightest of all the planets. Her diameter is 7687 miles; her mean distance from the sun 68 millions of miles; and her periodic revolution is performed in 224 days 17 hours. When this planet is in that part of her orbit which is west of the sun, she rises before him in the morning, and is called the morning star; when she is in the eastern part of her orbit, she shines in the evening, after he sets, and is called the evening star.

The next planet is the EARTH, the diameter of which is 7964 miles, the distance from the sun 95 millions of miles, and the time of revolution round

PRINCISHED BY EDM M. PLANT, 1976.

Plate 1



the sun one year. The earth turns round its axis from west to east in 23h. 56m. which occasions the apparent diurnal motion of the sun and all the heavenly bodies round it from east to west in the same time, and is, of course, the cause of their rising and setting, of day and night. The axis of the earth is inclined about 25° 28' to the plane of its orbit, and keeps nearly in a direction parallel to itself, throughout its annual course, which causes the return of spring and summer, autumn and winter. Thus the diurnal motion gives us the grateful vicissftude of night and day, and the annual motion the regular succession of the seasons. The earth is attended by a satellite called the Moon, whose diameter is 2161 miles; her distance from the centre of the earth is 240,000 miles: she goes round her orbit in 27 days 8 hours; but, reckoning from change to change, in 294 days. Her orbit is inclined to the ecliptic in an angle of 50 9', cutting it in two points diametrieally opposite to each other, called her nodes. As the moon shines only by the reflected light of the sun, she must appear different when in different situations with respect to that luminary. When she is in conjunction with the sun, her dark side is turned towards the earth, which renders her invisible; this is called new moon: when she is in opposition, her light side is wholly visible from the earth; this is called full moon.

If at the time of new moon she is near to either of her nodes, she may intercept a part of the sun's light, and thus cause an eclipse of the sun; and if she is near either of her nodes at the time of full moon, she may pass into the shadow of the earth, and cause an eclipse of the moon. In a similar manner, when the moon passes between an observer on the earth and a star, it is called an Occultation of the star. The instant when the moon's limb first covers the star is called the Immersion, and the moment of its re-appearance is called the Emersion. When Mercury or Venus passes between the sun and an observer, and appears to pass over the sun's disk, it is called a Transit of Mercury or Venus. Eclipses, Occultations and Transits, are of great importance in ascertaining the longitudes of places on the earth. Eclipses of the moon furnish a convincing proof of the rotundity of the earth, since the shadow of the earth, seen upon the moon when eclipsed, is always circular. This is farther confirmed by the appearance of objects at sea; for when a ship is making towards the land, the mariners first descry the tops of steeples, trees, &c. pointing above the water; the lower parts being hid, by reason of

the curvature of the earth.

The earth is not a perfect globe or sphere, but is a little flatted at the poles, being nearly of the figure of an oblate spheriod, the equatorial diameter being about 26 miles longer than the polar: but since this difference bears but a small comparison to the whole diameter, we may, for all the practical purposes of navigation, consider the earth as a perfect sphere, as will be done in the rest of this work. The natural divisions of the earth will be given hereafter.

MARS is the next planet to the earth; his diameter is 4189 miles, his distance from the sun 144 millions of miles, and his periodic revolution is performed in about 687 days. He revolves round his axis in 24 hours 40 minutes, appearing of a dusky reddish hue, and is supposed to be encompassed

with a very great atmosphere.

Between Mars and Jupiter are situated the four lately discovered planets, Vesta, Juno, Pallas, and Ceres, named Asteroids by Doctor Herschel. The elements of their orbits have not been accurately ascertained, but are nearly, as in the following description.

VESTA, was discovered by Doctor Olbers of Bremen, on the 29th of March, 1807. Its mean distance from the sun is about 205 millions of miles; its

periodic revolution is performed in about 3 1-6 years.

Juno, was discovered by Mr. Harding, of Lilienthal (near Bremen) on the first of September, 1804. It appears like a star of the eighth magnitude. Its distance from the sun is about 255 millions of miles; its periodic revolution is performed in 1582 days. The inclination of its orbit to the ecliptic is 15° 50′ and the eccentricity of the orbit 0.25.

The Inclination decreases at present about 80" in 100 years, by reason of the attraction of the planets on the earth. It is also affected by the Mutation given in Table XLIII, which sometimes amounts 16 9". In estimating the eccentricities of the planets, their mean distance from the gam is put equal to unity.

Its diameter. Pallas, was also discovered by Dr. Olbers, March 28, 1802. according to Doctor Herschel, is only 110 miles; it appears like a star of the eighth magnitude. Its mean distance from the sun is about 266 millions of miles. Its periodic revolution is performed in 1683 days. The inclination of its orbit to the ecliptic is 34° 39' and the eccentricity of the orbit 0.2468.

CERES, was discovered by Mr. Piazzi of Palermo on the first of January, 1801. Its diameter, according to Dr. Herschel, is only 160 miles. It appears like a star of the seventh or eighth magnitude. Its distance from the sun is about 266 millions of miles, and its periodic revolution is performed in 1683 days, being at nearly the same distance from the sun as Pallas. The inclination of the orbit of Ceres to the ecliptic is 10° 37' and the eccentricity 0.097. The situations of the nodes of the two planets Ceres and Pallas, and the inclinations of their orbits, are very different from each other, so that when those planets are in the same plane, they are at a great distance from each other, notwithstanding their mean distances from the sun are nearly equal. It has been supposed by some, that these small bodies are fragments of a former planet.

JUPITER is situated still higher in the system, and is the largest of all the planets, being easily distinguished from them by his peculiar magnitude and light. His diameter is 89,170 miles, his distance from the sun 490 millions of miles, and the time of his periodic revolution is 43321 days. Though Jupiter is the largest of all the planets, yet his diurnal revolution is the

swiftest, being only 9 hours and 56 minutes.

Jupiter is attended by four satellites, invisible to the naked eye, but through a telescope they make a beautiful appearance. In speaking of them, we distinguish them according to their places, into the first, second, and so on; by the first we mean that which is nearest to the planet. The appearance of these satellites is marked in the XIIth. page of the Nautical Almanac, for some particular hour of the night; the times when they are eclipsed, by passing into the shadow of Jupiter, are also given in the Nautical Almanac; these eclipses are of considerable use in determining the longitudes of places on the earth.

Before the discovery of the planet Uranus, Saturn was reckoned the most remote planet of our system. He shines with but a pale and feeble light. His diameter is 79,042 miles, his distance from the sun 900 millions of miles, and his periodic revolution in his orbit is performed in about 29 years 167 days. This planet is surrounded with a broad flat ring, has a diur-

nal revolution round its axis, and is attended by seven satellites.

By some observations made by Dr. Herschel, it appeared that the largest diameter of Saturn corresponds to the latitude of 450, but from later observations he has been induced to believe, that this irregularity is owing to an optical deception, arising from the refraction of the light in passing through

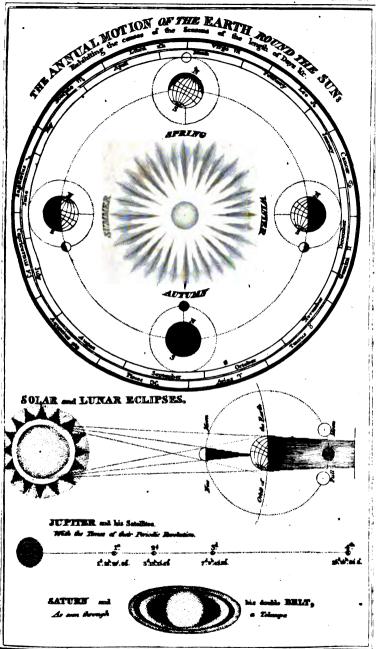
the atmosphere of the ring.

URANUS, Herschel, or Georgium Sidus, is the most remote planet of our system. It was discovered in the year 1781, by Dr. Herschel, though it had been seen several times, but had been considered as a fixed star. Its diameter is 35,109 miles, its distance from the sun 1800 millions of miles, and its periodic revolution in its orbit is performed in 83½ years. Dr. Herschel has

also discovered six satellites attending this planet.

The astronomy of comets is yet in its infancy. The return of one of them in the year 1758 was foretold by Dr. Halley, and it happened nearly as he predicted. He also foretold the return of another in the year 1790, but it never appeared. This was owing to the inaccuracy of the observations of the comet at its former appearance; for Mr. Mechain, having collected all the observations, and calculated the orbit again, found it to differ essentially from that determined by Dr. Halley. Olber's comet, which appeared in 1815, has a revolution of 72 years; and Encke's comet, which has been observed in several successive approaches to the perihelion, completes its revolution in the short period of 1204 days.

Comets move round the sun in all directions, but the planets and satellites, except one of the satellites of Uranus, move from west to east when seen from the sun; but if viewed from any other of the planets, as the earth, they would appear to revolve round it as a centre; but the sun would be the only one that moved uniformly the same way, for the other planets would some-



I JELISHED BY ELM.M.ELUNT. 1826.

times appear to move from west to east, and then to stand still; then they would seem to move from east to west; and after standing some time, they would again move from west to east; and so on continually. The motion of a planet from west to east is called the direct motion, or according to the order of the signs. The contrary motion from east to west, is called retrograde. When the planet appears to stand still, it is said to be stationary.

To illustrate what has already been said relative to the motions and distances of the planets and satellites, we have given the adjoining Plates III.

and IV. which require no explanation.

In noting the situations of the stars and planets, astronomers have been under the necessity of imagining various lines and circles on the sphere; and geographers have done the same for fixing the situation of places on the earth. The most remarkable of these are the following.

A great circle is that whose plane passes through the centre of the sphere;

and a small circle is that whose plane does not pass through that centre.

A diameter of a sphere, perpendicular to any great circle, is called the axis of that circle; and the extremities of a diameter are called its poles. Hence the pole of a great circle is 90° from every point of it upon the surface of the sphere; but as the axis is perpendicular to the circle when it is perpendicular to any two radii, a point on the surface of a sphere 90° distant from any two points of a great circle will be the pole.

All angular distances on the surface of a sphere, to an eye at the centre, are measured by arcs of great circles. Hence all triangles formed upon the surface of a sphere, for the solution of spherical problems, must be formed

by the arcs of great circles.

Secondaries to a great circle are great circles which pass through its poles,

and consequently must be perpendicular to their great circles.

The aris of the earth is that diameter about which it performs its diurnal motion; and the extremities of this diameter are called the poles.

The terrestrial equator is a great circle of the earth perpendicular to its axis. Hence the axis and poles of the earth are the axis and poles of its equator. That half of the earth which lies on the side of the equator, in which Europe and the United States of America are situated, is called the northern hemisphere, and the other the southern; and the poles are respectively called the north and south poles.

The latitude of a place upon the earth's surface is its angular distance from the equator, measured upon a secondary to it. These secondaries to the

equator are called meridians.

The longitude of a place on the earth's surface is an arc of the equator intercepted between the meridian passing through the place, and another, called the first meridian, passing through that place from which you begin to measure, or it is the angle formed at the pole by these two meridians. The Americans and English generally place the first meridian at London or Greenwich, the French place it at Paris, the Spaniards at Cadiz; some Geographers place it at Tenerifie, and others at other places. Throughout this work Greenwich will be reckoned as the first meridian. The longitude is counted from the first meridian, both eastward and westward, till it meets at the same meridian on the opposite point; therefore the longitude (and also the difference of longitude between any two places) can never exceed 180°.

If the plane of the terrestrial equator be produced to the sphere of the fixed stars, it marks out a circle called the celestial equator; and if the axis of the earth be produced in like manner, the points of the heavens, to which it is produced, are called poles, being the poles of the celestial equator. The

star nearest to each pole is called the pole star.

Secondaries to the celestial equator are called *circles of declination*; of these 24, which divide the equator into equal parts, each containing 15°, are called *hour* circles.

Small circles parallel to the celestial equator are called parallels of declination.

The sensible horizon is that circle in the heavens whose plane touches the earth at the spectator. The rational horizon is a great circle in the heavens, passing through the earth's centre, parallel to the sensible horizon.

If the radius drawn from the centre of the earth to the place where the spectator stands be produced both ways to the heavens, the point vertical to him is called the zenith, and the point opposite, the nadir. Hence the zenith and nadir are the poles of the rational horizon.

Secondaries to the horizon are called vertical circles, because they are perpendicular to the horizon. On these circles, therefore, the altitude of a

heavenly body is measured.

The secondary common to the celestial equator, and the horizon of any place, is the celestial meridian of that place. This meridian corresponds with the terrestrial meridian of the same place, which passes through the poles of the earth, the zenith and nadir crossing the equator at right angles, and cutting the horizon in the north and south points; that point being called north which passes through the north pole, and the opposite direction is called north. The vertical circle which cuts the meridian of any place at right angles is called the prime vertical; the points where it cuts the horizon are called the east and west points, and to an observer, with his face directed towards the south, the east point will be to his left hand, and the west to his right hand. Hence the east and west points are 90° distant from the north and south. These four are called the cardinal points. The meridian of any place divides the heavens into two hemispheres lying to the east and west; that lying to the east is called the eastern hemisphere. When the sun is at its greatest altitude on the meridian of any place, it is noon or mid-day.

The azimuth of an heavenly body is its distance on the horizon, when referred to it by a secondary, from the north or south points. The amplitude is its distance from the east or west points, at the time of rising or setting.

The ecliptic is that great circle in the heavens which the sun appears to describe in the course of a year. The ecliptic and equator, being great circles, must bisect each other, and their angle of inclination is called the obliquity of the ecliptic; and the points where they intersect are called the equinoctial points. The times when the sun comes to these points are called the equinoctial points. The ecliptic is divided into 12 equal parts, called signs;—viz. Aries \(\Tau\), Taurus \(\Sigma\), Gemini II, Cancer \(\Sigma\), Leo \(\Sigma\), Virgo \(\mathbb{T}\), Libra \(\Sigma\), Scorpio \(\mathbb{M}\), Sagittarius \(\frac{\pi}{\pi}\), Capricornus \(\Sigma\), Aquarius \(\Sigma\), Pisces \(\frac{\pi}{\pi}\). The order of these is according to the apparent motion of the sun. The first point of Aries coincides with one of the equinoctial points, and the first point of Libra with the other. The first six signs are called northern, lying on the north side of the equator; and the last six are called southern, lying on the south side.

The zodiac is a space extending eight degrees on each side the ecliptic, within which the motion of all the planets is contained, except the newly

discovered planets.

The right ascension of a body is an arc of the equator intercepted between the first point of Aries and a circle of declination passing through the body, massured seconding to the order of the signs.

measured according to the order of the signs.

Right ascension of the meridian or mid-heaven, is the distance of the meridian, from the first point of Aries, and is found by adding the apparent time past noon, to the sun's right ascension.

The ascensional difference of any object is the difference between the right ascension of the object and that point of the equator which rises or sets with it.

The declination of a star or any celestial object is its angular distance from

the equator, measured upon a secondary to it passing through the object.

The longitude of a star or any celestial object is an arc of the ecliptic intercepted between the first point for Aries and a secondary to the ecliptic passing through the body, measured according to the order of the signs.—If the observer be on the earth, the longitude is called the geocentric longitude; but if seen from the sun it is called the heliocentric longitude; the body in each case being referred perpendicularly to the ecliptic in a plane passing through the eye.

Nonagesimal degree of the ecliptic is its highest point at any given time,

and is 90° from the points where the ecliptic intersects the horizon.

The latitude of a star or any celestial object is its angular distance from the ecliptic, measured whon a secondary to it drawn through the body.—If the body be observed from the earth, its angular distance from the ecliptic is called the geocentric latitude; but if observed from the sun it is called the

heliocentric latitude. The secondary circle drawn perpendicular to the eclip-

tic is called a circle of latitude.

The tropics are two parallels of declination touching the ecliptic. One, touching it at the beginning of cancer, is called the tropic of cancer; and the other, touching it at the beginning of capricorn, is called the tropic of capricom. The two points, where the tropics cut the ecliptic, are called the solstitial points.

Columns are two secondaries to the celestial equator, one passing through the equinoctial points, called the equinoctial colure; and the other passing through the solstitial points, are called the solstitial colure. The times when

the sun comes to the solstitial points are called the solstices.

Aberration of a star or any heavenly body, is a small apparent motion, occasioned by the progressive velocity of light. This is calculated by means

of Tables XXXIX. XLI. or XLII.

Nutation is a small apparent motion of the heavenly bodies, occasioned by a real motion of the earth's axis, arising from the attractions of the sun and moon on the spheroidal form of the earth. The effect of this on the right ascension and declination is given in Table XLIII. and on the longitude in Table XL. The correction in this last Table being generally called the equation of the equinoxes in longitude.

Precession of the equinoctial points is a small motion of about 504" per year, occasioned by the same cause as the nutation. By this motion the equinoctial points are carried backward from east to west; consequently, the heavenly bodies appear to move forward the same quantity from west to east. The annual variations of the places of the stars from precession, and the secular equations arising from the change of the earth's orbit by the

attraction of the planets, are given in Tables VIII. and XXXVII.

The arctic and antarctic circles are two parallels of declination, the former about the north, and the latter about the south pole, the distance of which from the two poles is equal to the distance of the tropics from the equator, which is about 23° 28'. These are also called polar circles. The two tropics and two polar circles, when referred to the earth, divide it into five parts, called zones; the two parts within the polar circles are called the frigid zones; the two parts between the polar circles and tropics are called the temperate zones; and the part between the tropics is called the torrid zone.

Besides the imaginary divisions of the earth, there are various natural divisions of its surface, formed by nature, such as continents, oceans, islands,

seas, rivers, &c.

A Continent is a large tract of land, wherein are several empires, kingdoms, and countries conjoined—as Europe, Asia, Africa, and America.

An *Island* is a part of the earth that is environed or encompassed round

by the sea, as Long Island, Block Island, &c.

A Peninsula is a portion of land almost surrounded with water, save one narrow neck which joins it to the continent, as the Morea.

An Isthmus is a narrow neck of land joining a peninsula to the adjacent land;

by which the people may pass from one to the other, as the isthmus of Darien.

A Promontory is a high part of land stretching itself into the sea, the ex-

tremity of which is called a Cape or Headland. A Mountain is a rising part of dry land, over-topping the adjacent coun-

try, and appearing first at a distance.

An Ocean is a vast collection of water, separating continents from one another, and washing their borders or shores, as the Atlantic and Pacific Oceans.

A Sea is part of the ocean, to which we must sail through some strait, as the Mediterranean and Baltic seas. This term is sometimes used for the whole body of salt water on the globe.

A Strait is a narrow part of the ocean lying between two shores, and opening a way into some sea, as the Straits of Gibraltar that lead into the Medi-

A Creek is a small narrow part of the sea or river, that goes up but a little way into the land.

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A Bay is a great inlet of the land, as the Bay of Biscay, and the Bay of

Mexico; otherwise a bay is a station of road for ships to anchor in.

A River is a considerable stream of water issuing out of one or various springs, and continually gliding along in one or more channels, till it discharges itself into the ocean: the lesser streams are called rivulets.

A Lake is a large collection of waters in an inland place, as the lakes Su-

perior and Huron in America.

A Gulf is a part of the ocean or sea, nearly surrounded by the land, except

where it communicates with the sea, as the Gulf of Venice.

Thus we have given the most useful definitions of Astronomy and Geography, and to assist the learner there is also given Plate V. in which those terms are explained at one view. We may farther observe, that as the latitude of any place upon the earth is counted from the equator upon an arch of the meridian, the difference of latitude between two places, both north, or both south, is found by subtracting the less latitude from the greater; but if one latitude be north, and the other south, the difference is found by adding both latitudes together.

1. Consequently, if a ship in north latitude sails northerly, or in south latifude southerly, she increases her latitude; but in north latitude sailing southerly, or in south latitude sailing northerly, she decreases her latitude, because she sails nearer to the equator, from whence the latitude is reckoned.

Wherefore, in north latitude sailing northerly, or in south latitude sailing southerly, the difference of latitude, added to the latitude left, gives the latitude in.

3. In north latitude sailing southerly, or in south latitude sailing northerly, the difference of latitude, subtracted from the latitude left, gives the latitude in.

4. When the latitude decreases, and the difference of latitude is greater than the latitude sailed from, subtract the latitude left from the difference of latitude, and the remainder will be the latitude in, but of a different name, for it is evident in this case, that the ship has crossed the equator.

5. The difference of longitude between two places, being both east or west, is found by subtracting the less longitude from the greater; but if one be in east longitude and the other in west, their sum is the difference of longitude, when it does not exceed 180°, but if it exceeds 180°, that sum must be subtracted from 360°, and the remainder will be the difference of longitude.

6. Therefore in east longitude sailing easterly, or in west longitude sailing westerly, the difference of longitude added to the longitude left, gives the longitude in, when that sum does not exceed 180°; but if it exceeds 180°, the sum, subtracted* from 5600, leaves the longitude in, but of a different name from that left.

7. In east longitude sailing westerly, or in west longitude sailing easterly, the difference of longitude, subtracted from the longitude left, gives the longitude in; but when the difference of longitude is greatest, the longitude left must be subtracted from that difference, and the remainder will be the longitude in, but of a different name from the longitude left.

What has been said will be rendered familiar to the learner by the following

examples.

Exam. I. What is the different between Boston, in the latit N. and Richmond (Virgini of 37° 30' N.?	ude of	12° 23	Exam. II. A ship from latity sails southward until her latitude is 374 miles; what come to?	differ	ence	e oí
From Boston's lat. Subtract Richmond's lat.	42° 2 37 3		Latitude sailed from Diff. of lat. 374-68-	59° 6	27' 14	
Remains the diff. of lat.	4 5 60	3	Lat. in	65	41	s.
In Miles	293					

[•] In this rule it is supposed, that the sum of the longitude left, and the difference of longitude, is less an 380°, which is always the case when the difference of longitude is less than 180°, which we have merally supposed to be the case in these rules.



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Er. III. Required the differ between Georgetown and		Ex. IV. A ship from latitude sails south 1800 miles; wha			
Georgetown's lat.	33° 25' N.	she in?			
Cape Frio's lat.	23 1 5.	From diff. of lat. 1800 miles, or	30°	00'	S.
•		Sub. lat. left	28	25	N.
Diff. of lat.	56 24				
•	60	Diff, = lat. in	1	3 5	S.
ln miles	3384				

In the last example it is evident that as the difference of latitude is more than the latitude left, the ship must have crossed the equator, and consequently come into south latitude.

When one of the places has no latitude, or is on the equator, the

latitude of the other place is their difference of latitude.

Ex. V. What is the differen- between Cape Ann light-h bon? Cape Ann light-house's long.	ouse and I	.is-	Ex. VI. A ship from Cap Virginia, sails eastward till of longitude is 400 miles;	her difference	
Lisbon's long.	9 9 V		Cape Charles' long,	76° 04′ W.	
Diff. of long.	61 25	••	Diff. of long. 400 miles =	6 40 E.	
In miles	60 3685		Long. in	· 69 24 W.	
Ex. VII. What is the difference of longitude between Barcelona and Salem? Ex. VIII. A ship from 15° 40′ E. long. tade between Barcelona and Salem?					
Barcelona's long.	2° 12′	E.	27° 15', what longitude is	she in?	
Salem's long.	70 52	W.	Long. left	15° 40′ E.	
Diff. of long.	73 4		Diff. of long. Long. in	27 15 W.	
Ex. IX. What is the difference of longitude 160° 20′ W. tade between Manilla and New-York sails westward until she differs her longitude 11° 20′; what longitude is she in?					
Manilla's long.	121° 02′ 1	E.	Long. left	160° 20′ W.	
New-York light-house	74 01	W.	Diff. of long.	41 20 W.	

195 Sum exceeds 180° 201 Subtract it from 360 00 360 00 Diff. of long. 164 57 Long. in 158 In the last example the ship has crossed the opposite meridian, and there-

03

fore has come into a longitude of a different name.

PLANE SAILING.

PLANE SAILING is the art of navigating a ship upon principles deduced from the supposition of the earth's being an extended plane, on which the meridians are all parallel to each other.* A map of the several parts of the earth, constructed upon these principles, is called a Plane Chart. the parts of the earth are thus delineated on a plane, it is easy to see the track by which a ship may go from one place to another, and also what angle this track makes with the meridian. † Ships at sea are kept in this track by means of an instrument called the mariner's compass.

The MARINER'S Compass is an artificial representation of the horizon of any Place. It consists of a circular piece of paper (see Plate VI. fig. 1) called a card, divided (like the horizon) into 360 degrees or 32 points. This is fixed on a piece of steel, called a needle, to which the magnetic virtue has been communicated by means of a loadstone, which has the property of pointing steadily towards the north, and carrying the card with it, when turning freely on a pivot or any thing to support it. Thus all the points of the card will be

eath, will be given hereafter.

The explanations of Plane Sailing, and the definitions of this page (and in the former editions of this work) are searly the same as those given by Moore, in his Practical Navigator; by Robertson in his Plemants of Navigation, and by most writers on Navigation.

The method of calculating this angle on the true principles of sailing on the spherical surface of the

directed towards their corresponding points of the horizon,* consequently, by

help of the compass a ship may be kept in any proposed track or course.

The Course is the angle which the line described by a ship makes with the meridian, being sometimes reckoned in points, half points, &c. and

sometimes in degrees.

DISTANCE is the way or length a ship has gone on a direct course in a The method of measuring this distance by the log will be exgiven time. plained hereafter.

DIFFERENCE OF LATITUDE is the distance which the ship has made north or south of the place sailed from, or the portion of the meridian contained between the parallels of latitude sailed from and come to.

DEPARTURE is the east or west distance a ship has made from the meri-

dian, or the whole easting or westing made by the ship.

If a ship sails due north or south, she sails on a meridian, makes no departure, and her distance and difference of latitude are the same. sails due east or west, she goes on a parallel of latitude, makes no difference

of latitude, and her departure and distance are the same.

The difference of latitude and departure make the legs of a right-angled triangle, the hypotenuse of which is the distance the ship has sailed; the perpendicular is the difference of latitude counted on the meridian; the base is the departure, which is easting or westing counted from the meridian; the angle opposite to the base is the course, or angle, that the ship makes with the meridian; and the angle opposite the perpendicular is the complement of the course, which being taken together, make always 8 points or 90 degrees.

In constructing figures relating to a ship's course, let the upper part of the paper, or what the figure is d

South.

North.

lower part will be the south; the right hand east, and the left west.

Draw the north and south line to represent the meridian of the place the ship sails from; then, the ship's course is to the south ward, mark the upper end of th line for the place sailed from; bu if the course is northward, mar the lower end for that place.

When the course is easterly describe the arch, and lay off th course and departure on the rigl hand side of the meridian; b when westerly, on left hand side

When the course is given in de grees, they must be taken from the line of chords; but when i points, from the line of rhumbs and must always be laid off upo the arch, beginning at the mer

When the course is given points, the log-sine, log. co-sine &c. may be found in Table XX otherwise in Table XXVII.

In all cases, where the comple ment of course, or co-sine, &c. _used, the degrees or points p down, are the course itself, but the logarithms belonging to the con plement or co-sine, &c. of the course are taken.

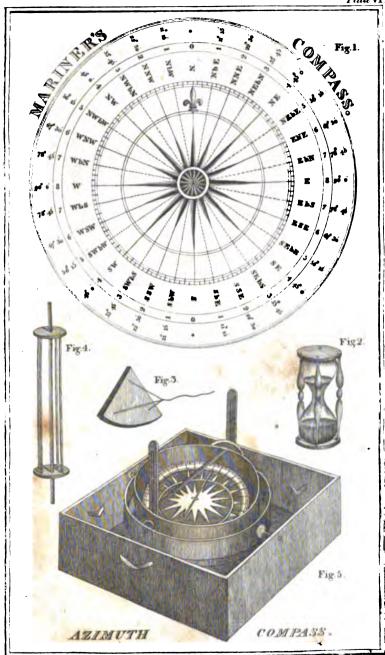
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A Table of the				f the Com	pass
•	makes wit	h the Meridi	an.		

North.

South.

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h-			4	8.26		ا ـــا
be	N. by E.	8. by E.	1	11.15	N. by W.	S. by W.
ut			14	14. 4		
rk			15	16.52		
1			12	19.41		
	N. N. E.	S. S. E.	2	22.30	N. N. W.	S. S. W.
7-			21	25.19		
he			24	29. 7]
bt			24	30.56		1
uı	N. E. by N.	S. E. by S.	3	33.45	NW. byN.	S.W. by S.
le.			31	36.34		
e-			31	39.22		
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in	N. E.	S. E.	4	45. 0	N. W.	S. W.
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is			64	75.56		
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he		!	71	81.34		
m-		l '	74	84.22		Į.
at			71	87.11		Į.
		East.	8	90. 0	West.	<u> </u>

^{*} It is here supposed that the needle points to the true north, but if it varies therefrom, allow must be made for the variation by the rules which will be given in this work.



FULLWHED BY EDM.M.BLTHT, 1896.

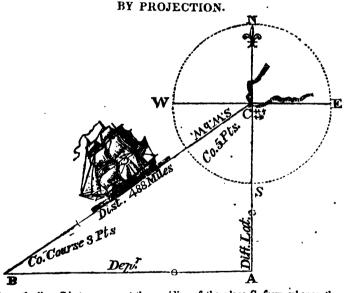
In the following Table, the rules for solving the various cases of Plane Sailing are collected.

FLANE SAILING.					
Case	Given.	Required,	Solutions.		
1	Course	Diff. of Lat.	Rad. : Dist. : : Cos. Course : Diff. of Lat.		
	and Distance	Departure.	Rad. : Dist. : : Sine Course : Departure.		
2	Course and	Distance	Co-sine Course : Diff. Lat. : : Rad. : Distance.		
	Diff. of Lat	Departure	Radius : Diff. Lat. : : Tang. Course : Departure.		
3	Course and	Distance	Sinc Course : Departure : : Rad. : Distance.		
	Departure.	Diff. of Lat.	Radius : Departure : : Co-tang. Course : Diff. Lat.		
4	Distance and Diff. of Lat.	Course Departure	Distance : Radius : : Diff. Lat. : Cos. Course. Radius : Distance : : Sine Course : Departure.		
5	Distance and	Course	Distance : Itadius : : Departure : Sine Course.		
	Departure.	Diff. of Lat.	Radius : Distance : : Cos. Course : Diff. Lat.		
6	Diff. of Lat. & Departure.	Course Distance.	Diff. Lat.: Radius :: Dep.: Tang. Course. Sine Course : Departure :: Rad.: Distance. Radius : Diff. Lat. :: Secant Course : Distance.		

CASE I.

Course and distance sailed given, to find the difference of latitude and departure from the meridian.

A ship from the latitude of 49° 67' N. sails S. W. by W. 488 miles; required the latitude she is in, and her departure from the meridian sailed from?



Draw the line CA to represent the meridian of the place C, from whence the ship sailed. With the chord of 60° in your compasses, and one foot in C as a centre, describe the compass N. W. S. E. Take 5 points in your compasses from the line of rhumbs on the plane scale, and set it off on the arch from S. towards W. for the course; through this point and C draw the line CB, which make equal to the distance 488; draw BA parallel to the E. and W. points W. E. to cut the meridian in A. Then will CA be the difference of latitude 271.1, and AB the departure 405.8.

BY LOGARITHMS.

By making the distance radius.						
To find the departure	. .	To find the difference of la				
As radius 8 points	10.00000	As radius 8 points -	10.00000			
Is to the distance 488	2.68842	Is to the distance 488	2.68842			
So is the sine course 5 points	9.91985	So is co-sine course 5 points	9.74474			
To the departure 405.8	2.60827	To the difference of lat. 271.1	2.43316			
Now as the ship is in north latitude sailing southerly; from the latitude left 49° 57° N.						

Take the difference of latitude 271.1=4 31 8.

Gives the latitude in 45 26 N.

And the departure from the meridian is 405.8 miles.

BY GUNTER.

Extend from radius or 8 points* to 5 points on the line marked SR; that extent will reach from the distance 488 to the departure 405.8 on the line of numbers.

2dly. Extend from radius or 8 points to 5 points, the complement of the course, on the line SR; that extent will reach from the distance 488 to the difference of latitude 271 on the line of numbers.

Thus may all the operations be performed in the several cases of Naviga-

tion.

By this case are calculated the tables of latitude and departure (Tables I. and II.) for every degree, point, and quarter point of the mariner's compass, to the distance of 300 miles. By the inspection of these Tables, a day's work may be calculated in a much more expeditious manner than by logarithms or by Gunter's Scale. In consequence of this, the method by inspection is generally used at sea in preference to every other method.

BY INSPECTION.

Find the given course at the top or bottom of the tables, either among the points or degrees, and in that page, against the distance taken in its column, will stand the difference of latitude and departure in their columns.

It must be observed, that in using these tables, the names Dist. Lat. Dep. must be found at the top if the course is found there, but if the course is

found at the bottom, those names must be found at the bottom.

Thus the course S. W. by W. or 5 points, is found at the bottom of the table of difference of latitude and departure for points; and as the distance 488 is too great to be found in the tables, divide it by 2 (or any other convenient number) and that gives 244, which look for in the distance column, and against it stand 135.5 for the difference of latitude, and 202.9 for the departure, which being doubled (because divided by 2) give 271 for the difference of latitude, and 405.8 for the departure, the same as before.

CASE II.

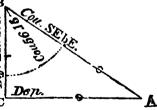
Course and difference of latitude given, to find the distance run, and departure from the meridian.

If a ship runs S. E. by E. from 1° 45' north latitude, and then by observation is in 2° 46' south latitude, what is her distance and departure?

In this case, as the ship has crossed the equator, the sum of the two latitudes 1° 45' and 2° 46' is the difference of latitude 4° 51'=271 miles.

BY PROJECTION.

Draw BC=271, and BA making an angle with BC equal to the course 5 points, or 56° 15'; draw CA perpendicular to BC to cut BA in A, and it is done; for the CA will be the departure=406, and AB at the distance=488.



	BY LOGA	RITHMS.		
By making diff. of lat. BC rad	ius.	By making the Dep. AB radius.† To find the distance.		
To find the departure.		To find the distance.		
As radius 4 points Into diff. of lat. 271	10.00000	As co-sine course 5 points	9.74474	
		Is to the diff. of lat. 271	2.43297	
Se is tang. course 5 pts.	10.17511	So is radius	10.0000 0	
Te the departure 405.6	2.60808	To the distance 487.8	2.68823	

[&]quot;When the course is given in points, make use of the lines marked sine rhumbs, and tangent rhumbs, on the upper side of the scale; when in degrees, make use of the lines marked sine and tangent.

I When the distance is too great to be found in the tables, you must divide it by 2, 3, 4, or any convenient number, the numbers corresponding to the quotient being multiplied by the divisor will give the

sought numbers.

† Sy making BC radius you would have Rad. : Diff. Lat. : : Secant Course : Distance but
this canon would not do for a common scale on which there is no line of secants. The same thing is to

he observed in the following cases.

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Hence the ship's distance run is 487.8 miles, and her departure from the meridian is 405.6 easterly.

BY GUNTER.

Extend from radius or 4 points to the course 5 points on the line marked TR, that extent will reach from the difference of latitude 271 to the departure 405.6 on the line of numbers.

2dly. Extend from the complement of the course 3 points to the radius 8 points on the line SR, that extent will reach from the difference of latitude 271 to the distance 488 on the line of numbers.

BY INSPECTION.

Find the course among the points or degrees, and the difference of latitude in its column, against which will stand the distance and departure in their columns.

Now, as the difference of latitude 271 is two great to be found in the tables, I divide it by 2, and that gives 135.5 which I find over S. E. by E. or 5 points in the latitude column; against that stand 244, for the distance, and 202.9 for the departure, which multiplied by 2 give the distance 488, and the departure 405.8.

CASE III.

Course and departure from the meridian given, to find the distance and difference of latitude.

If a ship sails N. E. by E. ‡ E. from a port in 3° 15' south latitude, until she depart from her first meridian 406 miles, I demand the distance sailed, and the latitude she is in?

BY PROJECTION.

Draw the meridian AB, upon which B erect the perpendicular BC, and set off thereon the departure 406 easterly from B to C; with the chord of 60°, on C as a centre, describe an arch, and set off thereon the complement of the course DE; through D and C draw the line CDA, cutting the meridian in the point A

B Dep. 496 E. C

A; then AC measured on the same scale before used, gives the distance 449, and AB 192, the difference of latitude.

BY LOGARITHMS.

By making the departure BC rad	lius.	By making the distance AC	rādius.		
As radius 4 points	10.00000	As sine course 5% pts.		9.95	616
Is to the departure 406	2.60853	Is to the departure 406		2.60	853
So is co-tang. course 5% pts.	9.67483	So is radius	1	0.00	000
To the diff. of lat. 192	2.28336	To the distance 449,1		2.60	
From the latitude left		•	3 º	15	8.
Subtract the difference of latitud	le 192 miles	s, or	3	12	N.
The remainder being 3, shows the ship is in the latitude of			0	3	ß.

BY GUNTER.

Extend from radius or 4 points to the co-course 21 points on the line marked TR; that extent will reach from the departure 408 to the difference of latitude 192 on the line of numbers.

2dly. Extend from the course 51 points to radius on the sines, that extent will reach from the departure 406 to the distance 449 miles on the line of numbers.

BY INSPECTION.

Find the course either among the points or degrees, and the departure in its column, against which will stand the distance and difference of latitude in their respective columns.

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Thus with the course 53 points, and half the departure 203* I find 224.5 for the distance, and 96.0 for the difference of latitude, which, being doubled, give the distance 449, and the difference of latitude 192.0, as before.

CASE IV.

Distance and difference of latitude given, to find the course and departure.

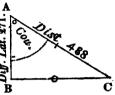
Suppose a ship sails 488 miles, between the south and the east, from a port in 20 52' south latitude, and then by observation is in 70 23' south latitude; what course has she steered, and what departure has she made?

From the latitude by observation 7° 23' take 2° 52' the latitude left, the re-

mainder 40 31'=271 miles, is the difference of latitude.

BY PROJECTION.

Draw the meridian AB=271; upon which erect - > the perpendicular BC; take 488 in your compasses, & and with one foot on A, as a centre, describe an arch cutting BC in C; join A and C; then will BC be the departure 406, and the angle BAC the course=560 16' or five points nearly.



BY LOGARITHMS.

To find the course. As the distance 483 Is to radius So is the diff. of lat. 271	To fit 2.68842 As radius 10.00000 Is to the distance 2.43297 So is sine course	e 489	10.000 0 0 2.68842 9.91993
To co-sine course 56° 16′	9.74455 To the departur	e 405.8	2,60835
Hence the course is	S. E. by E. and the d	eparture 405.8.	

BY GUNTER.

The extent from the distance 488, to the difference of latitude 271 on the line of numbers, will reach from radius, or 90° to 33° 44', the co-course on the line of sines.

And the extent from radius, to 56° 16' on the line of sines, will reach from the distance 488 to the departure 405.8 on the line of numbers.

BY INSPECTION.

Seek in the tables till against the distance taken in its column is found the given difference of latitude in one of the following columns, adjoining to it will stand the departure; which, if less than the difference of latitude, the course is to be found at the top; but if greater, the course is to be found at the bottom.

Thus half the distance 244, and half the difference of latitude 135.5, are found to correspond to a course of 5 points or S. E. b. E. and to the departure 202.9, which being doubled, gives 405.8, as before.

CASE V.

• Distance and departure given, to find the course and difference of latitude.

Suppose a ship sails 488 miles between the north and west, from the latitude of 320 25' north, until her departure is 405 miles, what course has she steered, and what latitude is she in?

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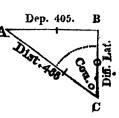
The nearest numbers in the table are 202.5 and 203.4, and as the number 203 is nearly a mean of these two values, I take the mean 224.5 of the corresponding distances 224, 225, and the mean 36 of the corresponding departures 36.8 and 96.2; these doubled give the true distance 448, and departure 192.

It may also be known whether the course be marked at the top or bottom of the table, by observing whether the difference of latitude and departure correspond with the marks at the top or bottom. Thus the half distance 244, and half difference of latitude 135.5 correspond to the course 5 points, because the column in which 135.5 is found, is marked latitude at the bottom; the same may be observed in the following cases. following cases.

2.43486

BY PROJECTION.

Draw the line AB equal to the departure 405, and perpendicular thereto the line BC to represent the meridian, then take the distance 488 in your compasses, and fixing one foot in A as a centre, describe an arch cutting BC in C, join AC and it is done; for the angle ACB will be the course, and BC the difference of latitude.



BY LOGARITHMS.

To find the course.

As the distance 488

Is to radius

10.00000 Is to the distance 488

2.68342 As radius

10.00000 Is to the distance 488

2.68342 So is the departure 405

2.66746 So is co-sine course 56° 6'

2.74644

To the sine of course 56° 6′ 9.91904 To the diff. of latitude 272,2 Hence the course is N. 56° 6′ W. or N. W. by W. nearly.

To the latitude sailed from 32° 25' add the difference of latitude 272, or 4° 32', the sum 36° 57' is the latitude the ship is in.

BY GUNTER.

Extend from the distance 488 to the departure 405 on the line of numbers, that extent will reach from radius to the course 56° 6' on the line of sines.

2dly. Extend from radius to the complement of the course 35° 54' on the line of sines, that extent will reach from the distance 488 to the difference of latitude 272 on the line of numbers.

BY INSPECTION.

Seek in the tables till against the distance, taken in its column, is found the given departure in one of the following columns; adjoining to it will stand the difference of latitude; which, if greater than the departure, the course is to be found at the top; but if less, the course is to be found at the bottom.

Thus half the distance 244, and half the departure 202,5, agree nearly to a course of 5 points or N. W. by W. and a difference of latitude 135,5, which being doubled, is 271, the difference of latitude, nearly as before.

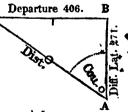
CASE VI.

Difference of Latitude and Departure given to find the Course and Distance.

A ship sails between the north and west till her difference of latitude is 271 miles, and her departure is 406 miles; I demand her course and distance?

BY PROJECTION.

Draw AB=271, and perpendicular to it C BC=406; join C and A; then will the angle CAB be the course=56° 17′, and AC the distance=488 miles.



BY LOGARITHMS.

To find the course.
As the diff. of latitude 271
Is to radius
10.00000 Is to the diff. of latitude 271
2.43297 As radius
10.00000 Is to the diff. of latitude 271
2.60853 So is sec. of course 56° 17'
10.25554

To tang. of course 56° 17′ 10.17556 To the distance 489.2

Hence her course is N. 560 17' W. or N. W. by W. and the distance saffett is 488.2 miles.

2.68861

BY GUNTER.

Extend from the difference of latitude 271 to the departure 406 on the line of numbers, that extent will reach from radius to 56° 17' the course on the

line of tangents.

2dly. For the distance we must consider it as radius (unless there is a line of secants on the scale) and extend from the course 560 17' to the radius or 90° on the line of sines, that extent will reach from the departure 406, to the distance 488 on the line of numbers.

BY INSPECTION.

Seek in the tables till the given difference of latitude and departure are found together in their respective columns, then against them will be the distance in its column, and the course will be found at the top of that table if the departure be less than the difference of latitude, otherwise at the bottom.

Thus with half the difference of latitude 135.5, and half the departure 208, enter the tables, and these numbers will be found to correspond nearly to 5 points or N. W. by W. course, and a distance equal to 244 miles, which being

doubled gives the sought distance, 488.

Questions to exercise the learner in the foregoing Rules.

Question I. A ship in 20 10' south latitude, sails N. by E. 89 leagues; what latitude she in, and what is her departure?

Answer. Latitude in 20 12' N. and departure 17,36 leagues.

Question II. A ship sails S. S. W. from a port in 410 30' north latitude, and then by observation is in 36° 57' north latitude; I demand the distance run and departure?

Answer. Distance run 98,5 leagues, departure 37,7 leagues.

Question III. A ship sails S. S. W. & W. from a port in 20 30' south latitude, until her departure be 59 leagues; I demand her distance run and latitude in?

Answer. Distance run 125,2 leagues, latitude in 80 1' south.

Question IV. If a ship sail 360 miles south westward from 210 59' south latitude, until by observation she be in 24° 49' south latitude, what is her course and departure?

Answer. The course is S. W. by W. half W. or S. 610 49' W. and her

departure from the meridian is 317,3 miles.

Question V. Suppose a ship sails 354 miles north eastward from 20 9' south latitude, until her departure be 150 miles, what is her course and latitude in?

Answer. Her course is N. 250 4' E. or N. N. E. & E. nearly, and she is

in latitude 30 12' N.

Question VI. Sailing between the north and the west, from a port in 10 59' south latitude, and then arriving at another port in 40 8' north latitude, which is 200 miles to the westward of the first port—I demand the course and distance from the first port to the second?

Answer. The course is N. 29° 40' W. or N. N. W. & W. nearly, and the

distance of the ports is 422,4 miles, or 140,8 leagues.

Question VII. Four days ago we were in lat. 3° 25' S. and have since that time sailed in a direct course N. W. by N. at the rate of 8 miles an hour; required our present latitude and departure?

Annær. Latitude in 7º 14' N. Departure 426,7 miles.

Question VIII. A ship in the latitude of 30 52' S. is bound to a port bearing N. W. by W. 1 W. in the latitude of 40 30' N. how far does that port lie to the westward, and what is the ship's distance from it?

Answer. The port lies 939,2 miles to the westward, and the direct distance

1065 miles.

Question IX. A ship from the latitude of 48° 17' N. sails S. W. by S. until she has depressed the north pole 2 degrees; what direct distance has . she sailed, and how many miles has she got to the westward?

Answer. Distance run 144,3 miles, and has got to the westward 80,2 miles.

TRAVERSE SAILING.

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A TRAVERSE is an irregular track which a ship makes by sailing on several different courses; these are reduced to a single course by means of two or more cases of Plane Sailing, either by geometrical construction, or

by arithmetical calculation.*

The geometrical construction is performed as follows: Describe a circle with the chord of 60°, to represent the compass, and lay off on its circumference the various courses sailed. From the centre, upon the first course set off the first distance, and mark its extremity: through this extremity, and parallel to the second course, draw the second distance of its proper length; through the extremity of the second distance, and parallel to the third course, draw the third distance of its proper length; and thus proceed till all the distances are drawn. A line, drawn from the extremity of the last distance to the centre of the circle, will represent the distance made good; a line, drawn from the same point, perpendicular to the meridian, will represent the departure; and the part of the meridian intercepted between this and the centre, will represent the difference of latitude.

The arithmetical calculation to work a traverse is as follows; Make a traverse table consisting of six columns; title them, Course, Dist. N. S. E. W. begin at the left side, and write the given courses and distances in their respective columns. Find the difference of latitude and departure for each of these courses, by Gunter's Scale, or by Tables I. or II. (as in Case I. Plane Sailing) and write them in their proper columns; that is, when the course is southerly, the difference of latitude must be set in the column S. when northerly in the column N. The departure, when westerly, in the column W. and when easterly in the column E. Add up the columns of northing, southing, easting, and westing; take the difference between the northing and southing, and also between the easting and westing; the former difference will be the difference of latitude, which will be of the same name as the greater; and the latter will be the departure, which will be also of the same name as the greater. With this difference of latitude and departure, the course and distance made good are to be found as in Case VI. Plane Sailing.

EXAMPLE I.

Suppose a ship takes her departure from Block Island, in the latitude of 41° 10′ N. the middle of it bearing N. N. W. distance by estimation 3 leagues, and sails S. E. 34, W. by S. 16, W. N. W. 39, and S. by E. 40 miles; required the latitude she is in, and her bearing and distance from Block Island?

^{*} This method of reducing compound courses to a single one is perfectly accurate in sailing on a plane, and is nearly so in sailing a short distance on the spherical surface of the earth; and though in this case it is liable to a small error in high latitudes, yet in general the rule is sufficiently accurate for reducing the several courses and distances sailed in one day to a single course and distance.

BY PROJECTION.

Let L represent the middle of Block-Island; draw the meridian LM, and on L as a centre, with a chord of 60°, describe a circle to represent the compass, on which mark the various courses sailed, and the bearing of the land at the time of taking the departure; opposite to this bearing draw the S. S. E. line LA, which make equal to 15 miles, the estimated distance of the land; then will A represent the place of the ship at the time of taking the departure: through A draw AB= **54** miles parallel to the . S. E. line; then will B be the place of the ship after sailing her first course; in like manner draw BC=16 miles parallel to the W. by S. line: CD= 39 miles parallel to the W. N. W. line and DE

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=40 miles parallel to the S. by E. line; then will E represent the place of the ship after sailing her several courses. Join EL, and draw EM perpendicular to LM; then will LE be the distance of Block-Island 66.8 miles, and the angle ELM=12° 16' will be the course made good, LM the difference of latitude, and EM the departure.

TO FIND THE SAME BY LOGARITHMS.

IO FIND I	LIE OAN.	o bi noominimo.	
		e S. S. E. 15 miles.	
To find the difference of is	titude.	For departure.	
As radius 90°	10.00000	As radius 90°	10.00000
Is to distance 15	1.17609	is to distance 15	1.17609
So is co-sine course 2 points	9.96562	So is sine course 2 points	9.58284
To Diff. lat. 13.9	1.14171	To departure 5.7	0.75893
		S. E. 34 miles.	
· For difference of latitu	de.	For departure.	
As radius 90°	10,00000	As radius 90°	10.00000
Is to co-sine course 45°		Is to sine course 45°	9.84949
So is distance 34.		So is distance 34	1.53148
To diff. latitude 24	1.38097	To departure 24	1.38097
		by S. 16 miles.	
For difference of latitude		For departure.	
As radius 90°	10.00000	As radius 90°	10.00000
Is to co-sine course 780.45'		Is to sine-course 78° 45'	9.99157
So is distance 16		So is distance 16	1.20412
To diff. latitude 3.1	0,49436	To departure 15.7 Digitized by	1,19569

Fourth course W. N. W. 59 miles.

For difference of latit	ade.	For departure.	
As radius 90°		As radius 90°	10.00000
Is to co-sine course 67° 36'	9.58284	Is to sine course 67° 30'	9.96562
So is distance 39		So is distance 39	1.59106
		}	
To diff. lat. 14.9	1.17390	To departure 36	1.55668

F ifth	course S. by E. 40 miles.	
For difference of latitud	le. For departure.	
As radius 90°	10.00000 As radius 90°	10.00000
Is to co-sine course 11° 15'	9.99157 Is to sine course 11° 15'	9.29024
So is distance 40	1.60206 So is distance 40	1.60906
To diff, lat. 39.2	1.59363 To departure 7.8	0.89230

Though this method of finding the difference of latitude and departure by logarithms is accurate, yet the calculations may be more easily made by the tables of difference of latitude and departure, as in Case I. Plane Sailing.

TRAVERSE TABLE.

Place all these courses, distances, &c. in the traverse table, then add up all the westings, eastings, northings, and southings, separately, and set down their respective sums at the bottom of each column; and as the westing is greater than the easting subtract the easting therefrom; the difference 14,2 shews that the ship's departure is so much west of her first meridian.

Camara	Dist			Dep	Departure.		
Courses.	Dist.	N. I		E.	W.		
S. S. E. S. E.	15 34			5.7			
W. by S. W. N. W. S. by E.	16 39 40	14.9	3.1	7.8	15.7 36.0		
From sum Take	C	=	80.2 14.9	37,5	51.7 37.5		
Rests		_	65.3		14.2		

Again, the southing being greater than the northing, subtract the northing from it, and the remainder 65.3 shews how far the ship is to the southward of her first place.

To find the direct course of Block Island from the		To find the distance of th	e Island.
As the diff. lat. 65.3		As sine of course 12° 16'	9.32723
Is to radius 45°		Is to the departure 14,2	1.15229 -
So is the departure 14.2	1.15229	So is radius 90°	10.00000
		,	
To tang. course 12° 16'	9.33738	To the distance 66.8	1.82501

Which, because the difference of latitude is southerly, and the departure westerly, is S. 12° 16' W. Whence Block Island Find the course and distance by Case VE. bears from the ship N. 12° 16' E. or N. by of Plane Sailing. E. 1º 1' E.

EXAMPLE II.

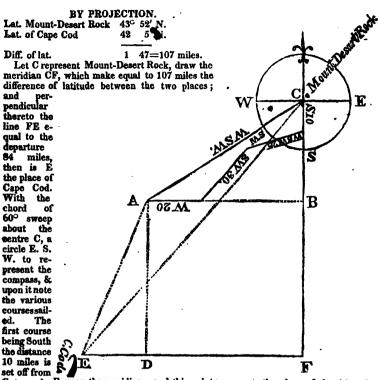
A ship from Mount-Desert Rock, in the latitude of 45° 52' N. sails for Cape Cod in the latitude of 42° 5' N. its departure from the meridian of Mount-Desert Rock being supposed to be 84 miles west; but by reason of contrary winds, she is obliged to sail on the following courses, viz. South 10 miles, W. S. W. 25 miles, S. W. 30 miles, and W. 20 miles. Required the bearing and distance of the two places, the course and distance sailed by the ship, and the bearing and distance of her intended port?

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set of from Towards F upon the meridian, and this point represents the place of the ship after sailing her first coure; continue setting off the various courses and distances as in the last example, viz. W. S. W. 25 miles, S. W. 30 miles, and West 20 miles to the point A; then will A represent the place of the ship after sailing these courses. Join CE, AC, AE; draw AB perpendicular to the meridian CF, and AD parallel thereto: then will AC=76,2 miles be the distance made good, AE=69,1 miles the distance of Cape Cod from the ship; CE the distance of the two places=136 miles; ACB=57° 36', the course made good; EAD=16° 34' the course to Cape Cod, and ECF the course from Mount-Desert Rock to Cape Cod=38° 8', &c.

BY LOGARITHMS.

To find the bearing and distance of the two places by Case VI. Plane Sailing.

To find the bearing.		To find the distance.	_
As diff. of lat. 107		As radius 90°	10.00000
Is to radius 45°		Is to diff. of lat. 107	2.02935
So is departure 84	1.92428	So is sec. course 38° 8'	10.10426
	0.00.400		
To tang. course 39° 8'	9.59490	To the distance 136	2.13364

Whence the course from Mount-Desert Rock to Cape Cod is S. 38° 8' W. distance 136 miles. The same may be found by the scale or by inspection.

The difference of latitude and departure for the several courses being calculated, by Case I. Plane Sailing, and arranged in the traverse table, it appears that the difference of latitude made good by the ship is 40,8 miles; and the departure 64,3 miles; then by Case VI. Plane Salling, these numbers are found to correspond to a course of S. 57° 36′, W. and distance 76,2 miles.

TRA	VERSE	TABLE.

Courses.	Ti-A	Diff.	Lat.	Depa	rture
Courses.	Dist.	N.	S.	E.	W.
South.	10		10.0		
W. S. W.	25	l	9.6		23.1
s. w.	30		21.2		21.2
w.	20			1	20.0
	Diff.	Let.	40.8	Dep.	64.3

Subtract the difference of latitude made good by the ship 40,3 miles, from the whole difference of latitude 107 miles, and there remain 66,2 miles, which is the difference of latitude between the ship and Cape Cod. In the same manner by subtracting the ship's departure, 64,3 miles, from the whole departure, 84 miles, there remain 19,7 miles for the departure between the ship and Cape Cod. With this difference of latitude, 66,2, and departure, 19,7, the bearing of Cape Cod is found, by Case VI. Plane Sailing, S. 160 34' W. and its distance 69,1 miles.

All the preceding calculations may be made by logarithms, by the scale, or by inspection. But we shall leave them to exercise the learner; and for the same purpose shall add the following example.

EXAMPLE III.

A ship in the latitude of 37° 10' N. is bound to a port in the latitude of 350 0' N. which lies 180 miles west of the meridian of the ship; but by reason of contrary winds she sails the following courses, viz. S. W. by W. 27 miles—W. S. W. 4 W. 30 miles—W. by S. 25 miles—W. by N. 18 miles—S. S. E. 32 miles—S. S. E. 3 E. 27 miles—S. by E. 25 miles—S. 31 miles, and S. S. E. 39 miles. Required the latitude the ship is in, and her departure from the meridian, with the course and distance to her intended port?

The difference of latitude and departure made on each course, are given in the adjoined traverse table: hence it appears that the difference of latitude made good is 169.4 miles, the departure 47.4 miles, and by Case VI. Plane Sailing, the course S. 15° 38' W. and distance 175.9 miles; and the course to the intended port S. 580 42' W. distance 155.2 miles; the latitude being in 34° 21' N.

TRAVERSE TABLE.						
Courses.	Dist.	Diff	Lat.	Depar	ture.	
Courses.	DBC.	N.	S.	E.	w.	
S. W. by W. W.S.W. W. W. by S. W. by N.	27 30 25 18	3.5	15.0 8.7 4.9		99.4 98.7 94.5 17.7	
S. S. E.* S. S. E. § E. S. by E.	32	3.5	29.6 23.2 24.5 31.0	13.9 4.9		
S. 8. E.*	39		36.0		_ 1	
		3.5	172.9 3.5	45.9	93.3 45.9	
	Diff.	Lat.	169.4	Dep.	47.4	

PARALLEL SAILING.

IN Plane Sailing the earth is considered as an extended plane, but this supposition is very erroneous, because the earth is nearly of a spherical figure, in which the meridians all meet at the poles, consequently the distance of any two meridians measured on a parallel of latitude (which distance is called the meridian distance) decreases in proceeding from the equator to the poles. To illustrate this, let PB represent the semi-

axis of the earth, B the centre, P the pole, PCA a quadrant of the meridian, AB the radius of the equator, and CD (parallel thereto) the radius of a parallel of latitude. Then it is evident that CD will be the co-sine of AC or the co-C sine of the latitude of the point C, to the radius AB: now if the quadrantal arch \mathbf{PCA} be supposed to revolve round $oldsymbol{A}$

the axis PB, the point A will describe the circumference of the equator, and C the circumference of a parallel of latitude; and the former circumference will be to the latter as AB to CD (as may easily be deduced

^{*} Instead of putting the course S. S. E. 32 miles, and S. S. E. 39 miles, you might, make one entry only, calling it S. S. E. 71 miles. Digitized by GOOGIC

from Art. LV. Geometry;) that is, as radius to the co-sine of the latitude or the point C: hence it follows, that the length of any arch of the equator intercepted between two meridians, is to the length of a corresponding arch of any parallel intercepted between the same meridians, as radius is to the co-sine of the latitude of that parallel. Hence we obtain the following theorems.

THEOREM I.

The circumference of the equator is to the circumference of any other parallel of latitude, as radius is to the co-sine of that latitude.

THEOREM II.

As the length of a degree of the equator is to the meridian distance corresponding to a degree on any other parallel of latitude, so is radius to the co-sine of that parallel of latitude.

THEOREM III.

As radius is to the co-sine of any latitude, so are the miles of difference of langitude between two meridians (or their distance in miles upon the equator) to the distance of these two meridians on that parallel of latitude in miles.

THEOREM IV.

As the co-sine of any latitude is to radius, so is the length of any erch on that parallel of latitude (intercepted between two meridians) in miles to the length of a similar arch on the equator, or miles of difference of longitude.

THEOREM V.

As the the co-sine of any latitude is to the co-sine of any other latitude, so is the length of any arch on the first parallel of latitude in miles, to the length of the same each on the other in miles.

By means of Theorem II. the following Table was calculated, which shows the meridian distance corresponding to a degree of longitude in every latitude: and may be made to answer for any degree or minute by taking proportional parts.

The following Table shews for every degree of latitude how many miles distant the two meridians are, whose difference of longitude is one degree.

-	_					<u> </u>		W		-					
D.	L.	MI	LES.	D. L.	MII	ES.	D. L.	MII	Æ8.	D. L.	MII	LES.	D. L.	MII	ES.
_	1	59.	. 99	19	56.	75	37	47.	92	55	34.	41	73	17.	54
	2	59.	96	20	56.	38	58	47.	28	58	33.	55	74	16.	54
	3	59.	. 92	21	56.	01	39	46.	63	57	32.	68	75	15.	58
	4	59.	. 85	22	55.	63	40	45.	96	58	31.	80	76	14.	52
	5	59.	. 77	23	55.	23	41	45.	28	59	30.	90	77	13.	50
	6	59.	67	24	54.	81	42	44.	59	60	50.	00	78	12.	47
	7	59.	55	25	54.	38	43	43.	88	61	29.	09	79	11.	45
	8	59.	42	26	53.	93	44	43.	16	62	28.	17	80	10.	42
,	9	59.	26	27	53.	46	45	42.	45	63	27.	24	81	9.	59
1	0	59.	09	28	52.	98	.46	41.	68	64	26.	50	82	8.	35
1	1	58.	90	29	52.	48	47	40.	92	65	25.	36	83	7.	31
1	2	58.	69	30	51.	96	48	40.	15	66	24.	40	84	6.	27
1	5	58.	46	31	51.	43	49	39.	36	67	23.	44	85	5.	25
1	4	58.	22	32	50.	88	50	58.	57	68	22.	48	86	4.	19
1	5	57.	96	53	50.	32	51	37.	76	69	2 1.	50	87	3.	14
1	В	57.	68	34	49.	74	52	36.	94	70	20.	52	88	2.	09
1	7	57.	58	85	49.	15	58	36.	11	71	19.	58	89	1.	05
1	В	57.	06	36	48.	54	54	35.	27	72	18.	54	90	0.	00
1	В	57.	06	36	48.	54	54	35.	27	72	18.	54	90	0.	0

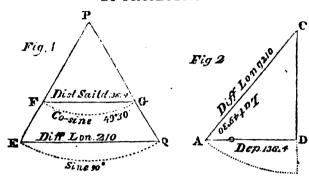
When a ship sails east or west on the surface of the earth supposed to be spherical, she describes a parallel of latitude, and this is called Parallel Sailing. In this case, the distance sailed (or departure) is equal to the distance between the meridians sailed from and arrived at in that parallel, and it is easy, by Theorem IV. (preceding) to find the difference of longitude from the distance, or the distance from the difference of longitude, as will appear plain by the following examples.

CASE I.

The difference of longitude between two places in the same parallel of latitude being given, to find the distance between them.

Suppose a ship in the latitude of 49° 30' north or south, sails directly easi or west until her difference of longitude be 80 80', required the distance sailed?

BY PROJECTION.



Take the sine of 90° from the Plane Scale, and with one foot of the compasses on P (Fig. 1) as a centre, describe the arch EQ; with the difference of longitude 210 miles in the compasses, and one foot in E, as a centre, describe an arch cutting EQ in Q; join PE, PQ. Take the sine of the complement of the latitude 40° 50′ in your compasses, and with one foot in P, as a centre, describe the arch FG, cutting PE, PQ, in FG; then the length of the chord FG being measured on the scale of equal parts will be the departure 136.4 miles.

Or this projection may be made in the following manner. Draw AD (Fig. 2) of an indefinite length, make the angle DAC equal to the latitude 490 30%, and AC equal to the difference of longitude 210 miles; draw CD perpendicular to AD; then will the line AD be the distance or departure required.

BY LOGARITHMS.

To find the departure or distance.	
As radius 90°	10.0000
Is to the difference of longitude 210	2.3222
So is co-sine latitude 49° 30'	9.8125

To the distance or departure 136,4

BY GUNTER.

The extent from radius to the complement of the latitude 40° 50' on the line of sines, will reach from the difference of longitude 210 to the distance 136,4 on the line of numbers.

BY INSPECTION.

Find the latitude among the degrees in Table II. and in the distance column the difference of longitude, opposite to which in the column of latitude will be the distance required.

In the present example the latitude is 490 80', and as the table is only calculated to single degrees, I find the numbers in the tables of 490 and 500, and take the mean of them; the former is 137,8, the latter 135,0, the mean of which is the sought distance or departure 136,4.

CASE II.

The distance between two places on the same parallel of latitude given, to find

their difference of longitude.
Suppose a ship in the latitude of 49° 30' N. or S. and long. 38° 40' W. sails directly west 136,4 miles; required the difference of longitude, and longitude in?

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2.13476

BY PROJECTION.

With the sine of the complement of the latitude 40° 30' in your compasses, and one foot in P, as a centre, (Fig. 1st. of the preceding case) describe the arch FG, upon which set off the departure 136,4 miles upon the chord FG, and through the points F and G draw the lines PE and PQ—then with the sine of 900 in the compasses, and one foot on P as a centre, describe an arch to cut PE, PQ, in E and Q; then the chord EQ being measured upon the same scale of equal parts that the departure was, will be the difference of longitude 210 miles.

Or thus, draw the line AD (Fig. 2d.) which make equal to the given distance 136,4, at D erect DC perpendicular to DA, make the angle DAC equal to the latitude; then will AC be the sought difference of longitude 210 miles.

	BY LOGARITHMS.	
As co-sine of latitude 49° 30'	9.81254 Long. left	36° 40′ W.
Is to the distance 136,4	2.13481 Diff. long.	3 30 W.
So is radius	10.00000	-
	Long. in	40 10 W.
To the diff. of long, 210	2.32227	

BY INSPECTION.

Look for the latitude among the degrees, as if it was a course, and the departure in the column of latitude;against which will stand the difference of

longitude in the distance column.

Thus in the course 49°, I seek for 136,4 in the latitude column, and find it corresponds to the distance 208; and in the course 50°, I find it nearly corresponds to 212; half the sum of 208 and 212 is 210, which is the sought difference of longitude.

Question I. A ship in the latitude of 32° N. sails due east till ber difference of longitude is 534 miles; required the distance sailed?

Annaer. 325.7 miles. Question II. A ship from the latitude of 53° 33' S. longitude 10° 18' E. salls due west 236 miles-

Question II. A ship from the latitude of 63° 35' S. longitude 10° 18' E. salls due west 236 miles—required her present longitude?

Answer. 3° 40' E.

Question III. If two ships in the latitude of 44° 30' N. distant 216 miles, should sail directly south until they were in the latitude of 32° 17' N. what distance are they from each other?

Answer. By Theorem V. 256 miles.

Question IV. A ship having run due east for three days, at the rate of 5 knots an hour, finds she las altered her longitude 8° 16; what parallel of latitude did she sail in?

Answer. 43° 28' N. or S.

MIDDLE LATITUDE SAILING.

IN salling north or south (or on a meridian) the difference of longitude is nothing, and the difference of latitude is equal to the distance sailed; but in sailing east or west (or on a parallel of latitude) the difference of latitude is nothing, and the difference of longitude may be calculated by the foregoing Theorems of Parallel Sailing. In sailing on any other course, the ship changes both her latitude and longitude; in this case, the difference of latitude, departure, and difference of longitude may be calculated by a proper application of the principles of Plane Sailing to the sailing on a spherical surface; to do which, the surface of the globe may be supposed to be divided into an indefinite number of small surfaces, as square miles, furlongs, yards, &c. which on account of their smallness, in comparison with the whole surface of the earth, may be esteemed as plane surfaces, and the difference of latitude and departure (or meridian distance) made in sailing over each of these surfaces, may be calculated by the common rules of Plane Sailing, and by summing up all the differences of latitude and departures made on these different planes, we shall obtain the whole difference of lati-tude and departure nearly.* Now, by Case I. of Plane Sailing, the distance described on any one of these small surfaces is to the corresponding differonce of latitude as radius is to the co-sine of the course, and as the course is

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I he error arising from this supposition will be decreased by increasing the number of the planes, a tlast, by increasing the number indefinitely, the error may be made less than any assignable quantity.

the same on all these surfaces, it follows that the sum of all the distances described thereon is to the sum of the corresponding differences of latitude as radius is to the co-sine of the course; that is, the whole distance sailed on the globe is to the corresponding difference of latitude as radius is to the cosine of the course. In a similar manner it appears, that the distance described on the globe is to the sum of all the corresponding departures (or meridian distances) described on these different surfaces, as radius is to the sine of the course. So that the canons for calculating the whole difference of latitude and departure from the course and distance are the same, whether the earth be esteemed as an extended plane or a spherical surface, and the same is to be observed with respect to the other cases of Plane Sailing.

We shall, therefore, in all the calculations of sailing on the spherical surface of the earth, in which the course, distance, difference of latitude and departure occur, make use of the canons already taught in Plane Sailing, and shall construct the schemes exactly in the same manner. The only additional calculation in sailing on a spherical surface consists in determining the longitude from the departure: for in sailing on a plane, the departure and longitude are the same, but in sailing on a spherical surface, the whole departure (as was observed above) is equal to the sum of all the meridian distances made in sailing over the indefinite number of small surfaces, into which we have supposed the spherical surface to be divided, and the whole difference of longitude corresponding is equal to the sum of all the differences of longitude, deduced from each of these small meridian distances by Theorem IV. of Parallel Sailing. Several methods have been proposed for abridging the calculation of the difference of longitude from the departure, the most noted of which are those known by the names of Middle Latitude Sailing and Mercator's Sailing, the latter (which will be hereafter explained) is perfectly accurate, the former is only an approximation, but it is very much used in calculating short runs and days works, but in calculating large distances across distant parallels it is liable to error. The principle on which the calculations of Middle Latitude Sailing is founded, is this:-Instead of calculating the difference of longitude corresponding to the departure made on each of the small surfaces, into which we have supposed the sphere to be divided, and adding them together, the whole departure (or sum of the meridian distances) is calculated, and the longitude deduced therefrom by the rules of Parallel Sailing, using for the latitude the arithmetical mean between the latitude sailed from and that On this supposition, we have the two first of the following theorems for calculating the departure from the difference of longitude, or the difference of longitude from the departure, which are the same as Theo. III. and IV. of Parallel Sailing, except in writing departure for distance, and middle latitude for latitude: the other theorems are easily obtained by combining the two first with the common theorems of Plane Sailing: observing that the Middle Latitude is half the sum of the two latitudes, if they are of the same name, or half their difference if of contrary names.

THEOREM I.

As radius is to the co-sine of the middle latitude, so is the difference of longitude to the departure.

THEOREM II. As the co-sine of the middle latitude is to the radius, so is the departure to the difference of longitude.

Now by case I. of Plane Sailing, the radius is to the sine of the course, as the distance sailed is to the departure, and, if we combine this analogy with Theorem II. we shall have,

Oling (in estimating the difference of longitude corresponding to each of these small meridian distances) the latitude corresponding to the middle point of the surface on which these small meridian distances are respectively made.
I This is true in theory, and would be so in practice, if the meridional difference of latitude in Table III, were given to a sufficient number of decimals, but being only given to the nearest mile or minute, the error arising from this cause, when the difference of latitude is small, is greater than the error in middle latitude sailing; in consequence of this, the method by middle latitude is almost exclusively used in the common operations on shipboard.

THEOREM III.

As the co-sine of the middle latitude is to the sine of the course, so is the distance sailed to the difference of longitude.

By Case II. of Plane Sailing, we have this analogy; as radius is to the tangent of the course, so is the difference of latitude to the departure; by combining this with Theorem II. we have

THEOREM IV.

As the co-sine of the middle latitude is to the tangent of the course, so is the difference of latitude to the difference of longitude.

Whence we easily deduce the following,

THEOREM V.

As the difference of latitude is to the difference of longitude, so is the co-sine of the middle latitude to the tangent of the course.

By means of the preceding theorems we have formed the following Table, which contains all the rules necessary for solving the various cases of Middle Latitude Sailing.

MIDDLE LATITUDE SAILING.

Case.	Given.	Sought.	SOLUTION.
	Olvein .		BOAD TOX.
1	Both Latitude: and Longitude.		Rad. : Diff. Long. : : Co-sine Mkl. Lat. : Dep. { Diff. Lat. : Rad. : : Dep. : Tang. Course. } Diff. Lat. : Diff. Long. : : Cos. Mid. Lat. : Tang. Course. { Rad. : Diff. Lat. : : Secant Course : Distance. } Sine Course : Depart. : : Rad. : Distance.
2	Both Latitudes and Departure.	Distance.	Diff. Lat Rad. : : Dep. : Tang. Course. Sine Course : Dep. : : Rad. : Distance. Co sine Mid. Lat. : Dep. : : Rad. : Diff. Long
3	One Latitude, Course and Distance.		Rad.: Dist.:: Co-sine Course: Diff. Lat. Hence: the other Latitude and Middle Latitude are found. Rad.: Dist.:: Sine Course: Departure. Co-Sine Mid. Lat.: Dep.:: Rad.: Diff. Long. Co-Sine Mid. Lat.: Sine Course:: Dist.: Diff. Long.
4	Both Latitudes and Course.		Rad. : Diff. Lat. : : Tang. Course : Departure. Co-sine Course : Diff. Lat. : : Rad. : Distance. { Co-sine Mid. Lat. : Dept. :: Rad. : Diff. Long. { Cos. Mid. Lat. : Tang. Course : : Diff. Lat. : Diff. Long.
5	Both Latitudes and Distance.	Departure.	Dist.: Rad.:: Diff. Lat.: Co-sine Course. Rad.: Dist.:: Sine Course: Departure. Co-sine Mid. Lat.: Dep.:: Rad.: Diff. Long.
6	One Latitude, Course and Departure.	Distance.	Rad. : Dep. : : Co-tang. Course : Diff. Lat. Hence the other Latitude and Middle Latitude are known. Sine Course : Departure : Rad. : Distance. Co-sine Mid. Lat. : Dep. :: Rad. : Diff. Long.
7	One Latitude, Distance and Departure.	Diff. Lat.	Dist.: Rad.:: Dep.: Sine Course. Rad.: Dist.:: Co-sine Course: Diff. Lat. Hence we obtain the other latitude and middle latitude. Co-sine Mid. Lat.: Dep.:: Rad.: Diff. Long.

We shall now proceed to illustrate these rules, by working an example in every case.

CASE I.

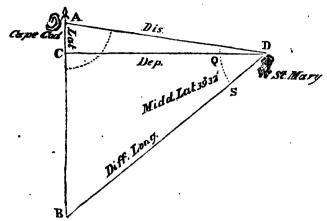
The latitudes and longitudes of two places given, to find their bearing and distance.

Required the bearing and distance between Cape Cod light-house, in the latitude of 42° 5′ N. longitude 70° 4′ W. and the island of St. Mary, (one of the Western Islands) in the latitude of 36° 59′ N. and longitude 25° 10′ W.

	t. 42° 5′ N. . 36 59 N.	42° 5′ 36 59	Long. 70° 4'W. Long. 25 10 W.
· Diff. lat.	5 6 60	Sum 79 4	44 54 60
•	-	Mid. lat. 39 32	
In miles	306	-	Diff. long. 2694 miles.

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BY PROJECTION.



Draw the east and west line DC, with the chord of 60° describe the arch QS about the centre D, to cut DC in Q; upon this arch, set off, from Q to S, the middle latitude 39° 32'; through D and S draw the line DB, which make equal to the difference of longitude 2694 miles; from B let fall upon DC the perpendicular BC, which continue towards A making AC equal to the difference of latitude 506 miles; join AD, and it is done. For by this method of construction, on the principles before explained, A will be the situation of Cape Cod, D the situation of St. Mary; CD will be the departure, which being measured will be found to be 2078 miles; the distance will be represented by AD, which being measured will be found to be 2099 miles, and the course from Cape Cod to St. Mary, will be represented by the angle CAD=31° 37'; therefore the course will be S. 81° 57' E. or E. § S. nearly.

Note. The course is put S. 81° 37′ E. because St. Mary being in a less northern latitude than Cape Cod, is to the southward of it; it is also to the castward of Cape Cod, because it is in a less western longitude.

В	Y LOGA	RITHMS.	_	
To find the departure (by The	orem I.)	To find the course.	€.	
As radius 90°		As diff. of lat. 306	2.4857 2	
Is to diff. of long. 2694	3.43040	Is to radius 45°	10.00000	
Se is co-sine mid. lat. 39° 32'	9.88720	So is the departure 2078	3.31760	
To the departure 2078	3.31760	To tang. of course 81° 37'	10.83188	
To find the distance. As radius 90° Is to the diff. of lat. 306	10.00000 2.48572	Note. The course may be found without the departure, by Theo. V. Middle Latitude Sailing.		
80 is sec. of course 81° 37'	10.83626	As the diff. of lat. 306 Is to the diff. of long. 2694	2.48572 3.43040	
To the distance 2099 Note. The log. of the depart	3.32198	So is co-sine mid. lat. 39° 32'	9.88720	
found 3.31760 is rather less than the log. of 2078=3.31765; but in finding the course by		•	13.31769 2.48579	
the departure, I have used the qui	e course by		2,40372	
at the first operation, and shall din all future calculations.	io the same	To tang. of course S1° 37'	10.83188	

Extend from the radius or 90° to 50° 28' the complement of the middle latitude, on the line of sines; that extent will reach from the difference of longitude 2694, to the departure 2078, on the line of numbers.

2dly. Extend from the difference of latitude 306, to the departure 2078,

^{*} If the place A be to the southward of D, the line AC should be set off upon the line CB, from C towards B.

on the line of numbers; that extent will reach from radius or 45°, to the

course 81° 37' on the line of tangents.

3dly. Extend from the course 81° 37' to the radius 90° on the line of sines, that extent will reach from the departure 2078 to the distance 2099 miles on the line of numbers.

BY INSPECTION.

RULE. Look for the middle latitude, as if it was a course in Plane Sailing, and the difference of longitude in the distance column, opposite to which, in the column of latitude, will stand the departure; having the difference of latitude and departure, the course and distance are found (as in Case VI. Plane Sailing) by seeking in Tab. II. with the difference of latitude and departure, until they are found to agree in their respective columns; opposite to them will be found the distance in its column, and the course will be found at the top of that table, if the departure be less than the difference of latitude, otherwise at the bottom.

Thus with one tenth of the difference of longitude 269.4 or 269, I enter Table II. and opposite to it, in the distance column of the Tables of 39° and 40°, I find 209.1 and 206.1 in the latitude column; now the middle latitude being nearly 39½°, I take the mean of these, 207.8 for the departure, which being multiplied by 10, gives the whole departure 2076. Again, I enter Table I, with one tenth of the departure 207.6, and one tenth of the difference of latitude 30.6 and find that they agree nearly to a course of 7½ points, and a distance of 210, which multiplied by 10, gives the sought distance 2100 miles nearly.

CASE II.

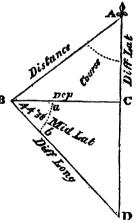
Both latitudes and departure from the meridian given, to find the course, distance, and difference of longitude.

A ship in the latitude of 49° 57′ N. and longitude of 15° 16′ W. sails southwesterly till her departure is 789 miles, and latitude in 39° 20′ N. Required the course, distance and longitude in?

Latitude left	49° 57′ N.
Latitude in	39 2 0 N.
Diff. of lat.	10 37=637 miles.
Sum of lats.	79 17
Middle lat.	44 38

BY PROJECTION.

Draw the meridian ACD, on which take AC equal to the difference of latitude 637 miles; draw CB perpendicular to AC, and make it equal to the departure 789 miles; about B as a centre describe an arch ab, on which set off the middle latitude 44° 58′; through B and b draw the line BD, meeting ACD in D; join AB and it is done; for AB will be the distance sailed, which being measured will be found=1014 miles; BD will be the difference of longitude=1109 miles, and the angle CAB will represent the course from the meridian 51° 5′.



sent the course from the men	iman or -	J •	
В	Y LOGA	RITHMS.	
To find the course.		To find the distance.	.
As the diff. of lat. 637	2.80414	As sine course 51° 5'	9.89101
Is to radius 45°	10.00000	Is to the departure 789	2.89708
So is the departure 789	2.89708	So is radius 90°	10,00000
To tang. course 51° 5'	10.09294	To the distance 1914	3.00607
To find the difference of lon As co-sine mid. latitude 44° 38' Is to the departure 789	9.85225	Longitude sailed from Diff. Long. 1109 miles	15° 16′ W. 18 29 W.
So is radius 90°	19.00000	Longitude in	33 45 W.
To diff. of long. 1109	3.04483	Digitized by 🗗 C	ogie

BY GUNTER.

1st. The extent from the difference of latitude 687 to the departure 789, on the line of numbers, will reach from radius, or 45°, to the course 51° 5′ on the line of tangents.

2dly. The extent from 51° 5' to radius, or 90°, on the line of sines, will reach from the departure 789, to the distance 1014 on the line of numbers.

3dly. The extent from the complement of middle latitude 45° 22', to radius, or 90°, on the line of sines, will reach from the departure 789 to the difference of longitude 1109 on the line of numbers.

BY INSPECTION.

Rule. With the difference of latitude and departure, find the course and distance (as in Case VI. of Plane Sailing) by seeking in Tab. II. until the difference of latitude and departure are found to correspond, against which in the distance column will be the distance; and if the departure be less than the difference of latitude, the course will be found at the top of that table, otherwise at the bottom.

Then take the middle latitude as a course, and find the departure in the latitude column, the number corresponding in the distance column will be

the difference of longitude.

In the present example, I take one tenth of the difference of latitude 657, and the departure 789; that is 63,7 and 78,9 the nearest numbers to these are 63,6 and 78,5, standing together over 51°, against the distance 101, which being multiplied by 10 gives 1010; whence the course by inspection is 8.51° W. and the distance 1010. Then I take one tenth of the departure, 78,9 and seek it in the column of latitude of 45° (which is the nearest to the middle latitude 44° 38'), the nearest number I find is 79.2, opposite which in the distance column stands 112, which being multiplied by 10 gives 1120 for the difference of longitude; this value differs a little from that found by logarithms, owing to the miles of middle latitude neglected, for if we were also to find the difference of longitude for the middle latitude 44° and proportion for the minutes, the result would come out nearly the same as by logarithms.

CASE III.

One latitude, course and distance given, to find the difference of latitude and difference of longitude.

A ship in the latitude of 42° 30′ N. and longitude 58° 51′ W. sails S. E. by S. 591 miles. Required the latitude and longitude in?

BY PROJECTION.

Draw the meridian ADE (as in case I. Plane Sailing) upon A as a centre describe an arch with the chord of 60°, and upon it set off, from where it cuts AD, the course S. E. by S. or 3 points, through that point of the arch, and the point A, draw the line AC, which make equal to the distance 591 miles; from C let fall upon AD the perpendicular CD; then will CD be the departure 328 miles, and AD the difference of latitude 491 miles. Hence we obtain the latitude arrived at, and the middle latitude; draw the line CE making an angle with DC of 38° 24′—the middle latitude; and the distance CE will be the difference of longitude 419 miles. hence the longitude in is obtained.



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BY LOGARITHMS. .

To find the	he difference of	latitude.	To find the departure	3.	
As radius 8 pts	•	10.00000	As radius 8 pts.	10.00000	
Is to the distan	ce 591	2.77159	Is to the distance 591	2.77159	
So is co-sine co	ourse 3 pts.	9.91985	So is sine course 3 pts.	9.74474	
To the diff. of	lat. 491.4	2.69144	To the departure 328.3	2.51633	
Latitude left	42° 30′ N.		To find diff, long. with departure.		
Diff. of lat.	8 1 ₁ S.		As co-sine mid. lat. 38° 24'	9.89415	
Latitude in	34 19 N.		Is to the departure 328.3*	2.51633	
Sum of lats.	76 49		So is radius 90°	10.00000	
Middle lat.	38 24		To diff. of long. 419 miles	2.62218	
Long. left	59° 51′ W.		Without the departur		
		•	As co-sine mid. lat. 38° 24' ar.	co. 0.10585	
Diff. of long. 41			Is to sine course 3 pts.	9.74474	
Longitude in	51 52 W.		So is distance 591	2.77159	
			To diff. of long. 419 miles	2.62218	

BY GUNTER.

1st. The extent from radius 8 points to the complement of the course 5 points on the line marked SR, will reach from the distance 591 to the difference of latitude 491 on the line of numbers.

2dly. The extent from radius 8 points to the course 3 points on the line SR, will reach from the distance 591 to the departure 328 on the line of numbers.

3dly. The extent from the complement of middle latitude 51° 36' to radius 90° on the line of sines, will reach from the departure 528 to the difference of longitude 419 on the line of numbers.

BY INSPECTION.

Rule. With the course and distance find the difference of latitude and departure (as in Case I. of Plane Sailing) by finding the given course at the top or bottom of the Tables, either among the points or degrees; in that page and opposite the distance taken in its column, will stand the difference of latitude and departure in their columns. Then take the middle latitude as a course and find the departure in the latitude column, against it, in the distance column will stand the difference of longitude.

Thus, under the course three points, or S. E. by S. and against the tenth of the distance 591=59,1 or 59 stand 49,1 and 32,8; these multiplied by 10 give 491 for the difference of latitude and 328 for the departure. Now taking the middle latitude 38° 24' or 58° as a course, and a tenth of the departure 328=32,8 in the column of difference of latitude (the nearest is 33,1) against which stands 42 in the distance column; this multiplied by 10 gives 420 for the difference of longitude nearly.

CASE IV.

Both latitudes and course given, to find the departure, distance and difference of longitude.

Suppose a ship sailing from a place in the latitude of 49° 57' N. and longitude of 30° W. makes a course good of S. 39° W. and then by observation is in the latitude of 45° 31' N.—it is required to find the distance run, and longitude in?

Latitude from	49° 57′ N
Lat. by observation	45 31 N
	4 26 60
Diff. lat.	266
Sum of lats.	95° 28'
Mid. lat.	47 44



[•] The logarithm of the departure was found by the preceding canon to be 2.01683, differing a little from the logarithm of 328.3.

BY PROJECTION.

Draw the meridian ACD, on which set off AC equal to the difference of latitude 266 miles; draw CB perpendicular to AC; draw the line AB, making an angle equal to the course 59° with AC, and meeting BC in B; through B draw BD, making an angle equal to the middle latitude 47° 44′ with the line BC, and it is done; for AB will be the distance 342.5 miles, BC the departure 215.4 miles, and BD the difference of longitude \$20.8 miles.

	BY LOG	ARITHMS.	
To find the departur	·e.	To find the difference of long	itude by the
As radius 45°	10.00000	departure.	
Is to the diff. of lat. 266	2.42488	As co-sine mid. lat. 47° 44'	9.82775
So is tang. course 39°	9.90837	Is to the departure 215.4	2,33325
•		So is radius 90°	10.00000
To the departure 215.4	2,33325		
To find the distance		To the diff. of long. 320.3	2.50559
As co-sine of the course 39°		The diff. of long. may be found	without the
Is to the diff. of lat. 266	2.42488		
So is radius 90°	10.00000		
		As co-sine mid. lat. 47° 44'	9.82775
To the distance 342.3		ls to tang. of course 39°	9.90837
To find the longitude		So is the diff. of lat. 266	9,42488
	30° 0′ W.	So is the dis. of lat. 500	4.22204
Diff. long. 320 miles or	5 20 W.		12.33325
Din. tong. 320 miles of	5 20 11.	1	9.89775
Paneltuda in	35 20		7.04110
Congitude in	33 ZU	To the diff. of long. 320.3	2,50550

BY GUNTER.

1st. The extent from radius 45° to the course 39° on the line of tangents, will reach from the difference of lat. 266 to the departure 215.4 on the line of numbers.

2dly. The extent from the complement of the course 51° to the radius 90° on the line of sines, will reach from the difference of latitude 266 to the distance 342.3 on the line of numbers.

3dly. The extent from the complement of the middle latitude 42° 16' to radius 90° on the line of sines, will reach from the departure 215.4 to the difference of longitude 320.3 on the line of numbers.

BY INSPECTION.

Find the course among the points or degrees (in Tab. I. or II. as in Case II. Plane Sailing) and the difference of latitude in its column, against which will stand the distance and departure in their columns; then take the middle latitude as a course, and find the departure in the latitude column, against which, in the distance column, will stand the difference of longitude.

Thus, with the course 39°, and half the difference of latitude 135, I enter Table II. the nearest number in the table is 132.9, which corresponds to the distance 171, and to the departure 107.6; these doubled give the distance 342,

and the departure 215,2 miles.

Diff. lat.

Then with the mid. lat. 47° 44′ or 48° as a course, I enter Table II. and seek for half the departure 107.6 in the lat. col. the nearest number to which is 107.7, which corresponds to the distance 161; this doubled gives the diff. of long. 322 miles nearly.

CASE V.

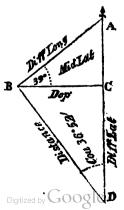
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Both latitudes and distance given, to find the course, departure, and difference of longitude.

Suppose a ship sails 300 miles north-westerly, from a place in the latitude of 37° N. and the longitude of 32° 16′ W. until she is in the latitude of 41° N.—required her course and longitude in?

Latitude left Latitude in	37° 41	0' 0	N.	, on Broad	0' N. 0
	4 60	0		Sum Mid. lat.	

240



BY PROJECTION.

Draw the meridian ACD, on which set off DC equal to the difference of latitude 240 miles; draw the line CB perpendicular to DC; take the distance 300 in your compasses, and with one foot in D, as a centre, sweep an arch cutting CB in B; join DB; make the angle CBA equal to the middle latitude 39° and draw BA cutting DCA in A, and it is done; for BC will be the departure 180 miles, BA the difference of longitude 231.6 miles, and the angle BDC will represent the angle of the ship's course with the meridian, which will therefore be N. 36° 52′ W.

BY LOGARITHMS.

		1141 1 11110			
To find the course.		To find the difference of	longitue	de t	y the
As the distance 300	2.47712	departure.	-		
Is to radius 90°	10.00000	As co-sine mid. lat. 39°		9.	89 050
So is diff. lat. 240	2.38021	Is to the departure 180.0		*2 .	25524
To co-sine course 36° 52'	9.90309	So is radius 90°			00000
To find the departure.		To diff. of long. 231.6			36474
As radius 90°	10.00000		ude in.		
Is to the distance 300	2.47712	Longitude left	32°	16'	w.
So is sine course 36° 52'		Diff. of long.	3	5 2	w.
To the departure 180.0	2.25524	Longitude in	36	8	w.
•	BY GU	NTER.			

1st. The extent from the distance 300 to the difference of latitude 240 on the line of numbers, will reach from radius 90° to the complement of the course=55° 8′, on the line of sines.

2dly. The extent from radius 90° to the course 36° 52' on the line of sines, will reach from the distance 300 to the departure 180 on the line of numbers.

Sdly. The extent from the complement of the middle latitude 51° to the radius 90° on the line of sines, will reach from the departure 180 to the difference, of longitude 231.6 on the line of numbers.

BY INSPECTION.

Find the course (as in Case IV. Plane Sailing) by seeking in Table II. till against the distance taken in its column, is found the difference of latitude in one of the following columns; adjoining to it will stand the departure; which if less than the difference of latitude, the course is to be found at the top of the Table, but if greater, at the bottom; then take the middle latitude as a course, and find the departure in the column of difference of latitude, against which, in the distance column, will stand the difference of longitude.

Thus the distance 300, and the difference of latitude 240, are found to correspond nearly to a course of 37°, and a departure of 180.5: then taking the middle latitude 39° as a course, I seek the departure 180.5 in the latitude column, corresponding to which, in the distance column, is the difference of

longitude 232.

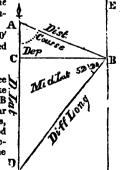
CASE VI.

One latitude, course, and departure given, to find the difference of latitude, distance, and difference of longitude.

. A ship in the latitude of 50° 10' S. and longitude of 30° 00' E. sails E. S. E. until her departure is 957 miles; required her distance sailed, and latitude and longitude in?

BY PROJECTION.

Draw the meridian ACD, and parallel thereto at a distance equal to the departure 957 miles, draw the line EB; make the angle CAB equal to the course 6 points, and draw AB meeting EB in B; from B let fall upon AD the perpendicular BC; then will AC be the difference of latitude 396.4 miles, and AB the distance sailed 1036 miles; having thus obtained the middle latitude 53° 28', make the angle CBD equal thereto, and draw BD meeting ACD in D; then will BD be the difference of longitude 1608 miles.



^{*}This logarithm, by the preceeding operation, was found equal to 2.25524, differing a little from the logarithm of 1800.

BY LOGARITHMS.

To find the diff.	of latitude.	To find the distance.			
As radius 4 points	10.00000	As sine course 6 points	9.9656%		
Is to the departure 957	2.98091	Is to the departure 957	2.98091		
So is co-tang. course 6	points 9.61722	So is radius 8 points	10.00000		
To the diff. of lat. 396,4	2.59813	To the distance 1036	3.01529		
Latitude left	50° 10′ S.	To find the diff. of long	ritude. 🤄		
Diff. of lat. 396 miles	6 36 S.	As co-sine mid. lat. 53° 28'	9.77473		
Latitude in	56 46 S	Is to the departure 957	2.98091		
Sum of latitudes	106 56	So is radius 90°	10.00000		
Middle latitude	53 2 8	To the diff. of long. 1608	3.20618		
	Longitude lest	30° 00′ Ĕ.	·		
	Diff. of long. 1608	= 26 48 E.			
	Long. in	56 48 E.			
	BY GU	NTER.			

1st. The extent from the course 6 points to the radius 4 points, on the line marked TR, will reach from the departure 957, to the difference of latitude 396,4 on the line of numbers.

2dly. The extent from 6 points to the radius or 8 points, on the line marked SR, will reach from the departure 957, to the distance 1036, on the line of numbers.

3dly. The extent from the complement of the middle latitude 36° 32' to the radius 90°, on the sines, will reach from the departure 957, to the difference of longitude 1608, on the line of numbers.

BY INSPECTION.

Find the course among the points or degrees, Tab. I. or Tab. II. (as in Case III. Plane Sailing) and the departure in its column, corresponding to which, in the columns of distance and difference of latitude, will be found the distance and difference of latitude respectively; then with the middle latitude as a course, seek the departure in the column of latitude, corresponding to which, in the distance column, will stand the difference of longitude.

Thus, I enter Table I. above E. S. E. or 6 points, and seek for one-tenth of the departure 95,7, the nearest to which is 96,1, and the corresponding numbers are 104 and 39,8, which multiplied by ten gives the distance 1040 and the difference of latitude 398 nearly; the middle latitude being nearly 53½°, I seek in the Table of 55° for the distance corresponding to a tenth of the departure=95,7 and find it to be 159; then I seek for the same number 95,7 in the Table of 54°, and find the number corresponding in the distance column to be 165, half the sum of these two numbers is 161, which multiplied by 10 gives the difference of longitude 1610 nearly.

CASE VII:

One latitude, distance sailed, and departure from the meridian given, to find the course, difference of latitude, and difference of longitude.

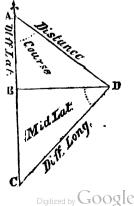
A ship in the latitude of 49° 30' N. and longi-

A ship in the latitude of 49° 30′ N. and longitude of 25° 0′ W. sails south-easterly 645 miles until her departure from the meridian be 500 miles; required the course steered, and the latitude and longitude the ship is in?

BY PROJECTION.

Draw the line BD equal to the departure 500 miles, and perpendicular thereto draw the meridian line ABC; take an extent equal to the distance 645 in your compasses, and with one foot in D as a centre, describe an arch cutting AB in A. join AD, then will AB be the difference of latitude 407,5 miles, and BAD the course, S. 50° 48′ E. Hence we have the latitude in, and middle latitude; make the angle BDC equal to the middle latitude and draw DC cutting ABC in C, then DC will be the difference of longi-

tude 721.1 miles.



DV TOGADIMING

.	I LUGA	MITHMS.				
To find the course.		Latitude left		490	30 ′	N.
As the distance 645	2.80956	Diff. of lat. 408		, 6	43	S_s
Is to radius 90°	10.00000	!				
So is the departure 500	2.69897	Latitude in		42	42	Ŋ.
To sine-course 50° 49'	9.88941	Sum of the latitudes		95	12	
To find the difference of la	ıtitude.	Middle latitude		46	6	
As radius 90°	10.00000	1				
Is to the distance 645	2.80956	ŧ				
So is co-sine course 50° 49'	9.80058	ļ				
To the diff. of lat. 407.5	2.61014				,	
To find the difference of los	ngitude.	Longitude left		25°	0'	W,
As co-sine mid. lat. 46° 6'	9.84098	Diff. long. 721	=	12	1	E.
Is to the departure 500	2.69897	1				
Se is radius 90°	10.00000	Longitude in		12	59	W,
To the diff. of long. 721.1	2.85799					

BY GUNTER.

1st. The extent from the distance 645, to the departure 500, on the line of numbers, will reach from the radius 90°, to the course 50° 49' on the line

2dly. The extent from radius 90°, to the complement of the course 59° 11', on the line of sines, will reach from the distance 645 to the difference

of latitude 407,5 on the line of numbers.

sdly. The extent from the complement of the middle latitude 43° 54', to the radius 900 on the line of sines, will reach from the departure 500, to the difference of longitude 721,1, on the line of numbers.

BY INSPECTION.

As in Case V. Plane Sailing, find the course by seeking in Table II. till against the distance, in its column, is found the given departure in one of the following columns, adjoining to which in the other column will be the difference of latitude, which if greater than the departure the course will be at the thp, but if less the course will be found at the bottom. Then take the middie latitude as a course, and find the departure in the column of difference of latitude, against which, in the distance column, will be found the difference of longitude.

Thus, one-third of the distance, 215, and one-third of the departure, 166,7, are found nearly to correspond to a course of 51 degrees, and a difference of latitude of 135,5, which multiplied by 3 gives the sought difference of latitude 406 nearly; then with the middle latitude, 46°, as a course, I enter the Table, and seek for one-fifth of the departure = 100, in the latitude column, the distance corresponding, 144, being multiplied by 5 gives the differ-

ence of longitude 720 nearly.

QUESTIONS FOR EXERCISE.

Question I. Required the bearing and distance between two places, one in the latitude of 37° 55′ N. and longitude of 54° 25′ W. the other in the latitude of 32° 58′ N. and longitude of 17° 5′ W.?

Answer. S. 80° 9′ E. and N. 80°, 9′ W. distance 1854 miles.

Question II. Required the direct course and distance, from a place in the latitude of 36° 55′ S. and longitude of 20° 0′ E. to another place in the latitude of 32° 88′ S. and longitude of 8° 54′ W.?

Answer. N. 790 46' W. distance 1447 miles.

Question III. A ship from the latitude of 37° 30' S, and longitude of 60° E. sails N. 79° 56' W. 202 miles; required the latitude and longitude in? Answer. Latitude 36° 55' S. longitude 55° 50' E.

Question IV. A ship from the latitude of 34° 35' N. and longitude of 45° 16' W. sails S. 85° 56' E. 101 miles; required her latitude and longitude?

Answer. Latitude 34° 24' N. longitude 48° 14' W.

MERCATOR'S SAILING.

THE calculations by Middle Latitude Sailing are sufficiently exact for a short run, or a day's work, and are to be preferred in all cases where the difference of latitude is small in comparison of the difference of longitude; but this method is liable to great errors in calculating the situations of places differing greatly in latitude and longitude, particularly in high latitudes; to remedy this inconvenience, a chart was invented and published in the year 1566, by GERARD MERCATOR, a Flemish Geographer, in which all the meridians are parallel to each other, but proportionally lengthened so as to conform to the spherical figure of the earth. The principles on which this chart is constructed were first explained in the year 1599, by Edward Wright, an

Englishman, and are as follows.

By Theorem II. of Parallel Sailing, the distance of two meridians corresponding to a degree or mile of longitude, in any latitude, is to the length of a corresponding degree or mile of the meridian, as the co-sine of the latitude is to the radius, that is (by Art. LVI. Geo.) as radius is to the secant of the latitude. Hence, if the meridians are supposed to be parallel to each other, or the distance of the meridians to remain the same in every latitude, the degree or mile of latitude must be increased in proportion to the secant of the Therefore, if the radius be supposed to be equal to one mile, the length of the first mile of latitude from the equator will be represented by the secant of 1', the second mile by the secant of 2', the third mile by the secant of 3', &c. Therefore the length of the expanded arch of the meridian may be found by a continual addition of secants, to every degree and minute of the quadrant, as in Table III. by means of which the chart (called Mercator's chart) may be constructed, and all the cases of Mercator's Sailing may be projected and calculated.*

In using this table, the degrees are to be found at the top or the bottom. and the miles at the side; in the angle of meeting will be the length of the corresponding expanded arch, usually called the meridional parts. If you wish to find the arch of the expanded meridian intercepted between any two parallels, or, as it is usually called, the meridional difference of latitude, you must, when both places are on the same side of the equator, subtract the meridional parts of the lesser latitude from the meridional parts of the greater, the remainder will be the meridional difference of latitude: but if they are on different sides of the equator, the sum of the meridional parts of both latitudes

will be the meridional difference of latitude required.

EXAMPLE I.

Required the meridional parts corresponding to the latitude of 42° 34'? Look in the bottom or top of the table for 42°, marked 42 d. and in the right or left hand column marked (M) for 34', under the former and opposite the latter stand 2828, the meridional parts corresponding to 42° 54'.

EXAMPLE II. Required the Meridional dif-EXAMPLE III. Required the meridional difference of latitude between Cape Cod, in the lat. of 42° 5' N. and the Island of St. Mary, in the latitude of 36° 59' N.? Cape Cod's lat. 42° 5' N. Mer. parts 2788 Lat. 35° 12' N. Mer. par. 2259 St. Mary's lat. 36° 59' N. Mer. parts 2391 C. of G. Hope's lat. 34° 26' S. Mer. par. 2203

ference of latitude between a place in the lat. of 35° 12' N. and the Cape of Good Hope, in the latitude of 34° 26' S.?

397 Sum is meridional difference of lat. Meridional difference of latitude

From these principles it follows, that in sailing upon any course, the true or proper difference of latitude is to the departure as the meridional difference of latitude is to the difference of longitude. Hence if MI (in the figure of Case I. following) be the proper difference of latitude, IO the departure,

The manner of constructing this chart will be particularly explained hereafter. It may be observed, that the smaller the subdivisions of the arch of the meridian are, the greater will be the accuracy of the relumined length of the expanded arch of the meridian. To be perfectly accurate the arch ought to be addivided into the smallest quantities possible. Attention was paid to this circumstance in calculating re above-mentioned Table.

MO the distance, the angle IMO the course, and we take MT equal to the meridional difference of latitude, and draw TH parallel to IO to cut MO continued in H; the line TH will represent the difference of longitude: for fby Art. LIII. Geom.) MI: IO:: MT: TH. Now in the triangle MTH, by making MT radius, we have MT: radius: TH: tang. TMH, that is the meridional difference of latitude is to radius, as the difference of longitude is to the tangent of the course. By making MH or TH radius we shall have other analogies, which combined with those in Plane Sailing, furnish the solutions of the various cases of Mercator's Sailing contained in the following table.

MERCATOR'S SAILING.

Case.	Given.	Sought	SOLUTIONS.
1	Both Latitudes and Longitudes.	Course. Distance. Departure	Having both lats. the mer. diff. lat. is found by Table III. Mer. Diff. Lat.: Rad.:: Diff. Long.: Tang. Course. (Rad.: Prop. Diff. Lat.:: Secant Course: Distance. (Co-Sine Course:: Prop. Diff. Lat.:: Rad.: Distance. (Rad.: Prop. Diff. Lat.:: Tang.: Course:: Departure. Mer. Diff. Lat.: Diff. Long.:: Prop. Diff. Lat.: Depart.
2	Both Latitudes and Departure.	Course. Distance. Diff. Long.	Merid. Diff. Lat. being found by Table III. we have Prop. Diff. Lat.: : Radius :: Departure: Tang. Course. (Radius: Prop. Diff. Lat.: : Sec. Course: Distance. (Sine Course: Distance. (Sine Course: Diff. Long. (Rad.: Mer. Diff. Lat.: Tang. Course: Diff. Long. Prop. Diff. Lat.: Dep.:: Mer. Diff. Lat.: Diff. Long.
•	One Latitude, Course and Distance.	Diff. Lat.	Radius : Distance : : Sine Course : Departure. Rad: : Dist. : : Oo-dine Course : Prop. Diff. Lat. Hence we have the other lattude and mer. diff. lat. by Table III. Rad: : Mer. Diff. Lat. : : Tang. Course : Diff. Long.
4	Both Latitudes and Course.	Departure.	Co-sine Course: Prop. Diff. Lat.:: Rad.: Dist. Rad.: Prop. Diff. Lat.:: Tang. Course: Departure. Mer. diff. lat. being found in Table III. we have Rad.: Mer. Diff. Lat.:: Tang. Course: Diff. Long.
5	Both Latitudes and Distance.	Departure.	Dist.: Rad.:: Prop. Diff. Lat.: Oo-sine Course. Radius: Distance:: Sine-Course: Departure. Rad.: Mer. Diff. Lat.:: Tang. Course: Diff. Long.
6	One Latitude, Course and Departure.		Rad.: Dep.:: Co-tang. Course: Prop. Diff. Lat. Hence we have the other latitude and mer. diff. latitude. Sine Course: Departure:: Radius: Distance. S Rad.: Mer. Diff. Lat.:: Tang. Course: Diff. Long. Prop. Diff. Lat.: Dep.:: Mer. Diff. Lat.: Diff. Long.
7	One Latitude, Distance and Departure.	Course. Diff. Lat. Diff. Long.	Dist.: Rad.:: Dep.: Sine Course. Rad.: Dist.:: Co-sine Course: Diff. Lat. Hence we obtain the other latitude and meridian difference latitude. { Rad.: Mer. Diff. Lat.: Tang. Course: Diff. Long. } Prop. Diff. Lat.: Dep.:: Mer. Diff. Lat.: Diff. Long.

CASE I.

The latitudes and longitudes of two places given, to find the direct course and distance between them.

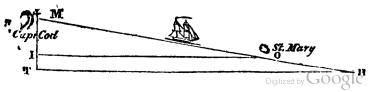
Required the bearing and distance from Cape Cod Light House in the latitude of 42° 5' N. and longitude 70° 4' W. to the island of St. Mary, one of the Western Islands; in the latitude of 36° 59' N. and longitude of 25° 10' W.?

Cape Cod's lat. 42° 5' N. St. Mary's lat. 36 59 N.	Meridional parts 2788 Meridional parts 2391	Long. 70° 4' W. 25 10 W.
5 6	Mer. diff. lat. 397	44 54
. 60		6 0

Difference of lat. 306 miles.

Diff. of long. 2694 miles.

BY PROJECTION.



Draw the meridian MT equal to the meridional difference of latitude 597 miles, set off also upon it MI equal to the proper difference of latitude 306 miles; perpendicular to MT draw TH and IO, make TH equal to the difference of longitude 2694 miles, draw MH cutting IO in O; then will the angle TMH be the course S. 81° S7′ E. and OM the distance 2099 miles.

BY LOGARITHMS.

To find the course.	To find the distance.					
As the mer. diff. of lat. 397	2.59879 As radius 90°	10.00000				
Is to radius 45°	10.00000 Is to the proper diff. of lat. 306	2.48572				
So is the diff. of long. 2694	3.43040 So is secant of course 81° 37'	10.83626				
To tang. of course 81° 37'	10.83161 To the distance 2099 miles	3.32198				

BY GUNTER.

1st. Extend from the meridional difference of latitude 397 to the difference of longitude 2694 on the line of numbers; that extent will reach from the radius or 45°, to the course 81° 37′ on the line of tangents.

2dly. Extend from the complement of the course 8° 23' to radius 90° on the line of sines, that extent will reach from the proper difference of latitude 506, to the distance 2099 on the line of numbers.

BY INSPECTION.

With the meridional difference of latitude and difference of longitude used as difference of latitude and departure, find the course, by inspecting the tables until those numbers are found to correspond; with this course and the proper difference of latitude, find the corresponding distance.

Thus one tenth of the merid. diff. lat. and diff. long. are found to agree nearly to a course of 74 points; this course and one tenth of the proper difference of latitude 30,6 is found to correspond to the distance 209; this multiplied by 10 gives the distance 2090, differing a little from the result by lagarithms, owing to the neglect of a few minutes in the course.

CASE II.

Both latitudes and the departure given, to find the course, distance, and difference of longitude.

A ship in the latitude of 49° 57' N. and longitude of 15° 16' W. sails south-westerly until her departure is 789 miles, and then by observation is in the latitude of 39°20' N. required her course, distance and longitude in?

Lat. left 49° 57' N. Mer. parts 3470 Lat. in 39 20 N. Mer. parts 2571

Diff. lat. 10 37=637 m. Mer. diff. lat. 899

Diff. Long. B

BY PROJECTION.

With the proper difference of latitude and departure, project as in Case VI. Plane Sailing, by drawing the meridian AEB, on which take AE equal to the proper difference of latitude 637 miles; erect ED perpendicular to AE and make it equal to the departure 789 miles; join AD and continue it towards C: make AB equal to the meridional difference of latitude 899 miles, and draw BC perpendicular to AB, to cut AC in C, and it is done. For AD will be the distance 1014 miles, BC the difference of longitude 1114 miles, and the angle BAC will be the course S. 51° 5′ W.

BY LOGARITHMS.

.55	I LUG	ARITHMS.	
To find the course.		To find the distance.	
As the proper diff. of lat. 637	2.80414	As radius	10.00 00 0
Is to radius 45°	10.00000	Is to prop. diff. of lat. 637	2.80414
So is the departure 789	2.99708	So is sec. course 51° 5'	10.20191
To tang. course 51° 5'		To the distance 1014	3.00603
To find the diff. of long.		Longitude lest	15° 16′ 📆.
As radius 45°	10.00000	Diff. of long. 1114 =	19 34 W.
Is to mer. diff. of lat. 899	2.95376	Longitude in	33 50 W.
So is tang. course 51° 5'	10.07274	The diff. of long, may also	be found by
To diff. of long. 1114	3.04670	saying, as prop. diff. of lat.	dep : mer.
-		diff. lat. : diff. of long.	00816

BY GUNTER.

The extent from the diff. of lat. 637 to the dep. 789 on the line of numbers, will reach from radius 45° to the course 51° 5' on the line of tangents.

The extent from the course 51° 5′ to radius 90° on the sines, will reach from the departure 789 to the distance 1014 on the line of numbers.

3dly. The extent from the radius 45° to the course 51° 5' on the line of tangents, will reach from the merid. diff. of lat. 899 to the difference of Iongitude 1114 on the line of numbers.

BY INSPECTION.

Find the course by Plane Sailing, Case VI. by seeking in the tables with the proper difference of latitude and departure till they are found to agree in their respective columns, corresponding to which will be the distance in its column, and the course will be found at the top of that column if the departure is less than the proper difference of latitude, otherwise at the bottom; with the same course, find the meridional difference of latitude in the latitude column, corresponding to which in the departure column will be the true difference of longitude.

Thus with one tenth of the true difference of latitude and departure 65,7 and 78,9, I find the course 510, and the distance 101, which multiplied by 10 gives nearly the true distance 1010; in the same table, opposite to one tenth of the meridional difference of latitude 89,9 I find the departure 111,1, which

multiplied by 10 gives the difference of longitude 1111 miles.

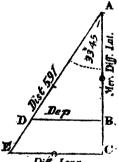
CASE III.

One latitude, course and distance given, to find the difference of latitude and difference of longitude.

A ship in the latitude of 420 30' N. and longitude of 580 51' W. sails S. W. by S. 591 miles; required the latitude and longitude in?

BY PROJECTION.

Draw the meridian ABC, and ADE, making an angle with it equal to the course 3 points, make AD equal to the distance sailed, 591 miles, and from D let fall upon AB the perpendicular BD: then will BD be the departure, and AB the difference of latitude, 491 miles. Hence we have both latitudes, and the meridional difference of latitude, to which make AC equal, and draw CE parallel to BD meeting ADE in E, then will CE be the difference of longitude, 419,6 miles.



BY LOGARITHMS. Diff. Long. To find the diff. of latitude. To find the diff. of longitude. As radius 8 points 10.00000 As radius 4 points 10.00000 Is to the distance 591 2.77159 is to the mer. diff. of lat. 628 2.79796 So is co-sine course 3 points 9.91985 So is tang. course 3 points 9.82489 To prop. diff. lat. 491.4 2.69144 To diff. of long. 419.6 2,62225 42° 30' N. Mer. parts 2822 Long. left Lat. left 58° 51′ W. Diff. Lat. 491=8 11 S. Diff. of long. 420= 7 00 W. 34 19 N. Mer. parts 2194 Lat. in Long. in 65 Mer. diff. lat. 628

BY GUNTER.

1st. The extent from radius 8 points to the complement of the course 5 points, on the line marked SR, will reach from the distance 591 to the difference of latitude 491.4 on the line of numbers.

The extent from the radius 4 points to the course 3 points, on the line marked TR, will reach from the meridional difference of latitude 628 14, the difference of longitude 419.6 on the line of numbers.

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BY INSPECTION.

As in Case I. Plane Sailing, find the course at the top or bottom of the tables, either among the points or degrees, and in that page, opposite the distance, will be found the difference of latitude and departure in their respective columns. Then in the same table find the meridional difference of latitude in the latitude column; corresponding to which, in the departure column, will be the difference of longitude.

Thus, under the course S. W. by S. or 3 points, and opposite one third of the distance 197, stands 165,8 in the latitude column, which multiplied by 3 gives the difference of latitude 491,4 miles; then find one fourth of the meridional difference of latitude 157 in the latitude column, against which stands 105 in the departure column, which multiplied by 4 gives 420 the dif-

ference of longitude.

CASE IV.

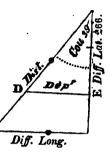
Both latitudes and course given, to find the distance and difference of longitude.

A ship from the latitude of 49° 57′ N. and longitude of 30° W. sails S. 59° W. till she arrives in the latitude of 45° 31′ N. Required the distance run and longitude in?

Lat. left 49° 57' N. Lat. in 45 31 N. Mer. parts 3470 Mer. parts 3074

Diff. lat. 4 26=266 m.

Mer. diff. lat. 396 miles.



2,50607

BY PROJECTION.

Draw the meridian AEB, on which take AE equal to the proper difference of latitude 266 miles, and AB equal to the meridional difference of latitude 396 miles; make the angle BAC equal to the course 39°, and draw ED, BC, perpendicular to AB, cutting ADC in D and C; then will AD be the distance 342 miles, and BC the difference of longitude 321 miles.

BY LOGARITHMS.

To find the distance. As the co-sine course 39° Is to the prop. diff. of lat. 266 So is radius 90°

To find the diff. of longitude.

9.89050 As radius 45°

2.42488 Is to mer. diff. of lat. 396
So is tang. course 39°

10.00000 So is tang. course 39°

2.59770
9.90837

To the distance 342.3

2.53438|To the diff. of long. 320.7 Longitude left 30° 0′ W. Diff. of long. 321 = 5 21 W.

Longitude in

35 21 W.

BY GUNTER.

1st. The extent from the complement of the course 51° to the radius 90° on the sines, will reach from the proper difference of latitude 266, to the distance 342,3 on the line of numbers.

2dly. The extent from radius 45° to the course 39° on the line of tangents, will reach from the meridional difference of latitude 396, to the difference of longitude 321 on the line of numbers.

BY INSPECTION.

As in case II. Plane Sailing, find the course among the points or degrees and the proper difference of latitude in its column, adjoining to which will be the distance and departure in their respective columns; then in the same table, find the merid. diff. of lat. in the lat. column, adjoining to which in the departure column, will be the difference of longitude.

Thus, under the course 390 and opposite the half diff. of lat. 133 (the

nearest to which is 132.9) stand 171 and 107,6, these doubled give the distance 342 and departure 215.2; and in the same table opposite the half mer. diff. of lat. 198 found in the latitude column, stands 160.5 in the dep. column, which doubled gives the difference of longitude 321 miles, nearly as before.

CASE V.

Both latitudes and distance given, to find the course, and difference of longitude.

A ship from the latitude of 37° N. and longitude of 320 16' W. sails 300 miles north-westerly, until she is in the latitude of 41° N. Required the course steered and longitude in?

Lat. left 37° N. Lat. in 410 N.

Mer. parts 2393 Mer. parts 2702

40=240 miles. Mer. diff. lat. 309 miles. Diff. lat.

C Diff. Long. R

BY PROJECTION.

Draw the meridian ABC; make AB equal to the proper difference of latitude 240, and AC equal to the meridional difference of latitude 309 miles, draw BD and CE perpendicular to ABC; with an extent equal to the dis-Cance 300 in your compasses, and one foot in A as a centre, describe an arch cutting BD in D; draw AD, which continue to cut CE in E, and it is done; for the angle BAD is equal to the course of 36° 52′, BD is the departure, and CE is the difference of longitude 231.7 miles.

BY LOGARITHMS.

To find the course. As the distance 300 Is to radius 90°

2.47712 As radius 45°
10.00000 Is to the mer. diff. of lat. 309 2.38021 So is tang. course 36° 52'

To find the diff. of longitude. 10.00000 2.48996 9.87501

to is prop. diff. of lat. 240 To comine course 36° 58'

9.90309 To the diff. long. 231.7 eft 32° 16' W. Longitude lest

2.36497

Longitude in

36 08 W.

BY GUNTER.

Diff. of longitude 232 = 3 52 W.

1st. The extent from the distance 300 to the proper difference of latitude 240, on the line of numbers, will reach from the radius or 90° to 55° 8′ the complement of the course on the line of sines.

2dly. The extent from radius 45° to the course 36° 52′ on the line of

tangents, will reach from the meridional difference of latitude 309 to the

difference of longitude 231.7, on the line of numbers.

BY INSPECTION.

As in Case IV. Plane Sailing, seek in the table till against the distance taken in its column is found the given difference of latitude in one of the following columns; adjoining to it will stand the departure, which if less than the difference of latitude, the course will be found at the top, other-, wise at the bottom; in the same table find the meridional difference of latitude in the latitude column, adjoining to which in the departure column will stand the difference of longitude.

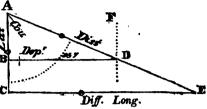
Thus the distance 300 and the difference of latitude 240, are found to correspond to a course of 570, and a departure of 180.5; and in the latitude column, opposite half the meridional difference of latitude 154.5 (the nearest to which is 154.1) stands 116.2 in the departure column, which doubled

gives the difference of longitude 252.4.

CASE VI.

One latitude, course and departure given, to find the distance, difference of latitude, and difference of longitude.

A ship from the latitude of 50° 10' S. and longitude of 30° E. sails E. S. E. until her departure is 957 miles; required the distance sailed, and the latitude and longitude in?



BY PROJECTION.

Draw the meridian ABC, and at a distance from it equal to the departure 957 miles, draw the line FD parallel to ABC; make the angle BAD equal to the course 6 points, draw AD to cut FD in D; from D let fall upon AB the perpendicular DB; then will AD be the distance 1056 miles, AB the difference of latitude 596 miles; hence we have both latitudes, and the meridional difference of latitude 667 miles, make the line AC equal thereto, and draw CE perpendicular to AC meeting AD continued in E; then will CE be the difference of longitude 1610 miles.

BY LOGARITHMS.

To find the distance.	Lat. left 50° 10' S. Mer. parts 3490
As the sine course 6 points	9.96562 Diff. lat. 396'=6 36 S.
Is to the departure 957	2.98091 Tak in FC AF C 36
So is radius 8 points	2.98091 Lat. in 56 46 S. Mer. parts 4157
To the distance 1036	
	3.01529 As radius 4 points 10.00000 10.00000
A 3' 4 i.A.	10.00000 So is tang. course 6 points 10.36278
As radius 4 points	10.00000 So is tang. course 6 points 10.36978
	2.98091 9.61722 To diff. long. 1610 = 26° 50' E. 3.90691
So is co-tang course 6 points	
To prop. diff. of lat. 396.4 miles	2.59813 Long. left 30 00 E.
20 propriation of the open among	2.59813 Long. left 30 60 E. Long. in 56 50 E.

BY GUNTER.

1st. The extent from the course 6 points to radius 8 points on the line marked S. R. will reach from the departure 957 to the distance 1036 on the line of numbers.

2dly. 'The extent from radius 4 points to the complement of the course 2 points, on the line marked T. R. will reach from the departure 957 to the difference of latitude 596 on the line of numbers.

3dly. The same extent (from the radius 4 points to the course 6 points on the line marked T. R.) will reach from the meridional difference of latitude 667, to the difference of longitude 1610, on the line of numbers.

BY INSPECTION.

As in Case III. Plane Sailing, find the course either in Table I. or Table II. and the departure in its column, corresponding to which will stand the distance and difference of latitude in their respective columns in the same Table find the meridional difference of latitude, in the latitude column, corresponding to which, in the departure column, will be found the difference of longitude.

Thus, over the course E. S. E. or 6 points, and against one-fifth of the departure 191.4 stand 79.2 and 207, which multiplied by 5 give the difference of latitude 596 miles, and the distance 1035 miles; then in the latitude column find a tenth of the meridional difference of latitude 66.7, the nearest to that is 66.8, against which, in the departure column, stands 160.8, which multiplied by 10 gives 1608, the difference of longitude.

MERCATOR'S SAILING.

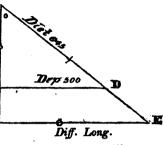
, CASE VII.

latitude, distance sailed, and departure given, to find the course, difference of latitude, and difference of longitude.

ship in the latitude of 49° 30' N. and the longitude of 25° W. sails southerly 645 miles, making 500 miles departure; required the course steered, the latitude and longitude in?

BY PROJECTION.

raw the meridian ABC, and on A point of it draw BD perpendicular eto, and make it equal to the depar-500 miles; with an extent equal to listance 645 miles in your compasuand one foot on D as a centre dee an arch to cut AB in A, join AD; will AB be the proper difference titude 407.5 miles, and the angle B) will be the course 50° 49'; hence aye the other latitude, and the menal difference of latitude, to which C e AC equal; and draw CE parallel



e AC equal; and draw CE parallel

D, meeting AD produced in E; then will CE be the difference of tude, 722.6 miles.

BY LOGARITHMS.

To find the course,		To find the diff. of lat.
e distance 645 radius 90° the deporture 500	10.00000	As radius 90° 10.00000 Is to the distance 645 2.90956 So is as sine source 50° 40′ 9.90058
the departure 500 fee of course 50° 49' To find the diff. of long. dius 45° ner. diff. of lat. 589 tang. course 50° 49' ff. of long. 722.6 Or thus: op. diff. of lat. 407.5* the departure 500 the mer. diff. of lat. 589	2.69897 9.88941 10.00000 2.77012 10.08879 2.85891 2.61014 9.60897	So is co-sine course 50° 49′ 9.80058 To diff. l. 407.5=6° 48′ S. 2.61014 Lat. left 49 30 N. M. par. 3428 Lat. in 42 42 N. M. par. 2839 Mer. diff. lat. Long. left 25 00 W. Diff. long. 12 3 E. Long. in 12 57 W. Hence the ship's course is S. 50° 49′ E. Lat. in 42° 42′ N. Long. in 12° 57′ W.
ff. of long. 722.7	2.85895	

BY GUNTER.

.. The extent from the distance 645 to the departure 500 on the line imbers, will reach from the radius 90° to the course 50° 49′ on the line nes.

ly. The extent from radius 90° to the complement of the course 59° n the line of sines, will reach from the distance 645 to the difference of 42 407.5 on the line of numbers.

ly. The extent from the radius 45° to the course 50° 48′ on the line of ents, will reach from the mer. diff. of lat. 589 to the difference of longi-722.6 on the line of numbers. Or, the extent from the proper differof latitude 407.5 to the departure 500, will reach from the meridionalence of latitude 589 to the difference of longitude 722.7 on the line of pers.

BY INSPECTION.

nd the course and difference of latitude, as in Case V. Plane Sailing, by ng in Tab. II. till the distance and departure are found to correspond in respective columns, adjoining to which, in the column of latitude, will

^{*} This log. was found above—it differs a little from the log. of 407.5.

be found the true difference of latitude, which if greater than the departure the course will be found at the top; but if less, the course will be found at the bottom: with this course seek the meridional difference of latitude in the latitude column, adjoining to which in the departure column will be found the difference of longitude.

Thus one-third of the distance 215, and one-third of the departure 166,7 are found nearly to correspond to a course of 51 degrees, and a difference of latitude of 135.3, which multiplied by 3, gives the true difference of latitude 406 nearly. Then one-fourth of the meridional difference of latitude 147, in the latitude column, is found nearly to correspond to the departure 131.9; this multiplied by 4, gives 727.6 the difference of longitude nearly.

Having explained the method of calculating single courses by Middle Latitude and Mercator's Sailing, it now remains to explain the method of calculating compound courses. To do this, you must construct a Traverse Table, and find the difference of latitude and departure for each course and distance, as in Traverse Sailing, and from thence the whole difference of latitude, departure, and latitude in, with the departure and latitudes, find the difference of longitude and longitude in, as in Case II. of Middle Latitude or Mercator's Sailing.

This method is exact enough for working any single day's work at sea, except in high latitudes, where it will be a little erroneous; in this case the difference of longitude and longitude in, may be calculated for every single course and short distance; but in general this nicety in calculation may be neglected.

and short distance; but in general this nicety in calculation may be neglected.

To illustrate the method of working compound courses, we shall here work an example, by Middle Latitude and Mercator's Sailing.

EXAMPLE.

TRAVERSE TABLE.

2312121-12 2212-						
A ship from Cape Henlopen, in the						
latitude of 380 47' N. longitude 750 17'						
W. sails the following true courses,						
viz. E. by S. 20 miles, E. N. E. 15						
miles, S. E. 26 miles, South 16 miles,						
W. S. W. 6 miles, N. W. 10 miles,						
and East 30 miles: required her lati-						
tude and longitude? .						

By constructing the Traverse Table with these courses and distances, it appears that the ship has made 27.8 miles of southing, and 69.3 miles of easting; and by subtracting the southing from the latitude of Cape Henlopen there remains the latitude in 58° 19' N.

Middle latitude

Courses.	Dis.	Diff. N.	Lat. S.	Depa E.	rture. W.
E. by S. E. N. E.	15	5.7		19.6 1 3. 9	-4
	26 16		18.4 16.0 2.5	18.4	
N. W.	10 30	7.1	2.5	30.0	5.5 7.1
		12.8	40.6 12.8	81.9 12.6	12.6
		D.Lat	.27.8	69.5	Dep.

Cape Henlopen's latitude Latitude in	38° 38		Meridional parts Meridional parts	
Sam of latitudes	77	6		36

39 33

By inspection of Table II. It appears that the difference of latitude 27.8 and departure 69.3 correspond to a course of 68° nearly, and a distance of 75 miles; and in the same page of the Table opposite to the meridional difference of latitude, found in the column of latitude, stands the difference of longitude 39 miles in the departure column; this subtracted from the longitude of Cape Henlopen 75° 17′ We leaves the longitude in 78° 38′ W. by Mercator's Sailing. Or, with the Middle Latitude 38° 33′ to 39° as a course, and the departure 69.3 in the latitude column, opposite to which is 89 in the distance column, which is the difference of longitude by Middle Latitude Sailing; consequently the longitude in is 78° 38′ W. as above.

Thus we see that such examples are performed as in Traverse Sailing and Case II. of Mercator's or Middle Latitude Sailing, either by Inspection, as

above, or by the scale of logarithms.

Having gone through the necessary problems in Mercator's Sailing, we shall now show how Mercator's Chart may be constructed by means of the Table of Meridional Parts.

To construct a Mercator's Chart to commence at the Equator.

Suppose it was required to construct the Chart in the Plate prefixed to this work which begins at the equator, and reaches to the parallel of 50 degrees, and contains 95 degrees of longitude west from the meridian of Greenwich?

Draw the line AD representing the equator, then take from any scale of equal parts the number of minutes contained in 95 degrees, viz. 5700, which set off from A to D; subdivide this line into 95 equal parts representing degrees of longitude. Through A and D draw the lines AB, DC perpendicular to AD, and make each of them equal to 3474 which are the meridional parts, corresponding to 50 degrees. Join BC which must be subdivided in the same manner as the line AD; and through the corresponding points of the lines AD, BC must be drawn (at the distance of 10° or 20°) the lines parallel to AB, representing meridians of the earth; these lines must be numbered 0, 10, 20, &c. beginning at the line AB which represents the meridian of Greenwich. Set off in like manner upon the meridians AB, DC, (beginning from the equator AD) the meridional parts corresponding to each degree of latitude from 0° to 50°; and through the corresponding points (at the distance of 10° or 20°) draw lines parallel to the equator AD, to represent the parallels of latitude. Then the upper part of the chart will represent the north, the lower the south, the right hand the east, and the left hand the west (which is generally supposed in charts, unless the contrary is expressly mentioned.)

If the Chart does not commence at the equator, but is to serve for a certain portion of the globe contained between two parallels of latitude on the tame side of the equator, you must draw the meridians as directed in the last example; then subtract the meridional parts of the least latitude of the chart from the meridional parts of the other latitudes, and set off these differences on the extreme meridians, draw lines through the corresponding

points, and they will be the parallels of latitude on the chart.

If the chart is to be bounded by parallels of latitude on different sides of the equator, you must draw a line representing the equator, and perpendicular to it draw the lines to represent the meridians, continuing them on both sides of the equator; then set off the parallels of latitude on both sides of the

equator, in the same manner as in the first example.

Take from the Table of latitudes and longitudes of places the latitude and longitude of each particular place contained within the bounds of the chart, and lay a rule over its latitude and another crossing that over its longitude; the point where these meet will represent the proposed place upon the chart. The most remarkable point of a sea coast being thus laid down, lines may be drawn from point to point which will form the outlines of the sea coast, islands, &c. to which may be annexed the depths of water expressed in common Arabian numbers, the time of high water on the full and change days expressed in Roman numbers: the setting of the tide expressed by an arrow; and whatever else may be thought convenient for the chart to contain.

This chart is not to be considered as a just representation of the earth's surface, for the figures of islands and countries are distorted towards the poles, as is evident from the construction; but the degrees of latitude and longitude are increased in the same proportion, so that the bearings between places will be the same on the chart as on the globe; and as the meridians are right lines, it follows, that the rhumbs, which form equal angles with the meridians, will be straight lines, which render this projection of the earth's surface much more easy and proper for the mariner's use than any other.

Having the latitude and longitude of a ship or place, to find the corresponding point on the chart.

RULE. Lay a ruler across the chart in the given parallel of latitude; take

in your compasses the nearest distance between the given longitude and the nearest meridian drawn across the chart; put one foot of the compasses in the point of intersection of the ruler and meridian, and extend the other along the edge of the ruler on the same side of the meridian as the place

lies, and that point will represent the place of the ship.

If the longitude on the chart be counted from a different meridian from that you reckon from, you must reduce the given longitude to the longitude of the chart, by adding or subtracting the difference of longitude of those meridians, and then mark off the ship's place as before directed. Or, you may draw a meridian line through the place you reckon your longitude from; then measure off the ship's longitude on the equator, and apply it to the edge of the ruler, from this meridian, and you will obtain the ship's place.

To find the bearing of any place from the ship.

RULE. Lay a ruler across the given place and the place of the ship; set one foot of the compasses in the centre of some compass near the ruler, and take the nearest distance to the edge of the ruler; slide one foot of the compasses along that edge keeping the other extended to the greatest distance from the ruler, and observe what point of the compass it comes nearest to, for that will be the bearing required.

To find the distance of any place from the ship.

RULE. Take the distance between the ship and given place in your compasses and apply it to the side of the chart or graduated meridian, setting one foot as much above one place as the other is below the other place, the number of degrees between the points of the compasses will be the distance nearly.

When the places bear north and south of each other this rule is accurate; but when they bear nearly east and west, and the distance is large, it will err considerably; but in general it is exact enough for common purposes; if greater accuracy is required, it is best to find the distance by calculation.

If any one wishes to estimate the distance accurately by the chart, he

must proceed in the following manner:

1. If the place be in the same longitude that the ship is in, then the pre-

ceding rule is accurate.

2. If the place be in the same latitude as the ship, or bear east or west, the distance cannot be obtained without calculating it by Case I. of Parallel

Sailing

3. If the place be neither in the same latitude, nor in the same longitude as the ship, the distance must be found in the following manner: Lay a ruler ever both places, and draw through one of them a parallel to the equator: take the difference of latitude between both places in your compasses from the equator; slide one foot on that parallel, keeping the other extended so that both points shall be on the same meridian, and note the point of the ruler which is touched by the other foot of the compasses, take the distance from this point to the given place through which the parallel was drawn and apply it to the equator, and you will have the sought distance.

The bearing and distance of any two places from each other may be found in the same manner as the bearing and distance of any place from the ship.

EXAMPLE.

Required the bearing and distance between the east end of Long-Island

and the north part of Bermudas?

A ruler being laid over both places as directed in the preceding rule, it will be found to lay parallel to the N.W. by N. and S. E. by S. line; and the distance between the two places being taken in the compasses, and applied to the graduated meridian, will measure about 10 degrees or 600 miles; therefore these places bear from each other N.W. by N. and S. E. by S. and their distance is 600 miles nearly.

OF THE LOG-LINE & HALF MINUTE GLASS.

VARIOUS methods have been proposed for measuring the rate at which a ship sails, but that most in use is by the Log and Half-Minute Glass.

The Log is a flat piece of thin board, of a sectoral or quadrantal form, (see Plate VI. Fig. 3) loaded on the circular side with lead sufficient to make it swim upright in the water: to this is fastened a line about 150 fathoms long, called the Log-line, which is divided into certain spaces called knots. and is wound on a reel (see Plate VI. Fig. 4) which turns very easily. The Half-Minute Glass is of the same form as an Hour Glass, (see Plate VI. Fig. 2) and contains such a quantity of sand as will run through the hole in

its neck in half a minute of time.

The making of the experiment to find the velocity of the ship is called heaving the log, which is thus performed. One man holds the reel, and another the half-minute glass; an officer of the watch throws the log over the ship's stern, on the lee side, and when he observes the stray line is run off (which is about ten fathoms, this distance being usually allowed to carry the log out of the eddy of the ship's wake) and the first mark (which is generally a red rag) is going off, he cries turn! the glass holder answers done! who watching the glass, the moment it is run out says stop! the reel being immediately stopt, the last mark run off shows the number of knots, and the distance of that mark from the reel is estimated in fathoms. Then the knots and fathoms together, show the distance the ship has run the preceding hour, if the wind has been constant. But if the gale has not been the same during the whole hour, or time between heaving the log, or if there has been more sail set or handed, a proper allowance must be made. Sometimes when the ship is before the wind, and a great sea setting after her, it will bring home the log; in such cases it is customary to allow one mile in 10, and less in proportion if the sea be not so great; allowance ought also to be made if there be a head sea.

This practice of measuring a ship's rate of sailing is founded upon the following principle: That the length of each knot is the same part of a sea mile, as half a minute is of an hour. Therefore the length of a knot ought to be the of a sea mile; but by various admeasurements it has been found that the length of a sea mile is about 6120 feet; hence the length of a sea knot should be 51 feet: each of these knots is divided into 10 fathoms of about 5 feet each. If the glass be only 28 seconds in running out, the length of the knot ought to be 47 feet and 6 tenths. These are the length generally recommended in books of navigation, but it may be observed, that in many trials it has been found, that a ship will generally over-run her reckoning with a log-line thus marked; and since it is best to err on the safe side, it has been generally recommended to shorten the above measures by 3 or 4 feet, making the length of a knot about 7½ fathoms of 6 feet each, to correspond with a

glass that runs 28 seconds.

In heaving the log you must be careful to veer out the line as fast as the log will take it; for if the log is left to turn the reel itself, the log will come home and deceive you in your reckoning. You must also be careful to measure the log-line pretty often, lest it stretch and deceive you in the distance. Like regard must be had that the half-minute glass be just 30 seconds, otherwise no accurate account of the ship's way can be kept. The glass is much influenced by the weather, running slower in damp weather than in dry. half-minute glass may be examined by a watch with a second hand, or by the following method—Fasten a plummet on a line and hang it on a nail, observing that the distance between the nail and middle of the plummet be 59} inches, then swing the plummet and notice how often it swings while the glass is running out, and that will be the number of seconds measured by the glass.

To correct the distance when the log-line and half-minute glass are faulty. If there be any error in the log-line or glass, the measured distance must be corrected in the following manner, supposing that a 30" glass requires 50 feet to a knot.

(1.) If the glass only is faulty, you must say, as the seconds run by the glass are to 30 seconds, so is the distance given by the log to the true distance.

Thus if a ship sails 8½ knots per hour, by a glass of 33 seconds, the true

number of knots per hour will be 7,1; for, 36: 30:: 8,5:7,1.

(2.) If the log-line only is faulty, you must say, as 50 feet is to the distance of a knot on the line, so is the distance run by the log to the true distance. Thus, if a ship sails 7 knots per hour, by a log-line measuring 53 feet, her true distance will be 7,4 miles per hour, because, 50:53::7:7,4.

(3.) If the log-line and glass are both faulty, you must say, as 50° multiplied by the length of the glass is to 30 multiplied by the length of the line, so is the measured to the true distance. Thus, if a ship sails 6 knots per hour with a glass of 24 seconds, and a log-line of 60 feet per knot, her true velocity will be 9 miles per hour, because $50 \times 24 : 30 \times 60 : : 6 : 9$.

DESCRIPTION AND USE

OF A

QUADRANT OF REFLECTION.

MR. JOHN HADLEY was the first who published a description of the Quadrant of Reflection, for measuring angular distances, and the instrument still bears his name, although it has been ascertained that Sir Israe Newton invented a similar one some years before, but never made it public: one of our countrymen, Mr. Thomas Godfrey, of Philadelphia, had also contrived an instrument on the same principles some time before Mr. Hadley made known his discovery.

Figure 1, Plate VII. represents a Quadrant of Reflection, the principal parts of which are, the frame ABC, the graduated arch BC, the index D. the nonius or vernier scale E, the index glass F, the horizon glasses G and

H, the dark glasses or screens I, and the sight vanes K and L.

The graduated arch BC is an octant or eighth part of a circle, but on account of the double reflection is divided into 90° numbered from 0° towards the left, and each degree is commonly divided into three equal parts of 20 minutes each. The graduation on the limb is continued a few degrees to the right of 0°; this portion is called the arch of excess, and is found very

convenient for several purposes.

The index D is a flat bar commonly made of brass, moveable round the centre of the instrument, and broader towards the axis of motion, where is fixed the index glass F; at the other end is fixed the nonius or vernier scale, used in estimating the subdivisions of the arch; at the bottom or end of the index there is a piece of brass which leads under the arch, having a spring to make the vernier lie close to the limb, and a screw to fasten it in any position. Some quadrants have a tangent screw affixed to the lower part of the index to adjust its motion. The vernier is a small narrow slip of brass or ivory fixed to that part of the index which slides over the graduated arch, and usually contains a space equal to 21 or 19 divisions of the limb, and is divided into 20 equal parts; hence the difference between a division on the limb, and a division on the dividing scale, is one twentieth of a division of the limb, or one minute; therefore, if any division on the vernier is in the same straight line with a division of the limb, then no other division on the

Instead of multiplying the length of the glass by 50, and the line by 80, you may multiply the former by 5, and the latter by 3. If any one chooses to mark the log line at less than 50 feet for a glass of 50 seconds, he must put its estimated length of the knot instead of 50, in all the above rules.

vernier can coincide with a division of the limb, the extreme divisions excepted. Some time ago it was usual to reckon the divisions on the vernier from its middle towards the right, and from the left towards the middle; but this being found inconvenient, a more commodious method has been introduced of numbering from right to left; hence the degree and minute, pointed out by the vernier, may be found thus: observe what minute on the vernier coincides with a division on the limb, then this minute, being added to the degree and parts of a degree on the limb immediately preceding the first division on the vernier, will be the degree and minute required. Thus, suppose 10' on the vernier coincided with a division on the limb, and that the division on the limb preceding the first division of the vernier, was \$0.20', the division pointed out by the vernier would be \$0.50'.

The index glass F, is a plane speculum or mirror of glass, quicksilvered and set in a brass frame; it is so placed that the face of it is perpendicular to the plane of the instrument, and is fixed to the index by the screw M; the other screw N serves to replace it in a perpendicular position, if by any accident it has been put out of order. The use of this mirror is to receive the rays from the sun, or other object observed, and reflect them towards

the horizon glasses.

The horizon glasses G and H, are two small speculums-G is called the fore horizon glass, from its being used in the common or fore observation, where the observer's face is turned towards the object; and H the back horizon glass, being used in the back observation, where the observer's back is turned towards the object; these mirrors receive the reflected rays from the index glass and reflect them to the eye of the observer. The horizon glasses are not entirely quicks ivered; the fore-horizon glass G is only silvered on the lower half, the other part being transparent, and the back part of the frame cut away, that the horizon or any other object may be seen through it; the back-horizon glass H, is silvered at both ends; in the middle is a transparent slit, through which the horizon may be seen: these two glasses are set in brass frames similar to that of the index glass, and fixed on moveable bases, which are adjusted by screws so as to set the glasses in their In general there are three dark glasses or screens I, two red ones of different shades, and one green; each is set in a brass frame, which turns on a centre that they may be used separately or together: they serve to defend the eye from the rays of the sun during an observation. The green glass is peculiarly adapted to take off the glare of the moon, but may be used for the sun when much obscured by clouds. When these glasses are used for a fore-observation, they are to be fixed as in fig. I, but when used for a back-observation they are to be placed at O.

The sight vanes, K and L, are pieces of brass standing perpendicular to the plane of the instrument; the vane K is called the fore-sight vane, and L the back-sight vane. There are two holes in the fore-sight vane, the lower of which, and the upper edge of the silvered part of the fore-horizon glass are equi-distant from the plane of the instrument, and the other hole is opposite to the middle of the transparent part of that glass. The back-sight vane has one perforation which is exactly opposite to the middle of the transpa-

rent slit in the back-horizon glass.

The adjusting-lever, (fig. 2,) which is fixed on the back of the quadrant, serves to adjust the horizon glass, by placing it parallel to the index-glass; when this lever is to be made use of, the screw B must be first loosened, and when by the adjuster A the horizon glass is sufficiently moved, the screw B must be fastened again, by which means the horizon glass will be kept from changing its position.

As the quadrant, from various accidents, is liable to be out of order, it is necessary that the mariner should be able to ascertain the errors, and readjust the several parts before he proceeds to make his observations. For this purpose he must examine whether the index glass and the horizon glasses be perpendicular to the plane of the instrument, and whether the plane of the fore-horizon glass be parallel, and that of the back-horizon glass

perpendicular to the plane of the index glass, when 0 on the vernier stands against 0 on the limb.

1st. To ascertain whether the index glass be perpendicular to the plane of the Quadrant.

Place the index on the middle of the arch, and hold the index glass near the eye, look into it, in a direction parallel to the plane of the instrument, and see if the reflected arch appear exactly in a line with the arch seen direct, or if the image of any point of the arch near B appear of the same height as the corresponding part of the arch near C seen direct, if so, the index glass is perpendicular to the plane of the quadrant; if not, the error must be rectified by the screws on the base behind the frame, by loosening the screw M, and tightening the screw N, or by loosening the screw N, and tightening the screw M.

2d. To ascertain whether the fore-horizon glass be perpendicular to the plane of the Quadrant.

Having adjusted the index glass, hold the instrument in a vertical position; look through the fore-sight vane, and move the index till the reflected and direct images of the horizon, seen in the horizon glass, coincide; then incline the instrument till its plane is nearly parallel to the horizon; if the images still coincide, the horizon glass stands perpendicular, otherwise it does not, and must be adjusted by the screws placed before and behind it, loosening one of them and tightening the other.

This adjustment may be made by the sun, moon, or star, by holding the quadrant in a vertical position, and observing if the object seen by reflection appears to the right or left of the object seen direct; and moving the screws

as above till both images coincide.

After having made the horizon and index glasses parallel, according to the directions in the following article, it will be best to re-examine this adjustment.

3d. To make the horizon glass parallel to the index glass when Q on the vernies stands on O on the arch.

Having fixed the index so that 0 on the vernier stands on 0 on the arch, look at any distant object and see if the image of it coincides with the object itself: if it does, the adjustment is complete; if not, they must be made to coincide by means of the adjusting lever. The horizon may be used for this purpose in the following manner: hold the plane of the instrument vertical, look through the lower hole in the vane K, and direct the sight through the transparent part of the Glass G to the horizon; then if the horizon line, seen in the silvered and transparent part, coincides, or makes one straight line, the horizon glass is said to be adjusted; but if the horizon lines do not coincide, slacken the screw B (fig. 2) in the middle of the adjusting lever, and turn the horizon glass on its axis until the horizon lines coincide, then fix the lever firmly by tightening the screw B. If this adjustment be again examined, it will perhaps be found imperfect; in this case, therefore, it remains either to repeat the adjustment, or find the error of it usually called the index error, which may be done thus-Let the horizon glass remain fixed, and move the index till the image and object coincide, then the difference between 0 on the vernier and 0 on the arch is the index error, which is to be added to the angle or altitude observed, if the 0 on the vernier be to the right hand of 0 on the arch, otherwise to be subtracted. Thus if the horizon was used, the instrument being held in a vertical position, you must look through the lower hole of the vane K, towards the horizon; then move the index till the reflected and direct images of the horizon coincide, the difference between 0 on the vernier and 0 on the arch will be the index error.

4th. To adjust the back-horizon glass, that it may be perpendicular to the plane of the index glass, when 0 on the vernier stands on 0 on the arch.

Set the index as far to the right of 0 on the arch as twice the dip of the horizon (taken from Table XIII.) hold the quadrant in a vertical position

look towards the horizon through the hole in the back horizon vane L, and the transparent slit of the back horizon glass H, then if the reflected horizon, which will appear inverted, coincide with that seen direct, the glass is truly adjusted, otherwise the screw, in the centre of the lever on the under side of the quadrant, must be slackened, and the glass turned on its axis till both horizons coincide, when the lever should be fixed by tightening the screw.

5th. To adjust the back horizon glass that it may be perpendicular to the plane of the quadrant.

Put the index on 0; hold the quadrant nearly parallel to the horizon; look through the hole on the back sight vane, and if the true and reflected horizons appear in the same straight line, the glass is perpendicular to the plane of the instrument; but if they do not coincide, the sunk screws before and behind the glass must be turned till both appear to form one straight line.

To take an altitude of the Sun by a Fore Observation.

If the sun is bright, turn down one or more of the dark glasses; hold the instrument in a vertical position; apply the eye to the upper hole in the fore-sight vane, when the image is so bright as to be seen in the transparent part of the fore-horizon glass, otherwise to the lower hole; direct the sight to that part of the horizon beneath the sun, and move the index till you bring the image of his lower limb to touch the horizon directly under him; but as this point cannot be exactly ascertained, the observer should move his instrument round to the right and left a little, keeping as nearly as possible the sun always in that part of the horizon glass, which is at the same distance as the eye from the plane of the quadrant,* by which motion the sun will appear to sweep the horizon, and must be made to touch it at the lowest part of the arch; the degrees and minutes pointed out by the index will be the observed altitude of the sun's lower limb at that instant.

To take an altitude of the Moon by a Fore Observation.

In the night when the moon is bright, her image may be seen in the transparent part of the fore-horizon glass, and the observation may be taken exactly in the same manner as an observation of the sun. If the image is so faint as not to be seen in the transparent part of the horizon glass, you must set the index to 0, hold the plane of the quadrant in a vertical position, direct the sight to the moon, and at the same time look for her reflected image in the silvered part of the horizon glass; move the index forward till the moon's image, which will appear to descend, just touches the horizon, then sweep the quadrant as in observing the sun, and bring her round limb into contact with the horizon, whether it be her upper or lower. The degrees and minutes pointed out by the index will be the observed altitude of that limb which was brought in contact with the horizon.

To take an altitude of a Star by a Fore Observation.

This is done exactly in the same manner as in observing the moon's altitude when her image is so faint as not to be seen in the transparent part of the horizon glass.

To take the Sun's altitude by a Back Observation.

Put the dark glasses in the hole O, and turn one or more of them down, according to the brightness of the sun; then, holding the instrument in a vertical position, look through the back sight vane towards that part of the horizon opposite the sun; move the index till the sun's image is seen in the silvered part of the glass; give the quadrant a slow vibratory motion and the sun will appear to describe an arch with its convex side upward; bring the upper limb, when in the upper part of this arch, in contact with that part of the horizon seen through the transparent slit, and the degrees and minutes

^{*} In common quadrants, if the upper hole be looked through, the sun's image must be made to appear a first the middle of the transparent part of the horizon glass, but if the lower hole be looked through, the image must be made to appear on the line joining the silvered and transparent parts of the horizon glass, as these parts of the horizon; glass are at the same distances from the plane of the instrument as the holes to the sight vanes respectively.

pointed out by the fidex, will be the altitude of the sun's lower limb. The altitude of the moon or a star, may be obtained in the same manner, only

observing to bring the round edge of the moon to the horizon.

The back observation is but little used on account of the difficulty of adjusting and observing: various remedies have been proposed for these defects, but none have yet been generally adopted. The back observation of the altitude of any object is useful only when there is not an open horizon for the fore observation; but even in that case the fore observation might often be used, if the distance of the horizon was known, as will be explained farther on.

To take the meridian altitude of any celestial object by a Fore Observation.

When the object rises and sets, it comes to the meridian above the horizon only once in 24 hours, and is then at its greatest altitude, by observing which, the latitude may be easily determined. The sun comes to the meridian exactly at noon or 12 o'clock: the moon and stars at various hours. To observe the meridian altitude, begin a few minutes before the time of passing the meridian; bring the object to sweep the horizon according to the preceding directions; this must be repeated until the object begins to descend below the edge of the sea; the degrees and minutes then shewn by the index will be the meridian altitude.

If the object does not set, it comes to the meridian below the pole, and is then at its least altitude; this altitude may be observed as above directed with this difference, that you must continue sweeping till the object begins to rise above the edge of the sea, instead of descending below it.

The meridian altitude of any object may be taken in a similar manner by

a back observation.

Strictly speaking, this method of finding the meridian altitude is absolutely accurate, only when the ship is at rest and the sun's declination constant. For if the ship is sailing towards the sun, the altitude will be increased, and it will be decreased if sailing from the sun; but the correction of altitude arising from this source is very small, and may be neglected in all nautical calculations, as will be shown hereafter.

Advice to Seamen in the choice of a Quadrant.

The joints of the frame must be close, without the least opening or looseness, and the ivory on the arch inlaid and fixed, so as not to rise in any place above the plane of the instrument; all the divisions of the arch and vernier must be exceedingly fine and straight, so that no two divisions of the vernier (except the first and last) coincide at the same time with the divisions of the All the glasses belonging to the quadrant should have their surfaces perfectly plane, and their fore and back surfaces exactly parallel; the first of these requisites in the horizon glass and index glass may be thus verified by means of two distant objects: move the index till both objects are exactly in contact, at the upper edge of the silvered part of the horizon glass, then move the quadrant in its own plane so as to make the united images move along the line, separating the silvered from the transparent part of the horizon glass, and if in this motion the images continue united, the reflecting surfaces are good planes, otherwise the planes are imperfect. The parallelism of the two surfaces of the reflecting glasses may also be examined by viewing the image of some object reflected very obliquely, for if that image appears single and well defined about the edges, it is a proof that the surfaces are parallel; on the contrary, if the edge of the reflected image appears as if it threw a faint shadow from it, or separated like two edges, it is evident that the two surfaces of the glass are inclined to each other; if the image be the sun, and viewed through a small telescope, the examination will be more perfect. To examine the dark glasses, you must bring the image of a distant object to coincide with the object seen direct; then turn the coloured glass so that the plane which was next to the index glass may now be next to the horizon glass, and if the direct and reflected images still coincide, Digitized by GOO the surfaces of the glass are parallel.

DESCRIPTION AND USE

OF A

SEXTANT OF REFLECTION.

A SEXTANT is constructed on the same principles, and may be used for measuring altitudes, in the same manner, as a quadrant. The arch of a sextant, as its name implies, contains 60°, but by reason of the double reflection, is divided into 120°. This instrument is particularly intended to measure the distance of the moon from the sun, or a fixed star, and as that distance is wanted as accurately as possible, to determine the longitude of the place of observation, the instrument is constructed with more care, and is provided with some additional appendages that are wanting in the quadrant. Fig. 3, Plate VII. represents a sextant, the frame of which is generally made of brass, or other hard metal, the handle at its back is made of wood; by this, when observing, the instrument is to be held with one hand, while the other is moving the index. The arch AA, is divided into 120°, each degree into three parts of 20 minutes each, and the vernier scale is in general so divided as to shew half minutes. In some sextants, the degree is divided into 6 equal parts of 10' each, and the vernier shows 10".

In order to observe with accuracy, and make the images come precisely in contact, a tangent screw B is fixed to the index, by which it may be moved with greater regularity than it can be by hand; but this screw does not act until the index is fixed by the screw C at the back of the sextant. Care should be taken not to force the tangent screw when it arrives at either extremity of its arch. When the index is to be moved any considerable quantity, the screw C must be loosened; but when the index is brought nearly to the division required, the back screw C should be tightened, and then the

index moved gradually by the tangent screw.

In many sextants, the lower part of the index glass, or that next the plane of the instrument, is silvered as usual, and the back surface of the upper part painted black; a screen, painted black, is fixed by its axis to the base of the index glass, and may be placed over the silvered part when the rays are strong; in which case the image is to be reflected from the outer surface of the upper part, and the error which might probably arise from the

planes of the glass not being parallel, is thereby avoided.

The coloured glasses are similar to those applied to a common quadrant, and are usually four in number, placed at D, to screen the eye from the solar rays and the glare of the moon; they may be used separately or together, as occasion requires. In addition to these, there are three similar glasses placed behind the horizon glass, to be used in finding the index error by means of the sun, and in observing the sun's altitude by an artificial horizon on land. The paler glass is sometimes used in observing altitudes at sea, to take off the strong glare of the horizon.

A sextant is generally furnished with a tube without glasses, and two tel-

^{*} There is not in general any apparatus for the back observation fixed to a sextant, but if the altitude of any celestial object be greater than 60°, the supplement of the altitude may be obtained by a back observation with a sextant with case and accuracy, and as this method may be often used with advantage when a force observation cannot besolvanized, we shall here point out the method of taking the observation, taken in this manner.—The back of the observer being turned to the sun, he must move the index till the image of the sun touches the edge of the back horizon, and then move the sextant a little to the right and left (as in a force observation) and the image will describe an arch with the convex side upward; move the index till the lower limb of the image, when in the upper part of the arch, just touches the horizon, and the observation will be complete; observing that, if the telescope be used, the image of the sun must be seen in the horizon glass at the same distance from the plane of the instrument as the eye of the observer. The altitude thus obtained will be the supplement of the altitude of the supplement. The corrections to be applied to obtain the true central altitude, will be given bereafter.

escopes, the one representing the objects erect or in their natural situation, the other inverting them, the eye glass being fixed in a moveable tube in order to adjust the telescope to a proper focus. By means of these telescopes the line of sight may be rendered parallel to the plane of the instrument, and the contact of the limbs of any two objects more accurately observed. The tube, or either telescope, is to be screwed into a brass ring, which is connected with another brass ring, by means of two screws, and by loosening one and tightening the other, the axis of the tube, or telescope, may be set parallel to the plane of the instrument. One of these rings is fixed to a brass tem that slides in a socket, and by means of the screw L at the back of the sextant, it may be raised or lowered so as to move the axis of the telescope to point to that part of the horizon glass judged the most fit for observation.

A circular head, containing a plate, in which there are three coloured glasses, and a part that is open, sometimes accompanies the sextant: this head is to be screwed on the eye end of the tube, or on that of either telescope. The edge of the plate projects a little beyond the head on one side, and is moveable by the finger, so that the open ring or any of the coloured glasses, may be brought between the eye glass of the telescope and the eye; this answers the purpose of the dark glasses placed at E, in adjusting by the

sun, or observing by an artificial horizon on land.

To these are added a small screw driver, to adjust the screws, and a mag-

nifying glass to read off the observation with greater accuracy.

The adjustments of a sextant are similar to those of a quadrant; the index and horizon glasses must be perpendicular to the plane of the instrument, and their planes parallel to each other when the index stands on 0; also the axis of the telescope must be set parallel to the plane of the instrument: each of these particulars must be examined before an observation is taken, and the adjustments, if requisite, made according to the following directions.

1st. To set the index glass perpendicular to the plane of the instrument.

Move the index forward to about 60°, and proceed exactly in the manner prescribed for the adjustment of the index glass of a quadrant, page 91.

2d. To make the horizon glass perpendicular to the plane of the Sextant.

This adjustment is made exactly in the same manner, as that of the quadrant described in page 91, except that instead of looking through the sight vane, you may use the tube or a telescope.

To make the horizon glass and index glass parallel when the index is on 0.

Having made the foregoing adjustments, set the first division on the index to 0 on the limb, fasten the index in this position, and make the coincidence of these divisions as perfect as possible, by means of the tangent screw, the eye being assisted by the magnifying glass; screw the tube, or telescope, into its support, and turn the screw L at the back of the instrument, till the line which separates the transparent and silvered parts of the horizon glass appears in the middle of the tube or telescope; having done this, hold the plane of the sextant vertically and direct the sight through the tube or telescope to the horizon; then if the reflected and true horizons do not coincide, turn the tangent screw at the back of the horizon glass till they are made to appear in the same straight line. Then will the horizon glass be adjusted.

After the screw that retains the horizon glass in its place, is fastened, it will be proper to re-examine this adjustment; if the coincidence of the horizons is not perfect, the adjustment must be repeated till it is so; but as it is difficult to obtain a perfect coincidence by this means, the horizons may be brought to coincide by turning the tangent screw of the index, and the difference between the 0 on the arch and the 0 on the vernier will be the index.

The property which is additive to all observations, if the 0 of the index stand on the

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extra arch, otherwise subtractive. The index error may also be found, very accurately, by measuring the diameter of the sun twice, with a motion of the index in contrary directions; that is, first bring the upper limb seen by reflection to coincide with the lower limb seen directly, then bring the lower limb by reflection to coincide with the upper seen directly. If both these measures are taken either to the right or left of 0 on the limb, half their sum will be the index error; additive if to the right of 0; subtractive, if to the left; but if one of the measures be taken to the right, and the other to the left of 0, half their difference will be the index error, which will be additive when the diameter measured to the right of 0 exceeds that measured to the left, otherwise subtractive. Thus if the measures were 38' to the left of 0 on the arch, and 26' to the right* on the extra arch, half the difference or 6' would be the correction, subtractive. In some sextants the horizon glass cannot be adjusted; the index error must in that case be found, and must be considered as a constant quantity to be applied to all angles measured with the same instrument.

To set the axis of the telescope parallel to the plane of the Sextant.

In measuring angular distances, the line of sight, or axis of the telescope. should be parallel to the plane of the instrument, as a deviation in that respect, in measuring large angles, would occasion a considerable error: to avoid which, a telescope is made use of, in which are placed two wires, parallel to each other, and equidistant from the centre of the telescope, by means of which, the adjustment may be made in the following manner.— Screw on the telescope and turn the tube containing the eye glass, till the wires are parallel to the plane of the instrument; then take two objects, as the sun and moon, whose angular distance must not be less than 900, because the error is more easily discovered when the distance is great; bring them exactly into contact at the wire nearest the plane of the sextant, and fix the index; then, by altering a little the position of the instrument, make the image appear on the other wire; if the contact still remains perfect, the axis of the telescope is in its right situation; but, if the limbs of the two objects appear to separate or lap over, at the wire which is farthest from the plane of the sextant, the telescope is not parallel, and it must be rectified by turning one of the two screws of the ring into which the telescope is screwed and fixed, having previously unturned the other screw: by repeating this operation a few times, the contact will be precisely the same at both wires. and the axis of the telescope will be parallel to the plane of the instru-

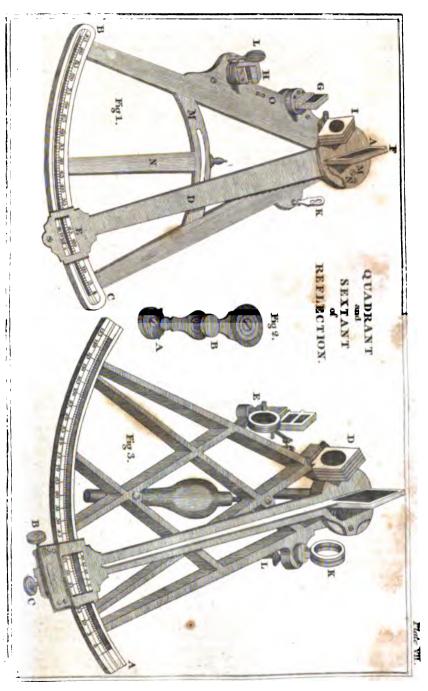
In order to estimate the error committed in not observing the contact of the objects in the middle between the two parallel wires of the telescope, it is necessary to know the angular distance of these wires: this may be found as follows: Turn round the eye piece of the telescope, till the wires are perpendicular to the plane of the instrument: hold the instrument in a vertical position, and move the index till the direct and reflected images of the horizon appear in the same line, which will happen when the index is at 0 if the instrument be well adjusted; then move the index till the reflected image of the horizon be at one wire and the directed image at the other; the angle moved through by the index, as shown by the divisions of the arch, will be the an-

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^{*} In reading off the measure on the extra arch, you must reckon the minutes on the vernier from left to right, counting 19 as 1', 18 as 2', &c. or else take the difference between the minutes denoted by the vernier and 20'. Thus if the angle on the extra arch appeared by the nonius to be 14' the real angle would be only 8'.

would be only 6°.

† This adjustment may be made in a manner similar to that by which the graduation on the frame of the telescope of a circular instrument is verified by using the adjusting tools of a circle or a ruser whose surfaces are perfectly parallel to each other. Thus, lay the sextant horizontally on a table, and place the ruler on the limb or plane of the instrument, and, at about 12 or 15 feet distance, let a well defined mark be placed in a range with the telescope, so as to be in the same straight line with the top of the ruler, then raise or lower the telescope by means of the screw L, till the entre of the eye piece of the telescope be at the same height as the top of the ruler, then if the mark be seen in the middle between the wires of the telescope, it is well adjusted; if not, it must be aftered by means of the screws of the ring thto which the telescope is svrewed.



gular distance of the two wires. This angular distance being obtained, the observer may, by means of it, estimate, at each observation, how much the place where the contact was observed, was elevated above, or depressed below, the plane passing through the eye and the middle line between the two parallel wires; the correction in Table XXXV. corresponding to this angle, is to be subtracted from the observed angular distance of the objects; thus if the distance between the wires was 5°, one of them would be elevated above the plane 1° 30′, and the other depressed as much below it; and if in taking an observation, the point of contact was estimated to be one-third part of the distance from the middle towards either wire, the angle of elevation or depression would be one-third part of 1° 30′ or 30′; and if the observed distance was 100°, the correction in Table XXXV. would be 19″, subtractive from the observed angle, which would therefore be 100°—19″=99° 59′ 41″. In general it will not be necessary to attend to this correction.

To measure the distance between the Sun and Moon.

Screw on the telescope, and place the wires parallel to the plane of the instrument; then if the index glass is half silvered and half blacked, and the sun very bright, raise the plate before the silvered part of the glass, and with the screw L raise the telescope to the transparent part of the horizon glass; turn down one or more of the dark glasses according to the brightness of the sun; then hold the sextant so that its plane may pass through the sun and moon; if the sun be to the right hand of the moon, the sextant is to be held with its face upwards; if to the left hand, the face is to be held downwards: with the instrument in this position, look directly at the moon through the telescope, and move the index forward till the sun's image is brought nearly into contact with the moon's nearest limb; then fix the index by the screw under the sextant, and make the contact perfect by means of the tangent screw; at the same time move the sextant slowly, making the axis of the telescope the centre of motion, by which means the objects will pass each other, and the contact be more accurately made, observing that the point of contact of the limbs must always be observed in the middle between the parallel wires. The observation being thus made, the index will point out the distance of the nearest limbs of the sun and moon,

To measure the distance between the Moon and a Star.

Furn down the green screen if the moon is bright, and direct the plane of the instrument through both objects, with its face upwards, if the moon is to the right of the star; but if to the left, the face is to be held downwards; look at the star through the telescope and transparent part of the horizon glass, and move the index till the moon's image appears nearly in contact with the star: fasten the index, move the sextant round the axis of the telescope as in measuring the distance of the sun and moon, and turn the tangent screw, till the coincidence of the star and the enlightened or round limb of the moon is perfect; observing that the point of contact of the limb of the moon and star must always be in the middle between the parallel wires. The observation being thus made, the index will point out the distance of the enlightened limb of the moon from the star, whether it be the farthest or nearest limb.

If the observer suspect that the mirrors, or coloured glasses, have not their surfaces exactly parallel, he may verify them as follows:—

Verification of the parallelism of the Index glass.

This verification is to be made ashore, by observing the angular distance of two well defined objects, whose distance exceeds 90° or 100° (having previously well adjusted the instrument) then taking out the central mirror and turning it, so that the edge which was formerly uppermost may now be nearest the plane of the instrument; rectify its position and again measure the distance of the two objects; half the difference between these two distances will be the error of the observed angle arising from the defect of pa-

rallelism of the central mirror. If the first distance exceeds the second, the error is subtractive, otherwise additive, the mirror being in its first position; but the contrary when in its second position. Thus, if the first distance was 119° 59' 21" and the second 120° 0' 39", the error would be 39", additive when the mirror was in its first position, subtractive for the second. The error for any other angle may be found by means of col. 2d. Table XXXIV. when the inclination of the plane of the horizon glass to the axis of the telescope is 80° , by saying, as the tabular correction corresponding to 120° (=4' 5") is to the error of the glass 59", so is the tabular error for any other angle as 85° (=1' 15") to the corresponding error of the glass 12". In this manner a table of errors may be made for all angles.*

The angle between the plane of the horizon glass and axis of the telescope produced, being constant in all observations and adjustments of the sextant,

no error can arise from the want of parallelism of its surfaces.

Verification of the parallelism of the surfaces of the coloured glasses.

Turn down the glass at D, which is to be examined, and another at E, to defend the eye from the sun; direct the telescope to the sun and move the index till its direct and reflected images coincide; then turn the dark glass at D, so that the surface which was farthest from the horizon glass may now be nearest to it, and if the contact of the same two limbs be complete, the surfaces of this glass are parallel, but if they lap over, or separate, the index must be moved to bring them again in contact, then half the arch passed over by the index will be the error, arising from the want of parallelism of the glass at D.

DESCRIPTION AND USES

OF THE

CIRCLE OF REFLECTION.

THE Circle of Reflection was invented by the celebrated professor MAYER of Groningen, and has since been greatly improved by the CHEVALIER DE BORDA, MR. TROUGHTON, and MR. MENDOZA Y RIOS. In its present improved state it has a decided superiority over the sextant in measuring the distance of the moon from the sun, or a star, on account of its correcting, in a great measure, the errors arising from a faulty division of the limb, want of parallelism in the surfaces of the mirrors and coloured glasses, and entirely avoiding the error which might arise in a sextant from the mirrors not being parallel when the index is on 0.

Fig. 1. Plate VIII. represents the Circle of Reflection, as given by DE BORDA; in fig. 2 is a section of the same instrument, marked with the same letters of reference as in fig. 1. The principal parts of this instrument are, the circular limb LMV; the central index EF; the horizon index MD; the central glass or mirror A; the horizon glass or mirror B; the telescope GH; the coloured glasses fig. 3, 4; the handle fig. 5; the ventelle fig. 6; and the

adjusting tool fig. 7.

The timb of the instrument LMV, is a complete circle of metal, and is connected with a perforated central plate by six radii; it is divided into 720° because of the double reflection; each degree is generally divided into three equal parts, and the division is carried to minutes, or lower, by means of the verniers of the two indices.

The two indices are moveable round the same axis, which passes exactly through the centre of the instrument; the central index EF carries the cen-

^{*} The method of calculating the above tabular sumbers when the angle of inclination of the telescope and horizon glass differs from 80° is given in the explanation of Table XXXIV. prefixed to the tables.

tral mirror A; and the horizon index MD carries the telescope GH and the horizon mirror B; both indices are furnished with verniers and tangent screws at O and N.

The central mirror A is placed on the central index immediately above the centre of the instrument; the plane of this mirror makes an angle of about 50° with the middle line of the index, and is adjusted perpendicular to the plane of the instrument, by means of the screws placed on the back part of the frame of the mirror.

The horizon glass B is placed on the horizon index near the limb, so as to interfere as little as possible with the rays proceeding from objects situated on the opposite side of that index with respect to the central mirror. The horizon glass is adjusted perpendicular to the plane of the instrument in a similar manner to that of the horizon glass of a sextant: and in some circles this mirror is moveable about an axis perpendicular to the plane of the instrument, by which means the situation with respect to the telescope may be adjusted.

The telescope GH, attached to the other end of the horizon index, is an astronomical one inverting the observed objects, and has two parallel wires in the common focus of the glasses, distant from each other between two and three degrees. These wires, at the time of observation, must be placed parallel to the plane of the instrument: To effect this, marks are made on the eye-piece, and on the tube at G, and by making them coincide, the wires may be brought to their proper position. The telescope may be raised or depressed by two screws I, K, so as to be directed to any part of the horizon glass; and, by means of the graduations on the swo standards i, k, the telescope may be rendered parallel to the plane of the instrument.

There are two sets of coloured glasses (fig. 3, 4) each set usually containing four glasses of different shades; the glasses of the larger set (fig. 4) which are placed before the central mirror at a, a, should have each about half the degree of shade with which the corresponding glasses (fig. 5) of the other set, placed at C, are tinged, because the rays from the luminous object pass twice through the coloured glass placed before the central mirror, and only once through the other. The glasses placed at a, a, are kept tight in their places by small pressing screws at their ends, or by slides passing in front, through perforations in the stems of their frames: when fixed for observation they make an angle of about 850 with the plane of the instrument, by which means the image from the coloured glass is not reflected to the telescope. When the angle to be measured is between 50 and 850, one of the largest set is to be fixed at a, a; in other cases, one of the smaller set is to be The reason of using the large glass is this—when placed in the socket C. the small glass is placed at C, it intercepts the direct light of the luminous object in its passage towards the central mirror, if the object happens to be situated within the angular space, included by the lines from the centre A, by the sides of the frame of the glass placed at C. This is avoided by using the larger glasses.

The kandle (fig. 5) is of wood, and is fixed to the back of the instrument immediately under the centre. By this it is held during the time of observation. The ventelle (fig. 6.) is used in terrestrial observations to diminish the light of the object seen directly, to render it equal in brightness to that of the objects seen by reflection: this is performed by putting the ventelle in the socket D,

and raising or depressing it till the objects appear of equal brightness.

There are two adjusting tools of the form represented in fig. 7; they are exactly of the same size, and their height is nearly equal to that of the central mirror; they may be used in adjusting the central mirror perpendicular to the plane of the instrument, and in making the axis of the telescope parallel to that plane.

The instrument, as we have now described it, is the same as it was left by De Borda; Mr. Troughton has since suggested the improvement of fixing to the horizon index the arch WSPR, and providing it with two sliding pieces U, X, in order to facilitate the fixing the indices at their proper

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angles with each other in taking successive observations. When the central and horizon glasses are parallel, the central index covers the space SP of the arch, and the spaces SW, PR, are each divided into degrees from S to W, and from P to R, and numbered 0 at S and P, and continued to 180° towards W and R. The use of this arch and sliding pieces will be explained

hereafter.*

That ingenious mathematician and navigator, M. Mendoza y Rios, has further improved the circular instrument by the substitution of a circular ring (moving round the centre of the instrument over or adjacent to the limb TMV) for a vernier instead of those attached to the indices by De Borda: and by fixing this circular vernier alternately to each of the indices it serves as a vernier for both, and after any number of observations, gives the compound motion of both indices, and thus double the number of distances are obtained by this instrument that can be obtained by De Borda's circle with the same number of observations. Mr. Rios has also improved the form of the handle for holding the instrument. In theory the instrument, as improved by Mr. Rios, appears to be superior to that of De Borda, but not having used one of the former kind, I cannot, from my own experience, decide whether it is so much superior in practice; but Mr. Rios says that he found it answered his expectations. As the method of taking the observation is nearly the same with both instruments, I shall confine myself to the explanation of the uses of De Borda's, from which the method of using the other will be easily discovered.

Adjustments of the Circle of Reflection.

Before entering upon an explanation of the adjustments of this instrument, it will be proper to premise that there are three different methods of observing the angular distance of two objects with this instrument, viz. (1) by what is called an observation to the right, (2) by an observation to the left,

and (3) by a cross observation.

An observation to the right is that, where the object, whose image is to be reflected and the central mirror are on the same side of the telescope. An observation to the lest, when the object to be reflected and the central mirror are on opposite sides of the telescope, which in both cases is supposed to be directed to the other object; and a cross observation is a combination of the fore-mentioned observations; the first being generally taken to the lest, and

the second to the right.

The adjustments of a circle consist in placing the mirrors perpendicular to the plane of the instrument, and in making the axis of the telescope parallel to that plane. These are all the adjustments necessary in measuring an angular distance by cross observations; but if one observation only be taken to the right, or to the left, it will be necessary to find the division, on which the horizon index must be placed, to make the horizon glass parallel to the central glass, when the central index stands on 0. These adjustments are similar to those of a sextant, but a particular explanation of each will here be given.

To set the central glass perpendicular to the plane of the instrument.

This adjustment may be made by placing the eye in front of the central glass at L, a little above the plane of the instrument, and observing if the reflected image of that part of the limb nearest the eye appears to make one continued circular line with the parts of the limb towards T, seen to the

^{*} Mr. Troughton suggested another alteration in the circle, but (as Mr. Rios justly observes) the instrument thus altered may be considered as a sextant, the limb of which is completed to the whole circumference. A circle of this description is anosally furnished with three indices and verniers, by each of which every observation must be read off. This is very troublesome, particularly in the night. It is true that this method corrects in a very great degree the error of not having the index fixed exactly on the centre, or that of not having an instrument perfectly circular; but errors of this kind in Borda's rircle may be reduced in any ratio by taking a number of observations, and the error will in general be extremely small in taking a sufficient number to bring the index nearly to the point ret out frum; so that in those important points I should, on the whole, prefer an instrument of Borda's construction.

right and left of the central glass; for in this case the glass is perpendicular to the plane of the instrument; otherwise it must be adjusted by means of

the screws till the two images coincide.*

By examining this adjustment in different parts of the limb, it will be known if the limb be in the same plane. If any difference should be found, the central glass must be so fixed that the reflected image of the limb may appear as much above the direct image in some places as below it in others.

To set the horizon glass perpendicular to the plane of the instrument.

The central glass being previously adjusted, and the telescope directed to the line separating the silvered from the transparent part of the horizon glass, hold the instrument nearly vertical, and move either index till the direct and reflected image of the horizon, seen through the telescope, coincide; then incline the instrument till it is nearly horizontal, and if the images do not separate, the horizon glass is perpendicular to the plane of the instrument; but if they do separate, the position of the glass must be rectified by means of the screws in its pedestal.

This adjustment may be also made by directing the sight through the telescope to any well defined object; then, if by moving the central index, the reflected image passes exactly over the object seen directly, the glass is perpendicular, otherwise its position must be adjusted by means of the screws

attached to the pedestal of the glass.

A planet, or star of the first magnitude, will be a good object for this pur- . pose. If the sun is used, one of the coloured glasses must be placed at C and another at D.

To make the axis of the telescope parallel to the plane of the instrument.

The telescope may be raised or depressed by means of two screws attached to the standards i, k, (fig. 2) and passing through two pieces of branconnected with the tube of the telescope. On each of these pieces is a mark or index by which the telescope is to be adjusted, for, by bringing the indices to the same mark on each standard, the telescope will be parallel to the plane of the instrument.

To find that division to which the horizon index must be placed to render the mirrors parallel when the central index is on 0.

Place the central index on 0; direct the telescope to the horizon glass, so that the line joining the silvered and transparent parts of that glass may appear in the middle of the telescope; hold the instrument vertically, and move the horizon index, till the direct and reflected horizons agree, and the division

shown by the horizon index will be that required.

This adjustment may also be made by measuring the diameter of the sun in contrary directions; thus, the central index being fixed on 0, place a dark glass at C and another at D; direct the telescope (through the transparent part of the horizon glass) to the sun, and move the horizon index, till his reflected image appear in the telescope; bring the upper edge of the direct image to coincide with the lower of the other, and note the angle shown by the index; then, by moving the horizon index, bring the lower edge of the

^{*} When the instrument is furnished with adjusting tools, this adjustment may be made in the following manner. Set the two tools on opposite parts of the limb at T and L; place the eye at r, at nearly the same height as the upper edge of the tools, so that part of the tool at T may be lid by the central glass may be used to the tool as a the same height as the other lool seen directly; then if the upper edges of the tools are apparently in the came straight line, the central glass is perpendicular to the plane of the instrument, otherwise its pusition must be adjusted by the screws at the back of the frame.

If you suspect that the marks on the standards are inaccurate, you may examine them in the following manner. Lay the circle borizontally on a table; place the two adjusting tools on opposite parts of the limb at T and L; and at about 12 or 15 feet distance let a well defined mark be placed, so so to be in the same straight line with the tops of the ions; then raise or lower the telescope till the mark is apparently in the middle between the two wires: then the axis of the telescope will be parallel to the plane of the instrument, and the difference (if any) between the divisions pointed out by the indices on the graduation of the standards i, k, (fig. 2) will be the error of the indices on the limiters of the limiters of the case in future adjustments to make allowance for the error.

direct image to coincide with the upper edge of the reflected one, and note also the angle pointed out by the index; half the sum of these two angles will be the point of the limb where the horizon index must be placed to render the mirrors parallel. Thus, if the index in the first observation stood on 473° 30′, and in the second on 474° 34′, the half sum of the two 474° 2′ would be the point where the horizon index must be placed to make the mirrors parallel.

These are all the adjustments necessary to be made preparatory to measuring any angular distance. When the angle is measured by cross observations, the error arising from the want of parallelism of the surfaces of the mirrors, and screens, will in general be very small; however, the method of verifying those glasses and making allowance for any error in them will be

given hereafter.

To observe the meridian altitude of any celestial object, either by un observation to the right or to the left.

The method of observing the meridian altitude of an object with a circle is exactly similar to that with a quadrant or sextant. The central index must be fixed on 0, and the horizon index on the point which renders the two mirrors parallel; then the altitude may be taken either by an observation to the right or to the left; but the former method, in which the large coloured glasses are not necessary, is in general to be preferred; because those large glasses are more liable to cause an error in the observation than the small ones.

If an observation to the right is to be taken, a small dark glass must be placed at C, if the object be bright, then hold the instrument in the right hand in a vertical position; move the central index, according to the order of the divisions of the limb, till the reflected image of the object, seen in the telescope, nearly touches the direct image of the horizon; tighten the index by the screw at the back of the instrument; make the contact complete in the middle between the parallel wires of the telescope, by the tangent screw, and by sweeping, exactly in the same manner, as when observing with a quadrant, and the central index will point out the altitude of the object.

If an observation to the left is taken, and the object be bright, a large dark glass must be placed at a a, if the altitude be between 50 and 350; otherwise a small glass at C; hold the instrument in the left hand, in a vertical position, move the central index contrary to the order of the divisions, and bring the reflected image in contact with the horizon as above: the angle shown by the central index being subtracted from 7200, will be the

sought altitude.

In both these methods of observing the meridian altitude of an object, the circle, the radius of which is only five inches, will hardly be so accurate as a good sextant of a larger radius; but, by the help of a well regulated watch, the meridian altitude may be obtained, by the circle, to a much greater degree of accuracy than by a sextant, by observing in the following manner. A few minutes before the object passes the meridian, begin to observe the altitude by cross observations (in the manner to be described in the next article) and note the time of each observation by the watch: continue to observe till a few minutes after the object has passed the meridian: then the angles shown by the central index being divided by the whole number of observations, will give the approximate meridian altitude; the correction to be applied to it to obtain the true meridian altitude, may be found by means

^{*} In some instruments there is an adjustment of the horizon glass, to place it at its proper angle with the axis of the telescope; if an adjustment of this kind is necessary, it ought to be made before the other adjustments, in so. h manner. that if a coloured glass be fixed at C, now of the rays from the central glass can be reflected to the telescope from the horizon glass without passing the coloured glass. To effect this, the vente'le must be placed at D, and lowered so as to intercept the direct light entirely; them place the coloured glass at O, and direct the telescope to the silveren part of the direct light entirely; them place the coloured glass at O, and direct the telescope to the silveren part of the prizon glass; move the central findex, and if no uncoloured images appear (reflected from the central glass) but all have the same tinge as that of the coloured glass used, the borizon glass is in its proper position; otherwise it must be turned on its axis till the uncoloured images disappear.

of Tables XXXII. and XXXIII. by a method which will be explained hereafter, when treating of finding the latitude by a single altitude of the sun.

In this article the meridian altitude has only been spoken of, though it is evident that the method is applicable to an object not on the meridian; but in this case the cross observations, which give to the circle all its advantages, may be used, and the mean of the altitudes taken instead of a single altitude; this method is peculiarly adapted to the taking of altitudes for regulating a watch, for which reason it will be particularly explained in the following article:-

To take altitudes of the sun or any celestial object, by cross observations, for regulating a watch.

Fix the central index on 0, and if the object be bright and the altitude between 5° and 35° place a large coloured glass before the central glass, at a a,

otherwise a small one at C; hold the instrument in the left hand, in a vertical position; move the hori- Times of obs. zon index till the image of the reflected object be brought in complete contact with the horizon, in the middle between the two parallel wires of the telescope, as directed in the preceding article, and note the time of observation by the watch; then fasten the horizon index; hold the instrument in the right hand, in a vertical position, move the central index 6) 26. 16. 40. according to the order of the divisions, till the reflected image be again brought into complete contact with the horizon* as above, and note the time of observation. Then half the sum of the times,

4b.20.' 0." 4, 21, 10, 4. 22. 15. 23. 0. 24. 45. 4. 25. 30. Angle 6)60⁵ 4. 22. 47. 10.

and half the angle shown by the index, will be a mean time, and a mean altitude corresponding thereto.

If greater accuracy be required, the observation must be repeated, setting out from the points where the indices then are, and observing in the same manner by moving first the horizon index, then the central one; continue, taking as many of these cross observations as are judged necessary, and note. the times of each observation: Then the sum of the times, divided by the whole number of observations, will be a mean time; and the angle shown by the central index, divided by the number of observations, will be a mean altitude corresponding thereto. Thus, if six observations were taken, and the times noted as in the adjoined table, the angle shown by the index being 60° 24', the mean time would be obtained by dividing the sum of the times 26h. 16' 40" by 6, and the mean altitude by dividing 60° 24' by 6; therefore the mean time would be 4h. 22' 47" and the mean altitude corresponding 100 4'

To measure the distance between the sun and moon, by a circular instrument.

The instrument being well adjusted, fix the central index on 0, and if the object be bright, place a small dark glass at C; hold the instrument so that its plane may be directed to the objects, with its face downwards when the sun is to the right of the moon: otherwise, with its face upwards; direct the sight through the telescope to the moon; move the horizon index according to the order of the divisions of the limb, till the reflected image of the sun appears in the telescope, and the nearest limbs of the sun and moon are almost in contact; fasten the index, and make the coincidence of the

The arch described on the limb by the central index will be equal to twice the altitude of the object. ^a The arch described in the limb by the central index will be equal to twice the altitude of the object, or twice the angle passed over b; the other index: if more cross observations be taken, each of the indices, when moved, whi describe an arch equal to double the altitude of the object; the same is to be observed in measuring any other angular distance. If the instrument is furnished with the arch WSR, and sliding pleers U, X, you must bring the slide X to the central index after taking the first observation to the left, and place the slide U at the same degree on the arch SW that X is on the arch PR; then, in the next observation to the left the slide X is to be brought to tuch the slide U; in the next observation to the left the slide X is to be brought to the central index, and so on for the other observations. Thus, by means of the observation, which will save considerable time. After being thus fixed, the contact must be completed by means of the taugent screw of the index, which is to be moved.

limbs perfect, in the middle between the two parallel wires of the telescope, by means of the tangent screw of the horizon glass, and note the time of observation, then invert the instrument, and move the central index, according to the order of the divisions of the limb, by a quantity equal to twice the arch passed over by the horizon index (or twice the distance of the sun and moon;*) direct the plane of the instrument to the objects; look directly at the moon, and the sun will be seen in the field of the telescope: fasten the central index, and make the contact of their nearest limbs complete, in the middle between the two parallel wires of the telescope, by means of the tangent screw of the central index, and note the time of observation; then half the arch shown by the central index will be the distance of the nearest limbs of the sun and moon, and half the sum of the times will be the mean time of observation.

Having finished these two observations, two others may be taken in the same manner, setting out from the points where the indices then are, and moving first the horizon index, then the central index: proceed thus, till as many observations as are judged necessary be taken, always observing that the number of them be even; then the angles shown by the central index (or that angle increased by 720°, or 1440°, &c. if the index has been moved once or twice, &c. round the limb) being divided by the whole number of observations, will give the mean distance, and the sum of all the times divided in

like manner will be the mean time of observation.

To measure the distance between the moon and star, by a circular instrument. Fix the central index on 0, and if the moon be bright, and the distance between 5° and 35°, place a large green glass before the central mirror at a, otherwise a small one at C; hold the instrument so that its plane may be directed to the objects, with its face downwards when the moon is to the right of the star, otherwise with its face upwards; direct the sight through the telescope to the star; move the horizon index, according to the order of the divisions of the limb, till the reflected image of the moon appears in the telescope, and the enlightened limb of the moon be nearly in contact with the star; fasten the index, and make the coincidence perfect, in the middle between the parallel wires of the telescope, by means of the tangent screw belonging to that index, and note the time of observation; then invert the instrument, and move the central index, according to the order of the divisions of the limb, by a quantity equal to twice the arch passed over by the horizon index;* direct the plane of the instrument to the objects, look directly at the star, and the moon will be seen in the field of the telescope; asten the central index, and make the contact of the enlightened limb of the moon and the star complete, in the middle between the two parallel wires of the telescope, by means of the tangent screw of that index, and note the time; then half the arch shown by the central index will be the distance of the star from the enlightened limb of the moon, and half the sum of the times will be the mean time of observation; these two observations being completed, others may be taken in the same manner, according to the directions above given for measuring the distance of the sun from the moon.

In continuing to take these cross observations by a circle furnished with the arch WSR, and slides U, X, it will be very easy to bring the reflected image into the field of the telescope; but if the instrument is not thus furnished, it will be often difficult to bring the image into the field of the telescope, and much time will be lost, and the observations rendered tedious by that means; to remedy this, a small table of the angles at which each index should be placed, ought to be made before beginning the observation: this table is easily formed, as follows: find roughly according to the directions heretofore given, the point at which the horizon glass must be placed to be parallel to the central glass, when the central index is on 0; then find what point of the arch the horizon index stands upon after measuring the first distance, as directed above; the difference between these

^{*} This may be done expeditionally by means of the slides U, X, as was explained in the preceding note.



two points will be the angular distance of the objects; the double of this distance being successively added to 0°, and to the angle i pointed out by the horizon index after the first observation, will give the points of the arch where the indices must be placed at the 2d. 3d. 4th. &c. observations. Thus, if the point of parallelism was 471° and the point where the horizon index was at the first observation, was 5250, the difference or 540 would be the angular distance; the double of which, or 108°, being added to 525° gives 635°, which is the point of the arch where that index must be placed at the third observation; 635° added to 108° gives 741° or 21° (be-

Central Index.	Horizon Index.
0 0	525
108	653
216	. 21
324	129
432	237
540	&c.
&c.	<u> </u>

cause the divisions recommence at 720°) which is the point where the index must be placed at the fifth observation, &c. as in the adjoined Table. central index being at first on 0°, after the second observation it will be on 108°, at the fourth on 108°+108°=216°, at the sixth on 216°+108°=324°, &c. Thus, by constantly adding 1080, or twice the distance of the objects, the angles, at which the indices must be placed, will be obtained; and by fixing them at these angles, the reflected image will be brought into the field of view without any trouble.*

Having explained the methods of adjusting and using the circle of reflection, it remains to show how to calculate the error arising from not observing the contact of the objects in the middle between the parallel wires of the telescope, and also to estimate the errors arising from the want of parallelism of the mirrors and coloured glasses. These verifications are much more necessary in a sextant than in a circle, and they may be in general neglected.

To estimate the error arising from not observing the contact of the objects in the middle between the parallel wires of the telescope.

To estimate that error, it is necessary to know the angular distance of the wires of the telescope, which may be thus determined.

Turn round the eye-piece of the telescope, till the wires are perpendicular to the plane of the instrument, and put the central index on 0; direct the telescope to any well defined object, at least 12 feet distant, and move the horizon index till the direct and reflected image of the object coincide; then make one of the wires coincide with the object, and turn the central index till the reflected image of the object coincides with the other wire—and the arch passed over by that index will be the angular distance between the wires. This angle being obtained, the observer must, by means of it, estimate, at each observation, how much the place where the contact was observed was elevated above, or depressed below the plane passing through the eye and the middle line between the two parallel wires of the telescope: the correction in Table XXXV. corresponding to this angle, is to be subtracted from the observed angular distance of the objects: thus if the distance between the wires was 20, one of them would be elevated above that plane 10 and the other depressed below it, by the same quantity; if, in taking an observation, the point of contact was estimated to be one-third part of the distance from the middle towards either wire, the angle of elevation or depression would be one-third part of 1° or 20'; and if the observed distance was 120° the correction in Table XXXV. would be 12" subtractive from the observed distance.

The correction for each observed distance being ascertained, in the above manner, the sum of them must be subtracted from the whole angle shown

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^{*} If the distance of the object varies during the observation, these angles will require correction as you proceed with the observations. Thus if the distance was increasing, and at the sixth observation it was found that the central index was on 326° instead of 324°, the increase being 2°, you must add 2° to the rest of the numbers in the Table, and place the horizon index at the seventh observation on 129°+2°=131° and the central index at the eighth observation at 4320+20=4340, &c.

by the central index, and the remainder, divided by the whole number of observations, will be the mean distance.

Verification of the parallelism of the surfaces of the central mirror.

This verification is to be made ashore, by observing the angular distance of two well defined objects, whose distance exceeds 90° or 100°, having previously well adjusted the instrument; after taking several cross observations and finding the mean distance, take out the central mirror and turn it so that the edge which was formerly uppermost may now be nearest the plane of the instrument; rectify its position, and take an equal number of cross observations of the angular distance of the same two objects; half the difference between the mean of these, and that of the former, will be the error of the observed angle, arising from the defect of parallelism of the central If the first mean exceeds the second, the error is subtractive, otherwise additive, the mirror being in its first position; but the contrary when in its second position. Thus, if 10 observations were taken at each operation, and in the first the angle shown by the index, was 11990 581, and in the second 1200° 64'; by dividing by 10 the mean angles are found to be 119° 59' 21" and 120° 0' 39", the difference of which is 78", the half of which or 39" is the error of the mirror additive when it is in its first position; subtractive in the second. The error for any other angle may be found by col. 4. Table XXXIV. when the inclination of the plane of the horizon glass to the axis of the telescope is 180, by saying, as the tabular error corresponding to 1200 that is 1' 30" is to the error found in the glass 39" so is the tabular error for any other at 85° which is 0'28", to the error of the glass corresponding 12"; and in this manner a table of errors may be made, not only for the cross observations, but also for the observations to the right or left.

It may be remarked that the errors are much less in the cross observations than in the observations to the right, which are those made with a quadrant or sextant, so that the circle has, in this respect, greatly the advantage of

those instruments.

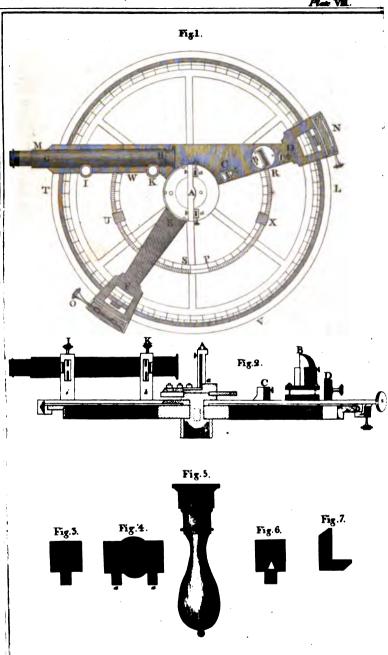
The angle between the plane of the horizon glass and axis of the telescope produced, being nearly the same in all observations and adjustments of the circle, no sensible error can arise from the want of parallelism in the surfaces of that glass.

Verification of the parallelism of the coloured glasses.

Place one of the dark coloured glasses at C and another at D, fix the central index at 0, direct the telescope to the sun, and move the horizon index till the limbs of the direct and reflected image coincide; then turn the dark glass placed at C, so that the surface which was farthest from the horizon glass may now be nearest to it, and if the contact of the same two limbs be complete, the surfaces of the glass placed at C are parallel: but if the limbs lan over or separate, the central index must be moved to bring them again in contact, then half the arch passed over by that index will be the error arising from the want of parallelism of the glass C. If great accuracy is required, the operation may be repeated, by setting out from the point where the indices then are, and taking 4 or 6, &c. observations, then the arch passed over by the central index being divided by 4, 6, &c. will be the sought error. The other small glasses may be verified in the same manner; and by placing one of the larger glasses before the central index at a a, and one of the smaller ones at D, the former may be verified as above. The green glasses may be verified by observing the diameter of the full moon, or by some bright terrestrial object.

It may be remarked as one of the greatest advantages of the circle, that in measuring an angle by the cross observations, no error can arise from the

^{*} If the inclination of the plane of the borizon glass and the axis of the telescope differ from 80°, you may find the tabular numbers by the method given in the explanation of Table XXXIV. prefixed to the Tables.



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want of parallelism in the surfaces of the smaller dark glasses; for if those glasses give too great an angle by an observation to the right, they give too little by the same quantity by an observation to the left. It is not so with the larger glasses placed at a a, because the incidence of the rays on those glasses is more oblique in one observation than in the other, so that the errors do not wholly balance each other; however, as those glasses are used only in measuring angles less than 350, in which the errors are nearly the same as if the incidence of the rays was perpendicular, the errors of those glasses will also nearly compensate each other in the cross observations; and if those observations only were used, it would be unnecessary to verify the dark glasses:-Even when taking observations to the right, or observations to the left, the error of the dark glasses would be destroyed, if the glass was turned at each observation, and the number of observations was even; but there are some cases in which an angle can only be measured by one observation, then it would be necessary to allow for the error of the dark glass, if the distance was required to be found within a few seconds.

ON PARALLAX, REFRACTION,

aka.

DIP OF THE HORIZON.

PARALLAX (or diurnal parallax) is the difference between the true altitude of the sun, moon, or star, if it were observed at the centre of the earth, and the apparent altitude observed at the same instant by a spectator at any point

on the surface of the earth.

Thus in Plate IX. fig. 8, let ABC be the earth, C its centre, A the place of a spectator, EDF part of the moon's orbit, e d G part of the orbit of a planet, and KZ part of the starry heavens. Then if at any time the moon be at D, she will be referred to the point H by a spectator supposed to be placed at the centre of the earth, and this is called the true place of the moon, but the spectator at A will refer the moon to the point b, and this is called the apparent place of the moon, the difference H b (or the angle HDb = ADC) is called the moon's parallax in altitude, which is evidently greatest when the moon is in the horizon at E, being then equal to the arch K I, and decreases from the horizon to the zenith and is there nothing. The parallax is less as the objects are farther from the earth: thus the parallax of a planet at d is represented by a b, being less than that of the moon at D; and the horizontal parallax K f of the planet is less than the horizontal parallax KI of the moon. As the parallax makes the objects appear lower than they really are, it is evident that the parallax must be added to the apparent altitude to obtain the true altitude. Having the horizontal parallax, the parallax in altitude is easily found by the following rule—As radius is to the co-sine of the apparent altitude, so is the horizontal parallax to the parallax in This rule may be easily proved: for in the triangle CAE we have CE: CA:: radius: sine CEA; and in the triangle CDA we have CD (or CE) : CA : : sine CAD : sine CDA; hence we have, radius : sine CEA : : sine CAD: sine CDA, but CEA = horizontal parallax, CDA = parallax in altitude, and sine CAD=co-sine app. alt. Hence we have radius: co-sine app. alt.:: sine hor. par.: sine par. in alt. but the parallaxes of the heavenly bodies being very small, the sines are nearly proportional to the parallaxes, hence we may say, as radius : co-sine app. alt. : : hor. par. : par. in alt.

The sum's mean parallax in altitude is given in Table XIV. for each 5° or 10° of altitude. The moon's horizontal parallax is given in the 7th page of the month of the Nautical Almanac, for every noon and midnight at the me-

dian of Greenwich.

REFRACTION OF THE HEAVENLY BODIES.

It is known by various experiments that the rays of light deviate from their rectilinear course in passing obliquely out of one medium into another of a different density, and if the density of the latter medium continually increase, the rays of light in passing through it will deviate more and more from the right lines in which they were projected towards the perpendicular to the surface of the medium: This may be illustrated by the following experiment: make a mark at the bottom of any bason or other vessel, and place yourself in such a situation that the hither edge of the bason may just hide the mark from your sight, then keep your eye steady, and let another person fill the bason gently with water: as the bason is filled, you will perceive the mark come into view, and appear to be elevated above its former In a similar manner, the light in passing from the heavenly bodies through the atmosphere of the earth deviates from its rectilinear course, by which means those objects appear higher than they really are, except when in the zenith; this apparent elevation of the heavenly bodies above their true places, is called the refraction of those bodies. To illustrate this, let ABC (Fig. 2, Plate IX.) represent the atmosphere surrounding the earth DEF, and let an observer be at D, and a star at a, then if there were no refraction, the observer would see the star according to the direction of the right line D a, but as the light is refracted, it will, when entering the atmosphere near A, be bent from its rectilinear course, and will describe a curve line from A to D, and at entering the eye of the observer at D will appear in the line D b, which is a tangent to the curve at the point D, and the arch ab will be the refraction in altitude or simply the refraction, which must be subtracted from the observed altitude to obtain the true.

At the zenith the refraction is nothing, and the lesser the altitude the more obliquely the rays will enter the atmosphere, and the greater will be the refraction: at the horizon the refraction is greatest. In consequence of the refraction, any heavenly body may be actually below the horizon, when appearing above it. Thus when the sun is at T below the horizon, a ray of light TI proceeding from T comes straight to I, and is there, on entering the atmosphere, turned out of its rectilinear course, and is so bent down towards the eye of the observer at D, that the sun appears in the direction of

the refracted ray above the horizon at S.

The mean quantity of the refraction of the heavenly bodies is given in Table XII. All observed altitudes of the sun, moon, planets, or other heavenly bodies, must be decreased by the numbers taken from that Table corresponding to the observed altitude of the object. The refraction varies with the temperature and density of the air, increasing by cold or greater density, and decreasing by heat and rarity of the atmosphere. The corrections to be applied to the numbers taken from Table XII. for different heights of Fahrenheit's Thermometer and the Barometer, are given in Table XXXVI.* Thus, if the refraction was required for the apparent altitude 5°, when the thermometer was at 20° and the barometer at 30,67 inches, we should have the mean refraction by Table XII. equal to 9′53″, and by Table XXXVI. the correction corresponding to the height of the thermometer 20° equal to +48″, and for the barometer 30,67 equal to +22″, hence the true refraction will be 9′53″+48″+22″=11′3″.

There is sometimes an irregular refraction near the horizon caused by the vapours near the surface of the earth; the only method of avoiding the error arising from this source, which is sometimes very great, is to take the observations at a time when the object which is observed is more than 10°

above the horizon.

The refraction makes any terrestrial object appear more elevated than it really is; the quantity of this elevation varies at different times from $\frac{1}{7}$ to $\frac{1}{3}$

^{*} This table is to be entered with the height of the Thermometer or Barometer at the top, and the apparent altitude at the side, under the former, and opposite the latter, will be the correction corresponding to the Thermometer or Barometer, which is to be applied to the mean refraction by addition or subtraction according to the signs at The top of the columns respectively.

of the angle formed at the centre of the earth, between the object and the observer, but in general this refraction is about $\frac{1}{14}$ of that angle.

DIP OF THE HORIZON.

Dip of the horizon is the angle of depression of the visible horizon below the true or sensible horizon (touching the earth at the observer) arising from the elevation of the eye of the observer above the level of the sea. in Plate IX. Fig. 1. let ABC represent a section of the earth, whose plane produced passes through the observer and the object, and let AE be the height of the eye of the observer above the surface of the earth, then FEG drawn parallel to the tangent to the surface at A, will represent the true horizon, and EIH, touching the earth at I, will represent the apparent horizon; therefore the angle FEH will be the dip of the horizon. Let M be an object whose altitude is to be observed by a fore observation by bringing the image in contact with the apparent horizon at H; then will the angle MEH be the observed altitude, which is greater than the angle MEF (the altitude independent of the dip) by the quantity of the angle FEH; so that in taking a fore observation the dip must be subtracted from the observed altitude to obtain the altitude corrected for the dip. In a back observation the apparent horizon is in the direction EK, and by continuing this line in the direction EL we shall have the observed altitude MEL, and it is evident that to this the dip LEF (=KEG) must be added to obtain the altitude corrected for the dip.

In Table XIII. is given the dip for every probable height of the observer expressed in feet. In calculating this table, attention was paid to the terrestrial refraction which decreases the dip a little, because IE becomes a curve line instead of a straight one, and EH is a tangent to that curve in the

point E.

What has been said concerning the dip of the horizon supposes it free from all incumbrances of land or other objects; but as it often happens when ships are sailing along shore, or are at anchor in a harbour, that an observation is wanted when the sun is over the land, and the shore nearer the ship than the visible horizon would be if it were unconfined; in this case the dip of the horizon will be different from what it otherwise would have been, and greater the nearer the ship is to that part of the shore to which the sun is brought down. For this reason Table XVI. has been inserted, which contains the dip of the sea at different heights of the eye, and at different distances of the ship from the land. This table is to be entered at the top with the height of the eye of the observer above the level of the sea in feet, and in the left hand side column with the distance of the ship from the land in sea miles and parts; under the former, and opposite the latter, stands the dip of the horizon, which is to be subtracted from the altitude observed by a fore observation instead of the numbers in Table XIII.

The distance of the land requisite in finding the dip from Table XVI. may be found nearly in the following manner—Let two observers, one placed as high on the main-mast as he can conveniently be, and the other on the deck immediately beneath him, observe at the same instant the altitude of the sun or other object that may be wanted, and let the height of the eye of the upper observer above that of the lower be measured in feet and multiplied by 0,56, and the product, divided by the difference of the observed altitudes of the sun in minutes, will be the distance in sea miles,

nearly.

Thus, if the eye of the upper observer was 68 feet higher than that of the lower, and the two observed altitudes of the sun 20° 0' and 20° 12' the distance of the land in sea miles would be 3.2. For $68 \times 0.56 = 38.08$ and this divided by the difference of the two observed altitudes of the sun 12' gives 3.2 nearly. Now if the lower observer was 25 fect above the level of the sea, the dip corresponding to this height and the distance 3.2 miles would be 6', which subtracted from 20° 0' leaves 19° 54' the altitude corrected for the dip.

The dip may be calculated in this kind of observations to a sufficient degree of accuracy without using Table XVI. in the following manner—Divide

the difference of the heights of the two observers in feet by the difference of the observed altitude in minutes, and reserve the quotient. Divide the height of the lower observer in feet by this reserved number, and to the quotient add one quarter of the reserved number, and the sum will be the dip in minutes corresponding to the lower observer. Thus in the above example $\frac{48}{3} = 5', 8$ is the reserved number, and $\frac{24}{5.8} = 4.4$, to this add one fourth of 5'.6 or 1'.4 and the sum will be the dip 5'.8 or nearly 6' corresponding to the lower observer, being the same as was found by the table.

TO FIND THE SUN'S DECLINATION.

THE declination of the sun is given to the nearest minute in Tab. IV. for every noon at Greenwich, from the year 1824 to 1838; and this table will answer for some years beyond that period, without any material error; if great accuracy is required, the declination may be taken from the second page of the month of the Nautical Almanac.* This declination may be reduced to any other meridian, by means of Table V. in the following manner.

To find the sun's declination at noon, at any place.

RULE. Take out the declination at noon at Greenwich from Table IV. (or from the Nautical Almanac;) then find the longitude from Greenwich in the top column of Table V. and the day of the month in the side column; under the former, and opposite to the latter, will be a correction in minutes and seconds, to be applied to the declination taken from Table IV: to know whether this correction be additive or subtractive, you must look at the top of the column where you found the day of the month, and you will see it noted whether to add or subtract, according as the longitude is east or west. This correction being applied, you will have the declination at noon at the given place. EXAMPLE I.

Required the declination of the sun at the end of the sea-day, October 10, 1824, in the longitude of 114° E. from Greenwich? Sun's declination Oct. 10, at Greenwich, at the end of the sea-day or beginning of the day in the N. A. by Tab. IV. 60 44' S.

Variation of Dec. Tab. V. Oct. 10, in 1140 E. long. sub. 🚅 🖎 0

True dec. noon, Oct. 10. in long. 1140 E.

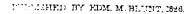
EXAMPLE II. Required the sun's declination at noon ending the sea-day of March 12, 1824, in the longitude of 75° W. from Greenwich? Sun's declination March 12, by Tab. IV. Var. Tab. V. March 12, long. 75° W.

su**b. . . .**

True declination, noon, March 12, long. 75° W.

The preceding correction ought always to be applied to the declination used in working a meridian observation to determine the latitude, though many mariners are in the habit of neglecting it.

^{*} In finding the declination, or any other quantity, in the Nautical Almanac, you must be careful to note the difference between the civil, nautical, and astronomical account of time. The civil day begins at midnight, and ends the following midnight, the interval being divided into 24 hours, and is reckoned in numeral succession from 1 to 12, then beginning again at 1 and ending at 12. The nautical or sea day begins at moon, 12 hours before the civil day, and ends the following moon; the first 12 hours are marked P. M. the latter A. M. The astronomical day begins at moon, 12 hours after the text day, and is divided into 24 hours, numbered in numeral succession from 1 to 24, beginning at noon, and ending the following noon. All the calculations of the Nautical Almanac are adapted to astronomical time; the declination marked in the Nautical Almanac, or in Table IV. is adapted to the beginning of the astronomical day, or to the end of the sea-day when mariners want the declination to determine their latitude. It would be nucle better if seamen would adopt the astronomical day, and wholly neglect the old method of country. By the sea-day.



To find the sun's declination at any time under any meridian.

RULE. Reduce the sun's declination at noon at Greenwich to noon under the given meridian, by the preceding rule. Then enter Table V. with the time from noon at the top, and the day of the month in the side column; under the former, and opposite the latter, will be the correction to be applied to that reduced declination. To know whether this correction be additive or subtractive, you must look at the top of the column where you found the day of the month, and you will find it noted whether to add or subtract, according as the time is before or after noon.

4 6				
EXAMPLE III. Required the sun's declination October 10, 1824, sea account in the forenoon, in the longitude of 114° E. from Greenwich? Sun's declination Oct. 10, at Greenwich at noon, by Tab. IV. Variation for 114° E. long sub.		60		
Declination at noon, October 10, in long. 114° E Variation of dec. for Sh. 39' from noon* Oct. 10, sub.	· ·	6	57 \$	s.
True dec. Oct. 10, sea acc. in long. 97° E. at 8h. 21' A. M.		6	54	s.
EXAMPLE IV.				
Required the sun's declination May 10, 1824, sea account, at in the longitude of 17° 50' E. from Greenwich?		50 ′		
Variation of declination, May 10, in long. 17° 50' E sub. Variation of declination for 5h. 80' P. M. additive		3'	48" 44	
Diff. is additive because the greatest number is so May 10, sea account, is May 9, by N. A. at which time sun's declination				
чествации	170	Z O	10	
'Frue declination May 10, 5h. 30' P. M. sea account in long. 17° 50' E. EXAMPLE V. Propried the sun's declination Mayeb 28, 1884, see account				
Required the sun's declination March 26, 1824, sea account, in the longitude of 120° E. from Greenwich?				
Variation of declination, March 26, in long. 120° E sub. Variation for 3h. P. M	• •	-	50' 5 6	
Diff. is subtractive because the greatest number is so	0	4 54	54 41	N.
True declination March 26, Sh. P. M. sea account	1	49	47	N.

VARIATION OF THE COMPASS.

IT was many years after the discovery of the compass, before it was suspected that the magnetic needle did not point accurately to the north pole of the world; but about the middle of the sixteenth century, observations were made in England and France, which fully proved that the needle pointed to the eastward of the true north. This difference is called the variation of the compass, and is named east when the north point of the compass (or magnetic north) is to the eastward of the true north; but vest when the north point of the compass is to the westward of the true north. The quantity of the variation may be found by observing with a compass the bearing of any celestial object when in the horizon (or, as it is called, the magnetic amplitude) the difference between this and the true amplitude found by calculation, will be the variation. The same may be obtained by observ-

a In the present example, the time is Oct. 10, 8h. 21'. A. M. which evidently wants Sh. 35" of the cyst of the sea day Oct 10, for which time the declination is marked in Table IV.

ing the magnetic azimuth of any celestial object (that is, its bearing by a compass when elevated above the horizon;) the difference between this and

the true azimuth found by calculation will be the variation.

Some years after the discovery of the variation, it was found that it did not remain constant: for the easterly variation observed in England gradually decreased till the needle pointed to the true north, and then increased to the westward, and is now above two points.

As all the courses steered by a compass must be corrected for the variation to obtain the true courses, it is of great importance to the navigator to know how to find the variation at any time; to do this it is necessary to find the magnetic amplitude or azimuth of a celestial object, which may be done as follows:

To observe an amplitude by an azimuth compass.*

When the centre of the sun is about one of his diameters; above the horizon, turn the compass round in the box, until the centre of the sun is seen through the narrow slit which is in one of the sight vanes, exactly on the thread which bisects the slit in the other ;‡ at that instant push the stop which is in the side of the box against the edge of the card, and the degree and parts of a degree which stand against the middle line on the top will be the magnetic amplitude of the sun at that time, which is generally reckoned from the east or west point of the compass.

To observe an azimuth by an azimuth compass.

Turn the compass round in the box until the centre of the sun is seen through the narrow slit which is in one of the sight vanes, exactly on the thread which bisects the slit on the other, or until the shadow of the thread falls directly along the line of the horizontal bar,‡ the card is then to be stopped, and the degree and parts of a degree which stand against the middle line of the stop, will be the magnetic azimuth of the sun at that time. which is generally reckoned from the north in north latitude, and from the south in south latitude.§ At the time of making this observation, you must also observe the altitude of the sun, in order to obtain the true azimuth.

What is here said of the sun, is alike applicable to the moon, planets, and · stars.

> TO FIND THE TRUE AMPLITUDE. RULE.

By logarithms.—To the log-secant of the latitude (rejecting 10 in the index) add the log. sine of the sun's declination; | the sum will be the log. sine of the true amplitude or distance of the sun from the east or west point, towards

the north in north declination, but towards the south in south declination.

By INSPECTION.—Find the declination at the top of Table VII. and the latitude in the side column; under the former, and opposite the latter, will be the _true amplitude. When great accuracy is required, you may proportion for the minutes of latitude and declination.

• The figure of an azimuth compass, furnished with sight vanes, is given in Plate VL fig. 5. The card

taking an amplitude.

If The declination of the sun at noon is given in the Nautical Almanac, and in Table IV. and must be corrected for the longitude of the ship and the hour of the day, by means of Table V.

[•] The figure of an azimuth compass, furnished with sight vanes, is given in Plate VL fig. 5. The card of fish so unpass is similar to that of a common compass.

† The observation is to be taken at that altitude on account of the dip, refraction and parallax, the correction of altitude depending on these causes being in general nearly equal to the sua's diameter. If the instrument is furnished with a magnifying glass fixed to one of the vanes, you may (instead of proceeding as above) turn the compass-box until the vane is directed towards the sun, and when the hright speck (or rays of the sun collected by the magnifying glass) falls upon the slit of the other vane, or upon the line in the horizontal bar, the card is to be stopped, and the divisions read off as above.

§ If the compass vibrate considerably at the time of making the observations, it would be conducive to accuracy to take several azimuths and altitudes, and work the observation with the mean azimuth and altitude. The same is to be observed in taking an amplitude.

EXAMPLE I.

Required the sun's true amplitude at rising, in the latitude of 890 0' N. on the 22d of December, 1820?

BY INSPECTION. BY LOGARITHMS. 0.10950 Under the declination 23° 29' and oppo-Latitude 390 0' log. sec. Sun's declin...23 28 log. sine 9.60012 ite the latitude 39 stands the true amplitude 300 49.

True ampli 30 49 log. sine 9.70962:

Hence the true bearing or amplitude of the sun at rising is E. 300 49' S. and at setting it is W. 300 49' S.

EXAMPLE II.

Required the moon's true amplitude at setting, in the latitude of 350 3' N. when her declination is 130 N.?

BY LOGARITHMS. BY INSPECTION. log. sec. 0.09734 Under the declination 13°, and opposite log. sine 9.35209 the latitude 35° stands 15° 56', which is Latitude35° 8' log. sec. Moon's declin. 13 0 nearly the true amplitude; the exact value 9.43943 may be found by finding the amplitude for 36° latitude, and proportioning the differ-True ampli. ... 15 59 log. sine ence for the miles in the latitude.

Hence the true amplitude at setting is W. 15° 58' North, and at rising E. 15° 58′ N.

EXAMPLE III.

Required the sun's true amplitude in the latitude of 420 50' N. when his declination was 20° N.?

BY LOGARITHMS. BY INSPECTION. Latitude 42° 30' log. sec. 0.13237 Under the declination 20° and opposite Sun's declin...20 N. log. sine 9.53405 the latitudes 42° and 43°, stand 27° 24' and 27° 53'; the mean of these gives the True ampli..... 27 38 log. sine 9.66642 true amplitude for the latitude of 42° 30'.

Hence the amplitude at setting is W. 27° 38' N. and at rising E. 27° 38' North.

To find the true azimuth at any time.

At the time of observing the magnetic azimuth, you must also observe the altitude of the object; this altitude must be corrected as usual for the dip. parallax, refraction,* &c. in order to obtain the true altitude; you must also find the declination of the object, and the latitude of the place of observation, and then the true azimuth may be calculated by the following

Add together the polar distance, the latitude, and the true altitude, take the difference between the half sum and the polar distance, and note the remainder. Then add together the log. secant of the latitude, the log. secant of the altitude (rejecting 10 in each index) the log. co-sine of the half sum, and the log. co-sine of the remainder; half the sum of these four logarithms will be the log. co-sine of half the true azimuth, which being doubled will give the true azimuth, reckoned from the north in north latitude, but from the south in south latitude.

R

^{*} In observations of the altitude of the sun's lower limb (by a fore-observation) it is usual to add 12 for the effect of dia parallax, and semi-diameter. The refraction is to be subtracted from the sum, and the remainder will be the true altitude nearly.

† The defination is to be found according to the directions in the note, in the last page.

† The polar distance of the sun, moon, or star, is the distance from the elevated pole, and is found by subtracting the declination of the object from 90°, when the latitude and declination are of the same name, but by adding to 9°° when of different names. Digitized by GOOGIC

EXAMPLE I.

In latitude 51° 32' N. the sun's true altitude was found to be 39° 28', his declination being then 16° 38' N.—required the true azimuth?

Polar distatice Latitude Altitude	51	32	secant
Sum	164	22	
Half sum			co-sine9.13355
Remainder	8	49	co-sine9.99484
			2)19.44695
Half sum log. co-sine	58 0	4′ 2	9:72347

True Azimuth116 8 from the north.

The logarithm 9.72347 of this example is also the co-sine of 121° 56', which doubled gives another azimuth 243° 52', the former being 116° 8'. One of these corresponds to an observation in the forenoon, the other to an afternoon observation.

EXAMPLE II.

In latitude 42° 16' S. the sun's true altitude was found to be 18° 40', his fectination being then 7° 58' N.—required the true azimuth?

Polar distance	6 secant0.13076
Sum:	4
Half sum 79 1 Polar distance 97 3	
Remainder 18 2	1 co-siné9.97734
***************************************	sum19.40097
Half sum log. co-sine 59 5	3 9,70049

True Azimuth119 46 from the south:

QUESTIONS TO EXERCISE THE LEARNER.

Question I. Given the sun's altitude corrected for dip, refraction, &c. 20° 46', his declination 17° 10' S. and the latitude of the place 40° 38' N. Required the true azimuth:

Answer. 137° 50' from the north.

Question II. What is the sun's azimuth in the latitude of 26° 30' N. in the forencon, when his correct central altitude is 24° 28' and his declination 22° 40' N.?

Answer. 75° 44' from the north.

Question III. At the Island of St. Helena the sun's true central altitude was found to be in the forenoon, his declination being then \$2° 58' S. Required theazimuth at that time?

Answer. 72° 21' from the south.

Question IV. What point of the compass did the star Aldebaran bear on, in the latitude of 34° 23′ S. on January 1, 1804, when the correct altitude of that star was 22° 26′?

Answer. 130° 16′ from the south.

Having the true magnetic amplitude or azimuth, to find the variation.

Having found the true and magnetic amplitude or azimuth, the variation may be easily deduced therefrom by the following rule, in which the amplitude is reckoned from the east or west point of the horizon, and is called north when to the northward of those points, but south when to the south-

Digitized by GOOSIC

ward. The azimuth is reckoned from the north in north latitudes, but from the south in south latitudes, and is named east when falling on the east side of the meridian, otherwise west. If the observed and true amplitudes be both north or both south, their difference will be the variation; but if one be north and the other south, their sum will be the variation. If the true and observed azimuths be both east or both west, their difference will be the variation, otherwise their sum; and the variation will be easterly when the point representing the true bearing is to the right hand of the point representing the magnetic bearing, but westerly when to the left hand; the observer being supposed to look directly towards the point representing the magnetic bearing.

EXAMPLE I.

Suppose the sun's magnetic amplitude at rising is E. 260 12' N. and the true amplitude E. 140 20 N. Required the variation?

From the greater E. 26° 12' N. E. 14 20 N. Take the lesser Remains variation 11 52 E.

The variation in this example is easterly, because the true amplitude falls fo the right of the magnetic.

EXAMPLE II.

The moon's true amplitude at rising was found to be E. 150 20' N. and her 800 E. and his magnetic azimuth N. magnetic amplitude E. 100 0' S. Re- 600 E. it is required to find the varia quired the variation?

True amplitude E. 15° 20' N. Magnetic amplitude E. 10 0 S. The sum is the variation ... 25 20 W.

EXAMPLE IV.

The star Aldebaran was observed at rising to bear by compass E. N. E. Jupiter was E. 100 N. when his magwhen the true amplitude was N. E. by netic amplitude was E. 200 S .- Re-E.—Required the variation?

True amp. N. Hoby E. or E. 33° 45' N. Mag. amp. E. N. E. or E. 22 30 N. 22 30 N.

Difference is the variation ... 11 15 W.

EXAMPLE III.

The sun's true azimuth being N. tion?

Magnetic azimuth N. 60 E. Diff. is the variation......20 E.

EXAMPLE V.

The true amplitude of the planet quired the variation?

True amplitude......E. 10° N. Magnetic amplitude E. 20 S. Sum is variation......30 W.

To calculate the variation by observing the sun's azimuth when at equal altitudes in the forenoon and afternoon.

The variation of the compass may also be determined by observing the magnetic azimuths of the sun in the morning and evening when at the same altitude, the observer being supposed to be at the same place at both observations; for it is evident that if the declination of the sun did not vary during the time elapsed between the observations, the middle point of the compass between the two bearings would be the bearing of the true north or south point of the horizon, at the place of observation, and the difference between that bearing and the north or south point of the compass would be the variation.

In this kind of observations it will be convenient always to estimate the magnetic azimuths from the south point of the compass, calling them east or west, as before directed, and this method is supposed to be made use of in the following rule. Then, if one azimuth be east and the other west, half their difference will be the variation, otherwise their half sum, and the variation will be of the same name as their greater azimuth, excepting, however, where the half sum is taken and exceeds 90°, in which case its supplement will be the variation of a different name from the azimuth. The variation being always supposed less than 90°.

If the declination of the sun varies during the elapsed time between the observations, (as is generally the case) an allowance may be made for that variation by applying a correction to the afternoon azimuth, calculated by the following rule.*

* The rule given in Doctor Mackey's "Complete Navigator" is insecurate.

Rule. Find from Table IV. the daily variation of the sun's declination on the day of observation. Then to the constant logarithm 9.1249 add the log-co-sine of the latitude of the place, the log-sine corresponding to the elapsed time between the observations found in the column P. M. the Prop. Log. of the daily variation of the sun's declination, and the Prop. Log. of the elapsed time*, estimating hours and minutes as minutes and seconds, the sum, rejecting 30 in the index, will be the Prop. Log. of the correction to be applied to the vestern azimuth, by subtracting when the sun is approaching towards the northern hemisphere, otherwise by adding. The azimuth thus corrected is to be used in estimating the variation instead of the observed azimuth.

It is not necessary in this calculation to find the latitude or declination to any great degree of accuracy, which is the greatest advantage of the method; another of the advantages consists in being able to take a great number of observations, and applying the correction at one operation to the variation deduced from the mean of all the observations, so that, when great accuracy is required, as in taking observations ashore, this method may be used with success; and it is evident that it is alike applicable to the moon or any heavenly body, but the observations must be taken in the same place, as it would increase the calculation considerably, to make an allowance for the change of place, as well for the change of declination; and it would be better in this case to calculate each observation separately by the rules before given.

EXAMPLE.

Suppose that on the 10th of April, 1820, in the latitude of 42° 29' N. long. 50° W. the sun's morning azimuth was observed to be S. 54° 24' E. and in the evening, when the sun was at the same altitude, was S. 39° 46' W. the elapsed time between the observations being 6h. 20m.—Required the variation?

Constant logarithm	. 9.1249
Latitude 42° 29' co-sine	. 9.8677
Elapsed time 6h. 20m. Sine	. 9.8676
Daily variation of declination 22' P. I	9128
Elapsed time 6h. 20m. taken as 6' 20" P. L	1.4536

Corrected azimuth S. 39 35 W. Morning azimuth S. 54 24 E.

Difference 14 49 The half of which 7° 24' is the variation, which is easterly, because the greater azimuth S. 54° 24' E. is easterly.

The variation, thus found, is to be allowed on all courses steered by the compass to obtain the true courses. To make this allowance, you must look towards the point of the compass the ship is sailing upon, and allow the variation from it towards the right hand, if the variation be east, but to the left hand, if the variation be west. Thus, if a ship steer S. E. with one point westerly variation, the true course will be S. E. by E. If the variation is one point easterly, the course will be S. E. by S.

In the following Table are collected a few observations of the variation, made at different times, and in different places.

of Table XXVII.

In this rule it is supposed that the bearing of the sun, by the afternoon observation, is to the westward of the meridian by compass; but if there be a great variation, that bearing might be to the eastward of the meridian by the compass, and in that case the correction of the western azimuth must be applied in a contrary manner to the above directions.

^{*} The elapsed time may be determined by any common watch, but if none was used in the observations, it may be determined as follows. If one of the observed azimuths was east and the other west, take half their sum, otherwise half their difference, and to the log, sine of this half sum (or half difference) add the log, secant of the sun's declination, and the log, co-sine of the sun's correct altitude at the time of taking the azimuth, the sum (rejecting 20 in the index) will be the log, sine to be used in the above calculation, and this logarithm will correspond to the elepsed time, marked in the column P. M. of Table XXVII.

					<u> </u>	,		
					gitude	Year of	Va	riation
Places observed at.	Latitude.			rom	Obser-	Observed		
<u> </u>				Gree	nwich.	vation.		
Cambridge, (Mass.)	420	. 23/	N.	710	8º W.	1708	90	
	ì					1742	8	0 W.
•	}		'			1757	7	20 W.
	ŀ			1		1761 1763	7	14 W. 0 W.
	l		i	i		1782	6	46 W.
Boston.	42	23	N.	71	4 W.	1742	Š	0 W.
Beverly (town.)	42	36	N.	70	52 W.	1781	7	2 W.
Salem.	42	33	N.	70	52 W.	1805	5	57 W.
1		٠.				1808	5	20 W.
London.	51	31	N.	0	5 W.	1580	11	15 E.
	l					1672 1780	2 22	30 W.
Paris.	48	50	N.	2	20 E.	1550	8	0 E.
1	•	-	• "	~	2 0 2 .	1660	ŏ	ŏ Z.
'			1			1769	20	0 W.
Funchal Road.	32	38	N.	17	5 W.	1792	18	35 W.
St. Croix Road.	28	27	N.	16	16 W.	1792	17	35 V.
Bonavista.	16		N.	22	58 W.	1792	12	36 W.
St. Jago (Praya Bay.)	14	52	N.	23	30 W.	1769	11	10 W.
Isle of May.	15	4	N.	22	46 W.	1792 1792	12 12	48 W.
Ascension.	7	56	Š.	14	21 W.	1678	1	0 E.
	1	-•	~.			1776	10	45 W.
St. Helcna.	15	55	S.	5	51 W.	1677	ő	40 E.
	1.			İ		1776	13	15 W.
				į.		1794	16	16 W.
Tristan d'Acunha.	37	7	Ş.	11	38 W.	1792	7	0 W.
Cape of Good Hope,	34	26 53	S. S.	18	23 E.	1776	21	0 W.
Cape Laguilas.	34	53	ъ.	20	10 E.	1600 1692	11	0 0 W.
	l			ł		1776	21	40 W.
	ŀ			l		1790	23	30 W.
Island St. Paul.	37	56	S.	77	28 E.	1677	23	30 W.
l				1		1803	19	30 W.
Isle of Bourbon,	20	52	S.	55	31 E.	1795	15	33 W.
Java Head.	6	46	s.	104	50 E.	1676	3	10 W.
Isle of Bourbon, Java Head. Batavia. At Sea. At Sea. At Sea. At Sea.	6	10	s.	106	51 E.	1786	0	54 W. 30 W.
At Sea.	29	10	N.	28	59 W.	1793 1795	15	00 W.
At Sca.	27	00	N.	23	43 W.	1795	15	44 W.
At Sea.	15	28	N.	20	48 W.	1795	12	5 W.
At Sea.	12	14	N.	20	5 W.	1795	11	39 W.
TAT SER	9	47		20	15 W.	1795	11	48 W.
At Sea.	8	54	N.	20	15 W.	1795	10	50 W.
At Sea.	5 3	46 16	N. N.	20	54 W. 27 W.	1795	111	00 W.
At Sea.	0	10	14.	21 24	27 W. 20 W.	1795 1795	10	47 W. 43 W.
At Sea.	2	33	S.	24	54 W.	1795	7	5 W.
At Sea.	5	48	Š.	26	54 W.	1795	5	24 W.
At Sea.	7	59	S.	27	54 W.	1795	4	14 W.
At Sea.	9	27	8.	27	55 W.	1795	3	33 W.
At Sea.	13	19	S.	26	58 W.	1795	3	54 W.
At Sea.	19	47	S.	25	56 W.	1795	2	50 W.
At Sea, near Trinidad. At Sea.	20 21	28 32	s. s.	28	44 W. 27 W.	1796	2	35 W. 25 W.
At Sea.	23	43	s. S.	25 23	45 W.	1795 1795	2	•31 W.
At Sea.	28	11	Š.	18	45 W.	1795	5	28 W.
At Sea.	35	5	Š.	6	0 W.	1795	11	10 W.
At Sea.	36	38	S.	Ŏ	15 E.	1795	13	40 W.
At Sea.	36	12	S.	4	16 E.	1795	15	19 W.
At Sea.	37	20	S.	7	25 E.	1795	16	57 W.
At Sea.	36	45	S.	19	27 E.	1795	24	233 W.
At Sea.	21	54	S.	53	41 E.	1795	13	59 W.
At Sea.	0	0		32	35 W.	1795	3	0 W

By the preceding table it appears that the variation at Cambridge, in the state of Massachusetts, is decreasing at the rate of about 1½ minutes per year, and by late observations at Salem, the needle there appears to be still approaching towards the true meridian. At London and Paris the variation formerly increased 10 or 11 minutes per year, but by some late observations made in London, it appears to be nearly stationary. Off the Cape of Good Hope the annual increase is about 7 minutes.

Besides this annual change of the variation, there is also a small diurnal change, which at London, Paris and Cambridge (Mass.) is from 10' to 15'. By this quantity the absolute variation at those places increases from about 8 A. M. to 2 P. M. when the needle becomes stationary for some time; after that, the variation decreases, and the needle comes back again to its for-

mer situation, or nearly so, in the night or by the next morning.

In addition to the observations contained in the preceding table, it may he observed, that the variation which (at present) is less than \$\frac{1}{2}\$ point W, near Cape Cod, decreases in going to the westward along the coast of the United States of America, so that near Cape Hatteras it is scarcely sensible, and farther to the westward becomes easterly. In the leeward West India islands it is about \$\frac{1}{2}\$ point E. and in the windward islands \$\frac{1}{2}\$ point E. Along the northern shore of the Brazils there is a small easterly variation which decreases in proceeding to the eastward towards Cape Roque, where it is scarcely sensible; in proceeding farther to the southward along the coast of America, the easterly variation increases so as to be above \$2\$ points E. near Cape Horn, and from thence gradually decreases along the coast of Chili and Peru, so as to be about \$\frac{1}{2}\$ point E. under the equator near Quito; but in proceeding to the northward toward the N. W. coast of America, the easterly variation increases to more than \$2\$ points.

On the contrary in proceeding to the eastward of the United States of America, the westerly variation increases, being nearly 1 point W. a little to the eastward of Cape Sable (Nova Scotia) and about 2 points W. on the E. part of Newfoundland, and at the Western Islands. At the Orkney islands it is 24 points westerly and is nearly the same in the English Channel, and on the coasts of England, Scotland and Ireland. On the coast of Holland, it is from 2 to 21 points W. In the Cattegat and Sound about 15 points W. In the Western part of the Baltic about 14 points. At the entrance of the Gulf of Finland 1 point W. In the Bay of Biscay about 24 points W. Near Cape St. Vincents 2 points W. In the Mediterranean from 14 to 14 points W. Near Cape Verd (Africa) 14 points W. and from thence gradually increases along the western shore of Africa towards the Cape of Good Hope, and is there above 2 points W. and from thence increases towards Cape Lagullas and a little to the eastward to 24 points or 24 points W. and then decreases in proceeding along the eastern shore of Africa, and is about 1 point westerly at the entrance of the Red Sea. In the Arabian Sea, Bay of Bengal, Java Sea, China Sea, and off the coast of Sumatra, it is very small, and on the S. E. part of New Holland is about 4 point E.

Ships sailing for India generally cross the equator between the longitudes of 18° and 24° west. The variation at the latter place was about 8° 48′ W. in the year 1795, as I have found by repeated observations, and the annual increase is about 6 minutes. If, in crossing the equator, you should find a greater variation, you would probably be to the eastward of 24° W. but if less, to the westward of 24° W. The alteration in the longitude is in that place about 2 degrees for 1 degree of variation. But there is always a great uncertainty attending this kind of observations, made with a common compass, since it is not uncommon to find 2 or 3 degrees difference between an azimuth in the morning and evening, when the ship during that time has been nearly stationary; the same difference will sometimes be found merely from making the observation when the ship is on a different tack. This is owing to the iron in the ship which attracts the compass by a force which is generally situated in a point near the centre of the ship. When this point and

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the compass are in the magnetic meridian of the compass, the true variation is obtained, but as soon as the position of the ship is changed so as to bring this point to the eastward or westward of the magnetic meridian passing through the compass, a corresponding change or observation in the variation to the eastward or westward is immediately perceived. This deviation sometimes amounted to 80 or 90 in the surveys of New Holland. This has since been confirmed by various observations in different places, particularly in the voyages towards the north pole lately made by order of the English government. method which was at first used to correct this error, which is sometimes of considerable importance in nautical surveys, where great accuracy is required, was to place the compass always in the same part of the ship, and to find by actual observation, the greatest deviation arising from this local attraction, which is when the ship's head is directed east or west. The deviation when the ship's head is in any other direction, is found by entering Table I. or Table II. in the page corresponding to that direction as a course, and with that greatest error in minutes in the distance column, the corresponding number in the departure column will be the required correction nearly. Thus if the deviation was 20 8' or 128' when the ship's head was directed towards the east, the deviation when in the direction of one point from the meridian, (that is N. by E. N. by W. S. by E. or S. by W.) would be found by entering Table I. in the page for one point, or with the distance 128', the corresponding departure 25' would be the correction to be applied on all bearings taken by the compass when in that situation. Mr. Barlow has invented a method of correcting this error, making use of a curious property of the attractive force of iron on the compass, it having been found that this force depends on the attractive surface, and not wholly on the quantity of iron; so that a solid globe of iron 30 inches in diameter, would affect the compass exactly in the same manner as a hollow shell of iron of the same diameter made of sheet iron only one tenth of an inch in thickness, though this shell could not contain but one hundredth part the quantity of iron which the globe does. Mr. Barlow therefore proposed to have a sheet of iron placed abaft the compass of such dimensions, and at such a distance, as should be found by experiment to bring the needle back to the magnetic meridian when the ship's head was east or west, then keeping the iron in that position it could correct the error of the local attraction of the ship in every direction of the ship's head. This method has been tested by experiment and found to succeed admirably. It has also been attended with the great advantage of leaving the compass free to act by the natural magnetism of the earth in high latitudes, where the force is much enfeebled by the obliquity of its direction on account of the greatness of the dip. In the voyages above named it was found that the compasses thus furnished traversed freely and accurately, when those of the common form moved very irregular, and were in some cases almost useless.

On the Dip of the Magnetic Needle.

If the needle of a compass be exactly balanced on its point in a horizontal position, and then the magnetic virtue be communicated, the needle will point towards the north, and will also be inclined to the horizon, the north point of the needle tending downwards, and the south point upwards in northern climates, and the contrary in southern climates. This inclination of the needle to the horizon is called *The Dip of the Magnetic Needle*, which is different in different places, though it has been found to remain nearly the same in the same place, since its discovery in the year 1576, in which year at London the dip was 71° 50', in 1723 it was 74° or 75°, and at present is about 72½°. Messrs. Humbolt and Biot published a method by which the dip may be calculated for any given place in north latitudes to a considerable degree of accuracy. This method is explained in the 22d vol. of Tilloch's Magazine, and is in substance as follows:

According to their theory there are two magnetic poles, one in the latitude of 79° 1' N. and in the longitude of 27° 42' W.* from Greenwich, the other diametrically opposite, in the latitude of 79° 1' S. and in the longitude of

[&]quot; Capt. Parry, in his voyage to the north, found the northern pole to be nearly in 70° N. and 30° W.

1520 18' E. The great circle of the earth 900 distant from these poles is called the magnetic equator. On the magnetic equator the dip is nothing, and at the poles is 90°, at any other point on the surface of the earth the dip varies with the distance from the magnetic pole: This distance may be calculated by common spherical trigonometry, or, (which is much more simple and sufficiently accurate for this purpose) by measuring the distance on a terrestrial globe from the magnetic pole to the place for which the dip is to he calculated; then to the log. co-tangent of this distance add the constant logarithm 0.30103, the sum will be the log. tangent of the dip. The dip was calculated on these principles for twenty-eight places in Europe, Asia, Africa and America, and in ten places the theory did not differ 10 from actual observations, and in five places did not differ 20, but at Spitzbergen the difference was between 40 and 50. Considering the difficulty of observing the dip with accuracy, the difference between the theory and observation may be considered as nearly within the limits of the errors of observation, and this difference may be rendered less by introducing a small correction depending on the longitude of the place of observation referred to the magnetic equator.

The methods proposed for finding the longitude by the variation and by

the dip, will be hereafter explained.



TO FIND THE LATITUDE BY OBSERVATION.

THE latitude of a place being its distance from the equator, is measured by an arch of the meridian contained between the zenith and the equator; hence, if the distance of any heavenly body from the zenith when on the meridian, and the declination of the object be given, the latitude may be thence found.

The meridian zenith distance of any object may be found by observing its altitude when on the meridian, or by observing one altitude taken at a given hour from passing the meridian, or by two altitudes taken out of the meridian and the elapsed time between the observations—each of these methods

will be explained by proper examples.

Altitudes of the sun and moon taken at sea require four corrections in order to obtain the true altitude of their centres; these are for Semi-diameter, Dip, Refraction, and Parallax.* When a planet or star is observed, the corrections for dip and refraction only are to be applied, as the semi-diameter and parallax of a planet are but a few seconds, and may be neglected in finding the latitude at sea.

In a fore-observation with a quadrant, sextant or circle, the semi-diameter is to be added if the lower limb was observed, but subtracted if the upper limb was observed. The dip and refraction are to be subtracted and the parallax to be added, and the central altitude will be thus obtained, which

being subtracted from 90° will give the true zenith distance.

In a back-observation with a quadrant, the semi-diameter is to be subtracted if the lower limb was observed, but added if the upper limb was observed. The dip and parallax are to be added, and the refraction subtracted, and the central altitude will be obtained, which being subtracted from 90° will give the true zenith distance.

In a back-observation with a sextant or circle, by measuring the supplement of the altitude by bringing the lower limb of the image of the object to touch the back horizon, the semi-diameter and refraction must be added to

^{*} The semi-diameter of the sun may be found in the 3d page of the month of the Nantical Almanac and is nearly 16. The sun's parallax is to be found in Table XIV. The refraction in Table XII. The dip in Table-XIII. The semi-diameter and parallax of the moon may be found from the Nautical Almanac, as will be explained hereafter. It may also be observed, that it is usual to add 12 for the correction for semi-diameter, dip and parallax, in a fore-observation of the sun's lower limb, taken on the deck of a common sized vessel, and by subtracting the refraction from the sun, the true shitude will be obtained nearly, and it ought always to be kept in mind that t'e refraction at low altitudes is of too much importunce to be neglected.

the altitude given by the instrument, and the dip and parallax subtracted therefrom, and by subtracting 90° from the remainder, the true zenith distance will be obtained.

To find the Latitude by the meridian altitude of any object.

Having obtained the true meridian zenith distance by either of these methods, you must then find the declination of the object at the time of obser-This may be found for the sun by the Nautical Almanac or by means of Tables IV. and V. in the manner before explained. The declination of a fixed star may be easily found by inspection in Table VIII. The declination of the moon or a planet may be found in the Nautical Almanac in a manner which will be hereafter explained. Having the meridian zenith distance and declination, the latitude is to be found by the following rules.

CASE I.

When the object rises and sets.

RULE. If the object bear south, when upon the meridian, call the zenith distance north; but if the bearing be north you must call the zenith distance Place the zenith distance under the declination, and if they are of the same name add them together, but if they are of different names, take their difference; this sum or difference will be the latitude which will be of the same name as the greatest number.

CASE II.

When the object does not set, but comes to the meridian above the horizon twice in 24 hours.

Many stars are always above the horizon of certain places of the earth, and in high latitudes the sun is sometimes above the horizon for several days, in which case the meridian altitude may be observed twice in 24 hours; that is, once at the greatest beight above the pole, and again at the lowest height upon the meridian below the pole. In the former case, the latitude is to be found by the preceding rule, but in the latter, by the following

Add the complement of the declination to the meridian altitude; the sum will

be the latitude, of the same name as the declination.

NOTE. When the sun or star is on the equator, or has no declination, the zenith dis-tance will be equal to the latitude of the place, which will be of the same name as the zenith distance.

When the sun or star is in the zenith, the declination will be equal to the latitude, and it will be of the same name as the declination.

To find the latitude by the meridian altitude of the sun or star.

EXAMPLE I. EXAMPLE II. Suppose that at the end of the sea day, Suppose that at the end of the sea day, June 21, 1824, in the longitude of 60° W. April 14, 1824, in the longitude of 140° E. Suppose that at the end of the sea day, the meridian altitude of the sun's lower from Greenwich, the altitude of the sun's limb bearing south, was found by a fore-lower limb by a fore-observation was 60° observation to be 40° 6'—required the la-25' when on the meridian and bearing south, titude, supposing the correction of the ob-the correction for dip, semi-diameter and served altitude for parallax, dip, and semi-parallax being 12 miles-required the latidiameter to be 12 miles? tude? Par. dip, and semi-dinm. add 12 Correction.....add 12 Sum40 18 True altitude † 60 37 Refraction subtract 1 Subtract from90 0 True altitude......40 17 23 N. True zenith distance29 Sun's declination, Tab. IV. cor. by Tab. V. for long. Subtract from90 True zenith distance49 Sun's declination, Tab. IV.....23 28 N. Latitude 44 Ń.

In this rule the sun is supposed to be the fixed point, and the zenith is referred to it. Thus, if the a.m bears south from an observer (or from his zenith) the zenith bears north from the sun, and it is this latter bearing which is used in the rule.

The gefraction being small is here neglected.

122 TO FIND THE	LATITUD
EXAMPLE . III. Suppose that at the end of	_
Suppose that at the end of May 15, 1824, in the meridia	the sea day,
wich, the meridian altitude of th	e sun's lower lo
limb bearing north was found l	
servation to be 30° 6', the coparallax, dip and semi-diam. being	ng 12 miles—re
required the latitude? Observed altitude	01
Par. dip, and sen:i-diamadd	30° 6′ Se
=	la:
Sum	
	Re
True altitude	00 00
	h
True zenith distance	59 44 S.
Latitude	40 49 S.
EXAMPLE IV.	fo
Suppose that at the end of	the sea day, is
Nov. 17, 1824, in the longitud	te of our E. su

Suppose that at the end of the Nov. 17, 1824, in the longitude from Greenwich, by a fore-obsermeridian altitude of the sun's lwas 50° 6′, bearing south; the observer being 17 feet above the the sea—required the latitude? Observed altitude	of treations of the contract o	on to the second	É. the ml
Dun's schil-dam and		10	
•	50	22	
Subtract dip and refraction			
True altitude*	50	17	
Subtract from			
PR 143 31 1	**	4 19	TA T
True zenith distance Sun's dec. cor. by Tab. V	. 19	43	S
Latitude	. 20	42	N

By a fore-observation, the meridian altitude of the sun's lower limb was found to be 40° 20' bearing south of the observer, the declination being 9° 56' N. and the eye 26 feet above the horizon-required the lati-

EXAMPLE V.

tude of the place? Observed altitude40° 20'

*	40	36	
Dip 5', refraction 1',sub.		6	
True altitude sun's centre* Subtract from		30	

DUDITUCE IT OFFICE TO THE PARTY OF THE PARTY	 	
Zenith distance Declination		

I The parallex being small is neglected.

Latitude.....

EXAMPLE VI.

By a back-observation with a quadrant of flection, the meridian altitude of the sun's wer limb was 25° 12' when the declination as $21^{\circ}14'$ S. and the eye of the observer 40 et above the horizon, the sun bearing S .-quired the lat. of the place of observation? 250 12/ beerved altitude 16

56 2 efraction sub. rue alt. of sun's centre*.....25 rue zenith distance65 eclination21 14 S.

EXAMPLE VII.

Suppose that on January 1, 1824, an oberver 17 feet above the water finds by a re-observation that the altitude of Sirius 53° 33' when passing the meridian to the uthward; required the latitude of the

place of observation? Observed altitude	.53°	33)	,
Dip of horizonsub.		4	_
1 ·	53	29	•
Refraction sub.		1	
	53	28	•
True zenith distance	. 36	32	N.
Sirius' declin. Tab. VIII.	. 16	29	S.
Latitude	. 20	3	Ŋ.

EXAMPLE VIII.

Suppose that on the 13th June, 1824, sex account, an observer in a high northern latitude and in the longitude of 65° W. from Greenwich, his eye being 20 feet above the surface of the water, observed by a fore-observation the altitude of the sun's lower limb on the meridian below the pole 8° 14': required the latitude?

The sun being below the pole at 12 hours. before the end of the sea day June 13, the correction of declination corresponding in Table V. is—1' 46", and the correction for 65° W. long. is +0' 38", hence both corrections make nearly 1' to be subtracted from the declination at noon 23° 14' N. which will give the declination at the time of observation 23° 13' N. the comp. of which is 66° 47'. 8º 14' Observed alt. sun's lower limb Semi-diameteradd 16

Dipsub.	4	
Refractionsub.	26 6	
True alt. of sun's centre 8 Complement of declination66	20	*

^{...... 59 26} N. Latitude..... • The parallex being small is here neglected, and the cun's semi-diameter is supposed to be 16'.
† The north polar distances of these bright stars are given for every 10 days in the Nautical Almança; the great accuracy is required, the declinations deduced from these may be used instead of the numbers in Table VIII.

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EXAMPLE IX.

Suppose that on January 10, 1924, an Suppose that by a back-observation will observer 18 feet above the water, finds the a sextant the lower limb of the sun's image the place of observation? Subtract dip. 4'. ref. 1'.....

True altitude		N
L'atitude	56	N
EXAMPLE X.		

Suppose that by a back-observation with a sextant the lower limb of the sun's image was brought to the back horizon, and the angle shown by the index was 110° 10', the sun being then on the meridian and bearing south, the declination being 200 5' N. the sun's semi-diameter 16' and the observer 20 feet above the horizon; required the latitude? Observed angle110° 10'

Semi-diameter.....add

Dipsub.	110	26 4	
	110	22	
Subtract	. 90	0	
Zenith distance† Declination	. 20 . 20	22 5	N. N.

EXAMPLE XI.

altitude of the north star, when on the me-was brought to the back horizon, and the ridian below the pole, to be 36° 23' by a angle shown by the index was 106° 12', the fore-observation; required the latitude of altitude of the observer being 22 feet and the correction for semi-diameter, parallax; and dip being (as usual) about 12; required the true latitude, supposing the de-clination to be 20° S, and that the sup boxes

Observed angle	106°	12 13	
Subtract	106 90		•
Zenith distance	16 20	24 0	s .
Latitude	36	24	s.

It was observed in the directions for finding the meridian altitude of an object, that an error would arise if the ship were in motion, or the sun's declination should vary. The amount of this correction may be estimated in the following manner.

Find the number of miles and tenths of a mile northing or southing made by the ship in one hour, and also the variation of the sun's declination in an hour expressed also in miles and tenths. Add these together, if they both conspire to elevate or depress the sun, otherwise take their difference, which call the arch A. Find in Table XXXII. the arch B, expressed in seconds, corresponding to the latitude and declination; then the arch A, divided by twice the arch B, will express the time in minutes from noon when the greatest (or least) altitude was observed. Moreover, the square of the arch A, divided by five times the arch B, will be the number of seconds to be applied to the observed altitude to obtain the true altitude which would have been observed if the ship had been at rest.

Thus if the ship sailed towards the sun south 12 miles per hour, the declination increasing northerly 1' per hour, we should have A=11+1=12. If the latitude was 42° N. declination 2° S. we should have by Tab. XXXII. B=2". In this case the time from noon is $\frac{1}{4}=3$ minutes, and the correction of altitude $^{1}\frac{4}{3}$ = 18 seconds only.

^{*} The north polar distances of these height stars are given for every 10 days in the Neutical Almanac; when great accuracy is required, the declinations deduced from these may be used instead of the numbers in Table VIII.

f The refraction and parallax being only a few seconds are neglected.

The refraction being small is neglected.

TO FIND THE LATITUDE

BY THE

MERIDIAN ALTITUDE OF THE MOON.

THE latitude may be found at sea by the moon's meridian altitude more accurately than by any other method, except by the meridian altitude of the sun; but to do this it is necessary to find the time of her passing the meri-dian, and her declination at that time. To facilitate these calculations we have given the Tables XXVIII. XXIX. and XXX. The uses of which will evidently appear from the following rules and examples.

To find the true time of the moon's passing the meridian.

In the sixth page of the Nautical Almanac, find the time of the moon's roming to the meridian of Greenwich for one day earlier than the sea account;* and also the time of her coming to the meridian of Greenwich the next day, when you are in west longitude, but the preceding day when in east longitude; take the difference between these times, with which you must enter the top column of Table XXVIII. and against the ship's longitude in the side column will be a number of minutes to be applied to the time taken from the Nautical Almanac, for the day immediately preceding the sea account, by adding when in west longitude, but subtracting when in east longitude; the sum or difference will be the true time of passing the meridian of the given place.

EXAMPLE.

Required the time of the moon's passing the meridian of Philadelphia;

April 19, 1820, sea account?

The day preceding the sea account is April 18, on which day the moon passed the meridian of Greenwich at 5h. Sm.; and being in west longitude, I find also the time of her passing the meridian the next day, 5h. 55m.; the difference of these two numbers is 52m.; with this I enter Table XXVIII. and at the top find 52; under this and opposite 75° (the longitude of Philadelphia) is the correction 11', to be added to 5h. 3m.; therefore the time of passing the meridian of Philadelphia is April 19d. 5h. 14m. sea account; or April 18d. 5h. 14m. p. m. civil account.

To find the moon's declination when on the meridian.

Find the time of the moon's coming to the meridian as above; turn the ship's longitude into time, (by Table XXI.†) and add it thereto if in west longitude, but subtract it in east, the sum or difference will be the time at Greenwich.—Take out the moon's declination from page 6th of the Nautical Almanac, for the nearest noon and midnight;‡ and note the difference of the declinations if of the same name, but their sum if of different names; enter Table XXX. with this sum or difference at the top, and the time at Greenwich in the side column, under the former and opposite the latter will be the correction to be applied to the declination which stands first in the

* Taking the time one day earlier than the sea account reduces it to astronomical time used in the Nautical Almanac.

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I Longitude may be turned into time without the help of Table XXI. by multiplying by 4 sexagesimally, and putting the product one denomination lower; and by dividing by 4, time may be turned into degrees, &c. Thus 80°×4=20'=5b. 20m. and 15° 16'×4=61'4'=1b. 1m. 4s.; in like manner 1b. 20m. or 50m. divided by 4, gives 20°, 3h. 16m. or 196m. divided by 4, gives 49°, which agree with the Table. If the ship be farnished with a chronometer regulated to Greenwich or mean time, this part of the operation will be saved, for by applying the equation of time, Table IV. A., with a contrary sign to that in the Table, the appearant time at Greenwich will be obtained, as in the explanation prefixed to the Tables. If the time at Greenwich be exactly noon or midnight, the true declination will be given by the Nautical Almanac, without the triubble of referring to Table XXX.

Nautical Almanac; additive, if that declination be increasing; subtractive, if decreasing; the sum or difference will be the true declination at the time of passing the meridian.

NOTES.

1. By the above rule, the day of the month on which the moon passes the meridian must be taken one less than the sea account: and when you add the longitude (turned into time) to the time of passing the meridian, and the hours of the sum exceed 24, you must subtract 24h. and add one to the day of the month; if the longitude be subtractive and greater than the time of passing the meridian, you must, previous to the subtrac-tion, add 24 hours to the time of passing the meridian, and subtract one from the day of the month; the sum or difference will be the time at Greenwich. If this time be less than 12 hours, you must take out the declination for the preceding noon and the following midnight; but if the time exceed 12 hours, you must take out the declination for the preceding midnight and the following noon.

When one of the declinations taken from the Nautical Almanac is north and the other south, the difference between the correction of Table XXX. and that declination which stands first in the Nautical Almanac, will be the true declination, which will be of the same name as that first declination, when the correction of Table XXX. is less than the first declination, but if greater of a contrary name.

· 3. In the same manner we may find the declination for any time in the day, by making use of the given time instead of the time of the moon's passing the meridian.

4. In the above rules the second differences of the moon's motion are neglected. In cases where very great accuracy is required, the calculation may be made as in Problem I. of the Appendix.

EXAMPLE.

Required the moon's declination at the time of her passing the meridian

of Philadelphia, April 19, 1820, sea account?

The time of passing the meridian of Philadelphia was found in the preceding Example to be April 19d. 5h. 14m. sea account, or April 18th. 5h. 14m. by Nautical Almanac account; this being added to the longitude of Philadelphia, in time 5h. 1m. nearly, the sum is the time at Greenwich, April 18th. 10h. 15m. The declination April 18th. at noon, was 28° 26' N. and on April 18th. at midnight 27° 48' N. the difference being 38', this being found at the top of Table XXX. and the time 10h. 15m. in the side column, the number corresponding is 35', which subtracted from the first declination 28° 26' leaves the declination required 27° 53' N.

At the time of the moon's passing the meridian you must observe the altitude of her upper or lower limb, and correct this altitude for semi-diameter, dip, parallax, and refraction, and you will obtain the central altitude, with which and the declination you may find the latitude by the rules before given. Or you may correct the observed altitude by the following approximate method which shortens the calculation, and is sufficiently accurate, especially when the dip is about 4' or 5', which is nearly the value in common observations at sea.

To find the latitude by the moon's meridian altitude, obtained by a fore-observation.

To the observed altitude of the moon's lower limb add 12 minutes, but if her upper limb was observed, subtract 20 minutes; with this altitude enter Table XXIX. and take out the minutes corresponding and add thereto, the sum will be the central altitude of the moon;* with this altitude and the moon's declination found as above, the latitude may be found as by a meridian altitude of the sun.

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^{*} In calculating accurately the moon's central altitude, you must proceed in the following manner: Find the time of the moon's passing the meridian reduced to Greenwich time as above, take out the moon's horizonial parallax and semi-diameter for this time, from the seventh page of the month of the Mautical Almanae, increase the semi-diameter by the correction in Table XV. add this augmented semi-diameter to, or subtract it from the observed altitude according as the lower or upper limb was observed (by a fore-observation) subtract the dip of the horizon taken from Table XIII. and add the correction for parallax and refraction (which may be easily found by Table XIX. by subtracting the correction found in that table from 89 42") and the sum will be the correct central affitude.

EXAMPLE I.

Suppose that on the 27th of June, 1820, sea account, in long. 80° W. from Greenwich, the meridian altitude of the moon's upper limb was observed to

he 40° 0' bearing south; required the true latitude?

June 27th, sea account, is June 26th, by Nautical Almanac, on which day the moon passed the meridian of Greenwich at 12h. 45m. and the next day at 13h. 46m. the daily difference being 61m. In Table XXVIII. under 60 (which is the nearest number to 61 in the table) and opposite to the long. 80°, stand 13m. which added to 12h. 45m. gives the time of passing the meridian, June 26d. 12h. 58m.

D. H. M.			
passes merid June 26 12 58			
Ship's long. 80° in time 5 20	Alt. D's upper limb40° Subtract	28)	
Time at Greenwich June26 18 18			
	39	40	
June 26, midnight decl 27° 24' S.	Add Tab. XXIX	43	
June 27, at noon			
·	D's true ait49	23	
Difference 1 14			
With this difference 1° 14', and the time) 's zen. dist49	37	Ň.
at Greenwich 18h. 18m. I enter Table XXX.	D's declination	45	Ş.
and find the correction 0° 39'			•
This subtracted from27 24 S.	Latitude	52	Ņ.
Gives the required decl26 45 S.			
All the pare reduction appril 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	•		

EXAMPLE II.

Suppose that on the 27th September, 1820, sea account, in long. 20°.E. the meridian altitude of the moon's lower limb was observed 50° 0′, bearing

south, required the true latitude?

September 27, sea account, is September 26, by the Nautical Almanac, on which day the moon passed the meridian of Greenwich at 16h. 17m. and the preceding day at 15h. 19m. differing 58m. in Table XXVIII. under 52 and opposite the long. 90° are 14m. which subtracted from 16h. 17m. leaves 16h. 5m. the time of passing the meridian of the place of observation.

D. H. M. Department of the passes merid. Sept)'s ohs. alt. lower limb56° Add	12	,
Time of Greenwich Sept 20 10 3	Sum50	<u> </u>	
Decl. Sept. 26 at noon 25° 11' N. at midnight 26 41 N.	Corr. Tab. XXIX. add		
Difference 1 30)'s correct alt 50	48	
With this difference 1° 30' and the time	D's zenith distance39	12	N.
at Greenwich 10h. 3m. I find the correction of Table XXX 1° 15'	D's declination26	26	N.
Declination at noon25 11 N.	Latitude65	3 8	Ŋ.
True declination 96 96 N	•		

EXAMPLE III.

Suppose that on the 29th November, 1820, sea account, in the longitude of 150° W. the meridian altitude of the moon's upper limb was observed 80° 26', bearing north, required the true latitude?

Nov. 29, sea account, is Nov. 28, by the Nautical Almanac, on which day the moon passed the meridian of Greenwich at 19h. 20m. and the next day at 19h. 59m. differing 59m. In Table XXVIII. under 59' or 40' and opposite the longitude of 150° stand 17m. which added to 19h. 20m. gives 19h. 57m. the time of passing the meridian of the place of observation.

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) passes merid. Nov	Obs. alt.)'s upper limb60° 26' subtract 20
Time at Greenwich Nov29 5 37	60 6 Corr. Tab. XXIX. add 28
J's decl. Nov. 29 at noon 0° 27' N. at midnight 2° 22 S.	D's corr. alt60 84
Sum2 49 With this sum 2° 49', and the time at	J's zen. dist
Greenwich 5h. 37m. I enter Table XXX. and find the corr. of decl 1° 19' Decl. Nov. 29 at noon 0 27 N.	Latitude
True declination 0 52 9.	In this example you must refer to notes. 1. and 2. page 125.

TO FIND THE LATITUDE

BY THE

MERIDIAN ALTITUDE OF A PLANET.

FROM page 4th of the month of the Nautical Almanac, take out the time of the planet's passing the meridian on the day nearest to that on which the observation was made; this will be nearly the time of passing the meri-

dian of the place of observation.

Turn the ship's longitude into time, and add it to the time of passing the meridian, when in west longitude, but subtract in east, the sum or difference will be the time at Greenwich nearly. Take out the planet's declination, from the Nautical Almanac, for the times immediately preceding and following the day of observation, and note the difference of the declinations when they are of the same name, but their sum when of different names, and find the interval between these times marked in the Nautical Almanac; take also the difference between the time first marked in the Nautical Almanac and the time of observation at Greenwich (remarking that this time is one day less than the sea account;) then as the former interval of time is to the latter, so is the sum, or difference of declinations, to the correction of the declination taken first from the Nautical Almanac, additive if that declination be increasing, but subtractive if decreasing; the sum or difference will be the declination of the planet at the time of observation. But you must observe that if the correction of declination be greater than the declination first marked in the Nautical Almanac, their difference will be the sought declination, which will be of a different name from the first declination.

From the observed altitude of the planet (taken by a fore observation) subtract the refraction and dip, the latter being in general about four minutes, and the remainder subtracted from 90° will give the correct zenith distance nearly; with which, and the declination, the latitude may be found as by an

observation of the sun.

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^{*} If you wish to find the time of passing the meridian more occurately, you must take a proportional part of the difference of the timer of coming to the meridian given in the Nautical Almanac, in the same manner as in finding the declination of the planet.

† This time is also given by a thronometer, as in note page 124, or in the explanation medical to the table.

EXAMPLE.

Suppose that on the 23d October, 1820,	, in long. 65° W. Jupiter passed the
meridian to the southward; his meridian a	altitude being observed was 45° 20',
and the dip 4'; required the true latitude?	

October 23, sea account, is October 22, by the Nautical Almanac;	now	on Oct	ober 19,
by the N. A. Jupiter passes the meridian at			\ '
To this add the long. 65° W. in time	4	2 U	•

Time at Greenwich, October 22d	13 46
Jupiter's declination, October 19	
October 25	7 46 S.

Then say, as 6 days (which is the interval between October 19 and October 25) is to 3 days 13\$ hours (which is the time elapsed between October 19th and October 22d. 13% h.) so is 7 minutes to 4 minutes, which added to 7° 39' S. gives 7° 43' S. the true declination at the time of observation.

Jupiter's observed altitude			
True altitude			
Zenith distance	44	45 43	N. S.
Latitude	37	2	N,

TO FIND THE LATITUDE BY DOUBLE ALTITUDES.

FORM I .- By double altitudes of the Sun.

WHEN by reason of clouds, or from other causes, a meridian altitude cannot be obtained, the latitude may be found by two altitudes of the sun, taken at any time of the day, the interval or elapsed time between the observations being measured by a good watch or chronometer, noticing the seconds, if possible, or estimating the times to a third or a quarter of a minute, if the watch is not furnished with a second-hand. The observed altitudes of the sun must be corrected, as usual, for the semi-diameter, dip, refraction and parallax, in the same manner as in finding the latitude by a meridian altitude. When great accuracy is required, the declination must be found at the time of each observation, using the third method of solution hereafter given, but when the sun's declination varies slowly, or the elapsed time is small, it will in general be sufficiently accurate to find the sun's declination for the *middle* time between two observations, and to consider it as invariable during the observations, computing the latitude by the first or second method.

This manner of finding the latitude is in general most to be depended upon where the sun's meridian zenith distance is great. If the sun passes the meridian near to the zenith, much greater care must be taken in measuring the altitudes and noting the times, than would be necessary under other circumstances. The nearer the sun is to the meridian at the time of one of the observations, the more correct the result will commonly be. In general the elapsed time ought to be as great or greater than the time of the nearest observation from noon. Similar remarks may be made upon every one of

the following forms.

In all these observations it is supposed that the watch moves uniformly according to apparent time, measuring twenty-four hours from the time of the sun's passing the meridian on two successive days at the same place of observation. If the watch gain or lose on apparent time, supposing the observer

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to be at rest, a correction must be applied for the gain or loss during the time elapsed between the observations, so as to obtain accurately the elapsed time or hour angle. It is not required that the watch should be regulated so as to give precisely the hour of observation; the only thing required is to find the elapsed time with all possible accuracy.

FORM II .- Double Altitudes of a Star.

Double altitudes of a fixed star may be used in finding the latitude, and the calculation is almost identical with that of double altitudes of the sun; the only difference consists in adding a small correction to the elapsed apparent time. between the observations, on account of the daily acceleration of 8' 56" in the time a star comes to the meridian on successive days. This correction is obtained to a sufficient degree of accuracy by adding one second for every six minutes of the elapsed time; the sums will be the corrected clapsed time or hour angle, to be used in the calculation, either by the first, second or third method. Thus if the elapsed time was Sh. or 180m. the correction would be 140 or 30", making the corrected elapsed time or hour angle 3h. 0' 30". If great accuracy is required, find the correction in Table XXXI. in the column marked at top 5' 58", and at the side with the elapsed time. In the precedings example, this Table would give 29" for the correction, instead of 50".

In observations of a fixed star the altitudes are to be corrected for dip and refraction, as in finding the latitude by a meridian altitude. The declination of the star is to be found in Table VIII.*. With these altitudes, the declination and the hour angle, the calculation is to be made by either of the three

methods hereafter given.

The chief difficulty in observations of this kind with a fixed star is the want of a good horizon in the night time. The method, however, might sometimes be used with success, soon after the dawn of day, or late in the evening twilight, at a time when the horizon is well defined, and the star sufficiently bright to bring its reflected image to the horizon. Sometimes a good horizon is produced by the aurora borealis, in which case a good observation might be made with stars in the northern horizon, but a single observation of the polar star will answer the same purpose, and be much more simple.

FORM III.—Double Altitudes of a Planet.

Double altitudes of a planet (particularly Jupiter and Venus, on account of their great brightness) might sometimes be used with success. The observed altitudes must be corrected for dip and refraction. The parallax and semi-diameter being small may be neglected, except in cases where extreme accuracy is required. The declination of the planet is to be found in page IV. of the Nautical Almanac, for the supposed time at Greenwich. The daily variation of the time of coming to the meridian is also to be found in the same page where the hour is marked at intervals of 6 days, and thus the time elapsed between the passage of the planet over the meridian on two successive days is found; then the corrected elapsed time or hour angle is obtained by the following

RULE. As the interval of time between two successive passages of the object over the meridian is to 24 hours, so is the apparent elapsed time between the observations, to the corrected elapsed time or hour angle: Or more simply by means of Table XXXI. finding the daily variation in the time of coming to the meridian at the top; and the elapsed time at the side, the corresponding correction is to be added to the elapsed time when the time of coming to the meridian is earlier on successive days, as is generally the case, but subtracted if later, the sum or difference will be the corrected elapsed time or

hour angle nearly.

With this hour angle, the declination and corrected altitudes, the latitude may be found by either of the three following methods of calculation.

whose place is given in that work.

† If the daily variation be less than 3 1-2 minutes, which is the smallest in the table, you may multiply the daily variation by 2 or 3, &c. and divide the result by the since number, and the correction will be obtained.

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^{*} Or more accurately in the Nautical Almanac, if any one of the twenty-four bright stars is observed,

FORM IV .- Double Altitudes of the Moon.

Double altitudes of the moon may also be used in finding the lati-These observations may be easily and very accurately made, but the calculation is much more complex than any of the preceding methods. on account of the great change in the moon's declination and right ascension. during the elapsed time between the observations. If, however, by the times of observation, and the longitude of the ship; or else by a chronometer, the time at Greenwich can be obtained within a few minutes; we may, from pages VI. VII. of the Nautical Almanac, find the corresponding declination, semidiameter and horizontal parallax of the moon for each of these observations. With the horizontal parallax and the moon's apparent altitude, find the correction in Table XIX. which being subtracted from 59' 42", leaves the correction of the moon's altitude for parallax and refraction,* which is to be added to the corresponding observed altitude corrected for semi-diameter and dip, and in this way the moon's correct central altitude is to be obtained at each observation. Lastly, the time of the moon's passing the meridian on successive days in page VI. of the Nautical Almanac, gives the interval of time between two successive passages of the moon over the meridian,† and this time is to 24 hours as the elapsed time between the observations is to the corrected elapsed time or hour angle. With this hour angle, the correct central altitudes and the declinations, the latitude may be found by the third of the following methods of calculation, it being very rare that the two first methods can be used, on account of the great change in the moon's declination.

FORM V.—By altitudes of two different objects, taken at the same time.

The latitude may be obtained by observing, at the same moment of time, the altitudes of two heavenly bodies: as for example, (1) The sun and moon. (2) The moon and a fixed star or planet. (3) A planet and a fixed star-(4) Two planets. (5) Two fixed stars. In these methods the altitudes are to be corrected as in the preceding Forms, for dip and refraction; also for parallax and semi-diameter when necessary, as is always the case in observations of the moon and sun. The declinations of the bodies are to be found for the supposed time of observation, reduced to the meridian of Greenwich. by means of the Nautical Almanac, or by Table VIII. for the fixed stars, as before taught. Then the difference of the right ascensions of the bodies (or that difference subtracted from 24 hours, if it exceed 12 hours) will be the hour angle which is to be used with these declinations and corrected altitudes in finding the latitude, by either of the two first methods if the declinations should be equal, or differ but one or two minutes, otherwise by the third method, which in fact may be considered as the only method to be used in this kind of observations, because, in almost all cases, the declinations of the objects differ considerably.

FORM VI.—By altitudes of two different objects, taken within a few minutes of each other, by one observer.

It may sometimes happen, for want of two good instruments, or from not having two observers, that the preceding Form V. cannot be employed. In this case the whole of the observations may be made by one person, noticing the interval between the observations, and making the calculation as in the following Form VII. But it is in general much better to make the observations as near to each other as possible, and then by a very simple process the calculation may be reduced to that of Form V. in which the observations are taken at the same moment. This is done by observing the first object twice, before and after observing the second object. For if the interval of time between these three observations are equal, as, for example, one minute, or

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^{*} When extreme accuracy is not required, we may find the correction for parallax and refraction from Table XXIX. which, if the altitudes are large, will not vary much from the truth.
† This time is given to minutes which in general is sufficient, because if the elapsed time is small, the effect of this correction would be only a few seconds. It might be obtained more accurately by means of the right accussions of the sun and moon, using the second differences, as taught in the Appendix.
† A particular case of this method occurs in taking a lunar observation, which will be treated of separately, because the distance of the two bodies being known, the calculation becomes more simple.

Two minutes, the half sum of the two altitudes of the first object may be taken for the altitude corresponding to the time of observing the second altitude, and the calculation may then be made as in Form V. Thus, suppose at 10h. 2m. A. M. per watch, the altitude of Sirius was 17° 54′, at 10h. 4m. per watch the altitude of Capella 60° 45′, and at 10h. 6m. per watch the altitude of Sirius was again observed and found to be 17° 58′. In this case the intervals of time are exactly two minutes, therefore the half sum of the altitudes of Sirius is to be taken 17° 56′, and combined with the altitude of Capella 60° 45′, supposing both to have been observed at 10h. 4m. per watch. This is the most simple form in which an observation of this kind can be made by one observer.

If, from any cause whatever, the observations cannot be taken at exactly equal intervals, the altitude of the first object, at the time of observing the second object, may be found by proportion, supposing the altitudes to vary uniformly during the few minutes of the observations. Thus in the preceding example, suppose the altitudes and the two first noted times to remain unaltered, but the last observation of Sirius to have been at 10h. 10m. per watch, (instead of 10h. 6m.) In this case, during the 8 minutes of time elapsed between 10h. 2m. and 10h. 10m. Sirius would have risen 4' (from 17° 54' to 17° 58'), therefore, by proportion, it is found that in 2 minutes (the time elapsed between 10h. 2m. and 10h. 4m.) the star would have risen 1', and the altitude would have increased from 17° 54' to 17° 55'; therefore at the time 10h. 4m. per watch, the altitude of Sirius must be taken at 17° 55', the altitude of Capella 60° 45', and with these quantities considered as observed at this last mentioned time 10h. 4m. the calculation must be made as in Form V.

There are several advantages attending these two last Forms V. VI. since no allowance is necessary for the change of place of the ship; the observations can be immediately made, in a short interval of fair weather, when the common method of double altitudes might fail from the intervention of clouds; the time can also be obtained at the same operation, &c.

FORM VII.—By altitudes of two different objects, taken at different times.

This method differs but very little from the two last; the altitudes are to be corrected in the same manner for dip and refraction, also for parallax and semi-diameter when necessary. The right ascension and declination of each object is to be found for the supposed time of observing that object reduced to the meridian of Greenwich. Then the apparent elapsed time between the observations, is to be turned into sidereal time, which may be done with sufficient accuracy as in the Form II. by adding one second for every six minutes of the elapsed time, the time thus corrected is to be added to the right ascension of the body first observed: the difference between this sum and the right ascension of the body last observed is the hour angle.* This, with the declinations and corrected altitudes, are to be used in finding the latitude by the third of the following methods of calculation, it being very rarely the case that the first or second methods can be used on account of the difference of the declinations. These three last forms, when a fixed star or planet. is used, are restricted very much from the want of a good horizon in the night; they are best adapted to the morning and evening twilight.

GENERAL REMARKS.

Having thus explained several of the different forms of making these observations, and the manner of finding in each form the hour angle, the declinations and the correct central altitudes, we shall now give three different methods of calculating the latitude, and shall illustrate the rules by proper examples. In the first and second methods the declination is supposed to be the same at both observations, which is true as it respects observations of a fixed star, and is in general sufficiently correct for common observations of double altitudes of the sun. The first of these methods is direct and simple, not embarrassed with much variety of cases, requiring only ten openings of the Table XXVII. without any halving or doubling of the logarithms, or the use

[.] If this difference exceed 12 hours, subtract it from 24 hours, and use the remainder as in form V.

of natural or versed sines. This method is in fact nearly, if not fully, as short as the second or approximative method invented by Mr. Douwes, and which was exclusively used in the former editions of this work. This second, or Douwes' method, is liable to the objection, that the calculation must sometimes be repeated several times before a true solution can be obtained, and then it becomes extremely troublesome. This difficulty does not occur in the first method, and on this account, as well as for its remarkable simplicity,

It is always to be preferred.

The third method embraces the general solution of the problem in the case where the variation of declination is noticed. This increases the labour considerably, and renders the solution more complex in its cases. It is, however, believed, that this method (drawn up in its present form by the author of this work) will be easily understood by navigators, and that they will thus be enabled to determine the latitude with considerable accuracy in cases where it might be of the utmost importance to know it, and where other methods could not be resorted to on account of bad weather. This method is nearly, if not quite, as short as that published by Dr. Brinkley in the Nautical Almanac of 1825, and does not require, like his method, a second or third, or even a greater number of operations.

If the observer should change his place or station during the elapsed time between the observations, a correction must be applied to one of the altitudes on this account. The manner of doing this is shown in the following exam-

ples V. and VI.

It may be observed, that in like manner as there are two latitudes corresponding to the same meridian altitude of the sun, according as the zenith is north or south of the sun when on the meridian; so in double altitudes there are generally two latitudes, corresponding to the proposed altitudes, according as the zenith and north pole are on the same side or on different sides of the arc or great circle passing through the two observed bodies, or through the two places of the same body; and it therefore becomes necessary to notice (at the time of observation) how the zenith and north pole are situated with respect to this great circle.

To estimate the effect of small errors in the observations.

When running in with the land, or crossing a dangerous parallel with no other means of obtaining the latitude than by double altitudes, it becomes a matter of great importance to ascertain the possible error of the latitude thus computed, arising from supposed errors in the observed altitudes, or in the elapsed time. The differential expressions in spherical trigonometry afford methods of doing this, but they are not adapted to the nature of this work, on account of the complication and variety of cases. The following method, though long, is general and infallible, and was once used by the

writer in a case of great anxiety and danger.

Rule. After having computed the latitude by either of the three following methods, using the observed altitudes* and elapsed time, repeat the operation, varying the altitude you suspect may be erroneous by 2' or 3', (or whatever you suppose the limit of the error in that altitude may be) the difference between this second latitude and that first computed, is the effect of the supposed error in that altitude. If you suspect the second altitude also to be erroneous, the operation may be again repeated, varying this second altitude 2' or 3' (or whatever the limit may be supposed) but using the first observed altitude and elapsed time, comparing this third computed latitude with the first, the difference is the effect of this supposed error in the second altitude. Finally, if the elapsed time is supposed to be erroneous, the operation may be again repeated, using the observed altitudes and varying the elapsed time by 20 or 30 seconds (or whatever the limit of this error may be supposed) the difference between this fourth latitude and that first computed, is the effect of this supposed error of the elapsed time.

Thus, suppose the first computed latitude was 30°, the second 50° 1', the

^{*} Mouning the observed altitudes, corrected as usual for dip, refraction, parallax and semi-diameter, if necessity:

third 30° 5', the fourth 30° 2'. The error arising from the first altitude would be 1', that from the second altitude 3' and that from the elapsed time 2'. If all these errors existed at the same time, the greatest limit of the error would be the sum of these quantities or 6', so that the true latitude would be 300 ± 6' or between 290 54' and 300 6'. In this way the limit of the error may be obtained in any case, and the degree of confidence that may be placed in the observation obtained. This examination is sometimes very necessary, because the objects may be so situated, that a small error in the observations might produce a considerable change in the computed latitude. It may be observed that the error of one observation is frequently corrected in whole or in part, by the error of the other; the one tending to increase the latitude, the other to decrease it.

To find the Latitude by Double Altitudes of the Sun (or any other object) the declination being invariable.

FIRST METHOD.

In this method the log-sines, co-sines, &c. of Table XXVII. are used, and for brevity the word log. is omitted in the rule. For the convenience of writing down at once in the same line all the logarithms which occur at the same opening of the book, they are arranged in three columns, as in the following formula, and it will be very convenient to have one of these blanks prepared at the commencement of the operation, and then the logarithms may be written down in their proper places with great rapidity.

F	ORMULA.		• '
Col. 1.	Col. 2.		Col. 3.
Elapsod time [rn] co-sec. Declinationsec.	.,	•••••	· co-sec.
Aco-sec.	co-sine	 · · · · · · · ·	. co-sine
å sum altsco-sine	co-sec.	В	co-sec.
4 diff. altssine	sec.		[B less than 90°, like decl. N. or S.]
Csine	co-sine		.co-sine
Z [Less than 90°, north or south like bearing of Zenith.]		z ·	
[E is Sum of B, Z, of same name, difference if of different name.]		E	sine
	Latitude.		sine

Rule. (Using Table XXVII.)

- Find the elapsed time* in column P. M. take out the corresponding co-secant and put it in Col. 1.
 - 2. Put the secant of the declination in Col. 1, its co-secant in Col. 3.
- 3. The sum of the logarithms in Col. 1, (rejecting 10 in the index) is the co-secant of the angle A, whose co-sine is to be put in Col. 2 and Col. 3.
- 4. The sum of the logarithms in Col. 3. (rejecting 10 in the index) is the . co-secant of the angle B (less than 900) which is to be named north or south, like the declination.
- 5. Find half the sum of the two altitudes; place its co-sine in Col. 1, its co-secant in Col. 2. Find also half the difference of the two altitudes; place its sine in Col. 1, its secant in Col. 2.
- 6. The sum of the three lower logarithms of Col. 1. (rejecting 20 in the index) is the sine of the angle C, whose co-sine is to be placed in Col. 2, and Col. 3,
- 7. The sum of the logarithms in Col. 2. (rejecting 30 in the index) is the secant of the zenith angle Z, which is to be taken out (less than 90°) and placed under B in Col. 3, naming it north if the zenith and north pole be

each example from 17 to 15.

If any other object than the sun is observed, the corrected enapsed time or hour angle, found as before taught, is to be used.

† The co-sines of A and O are each written down tastes, which reduces the number of logarithms in

situated on the same side of the arch or great circle passing through the two observed places (or objects), but south if the zenith and north pole be situated on different sides of that great circle.*

The angle E is found by taking the sum of the angles B, Z, if they are of the same name, or their difference if of different names, marking E north.

or south, like the greatest of the two angles B or Z.†

9. Put the sine of E in Col. 3. and the sum of the two last written logarithms of Col. 3. (rejecting 10 in the index) is the sine of the latitude, of the same name as E.

If the time of observation were required, it might be found by the following

rule (still using Table XXVII.)

RULE .- Add the tangent of C to the secant of E, the sum (rejecting 10 in the index) is the tangent of an angle. Take out half the corresponding time in Col. P. M. (or in Col. A. M. increased by 12 hours) and this will represent the horary distance of the object from the meridian, (upper or lower) at the middle time between the two observations. Take the sum and difference between this and half the elapsed time, or hour angle, and they will be the hours and minutes distance from the meridian corresponding to both observations, expressed in apparent solar time if the sun be observed, sidereal time if a star is observed, &c.

EXAMPLE I.

Being at sea in latitude 46° 50' N. by account, when the sun's declination was 11° 17' N. at 10h. 2m. per watch, in the forenoon, the sun's correct central altitude was 46° 55', and at 11h. 27m. per watch, in the forenoon, the correct central altitude was 540 9'. Required the true latitude?

Subtracting 10h. 2m. from 11h. 27m. gives the elapsed time 1h. 25m.

Col. 1.		Co	L. 2.	Col. 3.
Elap. time [P.M.] 1h.25m. co-sec. Declination 11° 17′ N sec.	10.73430 10.00849		•••••	co-sec. 10.70850
A	10.74278	co-sine	9.99278	co-sine 9.99278
sum alts. 50 39co-sine	9.80320	co-sec.	10.11239	B 11° 28' N. co-sec. 10.70128
4 diff. alts. 3 37sine	8.79970	sec.	10.00087	[B less than 90°, named N. or S. like decl.]
Csine	9.34568	co-sine	9.98907	co-sine 9.98907
Z [Less than 90° and N. or S. like Zenith.]		sec.	10.09511	Z 36 33 N.
[E is sum of B, Z, if of same name, if of different name.]	difference			E 48 01 N. sine 9.97119

Latitude 46 27 N. sine... 9.86026

If the sun had passed the meridian to the north of the observer, Z would have been 36° 33' S. and E. 25 $^{\circ}$ 5' S. whose sine 9.62730 added to cos. C 9.98907 gives the sine of the latitude 9.61637, corresponding to 24° 25' S.

In the first case (in north latitude) the tangent of C 9.35640 added to the secant E 10.17468 gives 9.58103, which, in the tangents, corresponds to 2h. 30m. nearly, whose half 1h. 15m. is the time of the middle observation from noon; adding and subtracting half the elapsed time 42m. 30s. gives the times from noon 1h. 57m. 30s. and 0h. 32m. 30s. of the observations, a small difference would be found if the calculation had been made to seconds instead of the nearest minute.

EXAMPLE II.

At sea in the latitude of 47° 19' N. by account, when the sun's declination was 12° 16' N. at 10h. 24m. A. M. per watch, the sun's correct central alti-

Rifferent and difference.

tude was 49° 9', at 1h. 14m. P. M. per watch, his correct central altitude was 51° 59'. Required the latitude?

Subtracting 10h. 24m. from 1h. 14m. increased by 12h. leaves the elapsed time 2h. 50m.

Col. 1.		Co	L. 2.		Cor. 3.	
Elap. time [P.M.] 2h.50m. co-sec.	10.44077	ï	***			
Declination 12º 16' N sec.	10.01003	•••••			· · · co-sec.	0.67272
A	10.45080	co-sine	9.97097	······	co-sine	9.97087
4 sum. alts. 50 34co-sine	9.80290	co-sec.	10.11218	B 15°08	N. co-sec.	10.64359
diff. alts. 1 25sine	8.39310	sec.	10.00013	Ì	[B less than N. or S. l	90°, named ike deci.]
Csine	8.64680	co-sine	9.99958		co-sine	9.99959
Z [Less than 90° and N. or S. like b Zenith.] [E is sees of B, Z, if of same name,			10.08276	Z 34 16	N.	
if of different name.]	aijjerence			E 47 24	N. sine	9.86694

If the sun had passed the meridian to the north of the observer, Z would have been \$4° 16' S. E. 21° 08' S. its sine 9.55695 added to cos. C 9.99958

gives 9.55653, the sine of the latitude 210 7'S.

If the observed object in this example had been a fixed star, with the same declination 12° 16' N. the same altitudes 49° 9', 51° 59', but the elapsed time 2h. 49m. 52s., the calculation would have been exactly as above. For by adding, according to the rule heretofore given, one second for every six minutes of elapsed time, which in this case would be 28 seconds, the corrected clarged time would be 2h. 50m, and every part of the work would be 28 above.

clapsed time would be 2h. 50m. and every part of the work would be as above. If the planet Mars had been observed, at the same corrected altitudes, on the 19th June, 1820, in a place where his declination at the middle time between the two observations was, by the Nautical Almanac, 12° 16' N. and the elapsed time 2h. 49m. 46s. the calculation would still be the same. For, by the Nautical Almanac, it appears that Mars passes the meridian on the 19th and 25th of June, at 4h. 21m. and 4h. 9m. accelerating 2 minutes per day. This being less than the numbers in Table XXXI. is to be doubled (as in note to Form III.) and the elapsed time being found at the side, the corresponding correction 28" halved and added to the elapsed time 2h. 49m. 46s. gives the hour angle 2h. 50m. to be used as above, all the work being the same. Proceed in like manner if the moon was observed at a time when the declination varies but littie.

EXAMPLE III.

Being at sea, in latitude 50° 40′ N. by account, when the sun's declination was 20° 0′ S. at 10h. 17m. A. M. per watch, the sun's correct central altitude was found to be 17° 13′, at 11h. 17m. per watch, the correct central altitude was found to be 19° 41′. Required the latitude?

Subtracting 10h. 17m. from 11h. 17m. gives the elapsed time 1h.

Col. 1.	Col. 2.	COL. 3.		
Elap. ti. [P.M.] 1h. 0m. co-sec. 10.98430 Declination 20° 0' S. sec 10.02701		co-sec. 10.46597		
A co-sec. 10.91131	co-sine 9.99670	co-sine 9.99679		
4 sum alts. 18 27 co-sine 9.97708	co-sec. 10.49966	B 20° 10' S. co-sec. 10.46265		
½ diff. alts. 1. 14 sine 8.33292	scc.' . 10.00010			
C sine. 9.22131	co-sine 9.99390	N. or S. like dec.] co-sine 9.99390		
Z [Less than 90° and N. or S. like bearing of Zenith.]		Z 71 8 N.		
[E is non of B, Z, of some name, difference of different name.]		E 50 58 N. sine. 9.89030		

If the sun had passed the meridian to the north of the observer, Z would have been 71° 08′ S. and E 91° 18′ S. whose sine 9.99989 added to 9.99590 gives the sine of the latitude 9.99579, corresponding to 80° 20′ S.

EXAMPLE IV.

Being at sea in the latitude of 60° 0' N. by account, when the sun was on the equator (or had no declination) at 1h. 0m. P. M. per watch, his correct central altitude was 28° 53', and at 3h. 0m. P. M. per watch, the correct central altitude was 20° 42'. Required the true latitude?

Col. 1.			L. 2.	Col. 3.			
Elap. time [P. M.] 2h. Om. co-sec. Declination 0 sec.	10.58700 10.00000	••••			[co-sec. Infinite.]		
A 30° 0' co-sec.	10.58700	co-sine	9.98494		[co-sine 9.98494]		
1 sum. alts 24 471 co-sine	9.95801	co-sec.	10.37745	B 0° 0′	[co-sec. Infinite.]		
4 diff. alts 4 51 sine	8.85340	sec	10.00110	B less that like dec	90° named N. or S.		
€ sine	9.39841	co-sine	9.98594		co-sine 9.98594		
Z [Less than 90°, and N. or S. like Zenith.]	bearing of	sec	10.34943	Z 63° 26′	N.		
[E is sum of B, Z, if of same name, of diff. name.]	, difference	ì		E 63 26	N. sine 9.95154		

Latitude 59 59 N. sine 9.93748

The calculations would have been the same for south latitude, which would be 59° 59′ S. The computation of A and B might have been dispensed with, for when the declination is nothing, B is nothing, and A is equal to the clapsed time 2h. turned into degrees by Table XXI. being in this example 30°; in this case all the terms included between the brackets [] might be omitted.

In the preceding examples both altitudes were supposed to be taken at the same place or station; but as that is seldom the case at sea, the necessary correction for any change of place must be made in the following manner.

Let the bearing of the sun be observed by the compass at the instant of the first observation; take the number of points between that bearing and the ship's course, corrected for lee-way, if she makes any; with which, if less than eight, or with what it wants of 16 points, if more than eight, enter the traverse table, and take out the difference of latitude corresponding to the distance run between the observations. Add this difference of latitude to the first altitude, if the number of points between the sun's bearing and the ship's course were less than eight; but subtract the difference of latitude from the first altitude, if the number of points were more than eight, and that altitude will be reduced to what it would have been if observed at the same place where the second was.* This corrected altitude is to be used with the second observed altitude in finding the latitude by the above rule. The latitude resulting, will be that of the ship at the time of taking the second altitude, and must be reduced to noon by means of the log.

EXAMPLE V.

In a ship running N. by E. ‡ E. per compass, at the rate of 9 knots per hour, at 10h. 0m. A. M. per watch, the sun's correct central altitude was found to be 13° 18' bearing S. ‡ E by compass, and at 1h. 40m. P. M. per watch, the sun's central altitude was found to be 14° 15', the latitude by account being 49° 17' N. and the sun's declination 25° 28' S. Required the true latitude?

^{*} This is the only correction necessary to make full allowance for the run of the ship; and the unexperienced calculator must take care not to fall into the error of applying a correction to the clauged time, as is directed in several works of note, particularly in the "Complete Newgater," by Dr. Macka. This will appear evident by supposing in the above Example V. that execond observer, with a watch regulated exactly like that used by the first, was at rest at the place of the second observation. Then at the first observation at the same moment of time by both watches, the first observation observations at the sun's altitude 35° 18°, and the second observer 12° 49°. At the second observation the times and altitudes would be altic, so that the elapsed time found by both observers would be the same, and the observations would require no correction, except what arises from reducing the altitude from 13° 18′ to 12° 43°, because the second observer is supposed to be at rest, and his observation requires no correction.

The correction to the first altitude.

The time elapsed between the observations was 3h. 40m. and in that time the ship sailed 33 miles upon the course N. by E. ‡ E. which makes an angle of 13½ points with the sun's bearing at the first observation S. ‡ E. the complement of which to 16 points is 2½ points. Now in Table I. the course 2½ points, and distance 33m. give 29 miles difference of latitude, which must be subtracted from the first altitude 13° 18' because the ship sailed above 8 points from the sun; therefore the first altitude corrected will be 12° 49', which must be used in the rest of the work.

Cor. 1.	Col. 2.	Cor. 3.
Elap. time [P.M.] 3h; 40m. co-sec. 10.33559 Declination 23° 28' S. sec 10.03749		co.sec. 10.39989
A co-sec. 10.37308	co-sine 9.95704	co-sine 9.95704
3 sum alts. 13 32 co.sine 9.99777	co-sec. 10.63076	B 26° 05'S. co-sec. 10.35692
½ diff. alts. 0 43 sine 8.09718	sec 10:00003	[B less than 90° and named N. or S. like declination.]
C sine 8.45803	co-sine 9.99982	co-sine 9.99982
	Z sec. 10.58765	Z 75 01 N.
Zenith.] [E is sum of B, Z, if of same name, difference if of different name.]	'	E 48 56 N. sine 9.97734
of different name.]	Latitude	48 54 N. sine 9.87716

If the sun had passed the meridian to the north of the observer, Z would have been 75° 01′ S. and E 101° 06′ S. whose sine 9.99180 added to 9.99982 gives the sine of the latitude 9.99162 corresponding to 78° 47′ S.

EXAMPLE VI.

Sailing N. E. ½ E. by compass, at the rate of 9 knots an hour, at 0h. \$1'40'' P. M. per watch, the altitude of the sun's lower limb, was 28°20' above the horizon of the sea, the eye being elevated 20 feet above the surface of the water, and the sun's bearing by compass S. by W. and at 2h. 58m. 20s. P. M. by watch, the altitude of the sun's lower limb was 16° 41' above the horizon, the eye being elevated as before, the latitude by account, at the time of the last observation, 48° 0' north, and the declination 13° 17' south. Required the true latitude at taking the last observation?

The correction of these altitudes for semi-diameter, parallax and dip, was 12 miles additive, which makes them 28° 32′ and 16° 53′; the refraction corresponding to the first was 2 miles, and for the second 3 miles, by subtracting which we have the true central altitudes 28° 30′ and 16° 50′. Now the elapsed time between the observations was 2h. 26m. 40s. during which the ship sailed 22 miles (at 9 miles per hour) in the direction of N. E. ½ E. per compass, the bearing of the sun at the first observation S. by W. being 12½ points distant from the ship's course, and as 12½ points want 3½ of 16 points, I enter Table I. and find the course 3½ points and distance 22, corresponding to which in the latitude column is 17 miles, which subtracted from the first altitude 28° 13′; with this and the second altitude 16° 50′, I calculate the latitude in the following manner:

Col. 1.		L. 2.	Col. 3.	
Elsp. ti. [P.M.] 2h. 26' 40" co-sec. 10.5	0232	1		
Declination 13° 17' S. sec 10.0	1178		co-sec.1	0.63871
A co-sec. 10.5	1410 co.sine	9.97861	co-sine	9.97861
1 sum alts. 22 311 co-sine 9.9	6553 co-sec.	10.41670 B		
	2011	10.00215	[B less than 90°, a N. or S.	nd named like decl.]
Csine 9.4	7603 co-sine	9.97962 .	co-sine	9.97962
Z [Less than 90°, and N. er S. like beau	ing of Z see.			
Zenith.		I	E 51 13 N. sine	9.89183
[is sum of B, Z, if of same name, differ of diff. name.]	ence III	Latitude	48 03 N. sine	9.87146

If the sun had passed the meridian to the north of the observer, Z would have been 65° 11' S. and E 79° 09' S. whose sine 9.99217 added to co-sine of C 9.97962 gives the sine of the latitude 9.97179, corresponding to 69° 54' S:

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EXAMPLE VII.—(Same as Dr. Brinkley's Nautical Almanac for 1800.)
The latitude by account 6° 30′ N. sun's declination 5° 30′ N. the sun's correct central altitudes were found to be 35° 21′ and 70° 01′, with an elapsed time between the observations of 2h. 20′. Required the latitude, the sun passing the meridian south of the observer?

Z [Less than 90°, and N. or 6. like bearing of Z sec. 10.00097 Z 3 50 N. Zenith.]

[E is sum of B, Z, if of same name, difference if of diff: name.]

[E is name.]

Latitude 7 38 N. sine .. 9.22211

If the sun had passed to the meridian north of the observer, Z would have been 3° 50′ S. and E=1° 56′ N. whose sine 8.52810 added to the co-sine of C 9.90170 is 6.42980, which is the sine of the other latitude 1° 32′ N. so that in this example both latitudes are north.

SECOND METHOD

of finding the latitude by double altitudes of the sun, when the variation of declination is neglected.

This method of finding the latitude depends on a set of tables, marked XXIII. in this collection, first prepared by Mr. Douwes, containing three logarithms titled Half Elapsed time, Middle time, and Log. rising. The two former are arranged together as far as six hours, the latter is placed at the end of the table, and is extended in the present edition as far as 12 hours. The table with the proper title must be entered at the top with the hour, at the side with the minute, and in the column marked at the top with the seconds, the corresponding number will be the sought logarithm, to which must be prefixed the index of the log. under 0" in the same horizontal line. to the time 3h. 52' 10" correspond the log. half-elapsed time 0.07138, log. middle time 5.22965, and log. rising 4.67274. In general it will be sufficiently exact to take these logarithms to the nearest 10 seconds, particularly when the sun's zenith distance is great; but if the log. to the nearest second is required, it may be found by taking the difference of the tabular logarithms corresponding to the next greater and next less time, and saying as 10" is to that difference, so are the odd seconds of time to the correction of the first tabular logarithm, additive, if increasing, subtractive, if decreasing. Thus if the log. & El. time corresponding to 3h. 52' 18" were required: the logs. corresponding to 3h. 52' 10" and 3h. 52' 20" are 0.07138 and 0.07119, whose difference is 19, then 10": 19:: 8": 15.—This subtracted from 0.07138 leaves 0.07123, the sought logarithm. By inverting the process we may find the nearest second corresponding to any given logarithm. We shall now give the rule for calculating the latitude adapted to double altitudes of the sun.

RULE.

To the log. secant of the latitude by account (Table XXVII.) add the log-secant of the sun's declination (Table XXVII.) rejecting 10 in each index, the sum is to be called the log. ratio.

From the natural sine of the greatest altitude (Table XXIV.) subtract the natural sine of the least altitude (Table XXIV.) find the logarithm* of their difference (in Table XXVI.) and place it under the log. ratio.

Subtract the time of taking the first observation from the time of taking the second, having previously increased the latter by 12 hours when the observations are on different sides of noon by the watch; take half the remainder, which call half the elapsed time.

With half the elapsed time enter Table XXIII. and from the column of half elapsed time take out the logarithm answering thereto, and write it under the log. ratio.

^{*} The index of this logarithm being as usual one test than the number of figures contained in the sufference of those natural sines. You must also observe that the ultitudes to be used are the correct central alditudes: that is, the observed altitudes corrected for dip, semi-diameter, parallax and refraction.

Add these three logarithms together, and with their sum enter Table XXIII. in the column of middle time, where, naving found the logarithm nearest thereto, take out the time corresponding, and put it under half the elapsed time. The difference between these times will be the time from noon when

the greater altitude was taken.

With this time enter Table XXIII and from the column of log. rising, take out the logarithm corresponding, from which logarithm subtract the log. ratio, the remainder will be the logarithm of a natural number, which being found in Table XXVI.* and added to the natural sine of the greater altitude, will give the natural co-sine of the sun's meridian zenith distance, which may be found in Table XXIV. Hence the latitude may be obtained by the rules of page 121.

NOTES.

1. If this computed latitude should differ considerably from the latitude by account, it will be proper to repeat the operation, using the latitude last found instead of the latitude by account, till the result gives a latitude nearly agreeing with the latitude used in the computation.

2. This method is best suited to situations where the sun's meridian zenith distance is not much less than half the latitude; for in latitudes where the sun approaches near to the zenith, the observations must be taken much nearer to noon: and the preceding rule, instead of approximating, will in some cases give the results of successive operations, wider and wider from the truth. To remedy this difficulty, a set of tables was published by Dr. Brinkley, at the end of the Nautical Almanac for 1799; but the great variety of cases incident to his method will hinder it from being generally used. Instead of Dr. Brinkley's method, we may generally use the method of arithmetical computation, called Double Position, which will frequently give, in a more simple manner, the required latitude, as will be shown in Example X. and in general it may be observed, that where Douwes' rule does not approximate, the object is most commonly so situated, as not to furnish the necessary observations to obtain a correct latitude, whatever method of computation might be used.

3. The operation is the same whether the sun has north or south declination; and also whether the ship is in north or south latitude. When the sun has no declination, the log, secant of the latitude (rejecting 10 in the index) will be the log, ratio: and when the latitude by account is nothing, the secant of the declination (rejecting 10 in the index) will be the log. ratio. This rule, as well as the former, is founded on the supposition that the declination is taken for the middle time between the observations, and that it does not vary during the elapsed time, which, however, rarely happens, and a correction ought to be applied to the latitude on this account, but this correction is generally small, and if it is large, the third method must be used, or the new method in the Appendix of this work.

EXAMPLE VIII.—(Same as Example L preceding.)
Being at sea in latitude 46° 30' N. by account, when the sun's declination was 110 17' N. at 10h. 2m. in the foregoon, the sun's correct central altitude was 46° 55', and at 11h. 27m. in the forenoon, his correct central altitude was 540 9'. Required the true latitude, and true time of the day when the greater altitude was taken?

	T	`ime	mes.				•			
		M.		Al	t.]	Nat.	Si.	Lat. by acc 46° 30' Sec	
2 obs.	11	27	0	64 ~	9′		310	5 5	Dec 11 17 Sec.	0.00848
1 obs.	10	2	0	46	55		730	3 6	Log. ratio	0.17067
Elap. time		25 42	-	Diff.	Nat	. Sines	80	19	Log. Diff. Nat. Sines Log. & Elap. time	
						1	о. м	. 5.		
		Mi	iddle	time		1	15	10		4.80908
		§]	Elap.	time			42	30		
		2 (Obs.	from	noon	0	32	40	Its log. rising Log. ratio sub.	3.00608 0.17067
		Na	t. ni	umb.					685 corresponding to log.	2.83541
Nat. sine greatest alt.				81035						
		Sum is nat. co-sine ⊕'s zen. dist. 81740 equal to 35° 10′ N. ⊙'s declination						•		
		La	titud	e in .			• • • •	• • • •		

[&]quot;Taking as usual a number of figures equal to the index of that logarithm increased by suity.

The latitude 46° 27' differing only 3' from the latitude by account, may be assumed as the true latitude.

By means of the time of the second observation from noon above found 32' 40", the error of the watch may be found; for in the present example, by subtracting 32' 40" from 12h. we have the time of the second observation 11h. 27' 20"; but the time of the watch was 11h. 27' 0"; therefore the watch was 20 seconds too slow; a small difference would be found in these numbers, if we were to proportion the logarithms of Tab. XXIII. to seconds. In the same manner the error of the watch may be found in the following examples.*

EXAMPLE IX .- (Same as Example V. before given.)

In this example the latitude by account was 49° 17′ N. The sun's declination 23° 28′ S. The first altitude corrected as before taught 12° 49′, the second altitude 14° 15′. From which the true latitude is required?

2 Obser. 1 Obser.	H. 13 10	40	0	Alt. Nat. 14° 15'=246 12 49=221	15	Lat. by acc. Declination	49° 23	17' 28		0.19554 0.03749
,				Diff. nat. si. 24		Log. ratio Its. log				
d Elap. time	1	50	0	:		lts log	••••	••••		0.33559
	0	10	10	Time correspond	ding (.o	••••	••••		3.94458
	1	39	50	Its log. in col. o Log. ratio	f ris	ng is	••••		•••••	3.97028 0.22303
. •				558 24 61		Nat. number of	. 		log.	3.74725
				f. zen. dist. 3020						
Latitude .	•••		•••			48 57 N.		•		
				e by computati st be repeated.		liffers consider	abl y	from	that	by ac:
						Lat. last found	48	0 57/	Sec.	0.18262

			•		Decimation	_z	zo	Sec.	0.03/49
Llapsed time	1	м. 5 0	()	•	Log. ratio Diff. N. sine 2432		Its		0.22011 3.38596 0.33559
Middle time Time from noon		40			Its log		· • • •	••	3.94166
THE HOM NOOL	•	20	•		Its log. in col. of ri	sing	••••	••	3.97170 0.22011
				5614 Nat.	number of		lo	g.	3.75159

4010				
0259	Nat. cos. mer. zen. distance	72 0	23/	N.
	Declination			

This latitude, differing only two miles from that used in the computation, may be depended upon as the true latitude of the ship at the time of the second observation. If the first altitude had not been corrected, the computed latitude would have been found—48° 40′ N.

^{*} When the middle time is greater than half the elapsed time, both observations are on the same aide of the meridian; otherwise, on different sides; whence it is easy to determine whether the greater airitade be observed before or after mon.

EXAMPLE X .- (Same as Example VII. before given.)

The latitude by account 6° 50' N. sun's declination 5° 30' N. the sun's correct central altitudes 350 21' and 700 01', elapsed time 2h. 20' are given to find the true latitude.

Making the calculations with the latitude by account 6° 50', the computed latitude by the first operation will be 80 17'. Repeating the operation with the latitude 8° 17', the second operation will give 7° 10'.* This must be used for a third operation, and by repeating the calculation accurately to seconds, it will require more than a dozen operations to obtain the true latitude 7° 58', which was found by the first method by a single operation. Dr. Brinkley made the latitude 7° 50' differing 8' from a strict calculation by spherical trigonometry. The detail of this calculation is not here given. but is left to exercise the learner. The object of the present example is to shew how the number of operations might be decreased by the arithmetical method of double position before mentioned.

Take the error or difference between the first ssumed latitude 6° 30' and the first computed latitude 8° 16' equal to 106'; also the $8 \cdot 16 \times 66 = 429 \cdot 00$ gror or difference between the second as-

Errors. Products.

172)1305 16(7°35 sumed latitude 8° 16' and second computed latitude 70 10' which is 66'. Multiply them crosswise as in the adjoined scheme, according to the usual rule of double position, dividing the sum of the products 1305° 16' by the sum of the errors 172, gives the corrected. latitude 7° 85' N. The sum of the products was taken in this case, because one of the assumed latitudes was greater and the other less than its corresponding computed latitude. If both computed latitudes had been greater or both less than the corresponding assumed latitudes, the differences of the errors and of the products ought to have been taken. It will rarely happen that more than one process of this kind will be required to give a correct result. In the present instance, however, it will be necessary; for, by repeating the operation with the assumed latitude 7° 35', the resulting computed latitude is 7° 41½, and the third error 6½. Repeating anew the computation, with this and the second latitude 8° 16' and second error 66', the resulting latitude is 70 38', the same as was found by the direct computation by the first method, and as accurately as could be obtained by repeating the operations about fourteen times by the second method.

In general, when such a large number of operations are required to produce a correct regult, it is a sure proof that the situation of the object is not well adapted to obtain an accurate latitude, and it would be lost labour, and lead to great mistakes to attempt it. Thus, in the present example, if the greatest altitude had been decreased only 12' 42", making it 690 48' 12", leaving unaltered the other altitude 350 21' and the interval 2h. 20m. the latitude of the place of observation would be 0, or under the equator, as is easily proved by computing the altitudes of the sun for the times 1h. 17m. 50s.8 and Sh. 57m. 50s.8, under the equator when the declination is 50 30' N. by the rules hereafter given. Hence it appears that a change of 12' 42" in the greatest altitude, would alter the computed latitude from 7° 38' to 0°. which makes an error of one degree of latitude for an error of 13 miles in that altitude, and as errors in the altitudes of this magnitude are easily committed at sea, even by very good observers, it shows very clearly the defect of the method of double altitudes when the sun approaches near to the zenith. This does not arise from any defect of the method of computation, but is an inherent defect of the method itself, which no process of spherics can remedy, and there is no other resource left, in such cases, than to make use of another object to determine the latitude.

Slight differences will be found in these calculations by using logarithms to seven places of figures i making the calculation accurately to seconds.

If the degrees of both latitudes are alike, the minutes only may be retained in these multiplications.

THIRD METHOD.

To find the latitude by double altitudes of the same or different objects, the declinations being different. [See Appendix to this work.]

This method, like the first, requires only the use of Table XXVII. In this rule the words sine, co-sine, &c. are written for log. sine, log. co-sine, &c. The logarithms are arranged in these columns as in the first method, according to the following formula, which ought to be written down before the calculation is commenced; this will simplify the operation, and may prevent mistakes. In this formula it is said that C is of the same affection as B, the meaning of which is, that if B is less than 90°, C also is less than 90°; and if B is greater than 90°. Likewise A is of the same affection as the hour angle, meaning that if the hour angle is less than 6 hours or 90°, A will be less than 90°; and if the hour angle exceed 6 hours, the angle A will exceed 90°.

		FORM	U LA.		•
	M.] sec.		Cel. 2.		Col. 3.
Decil.d [at gr. alti.]	· tan.	• · · · · · · ·	sine	i	
A [dir. name from d.]	tan-	A[same aff	.as H.]co-sec.		co-sine
D. Dec. [atle.alt.]				1	
в	****	ļ	co-sine	ļ	09-sec.
. c	co-sec	C [same at	Las B] co-sine	F	co-tan.
.7*		1.		z	[F less. 90°, diff.
		G	sin e	G	name from B
Least altitude	sec.		co-L		sino
Greatest altitude		I	tan.	I less 90	
Sum, 3 last num.		Dec. D	[at least elt.]	- [I pamed	as G. j
4 Sum	co-sine	K			sine
} S.—gr.alt.—Rem.	sine		. Latit	ude	sine
Sum of 4 logs.	2)				
₫ Z	sine				•
Z [named N. or S. like	bearing of Zen	ith.]		•	

In some late works on Navigation, no notice is taken of the cases where the hour angle exceeds 90°, or the distance of the objects exceeds 90°, and on that account the rules appear less subject to different cases than the following rule, which embraces all possible cases, and the apparent simplicity of the rules referred to, arises from their imperfections and incompleteness.

RULE.

1. Find the hour angle H, and take out the corresponding secant, which put in Col. 1. and its tangent in Col. 3.

2. Take the declination d, corresponding to the greatest altitude, place its

tangent in Col. 1. its sine in Col. 2.

3. The sum of the two logarithms in Col. 1. (rejecting 10 in the index) is the tangent of the angle A, which is less than 90° if the hour angle is less than 6 hours (or 90°) but greater than 90° if the hour angle is greater

^{*} The hour angle is the same as the elapsed time in double eltitudes of the sun.. This time is turned into degrees by Table *XXI. but it is more simple to double the hour angle and find it in Col. P.M. Table XXVII. and take out its corresponding tangent. If this double angle exceed 12h. reject 12h. and find the remainder in Col. A. M. and take out its corresponding tangent. In the following examples this double ongle is marked with the letters P. M. annexed.

than 6 hours. This angle is to be marked north or south, with a different name from the declination d, at the greatest altitude. The co-secant of A is

to be placed in Col. 2. its co-sine in Col. 3.

Place the declination D, corresponding to the least altitude, below the angle A, and if they are of the same name take their sum, but if of different names, take their difference, and call this sum or difference, the angle B, making it north or south like the greatest of the two quantities A, D. The co-sine of B is to be placed in Col. 2. its co-secant in Col. 3.

5. The sum of the three logarithms in Col. 5. (rejecting 20 in the index) is the co-tangent of an angle F, (less than 90°) which is to be taken out and marked north or south, with a different name from B.

The sum of the three logarithms in Col. 2. (rejecting 20 in the index) is the co-sine of the angle C, which is to be taken less than 900, if B is less than 900, but greater than 900 if B is greater than 900. The angle C and

its co-secant are to be placed in Col. 1.

7. Place the altitudes below C, take the half sum of these three quantities. subtract the greatest altitude from the half sum, and note the remainder. Place the secant of the least altitude in Col. 1. its co-tangent in Col. 2. its sine in Col. 3. The co-sine of the half sum in Col. 1. and the sine of the remainder in Col. 1. The sum of the four last logarithms of Col. 1. (rejecting 20 in the index) being divided by 2, gives the sine of an acute angle, which being found and doubled, gives the zenith angle Z, which is to be named north, if the zenith and north pole are on the same side of the arch or great circle, passing through the two objects, (or the two observed places of the same object) but south if the zenith and south pole, are on the same side of that great circle.1

Take the sum of the angles Z and F, if they are of the same name. but their difference if of different names; this sum or difference is the angle G, to be marked north or south like the greatest of the angles Z, F.6 The

sine of G is to be placed in Col. 2.

9. The sum of the two lower logarithms of Col. 2. (rejecting 10 in the index) is the tangent of an angle I, which is to be taken out (less than 90°) and marked north or south like G. The secant of I is to be placed in Col. S.

10. Write the declination D, corresponding to the least altitude below I, take their sum is of the same name, their difference if of different names. This sum or difference is the angle K, of the same name as the greater of these two quantities. The sine of K is to be placed in Col. 5.

The sum of the three last logarithms in Col. 3. is the sine of the

required latitude of the same name as K.

EXAMPLE XI.

Given the sun's correct central altitude 41° 33', and his declination 14° N. After an interval of 1h. 30m. by watch, his correct central altitude was 50°, and his declination 13° 58' N. Required the latitude, the sun being south of the observer when on the meridian?

§ If the sum be taken to find G, and it exceeds 199°, subtract it from 300° and call the remainder G with a different name from Z or F.

This rule is easily remembered in three places in which it occurs, from the circumstance that s is the first letter of sum and sums, and d the first letter of difference and different.
 If the sum be taken to find B and it exceed 180°, subtract it from 360°, and call the remainder B

if it it is sum be taken to find B and it exceed 180°, subtract it from 360°, and call the remainder B with a different name from A, D.

† This case occurs also in the first and second methods of solution, and it must be determined on the spot by the situation of the objects. In couble altitudes of the sun, moon, or planets, when the elapsed time is not very great, the angle Z is generally to be marked with the bearing of the zenith from the observed object, when at its greatest altitude on the meridian, which in north latitudes, without the tropics, is in general north; in south latitudes, without the tropics, south. Sometimes when the sun passes the meridian near the zenith, it may be doubtful whether the zenith be north or south; in which cases the problem may be solved for both cases (which increases the labour but lite) and that one of the two computed latitudes selected which agrees best with the ship's reckoning; but it is generally safest most to the observations of this kind, which are generally liable to great errors from small mistakes in the altitudes.

Col. 1. Hour ang. H 1b.30m.[P.M Decli. d [at gr. akti.] 13°.	- I.Sh.j sec. 10.03438 58' N. um. 9.58569	Col. 2.	Col. 3. 9.61722
A[dif. name from d.] 15 0	4 S. tan. 9.43007	A[same aff. as H.] co-sec. 10.5851:	co-sine 9.98481
D Dec. [at least alt.] 14 0	e N.	`	
В 1 0	ж s .	co-sine 9.9999	co-sec. 11.75012
C	49 co-sec 10.42988	O [mine aff. as B] co-sine 9.9877	F 2°40' N. co-tan. 11.83215
		•	Z 57 18 N. [F less 99°, diff. name fm. B.]
Least altitude 41	35 sec. 10.12588	co-tan. 10.0524	sine 9.82169
Greatest altitude 50	00	I 44° 20' N. tan. 9.9698	[I less 90°] sec. 10.14552 [I named as G.]
Sum		Dec. D 14 00 N. [at least alt.]	[1 1111111 11 11 11 11 11 11 11 11 11 11
1 Sum		K 58 20 N.	sine 9.92993
S-gralt-Rem. 6	11 sine 9.06589	Latitude	52 7 N. sine 0.89720
Sum 4 logs.	2)19.36143	· ·	
₫ Z28	39 sine 9.88071		
Z57	18 N. [named like	bearing of Zenith.]	•

If the latitude had been south, Z, instead of being 57° 18' North, would be 57° 18' South; G, 54° 38' S. I, 42° 37' S. K, 28° 37' S. and the latitude \$25° 34' S. The labour of making this extra calculation is but little, and where any doubt exists of the name of Z, it is best to make the computation both ways; this, however, will rarely happen. The calculations of this exsorple, and most of the following ones, are made to the nearest minute; where great accuracy is required, it will be proper to take the logarithms. and angles corresponding to seconds.

EXAMPLE XII.

The sun's correct central altitude was 32° 25', his declination 17° S. 8 hours afterwards, by a watch, the sun's correct central altitude was 30° 8', and declination 16° 55' S. the observer being in a high south latitude. Required the latitude?

-		
Col. 1. Hour H 8h. [P.M.16h.—4h.A.M.]sec. 10.30103	Col. 2.	Col. 3.
Decli. d [at gr. alti.] 17000' S. tan. 9.48534	1	1 .
A[dif.name from d.] 148 33 N. tan. 9.79637	A[same aff.as H.] co-sec. 10.28255	co-sine 9.93100
D Dec. [at least alt.] 16 55 S.		
B 151 S8 N.	co-sine 9.82240	co-sec. 10.12644
G 111 51 co-sec. 10.03288	O [same aff-as B] co-sine 9.57067	F 28° 50' S. co-tan. 10.29600
		Z 25 48 S. [F less 90°, diff. name fm. B.]
	• •	G 52 56 8.
Least altitude 30 08 sec. 10.06805	co-tan. 10.23623	i · _
Greatest altitude 32 25		[I less 90°] sec. 10,22922 [I named as G.]
Sum*174 24	Dec. D 16 55 S. [at least alt.]	
3 Sum 87 12 co-sine 8.69886	K 70 46 S	sine 9.97506
1 S.—qr. alti.—Rem. 54 47 sine 9.91221	Latitude	. 53 28 S. sine 9.90500
Sum 4 logs 2)18.69850		
Ł Z 12 53 sine 9.34825		•
Z 25 48 S. [named like	bearing of Zenith.]	

If the zenith had been north of the great circle passing through the sun and moon, we should have $Z=25^\circ$ 46' N. G 1° 64' S. I 1° 50' S. K 16° 45' S. and the latitude 90 18' S.

EXAMPLE XIII.

Suppose, at the same moment of time, the moon's correct central altitude was 55° 20', the moon's declination 0° 36' N. the sun's correct central altitude 37° 40', the sun's declination 0° 17' S. The hour angle, or difference of the right ascensions of the sun and moon, being by the Nautical Almanac 5 hours or 75°. Required the latitude, supposing it to be north?

Col. 1.	Col. 2.	Col. 3.
Hour angle H 5h. (P. M. 10h.) sec. 10.58700 Decli. d (at gr. alt.) 0° S6′ N. tan. 8.02004		
A(dif. name from d.) 2 19 S. tan. 8.60704	A(same aff.as H.) co-sec. 11.39538	co-sine 9.00064
D Decl. (at least alt.) 0 17 S.		
B 2 36 S.	co-sine 9.90955	co-sec. 11.34550
C 75 00 co-sec, 10.01506	C same aff. as B) co-sine 9.41295	F 0º 42' N. co-tan. 11.91489
		Z 29 40 N. (F less 90°, diff.
	G sine 9.70575	G 30 22 N.
Least altitude 37 40 sec. 10.10151	co-tan, 10.11211	sine 9.73609
Greatest altitude 55 20		[I less 90°] sec. 10.07748 [I named as G.]
Sum168 00	Dec. D 0 17 S. (at least alt.)	
3 Sum 84 00 co-sine 9.0132	K 32 56 N	sine 9.76533
2 S—gr. alt.=Rem. 28 40 sine 9.6803	Latitude	28° 24' N. sine 9.59890
Sum of 4 logs. 2)18.81673		•
½ Z 14 50 sine 9.4003		•

If the zenith had been south of the great circle passing through the objects, we should have Z 29° 40′ S. G 28° 58′ S. I 32° 6′ S. K 32° 23′ S. and the latitude 22° 44′ S.

EXAMPLE XIV.

Given the moon's correct central altitude 47° 37′, the moon's declination 17° 29′ S. the sun's correct central altitude, at the same time, 27° 22′, the sun's declination 8° 28′ S. the hour angle, or difference of right ascensions of the sun and moon, 5h. 40m. 28s. or 85° 7′. Required the latitude, supposed north?

Dan and moon, on.	10	MI. 200.010		Tred	an ca an	C MILLI	ac, supposed norur:
Col. 1 Hour H 85°7' (P.M.11b.2 Deeli, d (at gr. ait.) 17°	0′ 5 29′	6") sec. 11.06998 S. tan. 9.49829		c	ol. 2. sine	9.17774	Col. 3. tag. 11.06885
A (dif. name from d.) 74	53	N. tan. 10.56821	A(same	affas l	I.) co-sec.	10 01529	co-sine 9.41628
D Decl. (at least alt.) 8	22	s.					
В 66	25	N.		• • • • •	co-siue	9.60215	co-sec. 10.83788
C E2	51	co-sec. 10.00339	C (same	afi. us	B) co-sine	9.09518	F 16° 43' S. co-tan. 10.52251
			G		sine		Z 39 20 N. (F less 90°, diff. name (m. B.)
Least altitude 27	22	sec. 10.03155			co-tan.	10.28599	sine 9.66246
Greatest altitude 47	57		T :	3 6º 3 7'	N. tan.		[I less 90°] sec. 10.09548
Sum	50		Dec. D	8 28	S. (at leas	t alt.)	[I named as G.]
½ Sum 78	55	co-sine 9.28384	K	28 09	N		sine 9.67874
1 S gr. alt. = Rem. 31	13	sine 9.71660	,			Latitude	15° 41' N. sine 9.45163
Sum of 4 logs.		2)19.0543					
½ Z 19	40	sine 9.5271	9				
Z 39	20	N. (named lik	e bearin	g of Z	enith.)		

If the zenith had been south of the great circle passing through the objects, we should have Z 39° 20′ S. G 56° 3′ S. I 58° 2′ S. K 66° 30′ S. and the latitude 52° 46′ S.

13

FOURTH METHOD.

To find the Latitude from the altitudes and distances found in taking a lunar observation.

This is a particular case of Form V. and is more simple than the general solution, because the true distance of the objects, computed in working the lunar observation, may be used to shorten the calculation of the latitudes; we

shall therefore give a particular rule for this method.

Having the apparent altitudes and distance of the objects, find, by any of the methods of working a lunar observation hereafter given, the true distance. Find also the true altitudes, by correcting the apparent altitudes for parallax and refraction. The correction of the moon's altitude is equal to the difference between 59' 42" and the correction already found from Table XIX. in working the lunar observation; this difference, added to the moon's apparent altitude, gives her true altitude. In like manner the correction of the sun's altitude is equal to the difference between 60' and the correction already found in Table XVIII. (or in Table XVII. if a star is used): this difference is to be subtracted from the sun's (or star's) apparent altitude, to obtain his The time at Greenwich, corresponding to the true distance, true altitude. having been found in working the lunar observation, take from the Nautical Almanac, for this time, the declinations of the sun and moon, as was taught in pages 110, 124, and, if great accuracy is required, the correction for second differences of the moon's declination may be noticed, as in Problem I. of the Appendix to this work. If, instead of the sun, a star is used, its declination may be obtained from Table VIII. or more accurately from the Nautical Almanac, being one of the 24 bright stars whose places are now given for every ten days in that work. From these declinations, the north polar distances must be found, by adding the declinations to 90° if south, or subtracting from 900 if north.

Having thus obtained the *true* distance, the *true* altitudes, the declinations and north polar distances, the latitude may be computed by the following rule, adapted exclusively to Table XXVII. writing, as before, sine, co-sine, &c. for log. sine, log. co-sine, &c. the logarithms being arranged in three columns as in the former methods.

RIII.

RULE.

1. Place in Col. 1. the true distance and the polar distances. Take their half sum, subtract from this half sum the polar distance of the object which had the greatest altitude, and note the remainder. Put, in the same column, the co-secant of the true distance, the co-secant of the polar distance of the object having the least altitude, the sine of the half sum, the sine of the remainder. The sum of these four logarithms (rejecting 20 in the index) being divided by 2, gives the sine of an acute angle, which being found and doubled, is to be called the angle F.

2. Place in Col. 1. the true distance and the true altitudes. Take their half sum, and also the remainder or difference between the half sum and the greatest altitude. Place in the same column the co-secant of the distance (before found) the secant of the least altitude, the co-sine of the half sum, the sine of the remainder. The sum of these four logarithms (rejecting 20 in the index) being divided by 2, gives the sine of an acute angle, which being

found and doubled, is to be called the angle Z.

3. If the zenith and north pole be situated on the same side of the great circle, passing through the two objects, take the sum; of the angles F & Z for the angle G; but if the zenith and north pole be situated on different sides of that great circle, take their difference for the angle G. Place the co-sine of G in Col. 2.

1.4. Write in Col. 2. the co-tangent of the least altitude, and its sine in Col.

: If this sum should exceed 1809 subtract it from 3500 and call the rentainder the angle G.

^{*} In places without the tropics, the sum is used generally in northern latitudes, and the difference in southern latitudes.

The sum of the two logarithms in Col. 2. is the tangent of the angle I, which is to be taken less than 90°, and marked south if the angle G is less than 900, but north if G is more than 900. Place the secant of I in Col. 3.

Place the declination corresponding to the least altitude, below I; Take their sum if of the same name, but their difference if of different names; call this sum or difference the angle K, and mark it with the same name as the greatest of the two quantities. Place the sine of K in Column 3.

The sum of the three logarithms in Col. 3. (rejecting 20 in the index)

is the sine of the latitude of the same name as K.

Having found the latitude, the hour may be obtained by means of the true altitude and declination of the sun or star, by any of the usual methods hereafter given for that purpose; but, if the last of the observed altitudes was that of the sun or star, the horary distance of that object from the meridian might be obtained more simply by the following rule, adapted to Table XXVII.

RULE.—Add the tangent of the angle G, the sine of the angle I, the secant of the angle K, the sum, rejecting 20 in the index, is the tangent of an angle; take out the corresponding time in the Column P. M. or in the Column A. M. increased by 12 hours, half of either of these times is the horary distance of the lowest observed object from the upper or lower meridian, whence the hour may be obtained directly if it be the sun, but if it be the star (or moon) it is obtained by applying its horary distance to the hour of passing the meridian, according to the usual methods of finding the time from an altitude of a fixed star, or the moon.

EXAMPLE.—(Same as Dr. Brinkley's N. A. 1825.)
May 19d. 8h. 6m. P. M. in the longitude of 7h. 23m. west, it was found by working a lunar observation that the correct distance of the centres of the sun and moon was 90° 57′ 20"; true altitude of the sun's centre 11° 85′ 12"; true altitude of the moon's centre 27° 32' 18". At the same time by the Nautical Almanac the sun's declination was 190 56' 48" N. the moon's declination 13° 55' 48" N. Required the latitude and hour by this observation?

Col. 1.		Col. 2.	Col. 3.
	" co-sec. 10.000 06		
P. dist. at le. alt. 70 03 12	co-sec. 10.02687		
P. dist. at gr-alt. 76 04 12			
Sum237 04 44		G is sum of F, Z, if north pole ar of great circle, but their differen	
3 Sum 118 32 22	sine 9.94374		
\$ S p.d.atgr.al. 42 28 10			
• 2.0.00		f ·	
	2)19.8 00 10	ł	I is less than 900,—nanted
.½ F 52 36 00	sine 9.90005	z 61 56 52	south if G is less than 90°, north if G is more than 90°.
Augle F 105 12 00		F 105 12 00	•
True distance 90 57 20	co-sec. 10.00000	G 166 48 52 co-sine 9.98840	
Least alt 11 83 12		co-tan. 10.68947	6iue 9,30163
Greatest alt 27 32 18		78 09 33 N. tan. 10.67787	sec. 10.68728
Sum 130 02 50		Dec. 19 56 48 N. (at least alt.)	,
₫ Sum 65 01 25	co-sine 9.62557	K 98 05 21 N	sine 9,09566
& Sum-gr. alt. 37 29 07	sine 9.78480	Latitud	e 749 48 N. sine 9.98452 -
	2)19.41881		
17 90 40 90	sine 9.70940	1	To find the Hour.
Angle Z 61 36 62			tan. 2.36974
sangress and the de of	•		K
		,	sec. 10.65160
	Hour P. M.	7h, 47m, 424, or A, M,+12h,-16h	. 12m. 16s. tan. 10.21203

Divided by 2 gives the horary distance of the lowest object from the meridian, 51s. 53m. 51s. 8h. 06m. 09s.

The sun being at the lowest altitude, his distance from the upper meridian was 8h. 6m. 9s. being the hour of the day, and the sun's distance from the Iower meridian, or midnight, was Sh. 58m. 51s.

^{*} Both these logarithms may be taken out at the same time when the sine of the angle was found in the computation of the angle Z.

QUESTIONS FOR EXERCISE.

In the following questions the sun's semi-diameter is supposed to be 16'.

and the parallax nothing.

Being at sea, in latitude by account 390 28' N. when the sun's declination was 20° 41' N. at 11h. 30m. 15s. A. M. per watch, the altitude of the sun's lower limb was observed to be 68° 18' 45", and at 12h. 26m. 28s. P. M. was 70° 58', the height of the eye being 21 feet above the surface of the sea. Required the true latitude of the ship?

Answer, 390 28' N.

Being at sea in latitude 50° 40' N. by account, at 10h. 17m. 30s. A. M. per watch, the altitude of the sun's lower limb was observed to be 170 44' and at 11h. 17m. 30s. was 19° 31½, the declination being 20° S. and the height of the eye 21 feet above the sea. Required the latitude in?

Answer, 500 1' N.

Suppose a ship at sea, in latitude 47° 34' N. by account, and that at 9h. 55m. 30s. by watch, the altitude of the sun's lower limb was 17° 24'. bearing by compass S. by E. I E. and at 12h. 54m. 10s. the altitude of the same limb was 21° 45½ the declination being 19° 30' S. the height of the eye 20 feet above the sea, and the ship's course by compass E. & S. at the rate of 7 knots per hour. What was the true latitude?

Answer, 47° 24' N.

At 11h. 23m. 20s. A. M. per watch, the altitude of the sun's lower limb was 28° 18' the sun bearing S. by W. by compass. At 2h. 53m. 20s. P. M. the altitude of the same limb was 16° 40', the height of the eye 20 feet, his declination 130 17' S. and the latitude by account 47° 50' N. the ship's course during the elapsed time N. E. with her larboard tacks on board.* sailing at the rate of 6 knots, and making half a point lee-waywhat latitude was she in when the last altitude was taken?

Answer, 480 9' N.

To find the latitude by one Altitude of the Sun taken near noon, having the time of observation by a well regulated watch.

When the sun does not pass near the zenith, the meridian altitude and the latitude of the place may be accurately determined by observing his altitude when near the meridian, and noting the time by a watch regulated the preceding morning or following evening, by either of the methods given in this work.† To this time by the watch must be applied a correction equal to the difference of longitude made by the ship (turned into time) in the interval between the regulation and the observation near the meridian, by adding when the place of regulation was to the westward of the place of taking the other observation, otherwise by subtracting; the sum or difference will be the time of taking the observation, whence the time from noon will be obtained: with which and the observed altitude (corrected for semi-diameter, dip, &c. as usual) the sun's declination (found in Table IV. and corrected for the longitude of the ship) and the latitude by account; the latitude by observation may be found as follows:

the borizon.



^{*} The larboard side of a ship is the left side, when the observer is aft, looking towards her head, and the starboard is the right side. When a ship is sailing with her tarboard tacks on board, the lee-way is allowed to the right hand; but if her starboard tacks are on board, to the left hand.

In calculating the answers to these questions, proportional parts were taken for the seconds; a small difference would be to addit the nearest logarithms only were taken.

† The best time for regulating a watch is when the sun bears nearly east or west, and is above 10° from

RULE.

Add together the log. co-sine of the latitude by account (Table XXVII.) the log. co-sine of the declination (Table XXVII.) the logarithm in the column of rising (Table XXIII.) corresponding to the time from noon when the observation was taken; reject 20 in the index, the natural number of the remainder being found (in Table XXVI.) and added to the natural sine of the observed altitude (Table XXIV.) the sum will be the natural co-sine, of the meridian zenith distance, from which the latitude may be obtained by the common rules.

If the computed latitude differs considerably from the latitude by account, it is best to repeat the operation, using the latitude last found instead of the latitude by account. This method of finding the latitude by a single altitude of the sun, may be applied to any other celestial object.

EXAMPLE I.

Being at sea in latitude 49° 50′ N. by account, when the sun's declination was 20° S. at 11h. 29m. 20s. A. M. per watch, regulated the preceding morning, in a place 20 miles of longitude to the eastward. the sun's correct central altitude was 19° 41′* bearing south. Required the true latitude?

Time per watch11 29 26 20' in T. by Tab. XXI. 1 2		Co-sine 9.80957 Co-sine 9.97299
Time of observation11 28	Time from noon 0h. 32m. 0s. L	og. rising 2.98820

Nat. Num. 590 log. 2.77076

Sime from noon..... 32 0 Central altitude 19° 41′ Sine 33682

Mer. zen. dist. 69 57 N. Co-sine 34272 Declination 20 0 S.

60 N.

Latitude 49 57 N.

EXAMPLE II.

At sea in the latitude of 60 N. by account, the sun being on the equator, at 0h. 59m. 0s. P. M. per watch, regulated the preceding morning in a place 15 miles in longitude to the westward, the sun's correct central altitude† was 20° 53' bearing south. Required the latitude?

Latitude

Time per watch 0 59 0

15' long. in time	1	0 Declina	ation	0	Co-s	ine	10.00000
Time from noon	1 0	Time from	noon 11	ı. Oın.	Log.	rising	3.53243
,	Central	- I altitude····28°	53′	Nat. Numb. Sine		Log.	3.231 10
		enith distance60	0 N.	Co-sine	50007		

Latitude60

When the observation is taken a few minutes before or afternoon, the correction to be applied to the altitude, to obtain the meridian altitude, may be found by means of Tables XXXII. XXXIII. the first of which contains the variation of the altitude for one minute from noon, expressed in seconds and tenths—the other contains the square of the minutes and seconds of a minute contained in the top and the side columns. By these tables the correction of the observed altitude may be found by the following rule.

Parallax too small to be noticed.

† The observed altitude of the sun's lower limb being 292 43', TSS. D. 16'. Dip 4', Refraction 2'. Parallax too small to be noticed.

9.69997

Co-sine

^{*} The observed altitude of the lower limb being 190 32', G's semi-diameter 16', Dip 4'. Refruction S', Peraliax too small to be noticed.

RULE.

Enter Table XXXII. and find the latitude by account in the side column, and the declination at the top, opposite the former and under the latter will be the change of altitude in seconds and tenths for one minute from noon: then enter Table XXXIII. and find the minutes of the time from noon in the top column, and the seconds in the side column, under the former, and opposite the latter, will be a number which is to be multiplied by the number taken from Table XXXII. and the product will be the sought change of altitude, expressed in seconds and decimals.

In making use of Table XXXII. proportional parts may, if necessary, be taken for the miles of latitude and declination. The numbers in both these

tables are expressed in whole numbers and tenths.

EXAMPLE III.

Being at sea in the latitude of 40° N. when the sun's declination was 21° N. at 8' past noon the sun's correct central altitude was 70° 58'. Required

the meridian altitude and latitude?

In Table XXXII. opposite 40° lat. and under 21° dec. is 4'.3, and the number in Table XXXIII. corresponding to 8' is 64.0. By multiplying 64.0 by 4".3, the correction 275".2 (or 5' nearly) will be obtained; this quantity added to 70° 58' will give the meridian altitude 71° 3', and the latitude deduced therefrom will be 39° 57' N.

By observing several altitudes of the sun when near the meridian, and noting the times, the meridian altitude may be obtained, by the above method, to a great degree of accuracy; for by using this method, many observations may be taken on the same day, and the mean of the meridian altitudes deduced therefrom will in general be much more correct than that obtained by a single observation, by the usual method. To obtain the correction to be applied to the mean of all the observed altitudes, proceed thus:

Take from Table XXXIII. the number corresponding to each time from noon (the minutes being found at the top and the seconds at the side, the correction being under the former and opposite the latter) and divide the sum of these tabular numbers by the number of observations, the quotient being multiplied by the number taken from Table XXXII. will be the correction to be applied to the mean of the observed altitudes, to obtain the meridian altitude.

EXAMPLE IV.

. Being at sea in the latitude of 50° N. by account, when the sun's declination was 220 N. observed with a sextant, the altitudes of the sun's lower limb (bearing nearly south) as in the following table: the correction for semidiameter, dip, refraction, &c. being 12' additive. Required the meridian altitude and latitude?

windled the intitute.						
Obs. Alt.	Time from	Numbers Tab.				
61.45 61.46 61.46 61.47	4 15 3 2	38 0 18 1 9 2 4 7				
Sum 247.04		70 0				
Mean 61.46		17 5				

The mean of the numbers from Table XXXIII. is 17.5, this multiplied by the number of seconds from Table XXXII. viz. 2".5, gives the correction 43".75, or 44", which added to the mean of the observed altitudes 61° 46' gives the meridian altitude of the sun's lower limb 61° 46' 44" or 61° 47' nearly, to this add 12' for semi-diameter, &c. the sum 61° 59' will be the correct central meridian altitude, whence the latitude was 50° 1' N.

If the above altitudes had been taken with a circle, the calculation would have been exactly the same, except that each altitude would not have been given, but the sum of all of them 247° 4' would have been shown by the central index after finishing the observations.

^{*} The observed altitude of the sun's lower limb being 70° 46', Semi-diameter 16', Dip 4', Parallax first Refraction too small to be noticed:

EXAMPLE V.

Having regulated my watch, I found it to be 6'2" too slow for apparent time. I then sailed to the southward and eastward till the ship had made 60' difference of longitude, and was by account in the latitude of 40° N. the sun's declination being 20° S. The sun being then nearly on the meridian I observed ten altitudes of his lower limb by a circle of reflection and noted the times by the watch as in the following table, and the sum of all the altitudes taken from the circle was 295° 20'.—Required the true latitude, supposing the dip to be 4' and the semi-diameter 16'?

When it was 12 o'clock by the watch it was 12h. 6m. 2s. apparent time at the place where the watch was regulated, and 12h. 10m. 2s. apparent time at the place where the altitudes were taken to determine the latitude, because the former place was 60' or 4' in

time to the westward of the latter, consequently the watch was 10m. 2s. too slow for app. time at the place of taking the altitudes for determining the latitude. Hence we may determine the time from noon of taking each observation, as in the second column of the adjoined Table, and find the numbers corresponding in Table XXXIII. the mean of which is 6.97, this multiplied by the number in Table XXXIII. corresponding to the latitude 40° N. and declination 20° S. viz. 1".6 will give 11".152 or 11", which is the correction to be added to the mean of the observed altitudes to obtain the meridian altitude.

Time per	Time from	Numbers Tab. XXXIII.
1.45.43	4' 15"	18.1
46.58	3 0	9.0
47.52	26	4.4
48.50	18	1.3
49.28	0 30	0.2
50.48	0 50	0.7
51.10	1 12	1.4
52.13	2 15	5.1
53. 8	3 10	10.0
54.23	4 25	19.5
	Sı	ım 69.7
	N	lean 6.97

Now the sum of all the altitudes 298° 20', divided by 10, the	• • •	•	,
number of observations gives	50'	0	′
Add semi-diameter 16' and the above correction 11''+	16	11	
number of observations gives		8	
Subtract dip 4' and refraction 1' 39"	5	39	
Central Altitude30		40	
Zenith distance	59	20	N.
Declination20	0	0	5.
Latitude39	59	20	N.

When the meridian altitude of the object is small, the correction of altitude may be found by this method, for 12 or 15 minutes from noon, to a great degree of accuracy; but when the sun passes near the zenith, the time of observation must be proportionally nearer to noon. Thus in Example I. preceding, the time from noon was 32', and as the numbers in Table XXXIII. are the squares of the number of minutes, it follows, that the number corresponding to 32' would be the square of 32 or 1024.0. This multiplied by the number 1".3 of Table XXXII. corresponding to the latitude 50° N. and declination 20° S. will give the correction 1331".2 or nearly 22', which added to 15° 41' will give 20° 3' for the meridian altitude, or 69° 57' for the zenith distance, being the same as in that example.

It is very advantageous in this method to observe as many altitudes in the afternoon as before noon, and at nearly the same distances from noon, for in this case a small error in the regulating of the watch will not materially affect the calculation. This will appear evident by supposing, in the preceding example, that the watch was 11'2' too slow, instead of 10'2", by which means the times and numbers will be as in the adjoined Table, and the mean of all the numbers taken from Table XXXIII. will be 8.15, which multiplied by 1".6 will give 13" nearly, for the correction instead of 11", so that in this case an error of one minute in the regulation of the watch would only cause an error of 2 seconds in the meridian altitude.

But it must be carefully observed that in using this method you must not take the observation more than 2 or 3 minutes from noon when the sun passes within 10° or 12° of the zenith.

Times.	XXXIII.
3.15	10.6
2. 0	4.0
1. 6	1.2
0.8	0.0
0.30	0.2
1.50	3.4
2.12	4.8
3.15	10.6
4.10	17.4
5.25	29.3
Sun	n 81.5
Mean 8.15	

In Tab.

TO DETERMINE THE LATITUDE ON SHORE

BY MEANS OF AN ARTIFICIAL HORIZON.

IT frequently happens that the latitude of a place on shore cannot be determined by the usual methods, by a quadrant, sextant or circle, on account of not having an open In this case it is customary to make use of an artificial horizon formed by the surface of a vessel filled with water, mercury, Barbadoes tar, very clear molasses, or any other fluid of sufficient consistency not to be affected by the wind.* With this apparatus an observation may be taken on shore when the altitude of the object does not exceed 60°, with as much ease as at sea. Thus, if an altitude of the sun was required to be taken, the observer must place the vessel containing the mercury (or other fluid) in a firm position on the ground, and in a few minutes the surface of the liquor will attain a horizontal situation, the observer must then place himself in a situation so as to see the image of the sun formed by the fluid, which image will evidently be depressed as much below the horizon as the sun is elevated above, so that to obtain the double of the sun's altitude it is only necessary for the observer to bring the image of the sun formed by the instrument, down to the image formed by the artificial horizon, and the angle then pointed out by the index will be double of the altitude of the sun, the half of which will be the apparent If the nearest limbs of the two images are brought in contact, the half of the angle obtained by the instrument will be the altitude of the sun's lower limb, but if the farthest limbs are brought in contact, the half angle will be the altitude of the upper limb. The altitude thus obtained must be corrected for semi-diameter, parallax, and refraction, as usual, but not for dip, because a truly horizontal surface is obtained by means of the artificial horizon. In this manner the altitude of the sun, or any other bright object, may be obtained when the altitude is less than 60°; at higher altitudes the angle corresponding would be above 120°, which cannot be measured by a sextant on account of the length of the arch, nor by any other instrument of reflection, with a sufficient degree of accuracy. To illustrate this method, we shall here add the following examples.

EXAMPLE I.

EXAMPLE II.

The angular distance of the nearest limbs of the two images of the sun was found by the above me-liwe images of the sun, when on the meridian, was thou to be 689 10', when the declination was 10° S-lobtained by the above method, and found to be 34° and the sun's semi-diameter 16', the sun bearing south 0', when the declination was 10° N, and the semidiameter 16'; the sun bearing north of the observer-Required the latitude? of the observer. Required the latitude !

Half of 689 10' is the obs. alt Add semi-diameter		5′ 16
Subtract regraction	34	21
True altitude	34	20
Zenith distance		
Tatitude	45	40 N

of 34° 0' is the obs. alt Subtract semi-diameter		16 0'
Refraction sub	16	44 3
True altitude	16	41
Zenith distance		
Latitude	63	19 8.

The latitude may be determined on shore by this method to a great degree of accuracy by means of a circle of reflection, by taking several altitudes a few minutes before and after the sun passes the meridian, and estimating the correction to be applied to the altitude by means of Tables XXXII. and XXXIII. Thus, if in the example, page 151, the observations had been taken in this manner, the number of degrees denoted by the circle after taking ten observations, would have been 595° 20', this divided by 20, (twice the number of observations,) will give for the observed altitude 29° 46', and by adding the semi-diameter 16', parallax 8', and the correction found by Tables XXXII. and XXXIII. viz. 11 seconds, and subtracting the refraction 1' 39", the central altitude will be obtained 30° 0' 40" as in the page before mentioned.

molasses before using it.

If the instrument has an index error, it must be applied to the observed angle, or the half of the index error must be applied to the sun's altitude.

[•] In case the wind blows fresh, you must use a screen formed of two planes of talc or glass whose surfaces are ground perfectly parallel, and connected together in a frame so as to make an angle of about 30° with each other. This frame is to be placed over the box containing the fluid, and the rays of the sun passing through one of the plates, is reflected from the surface of the liquor, and passes through the other plate to the eye of the observer. The use of these plates is to be avoided, when it can possibly be done, on account of the defect of parallelism of the surfaces. This error is generally greats mear the border or the glass, so that it has been recommended to cover the edge of the glass with a paper or some paint, to the distance of \(\frac{1}{2}\) of \(\frac{1}{2}\) inch from the frame. If the surfaces of the glass are perfectly parallel, the observed angle will be the same as if the screen had not been used. If the reflecting fluid is molasses, air bubbles will sometimes class on the surface by the suris best this may in tome measure be added by besting the setting the series of the plane is the series of the series of the series of the plane and the series of will sometimes rise on the surface by the sun's heat, this may in some measure be avoided by heating the

Altitudes may be observed in this manner in taking an azimuth for determining the variation, or for regulating a watch in the manner explained in this work.--Observing in all cases, that the half of the observed angle is to be corrected for refraction, parallax, and semi-diameter, but not for the dip of the horizon; and that half the index error only is to be applied.

TO FIND THE LATITUDE BY AN ALTITUDE OF THE POLAR STARA

IN northern climates the latitude may be determined by means of an observed altitude of the Polar Star, provided the apparent time of observation can be ascertained within a few minutes.* This method might be frequently used at sea, when the horizon is well defined, if that star were of the first magnitude, but being only of the 2d. or 3d. magnitude, it is sometimes so dim that it is rather difficult to determine the altitude with precision. However, as there are times when it would be of great importance to determine the latitude within 10 or 12 miles. it was thought adviseable to explain this method, which may Busedwhen observations of the sun or moon cannotbe obtained.

Having therefore the apparent time of observation (which must be reckoned from noon to noon in numerical succession, that is 6h. A. M. must be called 18h. &c.) and the observed altitude of the star determined by a fore-observation, you must subtract from the altitude the dip, which is in general 4 minutes, and the refraction, and you will obtain the true alti-tude of the star. Then the sun's right ascension corresponding to the given day, must be found in Table VI.† and added to the apparent time of observation (rejecting 24 hours when the sum exceeds 24 hours) with that sum enter the adjoined table and take out the corresponding correction, which must be added to, or subtracted from the true altitude, according to the directions in the

Find in the side column the sum of the hour of the day nd the sun's right ascension, and the number in the middle column will be the correction of the true altitude. If the hour is found in Correction II the hour is tound these columns, the correction is additive. correction is additive. h. m. h. m. h. m. h. m. 0. 58 12. 58 0. 58 12. 58 1. 38 1. 3 0. 53 1. 38 12. 53 13. 1. 38 13. 13 1. 13 0. 43 12, 43 1. 23 12. 33 13. 23 0. 33 1. 38 1. 33 0. 23 1. 37 12. 23 13. 33 1. 43 0. 13 1. 36 12. 13 13. 43 53 0. 3 12. 13. 53 1. 1. 35 11. 2. 23. 3 53 53 1: 34 14. 2. 13 23. 43 11. 43 14. 13 1. 33 23. 33 11. 33 2. 23 14. 23 1. 31 2. 33 23. 23 11. 23 14. 33 1. 30 2. 23. 13 43 1. 28 11: 13 14. 43 23. 2. 53 1. 26 11. 3 14. 53 3 22. 53 24 10. 53 1. 15. 3. 22. 43 3. 13 1. 22 10. 43 15. 13 22. 33 3. 23 1. 19 10. 33 15. 23 22. 23 3. 33 10, 23 1. 16 15. 33 3. 10. 13 43 22. 13 1. 14 15. 43 3. 53 22. 10. 3 15. 53 1. 11 4. 3 21. 53 53 16. 1. 8 9. 43 4. 13 21. 43 16, 13 1. 5 4. 23 9. 33 21, 33 16. 23 1. 1 4. 33 21. 23 9. 23 16. 33 0. 58 9. 13 4. 43 21. 13 0. 54 16. 43 4. 53 21. 0. 51 9. 16. 53 8. 5. 3 20. 53 0. 47 53 17. 5. 13 20. 43 0. 43 8. 43 17. 13 5. 23 20. 33 0.39 8. 33 17. 23 33 20, 23 8. 23 17. 33 5. 0. 36 20. 13 8. 5. 43 0. 32 13 17. 43 8. 3 17. 53 5. 53 20. 3 0. 27 7. 6. 0. 23 53 18. 3 19. 53 6. 13 19. 43 0. 19 7. 43 18. 13 6. 23 19. 33 0. 15 7. 33 18. 23 6. 33 19. 23 0. 11 7. 23 18. 33 6. 43 7. 13 19. 13 0. 6 18. 43

2 7.

0

18. 53

18. 59

58

6.

0.

0.

table—the sum or difference will be the latitude of the place of observation.

6. 53

58 18.

19.

58

fif the star is not far from the meridian, an error of half an hour in the time would not affect the Satitude above 1 or 2 miles

fatitude above 1 or 2 miles.

It is accurate enough to take the numbers given in the table; but in strictness the right ascemsion ought to be taken from the Nautical Almanac, for the hoar of observation, reduced to Greanwich time, by adding or subtracting the longitude turned into times. In some of the copies of the second edition of this work, the words additive and subtractive were print ed in the wrong columns. This table will require a correction after a few years, on account of the variation of declination, and right ascension of the star. It corresponds nearly to the year 1824, for every year after that time you must add one quarter of a minute to the times in the slde columns, and decrease the tabular corrections of altitude about \(\frac{1}{100}\) part. Thus for the year 1836 the times must be increased. 3m. for the 12 years, so that 0h. 58' must be called 1h. 1m. and all the corrections of altitude mug be decreased \(\frac{1}{12}\) part, so that 1° 16' must be 1° 18' nearly and 0° 38' must be 0° 38' must be 0° 38' must be 0° 38' must be 0° 38' nearly.

154 TO FIND THE LATITUDE BY AN ALTITUDE OF, &c.

EXAMPLE I.

At 6h. 9m. P. M. June 3, 192 the dip 4'. Required the latit add			rved altitude of the Polar Star was	16°	10
	4h	. 44m.	Observed altitude16° Sub. Dip 4', Refrac. 3'		
Sum	10	53	True altitude		
•			Latitude	27	N.
	E	XAMI	PLE II.	•	

	otar '	Was	24.
titude ?			
Observed altitude	240	16'	
Dip 4', Ref. 2', Sub		6	
- '			
True altitude	24	10	
Corresponding corr. sub	1	37	•
	titude ? Observed altitude Dip 4', Ref. 3', Sub True altitude	titude ? 24° Observed altitude	Observed altitude 24° 16′ Dip 4′, Ref. 2′, Sub 6 True altitude 24 10 Corresponding corr. sub 1 37

TO FIND THE TIME AT SEA, AND REGULATE A WATCH.

WE have already noticed the difference between the civil, astronomical and nautical computation of time; but as it is a subject of great importance, it may not be unnecessary again to repeat, that a civil day is reckoned from midnight to midnight, and is divided into 24 hours; the first 12 hours are marked A. M. the latter 12 hours P. M. being reckoned from midnight in numeral succession from 1 to 12, then beginning again at 1 and ending at 12. Astronomers begin their computation at the noon of the civil day, and count the hours in numeral succession from 1 to 24, so that the morning hours are reckoned from 12 to 24. Navigators begin their computation at noon, 12 hours before the commencement of the civil day (and 24 hours before the commencement of the astronomical day) marking their hours from 1 to 12 P. M. and A. M. as in the civil computation.

There are two kinds of time, mean and apparent. Mean time is that shewn by a clock regulated to mean solar time. Apparent time is that shewn by the sun, estimating the apparent noon to commence at the passage of his centre over the meridian of any place. There is sometimes a difference of a quarter of an hour between mean and apparent time, owing to the unequal motion of the earth in its orbit, and the inclination of its axis. This difference is called the equation of time, and is contained Table IV. A. in page to the Nautical Almanac: It is necessary to take notice of the equation of time in determining the longitude by a time-keeper or by the eclipses of Jupiter's Satellites; but it is not necessary in any other nautical observation, because the calculations of the Nautical Almanac, except the times of the eclipses of Jupiter's Satellites, are adapted to apparent time.

We may obtain the apparent time at sea, when the ship makes no way through the water, by observing an altitude of the sun in the morning, and again in the afternoon, when at the same altitude, and noting the times by a watch; for the middle time between these two observations will be the apparent time of the sun's passage by the meridian; hence the error of the

^{*} There is also another method of computing the time, made use of by astronomers, called Siderian's time, in which the interval between two successive transits of a fixed star to the meridian is estimated as 24 hours, commencing the day at the time the first point of Aries is on the meridian, so that the hour tan affectial time is the same as the right aspersalon of the meridian.

watch may be found. A small correction is necessary for the variation of the sun's declination during the interval between the observations, and the method of calculating this correction will be given in this work, but this method cannot often be made use of at sea by reason of the motion of the vessel.

The best method of obtaining the apparent time at sea, is by observing, by a fore observation, the altitude of the sun's lower limb when rising or falling fastest, or when bearing nearly E. or W. to this altitude we must add the semi-diameter and parallax, and subtract the dip (or instead of these three corrections add 12,° which will answer very well for an observation taken on the deck of a common sized vessel;) subtract also the refraction taken from Table XII. and the remainder will be the correct altitude. The ship's latitude must be found at the time of observation by carrying the reckoning forward to that time. The declination must be taken from Table IV. or from the Nautical Almanac, and corrected for the ship's longitude, and the distance of the sun from the meridian by Table V. Then if the latitude and declination be both north or both south, subtract the declination from 90° and you will have the polar distance; but if one be north and the other south, add the declination to 90° and you will have the polar distance.

Having thus found the correct altitude, latitude, and polar distance, the apparent time of observation may be found by either of the three following methods, of which the first is the most simple, since it does not require the table of natural sines, all the logarithms being found in Table XXVII. This method is abridged by means of the table of hours affixed to the table of log. sines; in using which you must observe, that if the sine or co-sine of the logarithm sought is marked at the top of the table, the name of hour either A. M. or P. M. is also to be found at the top, and the contrary if the sine or co-sine is marked at the bottom.

To find the apparent time by the sun's altitude.

FIRST METHOD.

Add together the correct altitude of the sun's centre, the latitude and the polar distance; from the half sum subtract the sun's altitude, and note the remainder. Then add together the log. secant of the latitude (this and all the other logs being found in Table XXVII.) the log. co-secant of the polar dist. (rejecting 10 in each index) the log. co-sine of the half sum, and the log. sine of the remainder, half the sum of these four logarithms being sought for in the column of log. sines, will correspond to the hour of the day in one of the hour columns.

EXAMPLE I.

Suppose that on the 10th of October, 1824, sea account, at 8h. 21m. A. M. per watch, in the latitude of 51° 30' N. and long. 114° E. from Greenwich by account, the altitude of the sun's lower limb by a fore-observation was 13° 32', the correction for semi-diameter, parallax and dip 12'—Required the apparent time of observation?

By example III. page 111, the declination was 6° 34' S. this added to 90° gives the polar distance 96° 34'. To the sun's observed altitude 13° 52', I add 12 miles and subtract the refraction 4', the remainder is the correct altitude 13° 40'.

[→] The semi-diameter is in general about 16', the parallax never exceeds 9', and the dip is about 4'; and as the two former corrections are additive and the latter subtractive, the effect of all three corrections are subtractive, the effect of all three corrections will not differ materially from 12' additive.

156 TO FIND THE TIME AT SEA, AND REGULATE A WATCH!

©'s cor. alt 13° Lat 51 Polar dist 96	30	secant 0.20585 co-secant 0.00286
Sura161	44	
Half Sum 80 ⊙'s Alt 13	52 40	co-sine 9.20067
Rem 67	12	sine 9.96467
		2)19.37405

Watch too fast

13 51

Hence the time of taking this observation is Oct. 10, 8h. 7m. 9s. A. M. sea account, or which is the same thing, Oct. 10, 20h. 7m. 9s. reckoning from noon to noon; the time by the Nautical Almanac being October 9d. 20h. 7m. 9s.

EXAMPLE II.

Suppose that on the 10th of May, 1824, sea account, at 5h. 30m. P. M. per watch, in lat. 39° 54′ N. long. by account 17° 50′ E. from Greenwich, the altitude of the sun's lower limb by a fore-observation was 15° 45′, the correction for dip, parallax and semi-diameter being 12 miles, consequently the correct altitude 15° 54′—Required the apparent time of the observation?

By Example IV. page 111, the sun's declination was 17° 28' N. which subtracted from 90° leaves the polar distance 72° 52'.

⊙'a Alt.... 15° 54' Lat..... 39 secant 0.11511 54 Pol. dist 72 32 co-secant 0.02050 Sum 129 20 4 Sum 64 10 co-sine ... 9.63924 O's Alt. 15 54 Remainder 48 16 sine 9.87288

sine 9.82386 corresponding to which in the column P. M. is 5h. 34m. 26s. the true time of day.

Time per watch 5 30

2)19.64773

Watch too slow 4 26

EXAMPLES TO EXERCISE THE LEARNER.

1. In lat. 36° 39°S. ⊙'s declination 9° 27' N. the altitude of the ⊙'s lower limb in the morning was observed 10° 33'.* Required the apparent time?

Answer, 7h. 23m. 51s. 2. In lat. 36° 21' S. ②'s declination 9° 44' N. art. ⊙'s L. L. in morning 10° 48'.*

Required the apparent time?

Answer, 7h. 22m. 11s.

3. In lut. 29° 25' N. ©'s declination 23° 20' N. observed alt. of ©'s flower limb in

3. In lat. 29° 25 N. ②'s declination 23° 27 N. observed alt. of ②'s tower limb in the afternoon 14° 58'.* Required the time?
Answer, 5h. 41m.
4. In lat. 3° 31' S. ③'s declination 20° 3' S. observed alt. ⊙'s L. L. 38° 41* in the

afternoon. Required the time?

Answer, 3h. 18m. 47s.

5. In lat. 13°, 17' N. @'s declination 22° 10' S. in the morning observed alt. of @'s

L. L. 36° 26.* Required the time?

Answer, 9h. 17m. 8s.

6. In lat. 21° 36' S. ©'s declination 3° 37' S. in the morning observed altitude of

O's L. L. 35° 48'.* Required the time?

SECOND METHOD.

Find as in the former method, the sun's correct altitude, the ship's latitude, and the polar distance; thence the sun's correct zenith distance and the The correction for dip and semi-diameter being 12 additive, the correction for reduction is also to be applied as tatal.

complement of latitude; then add together the zenith distance, co-latitude, and polar distance, from half their sum subtract the zenith distance, and note the remainder; then add together the log. co-secant of the co-lat. (this and all the other logs, being found in Table XXVII.) the foliar co-cant of the polar distance (rejecting 10 in each index) the sine of the half sum and the sine of the remainder, half the sum of these four logarithms being found among the log, co-sines, will correspond in one of the adjoined columns to the time of day.

The two preceding examples are thus worked by this method.

	EXAMP	LE I.			
900 0		•	900 0'		900 ø/
©'s cor. alt. 13 40		latitude	51 30	⊙'s dec.	6 34
Zen. dist 76 20 Co-lat 38 30 Pol. dist 96 34	co-secant 0.20585 co-secant 0.00286	co-lat.	38 30	Pol. dist.	96 34
Sum 211 24					
J Sum 105 42 Zen. dist 76 20	sine 9.98349				
Rèm 29 22	sine 9.69055				
ı	2)19.89275			;	r F
A. M. is Sh. 7m. 9s. th		correspondingrees with th			column

		EXAMPL	E II.		
90°.0′ ⊙ 's cor. alt. 15 54			latitude	90°	_
Zen. dist 74	co-sec.	0.11511 0.02050	co-lat.	50	6
Supp 196 44					
4 Sam 98 22 Zen. dist 74 6	sine	9.99535	•		
Remainder 24 16	sine	9.61392			
	2)19.74478			

co-sine 9.87239 corresponding to which in the column P. M. is 5h. 34m. 27s. the time of day, which agrees nearly with the first method.

By the preceding method you may find the beginning or ending of the

twilight, by calculating the hour when the sun's zenith distance is 1080 (or when the sun is 180 below the horizon) for by observation it has been found. that the twilight begins or ends when the sun is at that distance from the zenith.

EXAMPLE.

Required the time of beginning and ending of the twilight, June 23, 1820, at Buston?

Co-lat 47 37	co-secant	0.13156
Pol. dist 66 33	co-secant	0.03744
Sum 222 10		
3 Sum 111 5	sine	9.96991
Zen. dist 108 0 Remainder 3 5	sine	. 8.73069
210220201101	sum	

O's dec. Pol. dist.

72 32

sum co-sine 9.43480 which converseponds to 2h. 6m. 20s. A. M. and 9h. 53m. 40s. P. M. Therefore, the first appearance of the twilight in the morning was at 2h. 6m. 20s. and the end of it in the evening at 9h.

158 TO FIND THE TIME AT SEA, AND REGULATE A WATCH.

THIRD METHOD.

If the sun or star's declination, and the latitude be both north or both south, take their difference, but if one be north and the other south, take their sum, and from the natural co-sine of this difference or sum subtract the natural sine of the true altitude (both being found in Table XXIV.) find the log. of their difference (in Table XXVI.) and thereto the log. secant of the latitude (from Table XXVII.) and the log. secant of the sun or star's declination (from the same Table) rejecting 10 in each index; the sum of these three logarithms being found in the column of rising (Table XXIII.) the hours, minutes, and seconds corresponding, will be the apparent time from noon, if an altitude of the sun was taken: but if a star was observed, the corresponding time will be the horary distance of the star from the meridian.

The two preceding examples are thus worked by this third method.

Latitude	EXAMPLE I. secant 0.20585 secant 0.00238
Sum 58 4 Sun's cerr. aktitude 13 40	Nat. co-sine
	Differ 29268 log. 4.46638
which in the column of rising is .	4.67507 corresponding to 3h. 82m. 51s.
Subfracted from 12h. leaves the true	time 8 7 9 Time per watch 8 21
mer methods.	Watch too fast 13 51 agreeing with the for-
mer metuous.	EXAMPLE II.
Latitude 89° 54′ N. Declination 17 28 N.	secant 0.11511 aecans 0.02050
Difference 22 26 Corr altitude 15 54	Nat. co-sinc 92432 Nat. sinc 27396
l	Difference 65086 log. 4.81315
which in the column rising is 51 Time per watch 5	34m. 27a. agreeing nearly with the other methods.
Watch too slow	4 27

To find the apparent time by an altitude of a fixed star.

Correct the observed altitude for the dip and refraction (the dip being generally 4 minutes when the observation is taken on the deck of a common sized vessel) find the ship's latitude at the time of observation, and the star's right ascension and declination in Table VIII.* then add together the star's correct altitude, the ship's latitude, and the polar distance; from the half sum subtract the star's altitude, and note the remainder. Then add together the log. secant of the latitude, the log. co-secant of the polar distance, rejecting 10 in each index, the log. co-sine of the half sum, and the log. sine of the remainder, half the sum of these four logarithms will be the log. sine of half the hour angle; take out the corresponding time in the column marked P. M. (Table XXVII.) and apply it to the star's right ascension, by subtracting when the star is east of the meridian, or adding when west of the meridian, the sum or difference will be the right ascension of the meridian. From the right ascension of the meridian (increased by 24 hours if necessary) subtract the sun's right ascension at noon at Greenwich, taken from

The right ascensions and declinations of the stars in Table VIII. are the mean values for January 1st, 1820, and must be reduced to the time of observation by means of the annual variation given in the same Table. When very great accuracy is required, the right ascension and declination thus obtained, must be corrected for the Aberration and Nutation as explained in the precepts of Tables XLIII. XLIII.; but in general these corrections may be neglected. These corrections are however all noticed in the places of 24 of the most noted fixed stars, given in the Nautical Almanac since the year 1822, for every ten days in the year, and when any of these stars are used, the places must be taken out, to the nearest day, from the Nautical Almanac, without any further correction. Thus in July 29, 1824, Precyon's right ascension was 7h. 50m. 7s. north polar distance 84° 18° 52" or 5° 40° 8" N. declination, corresponding to 35° 40° 8" south polar distance.

This adultional Table of the Nautical Almanac simplifies this kind of calculation considerably.

page 2d. of the month in the Nautical Almanac, the remainder will be the approximate time at the ship. To this time apply the longitude of the ship from Greenwich, turned into time, by adding in west longitude or subtracting in east longitude, the sum or difference will be the apparent time of the observation nearly, at the meridian of Greenwich.* Take from the Nautical Almanac the sun's right ascension for the preceding and following noons, and take the difference, which seek for at the top of Table XXXI. f and the hours and minutes of the time at Greenwich in the side column, the corresponding correction being subtracted from the approximate time at the ship. will give the apparent time required.

EXAMPLE I.

Ry Table VIII, for the year 1890.

Suppose that on September 9, 1820, sea account, in latitude 7° 45' S. and longitude 30° 18' E. from Greenwich, the altitude of the star Procyon, being then east of the meridian, was 28° 16' and the dip 4'. Required the time of observation?

Procyon's right ascension Variation in 82 months		Dec. Var.	5 °	41' 0	Ŋ.
Star's right ascension	7 29 54	Dec.	5 90	41	Ŋ.
2		Pol. dist	. 95	41	
Ref. Table XII 2					
Correct altitude 28 10					
Latitude 7 45	secant 0.00399 -				
Polar distance 95 41	co-sec 0.00214				•
Strin 2)131 36					
4 Sum 65 48 Alt 28 10	co-sine 9.61270				
Remainder 37 38	sine 9.78576				
	sum 2)19.40459				
	d sum sine 9.70229 correspo	nding t	o wh	ich.	ίπ
the column P. M. is			4h.	m.	
	Star's right ascension	• • • • • • • • •	7 2	9 8	14
	Right ascension of the meric Increased by		3 2°	7 :	52
•		2	7 2	7 1	 i2
Sent. 9. sen account, is September 8, b	v N. A. sun's R. A. moon			-	20
,	_	-	 -		-
	Approximate time at the ship Ship's longitude 30° 18' in the			_	32 12:
Sam's right ascension September 8, September 9,	Time at Greenwich nearly . 11h. 7m. 20s. 11 10 56	1	4 1	9 \$	20

36 The correction of Table XXXI. corresponding, is 2m. 2s. which being subtracted from the approximate time 16h. 20' 32" leaves the apparent time at the ship 16h. 18' 23" or 4h. 18' 23" A. M.;

the sum will be the sought correction.

† If the ship was furnished with a Chronometer regulated to Greenwich time, the 6's Right Ascendion might have been found at once from the Nautical Almanac to be the sum of 11h. 7m. 20s. and 2m. 9s. equal to 11h. 9m. 29s. which subtracted from 2th. 27m. 52s. gives the apparent time at the ship 18h. 18m. 2is. more simply than by the above method. The same remark will apply to other examples.

When the sum exceeds 24 hours, you must subtract 24 hours and add 1 to the day of the mouth, and then the hours to be subtracted are more than the hours of the time at the ship, you must add 24 hours of the inter, previous to the subtraction, and take 1 day from the day of the mouth.

† Table XXI. is only calculated to 12 hours. If the time at Greenwich exceeds 12 hours, you must take out the correction for 12 hours, and add it to the correction taken out for the rest of the time;

160 TQ FIND THE TIME AT SEA, AND REGULATE A WATCH:

EXAMPLE II.

Suppose that on April 16, 1820, sea account, in lat. 48° 57' N. and long. 66° W. the observed altitude of Aldebaran when west of the meridian was 22° 25', and the dip 4'. Required the apparent time at the ship?

Variation for 3d months Star's righ					
	t asc. 4 25 37	Dec.	16	8 1	N.
Obs. alt. Aldeb 22° 25'	11 400. 1 20 01		90	٠.	• • •
Dip 4	,	-		_	
		P. dist.	73	52	
22 21 Refraction 2					
Refraction 2					
Cor. alt. Aldeb 22 19					
	0.1826Z				
73 .52 co-sec	0.01745				
2)145 8					
72 34 co-sine ·	9.47654				
22 19					
Remainder 50 15 sine	9.88584				
sum	. 2)19.5 624 5				
••	0.50100				_
	e 9.78122 com				
	t ascension			3	
Right asce April 16, sea account, is April 15, by N. A. when (n. of the meridis D's rt. asc. noon		23 34	1	1 2
	ate time at ship W. in time			4	9
App. time Sim's right ascen. April 15, 1h. 34m. 12s. April 16, 1 37 54	at Greenwich	13	12	4	*

ponding, is 1m. 53s. which being subtracted from the approximate time at the ship,

7h. 48' 49'

leaves the apparent time at the ship 7 46 56 P. M.

This method of obtaining the time by the stars would be accurate if a good horizon could be obtained; but as that is not always the case, it is best to regulate your watch by the sun.

To regulate a watch by equal altitudes of the sun.

A watch may be regulated on shore by observing in the morning and evening the times when the sun is at the same altitude, * for the middle between these times would be the apparent time of noon by the watch if the declination of the sun remained the same during the observation; but if the declination varies, as is generally the case, the apparent time of noon determined in this manner (which for distinction we shall call the middle time) must be corrected for the change of declination by an equation, called the equation of equal altitudes, and the middle time thus corrected will be the

^{*} The altitudes should be taken when the sun rises or falls fast. The best time for observation is when the bearing of the sun is east or west. In general two or three hours from noon will be sufficient. An artificial horizon, formed by a vessel filled with mercury, may be used in taking these altitudes.

TO FIND THE TIME AT SEA, AND REGULATE A WATCH. 161

correct time of apparent noon by the watch. For greater accuracy, several altitudes should be taken in the morning, and corresponding ones in the afternoon, and the mean of the times of the morning and evening observations should be respectively taken, and the equation of equal altitudes corresponding to the mean of all the observations calculated and applied to the middle time, as if a single set of observations only had been taken.

In noting the times of observation, you must count the hours in numeral succession, so that if some of the observations are taken before 12h. by the. watch, and others after 12h, the next hour to 12h, must be called 13h, the next 14h. &c. Half the sum of the times of observation, corresponding to any set of observations (or the mean of a number of observations) will be the middle time, and the difference of the times of observation will be the elapsed time.

The equation of equal altitudes consists of two parts, which may be cal-

culated by the following rule:

RULE 1.—To the constant log. 8.9239 add the log. co-tangent of the latitude, the log. sine corresponding to the elapsed time found in the column P. M. of Table XXVII. the prop. log. of the hours and minutes of the elapsed time, reckoned as minutes and seconds, and the prop. log. of the daily variation of the sun's declination, the sum, rejecting 30 in the index, will be the prop. log. of the first part of the equation of equal aftitudes, reckoning minutes and seconds as seconds and thirds respectively.

2. To the constant log. S.8239 add the log. co-tangent of the sun's declination, the

log. tangent corresponding to the elapsed time found in the column P. M. of Table XXVII. the prop. log. of the hours and minutes of the elapsed time reckaned as minutes and seconds, and the prop. log. of the daily variation of the sun's declination, the sun, rejecting 30 in the index, will be the prop. log. of the second part of the equation of equal altitudes, reckoning minutes and seconds as seconds and thirds respectively.

The first part of the equation of equal altitudes is to be added to the middle time when the sun is receding from the elevated pole, otherwise subtracted; * and the second part is to be added when the declination is increasing, but subtracted when decreasing; these two corrections, being applied to the middle time, will give the apparent time of noon by the watch.

EXAMPLE.

Suppose that on the 9th. of May, 1820, civil account, in the latitude of 40° N. and long. 10° W. the following observations were taken at equal altitudes of the sun. Reonired the error of the watch?

Alt. O's L. L.	, watch ,	Ti		er wi M.	atch.			s per P. M.	watch
15° 35′		6b	. 29m	. 51s			17h.	32m.	184.
15 45		6					17	31	0
15 55		6	32	14			17	29	54
	Sum	_	93	12				93	12
	Mean	6	31	4			17	31	4
							6	31	4
	Differe	nce	is ela	apscd	time		11	0	0
						•••••		2	8
	Middle	tin	ne				. 12	1	4
Constant log								8.5	239
Lat. 40° co-tang					dec. 1	7.25 col	t	10.5	035
Elap. time 11h. Sine						•••••			
Elap. time 11h. or 11'								1.2	
Vari. Dec. 15' 48";	P. L		1.056	56					566
· · · · · · · · · · · · · · · · · · ·									
						44			

part 12" 15" P. L..... 1.1669 2d. part 0" 36" P. L. 2.4785
The first part of this equation 12" 15" is subtractive, because the sun is proceeding towards the elevated pole; and the second part 36" is additive, because the declination is increasing, so that the whole equation is about 12 seconds subtractive; this applied to the middle time 12h. 1m. 4s. gives the time of apparent noon by the watch 12h. 0m. 52s. so that the watch is 52 seconds too fast for apparent time.

Thus in north latitudes the first part is to be added from the summer to the winter solstics, and sub-

[&]quot; India in norm institudes the tirst part is to be added from the sammer to the winter costice, and suptracted the rest of the year.

† It is here supposed that the elapsed time is less than 12 hours, which is generally the case, but if that time exceeds '12 hours, the second part must be applied in a contrary manner, to the above rule,

‡ On May 9, at noon, by the Nautical Almanac, the declination was 17° 24′ 42″, and on the following neon 17° 40′ 30″, the difference 15′ 48″, being the daily variation. The declination corresponding in the long, of 19° W. being 17° 25′ N. nearly.

LUNAR OBSERVATIONS.

ALMOST all the methods of determining the difference of longitude between any two places, depends on the general principle of finding the difference between the times of taking any observation, estimated under the meridian of both those places. For in any place it is reckoned to be noon when the sun is on the meridian, and as the sun, by his diurnal motion, ap pears on the meridian of Greenwich (from which the longitude is reckoned) one hour earlier than in a place in 15° W. longitude. and one hour later than in a place in 15° E. longitude, and in proportion for a greater or less longitude, it follows, that if at the time of taking an observation, the corresponding time was known at Greenwich, the longitude of the place of observation would be found by allowing 15° for every hour of difference between those times, the longitude being East when the time at Greenwich is earlier than at the place of observation, otherwise West. Now an observer at any place, may determine the apparent time at any moment, by a watch regulated by any of the preceding methods; and if, at the same moment, the apparent time could be obtained at Greenwich, nothing more would be necessary for determining the ingitude. One method of determining the time at Greenwich is by a watch regulated to Greenwich time: for it is evident that if a watch could be so constructed as to go uniformly at all times, and in all places, an observer, furnished with a watch thus regulated, would only have to compare the time at the place of observation with the time at Greenwich shewn by the watch, and the difference of the times would give the difference of longitude. This method is useful in a short run, but in a long voyage, implicit confidence cannot be placed in an instrument of such a delicate construction, and liable to so many accidents. Another method of determining the longitude is by observing the beginning or end of an eclipse of the moon, or the satellites of Jupiter, and taking the difference between the time of observation and the time given in the Nautical Almanac for the meridian of Greenwich; it being evident that such an eclipse must be observed at both places at the same moment of absolute time, consequently the difference of the times will be the difference of longitude. An observation of an eclipse of the sun, or an occultation, after making allowance for parallax, &c. as taught in the Appendix to this work, may be used in like manner, and this is the most accurate method known. However, observations of eclipses are but of small practical utility at sea; for those of the sun and moon happen too seldom, and the difficulty of observing the eclipses of Jupiter's satellites prevents that method from being made use of. Other methods have been proposed, but among them all there is not one of such practical utility, as that by measuring the angular distance of the moon from the sun, or from certain fixed stars situated near the ecliptic, usually called a Lunar Observation. For observations of this kind may be taken in fair weather at all times (except near the time of new moon) when the objects are above 80 or 100 above the horizon, and as the moon moves in her orbit about 1'in 2 minutes of time, it follows that if her angular distance can be ascertained from the sun or star within 1' the time at Greenwich will be known within 2 minutes, and the longitude within 50 miles. To facilitate this method, there is annually published, by the Commissioners of Longitude in England, a Nautical Almanac, containing the true angular distances of the moon from the sun and certain fixed stars, for the beginning of every third hour, calculated for the meridian of Greenwich, and the time corresponding to any intermediate hour, may be found by proportional parts; hence an observation of these angular distances being taken in any place, and the corresponding time at Greenwich found by the Almanac, and compared with the time at the ship, their difference will be the longitude of the place of observation. But before the observed angular distance is compared with those in the Nautical Almanac, the corrections for parallax and refraction must be applied to obtain the true distance; for the moon being seen always lower than her true place, and the sun and stars higher, the true distance is almost always greater or less than the observed distance.

Because the sun, by hisupparent diurnal motion, describes 860 degrees in 24 hours, which makes 5 degrees in an hour.

The angular distances of the moon from the sun and proper stars, are generally given in the Nautical Almanae from one object on each side of her, to afford a greater number of opportunities of observation, and to enable the observer to correct in a great degree, the errors of the instrument, the adjustments, or a faulty habit of observing the contact of the limbs, because these errors have a natural tendency to correct each other, in taking the mean of observations made with stars on different sides of the moon. Previous to taking the observation, the Nautical Almanac must be examined; to see from what objects the distances are computed, and from those objects only must the distances be measured.

There are only nine stars from which the angular distances are computed in the Nautical Almanac; and as it is of the greatest importance to be able to discover them easily, I shall here add a number of remarks which will

be found useful for that purpose.

The best way of discovering any star, is by means of a celestial globe. If that cannot be obtained, the time of the star's passing the meridian, and its meridian altitude, may be calculated, and by observing at that time, the star may be easily discovered. The distances marked in the Nautical Almanac afford also to the observer an easy method of knowing the star from which the moon's distance is to be observed; for he has nothing to do but to set the sextant or circle to the distance computed roughly for the apparent time, estimated nearly for the meridian of Greenwich, and direct his sight to the east or west of the moon, according as the distance at Greenwich was found in the VIII. IX. X. and XI. pages of the month: and having found the reflected image of the moon upon the horizon glass, sweep the instrument to the right or left, and that image will pass over the sought star, if above the horizon, and the weather clear: the star is always one of the brightest, and is situated nearly in a line perpendicular to the moon's horns, or which is the same thing, in the line of the moon's shorter axis produced.

The computed distance made use of in sweeping for the star, may be found in this manner. Reckon the apparent time at the ship in the mainer of astronomers (by counting 24 hours from noon to noon, and taking the day one less than the sea account;) to this time apply the longitude turned into time, by adding in west, and subtracting in east longitude; the sum or difference will be the apparent time at Greenwich nearly. Take the distances from the Nautical Almanac for the time immediately preceding and following this estimated time, and note the difference of these distances:

Then say as Sh. or 180' is to the difference of the distances, so is the difference between the apparent time at Greenwich and the next preceding time set down in the Nautical Almanac, to a proportional part to be added to the next preceding distance taken from the Nautical Almanac if the distance be increasing, but subtracted if decreasing; the sum or difference will be the distance at which the quadrant or sextant is to be fixed.

In sweeping for the stars by this method, it will often happen that two or more are swept upon at once; this might cause some difficulty to an inexperienced observer, who would be at a loss to know which to make use of. To remove this, the following description of these stars is added.

ARIETIS.

**

**

ALDEBARAN.

A **

This star bears about west, distant 23° from the Pleiades or Seven Stars; it is of the second magnitude, and may be known by means of the star π of the third magnitude, situated S. W. from a Arietis at the distance of $3\frac{1}{2}$ degrees. South from the star π at the distance of $1\frac{1}{2}^{\circ}$, is the star ν , of the fourth magnitude. The northernmost of these stars is a Arietis.

About 35° E. from a Arietis, and 14° S. E. from the Pleiades or Seven Stars, is the bright star Aldebaran. Near this star, to the westward, are six or seven stars of the third or fourth magnitude, forming with Aldebaran a figure resembling the letter V, as is represented in the adjoined figure, where Aldebaran is marked a. At the distance of 23° from this star, in a S. E. direction, are three very bright stars, situated in a straight line near to each other, forming the belt of Orion.

	·
POLLUX.	At the distance of 45° from Aldebaran, in the direction of E. N. E. is the Star Pollux, which is a bright star, though not of the first magnitude. N. W. from it, distant 5°, is the star Castor, of nearly the same magnitude, and you will almost always sweep both at once; the southernmost is the one used.
REGULUS. * * * * * Regulus.	E. by S. & S. from Pollux, at the distance of $37\&^\circ$, is the star Regulus, of the first magnitude; to the northward of this star, (at the distance of 8°) is a star of the second magnitude; near to these are five stars of the third magnitude, the whole forming a cluster resembling a sickle, represented in the adjoined figure, Regulus being in the extremity of the handle. A line drawn from the northern polar star, through its pointers, passes about 12° to the eastward of Regulus.
* SPICA.	E. S. E. from Regulus, at the distance of 54°, is the star Spica of the first magnitude, with no very bright star near it: S. W. from this star, at the distance of about 16°, are twe stars of the third or fourth magnitude, situated as in the advice joined figure; the two northernmost of these stars η, ν, form a straight line with Spica, and by this mark it may be easily discovered. A line drawn from the northern polar star, through the middle star of the tail of the Great Bear, will pass near to Spica.
ANTARES.	E. S. E. from Spica, at the distance of 46°, is the star Antares, in 26 degrees of south declination; it is a remarkable star, of a reddish colour; on each side of it, to the W. N. W. and S. S. E. about 2° distant, is a star of the third or fourth magnitude, no very bright star being near.
* AQUILÆ.	N. E. from Antares, at the distance of 60°, is the very bright star a Aquilæ; N. N. W. from which, at 2° distance, is a star of the third magnitude, and S. S. E. at 3° distance, another star of a lesser magnitude. These three stars appear nearly in a straight line. The star a Aquilæ is nearly of the same colour as Antares.
FOMALHAUT.	S. E. from a Aquilæ, at the distance of 60°, is the star Fo-malhaut, which is a bright star of high southern declination, its altitude in northern latitudes being small, never exceeding 20° in the latitude of 40° N. This star bears nearly south from the star a Pegasi, distant 45 degrees. A line drawn from the pointers, through the northern polar star, and continued to the opposite meridian, will pass very near to a Pegasi and Fomalhaut.
B*	by N. from a Aquilæ, at the distance of 48°, and west from a Arietis, at the distance of 44°, is the star a Pegasi, which may be known by means of four stars of different magnitudes, situated as in the adjoined figure; in which a represents a Pegasi, β a star of the second magnitude, bearing north of it distant 13°: the others are of lesser magnitude, and two

It would be a great improvement in the Nautical Almanac if the distances of the moon were also given from the planets Jupiter, Venus, Mars, and Saturn, particularly the two first, as their distances from the moon on account of their brightness, can be measured much more easily than from a fixed star, such observations can be taken during the twilight. These distances have been published, for several years, by Mr. Schumacher, of Copenhagen; also in Baron Zach's Journal.

by this mark it may be easily discovered.

of it, distant 13°; the others are of lesser magnitude, and two of them, η , μ , form a straight line with the star α Pegasi, and

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General observations on the taking of a Lunar Observation.

The accuracy of a lunar observation depends chiefly on the regulation of the watch, and on the exact measurement of the angular distance of the moon from the sun or star; a small error in the observed altitudes of those objects, will not in general much affect the result of the calculation.

The best method of regulating a watch at sea is by taking an altitude of the sun when rising or falling quickly, or when bearing nearly east or west, and noting the time by the watch. With this altitude, the latitude of the place, and the sun's declination, find the apparent time of observation by either of the preceding methods: the difference between this time and that shewn by the watch, will shew how much it is too fast or slow. A single observation taken with care will generally be exact enough; but if greater accuracy is required, the mean of a number of observations may be taken. If the distance of the sun and moon be observed when the sun is three or four points distant from the meridian, the apparent time of observation may be deduced from the altitude of the sun taken at the precise time of measuring the distance; this will render the use of a watch unnecessary, and will prevent any irregularity* in its going, from affecting the result of the observation. If a night observation is to be taken, the watch should be regulated by an altitude of the sun taken the preceding evening, and its going examined by means of another observation taken the next morning; for the time found by an altitude of a star cannot be so well depended upon, except in the morning and evening twilight, as the atmosphere in the night is precarious, and the horizon generally ill defined; but the altitude may be sufficiently exact for finding the correction used in determining the angular distance.

Although all the instruments used in these observations ought to be well adjusted, yet particular care should be taken of the sextant or circle used in measuring the angular distance of the moon from the sun or star, since an error of 1' in this distance will cause an error of nearly 50' in the longitude deduced therefrom. When a great angular distance is to be measured, it is absolutely necessary to use a telescope, and the parallelism of it with respect to the plane of the instrument must be carefully examined; but in measuring small distances the use of the telescope is not of such great importance, and a sight tube may then be used, taking care, however, that the eye and point of contact of the objects on the horizon glass be equally distant from the plane of the instrument. But it ought to be observed that it is always conducive to accuracy to use a telescope, and after a little practice it is easily done.

Whilst one person is observing the distance of the objects, two others ought... to be observing the altitudes; and the watch either suspended near one of the observers, or put into the hands of a fourth person appointed to note the times; the observer, who takes the angular distance, giving previous notice to the others to be ready with their altitudes by the time he has finished his observation, which being done, the time, altitudes, and distance should be carefully noted, and other sets of observations taken, which must be done within the space of 15 minutes, and the mean of all these observations must

be taken and worked as a single one.

When a ship is close hauled to the wind, with a large sea, or when sailing before the wind, and rolling considerably, it is difficult to measure the distance of the objects; but when the wind is enough upon the quarter to keep the ship steady, there is no difficulty, especially in small distances, which are much more easily measured than large ones, and are not so liable to error from an ill adjustment of the telescope; an observer would therefore do well to choose those times for observation, when the distance of the objects is less than 70° or 80°. An observation of the sun and moon is generally much easier to take when the altitude of the moon is less than that of the sun, because the instrument will be held in a more natural and easy manner. When the moon is near the zenith, the observation is generally difficult to take, and liable to be erroneous, because the observer is forced to place himself in a disagreeable posture. For the same reason, an observation of the moon and

[•] It is not uncommon to find a difference in the regulation of a-watch in the forencen and afternoon; this difference generally arises from the irregularity in the going of the watch.
• If the distances are measured by a circular instrument, it will not be necessary to note the several distances measured, but only the times and altitudes, as the sum of all the distances measured by the circle will be given by the instrument at the end of the observations, and if the altitudes of the objects are also measured by circular instruments, it will not be necessary to note the several elititudes, but only the times of observation. the times of observation.



a star is generally much easier to take when the star is lower than the moon. This situation of the objects may in most cases be obtained by taking the observation at a proper time of the day. But it must be observed that neither of the objects pught to be at a less altitude than 100, upon account of the uncertainty of the refraction near the horizon; for the horizontal refraction varies from 35' to 36' 40" only by an alteration of 40° in the thermometer. This alteration might cause an error of two degrees in the longitude.

In taking the altitude of the moon, the round limb, whether it be the upper or lower, must be brought to the horizon. In damp weather it is rather difficult to observe the altitude of the stars, on account of their dimness, particularly a Pegasi and a Arietis. Sometimes they are so dim that they cannot be seen through the holes of the sight vane of a quadrant, particularly if the mirrors are not well silvered; in this case the vane must be turned aside, and the eye held in nearly the same place, or the altitude must be taken by a sextant furnished with a sight tube.

We have here supposed that there were observers enough to measure the altitudes when the distance was observed; but if that is not the case, the altitudes may be estimated by either of the methods which will be hereafter

given.

Preparations necessary for working a Lunar Observation.

Find the apparent time of observation by astronomical account, reckoning the hours from noon to noon in numerical succession from 1 to 24, and taking the day one less than the sea account; to this time apply the longitude turned into time by Table XXI.* by adding if in west longitude, but subtracting if in east; the sum or difference; will

be the supposed time at Greenwich or reduced time.

In page VII. of the month of the Nautical Almanac, find the moon's semi-diameter and horizontal parallax, for the nearest noon and midnight before and after the reduced time, and find the difference of the parallaxes and the difference of the semi-diameters; then enter Table XI. with these differences respectively in the side column, and the reduced time at the top, opposite the former and under the latter, will stand the corrections; to be applied respectively to the semi-diameter and horizontal parallax marked first in the Nautical Almanac, additive if increasing, subtractive if decreasing; the sum or difference will be the horizontal semi-diameter and the horizontal parallax respectively, at the time of observation. To this horizontal semi-diameter, must be added the augmentation from Table XV. corresponding to the moon's altitude, the sum will be the true semi-diameter of the moon.

The sun's true semi-diameter is to be found in page III. of the month of the Nantical

Almanac.

To the observed altitude of the sun's or moon's lower limb add 12, but if the upper fimbs were observed, subtract 20', and from the star's observed altitude subtract 4', and you will have nearly the apparent altitudes of those objects respectively.

To the observed distance of the moon from a star, add the moon's true semi-diameter, if her nearest limb was observed, but subtract that semi-diameter if her farthest limb was observed; the sum or difference will be the apparent distance. But to the observed dis-

part of the catculation win we unaccounty,
ty by the chronometer.

These corrections may be found easily without the table, by saying, as 12 hours are to the reduced time, (rejecting 12 hours when it exceeds 12) so is the difference of semi-diameter or parallax for 12 hours to the corresponding correction. If the reduced time cannot be found accurately in the table, you

If any one wishes to obtain the apparent akitudes strictly, he must, from the observed altitudes, subtract the dip of the borizon taken from Table XIII. and add or subtract the semi-diameter of the object according as the lower or upper limb was observed.

^{*} Or by multiplying by 4 sexagesimally, in the manner directed in the note, page 124.
† When the sum exceeds 24 hours, you must subtract 24 hours and add one to the day of the month; and when the time to be subtracted is greater than the apparent time, the latter must be increased by 24 hours, and one day taken from the day of the month, conformable to the usual rules of addition and subtraction. If the chronometer used in taking the observation be regulated to Greenwich time, this part of the calculation will be unnecessary, because the reduced time at Greenwich will be given direct.

time, (rejecting 12 hours when it exceeds 12) so is the difference of semi-diameter or parallax for 12 hours to the corresponding correction. If the reduced time camou be found accurately in the table, you must use the nearest numbers which will in general be sufficiently accurate.

§ These altitudes are supposed to be taken at see by a tore-observation; and the application of the above numbers will give the apparent altitudes corresponding to observations taken on the deck of a common sized vessel (where the dip is about 4 or 5') to a sufficient degree of accuracy: if the observer was 40 or 50 feet above the water, 1' or 2' might be taken from these altitudes. The propriety of using these numbers will appear by considering that every wave, by raising the ship above the level of the sea, will after the dip, and that an error of 1' or 2' in the altitudes, will in general came but a small error in the result of the calculation of a lunar observation, so that for all practical purposes the above numbers may be esteemed as sufficiently exact. It may also be observed, that the error arising from this source will not generally be greater than that arising from neglecting the equations depending on the spheroical form of the earth, and on the density and temperature of the air; equations which are almost always neglected.

tenes of the sum and moon's nearest limbs, add their true semi-diameters; the sum will be the apparent distance.

These preparations are necessary in every method of working a lunar observation. The most noted methods are those of Dunthorne, Borda, Maskelyne, Rios, Witchell.

Lyons, &c. and improvements thereon by various authors.

Dunthorne's and similar methods, have one great advantage, in not being liable to a variety of cases; but those methods are tedious, when tables of logarithms to minutes only are used, by reason of the great exactness required in proportioning the logarithms to seconds. This is obviated in the excellent methods published by Rios and Stansbury, but they require large and expensive tables, and on that account are not in very general

sec. Witchell's and Lyon's methods do not labour under the inconvenience of requiring large tables, nor do they require any particular notice of the seconds in finding the log. sines and log tangents, but these methods are embarrassed with a variety of cases: sometimes the corrections are additive, sometimes subtractive, and learners find a difficulty in rightly applying them. To remedy this, a method was published in the first edition of this work, in which two corrections were constantly additive, two subtractive. and one small correction was additive, when the distance was less than 90°, but subtractive when above 90°.

This method was further improved in the Appendix to that edition, by means of four new Tables, which are inserted in this edition, and numbered XVII. XVIII. XIX. and XX. by means of which the work is considerably shortened, and all the corrections rendered additive. This method will now be given, after making a few remarks on the manner of taking the corrections and logarithms from these new Tables.

Tables XVH, contains a correction and logarithm to be used when the moon's distance from a star is observed, and Table XVIII. is a similar one, to be used when the moon's distance from the sun is observed. Both these Tables are so extended, that no proportional parts are necessary in taking out the corrections and logarithms, except the altitude of the Sun or Star is less than 7° 30', and at such altitudes an observation is liable to error on account of the uncertainty of the refraction: so that, in using these tables, it is sufficiently accurate to find the number nearest to the given altitude of the sun or star, and make use of the corresponding correction and logarithm. Thus if the star's altitude is 12° 25', the nearest number in Table XVII. is 13° 24', corresponding to which are correction 55' 45", and logarithm 1.3161.

Table XIX. contains the corrections and logarithms corresponding to the moon's horizental parallax and altitude, both being found at the same opening of the book. The corrections for seconds of parallax and minutes of altitude, are easily taken out by means of Tables A, B, C, placed in the margin. The method of finding these correc-

tions is given at the bottom of the Table ; they are always additive.

Besides the two logarithms taken from Table XVII. (or XVIII.) and XIX. this new rule requires only four logarithms to be taken from Table XXVII. to four places of figures, and to the nearest minute, it being in general unnecessary to proportion for the seconds.

We shall now give the rule for correcting the distance, and shall, for brevity, use the word sine, secant, and co-secant, instead of log. sine, log. secant, and log. co-secant respectively, and the same will be observed in the second and third methods of correcting the distance.

Shortest method of correcting the Apparent Distance of the Moon from the Sun.* Add the apparent distance of the moon from the sun, to their apparent altitudes, and note the half sum. The difference between the half sum and the apparent distance, call the first remainder; and the difference between the half sum and the sun's apparent al-

titude, call the second remainder.

Take from Table XXVII. the following logarithms, which mark beneath each other in two columns, viz. the sine of the apparent distance to be marked in both columns, the co-secant of the second remainder to be marked also in both columns, the secant of the first remainder to be placed in the first column, and the secant of the half sum in the second column.

Enter Table XVIII. (or Table XVII. if a star was used) and take out the correction corresponding to the sun's altitude (or star's :) take also from the same Table the cor-

responding logarithm, which place in column 1st.

Enter Table XIX. with the moon's apparent altitude and horizontal parallax, find the corresponding correction, which place under the former correction, and the logarithm, which place in column 2d.

However, if the planet's horizontal parallax is known, the corresponding water many be found mently as follows:

Add together the 1st correction and the cor. Tab. XVII. call the sum S. Take the difference between S and 60 and to its prop. log. add the prop. log. of double of the planet's horizontal parallax, and the arithmetical complement of the key. found in Table XVII. the sum (rejecting 10 in the indus) will be the prop. log. of the sought correction, which is to be added to the computed distance if S is less than 67, but industriated if S exceed 69. See note page 169 for an example of this rule.

† Rejecting always the test in the indices.

[•] This rule is the same as that for correcting the distance of the moon from a star, except in reading star for sun, and using Table XVII. instead of Table XVIII. If the distance of the goon from a planet is used, its parallax may generally be neglected, considering it as a fixed star, and using Table XVII. However, if the planet's horizontal parallax is known, the corresponding correction of the distance

The sum of the four logarithms of column 1st. will be the proportional logarithm of the first correction, and the sum of the logarithms of column second* will be the proportional logarithm of the second correction; these corrections being found in Table XXII. are to be placed under the former corrections.

Enter Table XX. and seek in the side column the nearest minute of the correction of Table XIX. opposite thereto and under the distance in the column marked Table XIX. will be a number of seconds which reserve. Enter the Table again, and find at the sides the nearest minute to the sum of the correction of Table XIX. and the second correction opposite to this sum and under the distance in the column marked Table XIX. +2d. Cor. will be a number of seconds; the difference between this and the former reserved number of seconds will be a correction to be placed under these already found. Then by adding all these corrections to the apparent distance, decreased by 2°, the true distance will be obtained nearly.†

To determine the Longitude from the true distance.

If the true distance of the observed objects can be found in the Nautical Almanac, in either of the pages VIII. IX. X. XI. of the month, opposite to the given day, or to that which immediately precedes or follows it, the time will be found at the top of the page. If the true distance cannot be found in the Nautical Almanac, take out the two distances, one of which is the next greater, and the other the next less, than the true distance ; take the difference between these two distances, and also the difference between the distance which stands first in the Nautical Almanac and the true distance. Then the Proportional Logarithm of the former difference (Table XXII.) being subtracted from the prop. log. of the latter difference, the remainder will be the prop. log, of a portion of time to be added to the time standing over the first distance in the Nautical Almanac, and the sum will be the true time at Greenwich. The difference between this time and the apparent time at the ship turned into degrees and minutes by Table XXI. will be the true longitude of the ship from Greenwich at the time of observation, and the longitude will be East if the time at the ship be greater than that at Greenwich, otherwise West.1

To exemplify the preceding rules there will now be given six examples of correcting the apparent distance by this new rule, in which examples we shall also include the preparation, and the determination of the longitude from the true distance.

EXAMPLE I.

Suppose that on the 2d of January, 1820, sea account, at 6' past midnight, apparent time in the longitude of 121° 30' E. by account, the observed distance of the farthest limb of the moon from the star Aldebaran was 47° 48' 53", the observed altitude of the star 50° 35', and the observed altitude of the moon's lower limb 70° 35'. Required the true longitude?

Preparation. Sea account Jap. 2, is by N. A. Jan. 1d. 12h. 6'
Long. 1210 30' E. in time 8 6 Reduced time, Jan. 1, 4

S. D. Jan. 1, neon midn.		51 49	•	hor. par. Jan. 1, noon 54 25 midn. 54 17	*	obs. alt	····	50 35 4
Difference		2		Difference 8	*	app. ak	••••	50 51
Table XI.		1		Table XI. 3	D	obs. alt. L.L	••••	7 0° 85'
Difference	14	50) hor. par. 54 22		add		12
Aug. Tab. XV		15			D	app. alt.		70 47
) s. d.	15	5		Obs. Dist. 🖈	8	F. L. S. D. sub.		48' 55" 15
•				App. Dist. 🖈	D		47	53, 48

^{*} Rejecting always the tens in the indices.

† The distance obtained by this rule is not perfectly correct, since several small corrections must be applied to obtain the true distance to the nearest second, viz. (1) The refraction taken from Table XII. which was made use of in constructing Tables XVII. XVIVI. and XIX, ought to be corrected for the applied for the spheroidal figure of the earth. And (3) a very small correction equal to the fourth correction of whitchell's method, given in page 62 of the Appendix of the second edition of the Requisite Tables of Maskelyne, must also be applied.—But to notice all these corrections would increase the calculation every much, and the result of a single observation, in which all these things were noticed, would probably not be so accurate as the mean of two or three observations, taken at different times of the day, in which these corrections were neglected, and the time necessary to take and work the latter observations would not be much greater than to work a single observation, in which all the corrections were noticed.

It may be necessary to observe, that if the times at the ship and Greenwich fall on different days, the least day is to be reckoned the greatest, though the hour of the day may be the least; thus 17th day louir I to be esteemed greater than 16th day 22 hough

To find the true distance.

App. Dist. ** App. Alt. ** App. Alt.	47 34 sine 9.8681 same 9.8681 App. Dist. less 2=43 \$3 48 60 \$1 2 Rem. 35° 45° co-40 0.2534 same 0.2544 App. Dist. less 2=43 \$3 48 70 47 1 Ren. 56 52 sec. 0.0969 sum 84° 26° sec. 1.0127 Table XIX. 1 dg. 1.85°6 Fable XIX. log. 2258 Cor. 1 1 29
Saint.	168 52 Cor. 2. 7 50
d Sum App. Dist.	84 29 Cor. 1' 29" P. L. 2.0950/2 Cor. 7' 80" P. L. 1.3816 Table XX. 20 47 34
1st. Rem.	36 52
} Sum ★ App. Alt.	84 26 50 Si
2d Rem.	33 55 To find the longitude.
	0 , "
	True distance
•	Dist. by N. A. at 3h 46 56 11
	Difference 28 38 P. L. 7981
-	Dist, by N. A. at 3h 46 56 11 at 6h 49 23 59
•	Difference
	0 58 42 P. L. 4866
	add 3
	Time at Greenwich
	Longitude in time

EXAMPLE II.

Suppose that on the 20th September, 1820, sea account, at 7h. 23' 45", P. M. apparent time, in the longitude of 186° 30' W. by account, the observed distance of the nearest limb of the moon from the star Antares was 31° 5' 18", the observed altitude of the star 12° 34', and the observed altitude of the moon's lower limb 20° 26'. Required the true longitude?

Preparation.

Sea account Sept. 20, is by the N. A. Sept. 19d. 7b. 28' 45".

		mone	(4.12.12.12.12		. •			
		Reduce	d time Sept. 19	Is	29 45	* obs. ait sub	ů	કં-ફે 4
D 6. D.	Sept. 19, midn	. 16 35) hor. par. Sept. 19,	mide.	66 47	* app. alt	12	30
	Sept. 20, noon	16 40	Sept. 20,	, noon	61 3	D obs. ak. L. L	20	26
	Difference	5	Differen	ce	16	add		12
	Lable XI.	3	Table X	I.	9	D App. Alt	20	SB
Aug.	Sun: Table XV.	16 38 5) hør. p	mr-	60 56	Obs. Dist.) * N. L. 8	16	18 45
) S. D.	16 43				App. Dist.] * 8	1 20	-1

^{*} This Log.=Log. Table XIX. 2280-Log. Table C 9=2268.

This Corr == Corr. Table XIX. 41' 49"+Corr. Table A 13"+Corr. Table B 6"=42' 06".

[†] If in this case the object, instead of being the fixed star Antares, had been a planet, whose horizontal parallax is 20", the correction of distance arising from this parallax, might be found by the rule in the note page 157. Thus, the sum of the Oerr. Tab. XVII. (8" 14") and Oor. 1 (1" 29") is 8z. 26" 48". Rejecting 50" the prop. log. of the remainder 0" 45" is 2-4010, (8" 14") and Oor. 1 (1" 29") is 8z. 26" 48". Rejecting 50 the prop. log. of the remainder 0" 45" is 2-4010, the prop. log. of the double bottometal par. 40" is 2-4014. The sum of these three logs. (rejecting 50 in the index) is 2-3633 equal to the prop. log. of 22" the required correction, subtractive from 470 24" 45" above found, because 3 exceeds 60 and it would therefore breame 470 24" 37.

To find the true distance.

App. Dist. ** App. Alt. D App. Art. Sum 1 Rem. 2 Rem.	81 20 sine 9.9950 12 3092 Rem. 44° 44′ Co-se 0.1525 20 381 Rem. 24 08 Sec. 0.0395 sum Table XVII. Log. 1.3194 Table 57 141 Cor. 5′ 36″ P. L. 1.5055 2 Cor. 24 6 44 44	57° 1 X1X, 44′ 5	. Lo 5" I	Bee. g.* P. L.	8.9950 0.1525 0.2666 .1868 6029		Dist. less 2— Table XVII. Table XIX. Cor. 1 Cor. 2 Table XX. e distance	 55 5 44	01 47 09 36 55 23
	10 june one of	mg u	****	••					
	True distance Dist. by N. A. at 18h								
	Difference		14	23	P.	L.	1.0974		
	Dist. by N. A. at 18h at 21h								
	Difference	1	52	49	P.	L.	2029		
	. · · ∆ dd	0 18	22	57	P.	L.	8945		
	Time at Greenwich								
	Longitude in time	10	59	12:	=164°	48	w.		

EXAMPLE III.

Suppose on April 25th 1820, at 2h. A. M. sea account, in the longitude of 166° E. by account, the observed distance of the moon's farthest limb from Antares was 76° 32′ 15″, the observed altitude of the star 25° 34′, the observed altitude of the moon's lower limb 17° 59′. Required the true-longitude?

Preparation.

				A repuration.							
	Sea a Long	ccm . 166	mt Ap 50 E.	ril 25, or by N. A. April 24	th •	1 4b. 11	0′ 4				
	Redu	ced	time	April 24th			56				
D & S. D.	April 24, noon midn-	14" 14	51" 53)'s Hor. Par. noon midn.		23'' 31			ıb.		4
	Difference Table XI.		2 0	Difference Table XI.		8 2	* 6	p. Alt. s. Alt L. add	L.	23 17	30 59 12
	Sum Aug. Tab. XV.	14	51 5	D. Hor. Par.	54	25		pp. Alt.	_	18	
:-	Sub.) S. D.	14	56			-	Obs. Sub.	Dist.) 8. D.	76°	14	15'' 56
. ,			T	find the true distance.	٠,		Арр.	Dist.	76	17	59

App. Dist. ** Ap. Alt. D Ap. Alt. Sum	23 18		gine 2Rem. 35° 29° co 1 Rem. 17° 18° se Table XVII. log 1 Cor. 2° 42° P.	c. 0.0201 5. 1.5794	i sum Table	XIX. log.	0.2382 ec. 0.2879 ‡ 2432	App. Dist. less Table XVII. Table XIX.§ 1 Cor. 2 Cor. Table XX.	2 <u>—</u> 74	57 10	; 19 50 51 42 40 23	
1 Rem.	17 35	18 29			1		,	True distance	76	00	45	

^{*} This Log .= Log. Table XIX. 1883+Corr. Table C 5.=1888.

¹ This Corr.=Corr. Table XIX. 5' 5"+Corr. Table A 2"+Corr. Table B 2"=5' 9".

[†] This Log .- Log. Table XIX. 2426-| Corr. Tab. C 6-2452.

[§] This Corr.=Corr. Table XIX. 10' 19"+Corp. Tab. 4 32"+Corr. Tab. B 0"=10" 51".

LUNAR OBSERVATIONS.

To find the true longitude.

True distance 76° 00′ 45″ By N. A. distance at 3h. 76 00 57 6h. 74 30 53 Difference 1 29 59 P. L. 2.9542 P. L. 3011

Oh. 0′ 24″ P. L. 2.6531 add 3

Time at Greenwich 3 0 24

Time at Ship

14

Difference is long, in time 10 59 36=164° 54′ E. from Greenwich.

EXAMPLE IV.

Suppose that on the 31st October, 1820, sea account, at about 1h. P. M. in the longitude of 75° W. by account, the following observations of the sun and moon were taken. Required the true longitude?

Preparation.

Tim	e pe	r wai	ch.	Obse	rved	dista N.		Observed		Observed A) L. L.	lt.
	Oh.	58'	5'		690	43	49	45	57	17	19
	0	59	8	1		43	18		52	17	9
	1	Ö	10	1		42	47		48	16	59
	1	1	4	l		42	20		44	16	48
	1	1	53	1		41	56		39	16	36
5)	5	0	20			14	10		240	84	50
	1	0	4		68	42	50	4.	48	16	58
•	_		_	⊙ s.	D.	16	9		ſ		
A	p p. 1	time		D S.	D.	14	53	add	12	add	12
					69	13	52	40	5 0	17	10
•				l Ap	p. Di	st.	- 1	App. Alt.	⊙ . I	App. Alt. D)

Long. 75° W. 5 . 30d. 6h. nearly. Reduced time Oct. D S. D. Oct. 30, noon 14' 50") Hor. Par. Oct. 30, noon 54' 21" midnight.. 14 48 midn. 54 13 Différence Difference Table XI. 1 Table XI. 49 D Hor. Par. 54 Aug. Table XV. ..

1h.

Sea account 31 Oct. or N. A. Oct. 30d.

To find the true distance.

App. Dist. Ap. Alt. Ap. Ak.	69 14 sine 46 0 2 Rem. 200 12' co s. 17 10 1 Rem. 3 2 sec. Table XVIII. log.	0.0006 sum 66 12 sec.		7 13 45- 59 11 30 52 52
Sum } Sum 1 Rem. 2 Rein.	182 24 68 121 Cor. 52" P. L. 8 2 20 12	2.518 2 Cor. 15' 14" I	2 Cor.	15 M 22 88 40 28

^{*} This Log.=Log. Tab. XIX. 2454+Cor. Tab. C 3=2457,

D S. D. 14 53

[†] This Corr.=Corr. Tab. XIX. 19' 12"+Corr. Tab. A 40"+Corr. Tab. B 0":=10" 82".,

To find the true longitude.

By N. A. Dist. at 6h. 68 40 29 Differed By N. A. Dist. at 9h. 67 19 11	ence	0° 1	0' 6" 21 18	P. L. P. L.	3.2553 3452	
add			13"			
Time at Greenwich Time at Ship Difference is long, in time	1	0	4	'2' W.n	om Greenv	rictt.

EXAMPLE V.

Suppose that on the 5th May, 1820, sea account, at about 4h. 4' P. M. in the latitude of 50° 1' S. and in the longitude of 1° E. by account, the following observations of the sun and moon were taken. Required the true longitude?

Preparation.

	Observed Dist. O D N. L.	Observed Alt. O L. L.	Observed Alt. D's U. I.
	1010 42 35"	14 53	410 58
	41 30	15 21	· 34
	40 22	15 49	. 4
Sun's	3) 124 27	46 3	96
Mean	101 41 29	15 21	41 32
Index errors	3	- 3	+ 8
Cor. Index errors	101 41 26	15 18	41 40
	⊙ S. D. 15 52 D S. D. 16 14	add 12	sub. 20
·	102 13 32 App. Dist.	15 30 ⊙ App. Ait:	41 20) App. Alt.

Sea account, May 5, or N. A. May 4d. 4h. Longitude 1° E.

Reduced t	me .		May 4d. 4h.					
S. D. May 4, noon midnight	16′	3′′	Hor. Par. noon 58' midnight 58	48!! 58				
Difference		3 1	Difference	10 3				
Aug. Table XV	16		p Hor. Par 58	51				
▶ S. D	16	14						

To find the true distance.

App. Dist. App. Alt. App. Alt.	41	50 20	2 Rem. (54° 2' 22 42	sec.	0.0850	tom	same same 79° 32° sec XIX. log.	. O.	0462 .7407	Ap. Dis. less 2 Table XVIII. Table XIX.		58 18	35
Sum j Sum 1st. Rem	159 • 79	4						19' 12" P,	-		ist. Cor. 2d. Cor. Table XX.		19	43 12 16
2d. Rem.	64	2									frue distance	101	52	3

To find the apparent time and true longitude.

© Correct Alfitude* Lat_of Ship Poler Distance!	15° 30 106	27'	secent	0.06254 0.01750	True distance By N. A. dist. at 35.	101°	.52 25		•	
Crr.	151	52		************	Difference		33	8	P. E.	7361
Sum Half Sum	75	46	co-sine	9.39071	By N. A. dist. at Sh.	102	25 47	6 39		
lisif Sum—Altitude	60	19	sine	9.93891	Difference	1	37	27	P. L.	2685
Apparent time 4h. 5'		Sum	sine	9.70475	ł	1b	1'	3"	P. Z.	4696
•			1	·	Time at Greenwich Time at Ship Longitude in time	4.	3 2 2	3 12 0(from enwich.

EXAMPLE VI.

Suppose that on the 8th of February, 1820, sea account, at about Sh. 36' A. M. in the longitude of 21° W. from Greenwich by account, six distances of the sun and moon's nearest limbs were observed by a circle of reflection to be 464° 10' 12" the corresponding times and altitudes being as in the following Table. Required the true longitude?

Preparation.

Apparent watch,			Observed O (ved A L. L.	lt.		ved A U. L.	lt.
8h.	33' 34 35 36 37' 39	24" 36 18 36 4	tanc from cle s	es the che o	dis- aken e cir- e end bser-		34°, 34 34 34 34 34 35	1' 13 21 31 39 3	•	61° 61 61 61 61 60	47' 35 27 17 9 45
6)	36	0	4640	10'	12"		206	48		368	0
8 <i>Ар</i> р.	36 time	0	77 ⊙ S. D.) S. D.	21 16 15	49 14 56	add	34	28 12	sub.	61	20 20
			77 App. D	53 ist.	52	⊙ Ap	34 p. Alt.	40) A ₁	61 op. Alt	0

Feb. 8, sea account, or by N. A. Feb. Long. 21° W. 7d. 20h., 36'

Reduced time Fe

7.3 OOL

reduced time	red.	/ Q. ZZQ.	
D's S. D. Feb. 7, midnight Feb. 8, noon	15′ 36″ 15 43	D's Hor. Par. Feb. 7, midnight Feb. 8, noon	57′ 9″ 57 37
Difference Table XI.	7 6	Difference Table XI.	28 23
Aug. Tab. XV.	15 42 14)'s Hor. Par.	57 32
)'s S. D.	15 56	·	

To find the true distance.

App. Dist. Ap. Alt. Ap. Alt.	61	90 2 Rem. 52° 7' co-se. 01 Rem. 8 58 sec.	0.0052 4 st	98 (21C	1,2509	Ap. Dist. less 2=75 Fable XVIII. Table XIXŞ. 1st. Cor.	92	52 44 20 26
form ! Sum !st. Rem. 24. Rem.	173 86 8	84		or. 5' 7" P. L.	1.5464	2d. Cor. Table XX.	77 5 2	18

^{*} The correct altitude is found by subtracting the refraction 3' from the apparent altitude 15° 30'.

? The Poins Distance is found by adding the Declination 16° 4' N. (corresponding to the reduced time) to 30°.

by Google

This log:_Log. Table XIX. 2018+Log. Tab. C 9=2025.
This Corr. Corr. Table XIX. 32, 7"+Corr. Tab. A 18"+Corr. Tab. B 0"-22" 20" by

To find the true longitude.

770 32 47" Difference 0° 30' 54" P. L. True distance 7653 By N. A Dist. Feb. 7, 21h. 78 3 Difference 1 32 9 P. L. 2908 76 31 32 Feb. 8, 0h. 1h. 0' 22" P. L. 4745

1h. 0' 22" P. L. 4745 add 21

22

Time at Greenwich Time at Ship

22 0 20 36

Diff. is Long. in time

1 24 22=21° 5' W. from Greenwich.

Second method of finding the true distance of the Moon from the Sun or a Star.

From the sun's refraction (Table XII.) take his parallax in altitude (Table XIV.) the remainder will be the correction of the sun's altitude.

The star's refraction (Table XII.) is the correction of its altitude.

From the proportional logarithm of the moon's Horizontal Parallax* increasing the index by 10, take the sine of the moon's apparent zenith distance (Table XXVII.) the remainder will be the prop. log. of the parallax in altitude, which must be found in Table XXII. and the moon's refraction (Table XII.) subtracted therefrom, the remainder will be the correction of the moon's altitude.

Add together the apparent distance of the sun and moon (or star and moon) and their apparent zenith distances (or complement of their apparent altitudes) and note the half sum of these numbers: the difference between the half sum and the moon's apparent zenith distance call the first remainder; and the difference between the half sum

and the sun's (or star's) apparent zenith distance call the second remainder.

To the constant log. 9.6990 add the co-secant of the half sum and the sine of the apparent distance (both taken from Table XXVII.) the sum, rejecting 20 from the index,

will be a reserved logarithm.

To the reserved logarithm add the sine of the sun's (or star's) apparent zenith distance, the co-secant of the first remainder (both taken from Table XXVII.) and the prop. log. of the correction of the sun's (or star's) altitude (Table XXII.) the sum, rejecting 30 from the index, will be the prop. log. of the first correction to be found in Table XXII.

To the reserved logarithm add the sine of the moon's apparent zenith distance! the co-secant of the second remainder (Table XXVII.) and the prop. log. of the correction of the moon's altitude (Table XXII.) the sum, rejecting 30 from the index, will be the prop. log. of the second correction, to be found in Table XXII.

Then to the apparent distance add the correction of the moon's altitude, and the first correction, and subtract the sum of the second correction and the correction of the sum's

altitude, the remainder will be the corrected distance.

Add 60' to the correction of the moon's altitude, and 60' to the difference between the correction of the moon's altitude and the second correction; find both these sums in the side column of Table XX. and in either of the vertical columns, under the corrected distance, find the seconds corresponding, is the difference of these two numbers will be a number of seconds to be added to the corrected distance when less than 90°, but subtracted when above 90°, the sum or difference will be the true distance.

^{*} Instead of finding the moon's horizontal parallax from the Nautical Almanac, we may find the proportional logarithm thereof in the same page of the month of that work. Thus if we would work Example III. preceding, by this rule, we might take out the logs, 5193 and 5187, instead of the Hor. Par. 34''39" and 54' 51", and obtain by means of Table XI. the sought log. 5195 without referring to Table XXII.

[†] All these corrections may be found by means of Tables XVII. XVIII. and XIX. Thus the correction of Table XVII. subtracted from 60 minutes, will give the correction of the star's altitude. The correction of Table XVIII. subtracted from 60 minutes, will give the correction of the sun's altitude. The correction of Table XIX. subtracted from 55' 42' will give the correction of the moon's altitude. Perhaps the use of these Tables in this and in the following method, would not be inconvenient.

correction of Table XIX. subtracted from \$9' 42" will give the correction of the moon's altitude. Perhaps the use of these Tables in this and in the following method, would not be incurvenient.

† This logarithm was found before, in calculating the correction of the moon's altitude.

† Observing to take both numbers from the same vertical column. It may be observed that the numbers in one of the columns of Table XIX. of this collection, are the same as those of Table XIX. of edition 1, and the numbers in the other column differ 18" from the former; but the numbers in the side. column of Table XIX. differ 80' from those in Table XIX. so that in using Table XIX. deltion 1st. it is unflecessary to add 60' to the correction of the moon's altitude and the first correction; this renders that Table rather more convenient in this second method, than Table XX. of this collection.

EXAMPLE—(the same as Example 1. preceding.)

Suppose the apparent distance of the centre of the moon from the star Aldebaran was 47° 33′ 48″, the apparent altitude of the star 50° 31′, the apparent altitude of the moon's centre 70° 47′, and the proportional logarithm of the moon's horizontal parallax 5199. Required the true distance of the moon from the star?

5 4 40	30° 0,	90° 0' Hor. Par. P. Log. 10.5199 ** Ap. Alt. 50 S1 b Z. D. 19° 13' sine 9.5174 ** Refrac. 47''
App. Alt.	70 47	* Ap. Alt. 50 31) Z. D. 19° 13' sine 9.5174 * Refrac. 47'
) Len. Dist.	19 13	* Zen. Dist. 39 23 17' 54"P. L. 1.0025
		Refraction 20
-		Cor. D Alt 17 34
App. Bist.) Zen. Dist. * Zen. Dist.	47° 34' 19 18 59 29	Constant log 9.6990 § Sum 53° 8' co-sec 10.0969 Dist. 47 34 sine 9.8681
Sum	106 16	Reserved Log 9.5640 Reserved Log 9.6640
j Sum D Zen. Dist.	53 8 19 13	* Zen. Dist. 39° 29' sine 9.8034) Zen. Dist. 19° 13' sine 9.5174 1 Rem. 33 55 co-sec 0.2534 2d. Rem. 13 39 co-sec. 0.6271
1st. Rem.	33 55	* Cor. 0' 47" P. L 2.3613 D Cor. 17' 34" P. L. 1.0106
Half sum ★ Zen. Dist.	53 8 . 39 29	1 Cor. 1 29 P. L 2.0821 2d. Cor. 27 18 P. L. 0.8191
2d. Rem.	13 59	Apparent distance 47° 33′ 48″
		Add { First Correction 1 29 Cor. D Alt 17 34
•		Sub. {2d. Cor. 27' 18" } 47 52 51 Cor. * Alt. 47 } 28 5
•	•	Corrected distance 47 24 46 Correction Table XX
		True distance 47 24 47 differing 2' from the former method.

"We shall now give a third method of correcting the apparent distance, being an improvement on Witchell's method, which was published in the former edition of this work. This improvement was made in consequence of a suggestion from a gentleman eminently distinguished for his mathematical acquirements,* that, by a small variation in the calculation, the number of cases might be lessened: and, upon examination, it was found that by making other alterations, the number of cases might be farther decreased, and the manner of applying the corrections rendered more simple. The method thus improved is as follows.

Third method of finding the true distance of the Moon from the Sun or Star. From the sun's refraction (Table XII.) take his parallax in altitude (Table XIV.) the remainder will be the correction of the sun's altitude.

The star's refraction is the correction of its altitude.

From the proportional logarithm of the moon's horizontal parallax, increasing the index by 10, take the co-sine of the moon's apparent altitude (Table XXVII.) the re-mainder will be the proportional logarithm of the moon's parallax in altitude; from which, subtracting the moon's refraction (Table XII.) the remainder will be the correction of the moon's altitude.

1. Add together the apparent altitudes of the moon and sun (or star) and take the half sum; subtract the lesser altitude from the greater, and take the half difference; then add together

The tangent of the half sum,

The co-tangent of the half difference,

The tangent of half the apparent distance.

The sum, rejecting 20 in the index, will be the tangent of the angle A, which must be sought for in Table XXVII. and taken out less than 90° when the sun's altitude is less than the moon's, otherwise greater than 90°. The difference of the angle A,

The late Chief Justice Parsons.
These corrections may be found by Tables XVI(`XVIII. XIX. as was shown in the note to the cound method, page 174.
Every co-tangent in Table XXVII. corresponds to two angles, one greater than 90°, the other less.

and half the apparent distance, so to be called the first angle, and their sum the second

2. Add together the tangent of the first angle.

The co-tangent of the sun or star's apparent altitude.

The prop. log. of the correction of the sun or star's altitude.

The sum, rejecting 20 in the index, will be the prop. log. of the first correction.

Or the refraction (Table XII.) corresponding to the first angle or its supplement, will be the first correction nearly; particularly if the altitude of the sun or star be great, and the first angle be near 90°.

Add together, the tangent of the second angle,

The co-tangent of the moon's apparent altitude,

The prop. log. of the correction of the moon's altitude,

The sum, rejecting 20 in the index, will be the prop. log. of the second correction.

4. The first correction is to be added to the apparent distance when the first angle is less than 90°, otherwise subtracted; and in the same manner the second correction is to be added when the second angle is less than 90°, otherwise subtracted. By applying these two corrections, we shall obtain the corrected distance or third angle.

5. Add 60' to the correction of the moon's altitude and to the second correction; find both these numbers in the side column of Table XX. and in either of the vertical columns.* under the third angle, find the numbers corresponding; the difference of these two numbers will be a number of seconds to be added to the third angle when less than 90°, but subtracted when above 90°, the sum or difference will be the true distance.

Thus it appears that the first, second and third corrections, depend on the first, second and third angles respectively; if either of those angles be less than 90° the corresponding correction will be additive: but if more than 90°, subtractive. This rule being uniform for applying all three corrections, makes it more easy to be remembered.

EXAMPLE—(the same as Example I. preceding.)

Suppose the apparent distance of the centre of the moon from the star Aldebaran was 47° 33′ 48″, the apparent altitude of the star 50° 31′, the apparent altitude of the moon's centre 70° 47′, and the proportional logarithm of the moon's horizontal parallax 5199. Required the true distance of the centre of the moon from the star?

D App. Alt. 20 47 * App. Alt. 80 81		*	Hor. Par. P. L. 10.5139 App. Alt. 70° 47' cor. 8.5174 * Ref. 47'
Styn 121 18 Half Sum 60	9 59' (an.	10.25002	17' 84" P. L. 1.0025
			20 D Refraction
Diff 20 16 Half Diff. 10 Half Dist. 23	8 co-t. 47 tan.	10.74781 9.64415	17 34 cer. D Alt. Apparent Distance 47° 83′ 48° 1st cor. add 42°
Angle A 77	g tan.	10.64198	47 51 30
Differ. is 1st Angle 53 *Ap. Alt. 50 Cor. *Alt.	22 tan. S1 co-t. 47" P. L.	10.1287, 9.9158 2.9613	2d cor. sub
First Corr. Sum is second Angle 100		2.4058	True distance 47 24 47 Diff. 2" from the former method-
DApp. Alt. 70 Cor. D Alt.	47 co-t. 17' \$4" P. L.	9.5423 1.0106	•
Second cor.	9 44 P.L.	1.2663	

If the star's altitude had been greater than the moon's, the angle A would have been 162° 51'. The first angle in this example is 53° 22' and the refraction (Table XII.) corresponding, is 43", which is nearly equal to the first correction.

Method of taking a Lunar Observation by one observer.

Three observers are required to make the necessary observations for determining the longitude; one to measure the distance of the bodies, and the others to take the altitudes. In case of not having a sufficient number of instruments or observers to

^{*} Both numbers must be taken from the same vertical column, as was observed in the note to the second method, and the other remarks of that note are applicable to this method.

take the altitudes it has been customary to calculate them; there being given the latitude of the place, the apparent time, the right ascensions and the declinations of the objects. These calculations are long, when an altitude of a star is to be computed, and much more so when that of the moon is required; and a considerable degree of accuracy is required in finding, from the Nautical Almanac, the moon's right ascension and declination, which must be liable to some error on account of the uncertainty of the ship's longitude. The following method of obtaining those altitudes is far more simple, and sufficiently accurate. This method depends on the supposition that the altitudes increase or decrease uniformly.

Before you measure the distance of the bodies, take their altitudes, and note the times by a watch, then measure the distance and note the time (or you may measure a number of distances, and note the corresponding times, and take the mean of all the times and distances for the time and distance respectively)—after you have measured the distances, again measure the altitudes, and note the times: Then from the two observed altitudes of either of the objects, the sought altitude of that object may be

found in the following manner:

Add together the proportional logarithm (Table XXII.) of the variation of altitude* of the object between the two times of observing the altitudes, and the prop. log. of the time elapsed between taking the first altitude and measuring the distance; from the sum subtract the prop. log. of the time elapsed between observing the two altitudes of that object; the remainder will be the prop. log. of the correction to be applied to the first altitude, additive or subtractive according as the altitude was increasing or decreasing; to the altitude thus corrected; apply the correction for dip of the horizon and semi-diameter as usual.

EXAMPLE.

Suppose the distances and altitudes of the sun and moon were observed as in the following Table: It is required to find the altitudes at the time of measuring the mean distance?

Wen.	Time. 25. 3m. 20s. 4 20 5 50 2 4 50	Dist. @ and D N L. 402 0' 0'' 0 30 1 30 40 0 40	2h. 2m.		me. Alt. 2m. 50s. 7 0 4 50	Observed. ③'s L. L. 40° 20' 59 12 1 8
	's alt. L obs. D s. of dist.	34' 2b. 2' 0'' 2 4 30	P. L. 7238	Variation @ 's alt Time 1st obs. sun Time mean obs.	1° 8′ 2h. 2′ 30′′ 2 4 50	P. L. 4528
Differen	ce	2 50	P. L. 1.8575	Difference	2 0	P. L. 1.9642
two	time between observations ion of alt.	the } 4' 10"		Sum Elapsed time between the two observations	4' 80"	2.5770 P. L. 1.6021
ist alt		20 46	add	Correction of air. Sub. from sun's 1st. air.	0° 30′ 40 20	P. L. 7749
Alt.) i	s L. L. at gim can obs.	e of \{ 21 6		Alt. (a)'s L. L. at time of the mean obser.		•

Thus at the time 2h. 4'30", the mean observed distance of the sun and moon's nearcet limbs was 40° 0' 40", the altitude of the moon's lower limb 21° 6', and the altitude of the sun's lower limb 39° 50'; these altitudes must be corrected for dip and semidiameter as usual.

[•] Table XXII is only calculated as far as 3°, and if the variation of altitude exceed that quantity, you must exster the table with minutes and seconds, instead of degrees and minutes, and the correction of altitude at table out is minutes and seconds must be called degrees and minutes respectively.
† Or add its arithmetical complement, neglecting 10 in the index of the sum.

In this manner I have often obtained the altitudes in much less time than they could have been obtained by other calculations.

The same method may be used for finding the sun's altitude, when taking an azimuth. by noting the times of taking the observations by a watch, and taking two altitudes, the one before, the other after the observation, and proportioning the altitudes as above.

Any person who wishes to calculate strictly the apparent altitudes, may proceed according to the following rules.

The apparent time, the ship's latitude and longitude, and the sun's declination given, to find the apparent altitude of his centre.

RULE.

With the apparent time from noon, enter Table XXIII. and from the column of rising take out the logarithm corresponding, to which add the log. co-sine of the latitude, and the log. co-sine of the sun's declination; their sum, rejecting 20 in the index, will be the logarithm of a natural number, which being subtracted from the natural co-sine of the sum of the declination and latitude, when they are of different names, or the natural co-sine of their difference when of the same name, will leave the natural sine of the sun's true altitude at the given time. The refraction less parallax being added to the true altitude, will give the apparent altitude.

In general it will be near enough to take out the refraction only from Table XII. and

neglect the parallax.

EXAMPLES.

Required the true altitude of the sun's centre, in lat. 49° 57' N. and long. 75° W. July 26, 1820, at 6h. 56m. 30s. in the morning, sea account?

App. time	H. M. S. 12 0 0 6 56 30	•
Time from noon Latitude Declin. at that time	5 3 30 49 57 ON. 19 26 ON.	its log. in col. of rising 4.87859 its log. co-sine 9.90888 its log. co-sine 9.97458
Difference	80 31	nat. number
True alt. Refraction	259 45' 2	nat. sine
App. alt.	25 47	

EXAMPLE II.

What will be the true altitude of the sun's centre in the latitude of 39° 20' N. and the longitude of 40° 50' W. Nevember 26, 1820, at 3h. 21m. 30s. apparent time in the afternoon, sea account?

App. time from noon Latitude Declin. at that time	H. M. 8. 3 21 30 39 30 ON. 20 52 OS.	its log. in col. of rising
Sun ·	60 12	nat. number
True alt. Refraction	18 5 6 4	mat. sine
App. akt.	13 40 .	

The apparent time, the latitude and longitude given, to find the apparent altitude of a fixed star.

RULE.

Turn the longitude into time, and add it to, or subtract it from, the apparent time*

The apparent time must be taken (as usual) one day less than the sea account, and the hours must be reckoned from none to most in numerical succession from 3 to 24. It may also be observed, that if the observer be furnished with a chronometer, regulated to mean Greenwich thus, this part of the operation may be saved, reducing the seems time to apparent, by applying the equation Table IV. A, with a different sign from that in the Table, as is taught in the introduction to the tables.

at the ship, according as the longitude is west or east, the sum or difference will be the time at Greenwich.

Find, in the Nautical Almanac, the sun's right ascension for the noon preceding the time at Greenwich, and add thereto the correction taken from Table XXXI. corresponding to the hours and minutes of the time at Greenwich, the sum will be the sun's right ascension, which being added to the apparent time at the ship, will give the right ascension of the meridian, rejecting 24 hours when the sum exceeds 24 hours.

Find the star's right ascension and declination in Table VIII. for the year 1820, and correct them for the years clapsed since that time, by means of the annual variations given in the same table, and you will obtain the star's right ascension and declination at the time of observation.+

The difference between the star's right ascension and the right ascension of the meri-

dian, will be the distance of the star from the meridian.

Find in the column of rising of Table XXIII. the logarithm corresponding to the star's distance from the meridian, † and add thereto the log. co-sine of the latitude of the ship, and the log. co-sine of the declination of the star, the sum, rejecting 30 in the index, will be the logarithm of a natural number (Table XXVI.) which subtracted from the natural co-sine (Table XXIV.) of the sum of the declination and latitude when of different names, or the natural co-sine of their difference when of the same name, will leave the natural sine of the star's true altitude.

The refraction being added to the true altitude will give the apparent altitude.

EXAMPLE.

What was the apparent altitude of Aldebaran, at Philadelphia, April 12, 1820, sea

account, at 5h. 57m. 18s. in the afternoon, apparent time?
In Table VIII. the right ascension of Aldebaran for 1820, is 4h. 25m. 36s. and the variation for 31 months is 1" to be added because the time is after 1820; hence the right ascension at the given time 4h. 25m. 37s. The declination of the star for 1830 is 16° 8' N. its variation for 31 months being neglected.

Apparent time by N. A. April 11, Longitude 75° 9' W	H. M. S. 5 57 18 5 0 36	
Time at Greenwich, April 11	10 57 54	
© 'a R. A. April 11, at noon, by N. A. Var. for 10h. 57' 54" by Table XXXI.	1 19 28 · 1 41	
�'s R. A. at time of obs	1 21 9 5 57 18	
R. A. Mer	7 18 27 4 25 37	
*'s dist. from merid	2 52 50 Its log. in col. rising	9.88457
5.0	at. number 19966 its leg.	4.30030
True altitude 45 39 n Refraction 1	at. sine 71518	
Apparent altitude 45 40		

^{*} If any of the 24 bright stars are used, whose right accession or north polar distances are given for every 10 days in the year in the Nautical Almanac, we can obtain from it by inspection the right accession, and deduce the declination from the polar distance, and the numbers thus found having born corrected for aberration and nutation, are to be used as rather more accurate.

f If the distance from the meridian exceed 12 hours, you must subtract from 24 hours before entering table XXIII.

The apparent time, and the latitude and longitude of the ship given, to find the apparent altitude of the moon's centre.

Turn the longitude into time (by Table XXI.) and if in west longitude add it to, but in east longitude subtract it from the apparent time* at the ship, the sum or difference will be the time at Greenwich.

Take the sun's right ascension out of the N. A. for the preceding noon at creenwich, and add thereto the correction taken from Table XXXI. corresponding to the hours and minutes of the time at Greenwich, the sum will be the sun's right ascension, which being added to the apparent time at the ship, will give the right ascension of the meridian, rejecting 24 hours when the sum exceeds 24 hours.

Take from the N. A. the moon's right ascension and declination for the time immediately preceding and following the time at Greenwich, and proportion for the time at Greenwich, by means of Table XXX. and you will obtain the moon's right ascension

and declination at the time of observation.

Turn the moon's right ascension into time (by Table XXI.) and the difference between that time and the right ascension of the meridian, will be the moon's distance; from the meridian; with which enfer table XXIII. and take out the corresponding logarithm from the column of rising, and add thereto the log. co-sine of the latitude of the ship, and the log. co-sine of the declination of the moon, the sum, rejecting 20 in the index, will be the logarithm of a natural number (Table XXVI.) which subtracted from the natural co-sine (Table XXIV.) of the sum of the declination and latitude when of different names, or the natural co-sine of their difference when of the same name, will leave the natural sine of the moon's true altitude; from which, subtracting the correction corresponding to the altitude in Table XXIX.; there will remain the apparent altitude nearly.

EXAMPLE.

What was the moon's apparent altitude April 29, 1820, sea account, at 7h. 55m. 52s. P. M. in lat. 42° 34' S. long. 65° 7' 30" W. from Greenwich?

April 29, sea account is April 28 Long. 65° 7′ 30″ V	by N. A. a	t		7	M. 55	5. 52 30
Long. 03 7 30	v. m time					
App. time at Gree	nwich	••••••		12	16	22
Sun's Rt. ascen. April 28, at 12h. 16 Apparent time at the ship						
Right ascension of the meridian) 's right ascension in time						36 20
) 's distance from the meridian	••••••			4	27	44
Corresponding to which in the col. lo Latitude	g. rising is	•••••••	CO-8	ine	4.78 9.86 9.97	717
	nat. num. nat. co-sine	42092 92343	log.	_	4.62	490
p 's true alt 30 10 Cor. Table XXIX. 48	nat. sine	50251				
» Арр. alt 29 23						

The apparent time is to be counted from noon to noon, as directed in the preceding note.
† When the distance exceeds 12h, you must enter table XXIII. with the difference between that dispace and 24 hours.

1 In strictness you ought, instead of this correction, to use the correction of the moon's altitude, corresponding to her apparent altitude and horizontal parallax.

To find the Longitude by the Eclipses of Jupiter's Satellites.

The eclipses of the satellites are given in page III. of the mouth of the Nautical Alasanac for mean time at Greenwich. There are two kinds of these eclipses—an Imsersion, denoting the instant of the disappearance of the satellite by entering into the shadow of Jupiter, and an Emersion, or the instant of the appearance of the satellite is coming from the shadow. The immersions and emersions generally happen when the satellite is at some distance from the body of Jupiter, except near the opposition of Japiter to the sun, when the satellite approaches near to his body. Before the opposition they happen on the west side of Jupiter, and after the opposition on the east side; but if an astronomical telescope is used which reverses the objects, the appearance will be directly the contrary. The configurations, or the positions in which Jupiter's satellites appear at Greenwich, are laid down every night, when visible, in page XII. of the month of the Nautical Almanac.

As these eclipses happen almost daily, they afford the most ready means of determining the longitude of places on land, and might also be applied at sea, if the observations could be taken with sufficient accuracy in a ship under sail, which can hardly be done, since the least motion of a telescope which magnifies sufficiently to make these

observations, would throw the objects out of the field of view.

As these eclipses are given in the Nautical Almanac in mean time, it is necessary to regulate your watch to mean time;* this is easily obtained from the apparent time by applying to the latter the equation of time taken from the Nautical Almanac, by adding or subtracting according to the directions in the column from whence the equation was taken; hence the error of a watch with respect to mean time may be ascertained.

The watch being thus regulated, you must then find nearly the time at which the eclipse will begin at the place of observation; this may be done as follows: Find from the Nautical Almanac the time of an immersion or emersion, and apply thereto the longitude turned into time, by adding when in east, but subtracting when in west longitude, the sum or difference will be nearly the mean time when the eclipse is to be observed at the given place. If there be any uncertainty in the longitude of the place of observation, you must begin to look out for the eclipse at an earlier period; and when the eclipse begins, you must note the time by the watch, and after applying the correction for the error of the watch, if there be any, you will have the mean time of the eclipse at the place of observation; the difference between this and the time in the Nautical Almanar, being turned into degrees, will be the longitude from Greenwich.

EXAMPLE.

Suppose that on the 21st of August, 1820, sea account, in the longitude of 127° 55' W. by account, an immersion of the first Satellite of Jupiter was observed at 7h. 12m. 32s. P. M. mean time. Required the longitude?

By N. A. immersion	Aug. 20th.	15h. 47' 52"
By observ. Aug. 21, sea account, or by N. A	Aug. 20th.	7 19 32

To find the Longitude by Eclipse of the Moon. .

The determination of the longitude by an eclipse of the moon is performed by comparing the times of the beginning or ending of the eclipse, as also the times when any number of digits are eclipsed, or when the earth's shadow begins to touch or leave any remarkable spot on the moon's face; the difference of time between the like observations made at different places, turned into degrees, will be the difference of longitude of those places.

When the beginning or end of an eclipse of the moon is observed at any place, the longitude of that place may be easily found by comparing the time

w In the Almancs published before 1805, the apparent time of the eclipses was given instead of the



of observation with the time given in the Nautical Almanac, for the difference between the observed time of beginning or ending, and the time given in the Nautical Almanac, will be the ship's longitude in time, which may be turned into degrees by Table XXI. Thus if the beginning of an eclipse of the moon was observed March 30, 1820, sea account, at 9h. 59\frac{2}{3}m. the time by the N. A. being March 29 or March 30, sea account, at 5h. 16\frac{2}{3}m. their difference 4h. 43m. is the longitude of the place of observation ==70° 45', which is east from Greenwich, because the time at the place of observation is greatest.

To find the Longitude by a perfect time-keeper or chronometer.*

It was before observed, that if a chronometer could be made in so perfect a manner as to move uniformly in all places, and at all seasons, the longitude might easily be deduced therefrom, by comparing the time shown by the chronometer, regulated to the meridian of Greenwich (or some other known meridian) with the mean time at the place of observation. For the difference of these times would be the difference of longitude between that meridian and the place of observation. The moderate price of good chronometers now, in comparison with their values many years since, together with the various improvements in their construction, have caused this method of determining the longitude to be much more used within a few years, than it was when the first editions of this work were published: we shall therefore explain more

fully the use of this instrument, and the methods of regulation.

If a chronometer is to be used on a voyage, it must be adjusted, and its rate of going ascertained, before sailing. This may be done by taking altitudes of the sun or some other heavenly body, and finding therefrom the apparent time of observation, by any of the methods before given in pages 154—161. To this time must be applied the equation of time, found in page II. of the month of the Nautical Almanac, or in Table IV. A (reduced to the moment of observation by means of Table VI. A) by adding the equation to, or subtracting it from, the apparent time, according to the directions given in or at the top of each column of the table, the sum or difference will be the mean time of observation, being the same time as would be shown by a chronometer whose motion is perfectly uniform. Comparing this mean time of observation with the time by the chronometer, shows how much it is then too fast or too slow for the meridian of the place of observation; and by repeating the operation on a future day, the rate of going may be ascertained. If it is found to gain or lose a few seconds or parts of a second per day, that allowance must be made on all future observations at sea. Thus, if on the 1st of June, 1824, at 5h. 10m. 20s. by the chronometer, the mean time, deduced from an observation of the sun's altitude, was 5h. 12m. 40s. the chronometer would then be too slow by the difference of those times 2m. 20s. and if on the 21st of June following the time by the chronometer was 4h. 15m. 35s. when the mean time was 4h. 18m. 17s. the chronometer would then be too slow by the difference of those times or 2m. 42s. and the rate would have varied in 20 days from 2m. 20s. to 2m. 42s. which is a difference of 22 seconds in 20 days, being 1_{11}^{1} seconds per day, and this rate must be allowed on all future observations at sea, until a new regulation can be obtained, at some place whose longitude is known. It is best to have a considerable number of days interval between the two observations for fixing the rate, by which means it may be determined to tenths of a second, the absolute error of the observations being reduced in finding the daily rate, by dividing by the number of days. above difference of 22 seconds had been erroneous 2s. and the true value 20s. the daily rate would be one second instead of 1s.1, varying only one tenth of a second, notwithstanding the observations on which the rate was established, contained an error of 2 seconds.

^{*} The Chronometers most celebrate. for correciness are those made by Mr. French, London, and for gale by JANES LADD, No. 30 Wall-street, New-York, who mechanically understands that valuable instrument.
Propriete:

125

Having regulated a chronometer, in the manner first mentioned, at a place whose longitude from Greenwich is known, it is easy to find how much it is too fast or too slow for the meridian of Greenwich, by allowing for the difference of meridians. Thus, if the above mentioned observation of June 1, was made in place in 74° west longitude, corresponding in Table XXI. to 4h. 56m. the chronometer on that day would be too slow for Greenwich time by the sum of 4h. 56m. and 2m. 20s. or 4h. 58m. 20s. In general it will be full as simple. when thus regulating a chronometer, at a place whose longitude is known, to reduce at once the mean time at the place of observation to the meridian of Greenwich, by adding the longitude if west, subtracting if east, the sum or difference will be the mean time of observation upon the meridian of Greenwich, the difference between this and the time given by the chronometer. shows how much it is too fast or too slow for Greenwich mean time. Thus by adding the longitude 4h. 56m. to the mean time of the above observation 5h. 12m. 40s. the sum 10h. 8m. 40s. is the mean time at Greenwich, from which subtracting the time by the chronometer 5h. 10m. 20s. the remainder 4h. 58m. 20s. is what the chronometer is too slow for Greenwich time, as was found before.

The chronometer having been thus regulated to Greenwich time, and the daily rate of its going ascertained, if this rate should remain unaltered, the time at Greenwich will be known by it, at any moment at sea, and if at that moment by any observation of the sun, moon, planet or a fixed star, the opporent time be found by any of the methods explained in pages 154—161, and the mean time at the ship deduced therefrom, by applying the equation of time, as above explained, then the difference between this mean time at the ship, and the mean time at Greenwich shown by the chronometer, will be the longitude, which may be turned into degrees and minutes by Table XXI. We shall explain by a few examples the preceding remarks.

EXAMPLE I.

Wishing to regulate a chronometer, in a place whose latitude is 51° 30' N. and longitude 114° E. from Greenwich, I observed Oct. 10, 1824, at 8h. 21m. A. M. sea account per chronometer, the altitude of the sun's lower limb, by a fair observation 15° 32', the correction for semi-diameter, parallax and dip being 12'. It is required to find the error of the chronometer for seas time at Greenwich?

The apparent time of this observation, computed as in Example I. page 158, is 8h. 7m. 9s. A. M. corresponding to Oct. 9d. 20h. 7m. 9s. by the Nautical Almanac. From this subtract? the longitude 1140 turned into time 7h. 36m. by Table XXI. the remainder Oct. 9d. 12h. 51m. 9s. is the apparent time at Greenwich. Now by Table IV. A, the equation of time for Oct. 9d. at noon is sub. 12m. 44s. with a daily increase of 16s. and this in Table VI. A, under 16s. and opposite to 12h. 31m. 9s. gives 9s. to be added to 12m. 44s. (because it is increasing) the sum 12m. 53s. is the equation of time, which by the table is subtractive from the apparent time at Greenwich Oct. 9d. 12h. 31m. 9s. to obtain the mean time at Greenwich Oct. 9d. 12h. 18m. 16s. If the mean time at the place of observation is required, it would be found by subtracting the equation of time 12m. 53s. from the apparent time at the place of observation, 8h. 7m. 9s. and it would therefore be 7h. 54m. 16s.

EXAMPLE II.

May 10, 1824, at 5h. 39m. P. M. sea account per chronometer, in latitude 59° 54′, in a place whose longitude was known to be 17° 50′ E. from Greenwich, the altitude of the sun's lower limb by a fore observation was 15° 45′, the correction for dip, parallax and semi-diameter being 12′. It is required to find the error of the chronometer for mean time at Greenwich, and at the place of observation?

If the longitude had been east, it would have been subtractive.
 This is to be added if the ship's longitude is weet.



184 TO FIND THE LONGITUDE BY A CHRONOMETER.

The apparent time of this observation, computed as in Example II. page 156, is May 9d. 5h. 34m. 26s. by the Nautical Almanac. From this subtract? the longitude 170 30, turned into time, 1h. 10m. by Table XXI. the remainder May 9d. 4h. 24m. 26s. is the apparent time at Greenwich, By Table IV. A, the equation of time for May 9th. at noon is sub. 3m. 48s. with a daily increase of 3s. and this in Table VI. A, under 3s. and opposite 4h. 24m. 26s. is 1s. which added to 3m. 48s. (because it is increasing) gives the equation of time, at the moment of observation, sub. 3m. 49s. Subtracting this, according to the direction in the table, from the apparent time at Greenwich May 9d. 4h. 24m. 26s. leaves the mean time at Greenwich May 9d. 4h. 24m. 26s. leaves the mean time at Greenwich May 9d. 4h. 24m. 37s. Subtracting the same equation 3m. 49s. from the apparent time at the place of observation 5h. 54m. 26s. gives the mean time at the place of observation 5h. 30m. 37s. The difference between the mean time at Greenwich 4h. 20m. 37s. and the time by the chronometer 5h. 50m. is 1h. 9m. 23s. which is the time the chronometer is too fast for Greenwich mean time.

EXAMPLE III.

Suppose that July 27, 1820, sea account, the apparent time was found by an altitude of the sun to be 1h. 5' 8" P. M. when by a watch well regulated to mean Greenwich time, the time was 4h. 3' 8" P. M. Required the longitude?

Apparent time	1h.		8″ 8
			-
Mean time at place of observation		11	16
Time per watch	4	3	8

EXAMPLE IV.

Suppose that May 14, 1820, sea account, the apparent time was found by an altitude of the sun to be 4h. S' 5" P. M. when the time by the watch was 2h. P. M. the watch being too slow for mean Greenwich time 11' 9". Required the longitude?

Apparent time	4h .	3' 3	5″ 56	Time per watch Watch error add	2h. 0' 0" 11 9
Mean time at place of observ. Time at Greenwich	3 2	59 11	9 г. м. 9	Time at Greenwich	2 11 9 P. M.
Difference of longitude	1	48	0==97° 0′ I	ē.	

EXAMPLE V.

Suppose that on June 14, 1820, sea account, in a place whose longitude from Greenwich was known, a number of observations were taken to ascertain the going of the watch; and it was found that on that day it was 10° too slow for mean Greenwich time, and lost time 2" per day; and that on July 14, 1820, sea account, the time per watch was 6h. 0' 6' P. M. when, by an observed altitude of the sun, the apparent time was 1h. 16' 10' P. M. Required the longitude?

Apparent time		5	22	Error of watch, June 14, 0' 10" skt 30 days at 2" 1 0 skt	
Mean time at place of obser. Time at Greenwich	1 6	¥1 1	3 9 16	Error July 14,) V Ç
Longitude	4	39	44=	Time at Greenwich 6 1 16 =69° 56' W.	

^{*} This is to be added if the ship's longitude is west.

TO REGULATE A CHRONOMETER.

To regulate a Chronometer by Lunar Observations.

It sometimes happens that a chronometer is by accident suffered to down when at sea, in which case it can be regulated, by means of a graumber of lunar observations, which must be taken with the greatest cand with objects on different sides of the moon. These observations may made on the same day, or on several successive days, finding by each servation how much the chronometer is too fast or too slow for Greenw time, and taking the mean result for the error at the mean time of obvation.

. 4444	•	
EXAMPLE I.		ample III.
182Q.	m. s. 1819. m. s. 1217.	
April 6. Too fast 1st, obse	2 12 April 20. Too slow 1st. obs. 1 19 Aug. 5. Too	o slow 1st. obs.
,, ,, 2d. obs.	2 08 n 21 2d. obs. 1 21 n 6. Total	o fast 2d. obs -
Sd obe	0 64 80	slow 3d. obs
" Jeh aha	\$ 10 " 22 " 8. Te	o fast 4th. obs
200.00E	3) 66	0,200
		· • • • • • • • • • • • • • • • • • • •
	4) 36 Chec. too slow April 21, 1 22	-y-
37		
Mean error	2 9 Chro. too ale	ov, Aug. 7 -
Mean error	2 9 (Varo. too at	ow, Aug. 7 -

In the last example same of the observations made the chronometer slow, and others too fast; these are marked with different signs + and the sum is to be found, noticing the signs in the algebraical manner

taught in the introduction to the appendix of this work.

It has lately been discovered that chronometers generally go faster on board vessel than when on shore; this variation has been sometimes found to as much as 14 seconds per day, though in general not more than 1 or 2 coads. It is suspected that this arises from the attraction of the iron in remedy, the chronometer having acquired a small degree of magnetism remedy this inconvenience it has been recommended to keep the chronometers always in the same place on board the ship, and to regulate them we thus placed, before leaving the port, or by means of lunar observations ter the above manner) when at sen. Thus, in the first of the above exples, the chronometer was 2m. 9s. too fast for Greenwich time April 9, 1 suppose now by a set of lunar observations made April 30, 1320, it was for to be fast 2m. 30s. the variation would be 21s. in 21 days, which is 1 see per day for the acceleration of the chronometer.

It has also been found that chronometers generally gain by an increated density of the air, and lose by a decrease of density. The firing of guns board a vessel will sometimes alter the rate of going, unless the instrumble well suspended, or held in the hand during the firing. Any sudden will sometimes alter the rate. The imperfection of the oil used will a sometime impair the instrument. Finally, the mechanism used to core the change of temperature of the air may not do it completely, and so error may arise from this source. Notwithstanding these various car of error, it is wonderful to observe how accurately some of the

chronometers perform their office.

To find the Longitude by a Variation Chart,

In the year 1700, Dr. Halley proposed to find the longitude by a the published, on which the lines of the variation of the compass were drawn and since that time several similar charts have been published for the superpose; but the difficulty of determining the variation, combined with a crouses, will probably prevent this method from being sufficiently accurate be generally useful.

The method of using this chart is as follows: On the parallel of latit which you are in, find the observed variation, and that point will be

place of observation.

A chart, on which the lines of the dip of the magnetic needle are miled, might be used in the same manner for determining the longitude.

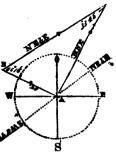
PROBLEMS USEFUL IN NAVIGATION.

PROBLEM I.

Coasting along shore, I saw a cape of land bearing N. N. E. and after sailing W. N. W. 20 miles, it bore N. E. by Required the distance of the ship from the cape at both stations ?

BY PROJECTION.

Describe the compass E. S. W. and let its centre A represent the place of the ship at the first station; draw the W. N. W. line AB equal to 20 miles, and B will represent the second station. Draw the N. N. E. line AC, of an indefinite length, and the line BC parallel to the N. E. by E. line of the compass; the point of intersection C will represent the place of the cape; and the distance BC being measured will be found 36 miles, and AC 30 miles.



BY LOGARITHMS-(by Case I. Obl. Trig.)

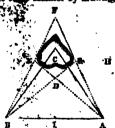
The difference between N. N. E. and W. N. W. is 8 points or 90°, therefore BAC is a right angle; also the difference between the N. E. by E. and N. N. E. is 3 points—angle ACB and the difference between the N. E. by E. point and the point opposite to W. N. W. is 5 points, equal to the angle ABC.

To find the distance	BC.		
As sl. ang. ACB 3 points .	87.	co. 0.25526	
Is to the distance AB 20 .		1.30108	
So is sine angle BAC 8 points	•	10.00000	
To the distance BC 36.0		1,58629	

So is sine angle ABC 5 points To the distance AC 29.98

The above solutions are by ease I. Oblique Trigonometry, though they might have been done in this example by case II. of Right-Angled Trigonometry, because the angle BAC is a right-angle. If the bearings of the middle point C ofant Island (or any remarkable peak) had been

taken and determined in this manner, you might have found at the same time the limit of the dimensions of the island, by measuring with a quadrant or extant, held in a horizontal position, the angular distances between that middle coint and the extremes of the island. For by drawing the lines ADE, AGF, making, the ingles DAC, GAC, with AC equal to the angular distances observed at A, and in the stage manner by drawing the lines BDG, BEF, making angles with BC equal to the angular distances observed at A, and in the stage manner by drawing the lines BDG, BEF, making angles with BC equal to the angular distances observed at B, you would obtain the guadrant. angular distances observed at B, you would obtain the quadrilateral figure DEFG, within which the island is to be placed. If similar observations could be procured at H, they would in general take off the corners at D and F; and observations at I, would generally take off the corners at E and G; and by observing the projecting points and cives in the island, while sailing round it, and drawing a figure conformable thereto within the limiting space thus found, the form and dimensions of the island may be obtained to a



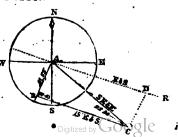
PROBLEM II.

Being at sea, I saw two headlands whose bearing from one another by the chart was W. by N. and E. by S. and distance 15 miles; the westernmost bore from me S. S. W. and the easternmost S. E. by E. Required my distance from each of them?

BY PROJECTION.

Draw the compass N. E. S. W. and through the centre A, draw the E. by S. line AR, the S. S. W. line AB, and the S. E. by E. line AC, and continue the two latter indefinitely, but upon the former AR take AD=15 miles; through D draw DC parallel to AB, to meet w AC in C, and draw CB parallel to AD. Then A will be the place where the headlands B and C were observed; and the distance AB of the westernmost headland being measured, will be found to be 5,8 miles, and the distance AC of the easternmost headland 15 miles.

considerable degree of accuracy.



BY LOGARITHMS.

Between the S. S. W. line AB and the S. E. by E. line AC, are 7 points, = < BAC; and between the S. E. by E. line AC, and the E. by S. line AD are 2 points = < CAD = < ACB (because AD, BC are parallel)—therefore ACB+BAC=9 points, and since all three sagles ACB, BAC, ABC are equal to 16 points, the angle ABC is also equal to 7 points, therefore (by art. 59 Geom.) the sides AC, CB are equal, being opposite to the equal angles ABC, BAC. If these angles had not been equal, the side AC mignature been calculated in the same manner as we shall now calculate the side AB.

To find the side AB.
As sine BAC 7 points co. ar.
Is to BC 15 miles 1.17609
So is sine ACB 2 points 9.58284
To AB 5.85 0.76786

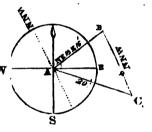
This Problem or the first may be used for finding the distance of a ship from any heldland, &c. when taking her departure from the land.

PROBLEM III.

Two ships sail from the same port, one sails N. E. \ E. 16 miles, the other sails easterly 20 miles, and then finds that the first bears N. N. W. Required the other ship's course, and the distance between the two ships?

BY PROJECTION.

Draw the compass ESW, and let its centre A represent the port sailed from; draw the N. E. & E. line AB = 16 miles, and through B, the line BC, parallel to the N. N. W. line, and continue it indefinitely; take 20 miles in your compasses, and putting one foot in A, describe with the other an arch cutting the line BC in C, and join AC. Then B will be the place of the first ship, C that of the second, and AC the course steered by the second ship, which will



be nearly E. S. E. & E. and BC the distance of the ships 174 miles.

BY LOGARITHMS.

The course from B to C is S. S. E. (opposite to N. N.W.) and from B to A is S. W. & W. (opposite to N. E. & E.) the difference between these bearings is 64 points=75° 7'=the angle ABC; having this angle and the sides AB, AC, the other angles and side may be found by Cases II. and III. of Oblique Trigonometry as follows:

As the side AC 20 miles
Is to sine ABC 78° 7'
Bo is side AB 16 miles

1.30103
Subtract

To sine aagtes C 49° 57'
For N. N W, add 22 30

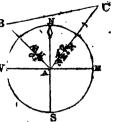
Bum makes N. 72 27 W, the bearing of A from C, whence the course of the ship from A towards C, h S. 72° 27' E, or E, S. E, ‡ E.

PROBLEM IV.

Two ships sail from the same port, one N. W. 30 miles, and the other N. E. by N. 40 miles. Required the bearing and distance of the ships from each other?

BY PROJECTION.

Braw the port sailed from; draw the N. W. me AB=50 miles, anothen N. E. by N. line AC=40 miles, join BC, which will be the bearing and distance of the two ships. Whence the bearing will be found to be W. S. W. W. and the distance 45.1 miles nearly.



BY LOGARITHMS (by Cases IV. and V. Ob. Trig.)

Between the N. W. line AB and the N. E. by N. Fine AC, there are 7 points—angle BAC, half the supplement of which to 180° is 50° 57½—half sum of the angles C and B.

To find the							ance &C.	
Assum of AB and AC		g. ar.				o 30/	ar. €0.	
Is to their difference				Is to side AC	40			1,69906
So is tang. 4 sum angle	s 50°	371	. 10.08583	So is sine angle A	1 78	45		9.99157
Tó tan. 🛓 diff.	9	521	9.24073	To the distance	BC	45.1		1.65393
Sum=ngle B Diff.=angle C	60 40	30 45						

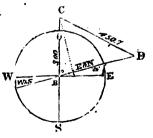
To the angle $C=40^{\circ}$ 45', add the course from C to $A=58^{\circ}$ 45', the sum is 74° 50', which is the bearing of B from C, viz. S. 74° 50' W. or W. S. W. 2 W. nearly.

PROBLEM V.

Two ports bear from each other E. by N. and W. by S. distance 400 miles; a ship from the custernmost sails northerly 450,7 miles, another from the posternmost sails 300 miles, and meets the first. Required the course steered by each ship?

BY PROJECTION.

Draw the compass ESW, and let the centre B represent the westernmost port; draw the E. by N. line BD=400 miles, and D will be the easternmost port; with 300 in your compasses and one foot in B, describe an arch; with 450,7 in your compasses, and one foot in D, describe another arch cutting the W former in C; join DC, BC. Then BC will be the course sailed by the westernmost ship, and DC the course sailed by the easternmost.



BY LOGARITHMS.

To find the angle CBD.

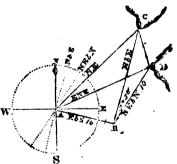
· ·	v	
By Theo. IV. Trig.	By Theo. V. Trig.	
Divide the triangle BOD into two right-angled	CD=460.7	
triangles by means of the perpendicular CA, and	BD=400 log. ar. co. 7	.39794
bisect BD in a, then	BC=300 log. ar. co. 7	-52288
As the base BD 400 ar. co 7.39794		•
Is to the sum of BC, CD, 750.7 2.87547	Sum 1150.7	
So is diff. of BC, CD, 150.7 2 17811		.75993
	sum less CD 124.65 lug. 2	.09569
To twice A a	· • • • •	
		.77644
Half or A a 141.4	Half sum 59° 22′ co-sine 9.	.88822
} BD=Ba=	2	
	D. M. I. Bo 44 - Lundo (TRD) - Handman	e
Diff. is BA	Doubled is 78 44=Angle CBD. Having	TOUNE
Then in the triangle AOB.	this angle, we may find either of the others	, mas,
As hypot. DC 300 247712	To find the angle CDB.	
Is to radius 90°		.34611
bo is AB 58.6 , 1.76790		.99155
•	So is BC 309 2	.47712
To en-sine CBD 72º 41' 9.29078	·	
	To sine CDB 40° 45′ 9.	.81478

As the angle CBD is 78° 44' or 7 points nearly, and the course from B to D is E. by N. the course from B to C must be north. The course from D to B being W. by S. or W. 11° 15' S. and the angle BDC=40° 45' the bearing of C from D must be W. 29° 80' N. because 40° 45'—11° 15'=29° 30'.

PROBLEM VI.

Coasting along shore, I saw two headlands, the first bore from me N. E. the second E. N. E.—after sailing E. by S. 10 miles, the first bore N. by E. and the second N. E. by N. Required the bearing of the two headlands from each other, and their distance?

Draw the compass NESW, and let its centure A represent the place of the ship at the first station; draw the E. by S. line AB == 10 miles, and B will bette place of the ship at the second station; draw the N. E. line AC, and the E. N. E. line AD; through the point B draw the lines BC, BD parallel to the N. by E. and N. E. by N. lines, and the points C and D where they intersect the lines drawn from A to the same headlands will be the points representing them respectively; join the points C and D;—then will CD be the



distance of the two headlands, and a line drawn through A parallel to CD will represent the hearing of those places from each other on the compass.

BY LOGARITHMS.
In the triangle ABC, we have all the angles and the side AB to End BC. For the bearings of B and C from A are E. by S. and N. E. the difference being 5 points—BAC; and the bearings of B and A from C, are E. by W. and S. W. the difference being 3 points equal to the angle ACB.

To find the side BC.									
As sine of ACB 3 pts. ar. co.			0.25526						
Is to the side AB 10			1.00000						
So is sine angle BAC 5 pts.		•	9.919 5						
To BC 14.97			1.17511						

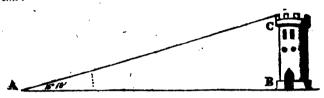
In the triangle ABD, we have all the angles and the side AB to find BD. For the bearings of B and A from D, are S. W. by S. and W. S. W. the difference being 3 points:::BDA; and the bearings of B and D from A, are E. by S. and E. N. E. the difference being also 5 points, equal to the angle BAD; therefore the a. le BAD:::BDA, and (by art. 39 Geom.) BD:::AB=10 miles. If these angles had not been equal, we might have calculated the side BD in the same manner as BC.

Now in the triangle CBD we have BD=10, BC=14.97, and the angle CBD=22° 30', for the bearings of C and D from B are N. by E. and N. E. by N. differing 2 points or 22° 30'; hence we may find the other angles and side CD as in case IV. Obl. Trig.

To find the angles BCD, BDC.	To find the distance CB.
As sam of BC, BD 24.97 ar. co	As sine ang. BUD 35° 44' ar. co 4.1 ls to side BD 10 1.9 So is sine angle UBD 22° 38' 3.5
To tang. ½ diff 45 1 10.00028	To the distance OD 6,29 6.5
Sum is angle BDC= 123 46 Diff. is angle BCD= 33 44 or nearly 3 points, and as the bearing of B from C is S. by W. The hearing of D from C must be S. S. E.	

PROBLEM VII.

Being 96 fathoms from the bottom of a tower, I found its altitude above the horizontal line drawn from my eye was 15° 10'. Required the elevation above that line?



BY PROJECTION.							
Draw the horizontal	line AB==96	fathoms, and					
perpendicular thereto, angle BAC=15° 10' an	the line BC	; make the					
angle BAC=15° 10' an	d draw AC to	cut BC in C,					
then will BC be the h	neight of the	tower 28 fa-					
fhoms.	_						

BY LOGARITE	1118		
As radius 90° . \.			10.09000
As radius 90° Is to the dist. AB 96 fathoms	٠	•	1.96321
So is tang. angle A 15º 10'	•	•	9.45304
To the height BC 26,0 fathour			1.4163

When an object, whose elevation above the horizon is to be determined, is at a very great distance, it will be necessary to notice the correction arising from the curvature of the earth and the refraction, and apply that correction to the height estimated by the above method. Thus if the angular elevation of a mountain whose base was more distant than the limit of the visible horizon, was observed by an instrument of reflection; the approximate height must first be obtained, as in the preceding example, and then the correction of that approximate height for the curvature of the earth, refraction, and dip, must be calculated by the following rule, and added to that height, the sum will be the true height above the level of the sea.

Rule. Find in Table X, the number of miles corresponding to the height of the observer above the level of the sea, and take the difference between that number and the distance of the mountain from the observer in statute miles; with that difference enter the same table and find the height in feet corresponding, which will be the correction to be added to the approximate height to obtain the true height of the mountain above the level of the sea.

EXAMPLE. Suppose the distance was 32 statute miles (or 188960 feet) and the observed altitude 1° 2′, the observer being 18 feet above the level of the sea. Required the height of the mountain above the same level?

As radius Is to distance 168960 So is elevation 1° 2' tang	log.	:	10.00000 5.22779 8.25616	Dist. of mountain Tab. X. 18 feet	:	:	:	:	M. 52: 5,61
Approx. height 3048 Correction 398	log.	•	3.48395	Difference Corresponding Corr-	Ťah	x.	:	:	26,39 396(t.

Sum S446 is the true leight above the level of the sea.

PROBLEM VIII.

Sailing towards Cape-Cod, I discovered the light-house just appearing in the horizon, my eye being elevated 20 feet above the sea; it is required to find the distance of the light-house, supposing it to be elevated 200 feet above the surface of the sea?

The solution of this problem depends on the uniform curvature of the sea, by means of which all terrestrial objects disappear at certain distances from the observer. These distances may be computed by means of Table X. in which the elevation in feet is given in one column, and the distance at which it is visible, is expressed in statute miles in the other column. If the place from which you view the object be elevated above the horizon, you must add together the distances corresponding to the height of the observer and the height of the object, the sum will be the greatest distance at which that object is visible from the observer.

In the present example the height of the observer was 20 feet, and the height of the object 200 feet.

In Table X. opposite 20 feet is 5,92 miles.

200 feet 18,71

Distance 24,63 statute miles of about 691 to a degree, the distance in nautical leagues, of 20 to a degree, being about 7.

PROBLEM IX.

A man being on the main-top-gallant-mast of a man of war, 200 feet above the water, sees a 100 gun ship she had engaged the day before, hull to; how far were those ships distant from each other?

A ship of 100 guns or a first rate man of war, is about 60 feet from the keel to the rails, from which deduct about 20, leaves 40 for the height of her quarter-deck above water. Now a ship is seen hull to when her upper works just appear.

In Table X. opposite 200 feet stand

18.71 8.37 40 feet

Distance 27.08 miles.

PROBLEM X.

Upon seeing the flash of a gun, I counted 30 seconds by a watch before I heard the report: How far was that gun from me, supposing that sound moves at

the rate of 1142 feet per second?

The velocity of light is so great, that the seeing of any act done even a rember of miles distance, is instantaneous; but by observation it is found that sound moves at the rate of 1142 feet per second, or about one statute maile in 4.6 seconds; consequently the number of seconds elapsed between seeing the fissh and hearing the report, being divided by 4.6 will give the distance in statute miles. In the present example the distance was about 64 miles, because 50 divided by 4,6 quotes 61 nearly.

PROBLEM XI.

To find the difference between the true and apparent directions of the wind.

Suppose that a ship moves in the direction CB from C to B, while the wind moves in its true direction from A to B, the effect on the ship will be the same as if she was at rest and the wind blew in the direction AC with a velocity represented by AC, the velocity of the ship being represented by BC. In this case the angle BAC will represent the difference between the true and apparent directions of the wind; the apparent being more a-head than the true, and the faster the vessel goes the more a-head the wind will appear to be. We must, however, except the case where the wind is directly aft, in which case the direction is not altered. It is owing to the difference between the true and appa-



rent directions of the wind, that it appears to shift its direction by tacking ship; and if the difference of the directions be observed when on different boards (the wind on both tacks being supposed to remain constant, and the vessel to have the same velocity and to sail at the same distance from the wind) the half difference will be equal to the angle BAC; by knowing which, together with the velocity of the ship BC, and the angle BCA, we may obtain the true velocity of the wind; or, by knowing the velocity of the wind and of the ship, and the apparent direction of the wind, we may calculate the difference between the true and apparent directions of the wind.

Thus if the velocity of a ship represented by BC be 7 miles per hour, that of the wind represented by AB 27 miles per hour, and the angle of the vessel's course with the apparent direction of the wind BCA=7½ points; the difference between the true and apparent directions of the wind would be obtained by drawing the line BC=7½ points, then with an extent equal to 27 miles, taken from the scale, and with one foot in B describe an arch to cut the line AC in A, join AB; then the angle BAC being measured, will be

the sought difference between the true and apparent directions of the wind. BY LOGARITHMS.

As AB Is to BCA So is BC	miles points miles	log. ar. colog. sine log.	8.56864 9.99790 0.84510
m n.a	 		0.4444

To BAC 14° 57′ log. sine

So that in this case the difference between the true and apparent directions of the wind is about 14 points, and by tacking ship and sailing on the other board as above mentioned, the wind would appear to change its directions above 24 points.

PROBLEM XII.

To measure the height of a mountain by means of the heights of two barometers taken at the top and bottom of the mountain.

Procure two barometers with a thermometer attached to each of them. in order to ascertain the temperature of the mercury in the barometers, and two other thermometers of the same kind to ascertain the temperature of the air. Then one observer at the top of the mountain, and another at the bottom, must observe at the same time the heights of the barometers and the thermometers attached thereto, and the heights of the detached thermometers, placed in the open air, but sheltered from the sun. Having taken these observations, the height of the upper observer above the lower may be determined by the following rule, which is adapted to a scale of English inches and to Fahrenheit's thermometer.

RULE. Take the difference of the logarithms of the observed heights of the barometer at the two stations, considering the first four figures, exclusive of the index, as whole numbers, the remainder as decimals; to this difference must be applied the product of the decimal 0,454, by the difference of the altitudes of the two attached thermometers, by subtracting if the thermometer was highest at the lowest station, otherwise adding: the sum or difference will be the approximate height in English fathoms. Multiply this by the decimal 0,00244, and by the difference between the mean of the two altitudes of the detached thermometers and 32°, the product will be a correction to be added to the approximate height when the mean altitude of the two detached thermometers exceeds 32°, otherwise subtracted; the sum or difference will be the true height of the upper above the lower observation English fathoms, which multiplied by 6 will be the height in feet.

EXAMPLE.

Suppose the following observations were taken at the top and bottoms of mountain. Required its height in fathous?

Attached Obs. at lower station Upper station	570	Detacl	hed thern 56° 42	29,68 inch. lo	g. 14724,6 g. 14027,8
Difference	14	Mean	49 52	Difference 0.454×14	696,8 6,4
		Diff.	17	Approx. height 690,4 × 17 × 0,00244	
		•		Height in fath.	719,0

--000

MENSURATION.

PROBLEM I.

To find the area of a Parallelogram.

RULE.

MULTIPLY the base by the perpendicular height, the product will be the area.

Norz. If both dimensions are given in feet, inches, &c. the product will be the area expressed in square feet, square inches, &c. respectively; if one of the dimensions be given in feet and the other in inches, the product divided by 12 will be the answer in square feet; if both dimensions are given in inches, the product will be square inches, which divided by 144 will be the answer in square feet. The same is to be understood in finding the area of other surfaces.

EXAMPLE I. Suppose the base DB of the rectangular parallelogram ACBD is 7 feet, and the perpendicular BC 3 feet; required the area? The product of the base 7 feet by the perpendicular 3 feet gives the area 21 square feet.

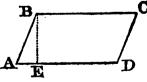
D B

EXAMPLE II. Suppose ACBD is a board whose D length DB is 22 feet and breadth BC is 14 inches; required the number of square feet? The product of the base 22 feet by the breadth 14 inches is 308,

this divided by 12 gives 25 square feet, the sought area.

EXAMPLE III. If DB be 25 inches and BC 20 inches; required the area in square feet? The product of the base 25 inches by the perpendicular 20 inches gives 500, which divided by 144 gives the area 5,47 or 3 47 square feet.

EXAMPLE IV. Given the base AD of the oblique angular parallelogram ABCD, equal to 30 feet, and the perpendicular height BE 15 feet; required the area of the parallelogram? Multiply the base 30 feet by the perpendicular 15 feet; the product 450 is the A area in square feet.



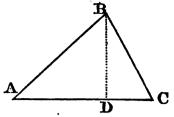
PROBLEM II.

To find the area of a Triangle.

RULE. Multiply the base by half the perpendicular height, and the product will be the area required.

C c

EXAMPLE. Given the base AC 30 feet, and the perpendicular DB 20 feet, required the area of the triangle? The base 30 multiplied by half the perpendicular 10 gives the area 300 square feet.



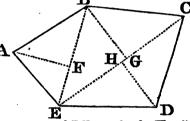
PROBLEM III.

To find the area of any irregular right-lined figure.

RULE. Reduce the figure to triangles by drawing diagonals therein; then find the area of each triangle, and the sum of them will be the area of the

find the area of each triangle, and the proposed figure. Or, instead of finding the area of each triangle separately, you may find at one operation the area of two triangles having the same diagonal by multiplying the diagonal by half the sum of the perpendiculars let fall thereon.

EXAMPLE. Required the area of the figure ABCDE, in which EC=38 feet, EB=22 feet, and



the perpendicular AF=13 feet, BG=14 feet, and DH=12 feet? The diagonal EB, 22 feet, multiplied by half the perpendicular AF, 6.5 feet, gives the area of the triangle AEB, 143 square feet; and the diagonal EC, 33 feet, multiplied by half the sum of the perpendiculars BG DH, 13 feet, gives the area of the figure BCDE, 429 feet; this added to the triangle AEB 143 feet, gives the whole area 572 square feet.

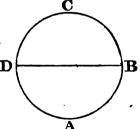
PROBLEM IV.

To find the area of a circle.

RULE. Multiply the square of the diameter of the circle by the quantity

0.7854, and you will have the sought area.

Note. Instead of multiplying by 0.7854 you may multiply by 11 and divide by 14, the quotient will be the area nearly. This quantity .7854 represents the area of a circle whose diameter is 1. The circumference of the same circle being 3.1416 nearly. The proportion of the diameter to the circumference is expressed in whole numbers by the ratio of 7 to 22 nearly; or more exactly by 113 to 355.*



EXAMPLE.

Required the area of a circle ABCD, whose diameter BD is 10.6 feet?

The diameter 10.6 multiplied by itself and by .7854 gives the sought area 88.247544 square feet.

PROBLEM V.

To find the area of an Ellipsis or Oval.

Rule. Multiply the longest diameter by the least, and the product by

^{*} This ratio may be easily remembered by observing that if the first three odd numbers 1, 3, 5, are repeated twice, they will produce the quantity 113355; the three first figures of which make the first term of the ratio, and the three last the last term of the ratio.

D

В

.7854, this last product will be the area required.

EXAMPLE. Required the area of an Ellipsis or Qval ABCD, whose longest diameter AC is 12 feet, and the shortest diameter BD 10 feet?

The product of the two diameters is 12×10=120, this multiplied by .7854 gives the sought area 94.2480 square feet.

The area of a sector of a circle may be found by means of the whole

area of the circle obtained in Problem IV. by saying, as 360 degrees is to the angle contained between the two legs of the sector, so is the whole area of

the circle to the area of the sector.

There are various regular solids, the most noted are the following.—(1) A Cube, which is a figure bounded by six equal squares. (2) A Parallelepped which is a solid terminated by six quadrilateral figures, of which the opposite ones are equal and parallel. (3) A Cylinder, which is a figure formed by the revolution of a rectangular parallelogram about one of its sides. (4) A Pyramid, which is a solid decreasing gradually from the base till it comes to a point. There are various kinds of Pyramids according to the figure of their bases: thus if the base be a triangle, the solid is called a triangular pyramid; if a parallelogram, a parallelogramic pyramid; and if a circle, a circular pyramid, or simply a Cone. The point in which the pyramid ends, is called the Vertex, and a line drawn from the vertex perpendicular to the base is called the height of the pyramid.

PROBLEM VI.

To find the solidity of a Cube.

RULE. Multiply the length of a side of the Cube by that length, and that product by the same length, and you will have the solidity required; which will be expressed in cubic feet if the dimensions were given in feet; but in cubic inches if the dimensions were given in inches,

EXAMPLE. If the side AB of the cube of 6.8 feet it is required to determine the solidity?

The product of 6.3 by 6.3 is 39.69, this multiplied again by 6.3 gives the solidity 250.047 cubic feet.

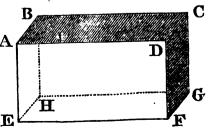
PROBLEM VII.

To find the solidity of a Rectangular Parallelepiped.

RULE. Multiply the length, breadth and depth, into each other; the product will be the solidity required.

EXAMPLE. Suppose in the parallelepiped ABCDFGHE, the length EF is 56 feet, the breadth GF 16 feet, and the depth FD A 12 feet; it is required to find the solidity?

The product of the length 36 by the breadth 16 is 576, this multiplied by the depth 12 gives the solidity 6912 cubic feet.



PROBLEM VIII.

To find the solidity of a Cylinder.

RULE. Multiply the square of the diameter of the base by the length and this product by the constant quantity .7854; this last product will be the solidity required.

EXAMPLE. Required the solidity of a cylinder ADHF, whose length, HD is 13 feet, and diameter of the base AD is 11 feet?

The diameter 11 multiplied by itself and by the length 13 gives 1573, which multiplied by 0.7854 gives the solidity in cubic feet 1235.4342.



PROBLEM IX.

To find the solidity of a Grindstone.

Grindstones in the form of cylinders are sold by the stone of 24 inches diameter and 4 inches thick; the number of stones that any one contains may be obtained by the following rule.

Rule. Multiply the square of the diameter in inches by the thickness in inches, and divide the product by 2304, and you will have the number of

stones required.

Example. Required the number of stones in a grindstone whose diame-

ter is 36 inches and thickness 8 inches?

The square of the diameter 36 is 1298, which multiplied by the thickness 8 gives 10368. This divided by 2304 gives 4.5 or 45 stones, the solidity required.

This Problem may be solved by means of the line of numbers on Gun-

ter's Scale, in a very expeditious manner, by the following rule.

RULE. Extend from 48 to the diameter, that extent turned over twice the same way, from the thickness, will reach to the number of stones required.

Thus in the preceding example, the extent from 48 to the diameter 36, turned over twice, from the thickness 8, will reach to 4.5, or 4½, which is the number of stones sought.

PROBLEM X.

To find the solidity of any Pyramid or Cone.

RULE. Multiply the area of the base by one third of the perpendicular height of the Pyramid or Cone, the product will be the solidity required.

EXAMPLE I. If the Pyramid has a square base, the side of which is 4 feet, and the perpendicular height 6 feet; it is required to determine the solidity?

The area of the base is $4 \times 4 = 16$ square feet, this multiplied by one third of the height or 2 feet gives

32 feet the solidity required.

EXAMPLE II. If the diameter of the base of a cone be 10,6 feet, and the perpendicular height 30 feet; it

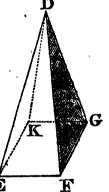
is required to find the solidity?

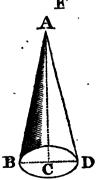
The area of this base was found in Problem IV. equal to 88,247544; this multiplied by one third of the height or 10 feet, gives the solidity required equal to 882,47544 cubic feet.

Having obtained, by the foregoing rules, the number of cubic feet in any body, you may find the corresponding number of tons by dividing the number of cubic feet by 40, which is the number of cubic feet on. Thus the solidity of the above-mentioned cone 882,47544, being divided by 40, quotes \$2,061886, which is the number of tons in that cone.

PROBLEM XI.

By a law of the congress of the United States of America, the tonnage of a ship is to be found in the following manner.





If the vessel be double-decked, take the length thereof from the fore part of the main stem to the after part of the stern-post above the upper deck; the breadth thereof at the broadest part above the main wales, half of which breadth shall be accounted the depth of such vessel; then deduct from the length three-fifths of the breadth, multiply the remainder by the breadth, and the product by the depth; divide this last product by ninety-five, and the quotient will be the true content or tonnage of such vessel.

If the vessel be single-decked, take the length and breadth as above directed, in respect to a double-decked vessel, and deduct from the length three-fifths of the breadth, and taking the depth from the under side of the deck plank to the ceiling of the hold; multiply and divide as aforesaid, the quo-

tient will be the true content or tonnage of such vessel.

EXAMPLE. Suppose the length of a double-decked vessel is 80 feet, and the breadth 24 feet, what is her tonnage? Three-fifths of the breadth 24 feet, is 14.4 feet, which, subtracted from the length 80 feet, leaves 85.6. This multiplied by the breadth 24 feet gives 1574.4, this multiplied by the depth 12 feet (half of 24) gives 18892,8, which divided by 95 quotes the tonnage 198.8.

Carpenters, in finding the tonnage, multiply the length of the keel by the breadth of the main beam and the depth of the hold in feet, and divide the product by 95; the quotient is the number of tons. In double-decked ves-

sels half the breadth is taken for the depth.

GAUGING.

HAVING found the number of cubic inches in any body by the preceding rules, you may thence determine the content in gallons, bushels, &c. by dividing that number of cubic inches by the number of cubic inches in a gallon, bushel, &c. respectively.

A wine gallon, by which most liquors are measured, contains 231 cubic inches. A beer gallon, by which beer, ale, and a few other liquors are measured, contains 282 cubic inches. A bushel of corn, malt, &c. contains 2150.4 cubic inches: this measure is subdivided into 8 gallons, each of which contains 268.8 cubic inches.

In all the following rules, it will be supposed that the dimensions of the body are given in inches, and decimal parts of an inch.

PROBLEM I.

To find the number of gallons or bushels in a body of a cubic form.

RULE. Divide the cube of the sides by 231, the quotient will be the answer in wine gallons; or by 232 and the quotient will be the answer in beer gallons; or by 2150.4, and the quotient will be the number of bushels.

EXAMPLE. Required the number of wine gallons contained in a cubic cistern, the length of whose side is 62 inches? Multiplying 62 by itself and the product again by 62, gives the solidity 258528; which divided by 251 gives the content 10812 wine gallons.

PROBLEM II.

To find the number of gallons or bushels contained in a body of the form of a rectangular Parallelepiped. See the figure of Problem VII. of Mensuration.

RULE. Multiply the length, breadth, and depth together; divide this last product by 231, for wine gallons; by 232 for beer gallons; and by 2150.4 for bushels.

EXAMPLE. Required the number of wine gallons contained in a cistern ABCDFGHE (see fig. Prob. VII. of Mensuration) of the form of a parallelepiped, whose length EF is 66 inches, its breadth GF 35 inches, and its depth FD 24 inches. Multiplying the length 66 by the breadth 35 gives 2310, this multiplied by the depth 24, gives the solidity 55440; which divided by 231 quotes 240 wine gallons.

PROBLEM III.

To find the number of gallons or bushels contained in a body of a cylindrical form.

RULE. Multiply the square of the diameter by the height of the cylinder, and divide the product by 294.12, the quotient will be the number of wine gallons; if you divide by 359.05 the quotient will be the number of ale gallons; and if you divide by 2738, the quotient will be the number of bushels.

Note. These divisors are found by dividing 231, 232, and 2150.4 by

.7854.

Required the number of wine gallons contained in the cylinder AFHD (See the fig. of Prob. VIII. of Mensuration) the diameter AD of its base being 26 inches, and length HD 18 inches? The diameter 26 multiplied by itself gives 676, this multiplied by the length 18, gives the solidity 12168, which divided by 294.12, gives the answer 41 wine gallons.

PROBLEM IV.

To find the number of gallons or bushels contained in a body of the form of a pyramid or cone. (See figures of Problem X. of Mensuration.)

RULE. Multiply the area of the base of the pyramid or cone by onethird of its perpendicular height; the product divided by 231 will give the answer in wine gallons; if divided by 282, the quotient will be the number of beer gallons; and if divided by 2150.4, the quotient will be the number of bushels.

Required the number of beer gallons contained in a pyramid Example. DEFGK (See fig. Prob. X. Example I.) whose base is a square EFGK, a side of which, as EF, is equal to 30 inches, and the perpendicular height of the pyramid is 60 inches? The square of 30 is the area of the base 900, this multiplied by one-third of the altitude 20, gives the solidity 18000, which divided by 282, gives the answer in beer gallons 63.8.

PROBLEM V.

To find the number of gallons or bushels contained in a body of the form of a frustrum of a cone. (See the figure below.)

Multiply the top and bottom diameters together, and to the product add one-third of the square of the difference of the same diameters; multiply this sum by the perpendicular height, and divide the product by 294.12 for wine gallons, by 359.05 for ale gallons, and by 2738 for bushels.

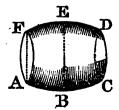
EXAMPLE. Given the diameter DC of the bottom of a frustrum of a cone 36 inches, the top diameter AB=27 inches, and the perpendicular height, FE 50 inches. Required the content in wine gallons? The product of the two diameters 36 and 27 is 972; their difference is 9, which squared and divided by 3 gives 27; this added to 972 gives 999, which multiplied by the height 50 gives the solidity . 49950; this divided by 294.12 quotes the content in wine gallons 169.



PROBLEM VI.

To gauge a cask.

To gauge a cask, you must measure the head diameters FA, DC, and take the mean of them when they differ; measure also the diameter EB at the bung, (taking the measure within the cask); then measure the length of the cask, making due allowance for the thickness of the heads. Having these dimensions, you may calculate the content in gallons or bushels, by the following rule.



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Rulz. Take the difference between the head and bung diameters, multiply this by .62 and add the product to the head diameter, the sum will be the mean diameter; multiply the square of this by the length of the cask, and divide the product by 294.12 for wine gallons, by 359.05 for beer gallons, and by 2738 for bushels.

The quantity .62 is generally used by gaugers in finding the mean diameter of a cask; but if the staves are nearly straight, it would be more accurate to take 55 or less;* if, on the contrary, the cask is full on the quarter, it

would be best to take .64 or .65.

EXAMPLE. Given the bung diameter EB=34.5 inches, the head diameter FA=DC=30.7 inches, and the length 59.3 inches; required the number of wine gallons this cask will hold? The difference of the two diameters 54.5, and 30.7, is 3.8; this multiplied by .62 gives 2.4 nearly, to be added to the head diameter 50.7 to obtain the mean diameter 53.1. The square of 55.1 is 1095.61, this multiplied by the length 59.3 gives the solidity 64969.673, which, divided by 294.12, quotes the content in wine gallons 220.8.

To gauge a cask by means of the line of numbers on Gunter's Scale, or on the calipers used by gaugers.

Make marks on the scale at the points 17.15, 18.95, and 52.35, which numbers are the square roots of 294.12, 359.05, and 2738, respectively. A brass pin is generally fixed on the calipers at each of these points, which are called the gauge points. Having prepared the scale in this manner, you may

calculate the number of gallons or bushels by the following

Rule. Extend from 1 towards the left hand to .62, (or less if the staves be nearly straight) that extent will reach from the difference between the head and bung diameters, to a number to the left hand, which added to the head diameter will give the mean diameter; then put one foot of the compasses upon the gauge point—which is 17.15 for wine gallons, 18.95 for ale gallons, and 52.33 for bushels—and extend the other to the mean diameter; this extent turned over twice the same way, from the length of the cask, will give the number of gallons or bushels-respectively.

In the preceding example the extent from 1 to .62 will reach from 3.8 to

2.4 nearly, which added to 30.7 gives the mean diameter 33.1.

Then the extent from the gauge point 17.15 to 38.1, turned over twice

from the length 59.3, will reach to 220.8 wine gallons.

If you had used the gauge point 18.95, the answer would have been in ale gallons; and if you had used 52.35, the answer would have been in bushels.

SURVEYING.

LAND is generally measured by a chain of 66 feet in length, divided into 100 equal parts called links, each link being 7.92 inches.

A pole or rod is 164 feet, or 25 links, in length; hence a square pole con-

tains 2721 square feet, or 625 square links.

An acre of land is equal to 160 square poles, and therefore contains 43580

square feet, or 100,000 square links.

To find the number of square poles in any piece of land, you may take the dimensions of it in feet, and find the area in square feet, as in the preceding Problems; divide this area by 43560, the quotient will be the number of acres: or by 272.25, and the quotient will be the number of square poles. If the dimensions be taken in links, and the area be found in square links, you may obtain the number of acres by dividing by 100000 (that is, by crossing off

^{*} In the example to Problem V. preceding (which may be esteemed as the half of a highered with staves perfectly straight) the multiplier is only 51. For this multiplied by 9, the difference between AB and DC, produces 4.50 or 4.5 nearly, which, added to 27, and the sum 51.8 squared, multiplied by 50 and divided by 394.12 quotes 189 gallons nearly.

the five right hand figures;) and the number of square poles may be obtained by dividing by 625.

PROBLEM I.

To find the number of acres and poles in a piece of land in the form of a rect-

angular parallelogram.

Multiply the base by the perpendicular height, and divide by 625, if the dimensions were taken in links, but by 272.25, if they were taken in feet; the quotient will be the number of poles, which, divided by 160, will give the number of acres.

EXAMPLE I. Suppose the base DB (see the figure of Ex. I. Prob. I. of Mensuration) of the rectangular parallelogram ACBD is 60 feet, and the per-

pendicular BC 25 feet; required the area in poles?

The product of the base 60 by the perpendicular 25 gives the content 1500 square feet, and by dividing by 272,25, we obtain the answer in square poles 5.5.

PROBLÉM II.

To find the number of acres and poles in a piece of land in the form of an oblique-angular parallelogram. (See the figure of Prob. I. Ex. IV. of Men-

suration.)

This area may be found in exactly the same manner as in the RULE. preceding Problem, by multiplying the base AD by the perpendicular height BE, and dividing by 625, when the dimensions are taken in links, but by 272.25 when taken in feet; the quotient will be the answer in poles, which divided by 160 will give the answer in acres.

Suppose the base AD is 632 links, and the perpendicular BE

328 links; required the number of poles?

Multiply the base 632 links by the perpendicular 326 links, the product 206032 divided by 625, gives the answer in poles 329.

PROBLEM III.

 T_0 find the number of acres and poles in a piece of land of a triangular form.

RULE. Multiply the base by the perpendicular height, and divide the product by 1250 when the dimensions are given in links, but by 544.5 when they are given in feet; the quotient will be the answer in poles.

Note. Instead of dividing by 1250, you may multiply by 8, and cross off

the four right hand figures.

EXAMPLE. Given the base AC (see fig. of Prob. II. of Mensuration) equal to 300 feet, and the perpendicular BD 150 feet; required the area in poles? Multiply the base 300 by the perpendicular 150, the product 45000, divided by 544.5, quotes the answer in poles 82.6.

PROBLEM IV.

To find the number of acres and poles in a piece of land of any irregular right-lined figure.

Rule. Find the area as in Problem III. of Mensuration, by drawing di-

agonals, and reducing the figure to triangles: the base of each triangle being multiplied by the perpendicular, (or by the sum of the perpendiculars falling on it) and the sum of all these products divided by 1250 when the dimensions are given in links, but by 544.5 when in feet, will give the area of the figure in poles.

EXAMPLE. Suppose that the piece of land is of the same form as the figure in Prob. III. of Mensuration, and that EB=22 feet, EC=33 feet, AF=15 feet, BG=14 feet, and DH=12 feet: it is required to find the area in poles? The product of EB 22 feet, by AF 13 feet, gives double the triangle EAB 286 square feet; and the diagonal EC 33 feet, multiplied by the sum of the perpendiculars BG, DH, 26 feet, gives double the figure BCDE, 858 square feet; the sum of this and 286, divided by 544.5 gives the area 2,1 or 21 poles.

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To find the content of a field by the Table of difference of Latitude and Departure.

This method is simple and much more accurate than by projection, the boundaries being straight lines, whose bearings and lengths are known. The rule for making these calculations is as follows.

RULE.

1. Begin at the western point of the field, as at the point A in the figure-Prob. III. of Mensuration, for a point of departure; and mark down in succession the bearings and lengths of the boundary lines AB, BC, &c. as courses and distances in a traverse table. Find the corresponding differences of latitude and departure by Table I. or II. (or by logarithms) and enter them in their respective columns N. S. E. W. as in the adjoined Table.

2. Find the departures or meridian distances of the points B, C, &c. from the point A, by adding the departures when east, but subtracting when west, and mark them respectively against the bearings, in the column of

meridian distance.

S. Place in the first line of the column M the first meridian distance 16.1, and in the following lines, the sum of the meridian distance which stands on the same line and that immediately above it. Thus on the second line I put 52.1 which is equal to the sum of 16.1 and 36.0. On the

Courses	Dis.	N.	8.	E.		Mer Dis	M.	North Areas.	South Areas.
N. 58° E. E. 6 S	19. 30.	10. 1		16 1 19 9	_	16.1 36.0	16.1 52.1	162.61	109.41
6 17 W. W.	20 20.	Ì	19. 1		5,8 20.0		66.2 #0 4		1264.42 0
N. 42º 35' W	15.1	11.1			10.2	0.0	1 .2	113.22	
		21 2	21. 2	36	36.0			2 75.85	1373.83 275.83
								Half.	:098. 549.

4. Multiply the numbers in the column M by the differences of latitude in the same horizontal line, and place the product in the column of areas marked north or south, according as the difference of latitude is north or south. Thus in the first number in the column M is 16.1 which multiplied by the corresponding difference latitude 10.1 N. produces the north area 162.61. The second value of M, 52.1, multiplied by the second difference of latitude 2.1 S. produces the south area 109.41. The third values 66.2 and 19.1 S. produces the south area 1264.42. The fourth difference of latitude is 0, which multiplied by the fourth meridian distance 40.4 produces 0 for the corresponding area, which is the case whenever the bearing is east or west, &c.

5. Add up all the north and all the south areas; half their difference will be the area of the field in square measures of the same name as those made use of in measuring the lines, whether feet, links, or chains, &c. Thus the sum of all the north areas is 275.23, the south 1373.83; their difference is 1098, half of which is 549 square feet the area of the given field.

It may be observed that the bearings and lengths of the boundary lines in this example are not exactly the same as those in Problem III. of Mensuration, which is the reason of the difference in the area above calculated and

that found in Problem III. by dividing the field into triangles.

If it had been thought necessary, the differences of latitude and departure might have been taken to one decimal place farther, by entering the table with ten times the length 19, 20, &c. and taking one-tenth of the corres-

ponding differences of latitude and departure.

In the above calculations we have supposed the survey to have been made with accuracy, in which case the sums of the differences of latitude in the columns N. S. ought to be equal to each other, also the sums of the department in the columns E. W. This is the case in the above example where

the sum of the Diff. of Lat. is 21.2, and the sum of the departure 38.0: but it more frequently happens that the numbers do not agree; in which case the work ought to be carefully examined, and if no mistake can be found, and the error is great, the place ought to be surveyed again; but if the error be small, it ought to be apportioned among all the differences of latitude and departure, in such manner as to produce the required correction with the least possible changes in the given numbers. The method of doing this was explained by me in the fourth number of the Analyst, in answer to a prize question of Professor Patterson, and is as follows. Find the error in latitude, or the difference between the sums of southing and northing, also the sum of the boundary lines, AB, BC, &c. Then say, as this sum is to the error in latitude, so is the length of any particular boundary to the correction of the corresponding difference of latitude, additive if in the column whose sum is the least, otherwise subtractive. The corrections of the departure are found by the same rule, except changing diff. of lat. into departure. Thus in the adjoined example, the sum of the boundary lines is 161.6, the error of latitude is 0.10 and of departure .08,

Begrings.	Langtha	N.	· .	E. W.		Согте	tions.	C	orrecte	d valu	CS.
manue.	Congras					N.	E.	N.	8.	E.	W.
N. 45° E. S. 30 W. S. 5 E. W. N. 20 E.	60. 25. 36. 29.6 31.	20.28 29.13 57.41	21.65 33.86 87.51 57.41	28.28 3.14 10.60 42.02	12.50 29.60 42.10 42.02	.10	23 42 43 43 43 43 43 43 43 43 43 43 43 43 43	28.50 0.69 29.15 57.47	21.63 35.84 87.47	28.50 3.16 10.62 42.98	12.49-29.60
		Error	.10	Errer	-08	<u> </u>			<u> </u>	1	

and the corrections of the diff. of lat. and departure are found by the following proportions:

Lat.	Dep.
161 6 : 0.10 : : 40 : 0.02*	16.16 : 0.08 : : 40 : 0.02
:: 25 :0.02	::25 :001
:: 96 : 0.02	::36 ;0.62
: : 29.6 : 0.0 2	z : 29.6 ; 0.01
:: 31 : 0.02	::31 :0.02

The first correction of lat. .02 is to be added to the first latitude, 28.28. because it is in the column whose sum 57.41 is less than the other 57.51, so that the first corrected diff. of lat. is 28.50. The second is the difference between 21.65, and the second correction .02, because 21.65 is in the greatest column, the corrected value is therefore 21.63. The third is found in the same manner to be \$5.86—.02=35.84. The fourth corrected difference of latitude is simply the fourth correction .02 placed in the column N, because the sum in that column, 57.41 is the least and the fourth diff. of latitude in the original table is 0. The fifth is the sum of 29.13 and the fifth correction 0.02, making 29.15. These are placed in their proper columns in the corrected values. In a similar manner the first departure is equal to the sum of 28.28 and the first correction 0.02, which is equal to 28.30. The second is the difference between 12.50 and the second correction .01, making 12.49; and so as for the others, taking the sum when the departure is in the column whose sum is the least, which in the present case is the east, and the difference when in the other column. In the traverse table thus corrected, the sum of the differences of latitude is 57.47 in both columns, and the sum of the departures 42.08. Having corrected the values of this traverse table, you must find the meridian distances, the column M, the north and south areas, &c. as in the former example.

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^{*} The boundary lines in this example are so nearly of an equal length that the correction of the difference of latitude (taken to the nearest decimal) is 0.02 for each of them, but in general, they will be frifferent. The table of difference of latitude and departure may be made use of in finding these corrections, thus: seek in the table till the first term 161.5 (or 162) is found in the distance column to corgraphed to the second term .10 (or 10) in the departure column, thus opposite the third term 40, 25, 36, -48. will be the sought corrections, as is evident.

SURVEYING. W.W. 6 Miles.

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In projecting a survey of this kind, where there is a small error, you must plot off as usual the boundary lines AB, BC, CD, &c. and it will be found that the termination of the last line EA will not fall exactly in the point A, but will be at a point near it, which we shall call a. To correct this error you must draw through the points B, C, D, &c. lines parallel to aA, in the direction from a to A, of such lengths as to be to Aa, as the distances of those points respectively from A (measured on the boundary ABCD, &c.) are to the whole length of the boundary line, through these points draw the corrected lines terminating on A. Want of room prevents entering into farther explanation of the subject.

The manner of surveying Coasts and Harbours.

From what has been said in the preceding problems, the intelligent reader will readily perceive the method of surveying a coast or harbour; but as this is an important subject, it was thought proper to enter more fully into an explanation of the different methods to be made use of.

To take a draught of a coast in sailing along shore.

Having brought the ship to a convenient place, from which the principal points of the coast, or bay, may be seen, either cast anchor, if it is convenient, or fie-to as steady as possible; or if the coast is too shoal, let the observations and measures be taken in a boat. Then while the vessel is stationary, take with an azimuth compass the bearings in degrees of such points of the coast as form the most material projections or hollows;* write down these bearings, and make a rough sketch of the coast, observing carefully to mark the points whose bearings are taken, with letters or numbers, for the sake of reference.

Then let the ship or boat run in a direct line, which must be very carefully measured by the log, or otherwise, one, two, or three miles, until she comes to a situation from which the same points before observed can be seen again with quite different bearings: then let the vessel lie steady as at the former station, and observe again the bearings of the same points, and make a rough sketch of the coast; this sketch may be made more accurately while the

vessel is running the base line.

To describe the chart from these observations, you must in some convenient part of a sheet of paper draw the magnetic meridian, and lay off the several bearings taken at the first station, marking them with their proper letters or numbers; lay down also the bearings taken from the second station. Draw a line to represent the ship's run both in length and course, and from that end of the line expressing the first station, draw lines parallel to the respective bearings taken from that end; also from the other end shaw lines parallel to the bearings taken at that end, and note the intersection of each pair of lines directed to the same point; and through these intersections draw by hand a curved line, observing to wave it in and out as near as can be like the tending of the coast itself. Then mark off the variation of the compass from the north end of the magnetic meridian towards the right hand, if it be west, or towards the left hand if it be east, and draw the true meridian through that point and the centre of the circle.

Against each part draw the appearance of the land marked in the sketches, distinguishing the rocky shore, highland, beach, &c. as in Plate V. or XI. Thus the sand beaches may be marked as in Plate XI. fig. 8, and the rocky shore as in fig. 9, &c. Put in the several soundings at low water,† in small figures, distinguishing whether they are fathoms or feet; show the time of high water on the full and change days by Roman figures, and tell the rise of the tide in feet. The direction and velocity of the flood tide are to be observed, which may be done by heaving the log when the ship or boat is at

^{*} In taking the bearings, if the vessel has much motion, the mean of several observations should be taken.

If the soundings were not taken at low water, they may be reduced thereto by a method which will be explained hereafter.

anchor, and the direction is to be represented by an arrow. Put in a compass and a scale of miles or leagues such as the vessel's run was laid down by; add the name of the place, and the latitude and longitude as true as can be obtained.

If there are shoals or sands on the coast, let them be taken in a boat, sailing round them, keeping account of the courses, distances, and soundings.* But to put them in the draught, the observer in the boat must take the bearings of two points on the coast (the bearings of which have been taken from the ship) from some part of each sand or shoal, so sailed round; or, the bearing of the boat at some part of the shoal, or of some beacon in that place, must be taken by the ship at each of the stations where the bearings of the shore were taken from the ship; for by either of these means, one point of the sand being obtained, the rest of it can be laid down from the observations taken in the boat. Rocky shoals may be marked on the chart, as in Plate XI. fig. 11, and sand banks as in fig. 10.

If the coast to be drawn is a bay or harbour winding in such manner that all its parts cannot be seen at two stations, let as many bases or lines be run and measured exactly as may be found necessary, observing that the several distances run should join to one another, in the nature of a traverse; that each new set of objects, or points observed, should be taken from two stations at the ends of a known distance, and that the objects whose bearings are taken do not so much extend beyond the limits of the base as to make angles with it less than about \(\frac{1}{2} \) of a point, but rather reserve such objects for the next measured base line; for \(\frac{1}{2} \) when lines lie very obliquely to one

another, their intersections are not easily ascertained.

If any particular parts of the harbour cannot be conveniently seen from rither of the stations, take the boat into those places, and having well examined them, made sketches thereof, estimating the lengths and breadths of the several inlets, either by the rowing or sailing of the boat, take as many bearings, soundings, and other notes, as may be thought necessary; then annex these particular views in their proper places, in the general draught.

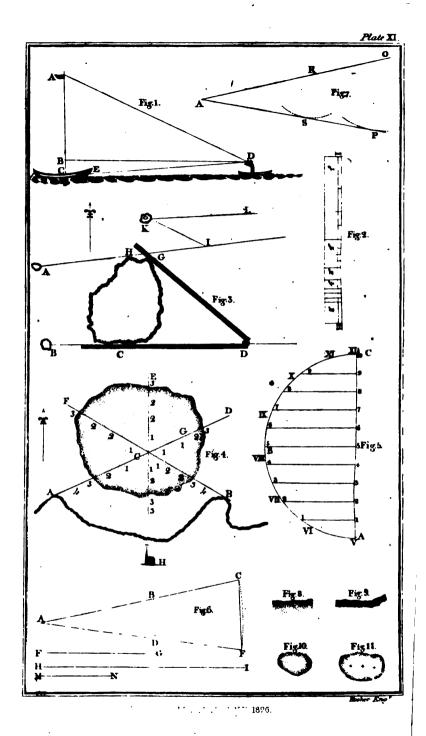
If there are any dangerous sands or rocks, besides inserting them in their proper places, you must see if there be any two objects ashore (such as a church, mill, house, noted cliff, &c.) which appear in the same right line when on the shoal; and these objects must be noted on your chart. If none can be found, you must take the bearings of some remarkable points, and note them on your chart; by which means it will be known how to avoid the danger.

It should be remarked in the draught the kind of bottom obtained in sounding, whether mud, sand, shells, coral, rocky ground, &c.; and where there is good anchorage draw the figure of an anchor. Also, if there is any particular channel more convenient than another, it is to be pointed out by lines drawn to its entrance from two or more noted marks ashore.

The positions of objects taken by a magnetic compass being liable to great uncertainties, as is well known to those who have had any experience, especially at sea; it has therefore been recommended to observe only the bearings of the station lines by the compass, and then measure the angles which the other objects make with these lines, by a quadrant or sextant, which for this purpose must be held in an horizontal position.

EXAMPLE I. (See fig. 1, Plate X. Suppose a ship at A observes the bearings of the most remarkable point of a bay, C, D, E, F, G, H, and I, and sails S. 64° E. 1½ miles to B; at B

^{*} It is difficult to ascertain correctly the courses and distances sailed by the boat, on account of the currents and other causes. This inconvenience may be obviated if the ship be at anchor, and not far from the boat, by observing in the boat the bearing of the ship by compass, and by measuring, with a quadrant, the angle contained between the top-gallant-mast head, and that part of the ship which is at the same height as the eye of the observer; for by this angle the distance of the boat from the ship may be determined, as will be explained in this work.



she also observes the bearings of the same points; hence it is required to construct the chart.

Bearing of	C from A	S. 36° W.	Bearing of C from	BS.	89° E.
0	D	N. 90 W.	D		48° W.
	E	N. 26° E.	E	N.	24° W.
	F	N. 55° E.	F	N.	13° E.
	G	East.	G	N.	47° E.
	H	S. 40° E.	H	S.	38° W.
	ī	S 19° E.	I	8.	46° W.

Draw the line AB, S. 64° E. 14 miles. Through the points A and B draw the lines AC, AD, AE, AF, AG, AH, AI, BC, BD, BE, BF, BG, BH, and BI, at their respective bearings, and where the corresponding lines cut each other, will be the points C, D, E, F, G, H, and I, required; through which the different curvatures of the land must be drawn, corresponding with your eye-draught. In this manner may a chart be constructed by observations taken upon the water. The manner of surveying upon land is exactly similar.

To survey a harbour by observations on shore.

Make an eye-draught of the place to be surveyed, and in going round the coast fix station staves, or straight poles, tall enough to be seen at a considerable distance, in the most remarkable points and bendings of the shore; but if at any of those places there is a noted tree, house, or any other remarkable thing, that object may serve instead of a station staff; and it will be convenient to black the staves, and tie a piece of white bunting at the top of each; then in the eye-draught put letters or numbers at the noted points or marks for the sake of distinction.

Choose the most extensive and level spot of ground you can meet with to measure your base line upon, which should not be less than a tenth part of the distance of the two extreme objects which are to be observed; and let the direction of the measured base line be such, that as many of the station staves as possible may be seen from each end of it. The bearing or position of this base must be well determined in degrees and minutes, and the length accurately measured, either by a measuring chain or a piece of log-line.

From each end of the base observe, with an azimuth compass, or with a Theodolite, if it can be procured, the bearings of each of the station staves; or else with a sextant measure the angles contained between the staves or remarkable objects, and the other end of the station line, and write them down in order in your book. These measures and angles being plotted down as before directed, will give the most conspicuous points of the shore, the intermediate spaces are to be filled up from the sketches made on the

But if either of these objects should spread on either hand so far beyond the limits of the base, that at either end thereof the other end and those objects should appear nearly in the same direction, or to make angles not exceeding 10°: or, if some of the remarkable objects could be seen only from one end of the base; then let the bearings of such objects be taken from a place whose position has been determined from both ends of the measured base; or, if there are several remarked objects which cannot be seen from either end of the base lines, let the bearings of such objects be taken from each of two points whose positions have been determined by bearings taken from both ends of the base, or it may, on some occasions, be proper to choose another place on which another base of a convenient length may be measured, and from the extremities of which the ends of the first base may be seen; and as many as can be of the remaining objects which lay too obliquely, or which could not be seen from the first base. In such manner proceed until the bearings are taken of all the points judged necessary for completing the survey of the limits of the harbour.

If a right line of a sufficient length for a base line cannot be measured, it may be taken in two adjoining lines as the two sides of a triangle, the included angle being accurately taken, and the bearing of either line OOGIC

When the outlines or limits of a harbour, bay, road, &c. are delineated, by the preceding precepts, let a small vessel go out to sea to take drawings of the appearance of the land and its bearings; sail likewise into the harbour, and draw the appearance of its entrance; take particular notice if there be any false resemblance of the entrance, by which ships may be deceived and run into danger; or when any two objects being brought in a line, or in one, will lead into the harbour without danger; search for the best anchoring places, and if possible, denote those places by bringing two objects in one, if not, take the exact bearings of two or three other objects, so that the places may be easily determined. The chart being correctly drawn, a compass, with the variation, and scale properly fitted to the plan, the islands, rocks, sands, &c. must be marked in their proper places, with their soundings at low water, the anchoring places, with the best track to get to them; the proper sailing marks to avoid dangers; the places where fresh water can be got; the name of the place, the country in, on what sea; the latitude and longitude; a sketch of the appearance the place makes at sea upon a known bearing, and at an estimated distance; and whatever else a judicious seaman shall think proper may be inserted. Then will the plan be fit for all nautical purposes, and may be embellished with proper colours if necessary.

EXAMPLE II. (See fig. 2. Plate X.)

From each end of a base line AB of 1200 fathoms, were observed the points C, D, E, F, and G; and as the points I, K, and L, were not visible from the extremities of the base line, another base line was measured from the point D to H of 650 fathoms, from which points the bearings of I, K, and L, were observed: hence it is required to construct a chart of the place.

Bearing of B from A East.	Bearing of C from B N. W. b. W.
C North.	D N. N. W.
D N. E. b. N.	E North.
E N. E. 1 N.	F N. b. E.
F. N. E. b. E. 1 E.	G N. E.
G E. b. N. 1 N.	,
Bearing of H from D N. W.	Bearing of I from H N. E. by N.
1 N. b. W.	K N. E. 4 E.
K N. b. E. & E.	L E. N. E.
L N. N. E. 🖟 E.	

Draw the east line AB=1200 fathoms; from each end of this line draw the lines AC, AD, AE, AF, AG, BC, &c. at their respective bearings; the points of intersection will give the points C, D, E, F, and G. From the point D (which was found in this manner) draw the N. W. line DH=650 fathoms: and through these points draw the lines DI, DK, DL, HI, &c. at their respective bearings; the points of intersection of the corresponding lines will be the situation of the points I, K, L. Between these remarkable points, draw the outlines of the land, conformable to your rough draught.

In order to determine the situation of the point M, which was seen too obliquely from the bases AB, DH, you may take the bearing of that point from B, and then from G (whose situation has been determined by bearings taken from the points A, B,) the intersection of the lines BM, GM, will de-

termine the situation of M.

Method of surveying a small bank or shoal where great accuracy is required. The method of determining the extent and situation of shoal ground by sailing round it and keeping an account of the courses and distances sailed, is well adapted to the taking of an extensive survey, or to the exploring of a large bank, where great accuracy is not required; but the difficulty of ascertaining with precision the courses and distances sailed (which are liable to error on account of the tides, currents, and the different velocity of the boat at different times owing to the unsteadiness of the wind) prevents this method from being sufficiently accurate to he used in exploring a dangerous shoul or bank at the entrance of a narrow channel of a harbour, or any other place where the exact form of the shoal is to be found; and if to ob-

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tain the necessary degree of correctness, the bearings of two remarkable objects were taken at every time of sounding, the time expended in taking the observations would be increased beyond all reasonable bounds. viate these difficulties, the following methods may be made use of, by either of which the necessary observations for determining the situation of the boat. may be made as fast as the soundings can be taken.

First Method. Procure a large sail-boat with a high mast, and a small row-boat; bring the sail-boat to anchor on the bank which is to be explored. and take accurately the bearings of two remarkable points of land, or other objects whose situation has already been determined by observations taken on shore, or in sailing along the land; by this means the situation of the sailboat may be accurately marked on the chart; then enter the small boat and row from the other in any particular direction, observing to keep the mast of the boat to bear upon any point of the compass, or (which is much more accurate) to keep the mast of the boat to range on any particular point of land, or other object marked on the chart; by these means any errors which might arise in the course of the boat will be entirely prevented. While proceeding in this direction let one person take the soundings, while another observes with a quadrant or sextant the angular elevation of the top of the boat's mast above the horizontal line drawn from the eye of the observer, and a third person notes the observations in the minute book, and the time of observation, in order to make the necessary reduction in the soundings to reduce them to low water. Proceed in this manner from the sail-boat till you get off the bank into deep water, or till the elevation of the mast is not much less than one degree, then row across the bank till the bearing of the mast is altered considerably, or till it appears in a range with another point of land at a considerable angular distance from the point with which the mast ranged in the first observations; then row towards the boat, sounding and observing the angular elevation of the mast as before. Proceed in this manner in sounding to and from the sail-boat till you have procured a suffi-cient number of soundings in every direction. Then go on board the sailboat and shift her birth to another part of the bank where soundings have not been taken and proceed to sound as before. Continue sounding and shifting the situation of the boat, till the whole bank has been explored, and then the observations may be plotted off by the directions in the following

Let ABC (Plate XI. fig. 1.) be the mast of the sail-boat, D the situation of the eye of the person who observes the angular elevation of the mast. Draw the line DB parallel to the horizon and join DA. Then the height AB must be measured accurately, and that being given and the observed angle ADB the corresponding distance BD may be obtained by the usual rules of trigonometry by saying as radius : AB : : co-tangent ADB : BD. Thus.if the height AB was 50 feet, and the angle ADB 10, the distance BD would be 1719 feet (being 57,3 times as great as AB). The distances corresponding to 2°, 3°, &c. are given in the adjoined table, by examining which it will appear that the distance BD corresponding to any angle ADB (less than 30°) may be obtained nearly by dividing 1719 by the angle ADB in degrees. Thus for 4 degrees the distance by feet. 101719 859 3 572 this rule would be 1410 = 4291, nearly, as in the table. The great-429 est difference between the distances determined by the rule and by 170 the table being 5 feet, corresponding to the angle 500, for 1719 = 10 82 57; whereas, by the table the distance is 52. In taking soundings by this method it will be very rarely necessary to measure an angle so great as 300, so that for all practical purposes the distance may be de-

^{*}A mark may be made at B, and a vane placed at the top of the mast at A, to enable the observer to distinguish those objects when at a great distance. If the height of the observer above the horizon be small in comparison to the height of the mast, the angular distance ADE between the surface of the sea and the top of the boat's mast might be measured, instead of ADE; for W the distancer BC and CE remained the same in all observations, it would be immaterial which angle was

termined in this example to a sufficient degree of accuracy by dividing 1719 by the observed angular elevation in degrees. On these principles we have the following rule for calculating the distance corresponding to a mast of any given height, and to any observed angular elevation.

Multiply the height of the mast above the eye of the observer by 57.5, and the product will be a constant quantity* which being divided by the observed angle of elevation expressed in degrees and decimals of a degree, the quotient will be the sought distance nearly.

If the height of the mast be expressed in equal parts taken from the scale by which the chart was plotted off, the distances found by the above rule would be expressed in the same equal parts; so that if the distances thus expressed corresponding to 1°, 2°, 3°, &c. were calculated and marked on a slip of paper (fig. 2, Plate XI.) from H to 10, from H to 20, from H to 30, &c. respectively, the slip H I thus marked, would be a very convenient scale for

plotting off such distances.

For further illustration of this method, we have given an example in fig. 4, Plate XI. in which C represents the place at which the sail-boat was at anchor, A and B the points observed, in order to ascertain her position on the chart, by drawing thereon the lines AC, BC, in opposite directions to the bearings of the points A, B, observed from the boat; for the point of intersection C will evidently be the place of the boat upon the chart. Suppose now that in the first set of observations the mast of the sail-boat was made to range on the point A, in this case the course of the boat must be on the continuation of the line AC towards D, then the slip H I (fig. 2, Plate XI.) is to be laid upon the line CD (fig. 4, Plate XI.) with the point H upon C, and the angular elevation being found on the slip the sounding corresponding (reduced to low water) is to be marked on the line CD, immediately under the mark on the slip. Thus if the angle was 4°, the point corresponding would Having plotted off the soundings taken in the direction CD, proceed in the same manner with the others, viz. those in the direction CE found by keeping the boat's mast in a range with the church at H, those in the direction CF found by keeping the boat's mast in a range with the point B, those in the direction CA found by keeping the mast to bear E. N. E. and so on with the other observations; and when all the soundings are marked on the chart, dotted lines are to be made round the shoal soundings, and thus the true figure of the shoal part of the bank will be obtained.

This method I have frequently used in taking a survey of the part of the coast of Massachusetts' Bay, included between Manchester and Lynn. The height of the mast of the boat used on the occasion was about 30 feet, and it was found that distances less than a third of a mile could be obtained in

this manner to a great degree of precision.

This method of determining the place where soundings Second Method. were taken consists in keeping (while sailing in a boat and sounding) a particular point of land, or any other object, to bear always in the same direction, and measuring with a quadrant or sextant held in a horizontal position the angular distance between that object and another object (making a considerable angle with the former) for by this means the situation of the boat at the time of sounding may be determined. Instead of bringing the object to bear upon a particular point of the compass, you may (when it can be

If AB represented the known vertical height of the summit of an island above the eye of an observ-er, the distance from the island might be determined by measuring the angular elevation ADB, as is evi-

measured; observing however that different scales must be used for plotting off the angles ADB and ADE

er, the distance from the Island might be determined by measuring use august everywhear, — dent from what has been said above.

This constant quantity may be determined without actually measuring the altitude AB, if the angular elevation can be measured at a place D where the distance BD is known. Thus is the example (Plate XI. fig. 4) the distance AC being known, and the angular elevation of the mast at O being cohern-ced at A in degrees and decimals of a degree, and multiplied by the distance AC, the product will be the constant quantity mentioned in the rule. This method may be used in determining the distance from an island by the method mentioned in the last note.

done) bring the object in a range with another remarkable object, and by this means you will avoid the error which might arise from the use of a

compass.

For an example of this method, suppose that a survey of the small islands A, B, K, (Plate XI. fig. 3) and the large one CHG had been taken and plotted off as in the figure. Then soundings may be taken in the direction BCD by bringing the small island B in a range with the southern part of the great island, and measuring the angle CDG formed by the extremes of the great island; or by keeping the small island A to range with the northern part of the great island, and measuring the angle HIK formed by the northern extreme of that island, and the small island K; or by running in the direction KL so as to keep the island K to bear W. 4 S. and measuring the angle formed by that island and the northern extreme of the great island, &c.

The method I have generally used for plotting off such angles is by means of a sector; and as that instrument is more easily procured than others better adapted to the purpose, I shall explain the method by showing how the angle CDG, measured as above, may be plotted off so as to determine the point D where that angular distance was observed. To do this, you must draw the line DC, and open the sector till the two legs form with each other an angle equal to the observed angle CDG, then slide one leg of the sector on the line DC till the other leg touches the northern extreme of the island at the point G, and the point directly under the centre of the joint of the sector will be the point of observation: as this point cannot be exactly marked on account of the size of the joint of the instrument, you may mark with a pencil on the line DC the top points where the circumference of the joint touches that line, and note the sounding in the middle between those two marks.

If a quadrant of a circle be described on a piece of paper, with a radius equal in length to one of the legs of the sector, and then divided into 90°, the sector may, by means of that quadrant, be opened to any angle in a very expeditious manner.

This method of obtaining distances when sounding, I have frequently used

with success.

To reduce soundings taken at any time of the tide to low water.

The soundings at low water are always to be marked on a chart, and if they are taken at any other time of the tide, a proper allowance must be made to reduce them to low water. This allowance may be made if the whole vertical rise of the tide from low to high water be known, and the time of high and low water, as in the following example.

Suppose the vertical rise of tide from low to high water, to be 10 feet, the time of low water 5h. A. M. and the time of high water 11h. 30m. A. M.; required the allowance to be made on an observation taken at 8 A. M.?

Draw the line AC (Plate XI. fig. 5) and make it equal to the whole rise of the tide 10 feet, taken from any scale of equal parts, and divide the line into equal parts representing feet, at the points 1, 2, 3, &c. to 10, the mark 10 (corresponding to the whole rise of the tide) being at the point C, and through these points draw lines 11, 22, 33, &c. perpendicular to AC, to meet the circumference of a circle drawn on the diameter AC. Divide the semi-circumference ABC of that circle into a number of equal parts representing the number of hours elapsed from low to high water, which in this case is 64h. the hour of low water being marked at A, and that of high water at C, the intermediate hours being marked in succession as in the figure;

This division of the semi-circle may be made by means of a line of chords. The number of degrees corresponding to one hour being found by saying, as the whole elapsed time from low to high water (sold hours) is to 180° so is one hour to the arch corresponding to 1 hour 27° 42°, which being taken figure a line of churds and hid off from 5h. will reach to 6h. &c.

then any hour being found on the arch, the number of the line drawn perpendicular to AC, and passing through the hour, will represent nearly the number of feet to be subtracted from a sounding taken at that time to reduce that sounding to low water. Thus the number of feet corresponding to 8h. is between 4 and 5, because the mark 8h. falls between the lines marked 4 and 5, so that the reduction is between 4 and 5 feet, on soundings taken at 8 A. M. to reduce them to low water on the day of observation; and if on that day the tide does not ebb so much as on a spring tide, the reduction must be increased by the difference in the ebbing of the two tides. Thus if on the day of observation the tide did not ebb so much by two feet as on a spring tide, the reduction corresponding to 8h. ought to be increased two feet, and would therefore be between 6 and 7 feet. Allowance may be made for this by increasing the number of feet marked in fig. 5, by marking 2 feet at A, 3 feet at 1, 4 feet at 2, &c. as is evident.

To reduce a Draught to a smaller Scale.

With a black-lead pencil draw on the draught to be reduced, cross lines, forming exact squares, and on the clean paper for the copy draw the same number of squares, making their sides larger or smaller in proportion to the intended size of the scale, such as \(\frac{1}{2}, \frac{1}{2}, \frac{1}{2}c. \) the length of the other; distinguish by a stronger mark every fifth or sixth row of squares in both, so that the several corresponding squares may be readily perceived; then, in each of the squares of the draught, draw, by the eye, a curve on the paper, similar to that in the square of the copying draught, till the whole is copied; when the black-lead lines may be rubbed out with bread or india ruber.

A chart may also be reduced in the following manner: thus, suppose you would reduce a chart in the ratio of the line MN (Plate XI. fig. 6) to HI. Draw the line AC, which make equal to HI, upon A as a centre, describe the arch CF, and make the chord CF=MN, join AF; then if you take any distance, AB you wish to reduce, and upon A, as a centre, describe an arch BD; the chord BD, intercepted by the lines AC, AF, will be the reduced distance corresponding to AB. This reduced distance may also be obtained by another method, which is more simple than the former: Take any extent from the large chart, which is to be reduced to a smaller scale, and apply it from A to O (Plate XI. fig. 7): take in your compasses the corresponding distance on the small chart, and with one foot in O sweep an arch P; draw the line AP just touching the arch in P; then if you take any distance from the great chart, and apply it from A to R, and at the point R sweep an arch S to touch the line AP; the extent RS will be the reduced distance corresponding to the line AR.

OF WINDS.

THE earth is surrounded by a fine invisible fluid, called Air, which by its weight is capable of supporting the vapours raised by the sun, and by its elasticity is capable of expanding or spreading itself, so as to fill up a larger space. When the elasticity of any portion of the air is changed, by the heat of the sun or by other causes, the neighbouring parts are put in motion to restore the equilibrium; in this manner a current of air is formed, called the Wind, which is distinguished by several names, viz. trade winds, monsoons, variable winds, &c. The trade winds blow constantly from the same part; the monsoons blow half the year one way, and half the other; and the variable winds are such as blow without any regularity either as to time, place, or direction. The following observations on the wind have been made by Dr. Halley and others.

There are constant trade winds, blowing from the east, in most parts of the Atlantic and Pacific Oceans, between the latitudes of 30° N. and 30° S.

Near the northern limits of these winds, they blow between the north and east; and near their southern limits, between the south and east.

In the Atlantic Ocean, at about 100 leagues from the coast of Africa, between the latitudes of 28° and 10° north, there is generally a fresh gale of wind blowing from the N. E.

Those bound to the Caribbee Islands across the Atlantic, find, as they approach the American side, that the N. E. wind becomes easterly, or seldom blows more than a point from the east, either to the northward or southward.

These trade winds on the American side are sometimes extended to 50°, \$1°, or even to 52° of north latitude, which is about 4° farther than what they extend to on the African side; also to the southward of the equator, the trade winds extend 3 or 4 degrees farther towards the south on the coast of Brazil on the American side, than they do towards the Cape of Good Hope, on the African side.

But we must not conclude that the above limits are without exception; for both their extent and direction vary considerably with the season of the year. When the sun approaches the tropic of cancer the S. E. trade winds prevail farther to the northward of the line, and incline more to the southward of S. E. and the N. E. trade wind inclines more to the eastward; and

the contrary at the opposite season of the year.

On the African coast, from Cape Blance to Sierra Leone, the winds in general blow from the north, inclining from the westward rather than from the eastward. From Sierra Leone to Cape Palmas, the ordinary course of the winds is from W. N. W. and beyond Cape Palmas, as far as 28° south latitude, from S. W. to S. inclining more to the southward or westward according to the particular situation or bearing of the shores and lands. And the part of the ocean extending along this coast to the distance of 80 or 100 leagues from the shore, is much more troubled with frequent calms, and with sudden and violent gusts of wind, known by the name of Tornadoes, which blow from all parts of the horizon. The reason of this change in the direction of the trade wind near the land is probably owing to the nature of the coast, which being violently heated by the sun, rarefles the air exceedingly, consequently the cool air from the sea will keep rushing in to restore the equilibrium.

In the Gulf of Guinea there is a periodical wind, called *Harmattan*, which blows in a N. E. direction from the interior parts of Africa. The season in which this wind prevails is during the months of December, January, and

February.

Between the 4th and 10th degrees of north latitude, and between the longitude of Cape Verd and the easternmost of the Cape Verd Islands, there is a tract of sea, which seems to be condemned to perpetual calms, attended with terrible thunder and lightning, and frequent rains. The cause of this seems to be, that the westerly winds, setting in on the coast of Africa, and meeting the general easterly winds in this tract, balance each other, and so cause the calms; and the vapours, carried thither by each wind, meeting

and condensing, occasion the almost constant rains.

These observations show the reason of the difficulty which ships find in sailing to the southward, between the coasts of Guinea and Brazil, particularly in the menths of July and August, notwithstanding the width of the sea is more than 500 leagues. For the S. E. winds at that time of the year commonly extend some degrees beyond the ordinary limits of 4° north latitude, besides coming so much southerly as to be sometimes south, sometimes a point or two to the west; it then only remains to ply to windward: and if on the one side they steer W. S. W. they get a wind more and more easterly, but then there is danger of falling in with the coast or shoals of Brazil: and if they steer E. S. E. they fall into the neighbourhood of the coast of Guinea, from whence they cannot depart without running easterly as far as the island of St. Thomas.

All ships departing from Guinea for Europe, their direct course is northward; but on this course they cannot go, because the coast, tending nearly

east and west, the land is to the northward; therefore as the winds on this coast are generally between the south and W. S. W. they are obliged to steer S. S. E. or south, and with these courses they run off the shore; but in so doing they always find the wind more and more contrary; so that, though when near the shore they can lie south; at a great distance they can make no better than S. E. and afterwards E. S. E. with which courses they generally fetch the island of St. Thomas or Cape Lopez, where finding the winds to the eastward of the south, they sail westerly with it, till coming to the latitude of 4 degrees south, they find the S. E. wind blowing perpetually.

On account of these general winds, all bound from Europe to the West Indies, or to the southern states of America, consider it most advantageous to get as soon as they can to the southward, so they may be certain of a fair and fresh gale, to run before it to the westward. For the same reason, those bound from America to Europe endeavour to gain the latitude of 30 degrees, where they first find the wind begin to be variable, though the most ordinary winds in the North Atlantic Ocean come between the south and

And for the same reasons those bound to India from America run to the eastward in the variable winds, so as to be in the longitude of 350 or 380 W. when in the latitude of 30° N. From thence they steer south-easterly towards the Cape de Verds, passing 40 or 50 to the westward of them, unless they wish to stop for supplies, or to correct their longitude. Being then in the common route of the European Indiamen, they steer south-easterly to cross the equator between the longitude of 18° W. and 25° W. where meeting the S. E. trade winds, they must brace up and sail upon a wind till they get through them, and come into the variable winds, where they may steer to the eastward. • Near the equator, the trade wind is generally stronger to the westward than to the eastward; and were it not for the fear of falling in with the Brazil coast, a ship might cross the line farther to the westward than what we have recommended above. Ships homeward bound, from the Cape of Good Hope towards America, may deviate a little to the westward of their straight course, and cross the equator in about 50° W. longitude, in order to take advantage of this fresher trade wind.

Between the southern latitudes of 10° and 30° in the Indian Ocean, the general trade wind about S. E. is found to blow all the year round, in the same manner as in the like latitudes in the south Atlantic Ocean; and during the six months, from May to November, these winds reach to within 2 degrees of the equator; but during the other six months, from November to May, a N. W. wind, called the little monsoon, blows in the tract lying between the 3d and 10th degrees of south latitude, in the meridian of the north end of Madagascar, and between the 2d and 12th degrees of south latitude, near the longitude of Sumatra and Java.

In the tract between Sumatra and the African coast, and from 80 of south latitude, quite northward to the Asiatic coast, including the Arabian Sea and the Bay of Bengal, the monsoons blow from October to April on the N. E. and from April to October on the S. W. In the former half year, the wind is more steady and gentle, and the weather clearer than in the latter six months. In the Red Sea the winds blow nearly nine months of the year from the southward, that is, from August to May, and the rest of the year from the N. and N. N. W. with land and sea breezes. In the Gulf of Persia the N. W. wind blows from October to July, and about three months from the opposite quarter. These winds being often interrupted by gales from the S. W. and by land breezes.

Between the island of Madagascar and the coast of Africa, and thence northward as far as the equator, there is a tract, wherein, from April to October, there is generally a S. S W. wind, and a contrary wind the rest of the

year, with regular land and sea breezes on both coasts.

To the eastward of Sumatra and Malacca, on the north of the equator,

and along the coasts of Cambodia and China, quite through the Phillippines as far as Japan, the monsoons blow N. E. and S. W. the N. E. setting in about October or November, and the S. W. about May.

Between Sumatra and Java to the west, and New-Guinea to the east, there are regular monsoons. The N. W. monsoon blows from October to April.

the S. E. monsoon the rest of the year.

The monsoons do not shift suddenly from one point of the compass to the opposite; in some places, the time of the change is attended with calms, in others by variable winds; and it often happens, on the shore of Coromandel and China, towards the end of the monsoons, that there are most violent storms, called Tuffons, greatly resembling the huricanes in the West Indies, wherein the wind is so vastly strong, that hardly any thing can resist its force; for this reason it is more dangerous to approach those shores at the time of the breaking up of the monsoon than at any other season of the vear.

The land and sea breezes prevail principally between the tropics. The sea breeze generally sets in about ten in the forenoon, and continues till about five or six in the evening: at seven the land breeze begins, and continues till about eight in the morning. The cause of these winds is this:—during the day the sea is not so much heated by the sun as the land, nor so much cooled at night: Hence, in the day time, the cooler air from the sea will rush towards the land to supply the deficiency occasioned by the greater rarefaction of the air, and hence arises the sea breeze. In like manner, during the night, the air at land, being more cooled than that at sea, will therefore blow from the land towards the sea, and hence occasion a land breeze.

A whirlwind is a dangerous phenomenon caused by the adjacent air, rushing in from all parts towards a centre with great rapidity, and destroying every thing it passes over in its progressive motion. A water spout and whirlwind arises from the same cause, the latter being formed at land, is composed principally of air, but the former being formed at sea, is composed of

water.

It was first observed by Doctor Franklin that the N. E. storms on the coast of the United States of America, frequently begin earlier in the southern states than in the northern. This he accounts for by supposing a great rarefaction of air in or near the gulph of Mexico; the air rising thence has tis place supplied by the next more northern, and therefore denser and heavier air; a successive current is thus formed, to which the coast and inland mountains give a N. E. direction.

Experiments have been made by several persons to determine the velocity of the wind, by observing the space passed over by a cloud or any light substance, and by other methods; and it has been found that the velocity of the

wind in a violent gale is about 50 or 60 miles per hour.

TIDES.

TIDE is a periodical motion of the water of the sea, by which it ebbs and flows twice a day. The flow continues about 6 hours, during which the water gradually rises till it arrives to its greatest height; then it begins to ebb or decrease, and continues to do so for about 6 more, till it has fallen to nearly its former level; then the flow begins as before. When the water has attained its greatest height it is said to be high-water, and when it is done falling it is called low-water.

The cause of the tides is the unequal attraction of the sun and moon upon different parts of the earth. For they attract the parts of the earth's surface nearest to them, with a greater force than they do its centre: and attract the centre more than they do the opposite surface. To restore this equilibrium the waters take a spheroidal figure, whose longer axes is directed

towards the attracting luminary. If the moon only acted upon the water. the time of high water would be when the moon was upon the meridian. above or below the horizon; or rather at an hour or two after, (because the moon continues to act with considerable force for some time after passing the meridian.) But the moon passes the meridian about 49' later every day; of course, if she only acted on the tides, they would be retarded every day 49', and it would be high water at the same distance from her passing the meridian; and it is upon this principle that the time of high water is calculated in most books of navigation, although the time thus calculated will sometimes differ an hour from the truth, owing to the neglect of the disturbing force of the sun. The effect of the moon upon the tides is greater than that of the sun, notwithstanding the quantity of matter in the latter is vastly greater than in the former: but the sun, being at a much greater distance from the earth than the moon, attracts the different parts of the earth with nearly the same force: whereas the moon, being at a much less distance, attracts the different parts of the earth with very different forces. According to the latest observations, the mean force of the sun for raising the tides is to the mean force of the moon as 1 to 24. By the combined effect of these two forces, the tides come on sooner when the moon is in her first and third quarters, and later in the second and fourth quarters, than they would do if caused only by the moon's attraction. The mean quantity of this acceleration and retardation is given in Table B, subjoined; the use of which will be explained hereafter.

The tides are greater than common about three days after the new and full moon; these are called spring-tides. And the tides are lower than common about three days after the first and last quarters; these are called the scop-tides. In the former case the sun and moon conspire to raise the tide in the same place, but in the latter the sun raises the water where the moon depresses it. When the moon is in her perigee, or nearest approach to the earth, the tides rise higher than they do, under the same circumstances, at other times; and are lowest when she is in her apogee, or farthest distance from the earth. The spring-tides are greatest about the time of the equinoxes, in March and September, and the neap-tides are less. All these things would obtain exactly, were the whole surface of the earth covered with sea; but the interruptions caused by the continents, islands, shoals, &c. entirely alter the state of the tides in many cases. A small inland sea, such as the Mediterranean or Baltic, is little subject to tides; because the action of the sun and moon is always nearly equal at the extremities of such seas.

In very high latitudes the tides are inconsiderable.

From the observations of many persons, the times of high-water on the days of new and full moon, in the most noted places of the globe, have been collected. These times are usually put in a table against the names of the places, arranged in alphabetical order as in Table XLVII. of the collection accompanying this work, by means of which the times of high-water may be found by various methods. The most common rule prescribed for this purpose, in books of navigation, is that depending on the golden number and epact, the tide being supposed to be uniformly retarded every day. This method will sometimes differ 2 hours from the truth, for which reason I shall not insert it: but shall proceed to explain the calculation by the adjoined tables A and B, and the Nautical Almanac; by means of which the time of high-water may be obtained to a greater degree of exactness than from our common Almanacs.

RULE.

Find the time of the moon's coming to the meridian at Greenwich on the given day, in page VI. of the Nautical Almanac. Enter Table A, and find the longitude of the given place, in the left hand column, corresponding

TIDES. 215

to which is a number of minutes to be applied to the time of passing the meridian at Greenwich, by adding when in west longitude, but subtracting when in east longitude; the sum or difference will be nearly the time that the moon passes the meridian of the given place. With this time enter Table B, and take out the corresponding correction, which is to be applied to the time of passing the meridian of the place of observation, by adding or subtracting, according to the direction of the table.

To this corrected time add the time of full sea on the full and change days; the sum will be the time of high-water at the given place, reckoning from the noon of the given day. If this sum be greater than 12h. 24m. you must subtract 12h. 24m. from it, and the remainder will be the time of high-water nearly, reckoning from the same noon; or if it exceed 24h. 48m. you must subtract 24h. 48m. from that sum, and the remainder will be the time of high water, reckoning from the same noon nearly.

EXAMPLE I.

Required the time of high water at Charleston, (S. C.) March 17, 1820, in the afternoon, civil account?

By the Nautical Almanac I find that the moon passes the meridian of Greenwich at 2h. 31m.; to this I add 11m. taken from Table A, corresponding to the longitude of Charleston. With the sum 2h. 42m. I enter Table B, and find (by taking proportional parts) that the correction is 45m. which is to be subtracted from 2h. 42m. (because immediately over it in the table it is marked Sub.); to the remainder 1h. 57m. I add the time of high water on the full and change days 7h. 15m. (which is found in the tide table following;) the sum 9h. 12m. is the time of high water on the afternoon of March 17, 1820, civil account.

EXAMPLE II.

Required the time of high water at Portland, (Maine) May 28, 1820, in the afternoon, civil account?

By the Nautical Almanac the moon will pass the meridian of Greenwich at 8 hours 49 minutes. The correction from Table A, corresponding to 70° the longitude of Portland is 9m. which added to 8h. 49m. gives the time of the moon's southing at Portland 8h. 58m. nearly. The number in Table B corresponding to 8h. 58m. is 28m. which is to be added to 8h. 58m. (because immediately over it, in the table, is marked Add.) To the sum 9h. 21m. I add the time of high water, on the full and change days, 10h. 45m. and the sum is 20h. 6m. consequently the high water is at 20h. 6m. past noon of May 25, that is, at 8h. 6m. A. M. of May 24. And by subtracting 12h. 24m. from 20h. 6m. we have 7h. 42m. which will be nearly the time of high water on the afternoon of May 25, 1820.

In this manner we may obtain the time of high water at any place, to a considerable degree of accuracy. But the tides are so much influenced by the winds, freshets, &c. that the calculated times will sometimes differ a little from the truth.

Many pilots reckon the time of high water by the point of the compass the moon is upon at that time, allowing 45 minutes for each point. Thus on the full and change days, if it is high water at noon, they say a north and south moon makes full sea; and if at 11h. 15m. they say a S. by E. or N. by W. moon makes full sea; and in like manner for any other time. But it is a very inaccurate way of finding the time of full sea by the bearing of the moon; except in places where it is high water about noon on the full and change days.

When you have not a Nautical Almanac, you may find the time of high water by means of the following tables C and D; and although the former method is the most accurate, yet the latter may be useful in many cases. To calculate the time of full sea by this method, observe the follow-

ing rule.

RULE.

Enter Table C, and take out the number which stands opposite to the year, and under the month for which the tide is to be calculated; this number, added to the day of the month, will give the moon's age, rejecting 50 when the sum exceeds that number. Against her age found in the left hand column of Table D, is a number of hours and minutes in the adjoined column, which being added to the time of high water at the given place on the full and change days, will give the time of high water required, observing to reject 12h. 24m. or 24h. 48m. when the sum exceeds either of those times.

By this rule I shall work the two preceding examples.

EXAMPLE III.

Required the time of high water at Charleston, (S. C.) March 17, 1820, in the afternoon, civil account?

In the table C, opposite 1820, and under March, stand 16, which, added to the day of the month 17, gives 33, and by subtracting 30, leaves 3, the moon's age: opposite 3 in Table D, is 1h. 46m. which added to 7h. 15m. the time of high water on the full and change days, gives 9h. 1m. for the time of high water; differing eleven minutes from the former method.

EXAMPLE IV.

Required the time of high water at Portland, (Mass.) May 23, 1820, in the

afternoon, civil account?

In the Table C, opposite 1820, and under May, stand 18; which added to the day of the month 23, gives (by neglecting 30) the moon's age 11; opposite to this, in Table D, is 9h. 19m. which added to 10h. 45m. the time of high water on the full and change days, gives 20h. 4m. from which subtracting 12h. 24m. there remains 7h. 40m. for the time of full sea May 23, 1820; this differs 2 minutes from the former method.

In the third column of Table D is given the time of the moon's coming to the meridian, for every day of her age: thus, opposite 11 days stand 8h. 57m. which is the time of her coming to the meridian on that day. This table may be of some use when a Nautical Almanac cannot be procured; but being calculated upon the supposition that the moon moves uniformly in the equator, the table cannot be very accurate. The numbers in this Table are reckoned from noon to noon; thus, 1h. A. M. is denoted by 15h.; 2h. A. M. by 14h. &c.

The time of new moon is easily found, by subtracting the number taken from Table C from So. Ex. Suppose it was required to find the time of new moon for May, 1820? By examining the table, we find the number corresponding to that time is 18; this subtracted from 30 leaves 12; therefore it will be new moon the 12th May, 1820.

When the time of high water is known for any day of the moon's age, we may from thence find the time of high water on the full and change days,

by the following

RULE.

Find the time of the moon's coming to the meridian of Greenwich, in page VI. of the Nautical Almanac; to this time apply the corrections taken from the tables A and B, (in the same manner as directed in the preceding rule for finding the time of high water) subtract this corrected time from the observed time of high water, and the remainder will be the time of high water, on the change and full days.

NOTE. If the time to be subtracted be greater than the observed time of full sea, you must increase the latter by 12h. 24m. or by 24h. 48m. nearly.

EXAMPLE.

Suppose that on the 17th March, 1820, the time of high water at Charles-

ton, (S. C.) was found to be at 9h. 12m. P. M. required the time of high ter on the full and change days?

I find, as in example 1st. preceding, that the number to be subtracted. 1h. 57m.—taking this from 9h. 12m. leaves 7h. 15m. which is the time

high water on the full and change days.

When you have not a Nautical Almanac, you may find the time of I water on the full and change by means of the Tables C and D. For in present example, I find by Table C, that the moon's age was S, corresponds to which, in the second column of Table D, is 1h. 46m. this subtraction 9h. 7m. leaves 7h. 21m. for the time of high water on the full change days.

7	AB.	A.	Tai	в. В.		• •		T	B.	Ċ							7	Pa:	B.]	D.
	Langlinde of the place.	Cor. of Mosa's passing the me-	Time of Moon's passing the me- ridien.	Corr.	Add	A TABLE FOR FINDING THE MOON'S AGE. Add the number taken from this Ta-								Moon's Age.	H.	w	H Moon passes			
	Deg.	-	Hours		su	ble to the day of the month; the sum (rejecting 30 or 60 if necessary) will be the Moon's age nearly.							0	0	 0 35	0				
1	0	9	_	Sub.			1:	,	7	-		7	1	-	·		2	_	10	1
	10 2 0	1	. 1 2	0 34	Year	Jan.	March	April	May.	June.	July	Aug.	Sept.	Oct.	Nov.	Dec.	3 4 5	_	46 22 1	2 5 4
	3 0	1 .1	4		1820	1517	.	 		-		-		-	-	-	6 7	-	44 35	4 5
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	60	8		Add o 2	1822	.7	7	9	10	11	12	1 1	15	16	17	17		_	19	8
	70	9	9	0 25	1823	1820	18	19	20	22	છ્ટ	24	26	26	28	28	13	10	10 54	
	80	11	. 10 11		1824	29	10	1	2	4	4	6	8	8	10	10	14 15	11 12	33 9	
	90	12	12			-	-	-		_		_	-	_	_	_	-			_
	100	14	15	Sub.	1825 1826	 	-	<u> </u>	-	 —	-		-	_	_	_	17	13	44 19	13
	110	15		0 34	1820	22 8	25	25	24	20	23	27	29	Z 8	1	32			54 81	
-	120	16	15 16		18 2 7	3 :	3	05	5	6	7	8	10	10	12	12		15	11 56	16
ļ	150	18	17	1 9	1828	141:	15	16	17	18	19	19	21	22	23	24	Q Q	16	49	17
	140	19	18 19	0 35	1829	25 26	25	27	28	29	ó	1	e	3	4	5	23 24	19	57 17	19
1	150	20	20	Add	1850		6	-	-	10	-	10	10	14	15	10			3₹ 3S	
	160	25		0 23			_ °	Ľ	_		Ľ	-	13	<u> </u>	Ľ	10			92 92	
1	170	25			1831	17 18	17	18	19	20	21	23	24	25	26				4	
1	180	94	7.23	0 14			1	_	-	_	Τ.	_	_				29	ะฮ	42	23

In all the preceding calculations of the time of high water, we have gleeted the correction arising from the variation of the distances of sun and moon from the earth, and from the different declinations of the editor. These causes might produce a correction of 10' or 12' in the of high water, but in general will be much less, and may therefore be gleeted.

210

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0 1832 28 29 28 0 1 5 3 5 7 7 8 9 29 24 0 24

CURRENTS.

A CURRENT is a progressive motion of the water, causing all floating bodies to move that way towards which the stream is directed. The set of a current, is that point of the compass towards which the waters run, and its drift is the rate it runs per hour. The most usual way of discovering the set and drift of an unknown current, is thus:

Let three or four men take a boat a little way from the ship: and by a rope fastened to the boat's stern, let down a heavy iron pot or loaded kettle to the depth of 80 or 100 fathoms; then heave the log, and the number of knots run out in half a minute will be the miles the current sets per hour.

and the bearing of the log will show the set of it.

There is a very remarkable current, called the GULF STREAM, which sets in a north-east direction along the coast of America, from Cape Florida towards the Isle of Sables, at unequal distances from the land, being about 75 miles from the shore of the southern states, but more distant from the shore of the northern states; the width of the stream is about 40 or 50 miles, widening towards the north; the velocity is various from one to three knots per hour, or more, being greatest in the channel between Florida and the Bahamas, and gradually decreasing in passing to the northward; but is greatly influenced by the winds both in drift and set.

We are chiefly indebted to Doctor Franklin, Commodore Truxton, and Mr. Jonathan Williams, for the knowledge we possess of the direction and velocity of this stream; its general course, as given by them, is marked on the chart affixed to this work. They all concur in recommending the use of the thermometer, as the best means of discovering when in, or near the stream. For, it appears by their observations, that the water is warmer than the air when in the stream; and that at leaving it, and approaching towards the land, the water will be found six or eight degrees colder than in the stream, and six or eight degrees colder still, when on soundings. Vessels coming from Europe to America, by the northern passage, should keep a little to the northward of the stream, where they may probably be assisted by a counter current, as is observed by Commodore Truxton. When bound from America to Europe, a ship may generally shorten her passage by keeping in the gulf. By steering N. W. you will generally cross the gulf in the shortest time, as the direction of the stream is nearly N.E. Those who wish for further information on this subject, may consult an ingenious treatise on "Thermometrical Navigation," published by Mr. Jonathan Williams, at Philadelphia, in 1799, and re-published by Edm. M. Blunt, to accompany his Chart of the Western Ocean, in 1819.

In the other parts of the Atlantic ocean the currents are variable, but are generally south-easterly, along the coast of Spain, Portugal and Africa, from the Bay of Biscay towards Madeira and the Cape de Verds. Between the

tropics there is generally a current setting to the westward.

There is also a remarkable current which sets through the Mozambique channel, between the Island of Madagascar and the main continent of Africa, in a south-westerly direction: in proceeding towards Cape Lagullas the current takes a more westerly course, and then tends round the Cape towards Ships bound to the westward from India, may generally shorten St. Helena. their passage, by taking advantage of this current. On the contrary, when bound to the eastward, round the Cape of Good Hope, they ought to keep far to the southward of it. However, there appears to be a great difference in the velocity of this current at different times; for some ships have been off this Cape several days endeavouring to get to the westward, and have found no current; others have experienced it setting constantly to the westward during their passage from the Cape towards St. Helena, Ascension and the West-India Islands. Digitized by Google

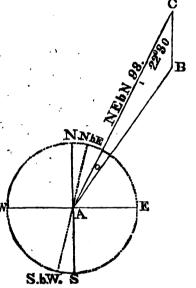
All cases of sailing in a current are calculated upon the principle, that the ship is affected by it in the same manner as if she had sailed in still water, with an additional course and distance exactly equal its set and drift: on this principle the projection and calculation of any problem of this kind may be easily made.

EXAMPLE.

If a ship sails 98 miles N. E. by N. in a current which sets S. by W. 27 miles in the same time; required her true course and distance?

BY PROJECTION.

Describe the compass NESW, through the centre A draw the N. E. by N. line AC=98 miles, through C draw the line CB parallel to the S. by W. line, and make CB=27 miles, and join-AB. Then AB will be the course and distance made good, which by measuring are N. E. I. N. 74 miles.



BY CALCULATION.

The shortest method of calculating this problem is by means of Table I. as in the adjoined Traverse Table; putting in it the course sailed by the ship, and the set of the current, and finding the difference of latitude and departure by that Table, then find the course and distance made good, as in Case VI. Plane Sailing. In the present example the course is N. E. 2 N. and the distance 74 miles nearly.

TRAVERSE TABLE.											
Courses.	Dist.	N.	8.	E.	w.						
N.E.byN. S. by W.		81.5	26.5	54.4	5.3						
		81.6 26.5	26.5	54.4 5.3	5.3						
Diff.	Lat.	55.0	Dep.	49.1							

METHOD OF KEEPING A SHIP'S RECKONING OR JOURNAL AT SEA.

A SHIP'S RECKONING is that account, by which it can be known at any time where the ship is, and on what course or courses she must steer to gain her port. Dead Reckoning is that account deduced from the ship's run from the last observation.

THE LOG-BOARD. .

The daily occurrences		_					
on board a ship are mark- ed on a board or slate,	н.	K.	F.	Courses,	Winds.	Leo- vay	Transactions.
called the log-board or log-slate, kept in the steer-	2	6		s. w.	N. E.		
age for that purpose, being usually divided into seven columns; the first contains the hours from noon to noon, being marked by some for every two hours, but by others for	4 6 8 10 12 4 6	5 5 4 4 4 4 4	5 5 5 5 5	E. N. E.	N.W.by W.		Moderate gales & fair weather; at 8 A. M. saw a ship to the north- ward.
every single hour; in the second and third columns	10	5 4	5	s. w.	W. N. W.	1	No observation.
are the knots and fathoms	12	1 4			<u> </u>	1	

the ship is found to run per hour, set against the hours when the log was hove. Some navigators do not divide the knot into ten fathoms, but into half knots only, marking the third column H. K. The fourth column contains the courses steered by compass; the fifth, the winds; the sixth, the lee-way,* and the seventh, the alteration of the sails, the business done aboard, and what other remarks the officer of the watch thinks proper to in-For it should be observed, that it is usual to divide a ship's company into two parts, called the starboard and larboard watches, who do the duty of the ship for four hours and four hours, alternately, except from 4 to 8 P. M. which is divided into two watches. The remarks made on the log-board are daily copied into a book called the Log-Book, which is ruled like the log-This book contains an authentic record of the ship's transactions, and the persons who keep a reckoning, transcribe them into their journals, and from thence make the necessary deductions relative to the ship's place, every day at noon, which operation is called working a day's work. While a ship is in port, the remarks entered in the log-book are called harbour work, or harbour journals, and the day is then estimated according to the civil computation as on shore, that is from mid-night to mid-night; but at sea the day's work ending at noon is dated the same as the civil day, so that the day's work marked Monday began on Sunday noon, and ended on Monday at noon; the day thus marked is called a nautical day; the first 12 hours being marked P. M. the latter A. M. There are various ways of keeping journals at sea, according to the different tastes of navigators. Some keep only an abstract of each day's transactions, specifying the weather, what ships or lands were seen, accidents on board, the latitude, longitude, course, and run: these particulars being drawn from the ship's log-book. Others keep a full copy of the log-book, and the deductions drawn therefrom, arranged in proper columns:—this is the most satisfactory method to these who may have occasion to inspect the journal; and we have adopted it in the following, but shall give an abstract at the end conformable to the other method.

When a ship is about losing sight of the land, the bearing of some noted place (whose latitude and longitude are known) must be observed, and its distance estimated and marked on the log-book: this is called taking a departure. In working this first day's work, the calculation is to be made in the same manner as if the ship had sailed that distance from that place upon a course opposite to that bearing, and that course and distance are to be entered accordingly into the traverse table, after allowing for the variation.

We have already taught the methods of finding the variation, which must be allowed on all courses steered, and on all bearings taken with the compass; to the right hand, if the variation be east; but to the left hand, if west; the observer being supposed to be placed in the centre of the compass, looking towards the point from which the variation is to be allowed.

The same of the leeway and manner of allowing for it, are explained in the following pages.

EXAMPLES.

		Points.	
Company	N. E. by E.	Variation 2 W.	True course N. E. by N.
	N. E.	11 E.	N. E. by E. 1 E.
	N. W.	3 W.	W. by N.
· · · · · · · · · · · · · · · · · · ·	8. E.	3 E.	S. by E.
	8. S. W.	11 W.	8. į W.
	E. S. E.	1} W.	E. 1 S.
•	8. W. į W.	į ₩.	8. W. i S.
	N. N. E. 3 E.	1] E.	N. E. J E.

To find the lee-way and allow for it.

The courses must likewise be corrected for lee-way, the nature of which may be thus explained. When a ship sails upon a wind, in a fresh gale, that part of the wind which acts upon the hull and rigging, together with a considerable part of the force exerted on the sails, tend to drive her immedistely from the direction of the wind, or, as it is termed, to leeward. But since the bow of a ship exposes less surface to the water than the side, the resistance will be less in the first case than in the second; the velocity therefore in the direction of her head will, in most cases, be greater than the velocity in the direction of her side, and the ship's course will be between the two directions, and the angle contained between the course towards which the ship's head is directed, and the course she really describes through the water, is termed her lee-way. The quantity of lee-way to be allowed will depend upon a variety of circumstances; as the mould and trim of the ship; the quantity of sail she carries; her velocity through the water, &c. hence no general rules can be laid down with accuracy that will determine the quantity of lee-way in all cases. The following have, however, been usually given by most writers on navigation.

When a ship is close hauled with all her sails set, the water smooth, and a light breeze of wind, she is then supposed to make little or no lee-

ē. When the top-gallant sails are handed, allow 1 point.

8. When under close reefed topsails, allow 2 points.

When one topsail is handed, allow 24 points.

When both topsails are handed, allow 32 points. When the fore course is handed, allow 4 points.

When under the mainsail only, allow 5 points.

When under a balanced mizen, allow 6 points. 2

8. When under bare poles, allow 7 points.

As these allowances depend entirely on the quantity of sail set, without regard to any other circumstance, it is evident that they can be considered only as probable conjectures, and may indeed serve to work up the day's work of a journal that has been neglected. But since the computation of a ship's way depends much upon the accuracy of this allowance, it would be proper for the officer of the watch to mark the lee-way on the log-board, in the column reserved for that purpose. The lee-way may be estimated by observing the angle which the wake of the ship makes with the point right astern, by means of a semi-circle marked on the taffrail, and divided into points and quarters; by means of which the angle contained between the direction of the wake and the point of the compass directly astern, may be easily ascertained.

The lee-way thus determined is to be allowed on all courses steered, to the right hand of the course steered, when the larboard tacks are aboard, but on the left hand, when the starboard tacks are aboard; the person making the allowance being supposed to be looking towards the point of the compass the ship

is sailing upon.

•	EXA	MPLES.	
Courses steered. N. W. E. N. E. E. S. E. W. by N. E. N. E. 4 E.	Wind. N. N. E. North. South. N. by W. S. E.	Lee-way. 1 point. 2 1	True course. N. W. by W. East. E. by S. W. j N. N. E. 4 N.

If the ship has been

When the variation and lee-way are both to be allowed on a course, you may do it at once, by allowing their sum when they are both the same way, or their difference when the allowance is to be made in different ways, taking care to make the allowance in the same way as the greater quantity ought to be, whether it be the variation or lee-way.

a sorp steers W. by N. with her larboard tacks aboard, and makes one point lee-way, there belong two points westerly variation; required the true course?

A ship steers E. S. E. with her starboard tacks aboard, and makes two points lee-way, there belong two points westerly variation; required the true course?

Lee-way to the right hand

1 point

Lee-way to the left Difference allowed to the left Whence the course is west. 1 point Sum allowed to the left
Whence the course is E. by N.

In a violent gale, with a head wind and heavy sea, when it would be dangerous to carry sail, it is usual to lie to under sufficient sail to prevent the vessel from rolling so much as to endanger the masts and rigging. ship is lying-to, the tiller is put over to leeward, and when the ship has head-way, the rudder acts upon her to bring her to the wind; the ship then loses her way in the water, which ceasing to act on the rudder, her head falls off from the wind, and the sail which is set fills and gives her fresh way through the water, which acting on the rudder, brings her head again to the wind. Thus the ship is kept continually falling off and coming to. In this case, you must observe the points on which she comes up and falls off, and take the middle between the two points for the apparent course, from which allow the variation and lee-way, and you will obtain the true course.

EXAMPLE. '

A ship lying-to under her mainsail, with her starboard tacks aboard, comes up E. by S. and falls off N. E. by E. there being one point westerly variation, and she makes 5 points lee-way-what course does she make good?

The middle between E. by S. and N. E. by E. is E. by N.; and by allowing 6 points to the left hand (viz. 5 for lee-way and 1 for variation) the true course will be obtained N. by E.

To exercise the learner we shall add the examples of correcting for variation and lec-way contained in the following Table.

THE TABLE.

acted upon by a current or a heave of the sea,		. Winds.	way	Varia- tion points	Courses corrected.
you must allow the set and drift as a course and distance in the Traverse Table, as directed in p. 219.	W. S. W. W. S. W. W. by N. S. W.	N. N. E. N. N. W. S. S. S. W. N. by W. W. N. W.	1 11	W. W. W.	N. 51 W. 8. 61 W. 8. 62 W. W. 8. 7 W. 8. 14 W.
Having corrected the courses for lee-way and variation, and estimated the distances sailed, the latitude and longitude in at noon are to be found by either of the preceding methods of sailing. The latitude	S. S. W. S. W. W. by N. S. E. by S. E. N. E.	W. S. W. W. W. by W. S. S. W. N. by W. E. S. E. S. \ \ L. N. N. S.	1 1 2 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	**************************************	S.S. E. 8. t. E. 8. S. W. † W. W. † N. W. S. W. † W. S. † W. E. N. E. † E. E. N. E. † E.
and longitude, thus cal- culated, are called the latitude and longitude by dead reckoning, and	E. S. E. W. S. W. W. by N. N. W.	E. S. E. N. E. S. S. W. by S. W. S. W.	1	14 W. 13 W. 13 W. 14 W. 14 W.	8. by E. ‡ E. E. ‡ S. S. W. by W. W. ‡ N. N. W. ‡ W.
if the real course and distance made good by the ship could be esti- mated accurately by the compass and log, no-	N. by E.	W. S. W. N. W. by W. W. by S. N. by E. N. W. by N.	1 11 14 14	E E E E	S. ‡ E. N. N. E. ‡ E. N. ‡ W. N. W. by W. ‡ W. W. ‡ S.

thing more would be necessary to determine the ship's place at any time; but by

reason of the various accidents that attend a ship's way, such as heave of the sea, unknown currents, different rates of sailing between the times of heaving the log, sudden squalls, improper allowance for lee-way and variation, the latitude and longitude of the ship as deduced from the reckoning, will frequently differ from the latitude and longitude by observation. In this case it will be proper to re-examine the calculation to see whether a iust allowance has been made for lee-way, variation, bad steerage, drift of the sea, error of the log-line and glass, &c. since it will sometimes be found that a different and more probable estimate of some of these quantities will make the dead-reckoning agree more nearly with the observations. Before the method of finding the longitude by lunar observations was introduced, the mariner had no other observation to be depended on except his latitude. and it was then usual to make allowances for supposed errors in the courses and distances, so as to make the latitude by observation and dead-reckoning agree. The method made use of by Robertson, Moore, and others, was divided into three cases, viz.

CASE I.

When the course was within three points of the meridian, the error was supposed to be wholly in the distance, on the principle that it would require a greater error in the course to cause the given error in the difference of latitude than could be supposed probable to have been committed. In this case the corrected departure, &c. were found, with the course, by deadreekoning, and the difference of latitude by observation, as in Case IV. of Middle Latitude or Mercator's Sailing.

CASE II.

When the course was between three and five points of the meridian, the error was supposed to be part on the course and part on the distance. In this case, the corrected departure was taken equal to the mean of the departure by dead-reckoning, and the departure which corresponds to the distance by dead-reckoning, and the difference of latitude by observation. With the corrected departure, and the difference of latitude by observation, the course, &c. were found as in Case II. of Middle Latitude or Mercator's Sailing.

CASE III.

When the course was more than five points from the meridian, the error was supposed to be wholly on the course, on the principle that it would require a greater variation in the distance to make the dead-reckoning and observation agree, than could be supposed probable, whereas it could require but a small change in the course to produce the sought effect. In this case, the corrected departure, &c. were found, with the distance, by dead reckoning, and the difference of latitude by observation, by Case V. of Middle Latitude or Mercator's Sailing.

This method was given in the former editions of this work, in conformity to custom, though I was decidedly opposed to making such corrections, being convinced that the difference between the dead-reckoning and observation is more owing to unknown currents than to errors in the courses and distances given by the log. Even admitting the principle that an arbitrary correction of this kind is proper, the preceding method does by no means appear to be the most probable. To show this, let us take the following

Suppose the course by dead-reckoning to be 35° 44′ 59″, the distance by the log. 100 miles, and the difference of latitude by observation 73.1 miles. This comes under Case I. and the error must be placed wholly on the distance, which is to be found with the course 33° 44′ 59″ (or 3 points hearly) and the difference of latitude 73.1, so that the corrected quantities are nearly by Table I. course 33° 44′ 59″, distance 88, and departure 48.8. Now by altering the course two seconds, making it 33° 45′ 1″, still retaining the distance by dead reckoning 100 miles, and the difference of latitude by observation 73.1, the example will come under Case II. and the corrected depar-

ture is the mean of the departure by dead-reckoning 55.6, and that corresponding to the distance 100, and the difference of latitude by observation 73.1; namely 68.2, so that this corrected departure is 61.9 miles, with which and the difference of latitude by observation 73.1, we obtain the corrected course 40° 15' and the distance 96. Thus we see that by altering the course by dead reckoning only two seconds, the corrected course varies from 830 44' 59" to 400 15' or above 6 degrees, and the departure varies from 48.8 to 81.9, both of which are highly improbable. This defect of the rule evidently arises from the sudden change of the method when the course is near 5 or 5 points; it being much more probable that the variations take place by small degrees, in such manner that when the course by dead-reckoning is exactly on the meridian, the error ought to be in the distance, and when the course is 8 points from the meridian, the error ought to be in the course, and at intermediate courses the errors in distance ought to be greater, the nearer the course is to the meridian, and the errors in the course greater, the far-ther it is from the meridian. Both these objects are attained in a very simple manner in the method proposed by the ingenious methematician, Dr. Adrian, late Professor of Mathematics and Natural Philosophy in Columbia College, New-York, which is somewhat similar to my method of correcting a survey. His method consists in finding, with the difference of latitude by observation, and the departure by account, the corrected course, distance, and difference of longitude by Case II. of Middle Latitude or Mercator's Sailing, so that no correction whatever is made in the departure. The propriety of this method will appear evident by observing that a change in the departure can have no tendency whatever in correcting an error in the latitude, and there can be no reason given why such change should be made to the eastward rather than to the westward, since it is supposed that all the allowances for heave of the sea, falling off the course, variation, error of the log, &c. have been previously taken into the calculation, and it seems to be contrary to sound reasoning to vary any of the elements when it will not serve to correct the known error of the latitude, particularly when there can be no reason given why the change should be made in one direction rather than another. In addition to this, the proposed method is not liable to the inconvenience of a sudden change in the rules when the course is near 3 or 5 points. It has also another advantage with respect to simplicity of calculation arising from the circumstance that the corrected difference of longitude is nearly the same as the difference of longitude by dead-reckoning. For the departure is not varied by the rule, and the middle latitude differs rarely more than a few minutes on account of the difference between the latitude by observation and account, so that in keeping a journal it will not be necessary to make any change in the longitude by dead-reckoning, even if you have not had an observation for several days. To illustrate this method I shall give the following

EXAMPLE.

Yesterday at noon we were in the latitude of 39° 18' N. and by an observation at noon this day are in the latitude of 57° 48' N. our dead-reckoning gives 107 miles southing and 64 miles westing. Required the course, distance, and difference of longitude?

With the difference of latitude by observation 90 miles (the difference of \$7° 48' and \$9° 18') and the departure by dead-reckoning 64 miles, I find by Case II. of Mid. Lat. Sailing, the course nearly \$5°, and the distance 110 miles; and with the middle latitude by observation \$8° 35', and the departure 64 miles, I find the difference of longitude to be 82 miles. If the middle latitude by dead-reckoning \$8° 41' had been taken, the result would have been nearly the same.

. If you have not had an observation several days, and then find an error in the latitude by account, you may on these principles correct the latitude

on the intermediate days, by saying, as the sum of all the distances sailed, since the first observation, is to the whole error in the latitude, so is the sum of the distances sailed from the time of taking the first observation, to the noon of any particular day, to the correction of the latitude by dead-reckoning on that day, southerly if the last latitude by observation is south of the latitude by dead-reckoning, otherwise northerly. Thus, if the latitudes by dead-reckoning at noon, on four successive days, were 41° 0′, 41° 30′, 42° 0′, 48° 0′, the latitude by observation on the first day 41° 0', and on the last day 45° 15'; differing 15 miles from the latitude by account; the distances sailed by the log, on the three days respectively, 30, 90, and 105 miles; we must say, as the whole sum of the distances 225 miles, is to the error of the latitude 15 miles, so is the first distance 30, to the correction of the second latitude 2', and so is the sum of 30 and 90 (=120) to the correction of the third latitude 8, so that the corrected latitudes will be $11^{\circ}0'$, $41^{\circ}30' + 2' = 41^{\circ}32'$, $42^{\circ}0' + 8' = 42^{\circ}8'$ and $43^{\circ}15'$, and the corrected differences of latitude on the successive days will be 32', 36', and 67', with which and the departure by dead-reckoning, the corrected courses, distances, &c. on each day may be found, if thought necessary; but as the corrected longitude is not sensibly altered by any of these corrections, it appears to be in general wholly unnecessary to make any alteration in the Journal on this account. But if it be thought proper to notice these corrections in platting off the track of a ship, it will be necessary first to plot off the courses by D. R. and then to place the points arrived at, at the end of each day, as much to the porth or south of the places by D. R. as will make the latitudes of those points agree with the corrected latitudes found by

The latitude and longitude by dead-reckening being found by the preceding methods, thence may be determined the bearing and distance of the place of destination; but when the mariner is fearful that his longitude by account is inaccurate, and he has no lunar observations to correct it, he must get into the latitude of the place, and (if possible) run east or west according

to his situation, and the prevailing state of the winds.

We have now given all the rules necessary for working a day's work and for the convenience of the learner (to enable him to refer to them easily) we have here collected them in the seven following articles.

Rules for working a day's work.

1. Correct the several courses sailed* for variation and lee-way, and enter them in a traversetable, and opposite to each course place the distance run on that course, found by summing up the knots and fathoms sailed by the ship on that course. Find in Table I. or II. the difference of latitude and departure corresponding to each course and distance, and set them in their respective columns: then the difference between the sums of the northings and southings will be the difference of latitude made good, of the same name with the greater; and the difference between the sums of the eastings and westings will be the departure made good, of the same name with the greater quantity.

2. Seek in Table I. or II. until the above difference of latitude and departure are found together in their respective columns; opposite to these will be the distance made good, and at the top or bottom of the page, according as the departure is less or greater than the difference of latitude, will be found

the course.

3. If the latitude from which the ship's departure is taken, or yesterday's latitude, be of the same name as the difference of latitude, add them together; but if of different names, take their difference; the sum or remainder will be the present latitude, of the same name as the greater.

4. Find the middle latitude between the latitude of yesterday and this

[•] The set and drift of a current (if there be any) is to be reckoned as a course and distance, and on the first day after losing sight of the land the bearing and distance of it are to be taken into account.

day, which take as a course in Table II. and seek for the departure in the column of Diff. Lat. then will the distance corresponding, be the difference of longitude, of the same name as the departure.

5. If the longitude in yesterday be of the same name as the difference of longitude, add them together; but if of different names, take their difference; the sum or remainder will be the long. in, of the same name as the greater.

6. If a lunar observation were taken at any time of the day, you must find, by the above method, the difference of longitude made since taking the observation for regulating the watch, and thence the longitude in at noon by that observation, and enter it in the Journal as the longitude by observation.

7. Find on a general chart the spot corresponding to the latitude and longitude by observation, and that place will represent the situation of the ship, whence the bearing and distance of the intended port may be found. The same may be obtained by middle latitude sailing, by inspection of Table II. thus: Find the middle latitude between the place of the ship and the proposed place, and seek for that latitude as a course in Table II. and find in the corresponding page of the Table, the difference of longitude (between the ship and the proposed place) in the distance column, opposite to which, in the latitude column, will be the departure. Seek in Table I. for this departure and the difference of latitude (between the ship and the proposed place) till they are found to agree, corresponding thereto will be the bearing and distance required. If the magnetic bearing be required, the variation must be allowed on the true bearing; to the right hand if the variation is westerly, or to the left hand if easterly.

We shall now proceed to exemplify the above rules; first by a few examples of separate day's works, and then by a Journal from Boston to Madeira,

kept in the usual form.

EXAMPLE I.

Yesterday, at noon, we were in the latitude of 48° 21' N. and the longitude of 36° 28' W. and have sailed till this day at noon, as per log-board; required the course and distance made good, with the latitude and longitude in?

1	LOG-BOARD.												
H	K.	F.	Courses.	Winds.	LW	Remarks.							
2 4 6 8 10 12 2	5 5	5		, N. N. W.	1	These 24 hours moderate gales and cloudy weather. At 4 P. M. speke ship Washing- ton, from New-York, bound to Cork.							
4 8 10 12			S. W. 1 S.	w.n. w	1	At 6 A. M. stowed the anchors and unbent the cables and coiled them between decks. Variation 21 points westerly.							

•							
TRAVERSE TABLE.							
Courses.	Dist.	N.	S.	E. ,	W.		
s. w. 🕯 s.	43		55.2		27.5		
SSW. 4 W.	39				18.4		
S PA M W	27		25.8		7.8		
	Diff.	Lat.	93.4	Dep.	58.5		
	Courses. S. W. J. S. SSW. J. W. S by W J. W.	Courses. Dist. 8. W. 1 S. 45 8. W. 1 W. 39 8. by W 1 W. 27	Courses. Dist. N. 8. W. 1 S. 45 8SW 1 W. 39 8 by W 1 W. 27	Courses. Dist. N. S. 8. W. 1 S. 45 8SW 1 W. 59 8 by W 1 W 27 27 85.8	Courses. Dist. N. S. E. S. W. 1 S. 45 SSW 1 W. 39 Sby W 4 W 27 25.8		

you must reckon 10 to a mile; and when the tenths are above 5, you must add one mile to the distance. Having found the distances you must find the corresponding differences of latitude and departures, in Table I. or II. and then with the whole difference of latitude and departure, find the course and distance made good, and the difference of longitude, by Case II. of middle latitude sailing.

In the present example, Yesterday's latitude	the	diff	erenc •	e of la	titude	is 95':	= 1 ⁰ 48	33′ 21	s. N.
The difference is the lati	itud	e in		·		•	46	48	Ŋ.
Sum of latitudes .	٠.		•				95	8	•
Middle latitude					•	•		54	
With the difference of latitud	le m	ade	good	93.4	S. an	d the d	eparti	are :	55.5
W I anter Table II, and find t	hev	COL	respo	nd ne	ariv t	o a cou	rse o	ıs.	30v
W and distance 108 miles. Th	nen '	with	ı the :	middle	e latit	ude 47	34 Q	r 48	٧, 1
enter Tuble II, and find the de	Dar	ture	55.5	in th	e lat.	columi	ı, opp	osit	e to
which, in the distance column,	is t	he d	iff. o	flong	. 80′	==	10	20'	W.
Longitude left							56	28	W.

^{*} As these examples were given only to illustrate the rules, we have not been attentive to mark the true variation.

† In India voyages it is customary to mark the log-board every hour; in that case, the distances marked on the log being summed up, will be the true distance sailed.

Sum is the longitude in

EXAMPLE II.

Yesterday at noon we were in the latitude of 350 46' N. and the longitude of 170 42' W. and have sailed till this noon as per log-board: required the

			G-BO	
HK.F.	Courses.	Winds.	LW.	Remarks.
2 6 6 3 5 8 4 5 8 5 5 8		S. W. & W.	14	These 24 hours moderate gales and clear weather.
7 5 8 8 5 8 9 5	ļ	s. w.	14	
0 5 1 5 2 2 5 2 1 5 8		•		At 8 A. M. saw a ship to wind- ward, steering east.
2 5 5 5 5 5 4 5 5	S. S.E. 4E.	S.W.byS.lW	13	
5 5 5 6 5 5 7 5 5 8 5 5		-	,	
9 5 6 0 5 6 1 5 4	S. E. by S.	S. W. by S.	14	
2 5 4	ł	 	<u> </u>	Variation & point easterly.
nd varia but not o d for ev oined T	tion, and the d doubled since t very hour) will raverse Table.	rrected for lee- listances summe he log-board is m I stand as in the Hence the di good is 105.4 S.	d up ark- ad- ffer-	5. E. § S. 31 24.9 18.5
е дера	rture 81.7 I 8. 38° E. a	consequently no the distance	the	3. E. 1 E. 22 14.8 16.5 Diff. Lat. 105.4 81.7 Dep
Diff	tude left of lat.	35° 46′ N. 1 45 S.	a	With the middle latitude 34° 53' or 35° and the departure \$1.7, the diff. of long. is out to be 100 miles 1° 40' E.
Sun	itade in 1 of lats. Idle lat.	34 1 N. 69 47 34 53		Longitude left 17 42 W. Longitude in 16 2 W.
atitude . St. Vi	To find the in 34 neent's lat. 37	1'N. Me	r. par	nce of Cape St. Vincents. ts 2173 Long. in 16° 2' W. ts 2394 C. St. Vin. lon. 9 2 W.
iff. of	lat. 3	0=180' Me	r. diff.	lat. 221 Diff. long. 7 0=420
	6 0 A 1 · ·		LOGAR	SITHMS.
	To find the diff. lat. 221	log. 2.3	4439	To find the distance. As radius 459 10.0000
s to Yad lo is dif		10.0 0 log. 2.6	0000 1	s to prop. diff. lat. 190 2.2559

2.62325 So is secant course 62° 15'

10.27886 To the distance 386.6

Hence the bearing of Cape St. Vincents is N. 62° 15' E. and distant 386.6 miles.

420 log.

So is diff. long.

To tang. course 62° 15'

10.33197

2.58724

EXAMPLE III.

Suppose that at the end of the sea day, March 10, 1824, we were in the latitude of 45° 34′ N. and the longitude of 55° E. and have sailed till next noon as per log-board; required the latitude and longitude in, and the variation of the compass?

		M-			LOG-BOARD.	
H	K.	F.	Courses.	Winds.	L.W	Remarks.
2 4 6 8 10 12	4	5 5 5	w.s.w.	South.		These 24 hours moderate gales, found a small current setting N. E. at the rate of one mile in 4 hours.
4 6 8 10 12	95 95 95 95		SW by W	S. by E.		At 8 A. M. sun's magnetic azimuth N. 125° 19' E. Alt. of ©'s L. L. 18° 40'; correction for dip and semi-diameter, 12' additive.

In calculating the variation from the above observation, it is necessary to find the declination and latitude at the time of observation. The former at noon ending the sea-day, March 11, 1824, was 3° 36'S. by Table IV. the correction for the long. 35° E. is+2'14"; and for the time from noon 4h. is +3 51", therefore the whole correction is nearly 6', which added to 3° 36' gives the declination at the time of observation 3° 42' S. consequently the polar distance 93° 42'. To find the latitude we must see by the log-board what courses and distances the ship has sailed from noon to the time of observation at 8 A. M. viz. W. S. W. 58 miles, and S. W. by W. 19 miles; the current setting in the same time N. E. 5 miles; these courses must be corrected for one point westerly variation, which is found to be nearly its value, by a rough calculation made with the latitude in, the preceding noon; and by arranging these courses and distances in a traverse table, we find that the difference of latitude made good at 8 A. M. is about 41 miles, consequently the latitude in at the time of observation is nearly 42° 53' N. the observed altitude of the sun's L. L. is 180 40'; the correction for dip and semi-diameter+12', and the refraction by Table XII.—3' nearly, consequently the sun's correct altitude is 180 49'. With these data, the true azimuth is calculated as in page 113. Dolar dist 090 40

Polar dist.	yy∵	42		
Latitude	42	55	Secant	0.13505
Altitude	18	49	Secant	0.02335
Sum	155	24.		
1 Sum	77	42	Co-sine	9.32844
Polar dist.	93	42	•	
Remainder	16	0	Co-sine	9.98284
			Sum	19.47018
Half sum is log.	co-si	ne 570	5′	9.73509
			2	
True azimuth	N.	114	10 E.	
Mag. azimuth	N.	125	19 E.	
Variation		11	9 W.	or nearly 1

				Pomi	•	
The variation being allowed on all the courses, and on the set of the current, and		RAVE	SE T	ABLE.		
the distances being summed up, the traverse		Dist.	N.	S.	E.	W.
table will be an adjoined : and the difference				-		
of latitude made good=49.8 S. departure	S.W.by W.	58		32.2		48.2
=67.5 W. Hence the course made good	S. W.	52		22.6		22.6
S. 5320 W. and distance=84 miles. And	N.E. by N.	6	5.0		3. 3	~~.0
by subtracting the difference of latitude 50'						
from latitude left 43° 34', there remains the			5.0	54.8	3.3	70.8
latitude in 42° 44' N. Hence we have the		!	• • • • • • • • • • • • • • • • • • • •	5.0	3.0	
middle latitude 43° 9', with which and the		1 1		y.0		3.3
departure 67.5, the difference of longitude		D:C	atized b)()()	
is 92' or 1° 32' W. nearly: and by sub-	1	DIII.	Lat.	49.8	Dep.	67.5

tracting it from the longitude left 35° E. we have the longitude in 33° 28' E.

EXAMPLE IV.

Yesterday at noon we were in the lat. of 40° 19' N. and in the long. of 67° 58' W. and have sailed till this moon as per log-book; required the bearing and distance of Cape Cod?

LOG-BOARD.

H	K	F.	Courses.	Winds.	LW	Remarks.
1	1		w.n.w.	North.	1	First part of these 24 hours light breezes
2	1				1	and fine weather; latter part pleasant
3	1					gales and cloudy.
4	1				l	·
5	2	5				
6	3	_			1	,
7	1	5			1	1
8	1	5	'		1	l t
9	1	Э			1	
10	1		AT UIT	N. N. E.	1	Saw great quantities of gulf weed, and rock
12	1		14. 44.	14. 14. 15.	.	weed.
1 7	o.	5	N.W.JN.	NEIE	1	weed.
2	2	5	74.44.324		1	, , , , , ,
5	2	5				
4	2	5				A & & A & directoring counded no
5	3	•	N. N. W.	NE by E	0	At 7 A. M. water discoloured, sounded no
6	3			, ,		bottom.
7	3				1	·
8	3			'	}	
	4				1	1
10	4				1	
11	4	5		E. N. E.		Latitude by observation 40° 52' N.
112	4	5		ļ	ł.	Variation 2 point W.

The distances are to be summed up, and marked in the traverse table without doubling, because the log-board is marked for every hour. By working this day's work like the others, we find the diff. of latitude made good=31. 6m. N. and the dep. 40. 5ms. W. hence the course N. 52° W. nearly, and distance 51 miles.

TRAVERSE TABLE.								
Courses.	Dist.	N.	S.	E.	W.			
W. 1 N.	15	0.7			15.0			
NWbyW#W	2	0.9	1		1.8			
NWbyW1W	10	5.1		•	8.6			
N.N.W. W.	29	24.9			14.9			
Diff.	Lat.	31.6		Dep.	40.5			

Latitude left Diff. of latitude	400	19' N. 32 N .	With the mid. lat. 4 departure 40.3 the diff.	of longitude
Latitude in by D. R.			is Long. left	0° 53′ W. 67 58 W.
Sum of lats. Middle latitude	81 40		Long. in	68 51 W.

To find the bearing and distance, of Cape Cod.

4.	Jum	ance occurring una	unitable of Cupe Cou.		
Lat. in by obs. Lat. of Cape Cod	40 ^C l 42	52' N. 5 N.	Long. in by D. R. Long. of Cape Cod	68 ⁰ 70	51' W. 4 W.
Diff. of lat.	1	15=73 miles.	Diff. of long.	1	13=75m.
Mid. lat.	41	28	,		

With the difference of longitude 75 miles, and the middle latitude 41° 28'. or 41½°, I find the depar. 54.6 nearly, with which, and the difference of lat. 73 miles, the bearing of Cape Cod is found to be N. 57° W. distant 91 miles.

JOURNAL.

OF A VOVAGE FROM BOSTON TO MADEIRA

H.	K.	F.	Cours	es.	Winds.	LW	Remarks on board, Thursday, Mar. 25, 1824
1 2 3 4 5							At noon got under way, with a fine breeze from the N. W.
678901	6 6 6	5 5 5 5	E. by	s.	N.W.	-	At 8 P. M. Cape Cod light-house bore S. S. E. 4 E. distant 12 miles; from which I take my departure.
1 2 3	6 6						
4 5 6 7	6 6	5			North.		. •
5 6 7 8 9 0	6	- 1					Variation ‡ point westerly.
5 6 7 8 9 10 11	6 6 6 6 7 7	5	Dist	In:	ff. Den	Lat. D.	by Lat. by Diff. Long. Possing and Diet

Cape Cod bearing from the ship S. S. E. ‡ E. distant 12 miles, is the same as if the ship had sailed from it 12 miles upon the opposite or N. N. W. ‡ W. point of the compass, and al-lowing for the variation, it becomes N. W. by N. this and the distance 12 miles, are to be in the traverse table as the first course and dis-Lance

5 94

The ship sailed all day upon an E. by S. course by sompass, which, by allowing the variation is E. ‡ S. The whole distance sailed (or the sum of all the distances) is 101 miles. With these course and distances, I find the corresponding differences of latitude and departures; and by subtracting the southing from the northing, and the westing from the easting, find that the difference of latitude made good is 5.0 N. and the departure 94.2 E, which correspond to a course of N. 860 58 E. and distance 94 miles.

	TRAVE	RSE T	ABLE		
Courses.	Dist.	N.	S.	E.	W
NW by N E. 4 S.	12 101 4	10.0	5.0	100.9	6.7
		10.0 5.0	5.0	100.9 6.7	6.7
e of N. 960 A8'	D.La				

2º 7' | 67º 57' distance 2493 miles.

420 10

Diff.	or Cape Cod's lat. of lat.	0 5 N. Then with	II. and against the	departure 94.2 for
Latite		42 10 M. the latitude	s the nearest tabular column, is 127—the di listance column.	number) found in Merence of longi
	of lats. e latitude		om, or Cape Cod's los Diff. Long.	ng. 70° 4′ W.
			Long. in	67 57 W.
Latitude in Funchal's lat.	To find the 42° 10' N. 32 38 N.	bearing and distance Mer. parts 2795 Mer. parts 2073		670 <i>57'</i> W. 16 54 W.
Diff. of lat.	9 32 60	Mer. diff. lat. 722	Diff. long.	61 3 60
In miles	572		In miles	9063

With the merick diff. lat. 722 miles, and diff. of long. 3063 miles, the bearing is found to be S. 760 44'

F. and with this bearing taken as a course, and the proper difference of latitude 872 miles, the distance is found to be 2493 miles, by Case L of Mercator's sailing.

Н	K	F	Cour	ses.	Winds.	LW.	Remarks on board, Friday, Mar	ch 26	, 18	24.
1	7		E. by	, S.	N.by E		Fresh gales and pleasant weather			
0	7									
5	7		178			1	Saw a number of fishing vessels t	o the	3 0 u	th-
4	7						ward.			
5	7				i i					
6	7					1	,	_		
7	7		E.by	SIS	N.N.E		At noon observed the altitude o	_		
8	7				1		the sun's lower limb bearing			
9	7		1		j		south		27	N.
10	7		1]	-	Add for semi-diameter, dip, &c.		12	
11	7		1		1		Refraction being small is neglected	l		
10	7				į	i	Correct altitude	50	3 9	
1	7		E. S	. E.		Ì	Subtract from	90	00	
2	7					1				•
3		1				i	⊙'s zenith distance	39	21	N.
4	6	1 '			1	1	⊙'s correct declination	2	22	N.
5	6		1			ļ	Latitude by observation	41	49	N
7	6	1 '					Latitude by observation	41	40	7.4
8	-	1			1	1				
9	0	1	1		·	1	` .			
10	0	1				1				
11	1 7		1			1				
12	6	1	1		1	1	Variation & points wasterly			
	1 0		1		<u> </u>	!	Variation 3 points westerly.			
C	ou	rse	Dist	t. Di	ff. Dep.	Lat.l	Dy Lat.by Diff. Long. in. Bearing.	ıg an	d D	ist.
		-		- 8		N.	N. E. W. Funcha	J 76	0 27	E
58	00	15	E 16	2 2			3 410 43 30 35 640 22 distanc			

The variation being allowed on each course, and the distances summed up, they will stand as in the adjoining traverse table; from hence, by means of Table I. I find the difference of latitude 27,5, and the departure 160,0, which corresponds to the course S. 30° 15′ E. and the distance 162 miles.

TRAVERSE TABLE.

Courses. Dist. N. S.- E. W.
E. ‡ S. 42 2.1 41.9
E. ‡ S. 42 6.2 41.5
E.S.E.‡E. 79 19.2 76.6

D.Lat. 27.5 160.0 Dep.

Yesterday's latitude Diff. of latitude		.,	420	10' 27		
Latitude in Sum of latitudes	,		41 83	43 63	N.	
Middle latitude			41	56		

with the muone latitude all 20 or 420 M acourse, I enter Table M. and seek for the departure 163,0 in the latitude column; the nearest number to which is 1583 corresponding to the distance 215, which is therefore the difference of longitude, equal to 30 M; Esterday's long. 67 57 W.

Long. in 64 22 W.

Latitude in Funchal's lat.	41° 43 ⁷ N.	he bearing and Mer. parts Mer. parts	2759	c of Funchal. Longitude in Funchal's long.	64° 22′ W. 16 54 W.
Diff. of lat.	9 5 60	M. D. lat.	686	Diff. of long.	47 28 60

In miles 545 In miles 2848

By Case I of Mercator's sailing, I find the bearing of Funchal to be S. 76° 27' E. and its distance
2826 miles.

H	K	F.	Cour	ses.	Win	ds.	LW	Remarks on board, Saturday, Mar. 27, 1824
1 2 3 4	8		E. S	· E.	N. by	E.		All these 21 hours fresh breezes and clear.
5 6 7 8 9	8							Mer. alt. sun's lower limb 51° 48' Add for semi-diam. dip, &c. 12
10 11	8							Sun's correct altitude 52 0 Subtract from 90 00
12 2 3	8	6			N. N	. Е		Sun's zenith distance Sun's correct declination 38 0 N 2 46 N
4 5 6	8	6						Latitude observed 40 46 N
8 9	8	6						·
10 11 12	7	-			NE	y N		Variation 3 point westerly, per amplitude
C	ou	rse	. Dist	Dil La	f. t. Der	$\mathbf{p} \cdot \begin{bmatrix} \mathbf{L} : \\ \mathbf{D} \end{bmatrix}$	it.by	Lat.by Diff. Long. in. Bearing and Dist.
E.	S.I	C4 F	E. 192	S 4			N. 0 56	N. E. W. Funchal S. 76° 40' 40° 46' 4° 8' 60° 14' E. dist. 2137 miles

TRAVERSE TABLE.							
Course.	Dist.	N.	S.	Ε.	W.		
E.S.E.IE.	192	D.Lat.	46.7	186.2	Dep.		

Middle latitude

The ship sailed all day upon the same course, which, corrected for the variation, is E. S. E. & E. the whole distance sailed is 192 miles, and the difference of latitude is 47

miles == Yesterday's latitude

0° 47′ S. 41 43 N.

64 22 W.

Latitude by D. R. 40

Hence the latitude by account differs 10 miles from the latitude by observation; but it will not be necessary to correct the longitude on account of this error.

Latitude yesterday by obs. 41° 43′ N. With the middle latitude With the middle latitude 41° 14' as a course, and the departure 186.2 as differ-Lat. by obs. this day, 40 46 N. ence of latitude, I find the corresponding Diff. of lat. by obs. 57 distance 248, which is equal to the differ-40 8/E. 82 Sum of latitudes 29 ence of longitude

60 14 W. Long. in Note. As this Journal is only designed to exemplify the rules of navigation, we have

Yesterday's long.

41

14

not endeavoured to give the true variation.

To find the bearing and distance of Funchal.

Latitude in 40° 46′ N. Mer. parts 2683 Longitude Longitude in 60° 14' W: Funchal's lat. 16 54 W. 32 39 N. Mer. parts 2073 Funchal's long. 43 Diff. of lat. 610 Diff. long. M. D. lat. 60 60

2600 In miles In miles 483 With the merid, diff. of lat, and diff. of long, the bearing is found to be 8, 769 46' E. with that and the proper diff. of lat, the distance is found to be 2137 miles,* by Case 1. Mercator.

II h

[&]quot;If the course was calculated to seconds, and the neritional parts taken to one or two places of definals, it would sometimes make a difference of a few miles in the calculated distance.

JOURNAL OF A VOYAGE

·F.	Courses.	Winds.	LW	Remarks on hoard, Sunday,M	ar.28, 1824.
7 7 6 6 6 6	1	NE byE.	1	Fresh gales with rain. At 4 A. M. spoke the ship Fra Philadelphia, bound to Lisbo	anklin, from on.
6 5 4 5 4 5 6		E. N. E.		Astronomod mon alt	
5 6 5 6 5 6		E. N. E.	1	At noon, observed mer. alt. sun's L. L. Add for semi-diameter, &c.	53° 58' 0 12
5 S 5 S 5 5				Sun's correct altitude Subtract from	54 5 90 00
5 5 6 6		E. by N.	1	Sun's zenith distance Sun's correct declination	35 55 N. 3 9 N.
6 6 6				Latitude observed	39 4 N.
6 5 5				Variation ‡ points westerly.	
ourse	Dist.	Diff. Dep.	Lat.	y Lat.by Diff. Long. Doggin	g and Dist.
	E 138	S. E.	$\frac{\mathbf{D} \cdot \mathbf{R}}{\mathbf{N} \cdot}$	l. Obs. Long. in.	al S. 79°7'E.
===	TRAVER	SE TABLE		The lee-way and var	iation being
urse	s. Dist.	N. S.	E	W. allowed on the course stand as in the adjoined	s, they will traverse ta-
5. 3	E. 50	29.	8 40.	ble. Then with the	lifference of
i. i	S. 44	32.	6 29.	latitude and departure t	ne course is 'E. and the
·E-3	E. 46	39.	5 28.	distance 138 miles.	
		Lat. 101.		3 Dep.	, -0 ==1
erda	cy's latitud ce of latitu	le 4 de 102'=	1 4	6' N. With the middle lat. 38 2 S. as a course, and the dep. as difference of latitude,	. 93.5, taken the diff. of
tude ı of	in latit ud es			4 N. long. is found to be 122 mile Vesterday's longitude	es = 2° 2′ E. 60 14W.
he contes,	as it was	de go <mark>od</mark> e calculate	ach e	5 Longitude in lay is marked in the journal to logarithms; but for practical p nearest degree by means of T	urposes, it is
		find the	bearii	ng and distance of Funckal.	
tude chal	in 's latitude	39° 4' N	•	Aiddle Latitude Sailing. Longitude in Funchal's longitude	58° 12′ W. 16 54 W.
	re of lat.		886 n	niles. Difference of long.	41 18
	latitudes atitu d e	71 42 35 51		In miles	2478
th th	e middle la	titude 35°	51' or	360 as a course, and the difference	e of longitude
us a the	uistance, l' distance an	caiculate ti d course, b	ne de _l 7 Case	parture; with that and the difference I. of Middle Latitude Sailing.	
		•		Digit	ized by Google

H.	K	F.	Co	urse	. Wi	nds.	LW	Rem	arks o	n board	i, Mond	lay, Mar	. 29,	18	24.
1 2 3 4 5 6 7	4 4 4 4 4 4		S	outh	E.\$	8.E.	1	The	se 24 h	ours m	oderate	, pleasar	nt w	eath	er.
8 9 10 11 12 1	4 4 4 4	6	s	. 4 E	. Ebs	54S	14	Add ⊙'s		ni-dian altitud	ver liml neter, d le		55° 0 55 90	32' 12 44 00	
2 3 4 5	5 5 5 5 5 5	4				İ			enith (correct				34 3	16 32	N. N.
6 7 8 9 10 11 12	3 3 3							V	ude ob	1 poir	nt weste		87	48	
C	ou	rse	е.	Dist.	Diff. Lat.	Dep	La D	t. by	Lat.by Obs.	Diff. Long.	Long. in.	Bearing	and	Di	вŧ.
:	Sou	ıth		86	S. 86	0		V. 9 58'	N. 37° 48'	0	W. 58° 12′	Funchal distance	S 81	O17 6 mi	ľE.

TRAVERSE TABLE.								
Course.	Dist.	N.	8.	E.	W.			
South.	86		86.0	Diff.	Lat.			

The lee-way and variation being allowed on both courses, they become south; the whole distance sailed or 86 miles, is therefore the difference of latitude by account, the departure being nothing; consequently the ship is

in the same longitude as yesterday.

Yesterday's latitude Difference of latitude 59° 4′ N. 86= 1 26 S.

Latitude in by D. R.

97 98 N.

The latitude by observation was 37° 48' N. differing 10 miles from the account; but this will not render it necessary to correct the longitude.

To find the bearing and distance of Funchal.

Latitude in Funchal's lat.		Mer. parts 2458 Mer. parts 2073	Longitude in Funchal's long	
Diff. of lat.	5 10 60	Mer. diff. lat. 380	Diff. of long.	41 18
In miles	510		In miles	9478

Hence the bearing is found to be S. 81° 17′ E. and the distance 2046 miles, by Case I. of Mercator's Sailing; and the same may be found by middle latitude, which is the most exact method when the two latitudes differ but little; and it is the way in which the calculation will be made in the rest of the journal.

HK.F. Courses. Winds. L.W	Remarks on board, Tuesday, Mar. 30, 1824.
1 3 East. N.NE. 3 3 4 3	These 24 hours fresh gales and squally. Handed the fore and main courses.
5 Lay to, up S. E. by E. 5 6 off S. E. by S. Drift 11 7 miles per hour.	
9 Up S. off S. W. Dein 5 10 1½ miles per hour.	·
12 1 2 5 E. by N. SEbyS 2 2 5 3 3 4 8 5 8 5	At midnight more moderate; wore ship and set the courses.
6 3 5 7 2 5 8 2 5 9 2 5 10 2 5	At 6 A. M. set the topsails close reefed.
$egin{array}{ c c c c c c c c c c c c c c c c c c c$	Variation 1 point westerly. by Lat.by Diff. Long. Bearing and Dist. Long. in.
N76°17′E 31 N. E. N. 37° 5	E. W. Funchal S. 80° 58'E. 5' 0° 38' 57°54' distance 2017 miles.

T	773 1.5					
Courses.	Dist.	N.	S.	E.	W.	Taking and S. S
E. S. E. South. W. S. W. N.E. & E.	12 6 6 32	20.3	6.0 2.3	24.7	5.5	the ship of in the rul as before the traver
D.	Lat.	12.9		35.8 5.5 30.3	5.5	

Taking the middle points, (viz. S. E. and S. S. W.) between the point to which the ship comes to and falls off, as taught in the rules of lying to, and then allowing as before for the variation and lee-way, the traverse table will stand as adjoined. With the difference of latitude and departure the course is found to be N. 76° 17' E. and the distance 31 miles.

Yesterday's latitude Difference of latitude	37° 48′ N. With the middle lat. 37 7 N. as a course, and the depar	ture 30.3 used
Latitude in Sum of latitudes Middle latitude	37 55 N as difference of latitude, ference of longitude to b Yesterday's longitude Longitude in	e 0° 58' E.

To find the bearing and distance of Funchal.

Latitude in 37° 55' N.	Longitude in	57° 54′ W.
Funchal's latitude 32 38 N.	Funchal's longitude	16 54 W.
Diff. of latitude 5 17=317 miles.	Diff. of longitude	40 40
Sum of latitudes 70 33	- ,	60
Middle latitude 35 16	In miles	2440

With the middle latitude 35° 16′ and the difference of longitude 2440, the departure is found to be 1992; with that and the difference of latitude 317, the bearing of Funchal is found to be S. 80° 58′ E. and the distance 2017 miles.

H.	K.	\mathbf{F} .	Cor	ırses.	Winds.	LW	Remarks on board, Wednesday, Mar. 31, 1824.
1 2 3 4 5 6 7	5 5 5	6 6 4	E. \$	S. E.	South.	1	Pleasant gales and fair weather.
8 9 10 11 12 1 2 3 4	6 6 7	4 5 5	ЕЬУ	SįS	S. 4 E.	ż	
5 6 7 8 9 10 11	7 7 7 7 7 7						Variation 1 point westerly per azimuth.
1-	ast	-	Dist	Diff. Lat		$\frac{\mathbf{N} \cdot \mathbf{R}}{\mathbf{N}}$	Lat. by Diff. Long. in. Bearing and Dist. E. W. Funchal S. 80° 12'

The variation and lee-way being allowed on both courses, it appears that the ship has made a due east course, the distance sailed 151 miles is the departure, and the difference of longitude is found by Case II. of Parallel Sailing. The latitude in is the same as yesterday's lat. 37° 55' N. Taking this as a course, and the departure 151 as difference of latitude, the distance which corresponds is the difference of longitude, 191 miles = 3° 11' E.

Yesterday's longitude	57 34 W.
•	
Longitude in	54 23 W.

To find the bearing and distance of Funchal.

Latitude in Funchal's lat.	37° 55′ N. 32	Longitude in Funchal's long.	
Diff. of latitude	5 17 = 317 miles.	Diff. of long.	37 29 W.
Sum of latitudes Middle latitude	70 3 3 35 16	In miles	2249 ·

Hence by case I. of Middle Latitude Sailing, the departure is found to be 1836 miles, the bearing of Funchal S. 80° 12' E. and the distance 1868 miles.

JOURNAL OF A VOYAGE

		7	Co	urs	es.	Winds	LW	Remarks on board, Thursday, April 1, 1	1824.
			E.	s.	E.	s.s.w		Fresh gales and pleasant weather.	
1 3 4 5 6 7 8 9 10 11 12	*********************	446655555	F	Zast	·•	S by W South. S.by E	ż	Sun's declination 4 4 Latitude observed 57 3 Variation 1 point westerly.	2 - 1
Co	ur	se.	. 1	Dist	. L	at. Dep	D.	by Lat.by Diff. Long. in. Bearing and I	Dist.
S85	೦೨	4/	E.	209	"	S. E. 6 201	N 370	N. E. W. Funchal, S. 79° distance 1658	

T	RAVEI	RSE T	ABLE			The courses being co	rreci	ed (for
Courses.	Dist.	N.	S.	E.	W.	lee-way and variation, the	trav	e rse 1	a-
E. by S. E. 4 S. E.N.E.4E.	100 70 35	10.2	19.5 6.9	•		ble will be as here given. Hence the course is S. distance 202 miles.		94']	
D.T.L.			26.4	201.3	Dep.	Yesterday's latitude Diff. of latitude	_	16	S.
			10.2			Lat. in by account	97	3 9	N.
Wish share	Diff.				- day	Yesterday's long.	54	25 7	w.
long. is			***		e dej	o. 201.3, the diff. of	. 4	15	E.
					1	Longitude in by account	50	8	w.

The latitude by observation differs 9 miles from the latitude by dead reckoning, but the longitude requires no correction on this account.

H	K.	F.	C	ourse	s. Wi	nds.	LW	Re	marks o	n boar	d, Friday	, April 2, 1824.
1 2 3 4 5	6 6 7 7	5 5 5 5	E	S. 1	E. So	uth.	Ą	F	resh ga	iles, w	ith rain.	
5 6 6 7 8 9 10 11 12 1 2 3 4 5 6 7	8888899999	5 5 5 5	Е.	s.]	E. S.	w.	0	T suri Pol	his daying the clux, the	took distanc longi	e of the	oservation, by mea- moon from the star oon, deduced from
8 9 10 11 12	9 9 9	5 5		-	(1):0:		15				nt wester	ly.
L	Cou	ırse	•	Dist.	Diff. Lat.	Dep	La D.	t.by R	Lat.by Obs.		Long. in.	Bearing & Dist.
87	90	56′	E	202	S. 35	E. 199	36°	7. 55'		E. 40 9'	W. 45° 59′	Funchal S. 79°50'É dist. 1456 miles.

	TRAV	ERSE T	ABLE		Ī	
Courses.	Dist.	N.	S.	E.	W.	The lee-way and variation being allowed on the courses, the traverse
E. 4 S. E. by S.	42 160			41.8 156.9		table will be as here given; hence the course was S. 79° 56' E. and the distance 202 miles.
1	1	D.Lat.	35.3	198.7	Dep.	

Yesterday's latitude Difference of latitude	37 ^C	30′ 35	N.	With the middle lat. 37 dep. 198.7, the difference is found to be 249 miles:	70 12' e of le	and	the tude
				is found to be 249 miles:	= 40	9′	Ε.
Latitude in	36	55	N.	Yesterday's longitude	50	8	w.
Sum of latitudes		25		, ,			
Middle latitude	37	12		Longitude in	45	59	W.

To find the bearing and distance of Funchal.

Latitude in Funchal's latitude		55' N. 58 N.	Longitude in Funchal's longitude		59′ W. 54 W.
Diff. of latitude	4	17=257 m.	Diff. of longitude	29 60	-
Sum of latitudes Middle latitude	69 54		In miles	1715	

Hence by Case I. of Middle Latitude Sailing, the bearing of Funchal is found to be S. 79° 50′ E. and its distance 1456 miles.

H	K.	F.	Cours	es Winds	LW	Remarks on board, Saturday, April 3, 1824.
1 2 3 4	9 9 9	6 4 4		E. West.		Fresh gales and rainy weather; latter part clear. A great swell from the N. E. for which I allow 9 miles.
6 7 8 9	9 9	5 5				Obs. alt. sun's lower limb at noon 58° 58' Correct. for semi-diam. &c. add 0 12
11 12 1 2	9	5		N. W.		Sun's correct altitude 59 10 Subtract from 90 00 Sun's zenith distance 50 50 N.
2 3 4 5 6 7	9 9 9					Sun's declination 5 27 N. Latitude observed 36 17 N.
8 9 10	9 9 9			North		
11	1 .]			Variation 14 point westerly per azimuth.
C	ou	rse	. Dist.	Diff. De	D.	by Lat.by Diff. R. Obs. Long. Long. in. Bearing and Dist.
S7	902	2'I	E. 217	S. E. 40 21		

7	TRAVI	ERSE	TABI	E.	
Courses.	Dist	N.	S.	E.	W.
E. 4 S.	220	•	32.3	217.6	
SSW.W	9		7.7		4.6
		D.tat	40.0	217.6	4.6
				4.6	
			Dep.	213.0	

In this day's work the swell is considered as a current setting the ship 9 miles per day; and since the swell comes from the N. E. it must set the ship S. W. and allowing the variation S. S.W. W. 9 miles, these are placed as a course and distance in the traverse table.

With the difference of latitude and departure the course is found to be S. 790 22' E. and the distance 217 miles.

With the middle lat. 36°38', and the dep. 213 miles, the diff. of long. is 40 25' E. found 265 miles==

45 59 W.

Yesterday's long. 34 W. Longitude in 41

Yesterday's latitude 36° 55' N. 40 S. Difference of latitude 15 N. Latitude in 36

To	find	the b	earing and	l distance of Funchal.	_		
Latitude in			N.	Longitude in	410	31'	W.
Funchal's latitude	32	38	N.	Funchal's longitude	16	51	W.
Difference of latitude	5	39=	=219 m.	Diff. of Long.	24	40	
Sum of Latitudes	68	55			60		
Middle latitude	34	27		In miles	1.18	0	

Hence by Case I. Middle Latitude Sailing, the bearing of Funchal is found to be S. 79° 50' E. and its distance 1240 miles.

to be S. 79° 50' E. and its distance 1240 miles.

To find the bearing and distance of Funchal by Mercator's Chart.

Having pricked off the place of the ship at noon, lay a ruler from the point to Funchal; take the nearest distance between the centre of the compass and the ruler; then slide one foot of the compasses along the edge of the ruler, keeping the other foot at the greatest distance from it, and it will be found to run nearly upon the E. by S. line, which is therefore the bearing of Funchal; then take in your compasses the extent from the place of the ship to Funchal, and apply it to the graduated meridian, setting one foot as much above one place as the other is below the other place, and the extent will be found to measure 204 degrees, or 1230 miles, which was the distance of the ship from Funchal nearly.

1 2 3	7 7 6 6	-	-	_	-	Winds N. E.	1	Remarks on board, Sunday, Apr First part fresh gales; latter part derate, a heavy sea running.	-	_
5 6 7 8 9	6 6 5 5 4 4	4 4 6 6	8	l. E		E.N.E		Mer. alt. sun's lower limb Correction for semi-diameter, &c. Sun's correct altitude Subtract from	61 61 90	12
11 12 1 2 5	4444				- 1	East. E.byS.	**	Sun's zenith distance Sun's declination Latitude observed	5	45 N. 50 N. 35 N.
5 6 7 8	4 4 4 4	5	8.	b y 1	Ξ.		14	-	•	
10 11	4 4 4	rs	e.	Di	st.	Diff.	ep.	Variation 14 points westerly. Lat.by Lat.by Diff. Long. Bearing D. R. Obs. Long. in.	- and	l Diet
S- 3	_	_	_	-		S.	E.	N. N. E. W. Funchs 40 55' 340 35' 10 18' 400 16' E. dist.	18.8	340 17

. TRAVERSE TABLE.										
	Dist.	N.	S.	E.	w.					
E. S. E. 4 E.			13.5	37.7						
9. E. į Ě.	20		13.4	14.8						
S. S. E. 4 E.	8		7.2	3.4						
S. by E.	16		15.7	3.1						
S. 1 E.	33		32.6	4.8						
	Diff.	Lat.	82.4	65.8						

The courses being corrected for lee-way and variation, will stand as in the adjoined traverse table.

Then with the difference of latitude 82.4 & the departure 63.8, I find the course S. 37° 45' E. Yesterday's latitude 360 17' N. Difference of latitude 1 22' S.

Lat. by account 84 55 N.

Yesterday's latitude Latitude in by obs.	36 ⁰ 54	9 17' N. 85 N.	With the dep. 63.8 mid. lat. 35°26', I find	miles, and the
Sum of latitudes Middle latitude	70 85	52	to be 78 miles = Yesterday's long.	1° 18′ E. 41 54 W.
			Longitude in	40 16 W.

To find the bearing and distance of Funchal.

Latitude in		9 55' N.	Longitude in	40° 16′ W,
Funchal's latitude		58 N.	Funchal's longitude	17 54 W.
Diff. of latitude Middle latitude	1 55	57=117 miles.	Diff. of longitude	23 22

In miles 1402

Hence by Case I. Middle Latitude Sailing, the bearing of Funchal is found to be S. 840 17' E. and its distance 1174 miles. Digitized by Google

H.	K.	F.	Cou	ırses	·W	inds.	LW	Remarks	on boa	rd, Mor	iday, Ap	ril 5	, 18	324.
1 2 3 4 5 6 7 8	3 3 2 2			E.	E.	N.E	1	and calm At 4 F current;	hours sm sh gales. e boat ar ing E. hip has b	nd tri I mi	ied le	the per		
9 10 11 12 1 2 3	S S 4	4	E. S	5. E	. N.	N.E		Mer. alt. Correction Sun's cor Subtract	on for ser rrect alti from	mi-dian tude		61° 0 61 90	12 51 00	
5	4 5 5 6	6 5 5						Sun's zenith distance Sun's declination				2 8 6	9 12	N.
6 7 8 9 10	6 7 7	5 5						Observed	l latitude	:		54	21	N.
11	8							Variat	ion 1½ p	oint we	sterly.			
C	oui	se.	Di	st.)iff. ⊿at.	Dep	Lat D.	by Lat. b	y Diff. Long.	Long. in.	Bearing	and	Di	st.
S.8	3 0	36′	- -	01	S. 11	E. 100	N	N.	E.	w.	Funchal distance			

TRAVERSE TABLE.										
Courses.	Dist	N.	S.	E.	W.					
S. E. 4 E.	10		6.7	7.4						
E. S.	70		10.3	69.2						
E.N.E.E.	24	5.8	i i	23.3						
		5.8	17.0	99.9	Dep.					
1	1		5.8							
+	Diff.	Lat.	11.2							

In addition to the courses sailed, I also allow 24 miles for the set of the current in the direction of east per compass, or E. N. E. & E. true course.

With the difference of latitude 11.2 With the middle lat. 340 28', and and the departure 99.9, the course is the dep. 99.9, I find the diff. of long. found to be S. 83° 36' E. and the dis-to be 121 miles= 20 1' E. 16 W. tance nearly 101 miles. Yesterday's longitude 40 Yesterday's latitude 34° 35' N. Difference of latitude 11 S. Longitude in 38 15 W. Latitude in by account 24 N. 34 To find the bearing and distance of Funchal. 38° 15' W. 34º 21' N. Longitude in 54 W. 38 N. Funchal's longitude 16

Latitude in \$4° 21' N. Longitude in \$8° 15' W Funchal's latitude 32 38 N. Longitude in Funchal's longitude

Difference of lat. 1 43=103 miles. Difference of long. 21 21 21 60

Middle latitude 33 30 nearly. In miles 1281

Hence by Case I. of Middle Latitude Sailing, the bearing of Funchal is found to be S. 84° 50′ F., and its distance 10°3 miles.

H	K	F.	Co	ur	ses.	Wind	sLW	Remark	s on h	oard, Tue	sday, Ap	ril 6	, 18	24.
1 2 3 4 5 6 7	99999		E.	s.	Ε.	North	1.	Fine fresh gales and clear weather.						
8 9	9 9									lower limi dip, &c.	b	60°	85′ 12	•
10 11 12	9 9				i			Sun's co Subtrac				62 90	47 00	
1	9							Sun's ze Sun's de				27 6		N. N.
4 5	9							Observe	ed latit	ude		53	48	N.
2 3 4 5 6 7 8 9 10														
11 12	9				ļ			Varia	tion p	er Amp. 1	4 point w	reste	ıly.	
Co	ur	se.	Dis	t.	Diff Lat	Dep.	D. R	y Lat.by Obs.	Diff.	Long. in.	Bearing	and	Di	st.
E.	4 5	3.	21	6	S. 32	E. 214	N. 53° 4	N. 9'53° 48'	E. 4º 18		Funchal E. dist.			

The course corrected for variation is E. § S. distance 216 miles; hence the difference of latitude is 31.7, and the departure 213.7 miles.

Yesterday's latitude Difference of latitude	3 40	21' N. 32 S.	
Latitude in Sum of latitudes by obs.	33 68	48 N.	
Middle latitude	34	4	-

With the middle latitude 34° 4′, and the departure 213.7 miles, I find the difference of longitude to be,258 miles = 4° 18′ E.

Yesterday's longitude 58 15 W.

Longitude in SS 57 W.

To find the bearing and distance of Funchal.

Latitude in Funchal's lat.		48' N. 38 N.	Longitude in Funchal's long.	93 ⁰		
Diff. of latitude	1		Diff. of long.	17	_	w.
Sum of latitudes	66	26		60	-	•
Middle latitude	33	13	In miles	1023		

Hence the bearing of Funchal is found to be S. 85° 19' E. and its distance 359 miles.

H	K	F.	Cours	es. V	Vinds.	LW.	Remark	s on bo	ard, Wed	inesday, A	p. 7, 18 2 4.
1 2 3 4 5	10 10 10 10 10 8 8 8	44665		E. P	INW.]	Fresh arge sea At 4] measuring the sun;	gales P. M. ng the	and pleas took a le	unar obser	er, with a rvation by noon from o noon by
9 10	8										
lii	8	ł									- 1
12							Varia	tion pe	r azimutl	h 14 point	westerly.
C	our	se.	Dist.	Diff. Lat.	Dep.	Lat.by D. R.	Lat.by Obs.	Diff. Long.	Long. in.	Bearing	and Dist.
S8	009	01	£ 210	8. 35	E. 207	N. 33° 13		E. 40 8'	W.	Funchal a	86 ⁰ 55' E 652 miles.

TRAVERSE TABLE.										
Courses.	Dist.	N.	S.	E.	W.					
E. 1 S. E. by S.	60 150			59.7 147.1						
			35.2	206.8						

By the adjoined traverse table, the difference of latitude is 35.2, and the departure 206.8; hence the course was S. 80° 20′ E. and the distance 209.8, or 210 miles.

Yesterday's latitude Difference of latitude	33 0 4		With the middle lat and the departure 206.8,		
Latitude in by account Sum of latitudes			of long. 248 miles, or Yesterday's longitude	40	
Middle latitude	55 5	10	Longitude in	29	49 W.

To find the bearing and distance of Funchal

_ · Ju		con units w	u ausumce oj rumcnu	•			
Latitude in Funchal's latitude	33° 32	13' N. 38 N.	Longitude in Funchal's longitude				W. W.
Diff. of latitude Sum of latitudes Middle latitude	65 32	35 51 55	Diff. of longitude	· ` ;	12 60	55	

In miles 775
Hence the bearing of Funchal is found to be S. 86° 55' E. and its distance 652 miles.

H	K	F.	Cou	ses.	Winds	LW	Remarks on board, Thursday, Ap. 8, 1824.
1 2 5 4	8 8	5		Sis	N.N.E		First part fresh gales and clear. Latter part rainy weather.
5 6 7 8 9	8 8 8 8 8	5	S.	E.	E.N.E	ł	
11 12 1 2 3 4 5	8 8 8						At 6 A. M. the wind hauled suddenly to the S. S. E.
7 8 9	777	5 5 5	_	st.	s.s.e	. 1	
10	7					ł	
12	ı	5			1		Variation 14 point westerly.
C	ou	rse	. Dis	t. Di La	ff. Dep	Lat. D. F	by Lat.by Diff. R. Obs. Long. Long. in. Bearing and Dist.
Sa	3 0,	45′	E 17	2 S		N. 520 5	

T	RAVE	RSE T	ABLE		
Courses.	Dist.	N.	S.	E.	W.
East.	50			50.0	
S. E. by E.	80	1	44.4	66.5	
NE.byElE	60	25.7		54.2	
		25.7	44.4 25.7	170.7	Dep
	Diff.	Lat.	18.7		

The leeway and variation being allowed on the courses, they will stand as in the adjoined traverse table; then with the difference of latitude 18.7, and the departure 170.7, the course is found to be S. 85° 45′ E. and the distance 172 miles.

Yesterday's latitude	33 0	13' N.	With the middle lat. 33	െ ഉ∕	and	the
Difference of latitude		19 S.	dep. 170.7, I find the di nearly 204 miles ==			g. is E.
Latitude in	32	54 N.	Yesterday's longitude			
Sum of latitudes	66	7				
Middle latitude	5 5	8	Longitude in	26	25	w.
To find	the be	aring a	ed distance of Funchal.			
Latitude in			Longitude in	260	25'	w.
Funchal's latitude	52	38 N.	Funchal's longitude	16	54	w.
			•			
Diff. of latitude		16	Diff. of longitude	9	81	
Sum of latitudes	65	52		60		
Middle latitude	32	46	·			•
			T	571		

In miles 571

Elence the bearing of Funchal is found to be S. 88° 5′ E. and its distance 480 miles.

H	K.	F.	Course	s. Winds.	LW	Remarks on board, Friday, April 9, 1824
1 2 3	7 7 8	5 5	E.byS.	South.		Fine breezes, with variable weather.
4	8					•
5 6	8	5 5		1		Mer. alt. sun's lower limb 64° 33'
7	9	1				Correction for dip, &c. 0 12
8 9	9					Sun's correct altitude 64 45
	9					Subtract from 90 00.
10 11	9					S
12	9					Sun's zenith distance 25 15 N. Sun's declination 7 41 N.
1	9					
1 2 3 4 5	9		E. by			Observed latitude 52 56 N.
4	8	1	ш. оу .	· *		
5	9				•	
6	9					•
7 8	9					•
9	9					
10	9	1				·
11 12	9			•		Variation 14 point westerly.
-	our	se	Dist.	Diff. Dep.	Lat.l D. F	by Lat. by Diff. Long. in. Bearing and Dist.
N	90	12	E 210	N. E. 209	N. 320 5	N. E. W. Funchal, S. 86° 11' E. 57' 32° 56' 4° 10' 22° 15' distance 270 miles.

TRAVERSE TABLE.										
Courses.	Dist.	N.	S.	E.	W.					
E. 4 S. E. 4 N.	120 90	8.8	5.9	119.9 89.6						
		8.8 5.9	5.9	209.5	Dep.					
Diff.	Lat.	2.9								

The variation being allowed on the courses, they will stand as in the adjoined table; then with the difference of latitude 2.9, and the departure 209.5, the course is found to be N. 89° 12′ E. and the distance 210 miles nearly.

Yesterday's latitude Difference of latitude	52° 54′ N. -3 N.			With the middle lat. the dep. 209.5, the dif- found 250 miles=		is	
Latitude by account	32 57		N.	Yesterday's longitude			
				Longitude in	22	15 V	v.
To find	the be	arinį	g an	d distance of Funchal.			
Latitude in	3 2 0	56'	N.	Longitude in	220	15' V	V.
Funchal's latitude	32	38	N.	Funchal's longitude	16	54 V	₹.
Difference of latitude		18		Difference of longitude	5	21	
Sum of latitudes	65	34			60		
Middle latitude	32	47					
, d				In miles	521		
Hence the bearing of Fun	chal is	found	đ to	be S. 86° 11' E. and its distar	ce 27	0 mil	es.

H	K.	F.	Co	ırses.	Winds.	LW.	Remarks on board, Saturday, Apr. 10, 1824.
1 2 3 4 5 6 7 8 9	9 9 9	5 5 5 5 5		by E	s.s.w.		All this day fine breezes, with very clear weather.
11 12 1 2 3 4 5 6 7 8 9	9 9 9 9 9 9 9 9 9		Eb	yS į S		•	At 10 A. M. made the land. The southern part of Madeira bearing per compass E. by S. I S. distant 19 leagues.
11							· · ·
12			_	. IDi	PE L	Lat.b	Variation 11 point westerly.
C	oui	se.	. D	ist. La	t Dep.	D. R	y Lat.by Diff. Long. Obs. Long. in. Bearing and Dist.
S	5 °:	57']	E 2	56 S		N. 320 2	E. W. 5° 5' 17°12'

TRAVERSE TABLE.										
Courses.	Dist.	N.	S.	E.	w.					
E.S.E.‡E. East. East.	111 90 57		27.0	107.7 90.0 57,0						
			27.0	254.7						

In the traverse table are placed the bearing and distance of the land at 10 A. M. (after allowing the variation.) Hence the whole difference of latitude is 27 miles, the departure 254.7, the course S. 85° 57' E. and the distance 256 miles.

		1		
Yesterday's latitude Difference of latitude	32° 56′ N. 27 S.	dep. 254.7, the diff. of l	ong. is found to b	16
Latitude by account Sum of latitudes	32 29 N. 65 25	303 miles = Yesterday's long.	5° 3′ E. 22 15 W.	
Middle latitude	32 42		17 19 W	

Therefore the latitude of the southern point of Madeira by my account is 32° 29' N. and its longitude 17° 18' W. these values differing but little from the values given in the table of latitudes and longitudes; I therefore conclude, that my journal was nearly exact; and that latitude and longitude of that part of Madeira were well laid down.

1 2				Pleasant gales and fair weather.
3 4 5				At 4 P. M. came to, off Funchal.
7 8 9 10 11				At 8 P. M. Went on ahore.
Cour	5es.	Dist.	Diff. Lat.	at. by Lat. by Diff. Long. in Bear. and Dist.

AN ABSTRACT OF THE FOREGOING JOURNAL.

Days.	Months.	. 1824.		Courses.	<u> </u>	Dist.	Lat. b	y D.R.	Lat	by 01	1 1	ng. in.	Be	rings and	Dist. Lat. by D. R. Lat. by Ob. Long. in. Bearings and Distances of Funchal at noon.	of Funcha	l at noon.
Thursday,	March	\$5	ż	880	860 58' E.	2	460	480 10 N.			670	57' W.	83	670 57' W. S. 780 44' E.	distant	9483	miles
Friday,	March	98	002	8	15 E.	162	41	43	410 45		Z.	3	20.	78 27 E.	distant	. 2526	mile
Saturday,	March	27	σż	75	56 E.	20	\$	28	2	8	8	14	83	76 48 E.	distant	2137	miles
Sunday,	Merch	83	oż.	94	29 E.	138	8	4	88	4	58 12	12	20.	79 7 E.	distant	2045	miles
Monday,	March	88		South.	4	88	87	38	37	84	28	18	20 20	81 17 E.	distant	2048	miles
Tuesday,	March	96	ż	ı	78 17 E.	31	37	55	_		57	\$	αż	80 58 E.	distant	2017	mile
Wednesday,	March	31		East		151	37	55			40	23	ω, ω	80 12 E.	distant	1869	miles
Thursday,	April	-	oż.	82	24 E.	202	87	39	37	8	20	80	30.	79 51 E.	distant	1658	· miles
Friday,	April	93	oż	82	56 E.	202	86	55			3	29	200	79 60 E.	distant	1456	mile
Saturday,	April	တ	σ'n	82	99 E	217	38	15	8	11	14	\$	S. 79	9 50 E.	distant	1940	miles
Sunday.	April	4	တ်	87	45 E.	104	\$	25	34 85	82	\$	16	αż	84 17 E.	distant	1174	miles
Monday,	April	2	øż	88	88 E.	100	\$	2	2	13	88	15	80 80	84 SO E.	distant	1073	miles
Tuesday,	April	8	<u> </u>	E. 4 3.	80	218	88	48	8	84	88	57	ος. O	85 19 E.	distant	828	miles
Wednesday,	April	7	σ'n	8	3 3	\$10	88	81.			83	48	æ ∞	86 55 E.	distant	652	miles
Thursday,	April	80	ø	83	45 E.	178	85	54			9	25	oó Oó	88 5 E.	distant	480	
Friday,	April	3	ż	88	18 E.	018	3	57	좕	28	22 15	15	88 .88	6 11 E.	distant	370	miles
Seturday.	April	10	200	83	57 E.	929	36	63	L	 	17 18	18	mad	e the land,	made the land, bearing E.by S. S. dist. 19 leag	5y 8.4 8.d	st. 19 lea
Sunday,	April	11						Z C	me to	anch	or at	4 P. M.	in F	Came to anchor at 4 P. M. in Funchal road.	id.		٠
												•					

EXPLANATION OF SEA TERMS.

4BACK. The situation of the sails, when their surfaces are pressed aft against the mast by the force of the wind.

Aboft er aft. The sternmost part of the ship; Carry aft any thing; that is, carry towards the stern. The mast rakes aft; that is, hangs towards the stern. How cheer ye

fore and aft? that is, how fares all the ship's company?

Abast the Beam denotes the relative situation of any object with the ship, when the chject is placed in any part of that arch of the horizon, which is contained between a line at right angles with the keel, and that point of the compass which is directly opposite to the ship's course. See Bearing. board. The inside of a ship. Aboard Main Tack! The order to draw the lower corner.

Aboard. The inside of a ship.

of the mainsail down to the chestree.

About. The situation of a ship as soon as she has tacked or changed her course.

About ship! the order to the ship's erew to prepare for tacking.

Abreast. The situation of two or more ships lying with their sides parallel, and their heads equally advanced; in which case they are abreast of each other.

Adrift. The state of a ship broken from her moorings, and driving about without control.

Affect. Buoyed up by the water from the ground.

All that part of a ship which lies forward or near the stem. It also signifies Afore. further forward.

After. A phrase applied to any object in the hinder part of the ship, as the after-hafth-way, the after-sails, &c. Aground. The situation of a ship when her bottom or any part of it rests on the grounds

A-head. Any thing which is situated on that point of the compass to which a ship's stem is directed, is said to be ahead of her. See Bearing.

A-hull. The situation of a ship when all her sails are furled and her belm is lashed to the lee side; by which she lies nearly with her side to the wind and sea, her head being somewhat inclined to the direction of the wind.

A-les. The position of the helm when it is put down to the lee side.

All in the wind. The state of a ship's sails, when they are parallel to the direction of the wind, so as to shake or shiver.

All-hands-ahoy! The call by which all the ship's company are summoned upon decks Alost. Up in the tops, at the mast heads, or any where about the higher rigging.

Along-side. Side by side, or joined to a ship, wharf, &c.
Along-shore. Along the coast; a course which is in sight of the shore, and nearly prerailel to it.

Alsof is distance. Keep aloof, that is, keep at a distance.

Amain. The old term for yield, used by a man of war to an enemy; but it now signi-

fies any thing done suddenly, or at once, by a number of men.

Amidships. The middle of a ship, either with regard to her length or breadth.

Anchor. The instrument by which a ship is held. The anchor is foul; that is, the califf has got about the fluke of the anchor. The machor is a peak; that is, directly under the hawse-hole of the ship. The anchor is a cock bill; that is, hangs up and down the ship's side.

An-end. The position of any mast, &c. when erected perpendicularly on the desk:
The top-masts are said to be an end, when they are hoisted up to their usual station.
A-peak. Perpendicular to the anchor; the cable having been drawn so tight as to bring

the ship directly over it. The anchor is then said to be a-peak.

Ashore. On the shore, as opposed to aboard. It also means aground

Any distance behind a ship, as opposed to ahead. See Bearing,

At Anchor. The situation of a ship riding by her anchor.

Atheort. Across the line of a ship's course. Atheort House; the situation of a ship when driven by accident across the fore part of another; whether they touch, or are at a small distance from each other, the transverse position of the former being principally understood. Atherest the Fore Foot; when any object crosses the line of a ship's course, but ahead of her, it is said to be athwart the Fore Foot. Atherest ships; reaching, or in a direction, across the ship from one side to the other.

When applied to the anchor, it means that the anchor is drawn out of the ground, and hangs in a perpendicular direction, by the cable or buoy rope. The top-sails are said to be atrip, when they are hoisted up to the mast head, or to their utmost extent.

Ausst. A term used for stop, or stay; as, west heaving, do not heave any more.

Avesigh. The same as atrip when applied to the anchor.

Auguing. A shelter or season of canvass spread over the decks of a ship, to keep of the K k

Spread the awning; extend it so as to cover the deck. Furl the the heat of the sun.

moning; that is, roll it up.

To back the Anchor. To carry out a small anchor ahead of the large one, in order to support it in bad ground, and to prevent it from loosening or coming home.

To back astern. In rowing, is to impel the boat with her stern foremost, by means of the

To back the sails. To arrange them in a situation that will occasion the ship to move astern.

To bagpipe the mizen. To lay it aback, by bringing the sheet to the mizen shrouds.

To balance. To contract a sail into a narrower compass, by folding up a part at one corner. Balancing is peculiar only to the mizen of a ship, and the mainsail of those vessels wherein it is extended by a boom.

Bale. Bale the boat; that is, to throw the water out of her.

Ballast is either pigs of iron, stones or gravel, which last is called shingle ballast; and its use is to bring the ship down to her bearings in the water, which her provisions and stores will not do. Trim the ballast; that is, spread it about and lay it even. The Ballast shoots; that is, it shifts, or runs over from one side of the hold to the

Bare poles. When a ship has no sail set, she is under bare poles.

Barge. A carval built boat, that rows with ten or twelve oars.

Batten. A thin piece of wood. Batten down the hatches, is to lay battens upon the tarpaulins, which are over the hatches in bad weather, and nail them down that they may not be washed off.

A post or stake crected over a shoal or sand bank, as a warning to seamen to keep at a distance. Also, a signal placed at the top of hills, &c.

Beams. Strong pieces of timber stretching across a ship's side to side, to support the decks, and retain the sides at their proper distance.

Bear a-hand. Make haste, despatch.

Bearing signifies the point of the compass which any two or more places bear from each other, or how any place bears from the ship by the compass; or it may be said to bear on the beam, abaft the beam, on the bow, the head or stern, &c.

Bearings of a ship, is that line which is formed by the water upon her sides when she is at anchor, with her proportion of ballast and stores on board. To bear to, is to sail into an harbour, &c. Bear round up; that is, put her right before the wind. Bring your guns to bear, is to point them to the object.

To bear in with the land, is when a ship sails towards the shore.

To Bear off. To thrust or keep off from the ship's side, &c. any weight when hoisting.

Bearing up or Bearing away. The act of changing the course of a ship, in order to make her run before the wind, after she had sailed sometime with a side wind, or close hauled; it is generally performed to arrive at some port under the lee, or to avoid some imminent danger occasioned by a violent storm, leak, or enemy in sight.

Beating to Windward. The making a progress against the direction of the wind, by

steering alternately close hauled on the starboard and larboard tacks.

To becalm. To intercept the current of the wind; in its passage to a ship, by any contiguous object, as a shore above her sails, a high sea behind, &c. and thus one sail is said to becalm another.

Before the Beam, denotes an arch of the horizon comprehended between the line of the beam (which is at right angles to the keel) and that point of the compass on which the

ship stems. See Bearing.

To make fast any running rope, as Belay the main brace, or make it fast.

Bend. To apply to and fasten; as, bend the sails, apply them to the yards and fasten them : unbend the sails, that is, cast them off, and take them from the yards; her sails are unbent, she has none fixed : bend the cable, make it fast to the anchor. Beneaped. See Neaped.

Between Decks. The space contained between any two decks of a ship.

Bight of a rope. The double part of a rope when it is folded. Bight, a narrow inlet of the sea.

Bilge. To break. The ship is bilged; that is, her planks are broken in by violence. Bilge-Water, is that which, by reason of the flatness of a ship's bottom, lies on her floor,

and cannot go to the well of the pump. Binnacle. A kind of box to contain the compasses in upon deck.

A place; as the ship's birth, the place where she is moored; an officer's birth, his place in the ship to eat or sleep in; birth the ship's company, that is, allot them their places to mess in; birth the hammocks, point out where each man's hammock is to

Very large pieces of timber in the fore part of a ship, round which the cables are fastened when the ship is at anchor. After-Bitts, a smaller kind of bitts upon the

quarter-deck, for belaying the running rigging to.

To Bitt the Cable, is to confine the Cable to the bitts, by one turn under the cross places.

and another turn round the bitt-head. In this position it may be either kept fixed. of it may be veered away.

The turn of the cable round the bitts. Bitter end; that part of the cable which stays within board, round about the bitts, when the ship is at anchor.

Block. A piece of wood with running sheaves or wheels in it, through which the running rigging is passed to add to the purchase.

To make a board is making a stretch upon any tack when a ship is working upon a wind. To board it up, that is, to turn to windward. The ship has made a stern board, that is, when she loses ground in working upon a wind.

Bostrosia. The officer who has charge of all the cordage, rigging, anchors, &c.

Bold shore. A steep coast, permitting the close approach of shipping.

Bolt rope. The rope which goes round a sail, and to which the canvass is sewed. side ropes are called leach ropes, that at the top the head rope, and that at the bottom the foot rope.

Bonnet of a sail, is an additional piece of canvass put to the sail in moderate weather to hold more wind. Lase on the bonnet, that is, fasten it to the sail. Shake off the bonnet,

take it off.

oot topping. Cleaning the upper part of a ship's bottom, or that part which lies immediately under the surface of the water, and daubing it over with tallow, or with a mix-Boot topping. ture of tallow, sulphur, rosin, &c.

Both sheets of: The situation of a ship sailing right before the wind.

Bow grace. A frame of old rope or junk, laid out at the bows, stems, and side of ships,

to prevent them from being injured by flakes of ice.

Bose lines. Lines made fast to the sides of the sails to has! them forward when upon a

wind, which being hauled taut, enables the ship to come nearer to the wind.

To besse. To pull upon any body with a tackle in order to remove it.

Bouspril. A large mast or piece of timber which stands out from the bows of a ship.

Boxhauling. A particular method of veering a ship, when the swell of the sea renders

tacking impracticable.

Boxing. An operation somewhat similar to Boxhauling. It is performed by laying the head sails aback, to receive the greatest force of the wind in a line perpendicular to their surfaces, in order to turn the ship's head into the line of her course, after she had inclined to windward of it.
races. The ropes by which the yards are turned about to form the sails to the wind.

To brace the yards. To move the yards, by means of the braces, to any direction required.

To brace about—to brace the yards round for the centrary tack.

To brace sharp—to To brace sharp—to brace the yards to a position in which they will make the smallest possible angle with the keel, for the ship to have head-way. To Brace to—to case off the lee braces, and round in the weather braces, to assist the motion of the ship's head in tacking.

A name peculiar only to certain ropes belonging to the mizen, used to truss it But it is likewise applied to all the ropes which are employed in hashing up the bottoms, lower corners and skirts of the other great sails. To broth up—to haul up a sail by means of the brails, for the more ready furling it when

To break bulk. The act of beginning to unload a ship.

To break sheer. When a ship at anchor is forced, by the wind or current, from that position in which she keeps her anchor most free of herself and most firm in the ground, so as to endanger the tripping of her anchor, she is said to break her sheer.

Breaming. Burning off the filth from a ship's bottom.

Breast-fast. A rope employed to confine a ship sideways to a wharf or to some other

To bring by the lee.—See to broach to.

To bring to. To check the course of a ship when she is advancing, by arranging the sails in such a manner that they shall counteract each other, and prevent her from either retreating or advancing. See to lie to.

To broach to. To incline suddenly to windward of the ship's course, so as to present her side to the wind, and endanger her oversetting. The difference between broaching to and bringing by the lee may be thus defined: Suppose a ship under great sail is steering south, having the wind at N. N. W. then west is the weather-side, and east the least If, by any accident, her head turns round to the westward, so that her sails are all taken a-back on the weather side, she is said to broach to. If, on the contrary, her head declines so far castward as to lay her sails aback on that side which was the lee side, it is called bringing by the lee.

Broadside. A discharge of all the guns on one side of a ship both above and below.

Broken backed. The state of a ship which is so loosened in her frame as to drop at each end.

By the board. Over the ship's side.

By the head. The state of a ship when she is so unequally loaded as to draw more wa-Digitized by GOOGIC ter forward than aft.

By the trind. The course of a ship as near as possible to the dimetion of the wind, which is generally within six points of it.

Anni-lines. Ropes fastened to the foot rope of square sails to draw them up to the middle of the yards for furling.

Buoy. A floating conical cask, moored upon shoals to show where the danger is; also used at anchors to show where they lie, in case the cable breaks.

Cap. A strong thick block of wood, having two large holes through it, the one square, the other round; used to confine the two masts together.

Copsize. Overturn—the boat is capsized, that is, overset. Capsize the quoil of rope; that is, turn it over.

Capstan. An instrument by which the anchor is weighed out of the ground, used also

for setting up the shrouds, and other work where a great purchase is required.

Corsen. To incline a ship on one side so low down, by shifting the cargo or stores on one side, that her bottom on the other side may be cleaned by breaming.

To carry away. To break—as a ship has carried away her bowsprit, that is, has broken

Casting. The motion of falking off, so as to bring the direction of the wind on either side of the ship after it had blown sometime right a-head. It is particularly applied to a ship about to weigh anchor.

Cit-Heads. The timbers on ship's bows with sheaves in them, by which the anchor is hoisted after it has been hove up by the cable.

To cat the anchor is to hook the cat-block to the ring of the anchor, and haul it up class to the cat-head.

Cat's paw is a light air of wind perceived at a distance in a calm, sweeping the surface of the sea very lightly, and dying away before it reaches the ship.

Caulking is filling the seams of a ship with cakum.

This word is applied to that squadron of a fleet, in a line of battle, which orcupies the middle of a line; and to that column (in the order of sailing) which is between the weather and lee columns.

Chains. A place built on the sides of the ship projecting out, and at which the shrouds are fastened, for the purpose of giving them greater angle than they could have if fastened to the ship's side, and of course giving them a greater power to secure the

Chain plates, are plates of iron fastened to the ship's sides under the chains, and to these

plates the dead eyes are fastened.

Chepelling. The act of turning a ship round in a light breeze of wind when she is close. hauled, so that she will lie the same way she did before. This is usually occasioned by negligence in steering, or by a sudden change of wind.

Chase. A vessel pursued by some other. Chaser-the vessel pursuing.

A phrase implying heartily, quickly, cheerfully. Cheerly.

To class off. The act of turning to windward from a lee shore, to escape shipwreck, &c. Clear is variously applied:—The weather is said to be clear when it is fair and open, the sea-coast is clear when the navigation is not interrupted by rocks, &c. It is applied to cordage, cables, &c. when they are disentangled, so as to be ready for immediate service. In all these senses it is opposed to foul. To clear the enchor is to get the cable off the flukes, and to disencumber it of ropes, ready for dropping. Clear house-When the cables are directed to their anchors without lying athwart the stam. To clear the house is to untwist the cables when they are entangled by heaving either a cross, an cibow or a round turn.

Clenched. Made fast, as the cable is to the ring of the anchor.

Close handed. That trim of the ship's sails, when she endeavours to make a progress in the nearest direction possible towards that point of the compass from which the wind

To club hand. A method of tacking a ship when it is expected she will miss stays on a ice shore.

Clue-lines, are ropes which come down from the mast to the lower corners of the sails, and by which the corners or clues of the sails are hauled up.

Clue of a sail. The lower corners of square sails, but the aftermost only of stay sails, the other lower corner being called the tack.

To clue up. To haul up the clues of a sail to its yard by means of the clue lines.

Goasting. The act of making a progress along the sea-ceast of any country.

To coil the cable. To lay it round in a ring, one turn over mother.

To come home. The anchor is said to come home when it loosens from the ground by the effort of the cable, and approaches the place where the ship floated, at the length of her moorings.

her moorings.

Contag is, denotes the approach of a ship's head to the direction of the wind.

Courses. The point of the compass upon which the ship sails. Courses, a ship's lower sails; as the fore-sail is the fore-course; the main-sail the main-course, &c.

The this is under her courses; that is, has no sail set but the main-sail, fore-sail and mizen.

Correcto. The person who steers the boat.

Crank. The ship is crank, that, is she has not a sufficient cargo or ballast to render her capable of bearing sail, without being exposed to the danger of eversetting.

Crow-foot, is a number of small lines spread from the fore parts of the tops, by means of the piece of wood through which they pass, and being hauled taut upon the stays, they prevent the foot of the topsails catching under the top rim; they are also used to suspend the awnings.

Cun. To direct. To can a ship is to direct the man at helm how to steer.

To cut and run. To cut the cable and make sail instantly, without waiting to weigh anchor.

Devit. A long beam of timber, used as a crane, whereby to hoist the flukes of the anchor to the top of the bow, without injuring the planks of the ship's sides as it ascends. There is always a Davit of a smaller kind fixed to the long-boat to weigh the anchor by the buoy-rope.

To deaden a ship's usey. To impede her progress through the water.

Dead eyes. Blocks of wood through which the laniards of the shrouds are reeved.

Dead lights. A kind of window shutter for the windows in the stern of a ship, used in very bad weather only.

Dead tester. The eddy of water, which appears like whirlpools, closing in with the ship's stern as she sails on.

Dead-wind. The wind right against the ship, or blowing from the very point to which she wants to go.

Dismasted. The state of a ship that has lost her masts.

Dog-same. A small vane with feathers and cork, and placed on the ship's quarter, for the men at cun and helm to see the course of wind by.

Dog-watch. The watches from four to six and from six to eight in the evening.

Doubling. The act of sailing round or passing beyond a cape or point of land.—Doubling upon—The act of inclosing any part of a hostile fleet between two fires, or of cannotading it on both sides.

Douce. To strike or haul down; as, douce the top-gallant sails, that is, lower them.

Down houl. The rope by which any sail is hauled down, as the jibb down-haul. To lower suddenly or slacken.

To drug the anchor. To trail it along the bottom after it is loosened from the ground. To draw. When a sail is inflated by the wind, so as to advance the vessel in her course,

the sail is said to draw; and so to keep all drawing is to inflate all the sails. Drift. The angle which the line of a ship's motion makes with the nearest meridian, when she drives with her side to the wind and waves, and not governed by the power of the helm. It also implies the distance which the ship drives on that line.

Driver. A large sail set upon the mizen yards in light winds. Drive. - The ship drives,

that is, her anchor comes through the ground.

Brop. Used sometimes to denote the depth of a sail; as the fore-top-sail drops twelve yards. To drop anchor. Used synonimously with to anchor. To drop astern. The retrograde

motion of a ship. Dunnage. A quantity of loose wood, &c. laid at the bottom of a ship to keep the goods from being damaged.

Earings. Small ropes used to fasten the upper corners of sails to the yards.

To ease, to ease away, or to ease off. To slacken gradually ; thus they say, ease the bowline, ease the sheet.

Ease the skip. The command given by the pilot to the steersman, to put the helm harda-lee, when the ship is expected to plunge her fore part deep in the water when close hauled.

To edge enosy. To decline gradually from the shore or from the line of the course which the ship formerly held in order to go more large.

To edge in with. To advance gradually towards the shore, or any other object.

Elbers in the Hause. Is when a ship, being moored, has gone round upon the shifting of the tides twice the wrong way, so as to lay the cables one over the other: having gone once wrong, she makes a cross in the hawse, and going three times wrong, she makes a round turn.

Endfor end. A term used when a rope runs all out of a block, and is unreeved, or in coming to an anchor, if the stoppers are not well put on and the cable runs all out, it is said to have gone out end for end.

End . When a ship advances to a shore, rock, &c. without an apparent possibility of preventing her, she is said to go end on for the shore, &c.

Engagement. Action or fight.

Ensign. The flag worn at the stern of a ship.

Entering port. A large port in the side of three deckers leading into the middle deck, to save the trouble of going up the ship's side to get on hoard. Digitized by

Epen keel. When the keel is parallel with the horizon, a ship is said to be upon an then

Fair. A general term for the disposition of the wind, when favourable to a ship's course. Fair-way. The channel of a narrow bay, river, or haven, in which ships usually advance in their passage up and down.

Fack, or Fake. One circle of any rope or cable quoiled.

Fag end. The end of any rope which is become untwisted by Trequent use ; to prevent which the ends of ropes are wound round with pieces of twine, which operation is

called whipping.

To fall a-board of. To strike or encounter another ship when one or both are in motion.

To fall astern—The motion of a ship with her stern foremost. To fall calm. To become in a state of rest by a total cessation of the wind. To fall down. To sail or be towed down a river nearer towards its mouth.

Falling off denotes the motion of the ship's head from the direction of the wind. It is

used in opposition to coming to.

Fall not off, or nothing off. The command of the steersman to keep the ship near the wind.

Fathom. A measure of six feet.

To fetch away. To be shaken or agitated from one side to another so as to loosen any thing which before was fixed.

Fid. A square bar of wood or iron, with shoulders at one end, used to support the weight of the topmast, when erected at the head of a lower-mast. Fid for splicing. A large piece of wood of a conical figure; used to extend the strands

and layers of cables in splicing.

To fill. To brace the sails so as to receive the wind in them, and advance the ship in her course, after they had been either shivering or braced a-back.

Fish. A large piece of wood. Fish the mast; apply a large piece of wood to it to strengthen it.

Fish-hook. A large hook by which the anchor is received and brought to the cat-head; and the tackle which is used for this purpose is called the fish-tackle.

To fish the anchor. To draw up the flukes of the anchor towards the top of the bow, in order to stow it, after having been catted.

Flag. A general name for colours worn and used by ships of war.

Flat aft. The situation of the sails when their surfaces are pressed aft against the most

by the force of the wind.

o flat in. To draw in the aftermost lower corner, or clue, of a sail towards the middle of the ship, to give the sail a greater power to turn the vessel. To flat in forward. To To flat in. draw in the fore sheet, jibb-sheet, and fore-stayasil sheet, towards the middle of the ship. Flaw. A sudden breeze or gust of wind.

Floating. The state of being buoyed up by the water from the ground.

The state of a tide when it flows or rises.

Flowing sheets. The position of the sheets of the principal sails when they are loosened from the wind so as to receive it into their cavities more nearly perpendicular than when close-hauled, but more obliquely than when the ship sails before the wind. going two or three points large has flowing sheets.

Fore. That part of a ship's frame and machinery that lies near the stem.

Fore end of.

Throughout the whole ship's length. Lengthways of the ship.

Fore-reach. To shoot a-head, or go past another vessel.

To force over. To force a ship violently over a shoul-by a great quantity of sail.

Forward. Towards the fore part of a ship.

Foul. Is used in opposition both to clear and fair. As opposed to clear we say, foul weather, foul bottom, foul ground, foul anchor, foul hawse. As opposed to fair, we say foul wind.

To Founder. To sink at sea by filling with water.

To free. Pumping is said to free the ship when it discharges more water than leaks

into her.

To freshen. When a gale increases it is said to freshen. To freshen the Hause. Veering out or heaving in a little cable to let another part of it endure the stress of the hawscholes. It is also applied to the act of renewing the service round the cable at the hawscholes.

Freshen the ballast. Divide or separate it.

Fresh away. When a ship increases her velocity, she is said to get fresh way.

Full. The situation of the sails, when they are kept distended by the wind.

Full and by. The situation of a ship, with regard to the wind, when close hauled and

sailing, so as to steer neither too nigh the direction, nor to deviate to leeward.

To furl. To wrap or to roll a sail close up to the yard or stay to which it belongs, and winding a cord around it, to keep it fast.

Gage of the ship. Her depth of water, or what water she draws.

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To rain the wind. To arrive on the weather side, or to windward of some ship or feet in sight, when both are sailing as near the wind as possible.

Gammon the binosprit. Secure it by turns of a strong rope passed round it, and into the cat-water, to prevent it from having too much motion.

Gengrosy. That part of a ship's side both within and without, by which persons enter

and depart.

Garboard streak. The first range or streak of planks laid in a ship's bottom next the keel. Gasket. The rope which is passed round the sail to bind it to the yard when it is furled.

To gather. A ship is said to gather on another as she comes nearer to her.

Gimbleting. The action of turning the anchor round by the stock, so that the motion of

the stock appears similar to that of the handle of a gimblet, when employed to turn the wire.

Girt. The ship is girt with her cables when she is too tight moored.

To give chase to. To pursue a ship or fleet.

Goove-wings of a sail. The clues or lower corners of a ship's mainsail or foresail, when the middle part is furled or tied up to the yard.

Grappling-iron. A thing in the nature of an anchor, with four or six flukes to it. Greave. To burn off the fifth from a ship's bottom.

Gripe of a ship. That thin part of her which is under the counter; and to which the stern post joins. The ship gripes, that is, turns her head too much to the wind.

Grounding. The laying the ship a-shore, in order to repair her. It is also applied to running aground accidentally.

Ground tackle. Every thing belonging to a ship's anchors, and which are necessary for anchoring or mooring; such as cables, hawsers, tow-lines, warps, buoy-ropes, &c.

Ground tier. That is, the tier of water casks which is lowest in the hold, and is among the shingle ballest.

Growing. Stretching out; applied to the direction of the cable from the ship towards the anchors: as, the cable grows on the starboard bow.

Grommet. A piece of rope laid into a circular form, and used for large boat's oars instead of rowlocks, and also for many other purposes.

Gunnel. The upper edge of a ship's side.

Gun-room. A division of the lower deck abaft, enclosed with network, for the use of the gunner and his stores.

Gybing. The act of shifting any boom-sail from one side of the mast to the other. Hall. To call to another ship.

Hallards. The ropes by which the sails are hoisted, as the top-sail hallards, or jibb haliards, &c.

Handing. The same as furling.

Hard a-weather. Put the tiller quite up to windward.

Haul. Pull.

To haul the wind. To direct the ship's course nearer to the point from which the wind blows.

Hause-holes. The holes in the bows of the ship through which the cables pass .- Freshen house, veer out more cable. Clap a service in the house; put somewhat round the cable at the hawse-hole to prevent its chafing. To clear house, is to untwist the cables where a ship is moored, and has got a foul hawse. Athwart hause, is to be across or before another ship's head.

Howser. A small kind of cable.

Houd fust. A rope employed to confine the head of a ship to a wharf or to some other ship. Headmost. The situation of any ship or ships which are the most advanced in a fleet. Head sails. All the sails which belong to the foremast and bowsprit.

Ilead-sea. When the waves meet the head of a ship in her course, they are called a head-sea. It is likewise applied to a single wave coming in that direction.

Head to-wind. The situation of a ship when her head is turned to the point from which the wind blows, as it must when tacking.

Head-way. The motion of advancing, used in opposition to stern-way.

To Heave. To turn about a capstan, or other machine of the like kind, by means of bars, handspikes &c. To Heave a-head, to advance the ship by heaving-in the cable or other rope fastened to an anchor at some distance before her. To Heave a-peck, to heave-in the cable, till the anchor is a-peck.

To Heave a-stern, to move a ship backwards by an operation similar to that of heaving a-head. To Heave down, to careen. To Heave in the cable, to draw the cable into the ship, by turning the capstan. To Heave in stays, to bring a ship's head to the wind, by a management of the sails and rudder, in order to get on the other tack. To Hesse out, to unfurl or loose a sail; more particularly applied to the stay-sails; thus we say, loose the topsails and heave out the staysails. To Hesse short, to draw so much of the cable into the ship, as that she will be almost perpendicularly over her anchor. To Heave tight or tout, to turn the capstan round till the rope or cable becomes straightened. To Heave the capstan, to turn it round. To Heave the lead, to throw the lead overboard, in order to find the depth of water. To Heure lie. log, to throw the log overboard, in order to calculate the 'velocity of the ship's way. Heave the capston, that is, turn it round with the bars. Heave handsomely, heave gently or leisurely. Heave hearty, heave strong and quick.

Heave of the sea, is the power that the swell of the sea has upon a ship in driving her out, or faster on, in her course, and for which allowance is made in the day's work.

Heel, or incline. She heels to port, that is, inclines or lays down upon her larboard or left side.

Helm. The instrument by which the ship is steered, and includes both the wheel and the tiller, as one general term. Helm's a-lee, that is, the tiller is quite down to leeward The situation of a ship when so far run aground as to be seen dry upon

Hitch. To make fast.

Hoist. To haul, sway, or lift up.

Hold, is the space between the lower deck and the bottom of the ship, where her stores,

To stow the hold, is to place the things in it.

To hold its own, is applied to the relative situation of two ships when neither advances. upon the other; each is then said to hold its own. It is likewise said of a ship which, by means of contrary winds, cannot make a progress towards her destined port, but which, however, keeps nearly the distance she had already run.

Home, implies the proper situation of any object; as, to had home the topsail sheets, is to extend the bottom of the topsail to the lower yard, by means of the sheets. In stowing a hold, a cask, &c. is said to be kome, when it lies close to some other

Hulk. A ship without masts or rigging; also a vessel to remove masts into or out of ships by means of sheers, from whence they are called sheer hulks.

Horse. A rope reaching from the middle of a yard to its arms or extremities, for the men to stand on when they are loosing, reefing, or furling a sail.

Hull of the ship, the body of it. To lay a hull, is to lay to with only a small sail in a gale

of wind. To hall a vessel, is to fire a shot into any part of her hull.

Hull down, is when a ship is so far off that you can only see her masts. To hall a shipto fire cannon balls into her hull within the point-blank range. Hull to-the situation of a ship when she lies with all her sails furled : as in trying.

In stays. See to heave in stays.

The act of enclosing any object between two bodies so as to render it im-Jamming. moveable.

Jeer blocks. The blocks through which jeers are reeved.

Jects. The ropes by which the lower yards are suspended.

Jibb. The foremost sail of a ship, set upon a boom which runs out upon the bowsprit. Jibb-boom. A spar that runs out upon the bowsprit. Jolly-boot. A small boat.

Junk. Old cable, or old rope.

Jury-mast. A temporary or occasional mast erected in a ship in the place of one which has been carried away by accident, &c.

Redge. A small anchor with an iron stock.

Keel. The principal piece of timber in a ship, which is usually first laid on the blocks in building.

Kee!-haul. To drag a person backwards and forwards under a ship's keel for certain

Any part of a cable, covered over with old ropes, to prevent its surface from

rubbing against the ship's bow or fore-foot.

To keep many. To alter the ship's course to one rather more large, for a little time, to avoid some ship, danger, &c. Keep sway is likewise said to the steersman who is ant to go to windward of the ship's course. To keep full—To keep the sails distended by the To Keep hold of the land—to steer near to or in sight of the land. To Keep of wind. to sail off or keep at a distance from the shore. To Keep the land abourd—the same as to keep hold of the land. To Keep the Luff-to continue close to the wind. To keep the wind—the same as to keep the luff.

Kelson. A piece of timber forming the interior of the keel; being laid on the middle of the floor timbers immediately over the keel, and serving to unite the former to the

latter.

Kentledge. Pigs of iron for ballast, laid upon the floor, near the kelson, fore and aft.

Kenk. A sort of twist or turn in a cable or rope.

Knippers. A large kind of plated rope, which being twisted round the messenger and cable in weighing, bind them together.

Knot. A division of the log-line, answering in the calculation of the ship's velocity, to one mile.

To Labour. To roll or pitch heavily in a turbulent sea. Digitized by GOOGLE

Luden in Bulk. Freighted with a cargo not packed, but lying toose, as corn, salt, &c.

Laid up. The situation of a ship when moored in a harbour, for want of employ.

Land fall. The first land discovered after a sea voyage. Thus a good land fall implies the land expected or desired; a bad land fall the reverse.

Land-locked. The situation of a ship surrounded with land, so as to exclude the prospect of the sea, unless over some intervening land.

Laniards of the shrouds, are the small ropes at the ends of them, by which they are hove

tant or tight.

Larboard. The left side of a ship, looking towards the head. Larboard Tack—the situation of a ship when sailing with the wind blowing upon her larboard side.

Launch-ho, signifies high enough, or lower.

Laying the land. A ship which increases her distance from the coast, as to make it appear lower and smaller, is said to lay the land.

ing-wind. A fair wind for a ship's course.

A chink or breach in the sides or bottom of a ship, through which the water emters into the hull.

Lee. That part of the hemisphere to which the wind is directed, to distinguish it from the other part which is called to windward. Lee Gage-a ship or fleet to leeward of another is said to have the lee-gage. Lee-Lurches—the sudden and violent rolls which a ship often takes to leeward, in a high sea, particularly when a large wave strikes her on the weather-side. Lee-Quarter—that quarter of a ship which is on the lee side. Lee-Shore—that shore upon which the wind blows. Lee-side—that half of a ship lengthwise, which lies between a line drawn through the middle of her length and the side which is farthest from the point of wind. To Leeward—towards that part of the horizon to which the wind blows. Leeward-Ship—a ship that falls much to lecward of her course, when sailing close hauled. Lectoard Tide-a tide that sets to lee-

The lateral movement of a ship to leeward of her course; or the angle which the line of her way makes with a line in the direction of her keel.

To Lie dong. To be pressed down sideways by a weight of sail in a fresh wind.

Leeches. The borders or edges of a sail.

To Lie to. To retard a ship in her course, by arranging the sails in such a manner as to counteract each other with nearly an equal effort, and render the ship almost itnmoveable with respect to her progressive motion or head way.

Lifts. The ropes which come to the ends of the yards from the mast-heads, and by which

they are suspended when lowered down.

Limbers, or Limber-holes. Square holes cut through the lower part of a ship's floor timbers, very near the keel; forming a channel for water, and communicating with the pump-well throughout the whole length of the floor.

List. Incline. The ship has a list to port, that is, she heels to the larboard.

Log, and Log line, by which the ship's path is measured, and her rate of going ascer-

Log-board, on which is marked the transactions of the ship, and from thence it is copied into the log-book every 24 hours.

A Long Sea. A uniform motion of long waves.

Look out. A watchful attention to some important object or event that is expected fo Thus persons on board of a ship are occasionally stationed to look out for sigarise. nals, other ships, for land, &c.

To Loose. To unfuri or cast loose any sail.

To Lower. To ease down gradually.

Laf. The order to the steersman to put the helm towards the lee side of the ship in

order to sail nearer to the wind.

To make a board. To run a certain distance upon one tack, in beating to windward. To make foul water—to muddy the water, by running in shallow places, so that the ship's keel disturbs the mud at the bottom. To make sail—to increase the quantity of sail already set, either by unrecfing or by setting others. To make stern way—to retreat or move with the stern foremost. To make the land-To discover it from afar. make water-to leak.

To man the yard, &c. To place men on the yard, in the tops, down the ladder, &c. to execute any necessary duties.

Mast. The upright timber or trees on which the yards and sails are set.

Masted. Having all her masts complete. Mend the service. Put on more service.

Messenger. A small kind of cable, which being brought to the capstan, and the cable by which the ship rides made fast to it, it purchases the anchor.

To middle grope. To double it into two equal parts.

Midshipp. See Amidships.

To miss stays. A ship is said to miss stays when her head will not fly up into the direc-

tion of the wind, in order to get her on the other tack.

Mizen mast. The mast which stands abast, and from which its rigging and sails are named; as of the sails, mizen, mizen-top-sail, &c. and so also are the other sails, &c. named from the other masts.

Moor, is to secure a ship with two anchors. Mooring-securing a ship in a particular station by chains or cables, which are either fastened to an adjacent shore or to anchors at the bottom. Mooring service—when a ship is moored, and rides at one cable's length, the mooring service is that which is at the first splice.

Mouse. A kind of ball or knob, wrought upon the collar of the stays.

Muster. To assemble.

A small passage between two lands. Narrows.

Neape Tides. The tides in the first and last quarter of the moon, and are not either so high, so low, or so rapid as spring tides. A ship is said to be beneaped when she has not water enough to take her off the ground, or over the bar, &c.

Near, or nonear. An order to the steersman not to keep the ship so close to the wind. Nippers. Certain pieces of cordage used to fasten the messenger to the cable in

heaving up the anchor.

Nothing off. A term used by the man at the cun to the steersman, directing him not to go from the wind.

Nun buoy. The kind of buoys used by ships of war.

Oslam. Old rope untwisted and pulled open.

Off and on. When a ship is beating to windward, so that by one board she approaches towards the shore, and by the other stands out to sea, she is said to stand off and en shore.

To seaward from the land. A ship is in the offing, that is, she is to seaward, at Offing. a distance from the land. She stands for the offing, that is, towards the sea.

Offward. From the shore, as when a ship lies aground and leans towards the sea, she is said to heel offward.

On board. Within the ship, as, he is come on board.

On the beam. Any distance from the ship on a line with the beams, or at right angles with the keel. See Bearing.

On the bow. An arch of the horizon, comprehending about four points of the compass on each side of that point to which the ship's head is directed. Thus, they say, the ship in sight bears three points on the starboard bow; that is, three points towards

the right hand, from that part of the horizon which is right a-head. See Bearing.

n the quarter. An arch of the horizon, comprehending about four points of the compass on each side of that point to which the ship's stern is directed. See On the boso. On the quarter. Open. The situation of a place exposed to the wind and sea. It is also expressed of

any distant object to which the sight or passage is not intercepted. Open Hawse. When the cables of a ship at her moorings lead straight to their respect-

ive anchors, without crossing, she is said to ride with an open house. Orlop. The deck on which the cables are stowed.

Over-board. Out of the ship; as, he fell over-board, meaning he fell out of, or from the ship.

Overgrown Sea, is expressed of the ocean when the surges and billows rise extremely high.

Overhaul. To clear away and disentangle any rope; also to come up with the chase;

as we overhaul her, that is, we gain ground of her.

Oper-Rake. When a ship at anchor is exposed to a head-sea, the waves of which break in upon her, the waves are said to over-rake her.

Over-set. A ship is over-set when her keel turns upwards.

Out-of-trim. The state of the ship when she is not properly balanced for the purposes

of navigation.

Parcel a rope. Is to put a quantity of old canvass upon it before the service is put on-Parcel a seam-is to lay a narrow piece of canvass over it after it is caulked, before it it is payed.

Parliament heel. The situation of a ship when she is made to stoop a little to one side, so as to clean the upper part of her bottom on the other side. See Boot-topping.

Parting. Being driven from the anchors, by the breaking of the cable.

Paul. A short har of wood or iron fixed close to the apetan or windlass of a ship, to prevent those engines from rolling back, or giving way, when they are charged with any great effort.

To Penol the capstan. To fix the pawls so as to prevent the capstan from recoiling,

during any pause of heaving.

pay. To daub or cover the surface of any body with pitch, tar, &c. in order to prevent it from the injuries of the weather.

To pay away or pay out. To slacken a cable or other rope, so as to let it run out for some particular purpose.

To pay off. To move a ship's head to leeward.

To peek the mizen. To put up the mizen yard perpendicular by the mast.

Peek. To ride a stay peck, is when the cable and the fore-stay form a line. a short peck, is when the cable is so much in as to destroy the line formed by the stay-peck. To rids with the yards a peck, is to have them topped up by contrary lifts, so as to represent St. Andrew's cross.

Pendant. The long narrow flag worn at the mast-head by all ships of the navy. Pendants—are those ropes which secure the brace-blocks to the yard arms, and are in general double, in case of one being shot away, the other may secure the yards in its

proper position-

Broad Pendant. A kind of flag terminating in a point used to distinguish the chief of

a squadron.

Pitching. The movement of a ship, by which she plunges her head and after part alternately into the hollow of the sea.

Point-blank. The direction of a gun when levelled horizontally.

Points. A number of plated ropes made fast to the sails for the purpose of reefing.

Poop. The highest and aftermost deck of a ship.

Pooping. The shock of a high and heavy sea upon the stern or quarter of a ship, when she scuds before the wind in a tempest.

Port. A name given on some occasions to the larboard side of the ship; as, the ship heels to port, top the yards to port, &c.-also, a harbour or haven.

The holes in the ship's sides from which the guns are fired.

Port the helm! The order to put the helm over to the larboard side.

Port-last. The gunnel.

Press of Sail. All the sail that a ship can set or carry.

Preventer. An additional rope employed at times to support any other, when the latter suffers an unusual strain, particularly when blowing fresh, or in a gale of wind.

Pudding and Dolphin. A large and lesser pad made of ropes, and put round the mast under the lower yards.

Purchase. Any sort of mechanical power employed in raising or moving heavy bodies. uarters. The respective stations of the officers and people in time of action. Quartering, distributing the men into different places. Quarter-bill, the list of the ship's company, with their stations for action noticed.

Quarter-wind, is when the wind blows in from that part of the horizon situated on the quarter of the ship. See On the quarter.

Quoil, is a rope or cable laid up round, one fake over another.

To Raise. To elevate any distant object at sea by approaching it; thus, to raise the land is used in opposition to lay the land.

To Rake. To cannonade a ship at the stern or head, so that the balls scour the whole length of the decks.

Range of Cable. A sufficient length of cable drawn upon deck before the anchor is cast loose, to admit of its sinking to the bottom without any check.

Radimes. The amail ropes fastened to the snrougs, by which the mean for Reach. The distance between any two points on the banks of a river, wherein the

Ready about! A command of the boatswain to the crew, and implies that all the hands are to be attentive and at their stations for tacking.

Rear. The last division of a squadron, or the last squadron of a ficet. It is applied likewise to the last ship of a line, squadron, or division.

Reef. Part of a sail from one row of eyelet-holes to another. It is applied likewise to a chain of rocks lying near the surface of the water.

Reefing. The operation of reducing a sail by taking in one or more of the reefs.

To pass the end of a rope through any hole, as the channel of a block, the To Reeve. cavity of a thimble, &c.

Rendering. The giving way or yielding to the efforts of some mechanical power. It is used in opposition to jambing or sticking.

Ribs of a skip. A figurative expression for the timbers.

Ride at anchor, is when a ship is held by her anchors, and is not driven by wind or tide. To ride athwart, is to ride with the ship's side to the tide. To ride house fallen, is when the water breaks into the hawse in a rough sea.

Rigging. A general name given to all the ropes employed to support the masts, to ex-

tend or reduce the sails, or to arrange them to the disposition of the wind. Righting. Restoring a ship to an upright position, either after she has been laid on a careen, or after she has been pressed down on her side by the wind.

To right the helm, is to bring it into midships, after it has been pushed either to starboard or larboard.

Rigging out a boom. The running out a pole at the end of a yard to extend the foot of a sail. Digitized by GOO

To rig the capstern. To fix the bars in their respective holes.

Road. A place near the land where ships may anchor, but which is not sheltered. Robands, or Ropebands. Short flat pieces of plaited rope, having an eye worked at one end; they are used in pairs to tie the upper edges of the square sails to their respective yards. Rolling.

The motion by which a ship rocks from side to side like a cradle.

Rough-Tree. A name applied to any mast, yard, or boom, placed in merchant ships, as a rail or fence above the vessel's side, from the quarter-deck to the fore-castle.

Rounding-in. The pulling upon any rope which passes through one or more blocks in a direction nearly horizontal; as, round in the weather-braces.

Rounding. Old ropes fastened on the cable, near the anchor, to keep it from chaing. Round-turn. The situation of the two cables of a ship when moored, after they have been several times crossed by the swinging of the ship.

Rounding-up. Similar to rounding-in, except that is applied to ropes and blocks which

act in a perpendicular direction.

O Row. To move a boat with oars. To Row.

Rossing. Pulling up a cable or rope without the assi tance of tackles.

The machine by which the ship is steered.

Ryllock. The niche in a boat's side, in which the oars are used.

Run. The aftermost part of a ship's bottom, where it grows extremely narrow as the stern approaches the stern-post. Run is also the distance sailed by a ship; and is likewise used by sailors to imply the agreement to work a single passage from one place to another.

To run out a warp. To carry the end of a rope out from a ship, in a boat, and fasten if to some distant object, so that by it the ship may be removed by pulling on it.

To sag to lecward. To make considerable lee-way.

Sailing-trim is expressed of a ship when in the best state for sailing.

She sands or sends. When the ship's head or stern falls deep in the trough of the sea. Scorting. The variation of the wind, by which it becomes unfavourable to a ship's making great progress, as it deviates from being large, and obliges the vessel to steer close hauled, or nearly so.

Scud. To go right before the wind; and going in this direction without any sail set

is called spooning.

Scattling. Cutting large holes through the bottom or sides of a ship, either to sink her,

or to unlade her expeditiously when stranded.

A large wave is so called. Thus, they say a heavy sea. It implies likewise the agitation of the ocean, as, a great sea. It expresses the direction of the waves, as, a head sea. A long sea means a uniform and steady motion of long and extensive waves; a short sea, on the contrary, is when they run irregularly, broken, and interrupted.

Sea-boat. A vessel that bears the sea firmly, without straining her masts, &c.

Sta-clothes. Jackets, trowsers, &c.

Sea-mark. A point or object on shore conspicuously seen at sea.

Sea-room. A sufficient distance from the coast or any dangerous rocks, &c. so that a ship may perform all nautical operations without danger of shipwreck.

To bind or make fast.

To wind something about a rope to prevent it from chafing or fretting. The service is the thing so wound about the rope.

Setting. The act of observing the situation of any distant object by the compass.

To set sail. To unfurl and expand the sails to the wind, in order to give motion to the

ship.

To set up. To increase the tension of the shrouds, back-stays, &c. by tackles, laniards, &c.
Settle. To lower; as, settle the topsail haliards, lower them.
To settle the land. To lower in appearance. It is synonymous with to laythe land.

Shank. The beam or shaft of an anchor.

Shank-painter. The rope by which the shank of the anchor is held up to the ship's side; is also made fast to a piece of iron chain, in which the shank of the anchor lodges. To shape a course. To direct or appoint the track of a ship, in order to prosecute a

voyage.

The sheer of the ship is the curve that is between the head and the stern upon Sheer.

her side. The ship sheers about, that is, she goes in and out.

To sheer off. To remove to a greater distance.

Sheers, are spars lashed together and raised up for the purpose of getting out or in a mast. Sheet. A rope fast ned to one or both of the lower corners of a sail, in order to extend and retain it in a particular situation. When a ship sails with a side wind, the lower corner of the main and fore-sails are fastened by a tack and a sheet, the former being to windward, and the latter to leeward; the tack is, however, only disused with a stern wind, whereas the sail is never spread without the assistance of one or both of the sheets; the stay-sails and studden-sails have only one tack and one sheet each, the stay-sail tacks are fastened forward, and the sheets drawn aft, but the

studden-sail tacks draw the outer corner of the sail to the extremity of the boom. while the sheet is employed to extend the inner corner.

To sheet home. To haul the sheets of a sail home to the block on the yard-arm. To shift the helm. To alter its position from right to left, or from left to right.

To ship. To take any person, goods, or thing on board. It also implies to fix any thing in its proper place, as, to ship the oars, to fix them in their rullocks.

Ship-shape. In a seaman like manner; as, that mast is not rigged ship-shape; put her about ship-shape, &c.

Shivering. The state of a sail when fluttering in the wind.

Shool. Shallow.

Shoe the enchor. A small block of wood, convex on the back, and having a hole sufficiently large to contain the point of the anchor-fluke on the fore side; it is used to prevent the anchor from tearing the planks on the ship's bow, when ascending or descending.

To shoot a-head. | To advance forward.

Shore. A general name for the sea coast of any country.

To shorten sail. Used in opposition to make sail.

Shrouds. A range of large ropes extended from the mast heads to the right and left sides of a ship, to support the masts, and enable them to carry sail.

Sinnett. A small plaited rope, made from rope-yarns.

Slack-water. The interval between the flux and reflux of the tide, when no motion is perceptible in the water.

Slatch, is applied to the period of a transitory breeze.

To slip the cuble. To let it run quite out when there is not time to weigh the anchor.

To slue. To turn any cylindrical piece of timber about its axis without removing it. Thus, to slue a mast, or boom, is to turn it in its cap or boom iron. Also to turn any package or cask round.

Sound. To try the depth of water.

Sounding-line. A line to sound with, which is marked in the following manner: - Black leather at 2 and 3 fathoms, white at 5, red at 7, black at 10, white at 13, (some seamen use black at 10 and 13) white at 15 as at 5, red at 17 as at 7, two knots at 10 fathoms, and an additional knot at every ten fathoms, with a single knot midway between each 10 fathoms, to mark the line at every five fathoms.

To spell the Mizen. To let go the sheet and peek it up.

To spill. To discharge the wind out of the cavity or belly of a sail, when it is drawn up in the brails, in order to furl or reef it.

Spilling lines, are ropes contrived to keep the sails from being blown away, when they are clewed up in blowing weather.

Splice. To make two ends of ropes fast together by untwisting them, and then putting the strands of one piece with the strands of the other.

Split. The state of a sail rent by the violence of the wind.

Spoon-drift. A sort of showery sprinkling of the sea water, swept from the surface of the waves in a tempest, and flying like a vapour before the wind.

Spray. The sprinkling of a sea, driven occasionally from the top of a wave, and not continual as a spoon-drift.

To spring a mast, yard, &c. To crack a mast, yard, &c. by means of straining in blowing weather, so that it is rendered unsafe for use. To spring a leak. When a leak first commences, a ship is said to spring a leak. To spring the luff. A ship is said to spring her luff, when she yields to the effort of the helm, by sailing nearer to the wind than before.

Spring-stays, are rather smaller than the stays, and placed above them, and intended to answer the purpose of the stay, if it should be shot away, &c.

Spring-tides, are the tides at new and full moon, which flow highest and ebb lowest. Spurling Line, is a line that goes round a small barrel abaft the barrel of the wheel, and coming to the front beam of the poop deck, moves the tell-tale with the turning of the wheel, and keeps it always in such position, as to show the position of the tiller.

Spur-shoes, are large pieces of timber which come abast the pump-well.

Squall. A sudden violent blast of wind.

Square. This term is applied to yards that are very long, as terms is to high masts. To square the yards. To brace the yards so as to hang at right angles with the keel.

To continue advancing. To stand in. To advance towards the shore.

To recede from the shore. To stand off.

Starboard. The right hand side of the ship when looking forward. Starboard-tack. A ship is said to be on the starboard tack when sailing with the wind blowing upon her starboard side.

Starbeard the helm! An order to push the helm to the starboard side.

To stay a ship. To arrange the sails and move the rudder so as to bring the ship's head to the direction of the wind, in order to get her on the other tack?

Stour. Large ropes coming from the mast heads down before the masts, to Brevent them from springing, when the ship is sending deep.

Steady! The order to the belmsman to keep the ship in the direction she is going at

that instant.

Steering. The art of directing the ship's way by the movement of the helm.

Steerage-way. Such degree of progressive motion of a ship as will give effect to the motion of the helm.

Stem. A circular piece of timber, into which the two sides of a ship are united at the fore end; the lower end is scarfed to the keel, and the bowsprit rests on the upper end.

To stem the tide. When a ship is sailing against the tide at such a rate as enables her to overcome its power, she is said to stem the tide.

Sleeve. Turning up. The bowsprit steeves too much, that is, it is too upright.

Sternfast. A rope confining a ship by her stern to any other ship or wharf.

The furthest astern, opposed to headmost. Sternmost.

The motion by which a ship falls back with her stern foremost.

Stiff. The condition of a ship when she will carry a great quantity of sail without

hazard of oversetting. It is used in opposition to crank.

Stoppers. Large kind of ropes, which, being fastened to the cable in different places abaft the bitts, are an additional security to the ship at anchor.

To stow. To arrange and dispose a ship's cargo.

Strand. One of the twists-or divisions of which a rope is composed. It also implies the sea beach.

Stranded. This term, speaking of a cable or rope, signifies that one of its strands is broken: applied to a vessel, it means that she has run aground and is lost.

To stream the buoy. To let it fall from the ship's side into the water, previously to cast ing ancher.

Stretch out. A term used to men in a boat when they should pull strong.

To strike. To lower or let down any thing. Used emphatically to denote the lowering of colours in token of surrender to a victorious enemy.

To strike sounding. To touch ground when endeavouring to find the depth of water. Sued, or Sewed. When a ship is on shore and the water leaves her, she is said to Be

sued; if the water leaves her two feet, she sues, or is sued two feet. Surf. The swell of the sea that breaks upon shore or on any rock. To surge the capstern. To slacken the rope heaved round upon it.

Sway away. Hoist.

Swell. The fluctuating motion of the sea either during or after a storm.

Sweeping. The act of dragging the bight or loose part of a rope along the surface of the ground, in a harbour or road, in order to drag up something lost.

Swinging. The act of a ship's turning round her anchor at the change of wind or tide.

To tack. To turn a ship about from one tack to another, by bringing her head to the wind.

Taffarel. The uppermost part of a ship's stern.

Taking in. The act of furling the sails. Used in opposition to setting.

Taking a-back. See a-back.

Tamkins, or Tomkins. The bung, or piece of wood, by which the mouth of a cannon is filled to keep out wet.

Turpendin. A cloth of canvass covered with tar or some other composition, so as to make it water proof.

Taught. Improperly though very generally used for tight.

Taunt. High or tall. Particularly applied to masts of extraordinary length.

Tell-tule. An instrument which traverses upon an index in the front of the poop-deck, to show the position of the tiller.

Tending. The turning or swinging of a ship round her anchor in a tide-way at the beginning of ebb and flood.

See a-thwart. Thwart ships. See a-thwart ships.

Thus. An order to the helmsman to keep the ship in her present situation, when sailing with a scant wind.

To tide. To work in or out of a river, harbour, or channel, by favour of the tide, and anchoring whenever it becomes adverse.

ide it up. To go with the tide against the wind.

Tide it up. To go with the tide against the winu.

Tide-way. That part of the river in which the tide ebbs and flows strongly.

Tide-way. a tier of casks, a tier of ships, &c. Tier Tier. A row; as a tier of gups, a tier of casks, a tier of ships, &c. Tier of a cable. A range of the fakes or windings of a cable which are laid within one another, in a horizontal position. Cable Trer—the space in the midst of a cable when it is colled;

also the place in which it is coiled.

Tiller. A large piece of wood or beam, put into the head of the rudder, and by messas of which the rudder is moved.

Topping. Pulling one of the ends of a yard higher than the other.

Tart or Taut, signifies tight.

To Tow. To draw a ship in the water by a rope, fixed to a beat or other ship which is rowing or sailing on.

Toto-line. A small hawser, or rope, used to remove a ship from one part of a harbour to another.

her after part, and to give it the figure most suitable to the service for which she is calculated. Transoms. Certain beams or timbers extended across the sternpost of a ship to fortify

Treverse. To go backwards and forwards.

Treenails or Trunnals. Long wooden pins employed to connect the planks of the ship's side and bottom to the corresponding timbers.

Trey-sail. A small sail used by cutters and brigs in blowing weather.

Trice, trice up. To haul up and fasten.

Trim. The state or disposition by which a ship is best calculated for the purposes of navigation. To trim the hold-to arrange the cargo regularly. To trim the sails-to dispose the sails in the best arrangement for the course which a ship is steering.

To trip the suchor. To loosen the anchor from the ground, either by design or accident.

Trough of the Ses. The hollow between two waves.

A round piece of wood put upon the top of flag-staves, with sheaves on each side for the haliards of the flags to reeve in.

Turning to soundward. That operation in sailing, whereby a ship endeavours to advance against the wind.

To Unballast. To discharge the ballast out of a ship.
To Unbend. To take the sails off from their yards and stays. To east loose the anchor from the cable. To untie two ropes.

To Unbit. To remove the turns of a cable from off the bits.

Under foot, is expressed of an anchor that is directly under the ship.
Under sail, or under way. When a ship is sailing she is said to be under way

Under the lee of the shore, is to be close under the shore which lies to windward of the ship.
Unfaul. Cast loose the gasket of the sail.

To Unmoor. To reduce a ship to the state of riding at single anchor, after she has been moored.

To Unreeve. To draw a rope from out of a block, timber, &c.

To Unrig. To deprive the ship of her rigging.

Urrou. The piece of wood by which the legs of the crow-foot are extended. Uorou.

Van. The foremost division of a fleet in one line. It is likewise applied to the foremost ship of a division.

Venc. A small kind of flag worn at each mast head.

To Veer or Weer the ship. To change a ship's course from one tack to the other, by turning her stern to windward.

Veer. Let out, as veer away the cable.

Veer. Shift. The wind veers, that is, it shifts, changes.

To Veer and Hand. To pull a rope tight by alternately drawing it in and stackening it. Viol or Voyal. A block through which the messenger passes in weighing the anchor.

A large messenger is called a viol.

Wake. The path or track impressed on the water by the ship's passing through it, leaving a smoothness in the sea behind it. A ship is said to come into the wake of another when she follows her in the same track, and is chiefly done in bringing ships to, or in forming the line of battle.

Wales, are strong limbers that go round a ship a little above her water-line.

Warp. A small rope employed occasionally to remove a ship from one place to another.

To Warp. To remove a ship by means of a warp.

That part of a ship contained between the quarter deck and the fore-castle. Waist.

Water line. The line made by the water's edge when a ship has her full proportion of stores, &c. or board.

Water-borne. The state of a ship, when there is barely a sufficient depth of water to float her off from the ground.

Water-logged. The state of a ship become heavy and inactive on the sea, from the great quantity of water leaked into her.

Water-tight. The state of a ship when not leaky.

Weather. To weather any thing is to get to windward of it. Synonymous with wind-

Weather-beaten. Shattered by a storm. Weather-bit. A turn of the cable about the end of the windhas. Weather-gage. When a ship or fleet is to windward of another, she is said to have the weather-gage of her Weather-quarter. That quarter of the ship which is on the windward side. Weather-stde. The side upon which the wind blows. Digitized by GOO

To Weigh Anchor. To heave up an anchor from the bottom.

To Wind a Ship. To change her position, bringing her head where her stern was.

Wind road. When a ship is at anchor, and the wind being against the tide, is so strong as to overcome its power and keep the ship to leeward of her anchor, she is said to be wind-road.

Wind's eye. The point from which the wind blows.

To Windward. Towards that part of the horizon from which the wind blows.

Windward Tide. A tide that sets to windward.

To Work a ship. To direct the movements of a ship by adapting the sails and managing the rudder according to the course the ship has to make.

To work to windward. To make a progress against the direction of the wind.

West. To bind round with ropes.

Yerds. The spars upon which the sails are spread.

Tasebug. The motion of a ship when she deviates from her course to the right or left.



EVOLUTIONS AT SEA.

Of the Ballast and Lading.

WHEN a ship is loading, it should be considered that her tendency to pitch or roll depends not alone on her form, but even more upon the distribution of the heaviest

parts of her cargo.

Particular attention is to be paid to moderate her pitching, as that is what most fatigues a ship and her masts: and it is mostly in one of these motions that masts are seen to break, particularly when the head rises after having pitched. Although the relling be proportionably a more considerable movement than pitching, it is seldom any accident is seen to arise from it, as it is always a slow one. It is however not less proper to prevent it as much as possible. This will, in general, be easily obtained, without being any way detrimental to the ship's stiff carrying of sail, if, when the ballast is iron, you stow it up to the floor heads; because it will recall the ship with less violence after her having inclined, and it will act on a point but little distant from the centre of gravity.

In the merchant service the stowage consists, besides the ballast, of casks, cases, bales, boxes, &c. which are all carefully wedged off from the bottom, sides, punpwell, Sc. and great attention paid that the most weighty materials are stowed nearest to the centre of gravity, or bearing of the ship, and higher or lower in the hold, agreeable to the form of the vessel. A full low built vessel requires them to be stawed high up, that the centre of gravity may be raised; to keep her from rolling away her masts, and from being too stiff and laboursome; as, on the contrary, a narrow, high built vessel requires the most weighty materials to be stowed low down, nearest the kelson, that the centre of gravity may be kept low, to enable her to carry sail, and to prevent her

oversetting.

To anchor in fine weather in a place where you will ride head to wind, being close hauled.

Being under the three topsails, fore-topmast stay sail, and mizen, stand on until you are within about two ship's lengths of the place where you mean to drop your anchor; then put the helm a-lee, and haul down the fore-topmast stay sail. As soon as the topsails shiver, clue them up briskly, before you lower, except the mizen topsail, which is to be laid to the mast, and the mizen sheet hauled flat aft, the instant the ship begins to have stern way, by reason of the wind being a-head. Then shift the kelm to windward, and let go the anchor, veering away the cable, to give it time to settle in the ground, until the vessel falls off, when she is to be checked, to bring ter head to the wind. When that is done, right the helm, and haul up the mizen.

To anchor in fine weather in a place where you will ride head to wind, he wind being large.

If you have the wind large, whether on the beam or more aft, the operation is still the same, only hauling up a little sooner to keep to windward, because it is in your power to drift as much as you think requisite, and because the ship will be entirely stepped as soon as all her sails begin to catch a back, and you will have done chaing them up when they begin to shake. The mizen topsail is next to be heaved to the mast, the helm put a-weather, and the anchor let go, as soon as the head way seases; then after giving her a sufficiency of cable, bring the ship up. If she has been going large she will not range precisely head to wind, since her headway coases as soon as the sails are taken a back, and the effort of the wind acts on all the rigging of the ship to impel her

both a stern and to leeward, which is indeed augmenting the effect of the rudder, as the helm is a-weather to bring the vessel to the wind; but as the power of the wind is very great to pay the ship's head off, it balances, wholly or partly, (according as the ship goes a stern with more or less velocity) the effort of the rudder and that of the mizen: thus she drifts, and remains as it were lying to with all her sails aback. This is the reason why we keep a little to windward, and let go the anchor, to bring the ship's head to wind at the proper time, which she will do the more readily as she is withheld forward by the cable, while the wind on her side forces her to leeward.

To anchor in fine weather in a place where you are to ride head to the stream and wind, the wind being large.

If you are obliged to ride with the head to the stream, you must, when it comes from the windward, put the helm a-lee in setting the mizen, then clue up the sails: and when the ship's head is right in the direction of the stream, let go the anchor, provided she has quite lost her headway; for else, you would get foul of the anchor stock by running over it. This must never be neglected, unless you find yourself under the necessity to bring up in any situation in which you may happen to be, which is almost always the case when you are taken too short to have time to stop the vessel: a reason why there is often a necessity of easting a second anchor, which generally eatches the ground by assistance of the first, which has begun to diminish the velocity of the ship; and as many of the sails are to be hauled down as you can, and as quick as possible.

To anchor in fine weather in a place where you will ride head to the stream, which Etimes from leeward, the wind being large.

When the current comes from the leeward, you must keep the ship away till her heaft comes to the set of the stream, and take in all the sails to diminish as speedily as possible her headway, which always continues of itself long enough when the wind is aft or very large; and when the ship is stopped by the effort of the water, let go the anchor without waiting for the vessel gathering stern-way, if the current is rapid; and in this case, as well as all those wherein there is a sea, or blowing fresh, the ship requires a great length of cable.

To come to an anchor with the wind aft.

First, hand the main topsail, and then lower the fore topsail down on the cap; and when you are within a reasonable distance of the place where you mean to drop anchor, (which distance is to be judged of from the readiness of the ship to obey the helm, and from her velocity) the tiller may be put either one way or the other, the fore topsail and fore topmast stay sail claed up and taken in, the mizen topsail braced sharp up, and the mizen sheet hauled flat aft. When the ship ranges close to the wind, she is, as it were, lying to under the mizen and mizen topsails, with the last mentioned sail full or a-back according as you may have occasion to shoot a-head or drop a-stern; so that if you are too much to windward of the spot where you mean to bring to, you drift till you arrive at it: if you are precisely in the proper birth, you let go the anchor in lowering down the mizen topsail, which is to be furled as soon as the vessel is brought up; then the ship will come head to wind by the power of the mizen, which must be brailed up as soon as it shakes.

Scudding under a foresail, to come to an anchor.

The foresail must be clued up when at some distance from your birth; and, some part of the way, run under bare poles. When near enough to sheer to the wind, you execute it by putting the helm hard a-lee; and as soon as the ship is come to, let go the anchor, giving her a large scope of cable, and observing to check her handsomely, in order to make her ride head to wind, as stopping her at first too short might endanger her cable or anchor. Should the first not bring her up, a second may be let go.

To anchor with a spring, in order to present the vessel's side to a place or ship you wish to cannonade.

This is executed when you know that the wind or current will bring your head, when at anchor, towards the object you mean to attack; for should the wind or tide bring your broadside to bear on the object you mean to cannonade, the spring would only be a precaution to get under way more quickly in case you were obliged to retreat, or in case the wind or tide should shift.

Get a large snatch-block in the aftermost port, or on the same side you wish to present to the wind or current, and on the same side with the anchor and cable with which you mean to bring up; then, through the block reeve a hawser, the end of which is to be eliached to the ring of the anchor you mean to lot go; the other part is to be brought to the capatan with necessary ranges of the cable and hawser on deck. That

ie, and the ship being arrived at the birth, you are to deaden her way according to sumstances; then let go the anchor, and veer away enough cable and hawser, now a le more of the one, and then a little more of the other, according as you wish to pret more head or stern; which you can do by heaving on the spring, or what is the se, veering away more cable. Should you find it requisite to shift your position, I have only to veer out more of the hawser.

come to anchor in roads that are often crowded with ships, and to leave clear births for others.

The best anchoring births in these places are mostly known by marks, and of course occupied by the first ships.

n a tide or trade wind road-stead, the next ship that comes should not anchor right and or a-stern of the first ship and so as to lie in the other's hawse, but should come on the bow and quarter at a sufficient distance to prevent other ships from coming ween, and in a slanting direction from the tide or wind. This might contribute to safety of ships when it blows strong upon a lee-tide or in strong sea breezes, as h single ship may then veer away what cable necessary, and keep clear of the other p's hawse a-stern; or, in case of driving or casting, they have a better chance of ring clear of each other.

To get up an anchor, in ships which have a main and jeer capstan.

n large ships which have a main and jeer capstan, and the strain is thought too great the messenger alone, the viol is used thus: Three or four turns are taken round jeer-capstan with one end, so as to leave that side clear on which the cable is coming and pass the other end through the viol block, which is lashed round the main-mast the lower deck. It is then carried forward and passed round the rollers in the mannear the hawse holes; then brought aft, and spliced to the other end with a short ice, and the ends marled down tight. That side of the viol on which the cable is ning in is fastened to the cable by nippers; and thus the continued efforts of the stan are conveyed to the cable, until it is hove in. The nippers are clapt on in the nger, from one to two futhoms asunder; and the viol is applied to the midships, or ide of the cable. Nippers are clapt on by taking three or four turns round the viol, r turns round the cable and viol, and then three or four turns round the cable. This thod is an exceeding good one, and very suitable to quick heaving; but when the ain is great, and the cable muddy, the nippers clapt on after this method will not nip ficiently, and some times recourse is had to the following method: Throw sand or es upon the cable, and take a long dry nipper, which middle and pass one half aft, king it in and out round the cable and viol; then worm its end round the viol only. er this pass the other half in the same manner forward, but worm its end round the ile only, and let each end of the nipper be held on. The advantages of this method , that as the strain of the cable lies forward and that of the viol aft, the nipper will drawn so tight as effectually to hold the cable till something gives way : also they never jamb, for both ends are clear for taking off. Another method, when the strain creat, is, to have nippers with an overhand knot made at one end; and with that end ound turn taken round the cable and viol, leaving three or four feet of the end; then h the other end, take three or four racking turns, and expend nearly the remainder h turns round the cable and viol, laying the knotted end under and over each of the turns; the end is then held fast. The men who clap on the nippers are attended boys, who hold the ends of them, and follow the progress of the cable as it is hove ; and, as the nippers arrive near the main hatch-way, they are taken off and carried ward, where they are again clapt on, and so in succession until the cable is hove in ficiently to raise the anchor above the water. It is then stoppered round all before bitts, that is, round the cable and viol. The anchor is then catted, and afterwards ed. To shift the viol for heaving in a second anchor, it must be unspliced, and the ns round the capstan reversed. When the strain is so great as to require other purises, the top tackles may be used thus: the double block is lashed to the main-mast topsail-sheet bitts, the treble block is lashed on the cable, and the fall is brought to the stan. If the top tackle falls are thought insufficient, any hawser may be used that I reeve through the blocks.

To get up an anchor in ships which have not a jeer-capstan.

ships without a jeer-capstan have no viol, but heave in their cables by a messenger, ich has an eye spliced in each end, one of which ends is passed with three or four ns round the capstan on the upper deck, and the other end passed forwards round rollers, at the fore part of the ranger; then brought aft to the other end, and labelthus: several turns are passed through the eyes crossing each other in the middle, na half hitch is taken round the parts, and the end stopped with spun-yarn. The rainder of the operation is performed as by the viol, with this exception; the mes-

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senger is applied to the outside of the cable, and when the nippers are insufficient, the messenger may be hitched thus: the bight of the messenger is fastened round the cable at the manger with a rolling hitch, and the bight seized round the cable before the hitch. This practice is by no means so good as the others.

When getting under way in a sea gale, the viol is better than a messenger, as the sending of the ship carries all the strain to the main capstan, and endangers the men at the bars, whereas with a viol, the strain is taken to the viol block, and the men at the

fore jeer-capstan heave in security.

To get up a second anchor.

Suppose by the former method, that the starboard anchor is got up, and that the cable of the second anchor enters the larboard hawse-hole, the operation of getting up the second anchor is the same, observing only, that the messenger must be shifted, and the turns on the capstan reversed, to change the disposition and side; and the men, who before held on the larboard side is the first operation, will hold on the starboard side now. The motion of the capstan is performed the contrary way, and the cable on the larboard side is fixed and hove in.

To get up an anchor in Merchant Ships.

Most merchant ships and small vessels heave up their anchors by a windlass, round which are taken three turns of the cable, and held on by hand, or by a jigger; thus—the end of the rope which has the sheave is passed round the cable, with a round turn, close to the windlass, the leading part of the rope coming over the sheave, and stretched aft, by means of the fall passing through the jigger block; the standing part of the fall is made fast round a stantion, at the fore part of the quarter deck, and the leading part is bowsed upon, which jambs the turns taken round the cable, and, when the jigger arrives abreast of the hatchway, it is removed forward, and the cable is jambed by a handspike at the windlass, until the jigger is refixed.

To weigh an anchor with the long boat.

This is done by taking the long boat to the buoy of the anchor, and putting the buoy-rope over the davit of the long boat, and a tackle on the buoy-rope; by which, with the assistance of men on the fall, the anchor is weighed out of the ground. This being accomplished, the cable is hove in on board; the buoy-rope and tackle being secured in the boat, they approach the ship as the cable is hove in, and the anchor catted and slowed. Small anchors and grapnels are got up by the davit, hauling upon the grapnel rope by hand.

To weigh an anchor by under-running. .

This is by placing the cable over the davit-head, and under-running it, till it is nearly apeck, when it is tripped by means of tackles, as before by the bouy-rope. This method is troublesome, and is only adopted when the buoy is gone, and a ship cannot get near her anchor for want of water.

To get under sail when the ship is swinging head to wind, and you want to cast either to starboard or larboard, in a place where there is no current.

To cast to starboard.

Heave short on your anchor till it is apeek; then haul in quite home, the larboard braces forward, and starboard braces abaft; loosen, sheet home, and hoist the topsails; put the helm a starboard, and heave till the anchor is aweigh. The moment the anchor quits the ground, the ship will begin to fall off to starboard. As soon as this movement is perceived, hoist the jib and fore-topmast staysail, if necessary, to help her; and when she has sufficiently fallen off, her sails abaft (which are trimmed sharp for the larboard tack) will fill. But, unless for very superior reasons, you had better continue lying-to till the anchor is catted, taking care to haul the mizen-sheet close aft, if the ship be inclined to fall off too much.

To cast to larboard.

Haul in the starboard braces forward and the larboard aft, and put the helm a-port. The rest of the operation is the same as the preceding, only changing the starboard for port.

To get under sail when the ship is riding head to wind and tide.

If a ship, riding head to wind and tide, wanted to get under sail, after having decided on which side it is best to have her cast, it must be performed according to one of the foregoing methods, except with regard to the helm, which must be put to starboard either before the anchor loosens, or while it does, if you wish to cast to port; because the water coming from forward, acts with the same force on the rudder as if the ship went with the current, impelling the rudder to starboard and head to port. Therefore it is evident in this case, the helm ought to be put to starboard; which, on the contrary, would be put to larboard, if the ship were to be cast to port.

If the ship, after the anchor is out of the ground, goes astern faster than the current runs, the helm must then be used as if there was no current, because the excess of ve-

locity, whereby the ship exceeds that of the water, acts upon the rudder.

If it blows fresh, so that you cannot set your topsails without reefing them, let that be done before they are sheeted home; and if it blows so hard as to be obliged to go only under a foresail, it would then be sufficient to loosen the fore-topsail, without sheeting it home, after having braced it quite close on the side opposite to that you want the ship to cast, not forgetting however to put the helm the same way as you cast, as soon as you perceive the ship going astern: and when the ship has fallen off sufficiently, then is the time to fill and trim the foresail.

To get under sail when the ship is swinging with her head to the current, and with the wind a point abaft the beam.

Heave short on your anchor till it is a-peek; next to this loosen, sheet home, and hoist the foresail and mizen-topsail, keeping the wind in, and heave vigorously at the clastan, till the anchor is a-weigh. At the same time hoist the jib and for topmast staysail, or haul out the mizen, according as circumstances may require. Whether you wish to come to windward, or fall off more quickly, you must still continue to heave round the capstan briskly, to get the anchor up, till you find yourself sufficiently offward to bring to, in order to stow it with ease, or to stand on under an easy sail with the anchor hanging out to windward, if the situation of things will admit of it. You may sometimes also hoist up both the main and fore-topsails, as soon as you get ready; but in certain cases, as when obliged to make the best of your way from an enemy, every sail possible must be set at once which the weather will admit of, especially when obliged to haul by the wind; in which case, the anchor must be got up and catted as well as it can; there are cases even when, without losing your time in weighing it, you crowd as many sails as you possibly can, and depart, in cutting or slipping the cable.

To get under sail with a spring.

If a ship be in a place too confined to cast under her sails only, or being obliged to but to sea in a gale of wind, without hoisting the anchors, you must, for greater safety, in casting the right way, get a spring out, to be clapped on the cable by which the ship swings, by pussing a howser or a stream cable through the aftermost port, on the opposite side to that you mean to cast; and after that spring is well have tight at the capstan, hoist the jib and fore-topmast-staysail, loose and sheet home the fore-topsail; when that is done, and if the weather permits; brace quite close the head sails, on the same side with the spring. When this is executed, slip or cut the cable, heaving briskly at the same time on the spring, till the ship has paid off sufficiently. Then fill the sails, by setting the mizen-topsail and every other sail you mean to employ, and slip or cut the spring, as circumstances may require. Care must be taken, not to let the ship fall off too much, before the spring is cut; because, having no way through the water, she will not come to the wind so soon as might be wished; and for the same reason the spring must not be cut, till she has fallen-off as much as is necessary; because, although the has no other motion but that of falling off, the vessel might perhaps not wear enough to answer the purpose.

To get under sail with a leading wind, in a title way.

If the ship to be got under sail has a leading wind, and is in the midst of vessels, or in a narrow channel, where it would be difficult to cast her upon the lee-tide, she should be got under sail before the weather-tide is done. Thus, the casting of the ship would

be avoided, and she may be steered through the fleet or channel with safety.

Should it, however, blow so fresh upon the windward-tide as to force the ship end on with her cable, it will be impossible to heave it in, without sheering the ship over from side to side, and heaving in briskly as the ship slacks the cable; but as this is attended with much danger, by the ship suddenly bringing up upon each sheer, it will be best to heave a-peck upon the first setting of the windward tide, before the ship swings to bring the wind abaft.

To cast a ship upon the larboard tack, and back her a-stern of danger.

We suppose the ship to lie at single anchor, with the wind and tide the same way, and ships or shoals right astern, in the intended course, and that to clear them, you

must cast upon the larboard tack, and make a stern board.

Make every thing as ready as possible before weighing: let the three topsails be hoisted, the yards braced up sharp with the larboard braces, and the mizen hauled out. Thus situated, when the anchor weighs, put the helm a-port. The tide, running aft, acts upon the starboard side of the rudder; and in that direction it will cast the ship the right way, and bring the wind upon the larboard bow. The wind being on the larboard bow, and the topsails a-back, will soon give the ship stern-way through the water; then the water will act against the larboard side of the rudder, and powerfully prevent

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the ship falling too fast off from the wind. Thus she will drive till the anchor is got quite up, and may be so continued till she has past the shoals and has room to veer, and

get upon her proper course.

It is advantageous to make a stern-board in getting under way from a single anchor in the above situation. The anchor heaves up more easily when the ship goes a-stern; and while heaving up it serves to keep the ship's head to the wind. A ship, however, cannot long be steered stern foremost when under sail, so as to keep the wind before the beam; but she will in a little time drive broadside through the water till she gets headway, and then it is proper to veer, provided the anchor be quite up.

To east a skip on the larboard tack, in a tide-way, with the wind two points on the starboard bow.

A ship, riding in a tide-way, with the wind two points on the starboard bow, and so near the shore on the larboard side, that she must be cast upon the larboard tack, to clear the shore, the three topsails must be hoisted, and the yards sharp braced up, with the larboard braces forward, and the starboard braces aft, with the starboard foretop bewline well hauled, putting the helm hard to port at the anchor's weighing; the tide acting upon the radder, and the wind upon the sails braced in that direction, brings the ship about with the wind on the larboard bow, before she gets stern-way, which should be always strictly noticed; for in all proceedings of this kind, if a ship gets stern-way before she brings the wind right ahead, she will not come about the right way. In that case, it is best to veer away the cable directly, and bring the ship up again: and carry out a kedge or small anchor on the larboard bow, hauling its cable or hawser in tight on the larboard quarter, when the bower anchor is a-peck. If this fail, the ship must lie till the windward tide makes, to bring the wind on the larboard bow, when the ship may be got under way, and clear the shore.

To cast a ship upon the larboard tack in a lee-tide; and shoot her by the wind a-head of danger.

If there be just room enough to go close by the wind to clear a danger lying to leeward, much depends on heaving up briskly the anchor after it is out of the ground, and The three topsails must be having proper sails ready to set to the best advantage. hoisted, and the yards sharp braced up, with the larboard braces forward and the starboard braces aft, when the anchor is at a long peek. At weighing the anchor, put the helm hard to port, then the action of the tide upon the rudder, and the wind on the foretopsail, will cast the ship off the right way, so as to fill the after-sails, when the foretopsail may be soon braced about and filled before she gets stern way. The helm will keep the ship under command sufficiently to steer her by the wind shead clear of danger; but if the ship gets stern-way in casting, the helm should be kept hard a-weather, to prevent her falling off too much from the wind; and when she gets headway again, be cautious how the weather-helm is eased with the anchor much below the bows, by which the resistance forward is increased, and the ship may be brought up in the wind, so as to prevent her shooting clear of the danger. This must be guarded against by the weather-helm and head sails, as jib, fore-topmast-stay sail, &c. As soon as the ship has shot far enough a-head to clear the danger to leeward, and there being but little room a-head, it is best to bring the ship to, and drive with the helm a-lee, with the main and mizen-topsail a-back, and the fore-topsail shivering till the anchor is up; then take proper time to veer.

To cast on the larboard tack, when riding with the wind right a-head, and to veer her short round before the wind in little room.

The head sails should only be loose, viz. the fore-topsail hoisted and the foresail loose; brace sharp up the larboard braces, the jib and the fore-topmast-staysail set, with the larboard sheets flat aft. When the anchor is a-peck and a lee-tide running, at weighing the anchor, the helm should be put to port so far as to bring the wind two points on the larboard-bow, which should be kept so by steering the ship till the tide ceases to run aft. Then put the helm hard to starboard, or a-lee; and when the ship gets stern-way, the water will act powerfully on the starboard, or lee-side of the rudder, turning the ship's stern to windward, whilst the wind, acting at the same time upon the head sails a-back, will box her round off upon her heel, so as to bring the wind nearly aft by the time she loses stern-way. hen the ship will cease falling off, and soon get head-way, which should be attended to, and the head sails braced about flat with the starboard braces, and the helm shifted hard to port at the same time.

When there is no tide, but still water, at weighing the anchor, the helm must be hard to starboard; and, as the ship gets stern-way, the water meets with so much resistance against the starboard side of the rudder in that direction, that the rudder acts with great power to turn the ship's stern round to port, and the head sails being set and frimmed as before mentioned, and the foresail let fall with the starboard bowline liquided

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close forward, will assist to east the ship so far round the right way, by the same time she loses her stern-way, as then to permit your proceeding as before directed. To insure success, heave the anchor up briskly. I'he same methods are adopted in casting the ship on the starboard-tack, only the helm and sails are managed the contrary way.

To tack a ship in getting to windward as much as possible.

To execute this with propriety, care must be taken that the ship does not yaw, that she is not too near or too far from the wind; because both situations are equally preindicial.

When this medium is obtained, haul the mizen out, while you put at the same time the helm a-lee, brace the sail to windward, in order that it may be as much as possible exposed to the wind. When the ship is come to the wind, so as to cause the square sails to shiver, let go the jib, and all the staysail sheets before the mainmast. At the moment when all the sails catch a-back, and particularly the mizen-topsail, let it be braced sharp about the other way, hauling up at the same time the weather-clew of the mainsail; and when the wind is right a-head, or even a little before, haul the mainsail, and trim sharp for the other tack as fast as possible. The jib and staysail sheets are also to be shifted over, at the same time righting the helm, whether the ship has lost her way, or even still advances a-head. Then as soon as she has passed the direction of the wind about 45°, in continuing her evolution, shift the foremast's sails, which are to be trimmed with celerity, at the same time putting the helm a-lee, if you fear the ship (which must still go a-stern if the operation be slowly executed) will not fall off sufficiently; for, if the sails are braced about briskly, she will never have stern-way; on the contrary, she will get a great deal to windward.

To tack a ship without endeavouring to get to windward.

There are circumstances sometimes when it is found necessary to tack, without caring much whether the ship loses to windward or not. For example: when a ship is found suddenly to be close to the land, in the night, or in foggy weather, near a danger or some vessel, which must instantly be avoided by staying the ship, because you find yourself to windward, and too near the object from which you wish to recede: in this case, when it is necessary to deaden the ship's way, and tack at the same time, you must suddenly put the helm hard a-lee, and in the same instant, let go the jib, fore, and staysail sheets, without touching the bowlines; and great care must be taken that the effect of the mizen is to be preserved as much as possible. When the sails begin to shiver, the mizen is hauled quite to windward; then, if the ship takes well the wind a head, the remainder of the operation must be executed as directed in the preceding case; but, if you should miss stays, you must proceed according to the second method of veering, called box-bauling.

To tack a ship in a dangerous rough sea, when her staying is doubtful.

Let every thing be got clear and ready; the hands at their proper stations, the sails trimmed fair, and the ship stecred just full, and close by the wind. Take the advantage of the smoothest time, when the ship has the most head-way. The other necessary precautions are, to haul down the jib, if set, and not to put the helm a-lee all at once, but to luff the ship up by degrees, to shake the sails. When they shake, give these orders: the helm hard a-lee! let go the lee sheets forward, but not the lee braces and fore-top bowline, as that usual practice backs the head sails too soon, and stops the ship's head-way, which ought to continue to give powen to the helm. till the wind is brought a-head, or the ship will not stay. Raise tacks and sheets and mainsail haul, when the wind is a point on the weather-bow; this swings the yards round sharp, that the main-tack may be got close down, whilst the head sails becalm the fore-leech of the main and main-topsails; while the wind, blowing aslant on the after-leech of these sails, acts jointly with the rudder to turn the ship's stern, so as to bring her about the right way. When she has failen off five or six points, let go and haul.

When a ship comes about, she is sure to have stern-way by the time the head sails are hauled, therefore the helm should not then be shifted a-lee, but should be kept hard a-weather, till her stern-way ceases. The water, acting upon the weather side of the rudder, prevents the ship falling around off from the wind, which the helm, when hard a-lee, occasions, while the stern-way continues. Notice should be made by the compass, that the ship continues coming about till the wind is on the other bow; for if she stops with the wind a-head, and her head-way is perceived to be done, the helm should be directly shifted to the other side, so that, by the stern-way, the water may act upon the rudder and bring her about, and then the helm should not be kept a-lee, but directly shifted and kept hard a-weather till her stern-way ceases. For the reason just given, the head sails may be hauled as soon as possible; for, the ship will be sure to fall off the faster and further in proportion to her stern-way; so that the weather-braces should be tended to prevent the head yards flying fore and aft, as they will do when it blows fresh; and

to keep the head sails shivering, that the fore-tack may be got close down easily, and the ship stopped the sooner from falling off. Shift the helm a-lee when the stern-way ceases, and the head sails may be trimmed sharp as the ship is perceived to come to.

On turning to windward in very narrow channels.

At weighing, if the wind is partly across the tide, it will cast the ship with her head towards the weather-shore, which she may be kept clear of, by driving with her sails a-back till the auchor is up and stowed; and, as the tack towards the weather-shore is the shortest, it is prudent to back as near the lee side as possible, in order to make the first board the longer; to get the three top-sails, jib, stay-sail, and mizen, properly set; and to get all ready in time for tacking. Make as bold as possible with the weather shore, because on that side a ship is always surest in coming about; and in case of missing stays, a ship may be backed off from the weather-shore, till she has room to fill and set the sails, and get sufficient head-way to try her in stays again without danger. But when the ship is got about, and standing towards the lee-shore, it may be necessary to put her in stays in good time, because she does not so certainly stay when going slanting with the tide, as when going across it.

By staying her thus in good time, if she even miss stays, there may be room enough to fill and try her the second time, or to use such means as may prevent her going on

shore.

But, when the wind is right against the tide which begins to make to windward, be cautious not to weigh the anchor till the ship swings end-on to the tide, and brings the wind so far aft, that she may be steered right against the tide, till the anchor is up and

stowed, and the sails with which the ship is to work are all ready.

Haul the wind and get ready for tacking, when you are close over to one side, to gain the whole breadth of the channel for getting under way. For this purpose let the first trip be made as short as possible, till it is found how the ship works upon both tacks; and then make longer or shorter boards accordingly, but take care not, to stand into an eddy tide on either side, which has often occasioned ships to miss stays and go on shore. If a ship will not stay, she must be veered, box-hauled or club-hauled.

To veer a ship without losing the wind out of her sails.

To execute this evolution both the main sail and mizen must be hauled up, the helm put a-weather, and the mizen-topsail a shivering, which will be kept so till the wind be right aft, suppressing for that purpose the effect of all the staysails abaft the centre of gravity. As the ship falls off, (which she will do very rapidly) round-in the weather-braces of the sails on the fore and mainmast, keeping them exactly trimmed to the direction of the wind, and remembering also that the bowlines are not to be started till the ship begins to veer. As she falls off, ease away the fore-sheet, raise the fore-tack, and get aft the weather-sheet, as the lee-one is eased off, so that when the ship is right before the wind, the yards will be exactly square. Then shift over the jib and staysail-sheets; and the ship continuing her evolution, haul on board the fore and main tacks, and trim all sharp fore and aft, remembering to haul aft the mizen and mizen-staysail sheets as soon as they will take the right way, or when the ship's stern has a little passed the direction of the wind. When the wind is on the beam, right the lelm to moderate the great velocity with which the ship comes to; the sails being trimmed, stand on by the wind.

To veer a ship that has lost her foremast.

Run out the end of a cable or hawser over the lee-quarter, and buoy it up from the ground with empty casks, &c. in case of coming into shoal water with little wind. This will assist the helm with such power, as to make the ship veer and steer at pleasure.

A spare yard or boom, rigged out abast the mizen shrouds, may guy the end of the cable or hawser more or less on either quarter, according as the ship may have occasion to sail. It may be easily shifted from side to side, and guyed out to leeward in proportion to the ship's griping to answer sailing upon both tacks; and, when sailing before the wind, it may be secured over the middle of the stern, which will prevent the ship's broaching to against the helm either way.

This would likewise much assist deep laden bad-steering ships, and prevent their broaching-to, to which they are liable, in spite of the best helmsmen, often occasioning them to lie to even with a fair wind. With a little contrivance by blocks lashed to the rails on the quarters, to lead the guys fair to the steering wheel barrel, it may be made

to steer a ship that has lost her rudder.

, To veer when lying-to under a mainsail .-

Advantage must be taken of the ship's falling off to put the helm a weather, and ease away the main-sheet roundly; and, when the ship has fallen off about 50°, let go the main-bowline, and round in the weather brace, taking care to keep the sail full. When

the ship is before the wind, get on board the main-tack, and right the helm to moder-

ate her coming to.

If, in the beginning, the ship is found difficult to veer, the fore stay-sail may be hoisted, and the sheets hauled well aft; but it is to be hauled down as soon as the ship is before

A second method.

Make fast a four inch rope to the strings of the main yard; and when the ship comesto, so as to shiver the main sail, bring it down before the sail to the topsail sheet bitts, and let it be hauled tight and belayed. Then, as soon as she falls off, put the helm aweather, and let go the main sheet. By these means, the lee part of the sail no longer has any power to keep the ship to the wind, and the weather part, acting before the centre of gravity, will cause her to veer faster than by the first method; though, in general the first method will answer the purpose.

To veer under bare poles.

The fore-staysail must, if circumstances will allow it, be hoisted. But if that cannot be done, the head yards are to be braced up as sharp as possible, and those abast pointed to the wind. Then, if the ship veers, she will steer under the masts and ropes only. A number of seamen sent up and placed close to each other in the weather fore-shrouds will be found also of very great service.

To box-haul a skip, or the second method of veering.

In this evolution the most rapid execution is necessary. Briskly, and at the same instant, haul up both the mainsail and the mizen; shiver the main and mizen topsails; put the helm hard a-lee; raise the fore tack; let go the head bowlines, and brace about the head yards sharp the other way; and let the jib and staysail sheets go in the same instant. When the ship has fallen off 90° , brace the after yards square in order to give the ship a little way, and to help her (with the rudder, the situation of which must be changed) to double the point where all the sails shiver; and when the wind is aft, you will proceed as in the method of "veering without losing the wind out of the sails."

If the circular motion of the ship, after she has fallen off 90 continues pretty rapid, the filling of the after sails, to give the ship headway. may be dispensed with; because

she continues to turn by the effect of her helm, which must not be shifted, since the vessel still continues her stern-way. Therefore, after having veered a few degrees more, the wind will fill all her sails, and the ship consequently will have head way. Then

change the situation of the rudder to bring her before the wind.

In a case of absolute danger, when it might be necessary to go a-stern and fall off more rapidly, put the helm a-lee, brace all the sails a-back, and observing not to brace the after sails more than square, that they may not counteract the head sails, which are braced sharp a-back to pay the ship's head off; because the effect of the after sails, in this situation, is to impel the ship abaft in the direction of her keel, which, with those forward, contribute to give her fresh stern-way, in order to cause the ship to veer with greater celerity. The jib and fore-topmast-staysail sheets being hauled over to windward, will assist the ship in falling off and going astern.

Box-hauling is deemed the surest and readiest way to get a ship under command of the helm and sails, with the least loss of ground to leeward, when a ship refuses stays. The masters of sloop rigged vessels, turning to windward in narrow channels, when they want but little to weather a certain point run up-in the wind till the head-way ceases, then they fill again upon the same tack; this they call making a half-board. Thus a then they fill again upon the same tack; this they call making a half-board. Thus a ship in box-hauling may be said to make two half-boards, first running with her head, then with her stern, up in the wind; by which two motions a ship rather gains to wind-

ward.

To club-hard a ship.

Club-hauling is practised when it is expected that a ship will refuse stays upon a leeshore: place the hands to their stations for putting the ship about, and some by the lee anchor; then put the helm down, and if the ship make a stand before she brings the wind a head, let go the anchor and haul the mainsail. When the wind is a head, cut the cable, and the ship will cast the way required. The after sails being full, let go and hauk

.Inother method.

Bend a hawser to the kedge-anchor on the lee-bow, and bring the end into one of the after ports, or over the taffarel. Let go the anchor, brace up all sharp the contrary way. put the helm a-lee, and haul in briskly on the hawser. As soon as the ship gets headway, cut or slip the hawser, and carry a press of sail.

To lie-to to windward of a ship, so as not to drift near her.

The main-topsail must be braced sharp a-back; keeping the fore and mizen topsails full; because the wind acts with a very small sine of incidence on a sail when full, in comparison to what it does when braced sharp a-back; so that the fore-topsail, being full, draws the ship a head, and the effect of falling off is opposed by the main and mizentopsails. She will of course not fall off much; nor will her lee way be very considerable; for the ship is well kept to the wind by the disposition given to her sails.

To lie-to under the lee of another ship.

The fore topsail ought to be braced sharp a-back, the main and mizen topsails kept full; because these two last mentioned sails tend to give the ship head-way, and keep her to the wind; they may be assisted by the mizen, which will oppose the falling off occasioned by the fore-topsail. Thus, should the ship to windward fall off violently, or drift too much, you are more ready to veer short round, and avoid being boarded; because the fore-topsail being braced sharp a back, the impulse of the wind on it is much greater than if it were full: and it is well disposed to veer suddenly, as soon as the power of the other sails is suppressed.

To bring-to with the fore or main topsails a back to the mast or filled.

Either the fore or main-top-sail must be braced sharp a back, and the lee-bowline hauled up a little; the other two top-sails trimmed sharp, with the mizen hauled out, and the helm a-lee.

If you bring to with the fore-topsail to the mast, the head yards may be only laid square. Then the wind will act obliquely on the sail, and the ship will fall off but little, because the effect is in the direction of the keel from forward aft, and the sails abaft, keep the ship to. The main-topsail may be worked in the same manner, if you wish not to expose yourself much to the wind.

To bring-to with the three topsails a-b ck.

The jib and staysails being hauled down, brace sharp round at once all the sails you wish to lie a-back in hauling up the lee-bowlines, the etter to expose the sails to the action of the wind; haul out the mizen, and put the heim hard a-weather.

To fill when lying-to with the fore-topsail to the mast.

Brail up the mixen, hoist the jib and foretopmast-staysail, shiver the main and mixen topsails, and when the ship has fallen op 20° or 30°, fill the fore-topsail, which was a-back before, and stand on.

To fill when lying-to with the main-topsail to the mast.

Brace sharp and briskly the fore-topsail a-back; shiver the main and mizen topsails; hoist the jib and fore-topmast staysails, and brail up the mizen, all at the same time; and when the ship has fallen off 20° or 30°, fill the fore-topsail and stand on

If you are obliged to keep the wind on the same tack as that on which you are lyingto, you have only to right the helm, fill the topsail which is a-back, and trim it sharp, to

continue your course.

A second method.

Trim the topsail which was to the mast, in order to give the ship way through the water, and be able to tack or run large, according as may be found necessary. But this method is very tedious, unless you mean to heave in stays; in which case it will be most expeditious.

A third method.

Shiver the main and mizen topsails, keeping the fore-topsail full, righting the helfa, and running up the jib and fore-topmast staysail at the same time. As soon as the ship has fallen off enough to get headway, fill the after sails, and keep the ship in the direction you mean to follow. It is easily seen that this method, though the most common, is not the most expeditious, when you have to veer considerably.

To fill when lying to with all the sails to the mast.

Brail up the mizen, lay the after yards square, and shift the helm a-lee. When the ship has fallen off sufficiently to fill the after sails, those forward are then to be braced about and trimmed full also, in order to stand on.

Of lying-to in a gale of wind.

To lie-to when it blows hard, keep as close to the wind as possible under some one sail well trimmed, with the helm lashed a-lee as much as may be requisite for the ship; and as ships commonly bring to from the stress of contrary winds, care should be taken to heave-to under such sail as will least strain the ship; because there are some ships

which lie-to better under the foresail than mainsail, others are more easy under the mainsail, some under a mizen, and many vessels lie-to best under a main staysail.

Lying-to under a foresail.

This is advantageous for veering when you are well to windward; but it augments the lee-way, and is more subject to break the sea on board, on account of the ship's continual falling off, because in that movement she gathers way by yielding to the impulse of the gale, and is afterwards recalled to the wind by the helm; so that in springing the luff she meets the wave which comes from to windward.

Lying-to under the mainsail.

The ship does not in this situation fall off so easily as in the last mentioned mode, because its effect passes abast the centre of gravity of the ship; but it keeps the ship more to the wind, and consequently occasions less lee-way.

Lying-to under the mizen.

Under the mizen, ships keep better to the wind than under any other sail, because it is farther abaft the centre of gravity than any of the rest, consequently ought to keep the vessel from drifting more than any of the others; but it is inconvenient should you have occasion to veer suddenly.

Lying-to under the main staysail.

Under the main staysail a ship will not make so much lee-way as under a foresail, because its efforts pass very near the centre of gravity; but it will, however, cause her to drift more than the mainsail; so that this mode of lying-to is a mean between the two others, and is preferable when it blows strong enough for that sail to support the rolling of the ship. It ought likewise to be preferred, because the ship will veer under that sail, the action of which passes at a small distance from the centre of gravity, and the power of the sail overcomes the resistance which all ships meet from the fluid under their lee; a resistance which always gives them a great inclination to fly up in the wind when it blows hard, or when under a heavy press of sail.

Lying-to under the fore, main, and mizen staysails.

All the preceding modes of lying-to have their peculiar faults; but the preferable way is under the fore staysail, the main staysail, and mizen staysail; because under these sails the ship will steer, and is in a better situation for veering than under any other sail; for only haul down the mizen staysail and put the helm a-weather, when the two other sails, being before the centre of gravity, will cause her to fall off; she will then soon

gather way, and steer easily.

Should the gale continue very hard, and one of those staysails be blown away, the loss is not of much consequence, as the courses, in case of an emergency, are ready to set; whereas the courses are not so readily replaced when lost. This node therefore appears preferable in every respect, whether you wish to veer or keep your wind; because if the ship does not sufficiently keep the wind, you may hanl out the belanced mizen, or take in the fore-staysail, or even the main staysail. One of these staysails, before the centre of gravity of the ship, is sufficient to make her veer as soon as the after ones are suppressed. There are besides, these following considerations for so doing: the ship will carry sail better; because, as the centre of effort of those on her is very low; she drifts less, holds a better wind, and goes faster through the water; and these three or four sails are so situated as to give the whole body of the ship play, which will strain her less than when under one single sail, which cannot by itself work it from aft forward.

Of sounding in fair weather, whether close-hauled, or going large.

Close-hauled.

If close-hauled, brail up the mizen and mizen-staysail, let go the main sheet that the sail may shiver, put the helm s-lee, and back the mizen-topsail by bracing it square. The head-sails as well as the jib and staysails, are to be keept in their first situation; recollecting to haul tight and belay the lee braces. When the ship has nearly lost her headway, though continuing still to come to the wind, yet catch that moment to heave the lead, and it is to be hauled in again with all possible despatch. To fill again, haul aft the main-sheet, trim the mizen topsail, and right the helm.

Going large.

In going large you have only to put the helm a-lee, to brail up the mizen, and being the lee braces quite tight, to prevent the yards having too much play when the sails are

^{*}Should the sea run too high for the lower stayedle to hosp the ship steady, a close-rected maintopsall will be found to answer the purpose admirably.

shivering. It is impossible to tack in this situation, as the jib and head-sails are always in action: and the square-sails soon coming to shake, on account of their sheets not being tacked, they lose all their power, and the ship is soon at a stand.

Another method preferable to the former.

Going large.

Brace the headsalls square, haul down the jib and staysails, without stirring the after sails, and put the helm a-lee. While the ship has still a little headway, heave the lead from the place where you haul it in: that lead will go first a little a-stern, but the ship being head to wind, will soon herself go a stern right upon the line; and as the helm is a-lee, the ship easily veers. But, if you wish to keep her to longer, right the helm and haul the mizen out, to prevent the ship's falling off.

If you have studding sails set, they must be hauled down, particularly the lower ones;

because, should the wind take them aback, their power on the boom might bring the ahip round entirely, for they act on a lever without the ship, the fulcrum of which is on the outside of the vessel before the centre of gravity. If, however, the helm is continued a-lee till the ship falls off, she will not come about, because then the vessel goes a-stern with great velocity, and the rudder acts powerfully to make her veer; but the fact is, that the ship will go a great deal a-stern, and will continue to do so much longer.

Close-hauled.

If close hanled, or a very little from the wind, the helm is to be put a-lee, and the instant the sails are taken a-back the headsails are to be filled by briskly bracing them square, without waiting for the wind being right a-head; then a little before the ship has lost her way, heave the lead from the place where you haul it in, and then proceed as before,

On ship's driving.

When it happens that there is not sufficient room to work in a tide's way, through a crowd of ships, or in a narrow channel, but that the ship must drive by the help of the tide, it may be done, provided the tide be strong enough in proportion to the wind. This art consists in keeping the ship in a fair way, by a management of the rudder and the sails.

To drive to windward, when the wind is against the tide.

If the channel is sufficiently broad, the ship should be drifted broadside to the wind, as the tide will then have the greatest power on her; and could the ship be backed a-stern or shot a-head at pleasure, she might be kept drifting upon the same tack with safety; but ships in a tide's way can never be backed so far a-stern as they will shoot a-head. At the first of a stern-board, a ship will go briskly a-stern, but will soon fall off, and drift with the wind abast the beam, forging a-head; for this reason she must be drifted with the helm a-lee. It follows, as a ship will shoot more a-head than she can be backed a-stern, that she will at length arrive at the opposite shore, when she must be stayed or veered and drifted upon the other tack. If she is to be stayed, (which is preferable, because less drift will be lost by it) let the sails be filled in time to give the ship sufficient headway to bring her about, then put the helm s-lee. Should she come but, the sails and helm having now a proper position for a sternboard upon the other tack, need not be touched till her sternway ceases, when the helm must be shifted a-lee: but should the ship refuse stays, then brace sharp round the headyards, and boxhaul her, by which method she will lose much less drift than by veering

If the ship now drifting broadside, is approaching a narrow channel, where drifting in this position, she must be veered and dropped, stemming the tide stern foremost. In this case, that the drift may be as much as possible, it will be necessary to take in sail, and reduce the ship's headway till she has only steerage way left; thus a vessel may be

dropped through a fleet of ships at anchor without danger.

To drive when the wind is across the tide.

Should the wind be a little scross the tide, a ship may be easily drifted in the fair way, with her head towards the weather shore; for thus it will be found that she can be backed and filled at pleasure, and generally be drifted with the sails shivering, in which position they oppose least power to prevent the drift.

It frequently happens, in serpentine rivers, that the tide sets across; in this case the ship must be drifted with her head to the side from which the tide sets. These sets are best discovered by observing the opening or shutting of two objects in the direction of the channel.

To bend a course in fair weather.

Stretch the sail a-thwart the deck, the starbeard side of the sail to the starboard side,

larboard to the larboard side; then bend yard ropes to the ear-ring cringles, and make the head ear-rings a few feet up upon the yard-ropes. The bunt-lines, leech-lines, a garnets, and all the geer bent, make fast a rope-band to each bunt-line and leech-line; leg, that the men may be enabled to catch the head of the sail from the yard. Now n well the yard ropes, bunt-lines, leech lines, and clue-garnets, and run the sail up to yard. The sail aloft, send the hands up to bring it to, and let them haul out the ather ear-ring first, then the lee; and, if it is a new sail, let them ride the head rope stretch it. The sail being hauled square out upon the yard, make fast the rope-bands, ping the head of the sail well upon the yard.

To bend a topsail in fair weather.

Overhaul the leeches of the sail, put in the ear-rings, bend the bow-line legs, lay out clues, and open them if necessary, and make the sail up snug again; then round wn upon the lee-topsail-haliards till the weather fly-block is high enough to bring the lup over the guard-iron: then rack the tie over the weather rigging. Now pile the lupon slings, with the lee-side uppermost: hook on the topsail haliards, and run the sail up into the top: then stretch the sail round the fore-part of the top, bend the r, and make fast the head ear-rings a few feet up upon the reef-tackle-pendants, with ope-band or two to each bunt-line leg. The jeer being bent, man the reeve-tackles, it-lines and clue-lines, and haul out the sail. Now let the hands lay out upon the d, and haul out the weather ear-rings first; then haul out to leeward, and ease off to idward till the sail is square, when make fast the rope-bands, keeping the head of the lwell up upon the vard.

To set a mainsail or foresail.

Before the sail is loosed, let the double block of a tackle be made fast to the weathere, and the single block be tooked low down upon the chess-tree, and the fall led aften man well the tack and fall at the same time; and when the sail is loosed, he away the weather-clue garnet let go the bunt-lines and leech-lines, bowse down on the tackle, and take in the main-tack; the main-tack being down, haul aft the set, brace up the yard, and haul the main-bowline.

To set a topsail.

Let a tackle be in readiness to clap on either sheet as may be required. First man bee-sheet; and, the sail being loosed, ease down the bunt-lines and lee clue-line, and all home the lee sheet; then haul home the weather sheet, hoist the sail, and brace up required.

Should the wind be quartering, the lower and topsuil yards should be braced well into wind, before the sail is sheeted home.

To take in a course.

Man well the weather clue-garnet, ease off the tack and bowline, and run it up; then, in the lee clue-garnet, bunt-lines, leech-lines, and weather braces; and being all ready, se away the sheet, haul up the clue-garnet, bunt-lines, and leech-lines, and round in weather-brace, till the yard is pointed to the wind. Then haul tight the trusses, braces, is and rolling tackle, and let the hands furl the sail.

To take in the foresail in the time of veering.

When the ship begins to veer, the yard being kept braced sharp up, let go the tack and wline, and haul up the weather clue-garnet. When the ship is nearly before the wind, bunt and leech-lines, and the other clue-garnet may be hauled up; and if the situanamits of it, and occasion requires, the ship may be steered with the wind on the arter, till the sail is secured.

To take in a topsail.

There are many opinions upon the best mode of performing this. Some approve of ting up to windward first, and others to leeward. If the weather-side is to be clued up st, the weather brace must be rounded well in, and the yard got close down upon the is, otherwise the lee rigging will be in danger of being carried away by the great presser of the lee yard-arm. If the weather brace can be rounded well in, and the yard be t close down, it will be best to clue up to windward first, for thus the sail may be taken without a shake, but, if the weather-brace cannot be hauled in to ease the yard off the rigging, recourse must be had to cluing up to leeward first. In this case, it will be it, if hands can be spared, to man both the clue-lines, bunt-lines, and weather-brace, the same time; thus, when the lee sheet is eased off, the weather-brace may be hauled with ease, and the yard laid to the wind; and, when the lee clue-line is half up, ease the weather-sheet, and run up the weather clue-line; then haul tight the lee brace, we tight the rolling tackle, and furl the sail.

To take in a jib.

Man well the down haul, let go the haliards, ease off the sheet, and haul down briskly; and, when the sail is close down, ease away the out-haul, and haul the sail into the bow-sprit cap; then let it be stowed away in the fore-staysail netting.

To hand in a lorder studding-sail.

To hand in a lower studding-sail, blowing fresh, lead one of the sheets clear aft, and man it well: then lower away briskly the outer haliards, to spill the sail; ease off the tack, run in upon the sheet, and lower away the inner haliards as required.

To have down a topmast studding-sail.

Man well the deck sheet and down haul, ease off the haliards, and haul the yard close out to the tack block; then ease away the tack, and haul down both upon the deck sheet and down haul.

To brail up and haul down a main-topmast-staysail.

Man well the lee brail and down haul, having a few hands to gather in the slack of the weather brail; then let go the hallards, ease off the sheet, and haul down and brail up as briskly as possible. When the sail is down, let go the tack, and stop the sail over to the lee fore rigging.

To brail up a mizen.

Man well the lee brails, ease off the mizen sheet, and brail up briskly, taking in at the same time the slack of the weather brails. After the sail is hauled up, stop its foot by passing the gasket round to leeward which will spill it.

To take in top-gallant-sail.

The lee sheet must be started first and clued up, and then the weather sheet.

To unbend a course.

First furl the sail, then cast off the rope-bands and make them fast round the sail, clear off the gaskets. When the rope-bands are all off, ease off the lee ear-ring, and lower down the sail; and, when the people upon deck have got hold of the lee part of the sail, ease away the weather ear ring.

Taunbend a top-sail.

First cast off the points of the reefs, keeping fast the ear rings; then furl the sail and east off the rope-bands, which make fast round the sail, clear off the gaskets. After this cast off the lee ear-rings and haul the lee side of the sail into the top; then haul in the weather side. Now unbend the reef-tackle, pendants, bunt lines, and bow lines; bight the sail snugly up together; and send it down by the clue-lines to windward or to lec-ward, as most convenient.

On scudding or bearing away in a storm.

When the waves run high, and sudden necessity requires to bear away, it should be considered that the lower sails forward, which the ship may be veered under when she comes before the wind, may be becalmed by the height of the waves breaking violently against the stern; and that therefore a close-reefed maintop-sail should be set to catch the wind, because it is a loftier sail, and may always be kept drawing full above the waves. This increases the ship's headway so much that the waves will not strike her abaft with so great a velocity as when her beadway is less.

Hence it follows, that when going to scud before high waves, the close-reefed-main-

topsail should be the last square sail taken in in a laboursome ship.

Of a ship overset on her side.

A common, but not always'a certain method to recover ships from this dangerous situation, is to cat away the masts: however, as this expensive method may fail, stop-waters only, on the lee quarter at sea, may cause the ship to veer; or, where there is ground, an anchor or anchors dropped from the lee bow, may bring the wind a-head and take the sails a-back, so as to east the ship on the other tack, and bring her upright.

To rig a main-topmast.

Tar the mast-head, get the cross-trees over, fix the bolsters and parcel them, put over burion pendants, then the shrouds, breast-back-stay, proper and spring-stay and cap, sway up the mast and fid it, seize in the dead eyes, stay the mast, set up the shrouds, rattle them down, lash the bullock blocks to the mast head.

To rig a topgallant mast,

Send down the top-rope, reeve it through the sheeve-hole, and make it fast round the hounds of the mast and standing part of the rope, leaving enough end to make fast to the cap, which done, sway-away, when the head is through the cap, make fast the spare end, or standing part of the top-rope to the cap, cut the seizing, clap on the grommet, then the shrouds, back stays and stay, sway up the mast, fid it, and set the rigging up.

To rig a boxesprit.

Lash the collar fore-stay for the bob-stays and bowsprit shrouds, then the collar for the spring stays, then the block for the topmast stay, fix the man-rope, gammon the bowsprit, and set bob-stays and shrouds up.

To rig a jib-boom.

Put over the traveller, horses, guys, the topgallant stay-block, and lash on the blocks, for the topgallant bowline and jib down-haul block to the traveller.

To rig a lower yard.

Get it athwart the gunwale, lash the jeers, quarter clue-garnets, bunt-lines, leechlines and slab-line blocks; then put over the yard arms, the horses, brace pendants, the yard-tackle pendants, then the top-sail sheet and lift-blocks, reeve the jeers, braces, lifts and yard tackle falls, truss parcels, sway the yard up, and haul all taut.

To rig a fore-topsail yard.

Reeve a top-rope through the bullock-block and send it down, and having put over the horses, make the top-rope fast to the middle of the yard, stopping it to the yard-arm, sway it up above the top, put over the brace pendants and lift blocks, reeve the lifts and braces, cut the yard-arm seizing and cross the yard, lash the tye, bunt-line and clue-line blocks, reeve the tye and haliards, sway it up above the cap, and parcel it, reeve the clue-lines, bunt-lines and reef-tackles.

To rig a topgallant-yerd.

Seize the clue-line blocks on, put the horses over the yard-arms, sway it upon the cap and rig the yard-arms, by putting on the brace-pendants and lifts, then cross the yard and parcel it.

To steer a ship when her rudder is lost.

To take a large spar, or part of a topmast, and cut it flat in the form of a stern-post, bore holes at proper distances in that part which is to be the fore part of the preventer or additional stern-post, then take the thickest plank on board, and make it as near aspossible into the form of a rudder, bore holes at proper distances in the fore-part of it, and in the after part of the preventer stern post to correspond with each other: and reeve rope grammots through those holes in the rudder, and after-part of the stern post for the rudder to play upon.

Through the preventer stern-post reeve guys, and at the fore-part of them fix tackles, and then put the machine overboard; when it is in a proper position, or in a line with the ship's stern-post, lash the upper part of the preventer post to the upper part of the ship's stern-post, then hook tackles at or near the main chains and bowse taut on the guys to confine it to the lower part of the preventer stern-post:—having holes bored through the preventer, and proper stern-post, run an iron bolt through both, taking care not to touch the rudder, which will prevent the false stern-post from rising up or falling down.

By the guys on the after part of the rudder, and tackles affixed to them, the ship may be steered, taking care to bowse taut the tackles on the preventer stern-post to keep it close to the proper stern-post.

CATALOGUE

TABLES, WITH EXAMPLES OF THE USES OF THOSE THAT ARE NOT EXPLAINED IN OTHER PARTS OF THIS WORK.

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TABLES I. and II. Difference of Latitude and Departure.-The first table contains the difference of latitude and departure corresponding to distances not exceeding 300, and for courses to every quarter-point of the compass. Table II. is of the same nature and extent, but for courses consisting of whole degrees. The manner of using these tables is particularly explained under the article of Inspection, in the different Problems of Plane, Middle Latitude, and Mercator's Sailing.

TABLE III. Meridional Parts.—An explanation of this table may be found in pages

77 and 79, and the uses of it are shown in all the Problems of Mercator's Sailing.

TABLE IV. The Sun's Declination.

TABLE IV. A. This table contains the equation of time for every noon at Greenwich and is to be reduced to any other hour by means of Table VI. A. Thus, suppose the equation of time was required for May 2, 1824, sea account at 10 A. M. apparent time, corresponding to May 1d. 22h. by the N. A. Table IV. A. gives the equation May 1, at noon, sub. 3m. 5s. and daily increase 8°. Find this at the top in Table VI. A. and 22h. at the side, the corresponding correction 7s. increases the equation 3m. 5s. to 3m. 1%s. which is the equation at the proposed time. Thus 7s. could have been subtractive if the equation had been decreasing, as it is in March. The equation of time being thus found, sub. 3m. 12s. is to be subtracted from the apparent time 32h. as in the table to get the mean time 21h. 56m. 48s. If the mean time 21h. 56m. 48s. had been given to find the apparent it must be applied differently from the direction in the table. and in this example must therefore be added to 21h. 56m. 48s. to obtain the apparent time 22h.

TABLE V. For reducing the Sun's Declination given for noon at Greenwich to any other time under any other meridian .- The manner of using the two preceding tables is

explained in pages 110 and 111.

TABLE VI. The Sun's Right Ascension.—The Sun's mean right ascension given in this table may be used when a Nautical Almanac cannot be procured, and no great ac-The Table is to be entered at the top with the month, and at the curacy is required. side with the day of the month.

TABLE VII. Amplitudes.—This table is explained in page 112.

TABLE VIII. Right Ascensions and Declinations of the principal fixed Stars.—
This table contains the right ascensions and declinations of the principal fixed stars, adapted to the 1st of January, 1820, and the annual variations in right ascension and declination; by means of which the right ascensions and declinations of any of these stars may be obtained for any time before or after the year 1820, by the rule at the end of the table. To illustrate the method of doing this, we shall here give the following examples.

To find the rig	ht asco	ms	ion of a star at any time.			
EXAMPLE I.			EXAMPLE II.			
Required the right ascension of Aldeba	ran. Ja	DU-	Required the right ascension of Aldebars	n. Jo	mn-	
ary 1, 18247	b. m	. s.	ary 1, 1800 ?	h. m		
R. A. by the Table in 1820	4 25	36	R. A. by the Table in 1820		5 36	
Variation to 4 years add		14	Variation in 20 years, subtract	1	1 9	
			•			
R. A. in January, 1824	4 25	50	R.A. on January 1, 1890	4 24	1 27	•
EXAMPLE III.			EXAMPLE IV.			
Required the right ascension of Spice	a, May	20,	Required the right ascension of Sirius	Nov	em-	
1826?			ber 6, 1807 ?	h. n		
R. A. by the Table in 1820	13 15	43	R. A. by the Table in 1820	6 57	7 18	
Variation in 6 years 43 months, add		20	Variation in 13 years subtract		34	
R. A. May 20, 1828	13 16	3	R. A. in January, 1807	6 30	6 5 9	
			Variation for 10 months and 6 days, add		2	
			R. A. November 6, 1807	6 36		
The sun's right ascension for any	time	ms	ry be found accurately by the Nautica	I Alı	ma-	

nac, by taking proportional parts of the daily difference, as will be explained in the precepts of Table XXXI. But in cases where no great accuracy is required, the right ascension may be obtained within 2 or 3 minutes, by means of Table VI.

To find the declination of a Star at any time.

EXAMPLE II. Required the declination of Aldebaran, January
Required the declination of Aldebaran, January
1824? Declination by the Table in 1920 Variation in 4 years 83" add nearly 180 8' N. Declination by the Table in 1820 Variation in 10 years 1'20' subtract Declination in 1824 16 9 N. Declination January 1, 1810

		EXAMPLE IV. Required the declination of Sirks, No.		6,
Declination by the Table in 1820 Variation in 6 years 43 months	10° 13′ B.	Declination by the Table in 1880 Var. in 22 years 1 month 24 days, is sub-	160 29 / 2	8.
Declination May 20, 1826	-	Declination November 6, 1797	100 27	8.

The right ascensions and declinations obtained by the preceding calculations, are the mean values, to which must be applied the corrections for the Nutation and Aberration Tables XLII. XLIII. in cases where great accuracy is required, as is now done in the Nautical Almanac for 24 of the brightest stars for every 10 days in the year, and those numbers in the Nautical Almanac are to be preferred.

To find when a star will be on the meridian.

RULE. Find the right ascension of the sun and star in the preceding tables VI. and VIII; subtract the sun's right ascension from the star's, having previously increased the latter by 24 hours when the sun's right ascension is the greatest; the remainder will be the time of the star's coming to the meridian. If the remainder be greater than 12 hours, the star will come to the meridian after midnight; but if less than 12 hours, before midnight.

EXAMPLE I. At what time will Aldebaran be on the January 1? Aldebaran's right ascension Add	ne meridian. At what time will Pollax be on the merid h. m. March 31? 4 28 Pollux's right ascension 24 Sun's right ascension	Man, 1 ID. 7 34 38
Sun's right ascension	28 26 Comes to the meridian in the evening 18 45	6 56
Aldebaran souths in the evening	9 41	
EXAMPLE III. At what time will the star Regulus be ridian, Desember 12? Regulus' right ascension Add	h m. on the meridian, June 1? 9 59 fromathaut's right ascension	ones h. m. 12 48 4 35
Sun's right ascension		8 13
After midnight Subtract	16 42 In the morning	6 13
In the morning	4 49	

To find what star will come upon the meridian at any given time.

RULE. Add the time from noon* to the right ascension of the sun, the sum (rejecting 24 hours when it exceeds 24) will be the right ascension of the star required to be known; with which enter the table of the star's right ascension, and find what star's right ascension agrees with, or comes the nearest to it, and that will be the star required, if the dec... tion of the star agrees with the table, which may be ascertained by observing the meridian altitude of the star, the latitude of the place being given.

EXAMPLE I.

EXAMPLE II.

	about 10 at What star will be upon the meridi	an 30 minotes
night, January 26 ?	h. m. past four in the morning, May 10?	
Sun's right ascension January 28	20 53	b, m.
Given time 10 hours P. M.	10 Such right ascension May 10	8 7
	- Give time 16 hours 30 minutes	16 30
	80 33	
Subtract	24 Right as easion of mid. heaven	19 37
	Newers nearly to Atair in the Engle	
Nearly answers to Sirius	6 SS	
EXAMPLE III.	EXAMPLE IV.	
	at 6h. 53m. What tar will be on the meridian	. September 1.
P. M. April 1?	h. m. at oh. 37m. P. M 3	h. m.
Sun's right ascension April 1	42 Sun's right ascension Sept. 1	10 41
Given time	6 53 Given time	5 57

Right ascension of the meridian
Answers nearly to Pollux.

In all the preceding examples, the right ascension of the sun ought to have been calculated for the moment of the star's passing the meridian, as will be more fully explained in the precepts of Table XXXI.

^{*} The time from noon must be reckoned, from the preceding noon, so that th. A. M. must be called 16h.

Semi-diurnal and Semi-nocturnal arches. - This table exhibits half TABLE IX. the time that a colestial object continues above the horizon when the latitude and declimation are of the same name, or below when they are of a contrary name; the former time being usually called the semi-diurnal arch, the latter the semi-nocturnal arch; whence the time of rising and setting may be computed, by the following rules.

To find the time of the sum's rising and setting, and the length of the day and night,

RULE. Find the sun's declination at the top of the page and the latitude in either side column, under the former, and opposite the latter, will be the time of the sun's acting if the latitude and declination are of the same name, but the time of rising if of different names.—The time of rising, subtracted from 12 hours, will give the time of setting; or the time of setting, subtracted from 12 hours, will give the time of rising.—The time of rising, being doubled, will give the length of the night; and the time of sets ting, being doubled, will give the length of the day.

EXAMPLE I.

Let it be required to find the time of the sun's rising and setting, with the length of the day and night in latitude 51° north, the 19th of July, 1820?

The sun's declination on the given day was 20° 51' north, or 21° nearly, under which, and against the latitude 51°, stand 7h. 53m. the time of the sun's setting on the given day, in lat. 51° north, which doubled, gives 15h. 46m. the length of the day; and by subtracting 7h. 53m. from 12h. the remainder 4h. 7m. is the time of the sun's rising, which doubled gives 8h. 14m. the length of the night.

But, when the sun has 21° south declination in this latitude, the time of sun setting

becomes 4h. 7m. the time of rising 7h. 53m. the length of the day 8h. 14m. and the length of the night 15h. 46m. as was the case nearly on the 26th November, 1820.

EXAMPLE III.

Let it be required to find the time of the sun's Required the time of the sun's rising, setting, and the length of the day and hight, and length of day in laitted 510 29 S. May 16th, at Boston, the 12th of July, 1820;

[1820] Under 220, which is nearly the declination on that I had an include 120 55 W at 100 W. h. h. of the declination on that I had an include 120 55 W at 100 W. h. h. of the declination on that I had an include 120 55 W at 100 W. h. h. of the declination 120 55 W at 100 W. h. h. of the declination 120 55 W at 100 W. h. h. of the declination 120 55 W at 100 W. of the declination 120 55 W at 100 W. of the declinatio

CROSE, 32-, AMICU IS DESTLY FIRST OCCUPATION ON EDGE OF ORGAN MICUITALISM 15- 02. (I	L 79-A 74- (10-142-)
day, and against 42° 23' or 42° N. the latitude offend against the lat. 34° S.	\$ 12 Ø
Boston, stands the time of the sun's ? h. m. Stands the sun's rining	6 54
setting \$ 7.25	
Subtracted from 12h. leaves sun-rising 4 35/Time of sun's setting	5 8
Sun-setting doubled is the length of day 14 50	2
Suz-rising doubled is the length of night 9 10	
The length of the day	10 12
And 6h. 54m, doubled is length of	f night 13 48
,	

When a great degree of accuracy is required, proportional parts may be taken for the minutes of latitude and declination.

To find the time of rising and setting of stars whose declination does not exceed 93° 28'.

Enter Table IX. and find the star's declination at the top, and the latitude at the side; under the former, and opposite to the latter, will be the semi-diurnal arch, when the latitude and declination are both north or both south; but if one be north and the other south, the difference between the Tabular number and 12 hours will be the semidiurnal arch. Find the time of the star's coming to the meridian according to the precepts of Table VIII. and subtract therefrom the semi-diurnal arch, the difference will be the time of rising; or by adding together the semi-diurnal arch, and the time of passing the meridian, the time of setting will be obtained.

EXAMPLE IV.		EXAMPLE V.		
Required when the star Arcturus rises a	ads sets	What time will the Dog-Star Sirius rise a	nd set	
December 1, in intitude 51° N.?	h. m.	at Philadelphia, Feb. 1?	b. 10.	
The time of the star's coming to the meridi-		Under the declination, which is nearly 160 8.		
an or southing in the morning, is nearly	9 38		12 .0	
Then under star's declination 200 nearly, and		which is nearly 40° N. stand	6 56	
against latitude 51° stand	7 47			
-		Subtracted from 12h. leaves half the time the		
Time of star's rising in the morning	1 51		5 4	
		The star comes to the meridian in the even-		
Added, gives the time of the star's setting	17 25	ing nearly at	9 39	
	12	,		
		Sum, rejecting 12 hours, is the time of setting		
Star sets 26 minutes after 5 in the evening	5 25	in the morning	2 43	
		Difference is the time of rising in the evening	4 35	

In like manner may the rising and setting of any planet be found when the decli-nation does not exceed 23° 28', and the time of the passage over the meridian is known. Suppose it was required to find the time of Jupiter's rising and setting, March 8, 1820, civil account, in the latitude of 520 N?

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In the Nautical Almanac for 1820, I find that Jupiter passes the meridian March 7d. 23h.'10m. or March 8d. 11h. 10m. A. M. civil account, his declination being 10° 55' S. or nearly 110. Under the declination 110, and opposite to the latitude 520 stand 6h. 58m, which is half the time Jupiter is below the horizon; this subtracted from 12h. leaves half the time that he is above the horizon, 5h. 2m.; this subtracted from 11h. 10m. A. M. leaves 6h. 8m. A. M. March 8, for the time of Jupiter's rising; and added to 11h. 10m. gives 4h. 12m. P. M. March S, for the time of Jupiter's setting.

Suppose it was required to find the time of the moon's rising and setting, May 5, 1820,

civil account, in the latitude of 52° N?

In the Nautical'Almanac, page VI. I find that the moon passes the meridian May 4d. 18h. 7m. or May 5d. 6h. 7m. A. M. civil account; her declination being about 210 S. Under the declination 21°, and opposite to the latitude 52°, stand 7h. 58°, half the time the moon is below the horizon, which subtracted from 12h. leaves half the time she is above the horizon, 4h. 2m.; this subtracted from 6h. 7m. leaves 2h. 5m. A. M. the time of the moon's rising, and added to 6h. 7m. gives 10h. 9m. A. M. the time of

her setting, nearly.

If greater accuracy is required, you must find the time at Greenwich corresponding to this approximate time of her rising and setting; then find the moon's declination, and the right ascensions of the sun and moon for that moment of time. The former subtracted from the latter leaves the corrected time of the moon's passing the meridian. With these data repeat the operation. In this way we may obtain the time of rising and setting to any degree of accuracy. Instead of taking the difference of the right ascension of the sun and moon, you may take the daily difference in the time of her coming to the meridian of Greenwich, and take a proportional part for the longitude of the place of observation (by means of table XXVIII.) and another proportional part, for the interval between the hour of passing the meridian, and the time of rising or setting.

It may be noted, that the numbers of Table IX. were calculated for the moment the sun's centre appears in the true horizon; allowance ought to he made for the dip, parallax, and refraction, by which the sun and stars, when near the horizon, appear in general to be elevated above half a degree above their true place, and the moon as much

below her true place.

TABLE X. For finding the distance of any terrestrial object at sea.—The explanation

and use of this table is given in Problems VII. and VIII. pages 190, 191.

TABLE XI. Table of Proportional Parts.—The method of using this table is given in page 166.
TABLE XII.

Table of Refraction.—Explained in page 108. Dip of the Horizon.—Explained in page 109. TABLE XIII. TABLE XIV. Sun's Parallax in altitude.—Explained in page 107.

TABLE XV. Augmentation of the moon's semi-diameter. - The moon's semi-diameter given in the Nautical Almanac is the same as would be seen by a spectator supposed to be placed at the centre of the earth, or nearly the same as would be seen by a spectator on the surface of the earth, when the moon is in the horizon. Now when the moon is in the zenith of the spectator placed at the surface, her distance from him is less than when at the horizon by a semi-diameter of the earth; consequently her apparent semi-diameter must be augmented in proportion as the distance is decreased, that is about one sixtieth part, or 16". At intermediate altitudes, between the horizon and zenith, the augmentation is proportional to the sine of the altitude, and the value for every 5° or 10° of altitude is given in Table XV. The augmentation corresponding to the altitude being found in the table, must be added to the semi-diameter taken from the Nautical Almanac for the time of observation reduced to Greenwich time, as was explained in page 166.
TABLE XVI. Dip of the sea at different distances from the observer.—Explained in

page 109.

TABLE XVII. For finding the difference between the refraction of a star and 60';

also a log. corresponding.

TABLE XVIII. For finding the difference between the correction of the sun's altitude for parallax and refraction and 60', also a logarithm corresponding thereto.—The manner of taking the numbers from the two preceding tables is explained in page 167, and the uses to which these tables may be applied are explained in pages 167 and 174.

TABLE XIX. For finding a correction and logarithm used in the first method of working a honor observation.—The correction found in this table being subtracted from 59' 42" will leave a remainder equal to the correction of the moon's altitude for parallex will leave a remainder equal to the correction of the moon's altitude for parallax

^{*} In strictness, this last correction, found by the table, ought to be decreased in the ratio of 2th to 24h. increased by the daily difference of the time of the moon's passing the meridian. Digitized by GOOGLE

and refraction. It will be unnecessary here to point out the method of taking out this correction, as it is fully explained in the first pages of the table. It may not, however, be amiss to observe, that after constructing the logarithms of this table, it was concluded to subtract therefrom the greatest correction of the Table C corresponding, in order to render those corrections additive. Thus the logarithm corresponding to the alt. 30° and hor. par. 54', was found at first to be 2372; and for the hor. par. 54' 10'' the correction was 2358 so that if these numbers had been published, the correction for seconds of parallax would have been subtractive; but as this would have been inconvenient, it was thought expedient to subtract from each of the numbers thus calculated, the greatest corresponding correction of Table C, which in the preceding example is 12; by this means the above numbers were reduced to 2360 and 2346 respectively, and the corrections of Table C were rendered additive. In a similar manner the rest of the logarithms of the table were calculated. It is owing to this circumstance that the corrections in Table C for 0" of parallax are greater than for any other number. Similar methods were used in calculating the other numbers of this table, and in arranging the Tables A and B.

TABLE XX. Third correction of the apparent distance—The method of finding

TABLE XX. Third correction of the apparent distance—The method of finding the correction from this table is explained in pages 168, 174, 176.

TABLE XXI. To reduce longitude into time, and the contrary.—In the first column

of this table are contained degrees and minutes of longitude, in the second the corresponding hours and minutes, or minutes and seconds of time; the other columns are a continuation of the first and second respectively. The use of this table will evidently appear by a few examples.

EXAMPLE I.	- 4- 200 044 0		APLE II.	
Required the time corresponding		to 6h. 33m. 20s. ?	na minutes corre	shouging
Opposite 50° in col. 1 is		Opposite 6h. 32m. 0s. 1 20	in col. 4 is in col. 2 is	98° 0' 20
Sought time	8 22 4	6 33 20		98 20

TABLE XXII. Proportional Logarithms.—These logarithms are very useful in finding the apparent time at Greenwich corresponding to the true distance of the moon from the sun or star, as is explained in page 168. They may be also used like common logarithms, in working any proportion where the terms are given in degrees, minutes, and seconds; or in hours, minutes, and seconds, as in the examples page 177. The table is extended only to 3° or 3h. and if any of the terms of a given proportion exceed 3° or 3h. you may take all the terms one grade lower; that is, reckon degrees as minutes, minutes as seconds, &c. and work the proportion as before; observing to write down the answer one grade higher; that is, you must estimate minutes as degrees, seconds as minutes, &c. Instead of taking all the terms one grade lower, you may change two of the terms only, viz. one of the middle terms and one of the extreme terms; thus the 1st. and 3d. or the 1st. and 2d. may be taken one grade less, and the fourth term will be given correctly; but if the fourth term be taken one grade less, you must, after working the proportion, write it one grade higher, as is evident. To illustrate this we shall give the following examples.

	EXAMPLE I.	•	i		EXAMI	PLE II.	
: If in 15' 10	of time the sun rise	8 20 40' how	If d	he sun's	declination	n changes 16" bange in 8b. 21	19" in 34
	ise in S' 10" at the same ' Prop Log. ar. co.		nours,	now muc a the let	and 3d t	erms must be	taken one
As 15' 10' Is to 2° 40'		0512			400 00 0		
Se is 3' 10'		1.7547	As	24' 0"	P. L.	ar. co.	9.1249
				16' 19'' 8' 2''	P. L. P. L.		1.0496 1.3504
To 33' 24'	r Prop. Log.	.7815	So is	8 2	F. L.		
			To.	5' 28"	P. L.	•	1.5179
			٠.				

	1.
in the in the 90m 1	EXAMPLE IV. If in 16' the sun rises 3° 27' how much will it rise in 3' 10'' 3
Here all the terms must be taken one grad	Here the 2d and 4th terms must be taken one
less. As 12' 0' P. L. ar. co. 8.823 18 to 7 1 P. L. 1.409 Bo ls 4 20 F. L. 1.618	grade less. 9 Ås 16' 0' ar. co. P. L. 8,9488 115 to 3 27 P. L. 1,7175 5 So is 3 10 P. L. 1,7547
To 2' 32" 2" P. L. 1.851 Which taken one grade higher is 2° 32' 2" th	5 To 0' 41" P. L. 2.4210 Which taken one grade higher is 41', the answer

TABLES XXIII. For finding the latitude by two altitudes of the sum. - The manner of using this table is explained in page 139, et seq.

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TABLE XXIV. Natural Sines .- This table contains the natural sine and co-sine for every minute of the quadrant to the radius 100000, and is to be entered at the top or bottom with the degrees, and at the side marked M. with the minutes, the corresponding numbers will be the natural sine and co-sine respectively, observing that if the degrees are found at the top, the name sine, co-sine, and M. must also be found at the top, and the contrary if the degrees are found at the bottom. Thus 43366 is the natural sine of 25° 42' or the co-sine of 64° 18'.

TABLE XXV. Logarithmic sines, tangents, and secunts to every point and quarter oint of the compass.—This table is to be used instead of table XXVII. when the course is given in points. The course is to be found in the side column, and opposite thereto will be the log. sine, tangent, &c. The names being found at the top when the course is

less than 4 points, otherwise at the bottom.

Logarithms of Numbers.-The explanation and uses of this table TABLE XXVI. are given in page 28, et seq. TABLE XXVII. Logari

Logarithmic Sines, Tangents, and Seconts. - This table is explained

in page 33, et seq. TABLE XXVIII. For reducing the time of the Moon's passage over the meridian of Greenwich, to the time of her passage over any other meridian.—The manner of doing this is explained in page 124.

TABLE XXIX. Correction of the Moon's altitude for parallax and refraction.—The mean correction of the Moon's altitude is given in this table for every degree of altitude

from 10° to 90°. The manner of using this table is explained in page 126.

TABLE XXX. For reducing the Moon's declination given in the Nautical Atmanac for noon and midnight at Greenwich, to any other time under any other meridian.—The manner of using this table is explained in page 125. In addition to which it may be observed that 12h, are marked both at the bottom of the left hand column and at the top of the right hand column; but this can cause no embarrassment, because when the time at Greenwich is 12h. the declination must be taken from the Nautical Almanac for mid-

night, without any correction.

For reducing the Sun's right ascension in time, as given in the Nau-TABLE XXXI. tical Almanac for noon at Greenwich, to any other time under any other meridian.—This table is useful in finding the Sun's right ascension at any time, by means of the right ascension given in the second page of the Nautical Almanac for noon at Greenwich. This table must be entered at the top with a daily variation of the sun's right ascension, and in the left hand column with the given time from noon, or in the right hand column with the longitude of the place; under the former, and opposite the latter, will stand a correction in minutes and seconds, to be applied to the sun's right ascension at noon at The correction found with the time from noon, is to be added in the afternoon, but subtracted in the forenoon; and the correction found with the longitude of the place, is to be added in west, but subtracted in east longitude.

Instead of finding the correction separately for the longitude of the place and the time from noon, you may find the whole correction at one entry, in the following manner: Turn the ship's longitude into time (by Tab. XXI.) and add it to the given time when in west longitude, but subtract the longitude when east; the sum or difference will be the time at Greenwich; find this time in the side column* and the daily variation at the top, corresponding to which will be the sought correction; which is to be added to the

sun's right ascension for the preceding noon at Greenwich.

24, 1820, sea account, in the longitude from Greenwich?	. 0	f 454 The	ייתוט פ	EXAMPLE II. Required the sun's right ascension a 24, 1920, sea account, in the longitude from Greenwich? June 24, sea account is June 23, by N. day at noon the sun's right asc. was Corr. Tab. XXXI. for 120° long, and daily variation, 4m. 9".4 sub.	• • •	(15° On 1	00 E.
Right Assession required	4	0	55	Right Ascen. required	.6	6	38
EXAMPLE III. Required the sun's right secension, M at th. P. M. sea account, in long, of 450 B. A. at noen in long, 450 W. by Ex. I. Corr. in Tab. XXXI. for th. P. M. add	W	77 '	. 55s.	at 9h. A. M. sea account, in long. of 120	E		
R. A. May 24, 1820, at 4h. P. M.	4	1	35	R. A. June 24, 1808, at 9h. A. M.	6	10	17

^{*} If the time at Greenwich be more than 12h, you must first take out the correction for 12h, and then for the rest of the time; the sum of these two will be the correction.

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The third and fourth examples may be worked by a single entry of Table XXXI. as follows.

EXAMPLE III.					EXAMPLE 1V			•	
Given time by N. & May Long. 45° in time add		ь. 4 3		•	Given time by N. A. June Long. 120° in time	23		m. 0	
Time at Greenwich	23	7	ě	. 10. s	Time at Greenwich	23	13	0	m. s.
Sun's R. A. May 23, at noon by N. A. Corr. Tab. XXXI. for 7h.			4	0 2	Sun's R. A. June 23, at noon Corr. Tab. XXXI. for 12h. for 1h.			6	\$ 1 2 5
Sun's R. A. at 4h. P. M. Biffering 1s. from the former metho	d.		4	1 36				-6	10 16

If you wish to find accurately the time that any star comes to the meridian, or the time of rising or setting, you must take the sun's right ascension for noon at Green-wich, from the Nautical Almanac; then the star's right ascension from Table VIII. and with these, find the approximate time of rising, setting, or coming to the meridian, by the method already given in the precepts for using Tables VIII. and IX. Then calculate the sun's right ascension for this approximate time, and repeat the operation till the assumed and calculated times agree, and you will have the true time required.

To explain this method, I shall give the following examples.

To find the time when a star comes to the meridian.

a place in the longitude of 79° 50′ W. Jan. 2, 1820, sea account?	EXAMPLE II. At what time was Pollux on the meridian of a place in the longitude of 70° 46′ W. March 31, 1880, sea account?
which day the sun's R. A. at noon at h. m. s.	March 31, sea account, is March 30, N. A. on which day, at noon, the sun's right assension was This, subtracted from R. A. of Pollux 7 34 17
28 25 36 Difference is the approximate time 9 41 48 Now calculating the sun's R. A. for this time is	
	The sum of these two corrections is 1 47 which subtracted from the approximate time of southing 6h. 58m. 56s. leaves the true time 6h.

The method (used in the last example) of applying the corrections to the approximate time, instead of applying them to the right ascension of the sun, will be found the most expeditions; but it must be noted, that the corrections to be applied to the approximate time must have a contrary sign to what they would have when applied to the right ascendon.

To find the time of rising or setting of a star.

RULE. Enter Table IX. with the declination of the star at the top, and the latitude of the place at the side; the corresponding number will be the time of the star's continuance above the horizon, when the latitude and declination are of the same name; but if they are of different names, the tabular number subtracted from 12h. will be the time of continuance above the horizon. Add this time to the star's right ascension, if you wish to find the time of setting; but subtract the former from the latter if you wish the time of rising. From this sum or difference subtract the sun's right ascension* corrected for the longitude of the place; the remainder will be the approximate time sought.† Enter Table XXXI. with the distance of this approximate time from noon, and the daily variation of the sun's right ascension: the correction corresponding is to be added to the approximate time in the forenoon, but subtracted in the afternoon, and you will have the corrected time of rising and setting.

[•] Increasing the number from which the subtraction is to be made, by 24 hours, when necessary, † Rejecting 24 hours when it exceeds 24 hours. If the time of rising or setting be more than 12h. it will be after midnight; but if less than 12h. it will be before midnight.

1329, sea account, in the latitude of 38° the longitude of 77° W? The star's declination was 16° 8′ N. at	53′l ndth in 6h	N. and ne lati	EXAMPLE II. At what time did the Dog-Star Sirius i latitude 39° 20′ N. and the longitude of ′ Jan. 2, 1830, sea account? The star's 'declination is 16° 28′ S. and the is 39° 20′ N. corresponding to which in is nearly. Which subtracted from	160 l te la Tah	titude le IX. 1. 56m.
Sum May 24, sea acc. or May 23 by N. A. at noon sum's R. A 4h. 0m. Corr. for long. 770 W. 1	11	20	Leaves the time of the star's being above the horizon Subtract from star's R. A. Remainder	5 6	4 87 33
Sum subtract	4,	. 1	Add	24	
Remains approximate time of setting Gorr. in Tab. XXXI. for 7h. 19m. sub. Corrected time of setting, P. M.	_	19	Sum Jan. 2, sea acc. or Jan. 1, by N. A. at noon sua's R. A. Corr. for long. 75° 50' W. 18h.	25	33
Confessed state of sessing) 2. see	Ī		Subtract the sum	18	45
		•	Remains approxim, time of rising Corr. in Tab. XXXI, for 6h. 48m. sub.	6	48
			Corr. time of rising in the afternoon	6	47

Variation of the sun's altitude in one minute from noon. TABLE XXXII. TABLE XXXIII. To reduce the numbers of Table XXXII. to other given intervals of time from noon.

The method of using the two preceding tables is explained in page 150 and 151.

TABLE XXXIV. Errors arising from a deviation of 1' in the surfaces of the central wirror.—This table shows the error arising in measuring an angle by an instrument of reflection from a deviation of 1' in the parallelism of the surfaces of the central mirror, the line of intersection of those surfaces (produced if necessary) being perpendicular to the plane of the instrument. If the line of intersection be inclined to that plane, the numbers in the table must, in general, be decreased in proportion to the sine of the angle of inclination.

The second, third, and fourth columns of the table are calculated upon the supposition that the surface of the horizon mirror is inclined 80° to the axis of the telescope, or that the angle intercepted between the ray incident on the horizon glass and the corresponding reflected ray passing through the telescope is 20°, which is the case in circular instruments of De Borda's construction, and on this supposition the errors of an instrument in measuring different angles may be ascertained by the rules in pages 98 and 106; when the intercepted angle is greater or less than 20°, which is the case in most sextants and quadrants, the error in any measured angle corresponding to an inclination of the surfaces of 1', may be obtained as follows.

Find in the first column the intercepted angle, and the sum of that angle and the observed distance; take the corresponding corrections from column 5th, and their difference will be the sought correction.

In a circular instrument you must find in the side column the sum and the différence of the intercepted angle and observed angle, and take out the corresponding corrections from column 5th, half their difference will be the sought correction. Having thus found the correction corresponding to 1', you may find the correction for other angles as in pages 98 and 106.

TABLE XXXV. Correction for a deviation of the telescope of an instrument of reflection from the parallelism to the plane of the instrument.—The uses of this table are

explained in pages 97 and 105.

TABLE XXXVI. Correction of the mean refraction for various heights of the barome-

ter and thermometer.—The use of this table is explained in page 108.

TABLE XXXVII. Latitudes and Longitudes of the fixed Stars.—This table contains the Latitudes and Longitudes of the principal fixed stars, adapted to the beginning of the year 1820, with the annual variations for precession and the secular equation, by which the mean values at any time may be obtained, in like manner as the Right Ascensions and declinations are from Table VIII.; by adding the correction of longitude after 1820, subtracting before 1820, and applying the correction of latitude with the same sign as in the table after 1820, but with a contrary sign before 1820.

EXAMPLE I. Required the Longitude and Latitude of a Pegasi, July 16, 1818?

11s. 200 58' 44" Latitude by Table XXXVII. 1' 13" Variation 1 year 5\frac{1}{2}m. sub. Long. by Table XXXVII. Variation 1 year 54m. sub.

Long. July 16, 1818 11 20 57 31 19 24 44 N.

EXAMPLE II. Required the Longitude and Latitude of a Pegasi, July 1, 1822? ong. by Table XXXVII. 11s. 20° 58′ 44″ Latitude by Table XXXVII. 2 5 Variation 2} years, add Variation 21 years, add 11 21 0 49 Latitude July 1, 1822 19 24 44 N Long. July 1, 1822

When great accu-The latitudes and longitudes, thus obtained, are the mean values. racy is required, the corrections for the equation of the equinoxes, Table XL. and aberration, Table XLI. must be applied.

TABLE XXXVIII. Reduction of latitude and horizontal parallax.-This table contains the corrections to be subtracted from the latitude of the place of observation, and from the horizontal parallax of the Moon, given in the Nautical Almanac, in calculating eclipses of the Sun or occultations. Thus, if the latitude of the place was 40° , and the Moon's horizontal parallax 57', the correction of latitude would be nearly—11' 15'', and that of parallax—4".7, so that the reduced latitude would be 39° 48' 42", and the reduced parallax 56' 55".3. These values are to be used in occultations, but in eclipses of the Sun, this parallax is to be further decreased by 8".8 for the Sun's parallax. When the latitude is not given exactly in the table, the two nearest numbers must be found, and a proportional part of their difference is to be applied to one of the numbers, as usual. In calculating this table, the ellipticity of the earth was supposed equal to $\frac{1}{355}$, as in the third edition of La Lande's Astronomy, and in Vince's Astronomy. This value the third edition of La Lande's Astronomy, and in Vince's Astronomy. This value differs but little from $\frac{1}{364}$, and $\frac{1}{365}$, deduced by La Place from two lunar equations in the third volume of his immortal work, La Mecanique Celeste. In the second volume of the same work he calculated the ellipticity to be $\frac{1}{236}$ from the lengths of pendulums observed in different latitudes, this calculation corrected for a small mistake in the numerical coefficient of y in the tenth of his equations A'' becomes $\frac{1}{3+1}$ which does not differ very much from the value assumed in this table.

TABLE XXXIX. Aberration of the Planets.—This table contains the aberration of the planets, to be applied to the true longitude or latitude, with the same sign as in the table. The argument at the side is the elongation of the planet from the Sun; that is, the difference of their geocentric longitudes, or its supplement to 360°. Thus, on July 19, 1820, the longitude of the Sun was 3s. 26° 38', the Geo. long. of Venus 4s. 13° 23', their difference 16° 45' is the elongation or distance from the inferior conjunction, corresponding to which is the aberration + 3" to be applied to the true longitude given by the tables to obtain the apparent longitude. The aberration of Mercury is given at its greatest, least, and mean distances from the Sun. At the intermediate places, a propor-

tional part of the differences of the nearest tabular numbers must be applied.

TABLES XL. & XLI. Equation of the Equinoxes and Aberration in Longitude. Table XL contains the equation of the equinoxes in longitude common to all the heavenly bodies. The argument is the longitude of the Moon's ascending node, given in page III. of the Nautical Almanac, the signs of longitude being found at the top or bottom, and the degrees at the side, the corresponding number with its sign is the equation

of the equinoxes in longitude.

Table XLI. contains the aberration of the stars in longitude and latitude, to be calculated by the rules at the bottom of the tables. The signs of the argument being found at the top, and the degrees at the side, * taking proportional parts for minutes. The corrections of longitude found in these tables, are to be applied, with their signs, to the mean longitude found in Table XXXVII. and the correction of latitude, Table XLL is to be applied to the mean latitude deduced from Table XXXVII. Thus on July 16, 1820, by the examples at the bottom of Tables XL. XLI. the equation of the equinoxes was +1''.2 and the aberration in longitude +11''.5, these corrections being applied to the mean longitude of the star deduced from table XXXVII. 11s. 20° 59′ 11″ gives its apparent longitude 11s. 20° 59′ 24″. In a similar manner the aberration in latitude -5''.6, found at the bottom of Table XLI. applied to the mean latitude 19° 24′ 44″ N. deduced from Table XXXVII. gives the apparent latitude of the star 19° 24' 38" N.

^{*} The degrees in this and the following tables are to be found in the column marked D on the same horizontal line with the signs. Thus, if the signs are at the top of the table, the degrees must be found on the left column, otherwise in the right.

TABLES XLII. XLIII. Aberration and Nutation in Right Ascension and Declination. Table XLII. contains the aberration, and Table XLIII. the Nutation in Right Ascension and Declination, to be found by the rules at the bottom of the tables, and applied with their signs to the mean values deduced from Table VIII. Thus by Table VIII. the Right Ascension of a Pegasi, July 16. 1820, was 22h. 55' 49".6, and its declination 14° 14' 10" N. The aberration of Right Ascension in time was nearly + 6".8, in declination —0".7. The Nutation in Right Ascension in time —0".1, in declination —2".2, as appears by the examples at the bottom of the tables. These corrections being applied to the mean values, give the apparent Right Ascension 22h. 55' 50".3, and the apparent declination 14° 14' 7" N. The equation of the obliquity of the ecliptic may be calculated by the rule at the bottom of the table. Thus, on July 16, 1820, the equation was +9".5, which applied to the mean obliquity 23° 27' 48".2, gives the apparent obliquity 23° 27' 57".7.

TABLE XLIV. Augmentation of the Moon's Semi-diameter.—This table is divided into four parts, and is useful in finding the augmentation of the moon's semi-diameter by means of the altitude and longitude of the nonagesimal when the moon's altitude is unknown. The precepts for this calculation are given at the bottom of the table, and for further illustration another example is added, in which it is required to find the augmentation at the commencement of the occultation calculated in Problem VII. of the Appendix, when the D's S. D. by the Nautical Almanac was 16' 18".9, her true latitude 1° 36' 11" S. parallax in lat. 10' 23".6, altitude of the nonagesimal 81° 17' 32", and the moon's apparent distance from the nonagesimal 51° 38' 26", as in Example III. Prob. V. Appendix. In this case the arguments of Part I. are 81° 17' 32" + 51° 38' 26" or nearly 4s. 12° 56' and 0s. 29° 39', and the corresponding corrections + 6".00, +4".05, whose sum is 10".05. This in Part II. gives +0'.10. In Part III. with the moon's true latitude 1° 55' 11" S. and her par. in lat. 10' 23".6, the correction is —0".10. The sum of these three parts is + 10".05, which being found at the side of Part IV. and the moon's horizontal S. D. 16' 18'.9 at the top, gives the corresponding correction + 0".40. This connected with the three fermer parts + 10".05, gives the sought augmentation 10".45, or 10''.4, as in the example Prob. VII. Appendix. It may be observed that the calculation by Problem IV. will sometimes produce the supplement of the altitude of the nonagesimal, but this requires no alteration in the rule, since the result is the same whether the altitude or its supplement is used.

TABLE XLV. Equation of Second Differences.—This table contains the equation of the second differences of the moon's motion, or the correction to be made on account of her unequal velocity between the times marked in the Nautical Almanac. The manner of applying this correction is taught in Problems I. II. III. of the Appendix.

ner of applying this correction is taught in Problems I. II. III. of the Appendix.

TABLE XLVI. Tables of Latitudes and Longitudes.—This table (as observed in the preface) has been completely revised for this edition, and the latitudes and longitudes of a great number of places are added to those given in the former editions of this work.

TABLE XLVII. Tide Table.—The explanation and uses of this table are given in page 913, et seq.

TABLE I.

Difference of Latitude and Departure for 1 Point.

1_		·	N.ĮE.			N.ĮW.			8.1E.		•	S. į W.	
П	Dist.	Lat.	1)ep.	Dist.		Dep.		Lat.	Dep.	Dist.	Lat.	Dep.	
	• 1	01.0	90.0	61	60.9	03.0	_	120.9	05.9		180.8	03.9	241 240.7
П	3	02.0	00.1	62	61.9	03.0		121.9	06.0		181.8	08.9	42 241.7
l	4	03.0 04.0		63 64	62.9 63.9	03.1		122.9	06.0		182.8	09.0	43 242.7
П	5	05.0		. 65	64.9	03.1 03.2		123.9 124.8	06.1		183.3	09.0	44 213.7
П	6	06.0			65.9	03.2		125.8	06.1 06.2		184.8 165.8	09.1 09.1	45 244.7
H	7			67	6 6 .9	03.8		126.8	06.2		186.8	09.1	47 216.7
H	8. 9	06.0		68	67.9	03.8		127.8	06.3		187.3	09.2	48 247.7
П	10	09.0 10.6	00.4	69 70	68.9 69.9	03 4		i28.8	06.3		188.8	09.3	49 248.7
H	111	11.0	00.6	71	70.9	03.4		129.8	06.4	_	189.8	09.3	. 50 249.7
Ш	12	12.0		72	71.9	03.0 03.5		130.8 131.8	06.4		190.8	09.4	251 250.7
ı	13	13.0		73	72.9	03.6		132.8	06.5 06.5		191.8 1 92 8	09.4 09.5	52 251.7 53 252.7
П	14	14.0		74	73.9	03.6		133.8	06.6		193.8	09.5	54 253.7
Н	15	15.0	00.7	75	74.9	03.7	3 5	134.8	06.6		194.8	09.6	55 254.7
Į. į	16 17	16.0 17.0		76 77	75.9	03.7	36	135.8	06.7		195.8	09.6	56 255.7
H	18			78	76.9 77.9	03.8 03.8		136.8	06.7		196.8	09.7	57 256.7
t	19				78.9	03.9		137.8 138.8	06.8 06.8		197.8 198.8	09.7	58 257.7
li	20	20.0	01.Q	80	79.9	03.9		139.8	06.9		199.8	09.8 09.8	59 258.7 60 259.7
ı	21	21.0	01.0	81	80.9	04.0		140.8	06.9		200.8	09.9	261 260.7
H	22	22.0	01.1		81.9	04.0		141.8	07.0		2 01.8	09.9	62 261.7
П	23	23.0	01.1		82.9	04.1		142.8	07.0		202.8	10.0	63 262.7
Ŀ	24 25	24.0 25.0	01.2	84	83.9	04.1		143.8	07.1	04	203.8	10.0	64 263.7
П	26	26.0	01.2 01.3		84.9 85.9	04.2		144.8	07.1		204.8	10.1	65 264.7
П	27	27.0	01.3		86.9	04.2 04.3		145.8 146.8	07.2 07.2		205.8 206.8	10.1	66 265.7
H	23	28.0	01.4		87.9	04.3		147.8	07.3		200.8 207.7	10.2 10.2	67 266.7
!	29	29.0	01.4		88.9	04.4		148.8	07.3		208.7	10.2	68 267.7 69 263.7
1	30	30.0	01.5	90	89.9	04.4		149.8	07.4		209.7	10.3	70 269 .7
l	31	31.0	01.6	• 91	90.9	04.5	161	150.8	07.4	211	210.7	10.4	271 270.7
ı	32 33	32.0 33.0	01.6	92	91.9	04.5		151.8	07.5		211.7	10.4	72 271.7
H	34	34.0	01.6 01.7	93 94	92.9 93.9	04.6		152.8			212.7	10.5	73 272.7
H	35	35.0		95	94.9	04.6 04.7		15 3 .8 15 4 .8	07.6 07.6		213.7 214.7	10.5	74 273.7
П	36	36.0			95 9	04.7		155.8	07.7		214.7 215.7	10.5 10.6	75 274.7 76 275.7
П	57	37.0	01.8	97	96.9	04.8		156.8	07.7		216.7	10.6	77 276.7
H	38 39	38.0	01.9		97.9	04.8		157.8	07.8		217.7	10.7	78 277.7
П	40	39.0 40.0		99 100	9 8.9	04.9		158.8	07.8		218.7	10.7	79 278.7
! !	41	41.0		i		04.9		159.8	07.9		219.7	10.8	80 279.7
ı	42	41.9	02.0 02.1		100.9 101.9	05.0 05.0		160.8	07.9		220.7	10.8	281 280.7
1	43	42.9	02.1		102.9	05.1		161.8 162.8	07.9 08.0		221.7 222.7	10.9 10.9	
	44	43.9	02.2		103.9	95.1		163.8	08.0		223.7		83 282.7 84 283.7
ı	45	44.9		05	104.9	05.2	65	164.8	08.1	25	224.7	11.0	85 284.7
1	46 47	45.9	02.3		105.9	05.2		165.8	08.1		225.7	11.1	86 285.7
	48	46.9 47.9	02.3		106 9 107 9	05.3		166.8	08.2		226.7	11.1	87 286 . 7
H	49	48.9	02.4		108.9	05.3 05.3		16 7 .8	08.2 08.3		227.7 228.7	11.2	
ı	50	49.9	02.5		109.9	05.4		169.8	08.3		228.7 229.7	11.2 11.3	89 288.7 90 289.7
H	51	50.9	02.5		110.9	05.4		170.8	08.4		230.7	11.3	291 290.6
l	52	51.9	02.6		111.9	05.5		171.8	08.4		231.7	11.4	92 291.6
lİ	53	52.9	02.6		112.9	05.5	-73	172.8	08.5		232.7	11.4	93 292.6
	54 55	53.9 54.9	02.6		113.9	05.6		173.8	08.5		233.7	11.5	94 293.6
П	56	55.9	02.7 02.7		114.9 115.9	05.6		174.8	08.6		234.7	11.5	95,294.6
H	57	56.9	02.8		116.9	C5.7 05.7		175.8 176.8	08.6 08.7		235.7 236.7	11.6	96 295.6
Н	58	57.9	02.8		117.9	05.8		177.8	08.7		236.7 237.7	11.6	97 296 .6 98 297 .6
l	59	58.9	02.9	19	118.9	05.8	79		08.8		238.7	11.7	
	60	59.9	02.9		119.9	05.9		179 8	08.8		239.7		
	Dist.	Dep.	Lat.	Dist.		Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep.
}_		E. N.			E.48.			W.IN			W.15		For 71 P
		-								Dist		~ (\V	10712

*

TABLE L. Difference of Latitude and Departure for 1 Point.

	N.JE.			N.JW.			S.JE.		s. j w.		
Lat.	Dep.	Dist.		Dep.		Lat.		Dist. Lat.	Dep.	Dist. Lat.	Dep.
01.0	00.1	61	60.7	06.0		120.4		181 180.1	17.7		23.6
02.0	00.2	62	61.7	06.1		121.4		82,181.1	17.8	42 240.8	
03.0	00.3	63		06.2		122.4		83 182.1			
04.0	00.4	64				123.4		84 183 . 1			
05.0	00.5	65				124.4		85 184 . 1			
06.0	00.6	66	65.7	06.5	26	125.4	12.4	86 185.1	18.2		24.1
07.0	00.7	67	66.7	06.6		126.4		87 186.1			
08.0 09.0	00.8	68 69	67.7 68.7	06.7 06.8		127.4 128.4		88 187.1 89 188.1			
10.0	01.0	70		06.9		129.4 129.4		90 189.1	18.6		
								191 190.1		II———	
10.9 11.9	01.1 01.2	71 72	70.7 71.7	07.0 07.1		130.4		92 191.1	18.7		
12.9	01.3	73	72.6	07.2		131.4 13 2 .4		93 192.1	18.9		
13.9	01.4	74	73.6	07.8		133.4		94 193.1	19.0		
14.9	01.5	75	74.6	07.4		134.3		95 194.1	19.1		
15.9	01.6	76	75.6	07.4		135.3	13.3	96 195.1	19.2		
16.9	01.7	77	76.6	07.5		156.3	13.4	97 196 . 1			25.2
17.9	01.8	78	77.6	07.6	38	137.3	13.5	98 197.0			
18.9	01.9	79	78.6	07.7		138.3		99 198.0			25.4
19.9	02.0	.80	79.6	07.8	40	139.3	13.7	200 199.0	19.6	60 258.7	25.5
20.9	02.1	81	80.6	07.9	141	140.3	13.8	201 200.0	19.7	261 259.7	25.6
21.9	02.2	82	81.6	08.0	42	141.3	13.9	62 201.0	19.8	62 260.7	
22.9	02.3	83	82.6	08.1		142.3		03 202.0			
23.9	02.4	84	83.6	08.2		143.3		04 203.0			
24.9	02.5	85	84.6	08.3		144.3		05 204.0			
25.9	02.5	86	85.6	08.4		145.3	14.3	06 205.0			
26.9 27.9	02.6 02.7	87	86.6	08.5		146.3	14.4	07 206.0	20.3		
28.9	02.8	88 89	87.6 88.6	08.6 08.7		147.3	14.5	08 207.0 09 208.0	20.4		
29.9	02.9	90	89.6	08.8		148.3 149.3	14.6 14.7	10 209.0	20.5 20.6		26.5
	 i										
30.9	03.0	91	90.6	08.9		150.3	14.8	211 210.0	20.7	271 269 . 7	26.6
31.8 32.8	03.1 03.2	92 93	91.6	09.0 09.1		151.3	14.9	12211.0 13212.0	20.8		26.7 26.8
33.8	03.3	94	92.6 93.5	09.2		15 2. 3 15 3 . 3	15.0 15.1	14 213.0	20.9 21.0		26.9
34.8	03.4	95	94.5	09.5		154.3	15.2	15 214.0	21.1		27.0
35.8	03.5	96	95.5	09.4		156.2	15.3	16 215.0	21.2		27.1
36.8	03.6	97	96.5	09.5		156.2	15.4	17216.0	21.3	77275.7	27.2
37.8	03.7	98	97.5	09.6	58	157.2	15.5	18217.0	21.4	78276.7	27.2
38.8	03.8	99	98.5	09.7	59	158.2	15.6	19217.9	21.5		27.3
39.8	03.9	100	99.5	09.8	60	159.2	15.7	201218.9	21.6	80 278.7	27.4
10.8	04.0	101	100.5	09.9	161	160.2	15.8	221 219.9	21.7	281 279.6	27.5
11.8	04.1	02	101.5	10.0	62	161.2	15.9	22 220.9	21.8	82 280.6	27.6
12.8	04.2	03	102.5	10.1	63	162.2	16.0	23 221.9	21.9	83 281.6	27.7
13.8	04.5		103.5	10.2		163.2	16.1	24 222 9	22.0		27.8
14.8	04.4		104.5	10.3		164.2	16.2	26 223.9	22.1	85 283.6	27.9
15.8	04.5		105.5	10.4		165.2	16.3	26 224.9	22.2	86 284.6	28.0
16.8 17.8	04.6		106.5	10.5		66.2	16.4	27 225.9	22.2	87 285.6	28.1
18.8	04.8		107.5 108.5	10.6 10.7		167.2 168.2	16.5 16.6	28 226 . 9 29 227 . 9	22.3 22.4	88 286.6	28.2 28.3
19.8	04.9		109.5	10.8		69.2	16.7	30'228.9	22.5	89287.6 90288.6	28.4
50.8	05.0									1	
51.7			110.5	10.9		70.2	16.8	231 229 .9	22.6	291 239.6	28.5
52.7	05.1 05.2		111.5	11.0		71.2	16.9	32 230.9	22.7	92 290.6	28.6
53.7	05.3		113.5	11.2		73.2	17.0 17.1	33 231.9 34 232.9	22.8 22.9	93 291.6 94 292.6	28.7
	05.4		114.4	11.3		74.2	17.2	35 233.9	22.9 23.0	95 293.6	28.9
	05.5		15.4	11.4		75.2	17.3	36/234.9	23.1	96 294.6	29.0
	05.6		116.4	11.5		76.1	17.3	37,235.9		97 295.6	29.1
	05.7		17.4	11.6		77.1	17.4	381236.9	23.2 23.3	98 296.0	29.2
18.7		19	118.4	11.7	.79	78.1	17.5	39 237.8	23.4	99 497.6	29.3
59.7				11.8	80 1	79.1	17.6	40 233.8	23.0	200 298 . 6	29.4
)ap. (Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep.		Dist. Dep.	i.at.
AN.			:.JS.		7	V.4N.		W. 15.		[For 7] I	
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Difference of Latitude and Departure for \$ Point.

1			N.E.			N.§W.			S.ŁE.		·	S. §W .		
1	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.		Dist.	Lat.		Dist. Lat.	Dep.
1:	1	01.0		61	60.3	09.0		119.7	17.8		179.0	26.6	241 238.4	35.4
H	2	02.0	00.3	62	61.3	09.1		120.7	17.9		180.0	26.7	42 239 .4	35.5
H	3	03.0 04.0	11	63	62.3 63.3	09.2		121.7 122.7	18.0 18.2		181.0 18 2 .0	26.9 27.0	43 240.4 44 241.4	35.7
П	4	04.9		65	64.3	09.5		123.6	18.3		183.0	27.1	45 242.3	35 .9
H	6	05.9		66	65.3	09.7		124.6	18.5		184.0	27.3	46 243.3	36.1
1!	7,	06.9	01.0	67	66.3	09.8		125.6	18.6		185.0	27.4	47 244.3	36.2
11	8	07.9	01.2	68	673	10.0		126.6	18.8		186.0	27.6	48 245.3	36.4
11	9	08.9		69 70	68.3	10.1		127.6	18.9		187.0 187.9	27.7 27.9	49 246 .3 50 247 .3	36.5
11.	10	09.9	01.5		69.2	10.3		128.6	19.1					36.7
П	11 12	10.9 11.9		71 72	70.2 71.2	10.4 10.6		129.6	19.2 19.4		188.9 189.9	28.0 28.2	251 248.3 52 249.3	36.8 37.0
П	13	12.9		73	72.2	10.7		130.6 131.6	19.5		190.9	28.3	53 250.3	37.1
Н	14	13.8		74	73.2			132.5	19.7		191.9	28.5	54 251 .3	37 .3
П	15	14.8		75	74.2	11.0		133.5	19.8		192.9	28.6	55 252.2	37.4
Н	16	15.8	L 1	76	75.2	11.2		134.5	20.0		193.9	28.8	56 253.2	37.6
H	17	16.8		77	76.2	11.3		135.5	20.1		194.9	28.9	57 254.2	37.7
П	18	17.8 18.8		78 79	77.2	11.4		136.5	20.2 20.4		196.9 196.3	29.1 29.2	58 255.2 59 256.2	37 .9 38.0
11	19 20	19.8		80	78.1 79.1	11.6 11.7		137.5 138.5	20.4		197.8	29.2	60 257,2	38.0
1	21	20.8	<u> </u>	81	80.1		·		20.7		198.8	29.5	261 258 .2	38.3
П	22	21.8		82	81.1	11.9 12.0		139.5 1 40 .5	20.8		199.8	29.6		38.4
H	23	22.8		83	82.1			141.5	21.0		200.8	29.8		
П	24	23.7		84	83.1			142.4			201.8	29.9	64 261 .1	38.7
П	25	24.7		85	84.1		45	143.4			2 02 . 8	30.1	65 262.1	38.9
П	26	25.7	1	86	85.1			144.4			203.8	30.2		39.0
Н	27	26.7		87	86.1	12.8		145.4			204.3	30.4		39.2
11	28 29	27.7 28.7		88 89	87.0	12.9		146.4 147.4			205.7 206.7	30.5 30.7	68 265.1 69 266.1	39.3 39.5
П	30	29.7		90	88.Q 89.0	13.1 13.2		148.4			207.7	30.8	70 267.1	39.6
H	31	30.7		91	90.0	13.4		149.4		t I	208.7	31.0	271 268 . 1	39.8
11	32	31.7		92	91.0			150.4			209.7	31.1	72 269.1	39.9
1!	33	32.6		93	92.0			151.3			210.7	31.3	73 270.0	
11	34	33.6		94	93.0			152.3		14	211.7	31.4	74 271.0	40.2
11	35	34.6		95	94.0			153.3			212.7	31.5	75 272.0	
11	36	35.6			95.0			154.3			213.7	31.7	76 273.0	
	37 38	36.6		97 98	96.0			155.3			214.7 215.6			
Н	39	38.6		99	96.9 97.9	14.4 14.5		156.3 157.3			216.6		78,275.0 79,276.0	40.9
	40	39.6		100	98.9	14.7		158.3			217.6	32.3	80 277.0	
	41	40.6		1		14.8		159.3	i	I—	218.6	32.4	281 278 .0	
	42	41.5			100.9	15.0	1 . 1	160.2	J 1		219.6	32.6	82 278 .9	
Į į	43	42.5	06.3	03	101.9	15.1	7	161.2		23	220.6		83,279.9	41.5
11	44	43.5			102.9	15.3	64	162.2			221.6			
1	45	44.5			103.9	15.4		163.2			222.6			
1 1	46 47	45.5 46.5			104.9			164.2			223.6 224.5		86,282.9 87,283.9	
11	48	47.5			105.8 106.8			165.2 166.2			225.5		88 284 .9	I
11	49	48.5			107.8			167.2			226.5		89 285 . 9	
П	50	49.5			108.8	16.1		168.2	4 : 1		227 5		90 286 .9	
11	51	50.4	07.5	11	109.8	16.3	1	169.1	<u> </u>	231	228.5	33.9	291 287 . 9	42.7
H	52	51.4			110.8			170.1			229.5		92 288 . 8	42.8
H	53	52.4	07.8	13	111.8	16.6	73	171.1	25.4	33	230.5	34.2	93 289 .8	
11	54	53.4			112.8			172.1			231.5			
H	55 56	54.4			113.8			173.1			232.5		95 291 .8 96 292 .8	
П	56 57	55.4 56.4		JI	114.7 115.7			174.1 175.1			233.4 234.4			
11	58	57.4			116.7			176.1			235.4		98 294	
$\{ \}$	59	58.4			117.7			177.1		1 1	236.4		99 295.8	
۱'	60	59.4			118.7			178.1	26.4		237 .4		300 296 .8	44.0
1	Dbt.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist	Dep.	Lut.	Dist. Dep.	iant.
1		EIN			E. (S.			W.IN	ī		WAS	. C-1	[For 74 F	oints.

	Difference of Latitude and Departure for 1 Point. N.b.E. N.b.W. S.b.E. S.b.W. Dist. Lat. Dep. Dist. Lat. Dep. Dist. Lat. Dep. Pist. Pi													
Dist.	Lat.		Diet !	Lat.					l)ist :			Dist.	Lat.	Dep,
1	01.0	00.2	61	59.8	11.9		118.7	23.6		177.5	35.3		236.4	47.0
2	02.0	00.4	62	60.8			119.7	23.8		178.5	35.5	42	237.4	47.2
. 3	02.9	00.6	63	61.8			120.6	24.0		179.5	35.7		238.3	
4 5	03.9 04.9	00.8	64 65	62.8 63.8	12.5 12.7		121.6 122.6	24.2 24.4		180. <i>5</i> 181.4	35.9 36.1		239.3 240.3	
6	05.9	01.2	66	64.7	12.9		123.6	24.6		182.4	36.3		241.3	
7	96.9	01.4	67	65.7	13.1	27	124.6	24.8	87	183.4	36.5	47	242.3	48.2
8	07.8	01.6	68	66.7	13.3		125.5	25.0		184.4	36.7		243.2	
10	08.8 09.8	01.8 02.0	69	67.7 68.7	13.5 13.7		126.5 127.5	25.2 25.4		185.4 136.3	36.9 37.1		244.2 245.2	48.6 48.8
11	10.8	02.1	71	69.6	13.9		128.5	25.6		187.3	37.3		216.2	49.0
12	11.8	02.3	72	70.6	14.0		129.5	25.8	92	188.3	37.5	52	247.2	
13	12.8	02.5	73	71.6	14.2		130.4	25 .9		189.3			248.1	49.4
14 15	13.7 14.7	02.7	74 75	72.6 73.6	14.4 14.6		131.4 132.4	26.1 26.3		190.3 191.3			249.1 250.1	49.6 49.7
16	15.7	03.1	76	74.5	14.8		133.4	26.5 26.5		192.2			251.1	
17	16.7	03.3	77	75.5			134.4	26.7		193.2			252.1	
18	17.7	03.5	78	76.5	15.2		135.3			194.2			258.0	
19 20	18.6 19.6	03.7	79 80	77.6	15.4		136.3	27.1		195. 2 196. 2			254.0 255.0	
		03.9	l	78.5	15.6	I	137,3	27.3				ii		
	22 21.6 04.3 82 80.4 16.0 42 139.3 27.7 02 198.1 39.4 62 257.0													
23	22.6	04.5	83	81.4			140.3			199.1	39.6		257.9	
24	23.5	04.7	84	82.4	16.4	44	141 .2	28.1	04	200.1			258.9	51.5
25	24.5	04.9	85	83.4			142.2			201.1			259.9	
26 27	25.5 26.5	05.1 05.3	86 87	84.3 85.3			143.2 144.2			202.0 203.0			260.9 261.9	
28	27.5	05.5	88	86.3			145.2			204.0			262.9	
29	28.4	05.7	89	87.3			146.1			205.0			263.8	
30	29.4	05 .9	90	88.3	17.6	50	147.1	29 .3	10	206.0	41.0	70	264.8	52.7
31	30.4	06.0	91	89.3	17.8		148.1	29.5		206.9		271		
32	31.4	06.2	92	90.2			149.1			207.9			266.8	
33 34	32.4 33.3	06.4 06.6	93 94	91.2 92.2			150.1 151.0	29.8 30.0		208.9 209.9			267.8 268.7	53.3 53.5
35	34.3	06.8	95	93.2			152.0			210.9			269.7	
36	35.3	07 .0	96	94.2			153.0			211.8		76	270.7	53.8
37	36.3 37.3	07.2	97 98	95.1	18.9		154.0		17	212.8			271.7	
38	38.3	07.4 07.6	99	96.1 97.1	19.1 19.3		155.0 155.9			213.8 214.8			272.7 273.6	
40	39.2	07.8	100	98.1	19.5		156.9		20	215.8			274.6	
41	40.2	08.0	101	99.1	19.7	161	157.9	31.4	11	216.8		281	275.6	54.8
42	41.2	08.2		100.0	19.9	62	158.9	31.6		217.7	43.3	82	276.6	
43	42.2	08.4		101.0			159.9			218.7			277.6	
44 45	43.2 44.1	08.6 08.8		102.0 103.0			160.8 161.8			219.7 220.7			278.5 279.5	
46	45.1	09.0		104.0			162.8			221.7			230.5	
47	46.1	09.2	07	104.9	20.9	67	163.8	32 .6	27	222.6	44.3	87	281.5	56.0
48	47.1	09.4		105.9			164.8			223.6			282.5	
49 50	48.1 49.0	09.6 09.8		106.9 107.9			165.8 166.7	33.0 33.2		224.6 225.6			283.4 284.4	
51	50.0	09.9		108.9	21.7		167.7	33.4		226.6			285.4	
52	51.0	10.1		109.8			168.7			227.5			286.4	
53	52.0	10.3	13	110.8	22.0	73	169.7	33.8	33	2 2 8.5	45.5	93	287.4	57.2
54	53.0	10.5	14	111.8			170.7			229.5			288.4	
55 56	53.9 54.9	10.7 10.9		112.8 113.8			171.6 172.6			230.5			289.3 290.5	
57	55.9	11.1		114.8			173.6			231.5 232.4			290.3	
58	56.9	11.3	18	115.7	23.0		174.6	34.7	38	233.4	46.4	98	292.5	58.1
59	57.9	11.5		116.7			175.6		39	234.4	46.6		293.5	
60				117.7			176.5		·	235.4			294.2	
Ulst.	Dep. E.b.N	Lat.			Lat.				'Dist.				Dep.	
	10.U.N	<u> </u>		E.b.S.			M.d.W			W.b.	tized by	J_T(ror7	Points.

TABLE I.

Difference of Latitude and Departure for 11 Points.

		N	.b.E.	E.	N	.b.W. <u>}</u>	W.	:	3.b.E. .	E.	9	.b.W.1	W.		
	Dist	Lat	Dep.		Lat.		Dist.		Dep.		Lat.	Dep.		Lat.	Dep.
1	1							117.4		181	175.6			233.8	
1	1 3			62				118.3 119.3			176.5 177.5			234.7 235.7	
ł	4	J 03.9	01.0	64	62.1	15.6	24	120.5	30.1		178.5		44	236.7	59.3
ł	1 6			65				121.5			179.5			237,7	
ŀ	1 2			66				122.2 123.5			180.4 181.4			238.6 239.6	
ļ	8	07.8	01.9	68	66.0	16.5	28	124.5	31.1	88	182.4	45.7	48	240.6	60.5
١	10		02.2	69 70				1 25 .1 126.1			183.3			241.5	
.}	11		02.7	71	68.9			127.1	'	:	134.3 185.3			242.5 243.5	
!	12		02.7	72	69.8			128.0			186.2			244.4	
ļ	13		03.2	73	70.8	17.7	33	129.0	32.3	93	187. 2	46.9	53	245.4	61.8
ł	14	1	03.4	74 75	71.8 72.8			1 30 .0 131.0			188. 2 189.2	47.1 47.4		246.4 247.4	
ı	16		03.9	76	73.7			131.9			190.1	47.6		248.3	
ı	17		04.1	77	74.7	18.7		132.9			191.1	47.9		249.3	62.4
ļ	18		04.4	78 79	75.7 76.6	19.0 19.2		133.9 1 34 .8			192.1 193.0	48.1 48.4		250.3 251.2	
l	20		04.9	80	77.6			135.8			194.0	48.6		252. 2	
١	21	20.4	0á.1	81	73.6		141	136.8	34.3	201	195.0	48.8		253.2	
l	22	21.3	05.3	82	79.5	19.9		137.7	34.5	02	195 9	49.1		254.1	63.7
١	23 24	22.3 23 3	05.6 05.8	80 84	80.5 81.5			! 38 .7 ! 39 .7	34.7 35.0		196. 9 19 7.9	49.3 49.6		255 . 1 256 . 1	63.9 64.1
۱	25	24.3	06.1	35	82.5			140.7	35 2		198.9	49.8		257.1	64.4
١	26	25.2	06.3	86	83.4	20.9		11.6	35.5	06	199.8	50.1	66	238.0	64.6
١	27 28	26.2 27.2	06.6 06.0	87 88	84.4 85.4			142.6 143.6	35.7 36.0		200.8 201.8	50.3 50.5		259.0 260.0	64.9 65.1
I	29	28.1	07.0	89	86.3			44.5	36 .2		202.7	50.8		260.9	65.4
I	30	29.1	07.3	90	87.3	21.9	0	145.5	36.4		203.7	51.0	70	261.9	65.6
ı	31	30.1	07.5	91	88.3			16.5	36.7		204.7	51.3		262.9	65.8
1	32 33	31.0 32.0	07.8 08.0	92 93	89.2 90.2	27.4 22.6		47.4 48.4	36.9 37.2		205.6 206.6	51.5 51.8		263.8 264.8	66.1 66.3
l	34	33.0	08.3	94	91.2			49.4	37.4		207.6	52.0		265.8	66.6
l	35	34.0	08.5	95	92.2			50.4	37.7		208.6	52.2		266.8	66.8
l	36 37	34.9 35.9	08.7 09.0	96 97	93.1 94.1			51.3 52.3	37.9 38.1		2 09.5 210.5	52.5' 52.7		267.7 268.7	67.1 67.3
l	38	36.9	09.2	98	95.1		51:	3.3ز.ا		18	211.5	53.0		269.7	67.5
l	39 40		09.5	99	96.0	24.1		54.2	38.6		212.4	53.2		270.6	67.8
l	41	38.8 39.8	10.0	100	970			56.2	38.9	·	213.4	53.5 53.7		$\frac{271.6}{272.6}$	68.0 68.3
l	42	40.7	10.0	02	98.9	24.5 24.8		57.1	39.4		215.3	53.7		272.6 273.5	68.5
ł	43	41.7	10.4	03	99.9	25.0	63 1	58.1	39.6	23	216.3	54.2	83	274.5	68.8
ı	44 45	42.7	10.7 10.9		100.9	25.3		59.1	39.8		217.3 218.3	54.4 54.7		275.5 276.5	69.0
l	46	44.6	11.2		101.9 102.8	25.5 25.8		60.1 61.0	40.1		19.2	54.9		277.4	69.2 69.£
l	47	45.6	11.4	07	103.8	26.0	67 1	62.0	40.6	27	20.2	55.2	87	278.4	69.7
l	48 49	46.6 47.5	11.7		104.8	26.2		63.0	40.8		21.2 22.1	55.4		279.4	70.(
١	50	48.5	12.1		105.7 106.7	26.5 26.7		63 . 9 64 . 9	41.1		23.1	55.6 55.9		280 .3 281 .3	70.5 70.5
	51	49.5	12.4		107.7	27.0		65.9	41.5		24.1	á6.1		282.3	70.7
	52	50.4	12.6		108.6	27.2		66.8	41.2		25.0	56.4		283.2	71.0
ı	53 54	51.4 52.4	12.9 13.1		109.6 110.6	27.5 27.7		67.8 68.8	42.0 42.3		26.0 27.0	56.6 56.9		2 84 . 2 28 5 . 2	71.4
	55	53.4	13.4		111.6	27.9		69.8	42.5	35	28.0	57.1		286.2	71.7
	56	54.3	13.6	16	112.5	28.2	76 1	70.7	42.8	36 2	28.9	57.3	96	287.1	71.5
	57 58	55.3 56.3	13.8 14.1		113.5	28.4 28.7		71 .7 72 .7	43.0 43.3		29.9 30.9	57.6 57.8		288.1 289.1	72.1 72.4
1	59	57.2	14.3		115.4	28.9		73.6	43.5	39 2	31 8	58.1		290.0	72.7
ł	60	58.2	14.6	20	116.4	29.2	80:1	74.6	43.7		32.8	58.3	300	291 0	72.5
1		Dep.	Lat.	Dist.	Dep.	Lat.	Dist. I	Dep.	Lat.						
_	E.	N.E.4E	i.	E.S	.E.4 E	٠	<u>1.W</u>	ı.W.	W	W.S	.W.	Cio	o G	or 61 I	oint

TABLE I.

Difference of Latitude and Departure for 14 Points.

	N	.b.E.41	E.	N.	b.W.4	W.	S	.b.E.4	E.	S.	ا دٍ.w .d	W.			1
Dist.	Lat.		Dist.				Lat.	Dep.	Dist.	Lat.	Dep.			Dep.	ار
1	01.0			58.4			115.8			173.2		11 -	230.6		11
2				59.3 60.3			116.7 117.7			174.2 175.1		11	231.6 232.5		П
4	03.8			61.2	1		118.7			176.1			233.5		П
5	04.8	01.5		62.2			119.6						234.5		П
6	05.7	01.7		63.2			120.6			178.0			235.4		П
7 8	06.7 07.7	02.0 02.3		64.1 65.1			121.5 122.5			178.9 179.9			236.4 237.3		П
9	08.6	02.6		66.0	20.0		123.4			180.9			238.3		Н
10	09.6	02.9	70	67.0	20.3		124.4			181.8			239.2		Н
11	10.5	03.2		67.9	20.6		125.4			182.8			240.2		11
12	11.5	03.5		. 68 . 9			126.3			183.7			241.1		П
13 14	12.4 13.4	03.8 04.1	73 74	69.9 70.8			127.3 128.2			184.7 185.6			242.1 243.1	73.4	Н
15	14.4	04.4		71.8	21.8		129.2	39.2		186.6			244.0		Ш
16	15.3	04.6	76	72.7	22.1	36	130.i	39.5		187.6	56.9	56	245.0	74.3	П
17	16.3	04.9	77	73.7	22.4	11	131.1	39.8		188.5			245.9		Н
18 19	17.2 18.2	05.2 05.5	78 79	74.6 75.6	22. 6 22. 9		13 2 .1 13 3 .0	40.1		189.5 190.4			246.9 247.8		
20	19.1	05.8	80	76.6	23.2		134.0			191.4			248.8		П
21	20. f	06.1	81	77.5	23.5					192.3		l	249.8		Ħ
22	21.1	06.4	82	78.5	23.8		135.9			193.3			250.7		Н
23	22.0	06.7	83	79.4			136.8			194.3			251.7		П
24 25	23.0 23.9	07.0 07.3	84 85	80.4 81.3	24.4 24.7		137.8 138.8			195. 2 196. 2			252.6 253.6		П
26	24.9	07.5	86	82.3	25.0		139.7	42.4		190.2			254.5		П
27	25.8	07.8	87	83.3	25.3		140.7			198.1			255.5		П
28	26.8	08.1	88	84.2	25.5		141.6			199.0			256.5		
29 30	27.8 28.7	03.4 08.7	89 90	85.2 86.1	25.8 26.1		142.6 143.5	43.3 43.5		200.0 201.0			257.4 258.4		
31	29.7	09.0	91	87.1	26.4		144.5	43.8		201.9	61.3		259.3		
32	30.6	09.3	92	88.0			145.5	44.1		201.9 202.9	61.5		260.3		П
33	31.6	09.6	93	89.0	27.0		146.4	44.4		203.8			261.2		
34	32.5	09.9	94	90.0	27.3		147.4			204.8			262.2	79.5	
35 36	33.5 34.4	10.2 10.5	95 96	90.9 91.9	27.6 27.9		148.3 149.3	45.0 45.3		205.7 206.7	62.4 62.7		263.2 264.1	79.8 80 I	П
37	35.4		97	92.8	28.2		150.2	45.6		207.7	63.0		265.1	80.4	
33	36.4	11.0	98	93.8	28.4	58	151.2	45.9	18	208.6	69.3	78	266.0		П
39	37.3		99	94.7	28.7			46.2		209.6			267.0		
40	38.3		100	95.7	29.0	1		46.4		210.5	63.9		267.9		
41 42	39.2 40.2	11.9 12.2	101	96.7 97.6	29.3 29.6		154.1 155.0	46.7 47.0		211.5 212.4	64.2 64.4		268.9 269.9	81.6 81.9	
43	41.1	12.5	03	98.6	29.9		156.0			213.4			270.8		
44	42.1	12.8	04	99.5	30.2		156.9	47.6		214.4	65.0		271.8		$\{ \}$
45	43.1	13.1		100.5	30.5		157.9	47.9		215.3			272.7		П
46 47	44.0 45.0	13.4 13.6		101.4 102.4	30.8 31.1		158.9 159.8	48.2 48.5		216.3 217.2	65.6 65.9		273.7 274.6	83.0 83.3	
48	45.9	13.9		103.3	31.4		160.8	48.8		218.2	66.2		275.6	83.6	il
49	46.9	14.2	09	104.3	31.6	69	161.7	49.1	29	219.1	66.5	89	276.6	83.9	
50	47.8	14.5		105.3	31.9		162.7	49.3		220.1	66.8		277.5	84.2	П
51	48.8	14.8		106.2	32.2		163.6	49.6		221.1	67.1		278.5	84.5	П
52 53	49.8 50.7	15.1 15.4		107.2 108.1	32.5 32.8		164.6 165.6	49.9 50.2		222.0 223.0	67.3 67.6		279 . 4 280 . 4	84.8 85.1	
54	51.7	15.7		109.1	33.1		166.5	50.2 50.5		223.0 223.9	67.9		281.3	85.3	
55	52.6	16:0	15	110.0	33.4	75	167.5	50.8	35	224.9	68.2	95	282.3	85.6	
56	53.6	16.3		111.0	33.7		168.4	51.1		225.8	68.5		283.3	85.9	
5 7 58	54.5 55.5	16.5 16.8		112.0 112.9	34.0 34.3		169.4 170.3	51.4 51.7		226 . 8 227 . 8	68.8 69.1		284.2 285.2	36.2 86.5	ĺ
59	56.5	17.1		13.9	34.5		171.3	52.0		228.7	69.4		286.1	\$ 6.8	!
60	57.4	17.4		114.8	34.8		172.2	52.3	40	229.7	69.7	300	287.1	87.1	ļ
	Dep.	Lat.		Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	1
E.	N.E.JE		E.	3.E.JE			N.W.		W.	.W.41	W.			Points.	

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Difference of Latitude and Departure for 11 Points.

		N.b.E.	E.	N	.b.W.	W.	:	S.b.E.	E.	S	.b.W.\$	W.	
Dis	iy Lat	. Dep.	l'ist	Lat.	Dep.	Dist	. Lat.	Dep	Dist	Lat.	Dep.	Dist. Lat.	Dep.
	1 00.						113.9			170.4			9 81.2
	2 01.						114.9			2 171 .4			
	3 02.						115.8			172.			
	4 03. 5 04.						116.8			173.9 174.9			
	05.						118.6			175.			
11 3							119.6			176.			
8							120.5			3 177.0			
!! !!							121.5			178.0			
10			-				122.4			178.9		!!	
11		-1 .					128.3			179.8 180.6			
12							124.3 125.2			181.7			
14							126.2			182.7	65.4		
15						35	127.1	45.5		183.6			
16						36	128.0			184.5			
17				72.5			129.0			185.5			
18				78.4 74.4			129.9			186.4 187.4			
20			80				130.9 131.8			188.5			
21			81	76.3	27.3		132.8			189.5		261 245.7	
22		07.4		77.2	27.6	42	133.7			190.2		62 246.7	
23			83	78.1	28.0		134.6			191.1	68.4	63 247.6	
24			84	79.1	28.3		135.6			192.1		64 248 6	
25				80.0	28.6		136.5			193.0			
26	24.5 25.4		86	81.0	29.0		137.5			194.0 194.9		66 250.5	
27 28	26.4		87 88	81.9 82.9	29.3 29.6		138.4 139.3			195.8		67 251.4 68 252.3	
29	27.3		89	83.8	30.0		140.3	50.2		196.8	70.4	69 253.3	
30			90	84.7	30.3		141.2	50.5		197.7	70.7	70 254.2	
31	29.2	10.4	91	85.7	30.7	151	142.2	50.9	211	198.7	71.1	271 255.2	91.3
32	30.1		92	86.6	31.0		143.1	51.2		199.6	71.4	72 256.1	
33	31.1		93	87.6	31.3		144.1	51.5		200.5	71.8	73 257.0	
34 35	32.0 33.0			88.5	31.7 32.0		145.0 145.9	51.9 52.2		201.5 202.4		74 258.0 75 258.9	
36	33.9		96	89.4 90.4	32.3		146.9	52 .6		203.4	72.8	76 259.9	
37	34.8		97	91.3	32.7		147.8	52.9		204.3	73.1	77 260.8	
38	35.8	12.8	98	92.3	33.0	58	148.8	53.2		205.3	73.4	78 261.7	
39	36.7		99	93.2	33.4		149.7	53.6		206.2	73.8	79 262.7	
40	37.7		100	94.2	33.7		150.6	53.9		207.1	74.1	80 263.6	1
41	38.6		101	95.1	34.0		151.6	54.2		208.1	74.5	281 264.6	
42	39.5 40.5	14.1 14.5	02)	96.0 97.0	34.4		152.5 153.5	54.6 54.9		209.0 210.0	74.8 75.1	82 265.5 83 266.5	
44	41.4		04	97.9	35.0		154.4	55.2		210.9	75.5	84 267 .4	
45	42.4		05	98.9	35.4	65	155.4	55.6		211.8	75.8	85 268.3	
46	43.3	15.6	06	99.8	35.7	66	156.3	55.9		212.8	76.1	86 269.3	
47	44.3	15.8		100.7	36.0		157.2	56.3		213.7	76.5	87 270.2	
48	45.2 46.1	16.2 16.5		101.7	36.4		158.2	56.6 56.9		214.7 215.6	76.8 77.1	88 271.2 89 272.1	
50	47.1	16.8		102.6 103.6	36.7 37.1		159.1 160.1	57.3		216.6	77.5	90 273.0	
51	48.0	17.2		104.5			161.0	57.6		217.5	77.8		98.0
52	49.0	17.5		105.5	37.4 37.7		161.9	57.9		218.4	78.2	291 274.0 92 274.9	
53	49.9	17.9		106.4	38.1		162.9	58.3		219.4	78.5	93 275.9	1
54	50.8	18.2	14	107.3	38.4	74	163.8	58 .6	34	220.3	78.8	94 276.8	99.0
55	51.8	18.5		108.3	38.7		164.8	59.0		221.3	79.2	95 277.8	
56	52.7	18.9 19.2		109.2	39.1		165.7	· 59.3		222.2 223.1	79.5 79.8	96 278.7 97 279.6	99.7
57 58	53.7 54.6	19.2		110.2	39.4 39.8		166.7 167.6	59.6 60.0		223.1 224.1	80.2	98 280.6	
59	55.6	19.9		12.0	40.1		168.5	60.3		225.0	80.5	99 281.5	100.7
60	56.5	20.2		13.0	40.4		169:5			226.0	80.9	300 282.5	
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.				Dist.	Dep.	Lat.	Dist. Dep.	Lat.
E	N.E.			.S.E.	Ē.	W.	N.W.	W.	W	.S.W.	W.	For 6	Points.

TABLE I.

Difference of Latitude and Departure for 2 Points.

		N.N.E.			N.N.W	·.		S.S.E			s.s.w.			٠.	1
Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.			Dep.	ī
1			61				111.8	46.3		167.2			222.7	92.2	1
2							112.7 113.6	46.7 47.1		168.1 169.1	69.6 70.0		223.6 224.5		H
4							114.6	47.5	84	170.0	70.4	44	225.4	93.4	H
. 5			65		24.9		115.5	47.8		170.9	70.8		226.4		П
6 7	05.5 06.5	02.3 02.7	66	1			116.4 117.3	48.2 48.6		171.8 172.8	71.2 71.6		227.3 228.2		П
8	07.4	03.1	68			23	118.3	49.0	88	173.7	71.9	48	229.1	94.9	Н
. 10			69				119.2	49.4		174.6			230.0		Н
10	09.2	03.8	70		26.8		120.1	49.7		175.5			231.0		11
11 12	10.2 11.1	04.2 04.6	71 72		27.2 27.6		121.0 122.0	50.1 50.5		177.4	73.1 73.5		231.9 232.8		П
13		05.0	73				122.9	50.9	93	178.3			233.7	96.8	Н
. 14	12.9	05.4	74				123.8			179.2	74.2		234.7		П
15 16	13.9 14.8	05.7 06.1	75 76				124.7 125.6	51.7 52.0		180. 2 181.1	74.6 75.0		235.6 236.5	97.6 98.0	П
17	15.7	06.5			29.5		126.6	52.4	97	182.0	75.4	57	237.4	98.3	П
18	16.6	06.9	78				127.5	52.8		182.9	75.8		238.4		11
19 2 0	17.6 18.5	07.3 07.7	79 80	1	30.2 30.6		128.4 129.3	53.2 53.6		183.9 184.8	76.2 76.5		239.3 240.2		П
21	19.4	08.0	81	74.8	31.0	i	130.3	54.0		185.7	76.9		241.1		11
22	20.3	08.4	82				131.2			186.6	77.3			100.3	П
23	21.2	08.8	83		31.8		132.1	54.7		187.5	77.7			100.6	П
24 25	22.2 23.1	09.2 09.6	84 85				135.0 134.0	55.1 55.5		188.5 189.4	78 · 1 78 · 5			101.0 101.4	П
26	24.0	09.9	86				134.9	55.9		190.3	78.8			101.8	П
27	24.9	10.3	87	80.4	33.3	47	135.8	56.3	07	191.2	79.2	67	246.7	102.2	
2 8 29	25.9	10.7	88				136.7	56 .6		192.2 1 93 .1	79.6			102.6	П
30	26.8 27.7	11.1 11.5	89 90				137.7 138.6	57.0 57.4		194.0	80.0 80.4			102.9 103.3	Н
31	28.6	11.9	91	84.1	34.8		139.5	57 .8		194.9	80.7			103.7	11
32	29.6	12.2	92	85.0	35.2	52	140.4	58.2	12	195.9	81.1	72	251.3	104.1	Н
33	30.5	12.6	93		35.6		141.4	58.6		196.8	81.5			104.5	
34 35	31.4 32.3	13.0 13.4	94 95		36.0 36.4		142.3 143.2	58.9 59.3		197.7 198.6	81.9 82.3			104.9 105.2	П
36	33.3	13.8	96	88.7	36.7	56	144.1	59.7	16	199.6	82.7	76	255 .0	105.6	П
37	34.2	14.2	97	89.6			145.0		17	200.5 201.4				106.0	
38 39	35.1 36.0	14.5 14.9	98		37.5 37.9		146.0 146.9	60.5 60.8		202.3	83.4 83.8			106.4	Н
40	37.0	15.3	100				147.8	61.2		203.3	84.2			107.2	
41	37.9	15.7	101	93.3	38.7		148.7	61.6		204.2	.84.6			107.5	11
42	38.8	16.1	02		39.0		149.7			205.1	85.0			1107.9	11
4 3 4 4	39.7 40.7	16.5 16.8	03 04	95.2 96.1	39.4 39.8		150.6 151.5	62.4 62.8		206.0 206.9	85.3 85.7			108.3 108.7	П
45	41.6	17.2	05		40.2		152.4	63.1	25	207.9	86.1	85	263.5	109.1	П
46	48.5	17.6	06		40.6		153.4	63.5		208.8	86.5			109.4	
47 48	43.4 44.3	18.0 18.4	07	98.9 99.8	40.9 41.3		154.3 155.2	63.9 64.3		209.7 210.6	86.9 87.3			109.8 110.2	
49	45.3	18.8		100.7	41.7		156.1	64.7	29	211.6	87.6			110.6	
50	46.2	19.1		101.6	42.1		157.1	65.1		212.5	88.0	90	267.9	111.0]
51	47.1	19.5		102.6	42.5		158.0	65.4		213.4	88.4			111.4	
52 53	48.0 49.0	19.9 20.3		103.5 104.4	42.9 43.2		158.9 159.8	65.8 66.2		214 3 215.3	88.8 89.2			111.7	
54	49.9	20.7	14	105.3	43.6	74	160.8	66.6	34	216.2	89.5	94	271.6	112.5	
55	50.8	21.0	15	106.2	44.0	75	161.7	67.0	35	217.1	89.9	95	272.5	112.9	1
56 57	51.7 52.7	21.4		107.2 108.1	44.4		162.6 163.5	67.4 67.7		218.0 219.0	90.3 90.7	96	273.5	113.3	1
58	53.6	22.2		109.0	45.2		164.5	68.1		219.0 219.9	91.1			114.0	1
59	54.5	22.6	19	109.9	45.5	79	165.4	68.5	39	220.8	91.5	99	276.2	114.4	1
60	55.4	23.0		110.9			166.3	68.9		221.7	91.8			114.8	4
	Dep.	Lat.		<u> </u>	Lat.		Dep.			Dep.					_
	₩.L.T.E.,			e.s.e.		W	N.W.		V	V.S.W.		Į.	OFO	Points.	

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Difference of Latitude and Departure for 21 Points.

	N.	N.E.	E.	N.	N.W.	w.	s	.S.E.	E.		.s.w. ₄	W.			
Dist.	Lat.	Dep.		Lat.	Dep.		Lat.	Dep.		~	Dep.		Lat.		_
1 2	00.9 01.8	00.4	61 62	55.1 56.0	26.1 26.5		$109.4 \\ 110.3$	51.7 52.2		163.6 164.5	77.4 77.8		217.9 218.8		
1 3		01.3	63	57.0	26.9		111.2	52.6		165.4	78.2		219.7		
4		01.7	64	57.9	27.4		112.1	53.0		166.3	78.7	44	220.6	104.3	3
5	04.5		65	58.8			113.0	53.4		167.2	79.1		221.5		
6 7	05.4 06.3	02.6 03.0	66 67	59.7 60.6			113.9 114.8	53.9 54.3		168.1 169.0	79.5 80.0		22 2 . 4 223 . 3		
8	07.2	03.4	68	61.5	29.1		115.7	54.7		169.9	80.4		224.2		
9	08.1	03.8	69	62.4	29.5		116.6	55.2	89	170.9	8.03	49	2 25 . I	106.5	5
10		04.3	70	63.3	29.9		117.5	55 .6		171.8	81.2		226.0		I
11	09.9	04.7	71	64.2	30.4		118.4	56.0		172.7	81.7		226.9		
12	10.8 11.8	05.1 05.6	72 73	65.1 66.0	30.8 31.2		119.3 120.2	56.4 56.9		173.6 174.5	82.1 82.5		227.8 228.7		
14	12.7	06.0	74	66.9	31.6		121.1	57.3		175.4	82.9		229.6		
15	13.6	06.4	75	67.8	32.1		122.0	57.7		176.3	83.4		230.5		
16	14.5	06.8	76	68.7	32.5		122.9	58.1		177.2	83.8		231.4		
17	15.4 16.3	07.3 07.7	77 78	69.6 70.5	32.9 33.3		123.8 124.8	58.6 59.0		178.1 179.0	84.2 84.7		232.3 23 3 .2		
19			79	71.4	33.8	39	125.7	59.4		179.9	85.1		234.1		
20	18.1	08.6	. 80	72.3	34.2	40	126.6	59. 9	200	180.8	85.5	60	235.0	111.9	2
21	19.0	09.0	81	73.2	34.6		127.5	60.3		181.7	85.9		235.9		
22	19.9	09.4	82	74.1	35.1		128.4	60.7		182.6	86.4		236.8		
24	20.8 21.7	09.8 10.3	·83	75.0 75.9	35.5 35.9		129.3 150.2	61.1		183.5 184.4	86.8 87.2		2 37 . 7 238 . 7		
25	22.6	10.7	85	76.8	36.3	45	131.1	62.0		185.3			239.6		
26	23.5	11.1	86	77.7	36 .8		132.0	62.4	06	186.2	88.1	66	240.5	113.7	7
27 28	24.4	11.5	87	78.6			132.9	62.9		187.1	88.5		241.4		
29	25.3 26.2	12.0 12.4	88 89	79.6 80.5	37.6 38.1		133.8 134.7	63.3 63.7		188.0 188.9	88.9 89.4		242.3 243.2		
30	27.1	12.8	90	81.4	38.5		135.6	64.1		189.8	89.8		244.1		
31	23.0	13.3	. 91	82.3	38.9	151	136.5	64.6	211	190.7	90.2	271	245.0	115.9	5
32	28.9	13.7	92	83.2	39 .3		137.4	65.0		191.6	90.6		245.9		
33 34	29.8 30.7	14.1	93 94	84.1	39.8 40.2		138.3 139.2	65.4 65.8		192.5 193.6	91.1 91.5		246 . 8 247 7		
35	31.6	15.0	95	35.0 85.9	40.6		140.1	66.3		194.4	91.9		248.ò		
36	32.5	15.4	96	86.8	41.0		141.0	66.7		195.3	92.4	76	249.5	118.0	0
37	33.4	15.8	97	87.7	41.5		141.9	67.1		196.2	92.8		250.4 251.3		
39	34.4 35.3	16.2 16.7	98 99	88.6 89.5	41.9 42.3		142.8 143.7	67.6 68.0		197.1 198.0	93.2 93.6		252.2		
40		17.1	100	90.4	42.8		144.6	68.4		198.9	94.1		253.1		
41	37.1	17.5	101	91.3	43.2	161	145.5	68.8		199.8	94.5		254.0		
42		18.0	02	92.2			146.4	69.3		200.7	94.9		254.9		
43 44	38.9 39.8	18.4 18.8	03 04	93.1 94.0	44.0 44.5		147.4 148.3	69.7 70.1		201.6 202.5	95.3 95.8		255 . 8 256 . 7		
45	40.7	19.2	05	94.9	44.9		149.2	70.5		203.4	96.2		257.6		
46	41.6	19.7	06	95.8	45.3	66	150.1	71.0	26	204.3	96.6		258.5		
47	42.5	20.1	07	96.7	45.7		151.0	71.4		205.2	97.1		259.4 260.3		
48 49	43.4 44.3	20.5 21.0	08 09	97.6 98.5	46.2 46.6		151.9 152.3	71.8 72.3	20	206.1 207.0	97.5 97.9		260.3 261 .3		
50	45.2	21.4	10	99.4	47.0		153.7	72.7		207.9	98.3		262.2		
51	46.1	21.8		100.3	47.5	171	154.6	73.1		208.8	98.8		263.1		
52	47.0	22.2	12	101.2	47.9		155.5	73.5		209.7	99.2		264.0		
53 54	47.9	22.7 23.1		102.2 103.1	48.3 48.7		156.4 157.3	74.0 74.4		210.6 211.5	99.6 100.0		264.9 265.8		
55	49.7	23.5		104.0	49.2		158.2	74.8	35	212.4	100.5	95	266.7	126.1	ı
56	50.6	23.9	16	104.9	49.6	76	159.1	75.2	36	213.3	100.9		267.6		
57 58	51.5 52.4	24.4 24.8		105.8 106.7	50.0 50.5		160.0 160.9	75.7 76.1		214.2 215.1			268.5 2 69.4		
59	53.3	25.2		106.7	50.9		161.8	76.5			102.2		270 .3		
60	54.2	25.7		108.5	51.3		162.7	77.0	40	217.0	102.6		27 1.2		
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.		Dep.				Lat.				
N.E	i.b.E.₫	E.	S.	E.b.E	.}E.	N.	W.b.W	. W.	S.	W.b.W	.4W.	[Fo	r 54	roint	В.

TABLE L

Difference of Latitude and Departure for 21 Points.

l		N	N.E.4	E.	N	.N.W.	w.		S.S.E.	ξE.	s	.s.w.	ιw.			
ī)ist.		Dep.					Lat.	Dep.	Dist.	Lat.	Dep.			Dep.	ī
11	ı	00.5	00.5	61	53.0			106.3				85.3	241	212.5	113.6	1
П	2						22	107.0	57.5		60.5		42	213.4	114.1	I
Н	3							108 .4				86.3 86.7	44	214.3	114.5 115.0	Ì
П	4							110.9			63.2				115.5	١
П	6							111.			64.0				116.0	ı
П	7					1 31.0	27	112.0	59.9		64.9		47	217.8	116.4	l
П	8							112.9			65.8		48	218.7	116.9	١
П	9							113.3			66.7 67.6				117.4 117.8	١
-	10			!!		. :		114.6				·				1
Ш	11			71			131	116.4			69.3				118.3	ı
П	12 18			73		34.4	33	117.3			70.2		53	223.1	119.3	l
Ш	14					3 34.9	lii 34	(18.2	1 .		71.1		54	224.0	119.7	l
П	10						35	119.1			72.0				120.2	١
Н	16					35.8	36		64.1		72.9				120.7	١
П	17							120.8			73.7 74.6				121.1 121.6	l
н	18 19							122.6			75.5				122.1	l
Н	20							123.					60	229.3	122.6	ı
-	21			il	·i		· !	124.4		- ,					123.0	۱
П	22							125.9			78.1				123.5	l
Н	23			il			1	126.			79.0				124.0	i
Н	24				74.1			127.0			79.9		64	232.8	124.4	١
Н	25					40.1		127.9			80.8				124.9	l
	26							128.8			81.7				125.4	1
П	27 28			87 88				129.6			82.6 83.4	97.6 98.1			125.9 126.3	l
П	20 29		1					131.4			84.3				126.8	l
П	30		1	90				132.5		10	85.2				127.3	ì.
-	31	27.3		91			·	133.2		211 1	86.1	99.5	:		127.7	l
П	32			92				134.1				99.9			128.2	ı
	33		15.6	93				134.9				100.4			128.7	ı
	34							135.8				100.9			129.2	l
П	35			95				136.7				101.4			129.6	ĺ
	36 37	31.7 32.6		96 97				137.6 138.5				101.8 102.3			130.1 130.6	ł
	38			98				139.5				102.8			131.0	ı
	39	34.4		99				140.2		1911	93.1	103.2			131.5	l
	4 G	35.3	18.9	100	88.2	47.1	60	141.1	75.4			103.7	80	246 . 9	132.0	ľ
Γ	41	36.2	19.3	101	89.1	47.6	161	142.0	75.9	221 19	94.9	104.2	281	247.8	132.5	l
	42	37.0	19.8	02	90.0			142.9				104.7			132.9	
1	43	37.9	20.3	03				143.8				105.1			133.4	
1	44	38.8	20.7 21.2	04	91.7			144.6				105.6			133.9	ı
1	45 46	39.7 40.6	21.7	05 06	92.6 93.5			145.5 146.4	77.8 78.3			106.1 106.5			134.3 134.8	ı
. [47	41.5	22.2	07	94.4			147.3				107.0			136.3	ı
	48	42.3	22.6	08				148.2	79.2			107.5			135.8	
1	49	43.2	23.1	09	96.1	51.4	69	149.0	79.7			107.9	89	254 .9	136.2	
L	50	44.1	23.6	10	97.0	51 .9	70	149.9	80 . 1	30 20	02.8	108.4	90	255.8	136.7	
Γ	51	45.0	24.0	111	97.9	52.3		150.8	80.6	231 20					137.2	H
	52	45.9	24.5	12	98.8	52.8		151.7	81.1			109.4			137.6	
1	53	46.7	25.0	13	99.7 100.5	53.3		152.6	81.6			109.8			138.1	Н
	54 55	47.6 48.5	25.5 25.9		100.5	53.7 54.2		153.5 154.3	82.0 82.5			110.3 110.8			138.6 139.1	I
	56	49.4	26.4		102.3	54.7		155.2	83.0			111.2			139.5	
	57	50.3	26.9		103,2	55.2		156.1	83.4			11.7			140.0	Н
1	58	51.2	27.3	18	104.1	55.6	78	157.0	83.9	38 20	9.9	12.2	98	262.8	140.5	
1	59	52.0	27.8		104.9	56.1		157.9	84.4	3921					140.9	
-	60	52.9	28.3		105.8	56.6		158.7	84.9			13.1			141.4	
Ш	15t.	Dep.	Lat.					Dep.		Dist. D						1
_	N.E	.b.E.	E.	8.	E.b.E.,	E	N.V	V.b.W	.JW.	S.W	.b.W	₩.	[Fo	r 5g]	Points.	

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TABLE I.

Difference of Latitude and Departure for 24 Points.

L		N	N.E.4	E.	N.	N.W.	W.	S.	S.E.4E		5.5	.W.3W	7.			
Γ	Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Der	D.
	1	00.9	00.5	61	52.3		121	103.8	62.2		155.2	93.1	241	206.7		
	2							104.6			156.1	93.6	42	207.6	124.	.4
L	3							105.5			157.0	94.1		208.4		
	4	1			1 .			106.4			157.8	94.6		209.3		
l					55.8			107.2			158.7	95.1		210.1		
i i	1				56.6			108.1		80	159.5	95.6		211.0		
ľ	١				57.5			108.9 109.8			16 0.4 161.3	96.1 96.7		211.9 212.7		
Į	1. 3							110.6			162.1	97.2		213.6		
ı	10			70				111.5			163.0			214.4		
1	11			71	60.9		'——	112.4		11		!	1	215.3	l	_
ı	12			72	61.8			113.2			163.8 164.7	98.2 98.7		216.1		
1	13			73	62.6			114.1			165.5	99.2		217.0		
H	14			74	63.5			114.9			166.4			1		
I	15			75	64.3	38.6		115.8				100.5		218.7		
H	16			76	65.2	39.1		116.7				100.8		219.6		
J	17	14.6	08.7	77	66.0		37	117.5	70.4			101.3	57	220.4	132.	1 :
1	18				66.9	40.1		118.4		98	169.8	101.8		221.3		
П	19			79	67.8	1		119.2				102.3		222.2		
	20			80	68.6	41.1	40	120.1	72.0	200	171.5	102.8	60	223.0	133.	7
	21			81	69.5	41.6		120.9		201	172.4	103.3		223.9		
H	22			82	70.3	42.2		121.8				103.8		224.7		
П	23			83	71.2	42.7		122.7			174.1			225.6		
П	24			84	72.0	43.2		123.5	74.0		175.0			226.4		
	25			85	72.9	43.7		124.4			175.8			227.3		
П	26 27	22.3 23.2		86 87	73.8	44.2		125.2			176.7			228.2		
П	28			88	74.6 75.5	44.7 45.2		126.1 126.9	75.6 76.1		177.5			229.0 229.9		
1	29	24.9		89	76.3	45 8		127.8	76.6		178.4 179.3			230.7		
H	30	25.7	15.4	90	77.2	46.3		128.7	77.1		180.1			231.6		
ŀ	31	26.6	15.9	91	78.1	46.8		129.5	77.6							
П	32	27.4	16.5	92	78.9	47.3		129.8 130.4	78.1		181.0 181.8		79	232. 4 23 3.5	139.	3 !
Н	33	28.3		93	79.8	47.8		131.2	78.7		182.7		73	234.2	140	4
П	34	29.2	17.5	94	80.6	48.3		132.1	79.2		183.6		74	235.0	140	9 :
П	35	30.0	18.0	95	81.5	48.8		132.9	79.7		184.4			235.9		
П	36	30.9	18.5	96	82.3	49.4		133.8	80.2		185.3			236.7		
	37	31.7	19.0	97	83.2	49.9		134.7	80.7	17	186.1	111.6	77	2 37 . 6 l	142	4
	38	32.6	19.5	98	84.1	50.4		135.5	81.2		187.0			238.4		
1	39	33.5	20.1	99	84.9	50.9		136.4	81.7		187.8			239.3		
1-	40	34.3	20.6	100	85.8	51.4		137.2	82.3		188.7			240.2		
1	41	35.2	21.1	101	86.6	51.9		138.1	82.8		189.6			241.0		
1	42	36.0	21.6	02	87.5	52.4		139.0	83.3		190.4			241.9		
1	43	36.9	22.1	03	88.3	53.0		139.8	83.8		191.3			242.7		
1	44	37.7 38.6	22.6 23.1	04 05	89.2 90.1	53.5 54.0		40.7 41.5	84.3 84.8		192.1 193.0			243.6 244.5		
1	46	39.5	23.6	06	90.9	54.5		42.4	85.3		193.8			245.3		
1	47	40.3	24.2	07	91.8	55.0		43.2	85.9		194.7			246.2		
1	48	41.2	24.7	08	92.6	55.5		44.1	86.4		95.6			247.0		
1	49	42.0	25.2	09	93.5	56.0		45.0	86.9		96.4			247.9		
ı	50	42.9	25.7	10	94.4	56.6		45.8	87.4		97.3			248.7		
1	51	43.7	26.2	111	95.2	57.1	1711	46.7	87.9	2311	98.1	18.8		249.6		
	52	44.6	26.7	12	96.1	57.6		47.5	88.4		99.0			250.5		
ı	53	45.5	27.2	13	96.9	58.1		48.4	88.9		99.9			251.3		
l	54	46.3	27.8	14	97.8	58.6	74 1	49.2	89.5	34	200.7	20.3		252.2		
	55	47.2	28.3	15	98.6	59.1		50.1	90.0		01.6			253.0		
	56	48.0	28.8		99.5	59.6		51.0	90.5		02.4			253.9		
	57	48.9	29.3		00.4	60.2		51.8	91.0		03.3			254.7		
-	58	49.7	29.8		01.2	60.7		52.7	91.5		04.1			255.6		
1		50.6 51.5	30.8		02.1	61.2		53.5	92.0		05.0			256.5 257.3		
-	!-					61.7		54.4	92.5							_ •
		Dep.		Dist. I		Lat.								Dep.		
_	N.E.	b.E.41	<u>ن.</u>	5.E	.b.E.}	Ľ.	N.V	V.b.W	₫W.	S.W	.b.W.	W.	$\subset \mathbb{R}$	or 51	roin	15.

Difference of Latitude and Departure for 3 Points.

	N	I.E.b.N	i.·	N	.W.b.N	ī.		5.E.b.S			.W.b.S				
Dist.				Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	J
1	00.8	00.6	61	50.7	38.9		100.6	67.2		150.5				133.9	
2	01.7	01.1	62	51.6	34.4		101.4	67.8		151.3 152.2				134.4 135.0	
3	02.5		63	52.4	35.0		102.3 103.1	68.3 68.9		153.0				135.5	
4 5	03.3 04.2		64 65	53.2 54.0	35.6 36.1		103.1			153.8				136.1	
6	05.0			54.9	36.7		104.8	70.0			103.3	46	204.5	136.7	J
7	05.8			55.7	37.2		105.6			155.5				137.2	- 1
8	06.7			56 .5	37.8		106.4			156.3				137.8 138.3	
9	07.5			57.4	38.3		107.3 108.1	71.7 72.2		15 7.1 158.0				138.9	
10	08.3	05.6	70	58.2	38.9									139.4	- 1
11	09.1	06.1	71	59.0	39.4		108.9 109.8	72.3 73.3		158.8 159. 6				140.0	ı
12 13	10.0 10.8		72	59.9 60.7	40.0 40.6		110.6	73.9		160.5				140.5	ı
14	11.6			61.5			111.4	74.4		161.3		54	211.2	141.1	
15	12.5	08.3	75	62.4	41.7		112.2			162.1				141.7	: 1
16	13.3			63.2	42.2		113.1	75.5		163.0				142.2	. 1
17	14.1			64.0	42.8		113.9 114.7	76.1 76.7		163.8 164.6				143.3	. 1
18 19	15.0 15.8		73 79	64.9 65.7	43.3 43.9		115.6	77.2		165.5				143.9	
20	16.6		80	66.5	41.4		116.4			166.3				144.4	
21	17.5		81	67.3	45.0	i	117.2	78.3	7	167.1	111.7	261	217.0	145.0	-
22	18.3			68.2	45.6		118.1	78.9		168.0				145.5	
23	19.1	12.8	83	69.0	46.1		118.9	79.4			112.8			146.1	
24	20.0		81	69.8	46.7		119.7	80.0	1	169.6				146.7	
25	20.8		85	70.7	47.2		120.6	80.5	1 1	170.5	113.9			147.2 147.8	
26 27	21.6 22.4		86 87	71.5 72.3	47.8 48.3		121.4 122.2	81.1 81.7			115.0			148.3	
23	23.3		88	73.2	48.9		123.1	82.2		172.9				148.9	
29	24.1	16.1		74.0	49.4	49	123.9	82.8		1 73 .8				149.4	
30	24.9	16.7	90	74.8	50.0	50	124.7	83.3	10	1 74 .6	M6.7			150.0	-' 1
31	25.8	17.2	91	75.7	50.6		125.6	83.9		175.4				150.5	11
32	26.6		92	76.5	51.1		126.4	84.4		176.3				151.1	П
3 3	27.4		93	77.3	51.7		127.2 128.0	85.0		177.1 177.9	118.3			151.7 152.2	
34 3 5	28.3 29.1		94 95	78.2 79.0	52.2 52.8		128.9	85.5 86.1			119.4			152.8	- 1
36	29.9			79.8	53 .3		129.7				120.0	76	229.5	153.3	
37	30.8		97	89.7	53.9		130.5				120.5			153.9	
38	31.6		98	81.5	54.4		131.4			181.3				154.4 155.0	
39 40	32.4 33.3		99 100	82.3 83.1	55.0 55.6		132.2 133.0	88.3 88.9		182.1 182.9				155.5	
		22.8					133.9			183.8		1		156.1	-: 1
41 42	34.1 34.9		101	84.0 84.8	56.1 56.7		134.7	90.0			123.3			156.7	
43	35.8			85.6	57.2		135.5				123.9			157.2	
44	36.6		04	86.5	57.8	64	136.4	91.1	24	186.2	124.4	84	236.1	157.8	: []
45	37.4		05	87.3	58.3		137.2	91.7			125.0			158.3	
46	38.2		06	88.1	58.9		138.0				125.5 126.1			158.9 159.4	
47 48	39.1 39.9		07	89.0 89.8	59.4 60.0		138.9 139.7				126.7			160.0	
49	40.7	27.2	09	90.6	60.6		140.5			190.4		89	240.3	160.5	į
50	41.6	27.8	10	91.5	61.1	70	141.3	94.4	30	191.2	127.8	90	241.1	161.1	_
51	42.4	28.3	111	92.3	61.7	171	142.2	95.0	231	192.1	128.3			161.7	
52	43.2	28.9	12	93.1	62.2		143.0	95.5	4		128.9			162.2	
53	44.1	29.4	13	94.0	62.8		143.8	96.1			129.4			162.8	
54 55	44.9 45.7	30.0 30.6	14 15	94.8 95.6	63.3 63.9		144.7 145.5	96.7 97.2		194.6 195.4	130.0			163. 3 163.9	
56	46.6		16	96.5	64.4		146.3			196.2				164.4	
57	47.4	31.7	17	97.3	65.0		147.2	98.3	37	197.1	131.7	97	246.9	165.0) . 1
58	48.2	32.2	18	98.1	65.5	78	148.0	98.9	38	197.9	132.2			165.5	
59	49.1	32.8	19	98.9	66.1		148.8	99.4		198.7				166.1	
60	49.9	33.3	20	99.8	66.7			100.0		199.6				166.7	4;
Dist.	Dep.	Lat.													
	.E.b.	<u> </u>		6.E.b.I	<u>ن</u> .	N.	W.b.V	٧	S	.W.b.1	V. _{Digit}	izad J	hor 5	Points	<u>-</u>

Difference of Latitude and Departure for 31 Points.

	N.	E.¶N		N	.w. _{\$} N	í.	1	S.E.4S.		8	5.W. § S		
Dist.	Lat. I	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat. 1	Эер.
1		00.6	61	49.0	36.3	121	97.2	72.1		145.4		241 193.61	
2		01.2	62	49.8	36.9	22	98.0			146.2		42 194 .41	
3		01.8 02.4	63 64	50.6 51.4	37.5 38.1	23 24	98.8 99.6			147.0 147.8		43 195.21 44 196.01	
5		03.0	65	52.2	38.7		100.4			148.6		45 196 .81	
6		03.6	66	53.0	39.3		101.2	75.1		149.4		46 197.6	46.5
11 7		04.2	67	53.8	39.9		102.0	75.7		150.2		47 198.41	
8 9		04.8 05.4	68 69	54.6 55.4	40.5		102.8 103.6	76.2 76.8		151.0 151.8		48 199 .2 1 49 200 .0 1	
10		06.0	70	56.2	41.7		104.4	77.4		152.6		50 200.81	
		06.6	71	57.0	42.3		106.2	78.0		153.4		251 201 .61	
12		07.1	72	57.8	42.9	-	106.0			154.2		52 202 .4	
13		07.7	73	58.6	43.5		106.8	79.2		155.0		53 203.21	
14		08.3	74	59.4	44.1		107.6	79.8		155.8		54 204 . 0 1	
15 16		08.9 09.5	75 76	60.2	44.7		108.4	80.4		156.6		55 204 .8 1 56 205 .6 1	
17		10.1	77	61.8	45.3 45.9		109. 2 110.0			157.4 158.2		57 206 .4 1	
ii		10.7	78	62.7	46.5		110.8			159.0		58 207.21	
19	15.3	11.3	79	63.5	47.1	39	111.6	82.8	99	159.8	118.5	59 208 .01	54.3
20		11.9	80	64.3	47.7	_	112.4	83.4		160.6		60 208.8 1	
21		12.5	81	65.1	48.3		113.3	84.0		161.4		261 209 . 6 1	
22		13.1 13.7	8 2 83	65.9	48.8		114.1	84.6		162.2		62 210.41 63 211.21	
24		14.3		66.7 67.5	49.4 50.6		114.9 115.7			$163.1 \\ 163.9$		64 212.01	
25		14.9	85	68.3	50.6		116.5			164.7		65 212.81	
26	20.9	15.5	86	69.1	51.2		117.3		06	165.5	122.7	66 213.71	
27		16.1	87	69.9	51.8		118.1	87.6		166.3		67 214 .51	
28 29		16.7 17.3	88 89	70.7 71.5	52.4 53.0		118.9 119.7			167.1 167.9		68 215.31	
30		17.9	90	72.3	53 .6		120.5	88.8 89.4		168.7		69 216.1 1 70 216.9 1	
31		18.5	91	73.1	54.2		121.3	90.0	211	169.5		271 217.71	
32		19.1	92	73.9	54.8		122.1	90.5		170.3		72 218.51	
33		19.7	93	74.7	55.4		122.9		13	171.1	126.9	73219.31	
34		20.3	94	75.5	56.0		123.7			171.9		74 220 . 1 1	
35 36		20.8 21.4	95 96	76.3	56.6 57.2		124.5 125.3			172.7 173.5		75 220.91 76 221.71	
37		22.0	97	77.1	57.8		126.1			174.3		77 222.51	
38		22.6	98	78.7	58.4		126.9			175.1		78 223.31	
39		23.2	99	79.5	59.0		127.7			175.9		79 224 . 1 1	
40		23.8	100	80.3	59.6		128.5			176.7		80 224 . 9 1	
41	,	24.4	101	81.1	60.2		129.3		221	177.5		281 225 .71	
42		25.6 25.6	02	81.9 82.7	60.8 61.4		130.1 130.9			178.3 179.1		82 226.5 1 83 227.3 1	
44		26.2	04	83.5	62.0		131.7			179.9		84 228.11	
45	36.1	26.8	05	84.3	62.5		132.5	98.3	25	180.7	134.0	85 228.91	69.8
46		27.4	06	85.1	63.1		133.3				134.6		
47		28.0 28.6	07 08	85.9 86.7	63.7 64.3		134.1	99.5 100.1		182.3 183.1		87 230.5 1 88 231.3 1	
49	39.4	29.2	09	87.5	64.9			100.1		185.9		89 232 . 1 1	
50		29.8	110	88.4	65.5			101.3	30	134.7		90 232 . 9 1	
51		30.4	111	89.2	66.1			101.9	231	185.5	137.6	291 233.71	
52	41.8	31.0	12	90.0	66.7	72	138.2	102.5	32	186.3	138.2	92 234 . 5 1	73.9
53		31.6	13	90.8	67.3		139.0			187.1		93 235 . 3 1	
54 55		32.2 32.8	14 15	91.6	67.9 68.5			103.7 104.2	34 35	188.0	139.4 140.0	94 236.1 1 95 236.9 1	
56		33.4	16	92.4 93.2	69.1	76		104.8	36			96 237 . 7 1	
57		34.0	17	94.0	69.7			105.4		190.4		97 238 . 6 1	
58	46.6	34.6	18	94.8	70.3	78	143.0	106.0	38	191.2	141.8	98 239 .4 1	
59		35.1	19	95.6	70.9			106.6		192.0		99 240.2 1	
60		35.7	20	96.4	71.5			107.2		192.8		300 241 .0 1	
Dist.		Lat.									INIL.	Dist. Dep.	LET.
	I.E. E.		S	E#E		1	i.W. į V	٧.		W.W	تكبط	For 44 P	DINTS.

TABLE I.

Difference of Latitude and Departure for 31 Points.

<u> </u>		N.E.J	N.	ľ	ı.w.4ı	₹.		S.E.1S			5. W. 45	3 .			
Dist	. Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	<u>.</u>
				47.2	38.7	121	93.5	76.8			114.8		186.3		
	01.5			47.9	39.3	22	94.3				115.5		187.1		
				48.7	40.0	23	95.1	78.0		141.5			187.8		
11 2	1 22 2			49.5 50.2	40.6 41.2	24 25	95.9 96.6				116.7 117.4		188.6 189.4		
11 8				51.0	41.9		97.4				118.0		190.2		
1 7			67	51.8	42.5	27	98.2				118.6		190.9		
8	06.2	05.1	68	52.6	43.1	28	98.9	81.2			119.3		191.7		
] , 9				53.3	43.8	29	99.7	81.8			119.9	49			
10	07.7	06.3	70	54.1	44.4	30	100.5	82.5			120.5	50	193.3	158.6	5
11				54.9	45.0	131	101.3				121.2	251	194.0		
1 12				55.7	45.7		102.0				121.8		194.8		
1 15				56.4	46.3		102.8				122.4		195.6		
14				57.2	46.9		103.6				123.1		196.3 197.1		
16				58.0 58.7	47.6 48.2		104.4 105.1	85.6 86.3			123.7 124.3		197.9		
1 17				59.5	48.8	1	105.1	86.9			125.0		198.7		
l is				60.3	49.5		106.7				125.6		199.4		
19				61.1	50.1		107.4				126.2		200.2		
20				61.8	50.8		108.2	88.8			126.9		201.0		
21	16.2	13.3	81	62.6	51.4	141	109.0	89.4	201	155.4	127.5	261	201.8	165.6	ៅ
22				63.4	52.0		109.8	90.1		156.1		62	202.5	166.5	įΙ
2.5	17.8	14.6	83	64.2	52.7		110.5	90.7			128.8		203.3		
24				64.9	53 .3	44	111.3			157.7	129.4		204.1		
20				65.7	53 .9		112.1	92.0		158.5			204.8		
1 26				66.5	54.6		112.9			159.2			205.6		
27			87	67.3	55.2		113.6	93.3		160.0			206.4		
28				68.0	55.8		114.4	93.9			132.0		207.2		
30				68.8	56.5		115.2			161.6			207.9 208.7		
1 1				69.6	57.1		116.0		-		133.2				_
31			111	70.3	57.7		116.7	95.8		163.1 163.9			209.5 210.3		
33				71.1 71.9	58.4 59.0		117.5 118.3	96.4 97.1		164.7			211.0		
34				72.7	59.6		119.0	97.7			135.8	74	211.8	173.8	i
35				73.4		1	119.8	98.3			136.4		212.6		
36	27.8	22.8	96	74.2	60.9		120.6	99.0			137.0	76	213.4	175.1	ı
37				75.0	61.5	57	121.4	99.6	17	167.7	137.7	77	214.1	175.7	1
38				75.8	62.2		122.1	100.2			138.3		214.9		
39				76.5	62.8		122.9			169.3			215.7		
40			1	77.3	63.4	'I	123.7		1	170.1			216.4		
41				78.1	64.1		124.5		221		140.2	281	217.2	178.3	3
49				78.8	64.7		125.2				140.8		218.0		
1 4				79.6 80.4	65.3 66.0		126.0 126.8			173.2	141.5		218.8 219.6		
45				81.2	66.6		127.5			173.2			220.3		
46				81.9	67.2		128.3				143.4		221.1		
47				82.7	67.9		129.1				144.0		221.9		
48	37.1		08	83.5	68.5	68	129.9	106.6	28	176.2	144.6	88	2 22 .6	182.7	, l
49			09	84.3	69.1	69	130.6	107.2	29	177.0	145.3		223.4		
50	38.7	31.7	10	85.0	69 .8	70	131.4	107.8	30	177.8	145.9	90	224 .2	184.0	<u> </u>
51	39.4	32.4	111	85.8	70.4		132.2			178.6			224.9		
52	40.2		12	86.6	71.1		133.0			179.3			225.7		
53	41.0		13	87.4	71.7		133.7			180.1			226.5		
54 55	41.7			88.1 88.9	72.3 73.0		134.5 135.3			180.9 181.7	148.4		227.3 228.0		
56	43.3		16	89.7	73.6		136.0			182.4			228.8 22 8.8		
57	44.1	36.2	17	90.4	74.2		136.8			183.2			229.6		
58	44.8		18	91.2	74.9		137.6			184.0			230 . 4		
59	45.6		19	92.0	75.5		138.4			184.7			231 . 1		
60	46.4	38.1	20	92.8	76.1		139.1			185.5			231.9		
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	\neg
	N.E.	E.		EJE			W.W			W.JW			OF 44		_
							_					45	V/ V//	6	_

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Difference of Latitude and Departure for 31 Points.

1		1	n.e.in	ī.	1	N.W. <u>‡</u>	19 .		S.E.3	3. ·		5.W. <u>1</u>	8 .
	Di	Lat.	Dep.	Dist	Lat.	Dep.	Dist.			Dist.	La	t. Dep.	Dist. Lat. Dep.
	!	1 00.7			45.9			89.7				1 121.0	
		2 01.5										9 122.9	
	1	3 02.2 4 03.0					23 24					6 122 . 9 3 123 .	
		5 03.7			1							1 124	
		6 04.4							84.6			8 124.	
		7 05.2	04.7	67	49.6	45.0	27	94.1	85.3	87	1 3 8.	6 125.6	47 183.0 165.9
		8 05.9										3 126.3	
	١,	9 06.7 0 07.4		69 70								0 1 26 .9 8 1 27 .6	
				71		!						5 123.3	
-		1 08.2 2 08.9										3 128.9	
		3 09.6										0 129.6	
1		4 10.4				49.7	34					7 130.5	
1		5 11.1		75				100.0				5 131 .0	
1		6 11.9 7 12.6			56.3 57.1			100.8 101.5				2 131 .6 0 132 .3	
1		7 12.6 8 13.3		77 78	57.8			102.3				7 133.0	
١		9 14.1		79	58.5			103.0				133.6	
1	2	0 14.8	13.4	80			40	103.7				134.3	
1	2	1 15.6	14.1	81	60.0	54.4	141	104.5	94.7	201	148.	135.0	261 193.4 175.3
ł	2		14.8	82	60.8			105.2				135.7	
i	2		15.4	83	61.5	55.7		106.0				136.3	
I	2		16.1 16.8	84 85	62.2 63.0	56.4 57.1		106.7 107.4	96.7			137.0 137.7	
1	2		17.5	86	63.7	57. 8		108.2	97.4 98.0			138.3	
I	2			87	64.5	58.4		108.9	98.7			139.0	
1	21		18.8	88	65.2	59.1		109.7	99.4	08	154.1	139.7	68 198.6 180.0
ı	2:		19.5	89	65.9	59.8		110:4				140.4	
i	30		20.1	90	66.7	60.4		111.1		1		141.0	70 200 . 1 181 . 3
ļ	31		20.8	91	67.4	61.1		111.9				141.7	271 200 .8 182 .0
1	35		21.5	92 93	68.2 68.9	61.8 62.5		112.6 113.4				142.4	72 201.5 182.7 73 202.3 183.3
1	34		22.8	94	69.6	63.1		114.1				143.7	74 203 . 0 184 . 0
1	30		23.5	95	70.4	63.8		114.8				144.4	75 203.8 184.7
l	36		24.2	96	71.1	64.5	56	115.6	104.8			145.1	76 204.5 185.4
ı	3		24.8	97	71.9	65.1		116.3				145.7	
ł	39		26.2	98 99	72.6	65.8 66.5		117.1 117.8				146.4 147.1	78 206 .0 186 .7 79 206 .7 187 .4
L	44		26.9	100	74.1	67.2		118.6				147.7	80 207 .5 188 .0
1	47		27.5	101	74.8	67.8		119.3		_		148.4	281 208 .2 188 .7
П	41		28.2	02	75.6	68.5		20.0				149.1	82 208 .9 189 .4
П	43		28.9	03	76.3	69.2	63	120.8	109.5	23	165.2	149.8	83 209 .7 190 . 1
1	44		29.5	04	77.1	69.8		21.5				150.4	84 210.4 190.7
i	45		30.2 30.9	05	77.8	70.5		22.3				151.1	85 211.2 191.4
ı	46 47		31.6	06 07	78.5 79.3	71.2 71.9		23.0				151.8 152.4	86 211 .9 192 .1 87 212 .7 192 .7
	48		32.2	08	80.0	72.5		24.5				153.1	88 213.4 193.4
1	49		32.9	09	80.8	73.2		25.2				153.8	89 214.1 194.1
П	50		33.6	10	81.5	73.9		26.0				154.5	90 214.9 194.8
Н	51		34.2	111	82.2	74.5		26.7				155.1	291 215.6 195.4
П	52		34.9	12	83.0	75.2		27.4				155.8	
Н	53 54		35.6 36.3	13 14	83.7	75.9 76.6		28.2				156.5 157.1	93 217.1 196.8 94 217.8 197.4
	55		36.9	15	85.2	77.2		29.7				157.8	95 218.6 198.1
1	56	41.5	37.6	16	86.0	77.9	76 1	30.4	18.2	36	74.9	158.5	96 219.3 198.8
1	57		38.3	17	86.7	78.6	77 1	31.11	18.9			159.2	97 220 . 1 199 . 5
1	58		39.0	18	87.4	79.2		31.91				159.8	98 220.8 200.1
	59 60		39.6 40.3	19 20	88.9 88.9	79.9 80.6		32.6 1 33.4 1				160 · 5 161 · 2	99 221.5 200.8 300 222.3 201.5
1		Dep.											Dist. Dep. Lat.
-		N.E.JE.			E. E.			W.JW			W. 1V		[For 44Points.
				3			41.			.ن		• •	To or Adv ours.

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TABLE I.

Difference of Latitude and Departure for 4 Points,

		N.E.			N.W.			S.E.			s.w.		
Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.
1	00.7	00.7	61	43.1	43.1	121	85.6	35.6			128.0	241 170.4	
2	01.4	01.4		43.8	43.8	22	86.3	86.3	82	128.7	128.7 129.4	42 171 . 1 43 171 . 8	
3	02.1 02.8	02.1 02.8		44.5 45.3	44.5 45.3	23 24	87.0 87.7	87.0 87.7			130.1	44 172.5	
4 5	03.5	03.5	65	46.0	46.0	25	88.4				130.8	45 173.2	
6	04.2	04.2	66	46.7	46.7	26	89.1	89.1			131.5	46 173.9	
7	04.9	04.9	67	47.4	47.4	27	89.8	89.8			132.2	47 174.7 48 175.4	
8	05.7 06.4	05.7 06.4	68 69	48.1 48.8	48.1 48.8	28 29	90.5 91.2	90.5 91.2			132.9 133.6	49 176.1	
9 10	07.1	07.1	70	49.5	49.5	30	91.9	91.9			134.4	50 176.8	
11	07.8	07.8	71	50.2	50.2	131	92.6	92.6			135.1	251 177.5	177.5
12	08.5	08.5	72	50.9	50.9	32	93.5	93.3	92	135.8	135.8	52 178.2	
13	09.2	09.2	73	51.6	51.6	33	94.0	94.0			136.5	53 178.9	
14	09.9	09.9	74	52.3	52.3 53.0	34	94.8 95.5	94.8 95.5			137.2 137.9	54 179.6 55 180.3	
15 16	10.6 11.3	10.6	75 76	53.0 53.7	53.7	35 36	96.2	96.2			138.6	56 181.0	
17	12.0	12.0	77	54.4	54.4	37	96.9	96.9			139.3	57 181.7	
19	12.7	12.7	78	55.2	55.2	38	97.6	97.6			140.0	58 182 . 4	
19	13.4	13.4	79	55.9	55.9	39	98.3 99.0	98.3 99.0			140.7 141.4	59 183.1 60 183.8	
20	14.1	14.1	80	56.6	56.6	40			ł	142.1		261 184.6	
21 22	14.8 15.6	14.8 15.6	81 82	57.3 58.0	57.3 58.0	141	99.7 100.4	99.7 100.4			142.1	62 186.3	
23	16.3	16.3	83	58.7	58.7		101.1				143.5	63 186.0	186.0
24	17.Q	17.0	84	59.4	59.4			101.8			144.2	64 186.7	
25	17.7	17.7	85	60.1	60.1		102.5				145.0	65 187 .4	
26	18.4	18.4		60.8	60.8 61.5	46	103.2 103.9	103.2			145.7 146.4	66 188.1 67 188.8	188.8
27	19.1 19.8	19.1 19.8	87 88	61.5	62.2		104.7				147.1	68 189 .5	
29	20.5	20.5	89	62.9	62.9	49	105.4	105.4			147.8	69 190.2	190.2
30	21.2	21.2	90	63.6	63.6		106.1				148.5	70 190.9	
31	21.9	21.9	91	64.3	64.3	151	106.8	106.8			149.2	271 191.6	
32	22.6	22.6	92	65.1	65.1		107.5 108.2				149.9 150.6	72 192.3 73 193.0	
33 34	23.3 24.0	23.3 24.0	93 94	65.8 66.5	65.8 66.5		108.9				151.3	74 193.7	
35	24.7	24.7	95	67.2	67.2	55	109.6	109.6	15	152.0	152.0		
36	25.5	25.5	96	67.9	67.9			110.3			152.7	76 195.21 77 195.91	
37	26.2	26.2 26.9	97	68.6 69.3	68.6 69.3	57	111.7	111.0			153.4 154.1	78 196 .6	
38	26.9 27.6		98 99	70.0	70.0	59	112.4	112.4			154.9	79 197.3	
40	28.3		100	70.7	70.7	60	113.1	113.1	20	155.6	155.6	80 198.0	198.0
41	29.0	29.0	101	71.4	71.4	161	113.8	113.8			156.3	281 198.7	
42	29.7	29.7	02	72.1	72.1			114.6			157.0	82 199.4 83 200.1	
43	30.4	30.4	03	72.8	72.8 73.5			115.3 116.0			157.7 158.4	84 200.8	
44	31.1 31.8	31.1 31.8	04 05	73.5 74.2	74.2	65	116.7	116.7	25	159.1	159.1	85 201.5	201.5
46	32.5	32.5	66	75.0	75.0	66	117.4	117.4	26	159.8	159.8	86 202.2 87 202.9	202.2
47	33.2	33.2	07	75.7	75.7			118.1			160.5	87202.9	202.9
48	33.9	33.9 34.6	08 09	76.4 77.1	76.4 77.1			118.8 119.5			161.2 161.9	88 203.6 89 204.4	
49 50	34.6 35.4	35.4	10	77.8	77.8		120.2				162.6	90 205 . 1	
51	36.1	36.1	111	78.5	78.5			120.9			163.3	291 205.8	
52	36.8	36.8		79.2	79.2	72	121.6	121.6	32	164.0	164.0	92 206.5	
53	37.5	37.5		79.9	79.9			122.3			164.8	93 207 .2 94 207 .9	
54	38.2	38.2	14	80.6 81.3	80.6			123.0 123.7	34	166.2	165.5 166.2	95 208.6	
55 56	38.9 39.6	38.9 39.6	15 16	82.0	82.0	76	124.5	124.5	36	166.9	166.9	96 209.3	209.3
57	40.3	40.3		82.7	82.7	77	125.2	125.2	37	167.6	167.6	97 210.0	210.0 j
58	41.0	41.0		83.4	83.4	78	125.9	125.9			168.3	98 210.7 99 211.4	
59	41.7	41.7	19 20	84.1 84.9	84.1 84.9		126.6 127.3	126.6 127.3			169.0 169.7	300 212.1	
Dist.	42.4 Dep.	42.4 Lat.	1):=+	Den	T.e+	Diet	Den	Let				Dist. Dep.	
1718(.)	N.E.	Lull.	IUBL.	N.W.	JJGL.	(D'481.)	S.E.	- Adt. I	. 20 40 6.1	S.W.		For 4 l	oints.
			-		<u></u>								

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Difference of Latitude and Departure for 1 Degree.

l								F -			6-				_
Dist		Dep.	-					Dep.			Dep.			Dep.	-
		00.0	61	61.0	01.1		121.0			181.0	03.2		241.0	04.2	١
3		00.0 00.1	62 63	62.0 63.0	01.1 01.1		122.0 123.0	02.1 02.1		18 2 .0 183.0	03.2		242.0 243.0	04.2 04.2	1
11 4			64	64.0	01.1		124.0			184.0	03.2		244.0	04.3	1
2		1	65	65.0	01.1		125.0	02.2		185.0	03.2		245.0	04.3	1
11 e	06.0	00.1	66	66.0	01.2		126.0			186.0		46	246.0	04.3	ı
11 3			67	67.0	01.2		127.0	02.2		187.0	03.3		247.0	04.3	1
	8 08.0 9 09.0		68 69	68.0 69.0			128.0 129.0	02.2 02.3		188.0 189.0	03.3 03.3		248.0 24 9.0	04.3 04.3	1
1 1			70	70.0	01.2		130.0	02.3		190.0	03.3		250.0	04.4	1
			71	71.0	01.2		131.0	02.3		191.0	03.3		251.0	04.4	-
ll is		00.2	72	72.0	01.3		132.0			192.0			252.0	04.4	1
1:			73	73.0	01.3		133.0	02.3		193.0			253.0	04.4	1
14			74	74.0			134.0			194.0			254. 0	04.4	•
15		00.3	75	75.0			135.0	02.4		195.0			255.0	04.5	١
16		00.3 00.3	76 77	76.0 77.0	01.3 01.3		136.0 137.0	02.4 02.4		196.0 197.0	03.4 03.4		256.0 257.0	04.5 04.5	1
18			78	78.0	1		138.0	02.4		198.0			258.0	04.5	1
19			79	79.0			139.0			199.0			259.0		
20	20.0	00.3	80	80.0		40	140.0	02.4	200	200.0	03.5	60	260.0	04.5	١
21	21.0	00.4	81	81.0	01.4	141	141.0	02.5	201	201.0	03.5	261	261.0	04.6	1
22		00.4	82	82.0			142.0			202 .0			262.0	04.6	1
23			83	83.0			143.0			203.0			263.0	04.6	ì
24			84 85	84.0 85.0	01.5 01.5		144.0 145.0			204.0 205.0			264.0 265.0		1
26			86	86.0			146.0			205.0 2 06.0			266.0	04.6 04.6	-
27			87	87.0			147.0			207.0			267.0	04.7	1
28			88	88.0			148.0		08	208.0		68	268.0	04.7	ı
29			89	89.0			149.0	02.6		209.0	03.6		269.0	04.7	1
30		00.5	90	90.0	01.6		150.0	02.6	11	210.0	03.7	·	270.0	04.7	
31	31.0	00.5	91	91.0	01.6		151.0	02.6		211.0			271.0	04.7	1
32 33	32.0 33.0	00.6 00.6	92 93	92.0 93.0	01.6 01.6		152.0 153.0			212.0 213.0			272.0 273.0		į
34	34.0	00.6	94	94.0	01.6		154.0	02.7		214.0			274.0		i
35	35.0	00.6	95	95.0	01.7		155.0	02.7		215.0			275.0		
36	36.0	00.6	96	96.0	01.7	56	156.0			216.0			276.0		-
37	37.0	00.6	97	97.0	01.7		157.0	02.7		217.0			277.0		
38	38.0	00.7 00.7	98 99	98.0	01.7 01.7		158.0	02.8 02.8		218.0 219.0			278.0	04.9	
39 40	39.0 40.0	00.7		100.0	01.7		159.0 160.0			219.0 2 2 0.0	03.8 03.8		27 9.0 280.0	04.9 04.9	-
41	41.0	00.7		101.0	01.8		161.0	02.8	·	221.0	03.9		281.0	04.9	-
42	42.0	00.7		102.0			162.0			222.0	03.9		282.0	04.9	-
43	43.0			103.0	01.8		163.0		23	223. 0	03.9		283.0		
44	44.0	00.8		104.0			164.0			224.0	03.9		284.0		
45	45.0			105.0			165.0			225.0			285.0		
46	46.0	8.00 8.00		106.0 107.0			166.0 167.0			226 .0 227 .0			286.0 287.0		
47 48	47.0 48.0	00.8		108.0	01.9 01.9		168.0			22 7.0			288.0		
49	49.0			109.0			169.0		29	229 .0	04.0		289.0		1
50	50.0	00.9		110.0	01.9		170.0		30	230.0	04.0		290.0	05.1	
51	51.0	00.9	111	111.0	01.9		171.0	03.0	231	231.0	04.0	291	291.0	05.1	-
52	52.0	00.9	12	112.0	02.0		172.0		32	232 .0	04.0	92	292.0	05.1	-
53	53.0	00.9		113.0			173.0			233 .0			293.0		
54	54.0			114.0 115.0	02.0 02.0		174.0 175.0			234.0 235.0			294 .0 2 95.0	05.1 05.1	
55 56	55.0 56.0			116.0			176.0			236.0 236.0	04.1 04.1		296.0 2 96.0	05.2	
57	57.0	01.0		117.0	02.0		177.0			237.0	04.1		297 .0	05.2	1
58	58.0	01.0	18	118.0	02.1	78	178.0	03.1	38	238.0	04.2	98	298.0	05.2	1
59	59.0	01.0		119.0	02.1		179.0			239.0	04.2		299.0	05.2	1
60	60.0	01.0		120.0	02.1		180.0		I	240.0	04.2		300.0	05.2	1
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist	Dep.	Lat.		Dep.	Lat.	1

[For 89 Degrees.

Difference of Latitude and Departure for 2 Degrees.

)ist.	Lat.	Dep.	Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
	01.0	00.0			02.1	121	120.9	04.2	181	180.9	06.3	241	240.9	08.4
2							121.9			181.9			241.9	08.4
3							122.9			182.9			242.9	
4 5							123.9 124.9			18 3 .9			243.9 244.9	1
6							125.9			185.9			245.9	
7	07.0						126.9			186.9			246.8	
8	08.0	00.3	68	68.0	02.4	28	127.9	04.5	88	187.9	06.6	48	247.8	08.7
9							128.9			188.9			248.8	
10							129.9			189.9			249.8	
11	11.0	00.4		71.0			130.9			190.9			250.8	8.80
12 13		00.4					131.9 1 32 .9			191.9 19 2 .9			251.8 252.8	8.80 08.8
14	14.0						133.9			193.9			253.8	08.9
15	15.0						134.9			194.9			254.8	08.9
16	16.0			76.0			135.9			195.9	06.8	56	256.8	08.9
17	17.0	00.6		77.0			136.9			196.9			256.8	09.0
18 19	18.0 19.0						137.9 138.9			197.9 198.9			257.8 258.8	09.0
20	20.0						139.9			199.9	06.9 07.0		259.8	09.0 09.1
21	21.0	00.7	1——	81.0			140.9			200.9			260.8	09.1
22	22.0	00.8		82.0			141.9			2 01.9			261.8	09.1
23	23.0						142.9			202.9		63	262.8	09.2
24	24.0						143.9	05.0	04	203.9		64	263.8	09.2
25	25.0	00.9		84.9			144.9			204.9	07.2		264.8	09.2
26 27	26.0 27.0	00.9		85.9			145.9			2 05.9			265.8	09.3
28		01.0		86.9 87.9			146.9 147.9			206.9 207.9			266.8 267.8	09.3 09.4
29	29.0	01.0		88.9			148.9			208.9	07.3		268.8	09.4
30	30.0	01.0		89.9	03.1		149.9			209.9	07.3		269.8	09.4
31	31.0	01.1	91	90.9	03.2	151	150.9	05.3	211	210.9	07.4	271	270.8	09.5
32	32.0	01.1	92				151.9			211.9	07.4		271.8	09.5
33	33.0	01.2	93		03.2		152.9	05.3		212.9			272.8	09.5
34 35	34.0 35.0	01.2	94 95		03.3 03.3		153.9 154.9	05.4 05.4		213.9 214.9			273.8 274.8	09.6 09.6
36	36.0	01.3			03.4		155.9	05.4		215.9 215.9		76	275.8	09.6
37	37.0	01.3	97	96.9	03.4		156.9	05.5		216.9	07.6		276.8	09.7
38	38.0	01.3		97.9	03.4	58	157.9	05.5	18	ž17.9	07.6	78	277.8	09.7
3 9	39.0	01.4	99	98.9			158.9	05.5		218.9			278.8	09.7
40	40.0	01.4	100	99.9	03.5		159.9	05.6		219.9			279.8	09.8
41 42	41.0 42.0	01.4		100.9	03.5		160.9	05.6		220.9	07.7		280.8	09.8
43	43.0	01.5 01.5		101.9 102.9	03.6 03.6		161.9 162.9	05.7 05.7		221.9 2 22 .9	07.7 07.8		281.8 282.8	09.8 09.9
44	44.0	01.5		103.9	03.6		163.9	05.7		223.9	07.8		283.8	09.9
45	45.0	01.6		104.9	03.7		164.9	05.8		224.9	07.9		284.8	09.9
46	46.0	01.6		105.9	03.7		165.9	05.8		225.9	07.9		285.8	10.0
47 48	47.0	01.6		106.9	03.7		166.9	05.8		226.9	07.9		286.8	10.0
49	48.0	01.7 01.7	08	107.9 108.9	03.8 03.8		167.9 168.9	05.9 05.9		2 27 .9 228.9	08.0		287.8 288.8	10.1
50	50.0	01.7	10	109.9	03.8		169.9	05.9		229.9	08.0 08.0	90	289.8	10.1
51	51.0	01.8		110.9	03.9		170.9	06.0		230.9	08.1		290.8	10.2
52	52.0	01.8		111.9	03.9		171.9	06.0		231.9	08.1		291.8	10.2
53	53.0	01.8	13	112.9	03.9	73	172.9	06.0		232.9	08.1		292.8	10.2
54	54.0	01.9		113.9	04.0		173.9	06.1		233.9	08.2	94	293.8	10.3
55	55.0	01.9		114.9	04.0		174.9	06.1		234.9	08.2	95	294.8	10.3
56 57	56.0 57.0	02.0	17	115.9 116.9	04.0 04.1		175.9 176.9	06.1		235.9	08.2		295.8	10.3
58	58.0	02.0		117.9	04.1		177.9	06.2 06.2		36.9 37.9	08.3 08.3		296.8 297.8	10.4
59	59.0	02.1		118.9	04.2		178.9	06.2		38.9	08.3		298.8	10.4
60	60.0	02.1	20	119.9	04.2	80	179.9	06.3	402	39.9	08.4		299.8	10.5
list.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.
												/For		

[For 88 Degrees

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Difference of Latitude and Departure for 3 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
1	01.0	00.1	61	60.9	03.2		120.8	06.3		180.8	09.5	241	240.7	12.6
2	02.0	00.1	62	61.9	03.2		121.8	06.4		181.8	09.5	42	241.7	12.7
3	03.0	00.2	63	62.9	03.3		122.8	06.4		182.7	09.6		242.7	12.7
5	04.0 05.0		64 65	63.9 64.9	03.3 03.4		123.8 124.8	06.5 06.5		183.7 184.7	09.6 09.7		243.7 244.7	12.8 12.8
. 6	06.0		66	65.9	03.5		125.8	06.6		185.7	09.7		245.7	12.9
. 7	07.0	00.4	67	66.9	03.5	27	126.8	06.6	87	186.7	09.8	47	246.7	12.9
8	08.0		1		03.6		127.8	06.7		187.7	09.8		247.7	13.0
10	09.0 10.0		69 70	68.9 69.9	03.6 03.7		128.8 129.8			188.7 189.7	09.9 09.9		248.7 249.7	13.0 13.1
1 11	11.0		71	70.9	03.7		130.8	06.9		190.7	10.0		250.7	13.1
12	12.0	00.6	72	71.9	03.8		131.8	06.9		191.7	10.0		250.7 251.7	13.1
13			73	72.9	03.8		132.8	07.0		192.7	10.1		252.7	13.2
14			74				133.8			193.7	10.2		253.7	13.3
15 16				74.9 75.9	03.9 04.0		134.8 135.8	07.1 07.1		194.7 195.7	10.2 10.5		254.7 255.6	
17			76	76.9	04.0		136.8			196.7	10.3		256.6	
18			78		04.1		137.8	07.2		197.7	10.4		257.6	
19			79		04.1		138.8	07.3		198.7	10.4		258.6	
20			80		04.2		139.8	07.3		199.7	10.5		259.6	13.6
21	21.0 22.0		81	80.9	04.2		140.8	07.4		200.7	10.5		260.6	
22	23.0		82 83	81.9 82.9	04.3 04.3		141.8 142.8			201.7 202.7	10.6 10.6		261.6 262.6	13.7
24					04.4		143.8			203.7	10.7		263.6	
25	25.0		85	84.9	04.4		144.8	07.6		204.7	10.7		264.6	
26	26.0	01.4	86	85.9	04.5		145.8	07.6		205.7	10.8		265.6	13.9
27	27.0 28.0		87 88	86.9 87.9	04.6 04.6		146.8 147.8	07.7 07.7		206.7 207.7	10.8 10.9		266.6 267.6	14.0
29	29.0		89	88.9	04.7		148.8			208.7	10.9		268.6	
30	30.0	01.6	90	89.9	04.7		149.8	07.9		209.7	11.0		269.6	14.1
31	31.0	01.6	91	90.9	04.8	151	150.8	07.9		210.7	11.0		270.6	
32	32.0		92	91.9	04.8		151.8	08.0		211.7	11.1		271.6	14.2
33 34	33.0 34.0		93 94	92.9 93.9	04.9 04.9		152.8 153.8	08.0 08.1		212.7 213.7	11.1 11.2		272.6 273.6	14.3
35	35.0	01.8	95	94.9	05.0		154.8	08.1		214.7	11.3		274.6	14.4
36	36.0	01.9	96	95.9	05.0		155.8	08.2		215.7	11.3		275.6	
37	36.9	01.9	97	96.9	05.1		156.8	08.2		216.7	11.4		276.6	
38 39	37.9 38.9	02.0 02.0	98 99	97.9 98.9	05.1 05.2		157.8 158.8	08.3 08.3		217.7 218.7	11.4 11.5		277.6 278.6	14.5
40	39.9	02.1	100	99.9	05.2		159.8	08.4		219.7	11.5		279.6	14.7
41	40.9	02.1	101	100.9	05.3	161	160.8	08.4	221	220.7	11.6	281	280.6	14.7
42	41.9	02.2	02	101.9	05.3	62	161.8	08.5		221.7	11.6		281.6	14.8
43	42.9 43.9			102.9	05.4		162.8	08.5		222.7	11.7		282.6	
45	44.9	02.3 02.4		103.9 104.9	05.4 05.5		163.8 164.8	08.6 08.6		223.7 224.7	11.7 11.8		283.6 284.6	14.9
46	45.9	02.4		105.9	05.5		165.8	08.7		225.7	11.8		285.6	
.47	46.9	02.5		106.9	05.6		166,8	08.7		226.7	11.9		286.6	
48 49	47.9	02.5 02.6		107.9			167.8	8.80.	28		11.9 12.0		287.6 288.6	15.1
50	48.9 49.9	02.6		108.9 109.8	05.7 05.8		168.8 169.8	08.8 08.9		228.7 229.7	12.0		289.6	15.1 15.2
51	50.9	02.7	١	110.8	05.8		$\frac{103.0}{170.8}$	08.9	231	230.7		291	290.6	15.2
52	51.9	02.7		111.8	05.9		171.8	09.0		231.7			291.6	15.3
53	52.9	02.8	13	112.8	05.9	73	172.8	09.1	33	232.7	12.2		292.6	15.3
54	53.9	02.8		113.8			173.8	09.1	1	233.7	12.2		293.6	15.4
55 56	54.9 55.9	02.9 02.9		114.8 115.8			174.8 175.8	09.2 09.2		234.7 235.7	12.3 12.4		294.6 295.6	15.4
57	56.9	03.0		116.8			176.8	09.3		236.7	12.4		296.6	15.5
58	57.9	03.0	18	117.8	06.2	78	177.8	09.3	3 8	237.7	12.5	98	297.6	15.6
59	58.9	03.1		118.8	06.2		178.8	09.4		238.7	12.5		298.6	15.6
60	59.9	03.1		119.8			179.8	09.4		239.7	12.6		299.6	15.7
Dist.	Dep.	Lat.	DIST.	Dep.	LBI.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.		Dep.	
												[ro	r 87 D	egrees.

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Difference of Latitude and Departure for 4 Degrees.

10 10 10 10 10 10 16 16	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.
30 0.0 0 0.0 2.6 43 62.8 04.4 23192.7 08.6 83182.6 12.8 4324.2 417.0 50.0 0.0 0.3 64 63.0 0.5 5.0 447 2473.4 708.7 8 68 68 67.6 06.0 0.4 46 65 65.8 04.6 25125.7 08.7 8 68 128.5 12.8 4424.4 17.1 68.0 0.0 0.0 4.6 66 65.8 04.6 25125.7 08.9 8 68 18.5 13.0 4724.4 17.1 68.0 0.0 00.4 66 66 67.6 04.7 23192.7 08.9 8 68 18.5 13.0 4724.4 17.2 8 08.0 00.6 69 68.8 04.8 29128.7 09.0 9.0 138.5 13.0 42724.4 17.2 10.0 0.0 0.0 0.0 6 69 68.8 04.8 29128.7 09.0 93 18.5 13.2 4824.3 17.4 17.3 10.0 0.0 0.7 70 69.8 04.3 30129.7 09.1 90.1 90.1 90.1 90.1 90.1 90.1 90.1		01.0	00.1		60.9	04.3								
\$\frac{4}{0}, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0, 0,														
50 5.0 0 0.0 4, 66 66.8 0 4.6 25/195.7 0 8.3 85/184.5 13.0 46/245.4 17.1 2 70.7 0 00.5 67 66.8 0 4.7 27/126.7 0 8.9 8 87/186.5 13.0 47/246.4 17.2 8 9 09.0 0 6.6 69 68.8 0 4.8 29/128.7 0 9.0 89/185.5 13.0 44/247.4 17.2 19.0 0 0 0 0 6 69 68.8 0 4.8 29/128.7 0 9.0 89/185.5 13.0 44/247.4 17.2 19.0 10.0 0 0 0 7 0 99.8 0 4.9 30/129.7 0 9.1 90/189.5 13.3 20/249.4 17.4 10.1 11.0 0 0 0 1.7 70 99.8 0 13/187.0 0 1.1 11.0 0 0 0 1.7 170.9 0 5.0 13/187.0 0 1.1 11.0 0 0 0 1.7 170.9 0 5.0 13/187.0 0 1.1 11.0 0 0 0 1.7 170.9 0 5.0 13/187.0 0 1.1 11.0 0 0 0 1.7 170.9 0 5.0 13/187.0 0 1.1 11.0 0 0 0 1.7 170.9 0 5.0 13/187.0 0 1.1 11.0 0 0 1.0 174 73.8 0 5.2 34/183.7 0 9.3 94/193.5 13.5 45/253.4 17.6 14.4 0 0 1.0 74 73.8 0 5.2 34/183.7 0 9.3 94/193.5 13.5 55/24.4 17.6 16.6 0 0 1.1 76 75.8 0 5.3 36/185.7 0 9.5 96/195.5 13.7 55/254.4 17.9 17.1 17.0 0 1.2 77.6 8 0 5.4 38/187.7 0 9.6 97/196.6 13.7 55/254.4 17.9 17.1 17.0 0 1.2 77.6 8 0 5.4 38/187.7 0 9.6 97/196.6 13.7 55/254.4 17.9 17.1 17.0 0 1.2 77.6 8 0 5.4 38/187.7 0 9.6 97/196.6 13.7 55/254.4 17.9 17.0 0 1.3 79 78.8 0 5.5 39/183.7 0 9.5 99/198.5 18.9 99/258.4 18.1 19.0 0 1.3 79 78.8 0 5.5 04/133.7 0 9.6 97/196.6 13.7 526.4 17.9 19.0 0 1.3 79 78.8 0 5.5 04/133.7 0 9.8 20/190.5 14.0 66/259.4 18.1 19.0 0 1.3 79 78.8 0 5.5 04/133.7 0 9.8 20/190.5 14.0 66/259.4 18.1 19.0 19.0 0 1.3 80.8 0 5.9 44/143.6 10.0 0 420.5 14.0 66/259.4 18.1 19.0 19.0 0 1.3 80.8 0 5.9 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.6 83 82.8 0 5.9 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.6 83 82.8 0 5.9 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.6 83 88.8 0 5.9 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.6 83 88.8 0 6.5 5/184.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.8 80.8 80.8 0 6.0 0 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.8 80.8 80.8 0 6.0 0 44/143.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.8 80.8 80.8 0 6.0 0 5/184.6 10.0 0 420.5 14.4 66/256.4 18.5 62/25.9 0 1.8 80.8 80.8 0 6.0 0 5/184.6 10.0 0 420.5 14.4 66/256.4 18.5 62/256.4 18.5 62/256.4 18.5 14.														
60 6.0 00.4 66 65.8 04.6 26 125.7 08.3 86 185.5 13.0 46 245.4 17.2 8														
7 07. 0 00. 5 67 66. 8 04. 7 27126. 7 08. 9 87 186. 5 13. 0 47 246. 4 17. 2 8 08. 0 00. 6 69 68. 8 04. 8 29128. 7 09. 0 89 188. 5 13. 1 49 247. 4 17. 4 17. 2 1 10. 0 00. 7 70 69. 8 04. 9 30 129. 7 09. 1 90 189. 5 13. 3 50 249. 4 17. 4 17. 2 12. 0 00. 8 72 71. 8 05. 0 131 130. 7 09. 1 90 189. 5 13. 3 50 249. 4 17. 4 17. 2 12. 0 00. 8 72 71. 8 05. 0 131 130. 7 09. 1 92 191. 5 13. 3 50 249. 4 17. 4 17. 2 14. 1 10. 0 01. 0 74 73. 8 05. 2 35 131. 7 09. 2 92 191. 5 13. 4 52 251. 4 17. 6 13. 1 10. 0 01. 0 74 73. 8 05. 2 35 134. 7 09. 3 93 192. 5 13. 5 52 252. 4 17. 6 15. 0 10. 0 75 74. 8 05. 2 35 134. 7 09. 3 93 192. 5 13. 5 52 252. 4 17. 7 15 15. 0 01. 0 75 74. 8 05. 2 35 134. 7 09. 4 95 194. 5 13. 6 252. 4 17. 7 17 17 17 10 11. 2 77 76. 8 05. 4 37 136. 7 09. 6 98 195. 5 13. 7 57 256. 4 17. 9 19 19. 0 01. 3 78 78. 8 05. 5 39138. 7 09. 7 09. 1 99198. 5 13. 7 57 256. 4 17. 9 19 19. 0 01. 3 79 78. 8 05. 5 39138. 7 09. 7 09. 8 2019. 5 14. 0 6025. 4 18. 1 12 12 12 12 12 12 12 12 12 12 12 12 1														
8 08.0 0 00.6 68 67.8 04.7 281 27.7 08.9 88 87.5 13.1 48 247.4 17.3 9 09.0 00.6 69.6 8.9 04.9 30 129.7 09.1 90 189.5 13.3 50 249.4 17.4 11 11.0 00.8 71 70.9 05.0 32 131.7 09.2 92 191.5 13.4 50 249.4 17.4 13 13.0 00.9 73 72.8 05.1 53 232.7 09.3 93 192.5 13.5 53 252.4 17.6 14 14.0 01.0 74.7 78.8 05.2 34 133.7 09.3 94 191.5 13.4 53 251.4 17.6 15 15.0 01.0 75 74.8 05.2 34 133.7 09.4 95 194.5 13.6 53 252.4 17.6 16 16.0 01.1 76.7 8.8 05.2 35 134.7 09.4 95 194.5 13.6 55 254.4 17.8 18 18.0 01.3 78 77.8 05.4 38 137.7 09.6 98 197.5 13.7 57 256.4 17.9 19 19.0 01.3 78 87.8 05.5 38 138.7 09.6 98 197.5 13.7 57 256.4 17.9 19 19.0 01.3 78 88.8 05.5 38 33.7 09.6 98 197.5 13.7 57 256.4 17.9 19 19.0 01.3 78 88.8 05.5 38 33.7 09.6 99 198.5 13.9 99 598.4 18.1 20 20.0 01.4 30 79.8 05.6 40 139.7 09.8 20 1200.5 14.0 60 259.4 18.1 21 20.9 01.6 83 82.8 05.8 45 144.7 09.8 20 200.5 14.0 60 259.4 18.1 22 21.9 01.6 83 82.8 05.8 45 144.6 10.1 05 204.5 14.3 66 256.4 18.5 22 22.7 01.6 83 88.8 05.9 44 443.6 10.0 04 203.5 14.2 66 265.4 18.5 22 27.9 01.6 83 88.8 05.9 45 144.6 10.1 05 204.5 14.4 66 256.4 18.5 24 23.9 01.7 86.8 06.1 47 146.6 10.3 08 07.5 14.4 66 256.4 18.5 25 24.9 01.7 86.8 06.1 47 146.6 10.3 08 07.5 14.4 66 256.4 18.5 26 25.9 01.8 86 8.8 06.7 56 156.6 10.9 16 16.5 15.1 76 275.3 19.5 28 27.9 02.0 89 88.8 06.9 56 146.6 10.5 20 11.5 14.7 72 77.3 19.5 29 28.9 02.0 89 88.8 06.9 56 146.6 10.5 20 11.5 14.7 72 77.3 19.5 20 29.0 01.0 03 07.0 05 146.6 05.5 05 14.6 05 25.5 05 25.6 05 25														
9 9.0 0.0 6.6 69 68.8 04.8 29128.7 09.0 89108.5 13.2 49248.4 17.4 10 10.0 00.7 70 69.8 04.9 30129.7 09.1 9018.5 13.3 249248.4 17.4 11 11.0 00.8 71 70.8 05.0 131130.7 09.1 191190.5 13.3 251250.4 17.5 12 12 0.0 00.8 72 71.8 05.0 32131.7 09.2 92191.5 13.4 53251.4 17.6 13 13.0 00.9 73.7 18.8 05.2 34133.7 09.3 94193.5 13.5 53252.4 17.6 18.1 19.1 19.0 10.0 74 73.8 05.2 34133.7 09.3 94193.5 13.5 53252.4 17.7 17.6 16.1 16.0 01.0 74 73.8 05.2 35134.7 09.4 95194.5 13.6 5225.4 17.7 17.0 10.0 1.2 77 76.8 05.4 37136.7 09.5 96195.5 13.7 552.54 17.9 17.7 17.0 01.2 77 76.8 05.4 37136.7 09.6 97196.6 13.7 562.55.4 17.9 19.0 01.3 79 78.8 05.5 39138.7 09.6 98197.5 13.8 82297.4 18.0 19.0 01.3 79 78.8 05.5 39138.7 09.6 98197.5 13.8 82297.4 18.0 19.0 01.4 30 79.8 05.5 40139.7 09.8 201190.9 11.5 82 82.8 05.8 39138.7 09.7 09.8 200199.5 14.0 06.259.4 18.1 20.2 20.9 01.6 83 82.8 05.7 42141.7 09.9 8 2011200.5 14.0 261 260.4 18.5 22.2 24.9 01.6 83 82.8 05.8 43142.7 10.0 03202.5 14.2 64263.4 18.4 22.2 29.9 01.6 83 82.8 05.8 43142.7 10.0 03202.5 14.2 65264.4 18.5 26.2 59.0 01.8 86 85.3 06.0 46145.6 10.2 06205.5 14.4 66265.4 18.6 22.2 24.9 01.7 86 88 87.8 06.0 46145.6 10.2 06205.5 14.4 66265.4 18.5 26.2 29.9 01.9 86 87.8 06.1 48147.6 10.3 08207.5 14.5 66265.4 18.5 29.2 29.9 02.1 90 88 87.8 06.1 48147.6 10.3 08207.5 14.5 66265.4 18.5 29.2 29.9 02.1 90 88.8 05.3 60.1 48147.6 10.3 08207.5 14.5 66265.4 18.5 29.2 29.9 02.1 90 89.8 06.5 53126.6 10.7 13212.5 14.7 271270.9 18.9 33.9 02.2 91 99.8 06.5 53126.6 10.7 13212.5 14.9 73277.3 19.6 33.3 9.9 02.2 91 99.8 06.5 53126.6 10.7 13212.5 14.9 73277.3 19.6 33.3 9.9 02.4 94.9 93.8 06.5 53126.6 10.7 13212.5 14.9 73277.3 19.6 44.9 03.9 02.9 90.10 0.7 07.5 66165.6 10.9 16216.5 15.1 76275.3 19.5 34.3 9.9 02.9 90.10 0.7 07.5 66165.6 10.9 16216.5 15.1 76275.3 19.5 34.3 9.9 02.9 02.0 10.0 03.0 07.0 161100.6 11.2 20219.5 14.5 14.8 92291.3 19.5 44.9 03.9 03.0 03102.7 07.5 67166.6 11.6 222221.5 15.5 38288.3 19.7 3277.3 19.6 44.9 03.9 03.0 03102.7 07.5 67166.6 11.6 2222221.5 15.5 82288.3 19.7 3288.3 19.9 03.5 10.0														
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42 41.9 02.9 02 101.8 07.1 62 161.6 11.3 22 221.5 15.5 82 281.3 19.7 43 42.9 03.0 03 102.7 07.2 63 162.6 11.4 23 222.5 15.6 83 282.3 19.7 45.4 49.9 03.1 04 103.7 07.3 65 164.6 11.4 24 223.5 15.6 84 283.3 19.8 45.4 49.9 03.1 05 104.7 07.3 65 164.6 11.5 25 124.6 15.7 85 284.3 19.8 46.4 5.9 03.2 06 105.7 07.4 66 165.6 11.6 26 225.4 15.8 86 285.3 20.0 47.7 46.9 03.3 07 106.7 07.5 67 166.6 11.6 26 225.4 15.8 87 286.3 20.0 10.0 10.0 10.0 10.0 10.0 10.0 10.0	40	39.9	02.8	100	99.8	07:0	60	159.6	11.2	20	219.5	15.3	80 279.3	19.5
43 42.9 03.0 03102.7 07.2 63 62.6 11.4 23222.5 15.6 83282.3 19.7 44.9 03.1 04.103.7 07.3 64.163.6 11.4 24.223.5 15.6 84.283.3 19.8 45.4 44.9 03.1 05.104.7 07.3 65.164.6 11.5 25.224.5 15.6 84.283.3 19.8 46.4 5.9 03.2 06.105.7 07.4 66.165.6 11.6 26.225.4 15.8 86.285.3 20.0 47.46.9 03.3 08.107.7 07.5 67.166.6 11.6 27.226.4 15.8 87.286.3 20.0 49.9 03.5 08.107.7 07.5 68.167.6 11.7 28.227.4 15.9 88.287.3 20.1 28.287.4 15.9 88.287.3 20.1 28.287.4 15.9 88.287.3 20.1 28.287.4 15.9 88.287.3 20.2 25.5 15.6 11.6 27.226.4 15.8	41	40.9		101	100.8	07.0	161	160.6						
44 43.9 03.1 04103.7 07.3 64 65.6 11.4 24 225.5 15.6 84 283.3 19.8 45.4 44.9 03.1 05104.7 07.3 65164.6 11.5 25224.5 15.7 85284.3 19.9 47 46.9 03.3 07106.7 07.5 67166.6 11.6 26 225.4 15.8 87286.3 20.0 17.0 17.0 17.0 17.0 17.0 17.0 17.0 1										22				
45 44.9 03.1 05104.7 07.3 65164.6 11.5 252224.5 15.7 85284.3 19.9 466 45.9 03.2 06105.7 07.4 66165.6 11.6 26 225.4 15.8 87286.3 20.0 49.9 03.5 07106.7 07.5 68 167.6 11.7 28 27.4 15.8 87286.3 20.0 19.9 19.9 03.5 10 109.7 07.6 69 168.6 11.6 29 228.4 16.0 89288.3 20.2 19.9 03.5 10 109.7 07.7 70 169.6 11.9 30 229.4 16.0 90 289.3 20.2 19.9 03.5 10 109.7 07.7 70 169.6 11.9 231 230.4 16.1 291 290.3 20.2 19.9 03.6 12111.7 07.8 72171.6 12.0 32 231.4 16.2 92 291.3 20.4 16.3 52.9 03.7 13 112.7 07.9 73 172.6 12.1 33 232.4 16.3 93 292.3 20.4 16.5 55.9 03.9 16115.7 08.0 75 174.6 12.2 35 23.4 16.3 93 292.3 20.4 16.5 55.9 03.9 16115.7 08.1 75 174.6 12.2 35 23.4 16.5 95 294.3 20.6 59 58.9 04.0 17116.7 08.2 78 177.6 12.3 36 235.4 16.5 96 295.3 20.6 59 59.9 04.0 18 117.7 08.2 78 177.6 12.3 37 236.4 16.5 96 295.3 20.6 59 59.9 04.0 18 117.7 08.2 78 177.6 12.4 38 237.4 16.5 98 297.3 20.8 19.9 19.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 99 298.3 20.9 16.5 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 99 298.3 20.9 16.5 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 99 298.3 20.9 16.5 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 16.7 16.7 16.7 16.7 16.7 16.7 16.7														
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47 46.9 03.3 07106.7 07.5 67 166.6 11.6 27 226.4 15.8 87 286.3 20.0 49.9 03.5 08107.7 07.5 68167.6 11.7 28 227.4 15.9 88 287.3 20.1 29 28.4 16.0 90 289.3 20.2 50 49.9 03.5 10109.7 07.7 70 169.6 11.9 30 229.4 16.0 90 289.3 20.2 51 50.9 03.6 111110.7 07.7 171 170.6 11.9 231 230.4 16.1 291 290.3 20.2 51 52 51.9 03.6 12 111.7 07.8 72 171.6 12.0 32 231.4 16.2 92 291.3 20.4 54 53.9 03.7 13 112.7 07.9 73 172.6 12.1 34 233.4 16.3 93 292.3 20.5 55 54.9 03.8 14 113.7 08.0 74 173.6 12.1 34 233.4 16.3 94 293.3 20.5 55 54.9 03.8 15 114.7 08.0 75 174.6 12.1 34 233.4 16.3 94 293.3 20.5 55 54.9 03.8 15 114.7 08.0 75 174.6 12.2 35 234.4 16.3 95 294.3 20.5 55 54.9 03.8 15 114.7 08.0 75 174.6 12.2 35 234.4 16.3 95 294.3 20.5 56 55.9 03.9 16 115.7 08.1 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 57.9 04.0 17 116.7 08.2 77 176.6 12.3 37 226.4 16.5 97 296.3 20.7 58 57.9 04.0 18 117.7 08.2 78 177.6 12.4 38 237.4 16.5 98 297.3 20.8 59 59.9 04.1 19 118.7 08.3 79 178.6 12.5 39 238.4 16.7 99 298.3 20.9 01st. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
49 47.9 03.3 08107.7 07.5 68 67.6 11.7 28 227.4 15.9 88 287.3 20.1 49.9 03.4 09108.7 07.6 69168.6 11.8 291228.4 16.0 89288.3 20.2 50 49.9 03.5 10 109.7 07.7 70 169.6 11.9 30 129.4 16.0 90289.3 20.2 50 12 51.9 03.6 1111110.7 07.7 171 171 170.6 11.9 231230.4 16.1 291290.3 20.3 52.9 03.7 13 112.7 07.9 73 172.6 12.1 33 231.4 16.2 92 291.3 20.4 53.9 03.8 14113.7 08.0 74 173.6 12.1 33 232.4 16.3 93 292.3 20.4 555 54.9 03.8 15 114.7 08.0 75 174.6 12.2 3523.4 16.3 93 292.3 20.4 555 55.9 03.9 16 115.7 08.0 75 174.6 12.2 36 235.4 16.5 96 295.3 20.6 55 55.9 03.9 16 115.7 08.1 75 174.6 12.2 36 235.4 16.5 96 295.3 20.6 57 56.9 04.0 17 116.7 08.2 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 59 58.9 04.1 19 118.7 08.2 78 177.6 12.3 37 236.4 16.5 97 296.3 20.7 58 57.9 04.0 18 117.7 08.2 78 177.6 12.3 37 236.4 16.5 98 297.3 20.8 60 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 99298.3 20.9 01st. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.										97	226 4			
49 48.9 03.4 09 108.7 07.6 69 168.6 11.8 29 228.4 16.0 89 288.3 20.2 50 49.9 03.5 10 109.7 07.7 70 169.6 11.9 30 229.4 16.0 90 289.3 20.2 51.9 03.6 12111.7 07.8 72 171.6 12.0 32 231 4 16.2 92 291.3 20.4 53.9 03.8 14113.7 08.0 73 172.6 12.1 33 232.4 16.3 93 292.3 20.4 54.9 03.8 14113.7 08.0 75 174.6 12.2 35 234.4 16.3 94 293.3 20.5 55 54.9 03.8 15 114.7 08.0 75 174.6 12.2 35 234.4 16.4 95 294.3 20.5 55 55.9 03.9 16115.7 08.1 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 55.9 03.9 16115.7 08.1 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 55 97 04.0 17116.7 08.2 77 176.6 12.3 37 236.4 16.5 97 296.3 20.7 58 57.9 04.0 18117.7 08.2 78 177.6 12.3 37 236.4 16.5 97 296.3 20.7 58 57.9 04.0 18117.7 08.2 78 177.6 12.4 38 237.4 16.6 99 297.3 20.8 59.9 04.1 19118.7 08.3 79 178.6 12.5 37 236.4 16.7 99 298.3 20.9 01.5 Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
50 49.9 03.5 10 109.7 07.7 70 169.6 11.9 30 229.4 16.0 90 289.3 20.2 51 50.9 03.6 111 110.7 07.7 171 170.6 11.9 231 230.4 16.1 291 290.3 20.3 52 51.9 03.6 12 111.7 07.9 73 172.6 12.1 332 32.4 16.2 92 291.3 20.4 53 52.9 03.7 13112.7 07.9 73 172.6 12.1 33 232.4 16.3 94 293.3 20.4 54 53.9 03.8 14113.7 08.0 75 174.6 12.1 34 233.4 16.3 94 293.3 20.5 55 54.9 03.8 15 114.7 08.0 75 174.6 12.2 35 234.4 16.3 94 293.3 20.6 56 55.9 03.9 16 115.7 08.1 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 57 56.9 04.0 18 117.7 08.2 </td <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td>														
52 51.9 03.6 12111.7 07.8 72 171.6 12.0 32231.4 16.2 92291.3 20.4 53.9 03.7 13112.7 07.9 73172.6 12.1 33232.4 16.3 93292.3 20.4 53.9 03.8 14113.7 08.0 74173.6 12.1 34233.4 16.3 94293.3 20.5 55 54.9 03.8 15114.7 08.0 75174.6 12.2 35234.4 16.4 95294.3 20.6 55 55.9 03.9 16115.7 08.1 76175.6 12.3 36235.4 16.5 96295.3 20.6 57.9 04.0 17116.7 08.2 77176.6 12.3 37236.4 16.5 97296.3 20.7 58 57.9 04.0 18117.7 08.2 78177.6 12.4 38237.4 16.5 97296.3 20.7 59 58.9 04.1 19118.7 08.2 79178.6 12.4 38237.4 16.6 99297.3 20.8 59.9 04.2 20119.7 08.4 80 179.6 12.6 40 239.4 16.7 99298.3 20.9 01.5 Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
52 51.9 03.6 12111.7 07.8 72 171.6 12.0 32231.4 16.2 92291.3 20.4 53.9 03.7 13112.7 07.9 73172.6 12.1 33232.4 16.3 93292.3 20.4 53.9 03.8 14113.7 08.0 74173.6 12.1 34233.4 16.3 94293.3 20.5 55 54.9 03.8 15114.7 08.0 75174.6 12.2 35234.4 16.4 95294.3 20.6 55 55.9 03.9 16115.7 08.1 76175.6 12.3 36235.4 16.5 96295.3 20.6 57.9 04.0 17116.7 08.2 77176.6 12.3 37236.4 16.5 97296.3 20.7 58 57.9 04.0 18117.7 08.2 78177.6 12.4 38237.4 16.5 97296.3 20.7 59 58.9 04.1 19118.7 08.2 79178.6 12.4 38237.4 16.6 99297.3 20.8 59.9 04.2 20119.7 08.4 80 179.6 12.6 40 239.4 16.7 99298.3 20.9 01.5 Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.	51	50.9	03.6	111	110.7	07.7	171	170.6		231	230.4		291 290.3	20.3
53 52.9 03.7 13112.7 07.9 73 172.6 12.1 33 232.4 16.3 93 292.3 20.4 53.9 03.8 14113.7 08.0 75 174.6 12.1 34 233.4 16.3 94 293.3 20.5 55 55.9 03.9 16115.7 08.1 75 174.6 12.2 35 23.4 16.5 96 295.3 20.6 55 57 56.9 04.0 17116.7 08.2 78 177.6 6 12.3 37 236.4 16.5 96 295.3 20.6 59 58.9 04.1 19118.7 08.2 78 177.6 12.3 37 236.4 16.5 97 296.3 20.7 59 59.9 04.2 20 119.7 08.2 80 179.6 12.6 39 293.4 16.7 99 298.3 20.9 04.1 19118.7 08.3 80 179.6 12.6 40 239.4 16.7 99 298.3 20.9 05.5 05.5 05.5 05.5 05.5 05.5 05.5 0				12						32	231.4		92 291.3	20.4
55 54.9 03.8 15114.7 08.0 75 174.6 12.2 35 234.4 16.4 95 294.3 20.6 55.9 03.9 16115.7 08.1 76175.6 12.3 36 235.4 16.5 96 295.3 20.6 57.9 04.0 17116.7 08.2 77 176.6 12.3 37 236.4 16.5 97 296.3 20.7 58 57.9 04.0 18117.7 08.2 78 177.6 12.4 38 237.4 16.6 98 297.3 20.8 59 58.9 04.1 19118.7 08.3 79 178.6 12.5 39 238.4 16.7 99 298.3 20.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 300 299.3 20.9 04.1 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.	53	52.9	03.7	13	112.7	07.9	73	172.6	12.1	33	232.4	16.3	93 292.3	20.4
56 55.9 03.9 16115.7 08.1 76 175.6 12.3 36 235.4 16.5 96 295.3 20.6 57.9 04.0 17116.7 08.2 77 176.6 12.3 37 (236.4 16.5 97296.3 20.7 58 57.9 04.0 18117.7 08.2 78 177.6 12.4 38 237.4 16.6 98 297.3 20.8 59.9 04.1 19118.7 08.3 79178.6 12.5 39 (238.4 16.7 99298.3 20.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 300 299.3 20.9 04.2 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.										34	233.4	16.3	94 293.3	
57 56.9 04.0 17116.7 08.2 77 176.6 12.3 37 236.4 16.5 97 296.3 20.7 58 57.9 04.0 18117.7 08.2 78 177.6 12.4 38 237.4 16.6 98 297.3 20.8 59 58.9 04.1 19 118.7 08.3 79 178.6 12.5 39 238.4 16.7 99 298.3 20.9 60 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 300 299.3 20.9 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
58 57.9 04.0 18 117.7 08.2 78 177.6 12.4 38 237.4 16.6 98 297.3 20.8 59 58.9 04.1 19 118.7 08.3 79 178.6 12.5 39 238.4 16.7 99 298.3 20.9 60 59.9 04.2 20 119.7 08.4 80 179.6 12.6 40 239.4 16.7 300 299.3 20.9 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
59 58.9 04.1 19118.7 08.3 79178.6 12.5 39238.4 16.7 99298.3 20.9 60 59.9 04.2 20119.7 08.4 80179.6 12.6 40239.4 16.7 300299.3 20.9 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
60 59.9 04.2 20119.7 08.4 80 179.6 12.6 40 239.4 16.7 300 299.3 20.9 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.														
	-131.	Dep.	Dat. 1	-11:01"	Dep. I	ardt.	D181.1	Dep.	Dat.	DIBL.	Dep.	nat. J		

TABLE II.

Difference of Latitude and Departure for 5 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.
1	01.0	00.1	61	60.8	05.3		120.5	10.5	181	180.3	15.8	241	240.1
2	02.0	00.2	62	61.8	05.4	22	121.5	10.6	82	181.3	15.9		241.1
3	03.0	00.3	63	62.8			122.5	10.7		182.3	15.9		242.1
4	04.0	00.3	64	63.8	05.6		123.5	10.8		183.3	16.0		243.1
5	05.0	00.4 00.5	65	64.8	05.7		124.5	10.9 11.0		184.3 185.3	16.1 16.2		244.1 245.1
6	06.0 07.0	00.6	66 67	65.7 66.7	05.8 05.8		125.5 126.5	11.1		186.3	16.3		246.1
8	08.0	00.7	68	67.7	05.8		127.5	11.2		187.3	16.4		247.1
9	09.0		69	68.7	06.6		128.5	11.2		188.3	16.5		248.1
10	10.0		70	69.7	06.1		129.5	11.3		189.3	16.6	50	249.0
11	11.0	01.0	71	70.7	06.2	131	130.5	11.4	191	190.3	16.6	251	250.0
12	12.0	01.0	72	71.7	06.3		131.5	11.5		191.3	16.7	52	251.0
13	13.0	01.1	73	72.7	06.4		132.5	11.6		192.3	16.8	53	252.0
14	13.9	01.2	74	73.7	06.4		133.5	11.7		193.3	16.9		253.0
15	14.9	01.3	75	74.7	06.5		134.5	11.8	95	194.3	17.0		254.0
16	15.9	01.4	76	75.7	06.6		135.5	11.9		195.3	17.1		255.0
17	16.9	01.5 01.6	77	76.7	06.7		136.5	11.9 12.0		196.3 197.2	17.2 17.3		256.0 257.0
18	17.9 18.9	01.7	78 79	77.7 78.7	06.8 06.9		137.5 138.5	12.1		198.2	17.3		258.0
20	19.9	01.7	80	79.7	07.0		139.5	12.2		199.2	17.4		259.0
21	20.9	01.8	81	80.7		141	140.5	12.3	·	200.2	17.5	1	260.0
22	21.9	01.9	82	81.7	07.1 07.1		141.5	12.4		201.2			261.0
23		02.0		82.7	07.2		142.5	12.5		202.2	17.7		262.0
21	23.9	02.1		83.7	07.3		143.5			203.2	17.8		263.0
25	24.9	02.2		84.7	07.4		144.4	12.6		204.2	17.9		264.0
26	25.9	02.3	86	85.7	07.5	46	145.4	12.7	06	205.2	18.0		265.0
27	26.9	02.4		86.7	07.6	47	146.4	12.8		206.2	18.0	67	266.0
28	27.9	02.4		87.7	07.7		147.4			207.2	18.1		267.0
29	28.9	02.5	89	88.7	07.8	49	148.4	13.0		208.2	18.2 18.3		268 0 269 0
30	29.9	02.6	90	89.7	07.8		149.4			209.2			
31	30.9	02.7	91	90.7	07.9		150.4	13.2		210.2	18.4		270.0 2 71. 0
32	31.9	02.8 02.9	92	91.6	08.0		151.4 152.4	13.2 13.3		211.2 212.2	18.5 18.6		272 D
33 34	32.9 33.9	03.0	93 94	92.6 93.6	08.1 08.2		153.4	13.4		213.2	18.7		273.D
35	34.9	03.1	95	94.6			154.4	13.5		214.2	18.7		274.0
36	35.9	03.1	96	95.6		56	155.4	13.6		215.2	18.8		274.9
37	36.9	03.2	97	96.6	08.5	57	156.4	13.7	17	216.2	18.9		275.9
38	37.9	03.3	98	97.6	08.5		157.4	13.8		217.2	19.0		276.9
39	38.9	03.4	99	98.6	08.6		158.4	13.9		218.2	19.1		277.5
40	39.8	03.5	100	99.6		·	159.4	13.9		219.2	19.2	l I	278.9
41	40.8	03.6		100.6	08.8		160.4	14.0		220.2	19.3		279.9
42	41.8	03.7		101.6	08.9	62	161.4	14.1		221.2	19.3		280.9
43	42.8	03.7 03.8		102.6 103.6			162.4 163.4	14.2 14.3		222.2 223.1	19.4 19.5		281.9 282.9
44	43.8 44.8	03.8		104.6			164.4			224.1	19.6		233.9
46	45.8	04.0		105.6			165.4			225.1	19.7		284.9
47	46.8	04.1		106.6	09.3		166.4	14.6		226.1	19.8	87	285.9
48	47.8	04.2		107.6	09.4		167.4	14.6		227.1	19.9		286.9
49	48.8	04.3		108.6	09.5		168.4	14.7		228.1	20.0		287.9
50	49.8	04.4	10	109.6	09.6	70	169.4	14.8	30	229 .1	20.0		288.9
51	50.8	01.4		110.6	09.7		170.3	14.9		230.1	20.1	291	289.9
52	51.8	04.5		111.6	09.8		171.3	15.0		231.1	20.2	92	290.9
53	52.8	04.6		112.6	09.8		172.3	15.1		232.1	20.3		291.9 292.9
54	53.8	04.7		113.6			173.3 174.3			233.1 234.1	20.4 20.5		292.9
55 56	54.8 55.8	04.8 04.9		114.6 115.6	10.0 10.1		175.3	15.3		234.1	20.6		294.9
57	56.8	05.0		116.6	10.1		176.3			236.1	20.7	97	295.9
58	57.8	05.1		117.6			177.3			237.1	20.7		296.9
59	58.8	05.1		118.5	10.4	79	178.3	15.6	39	238.1	20.8	90	297.9
60	59.8	05.2	20	119.5	10.5	-	179.3		40		20.9	1	298. 9
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.
												TFo	r 85 D

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TABLE II.

Difference of Latitude and Departure for 6 Degrees.

			V-24											
st.	Lat.	Dep.			رحند			Dep.	l					
1	01.0	00.1	61	60.7	06.4		120.3			180.0	18.9		239.7	25.2
2	02.0	00.2 00.3	62 63	61.7 62.7	06.5 06.6		121.3 122.3			181.0 182.0	19.0 19.1		240.7 241.7	25.3 25.4
4	04.0	00.4					123.3			83.0	19.2		242.7	25.5
5	05.0	00.5	65	64.6	06.8		124.3			84.0	19.3		243.7	25.6
6	06.0	00.6	66	65.6	06.9		125.3			85.0	19.4		244.7	25.7
7 8	07.0 08.0	00.7	67		07.0		126.3			186.0	19.5		245.6	25.8
9	09.0	00.8 00.9	68 69	67.6 68.6	07.1 07.2		127.3 128.3	13.4 13.5		187.0 188.0	19.7 19.8		246 . 6 247 . 6	25.9 26.0
10	09.9	01.0	70	69.6	07.3		129.3			189.0	19.9		248.6	26.1
11	10.9	01.1	71	70.6	07.4		130.3	13.7	I	90.0	20.0		249.6	26.2
12	14.9	01.3	72	71.6	07.5		131.3	13.8		90.9	20.1		250.6	26.3
13	12.9	01.4			07.6		132.3			191.9	20.2		251.6	26.4
14 15	13.9 14.9	01.5 01.6	74	73.6 74.6	07.7		133.3	14.0		192.9	20.3		252.6	26.6
16	15.9	01.7	75 76	75.6	07.8 07.9		134.3 135.3	14.1 14.2		93.9	20.4 20.5		253.6 254.6	26.7 26.8
17	16.9	01.8		76.6	08.0		136.2	14.3		95.9	20.6		255.6	26.9
18	17.9	01.9	78	77.6	08.2	38	137.2	14.4	98	96.9	20.7	58	256.6	27.0
19	18.9	02.0	79	78.6	08.3		138.2	14.5		97.9	20.8		257.6	27.1
20	19.9	02.1	80	79.6	08.4		139.2	14.6		98.9	20.9	1	258.6	·
21 22	20.9 21.9	02.2 02.3	81 82	80.6 81.6	08.5 08.6		140.2 141.2	14.7		199 9	21.0		259.6	27.3
2:3	22.9	02.4	83	82.5	08.7		142.2	14.8 14.9		200.9 201.9	21.1 21.2		260.6 261.6	27.5
24	23.9	02.5	84	83.5	08.8		143.2	15.1		02.9	21.3		262.6	
2.5	24.9	02.6	85	84.5	08.9		144.2	15.2	05	203.9	21.4	65	263.5	27.7
26 27	25.9 26.9	02.7 02.8	86	85.5	09.0		145.2	15.3		204.9	21.5		264.5	27.8
28	27 8	02.9	87 88	86.5 87.5	09.1 09.2		146.2 147.2	15.4 15.5		205.9 206.9	21.6 21.7		265.5 266.5	27.9
29	28.8	03.0	89	88.5	09.3		148.2	15.6		207.9	21.8		267.5	
- 30	29.8	03.1	90	89.5	09.4			15.7		208.8	22.0		268.5	28.2
31	30.8	03.2	91	90.5	09.5	151	150.2	15.8	211	209.8	22.1	271	269.5	28.3
32	31.8	03.3	65	91.5	09.6		151.2	15.9		210.8			270.5	28.4
3.1	32.8 33.8	03.4 03.6	93 94	92.5	09.7 09.8		152.2	16.0		211.8	22.3		271.5	
35	34.8	03.7	95	93.5 94.5	09.9		153.2 15 4.2	16.1 16.2		212.8 213.8	22.4 22.5		272.5 273.5	28.7
36	35.8	03.8	96	95.5	10.0		155.1	16.3		214.8	22.6		274.5	
37	36.3	03.9	97	96.5	10.1		156.1	16.4	172	215.8	22.7		275.5	
38 39	37.8 38.8	04.0 04.1	98		10.2		157.1			216.8	22.8		276.5	29.1
40	39.8	04.2	99 100	98.5 99.5	10.3 10.5			16.6 16.7		217.8 218.8	22.9 23.0		277.5 278.5	29.2 29.3
41	40.8	04.3	II	100.4	10.6		160.1	16.8	I	219.8	23.1		279.5	29.4
42	41.8	04.4		101.4	10.7		161.1	16.9		220.8	23.1		280.5	
43	12.8	04.5	03	102.4	10.8	63	162.1	17.0		221.8	23.3		281.4	29.6
₽r	43.8	04.6		103.4	10.9		163.1	17.1		222.8	23.4	84	282.4	
45 46	44.8 45.7	04.7 04.8		104.4 105.4	11.0 11.1		164.1	17.2		223.8	23.5		283.4	29.8
47	46.7	04.9		106.4	11.2		165.1 166.1	17.4 17.5		224.8 225.8	23.6 23.7		284.4 285.4	
1 8	47.7	05.0		107.4	11.3		167.1	17.6		226.8	23.8		286.4	
49	48.7	05.1		108.4	11.4		168.1	17.7		27.7	23.9	89	287.4	30.2
50	49.7	05.2		109.4	11.5		169.1	17.8		28.7	24.0	1	288.4	30.3
51 52	50.7 51.7	05.3 05.4	111	110.4 111.4	11.6 11.7	171	170.1	17.9		229.7	24.1		289.4	30.4
53	52.7	05.5		111.4	11.8		171 . 1 172 . 1	18.0 18.1		230.7 231.7	24.3		290.4	
54	53.7	05.6		113.4			173.0	18.2		232.7	24.4 24.5		291.4 292.4	
55	54.7	05.7		114.4	12.0	75	174.0	18.3		233.7	24.6		293.4	
56	55.7	05.9		115.4			175.0	18.4		234.7	24.7		294.4	30.9
57 58	56.7 57.7	06.0 06.1		116.4 117.4	12.2 12.3		176.0 177.0	18.5		235.7	24.8		295.4	
59	58.7	06.2		118.3	12.4		178.0	18.6 18.7		236.7 237.7	24.9 25.0		296.4 297.4	
50	59. 7	06.3	20	119.5	12.5	80	179.0	18.8	40	238.7	25.1	300	298.4	31.4
st.	Dep.	Lat.			Lat.			Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.
									•				- 81 D	

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TABLE II.

Difference of Latitude and Departure for 7 Degrees.

Dis	t., L	at.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.	Dist. Lat. Dep.
	1 0	1.0	00.1	61	60.5	07.4		120.1	14.7	181 179.7	22.1	241 239.2 29.4
1		2.0	00.2	62	61.5	07.6		121.1	14.9	82 180 . 6	22.2	42 240.2 29.5
1		3.0	00.4	63	62.5	07.7		122.1	15.0	83 181 .6	22.3	43 241.2 29.6 44 242.2 29.7
1		4.0	00.5	64	63.5	07.8		123.1	15.1 15.2	84 182.6 85 183.6	22.4 22.5	45 243.2 29.9
1		5.0	00.6	65	64.5	07.9 08.0		124.1 125.1	15.4	86 184 . 6	22.7	46 244.2 30.0
1		6.0 6.9	00.7	66 67	66.5	08.2		126.1	15.5	87 185.6	22.8	47 245 2 30 1
		7.9	01.0	68	67.5	08.3		127.0	15.6	88 186 . 6	22.9	48 246 .2 30 .2
1		8.9	01.1	69	68.5	08.4		128.0	15.7	89 187 . 6	23.0	49 247.1 30.3
1		9.9	01.2	70	69.5	08.5	30	129.0	15.8	90 188 . 6	23.2	50 248.1 30.5
-	1 1	0.9	01.3	71	70.5	08.7	131	130.0	16.0	191 189.6	23.3	251 249.1 30.6
		1.9	01.5	72	71.5	08.8		131.0	16.1	92 190.6	23.4	52 250.1 30.7
		2.9	01.6	73	72.5	08.9	33	132.0	16.2	93 191.6		53 251.1 30.8
1	4 1	3.9	01.7	74	73.4	09.0	34	133.0	16.3	94 192.6		54 252.1 31.0
		4.9	01.8	75	74.4	09.1		134.0	16.5	95 193.5		55 253.1 31.1
		5.9		76	75.4	09.3		135.0	16.6	96 194.5		56 254.1 31.2
		6.9		77	76.4	09.4		136.0	16.7	97 195.5	24.0 24.1	57 255.1 31.3 58 256.1 31.4
		7.9		78	77.4	09.5 09.6		137.0 138.0	16.8 16.9	98 196 :5 99 197 .5		59 257.1 31.6
		8.9 9.9		79 80	78.4	09.0		139.0	17.1	200 198.5		60 258.1 31.7
1-	_ _				79.4							261 259.1 31.8
		8.0 8.1		81 82	80.4	09.9		139.9 140.9	17.2 17.3	201 199.5 02 200.5		62 260.0 31.9
		2.8			81.4 82.4			141.9	17.4	03 201.5		63 261.0 32.1
		3.8		84	83.4			142.9		04 202.5		64 262.0 32.2
		4.8			84.4			143.9		05 203.5		
		5.8			85.4			144.9		06 204.5	25.1	
		6.8			86.4			145.9		07 205 . 5	25.2	67 265.0 32.5
!!!	28 2	7.8	03.4	88	87.3	10.7		146.9		08 206.4	25.3	68 266.0 32.7
		8.8			88.3			147.9		09 207.4		
-:	30 _2	9.8	03.7	90	89.3		50	148.9	18.3	10 208.4		70 268.0 32.9
		0.8			90.3	11.1		149.9				271 269.0 33.0
		11.8		92				150.9				72 270.0 33.1
		2.8			92.3	11.3		151.9		13 211.4		
		3.7			93.3			152.9				
	- 1	4.7 5.7		95 96	94.3 95.3		56	153.8 154.8		15 213.4 16 214.4	26.2 26.3	B - 1
		6.7					57			17215.4)
		7.7						156.8				78 275.9 33.9
		8.7		99				157.8		19217.4		79 276.9 34.0
		19.7				12.2		158.8		20 218.4		80 277.9 34.1
l	41	0.7	05.0	101	100.2	12.3	161	159.8	19.6	221 219.4	26.9	281 278.9 34.2
		11.7			101.2	1		160.8				82 279.9 34.4
		12.7			102.2			161.8	19.9	23 221.3	27.2	83 280.9 34.5
		13.7			103.2			162.8	20.0	24 222.5		
		4.7			104.2			163.8		25 223.3		
	46 4	15.7	05.6		105.2			164.8				
		6.6			106.2			165.8				
		17.6			107.2			166.7		28 226. 29 227.		
	1	18.6 19.6			108.2 109.2			167.7 168.7		30 228.		
ı —										I		
		0.6			110.2			169.7				
		51.6 52.6			111.2			171.7				
		53.6			113.2			172.7				0 1
		54.6			114.1			173.7				8
		55.6			115.1			174.7				
		56.6	06.9	17	116.1	14.3	77	175.7	21.6			97 294.8 36.2
	58	57.6	07.1	18	117.1	14.4	78	176.7	21.7	38 236 .9	29.0	98 295.8 36.3
		58.6			118.1			177.7				99 296.8 36.4
	60 4	59.6			119.1		80	178.7	21.9	40 238.5		300 297.8 36.6
Di	st. I	ep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep.	Lat.	Dist. Dep. Lat.
												[For 83 Degrees

TABLE II.

Difference of Latitude and Departure for 8 Degrees.

. Lat.	Dep.	Dist.	Lat.	Dep. 1	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.
01.0	00.1	61	60.4	08.5		119.8	16.8		179.2	25.2	241 238.7	·
02.0	00.3	62	61.4	08.6		120.8	17.0		160.2	25.3	42 239 .6	
1 03.0	00.4	63	62.4	08.8		121.8	17.1	83	181.2	25.5	43 240.6	
i, 04.0	00.6	64	63.4	08.9	24	122.8	17.3		182.2	25.6	44 241.6	
05.0	00.7	65	64.4	09.0		123.8	17.4		183.2	25.7	45 212.0	
05.9	00.8	66	65.4	09.2		124.8	17.5		184.2	25.9		
06.9	01.0	67	66.3	09.3		125.8	17.7		185.2	26.0		
07.9	01.1	68	67.3	09.5 09.6		126.8 127.7	17.8 18.0		186.2 187.2	26.2 26.3	48 245.0 49 246.0	
05.9	01.3 01.4	69 70	68.3 69.3	09.7		128.7	18.1		188.2	26.4	50 247.	
10.9	01.5	71	70.3	09.9		129.7	18.2		189.1	26.6	251 248.0	
11.9	01.5	72	70.3	10.0		130.7	18.4		190.1	26.7	52 249.	
12.9	01.8	73	72.3	10.2		131.7	18.5		191.1	26.9	53 250 .	
1 13.9	01.9	74	73.3	10.3		132.7	18.6		192.1	27.0	54 251 .5	
14.9	02.1	75	74.3	10.4		133.7	18.8		193.1	27.1	55 252.5	
15.8	02.2	76	75.3	10.6		134.7	18.9	96	194.1	27.3	56 253.5	35.6
16.8	02.4	77	76.3	10.7		135.7	19.1		195.1	27.4	57 254.8	
17.8	02.5	78	77.2	10.9		136.7	19.2		196.1	27.6	58 255 .	
18.8	02.6	79	78.2	11.0		137.7	19.3		197.1	27.7	59 256 .	
19.8	02.8	80	79.2	11.1		138.6	19.5		198.1	27.8	I	
1 20.8	02.9	81	80.2	11.3		139.6	19.6		199.0		261 258.	
2 21.8	03.1	82	81.2	11.4		140.6	19.8		200.0		62 259.	
3 22.8 1 23.8	03.2	83	82.2	11.6 11.7		141.6	19.9		201.0 202.0			
24.8	$03.3 \\ 03.5$	84 85	83.2 84.2	11.8		142.6 143.6	20.0 20.2		202.0		64 261.4 65 262.4	
25.7	03.6	86	85.2	12.0		144.6	20.3		204.0		66 263.4	
26.7	03.8	87	86.2	12.1		145.6	20.5		205.0		67 264.	
27.7	03.9	88	87.1	12.2		146.6	20.6		206.0		68 265.	
28.7	04.0	89	88.1	12.4		147.5	20.7	09	207.0		69 266.4	
29.7	04.2	90	89.1	12.5	50	148.5	20.9	10	208.0	29.2	70 267	37.6
30.7	04.3	91	90.1	12.7	151	149.5	21.0	211	208.9	29.4	271 268.4	37.7
31.7	04.5	92	91.1	12.8		150.5			209.9	29.5	72 269.4	
3 32.7	04.6	93	92.1	12.9		151.5			210.9			
H 33.7	04.7	94	93.1	13.1		152.5			211.9			
34.7	04.9	95	94.1	13.2		153.5			212.9			
35.6	05.0	96 97	95.1	13.4		154.5			213.9			
36.6 37.6	05.1 05.3	98	96.1 97.0	13.5 13.6		155.5 156.5			214.9 215.9		77 274.5	
38.6	05.4	99	98.0	13.8		157.5			216.9		79 276	
39.6	05.6	100	99.0	13.9		158.4			217.9		80 277	
40.6	05.7		100.0	14.1		159.4	22.4		218.8		281 278.	
41.6	05.8		101.0	14.2		160.4	22.5		219.8			
42.6	06.0		102.0	14.3		161.4			220.8			
43.6	06.1		103.0	14.5		162.4			221.8			
44.6	06.3	05	104.0	14.6	65	163.4	23.0	25	222.8	31.3	85 282	2 39.7
45.6	06.4		105.0	14.8		164.4			223.8			
46.5	06.5		106.0	14.9		165.4			224.8			
47.5	06.7		106.9	15.0		166.4			225.8			
48.5	06.8		107.9	15.2		167.4			226.8			
49.5	07.0		108.9	15.3		168.3		1!	227.8		·	
50.5	07.1		109.9	15.4		169.3			228.8			
51.5	07.2 07.4		110.9	15.6 15.7		170.3			229.7 230.7			
53.5	07.5		111.9 112.9	15.7		171.3 172.3			230.7			
54.5	07.7		113.9			173.3			232.7			
55.5			114.9			174.3			233.7			
56.4			115.9									
57.4	08.1	18	116.9	16.4	78	176.3	24.8	38	235.7	33.1	98 295.	1 41.5
58.4			117.8			177.3			236.7			
59.4	08.4	_		16.7		178.2			237.7		·	
Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep	. Lat.
			 -								[For 89]	Name = 1

[For 82 Degrees.

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• TABLE II.

Difference of Latitude and Departure for 9 Degrees.

ł														
Pi			Dist	_	Dep.						Dep.	Dist.		De
11	1 01.						119. 120.			178.8			238.0	37
11	2 02.						121.			179.8 180.7			239.0 240.0	37
	4 04						122.			181.7		44	241.0	38
11	5 04.					25	123.	19.6		182.7	! 2 8.9	45	242.0	38
	6 05.		1.				124.			183.7			243.0	
	7 06.1 8 07.1						125.4 126.4			184.7			244.0	
	8 07.1 9 08.1						127.			185.7 186.7			244 .9 245 .9	
	0 09						128.4			187.7			246.9	39
11-					11.1		129.4		IJ			·	247.9	39
II i					11.3		130.4			189.6			248.9	39
	3 12.				11.4		131.4			190.6			249.9	
!					11.6		132.4			191.6			250.9	39
;					11.7 11.9		133.5 134.5			192.6 193.6			251.9 252.8	39 40
ll i	-1 - a		77				135.5			194.6			253.8	40
	8 17.						136.3			195.6			254.8	40.
]]	9 18.		79	78.0	12.4		137.5		99	196.5	31.1	59.2	255.8	40.
2	0 19.	03.1	80	79.0	12.5	40	138.3	21.9	200	197.5	31.3	·	256.8	40.
2			81	80.0	12.7		139.3			198.5	31.4		57.8	40.
2			82	81.0	12.8		140.5			199.5	31.6		58.8	
2			83 84	82.0 83.0	13.0 13.1		141.2 142.2			200.5 201.5	31.8 31.9		59.8 60.7	41.
2			85	84.0	13.3		143.2			202.5	32.1		61.7	41.
20			86	84.9	13.5		144.2			203.5	32.2		62.7	41.
2			87	85.9	13.6		145.2		07	204.5	32.4	67 2	63.7	41.
25			88		13.8		146.2			205.4	32.5		64.7	41.
29			89	87.9	13.9		147.2			206.4 207.4	32.7 32.9		65.7 66.7	42. 42.
30	-		90	88.9	14.1		148.2							
31			91 92	89.9 90.9	14.2 14.4		149.1 150.1	23.6 23.8		20 8.4 2 09.4	33.0 33.2	271 2	68.7	42. 42.
33			•93	91.9	14.5		150.1 151.1	23.9		210.4	33.3		69.6	42.
34			94	92.8	.14.7		62.1	24.1		211.4	33.5		70.6	42.
35			95	93.8	14.9		153.1	24.2		212.4	33.6	75 2	71.6	43.
36			96	94.8	15.0		154.1	24.4		213.3	33.8	76 2	72.6	43.
37			97 98	95.8 96.8	15.2 15.3		55.1 56.1	24.6 24.7		214.3 215.3	33.9 34.1	782	73.6	43. 43.
39			99	97.8	15.5		57.0	24.9		216.3	34.3	79 2		43.
.40			100	98.8	15.6		58.0	25.0		217.3	34.4		76.6	43.
41	40.5	06.4	101	99.8	15.8	161	59.0	25.2	221	218.3	34.6	281 2	77.5	44.
45		06.6	02	100.7	16.0		60.0	25.3		219.3	34.7	82 2'	78.5	44.
43		06.7		101.7	16.1		61.0	25.5		220.3	34.9	83 2		44.
44		06.9 07.0		102.7 103.7	16.3		62.0 63.0	25.7 25.8		221.2 222.2	35.0 35.2	84 2 85 2		44.
46		07.2		104.7	16.6		64.0	26.0		223.2	35.4	8612		44.
47		07.4		105.7	16.7		64.9	26.1		24.2	35.5		83.5	44.
48	47.4	07.5	08	106.7	16.9	68 1	65.9	26.3		25.2	35.7	88 2		45.
49		07.7		107.7	17.1		66.9	26.4		26.2	35.8	89 2		45.
50		07.8		108.6	17.2		67.9	26.6		27.2	36.0	90 21		45.
51	50.4	08.0		109.6	17.4	171 1		26.8		28.2	36.1	291 28		45. 45.
52 53	51.4 52.3	08.1 08.3		110.6	17.5 17.7		69.9 70.9	26.9 27.1		29.1 20.1	36.3 36.4	92 28 93 28		45.
54	53.3	08.4		112.6	17.8		71.9	27.2		31.1	36.6	94 29		46.
55	54.3	08.6	15	113.6	18.0	75 1	72.8	27.4	35	32.1	36.8	95 29	91.4	46.
56	55.3	08.8		114.6	18.1		73.8	27.5		33.1	36.9	96 29		46.
57	56.3	08.9 09.1		15.6	18.3		74.8	27.7		34.1	37.1 37.2	97 29 98 29		46. 46.
58 59	57.3 58.3	09.1		116.5	18.5		75.8 76.8	27.8 28.0		36.1	37.4	99 29		46.
60	59.3	09.4		18.5	18.8		77.8	28.2		37.0	37.5	300 29		46.
Dist.	Dep.	Lat.			Lat.		Dep.	Lat.	Dist.	Dep.	Lat.	Dist. I	ep.	Lat
										,		[For 8		
						•						-	•	-

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Difference of Latitude and Departure for 10 Degrees.

ŀ	Di-d	. ¥ = 4	D	The state of	¥4	Des	Diet	Lat.	Den !	Dist.	Let	Den	n:	Tet	D-	_
ľ	Dist.	Lat.		Dist.		Dep.	Dist.	119.2	Dep. 21.0		Lat. 178.3	Dep.			Dep.	1
ı	1 2	01.0 02.0		61 62	60.1 61.1	10.8		120.1	21.0		178.3 179.2	31.4 31.6		237.3 238.3	41.8	-
l	3	03.0		63	62.0			121.1	21.4		180.2	31.8		239.3	42.2	l
ŀ	4	03.9		64	63.0	11.7	24	122.1	21.5		181.2	32.0		240.3	42.4	ı
l	5	04.9	00.9	65	64.0	11.3		123.1	21.7	85	182.2	32.1	45	241.3		١
	6	05.9		66	65.0	11.5		124.1	21.9		183.2	32.3		242 .3	42.7	ı
ŀ	7	06.9	01.2	67	66.0	11.6		125.1	22.1		184.2	32.5		243.2	42.9	١
l	8 9	07.9	01.4	68	67.0	11.8 12.0		126.1 127.0	22.2 22.4		185.1	32.6		244.2	43.1	١
1	10	08.9 09.8	01.6 01.7	69 70	68.0 68.9	12.0		127.0	22.4		186.1 187.1	32.8 33.0		245.2 246.2	43.2° 43.4	١
1				71	69.9	12.3		129.0	22.7					_		1
ı	11 12	10.8	01.9 02.1	72	70.9	12.5		130.0	22.7		188.1 189.1	33.2 33.3		247.2 248.2	43.6 43.8	l
ŀ	13	12.8	02.3	73	71.9	12.7		131.0	23.1		190.1	33.5		249.2	43.9	١
Ì	14	13.8	02.4	74	72.9	12.8		132.0	23.3		191.1	33.7		250.1	44.1	١
t	15	14.8	02.6	75	73.9	13.0		132.9	23.4		192.0			251.1	44.3	١
	16	15.8	02.8	76	74.8	13.2		133.9	23.6	96	193.0	34.0	56	252.1	44.5	1
l	17	16.7	03.0	77	75.8	13.4		134.9	23.8		194.0			253.1	44.6	1
l	18	17.7	03.1	78	76.8	13.5		135.9	24.0		195.0	34.4		254.1	44.8	ı
	19 20	18.7	03.3 03.5	79	77.8 78.8	13.7 13.9		136.9 137.9	24.1 24.3		196.0	34.6		255.1	45.0	١
ŀ	·	19.7		80			l				197.0	34.7		256.1	45.1	
	21 22	20.7 21.7	03.6 03.8	81	79.8 80 .8	14.1 14.2		138.9 139.8	24.5 24.7		197.9	34.9		257.0		l
ı	23	22.7	04.0		81.7	14.4		140.8	24.1		198.9 199.9	35.1 35.3		258.0		I
l	24	23.6			82.7			141.8	25.0		200.9	35.4		259.0 260.0		1
	25	24.6	04.3		83.7	14.8		142.8	25.2		201.9	35.6		261.6		
	26	25.6	04.5		84.7	14.9		143.8	25.4		202.9	35.8		262.0		I
1	27	26.6		87	85.7	15.1		144.8	25.5		203.9	35.9	67	262.9	45.4	I
	28	27.6			86.7	15.3		145.8	25.7		204.8			263.9		۱
ļ	29 30	28.6	05.0		87.6	15.5		146.7	25.9		205.8	36.3		264.9		١
l	1——	29.5	05.2	90	88.6	15.6		147.7	26.0		206.8	36.5	J	265.9		1
l	31	30.5	05.4	91	89.6	15.8		148.7	26.2	211	207.8	36.6		266.9	47.1	۱
١	32 33	31.5 32.5		92		16.0 16.1		149.7 150.7	26.4 26.6		208.8 209.8			267.9	47.2	
	34	33.5			92.6	16.3		151.7	26.7		210.7	37.2		268.9 269.8	47.4	1
ı	35	34.5		95	93.6	16.5		152.6	26.9		211.7	37.3		270 .8	47.8	ı
1	36	35.5	06.3	96	94.5	16.7	56	153.6	27.1	16	212.7	37.5	76	271.8	47.9	١
1	37	36.4			95.5	16.8		154.6	27.3	17	213.7	37.7	77	272.8	48.1	١
ŀ	38	37.4		98	96.5	17.0		155.6	27.4		214.7			273.8	48.3	١
	39 40	38.4 39.4	06.8 06.9	99	97.5 98.5			156.6	27.6		215.7			274.8		١
						17.4		157.6	27.8		216.7		l	275.7	48.6	
	41 42	40.4 41.4	07.1 07.3	101	99.5	17.5	161	158.6	28.0					276.7	48.8	١
	43	42.3			100.5 101.4	17.7 17.9	62	159.5 160.5	28.1 28.3		218.6			277.7		l
	44	43.3			102.4			161.5	28.5		219.6 220.6	38.7 38.9		278.7 279.7	49.1	l
	45	44.3			103.4	18.2		162.5	28.7		221.6	39.1		280.7	49.5	
ŀ	46	45.3	08.0	06	104.4	18.4		163.5	28.8		222.6	39.2		281.7	49.7	١
	47	46.3			105.4	18.6	67	164.5	29.0		223.6	39.4		282.6	49.8	1
ľ	48	47.3			106.4	18.8		165.4	29.2		224.5			283.6		1
ŀ	49 50	48.3 49.2	08.5 08.7		107.3	18.9		166.4	29.3		225.5	39.8		284.6	50.2	1
l					108.3	19.1		167.4	29.5		226.5	39.9		285.6	50.4	
	51 52	50.2 51.2	08.9		109.3	19.3		168.4	29.7		227.5	40.1		286.6	50.5	1
	53	51.Z	09.0 09.2		110.3 111.3	19.4 19.6		169.4 170.4	29.9 30.0		228.5	40.3		287.6		
	54	53.2	09.4		112.3	19.8		171.4	30.0 30.2		229.5 230.4	40.5 40.6		288.5 289.5		
	55	54.2	09.6		113.3	20.0		172.3	30.4		230.4 231.4	40.8		290.5		
	56	55.1	09.7	16	114.2	20.1		173.3	30.6		232.4	41.0		291.5		١
	57	56.1	09.9		115.2	20.3		174.3	30.7		233.4	41.2		292.5		l
	58	57.1	10.1		116.2	20.5		175.3	30.9		234.4	41.3	98	293.5	51.7	l
	5 9	58.1 59.1	10.2		117.2	20.7		176.3	31.1		235.4	41.5		294.5		١
1			10.4		118.2	20.8		177.3	31.3		236.4	41.7		295.4	52.1	1
_	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.		Lat	1
													[Fo	80 D	egrees.	_

fror an negrees

Difference of Latitude and Departure for 11 Degrees.

-	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat	Dep.	Dist	Lat.	Dep.	:Dist.	Lat.	Dep.
	1	01.0	00.2	61	59.9	11.6		118.8	~	·	177.7	34.5		236.6	46.(
l	. 2	02.0	00.4	62	60.9	11.8		119.8			178.7	34.7		237.6	46.
ı	3	02.9	00.6	63	61.8	12.0		120.7	23.5		179.6	34.9	43	238.5	46.4
l	4	03.9		64	62.8	12.2		121.7	23.7		180.6	35.1		239.5	46.0
ļ	5	04.9		65	63.8	12.4		122.7	23.9		181.6	35.3		240.5	46.
İ	6	05.9 06.9	01.1 01.3	66 67	64.8 65.8	12.6 12.8		123.7 124.7	24.0 24.2		18 2 .6 18 3 .6	35.5		241.5 242.5	46.
l	8	07.9	01.5	68	66.8	13.0		125.6	24.4		184.5	35.7 35.9		243.4	47.
l	9	08.8	01.7	69	67.7	13.2		126.6	24.6		185.5	36.1		244.4	
ı	10	09.8	01.9	70	68.7	13.4		127.6	24.8		186.5	36.3		245.4	47
l	11	10.8	02:1	71	69.7	13.5	131	128.6	25.0	191	187.5	36.4	251	246.4	47.
Ĺ	12	11.8	02.3	72	70.7	13.7	32	129.6	25.2	92	188.5	36.6		247.4	48.
ı	13	12.8	02.5	73	71.7	13.9		130.6	25.4		189.5	36.8		248.4	48.
İ	14	13.7	02.7	74	72.6	14.1		131.5	25.6		190.4	37.0		249.3	48.
	15 16	14.7 15.7	02.9 03.1	75 76	73.6 74.6	14.3 14.5		132.5	25.8 26.0		191.4 192.4	37.2		250.3 251.3	
ŀ	17	16.7	03.2	77	75.6	14.7		133.5 134.5	26.1		193.4	37.4 37.6		252 3	
l	18	17.7	03.4	78	76.6	14.9		135.5	26.3		194.4	37.8		253.3	
ľ	19	18.7	03.6	79	77.5	15.1	1	136.4	26.5		195.3	38.0		254.2	
ı	20	19.6	03.8	80	78.5	15.3		137.4	26.7	200	196.3	38.2	60	255.2	49.
ı	21	20.6	04.0	81	79.5	15.5	141	138.4	26.9	201	197.3	38.4	261	256.2	49.1
ŀ	22	21.6	04.2	82	80.5	15.6		139.4	27.1		198.3	38.5		257.2	50.0
١	23	22.6	04.4	83	81.5	15.8		140.4			199.3	38.7		258.2	
ı	24	23.6	04.6	84	82.5	16.0		141.4			200.3	38.9		259.1	50.
	25 26	24.5 25.5	04.8 05.0	85 86	83.4 84.4	16.2 16.4		142.3	27.7 27.9		201. 2 202.2	39.1		260.1	50.0
	27	26.5	05.2	87	85.4	16.6		144.3			203.2	39.3 39.5		261.1 262.1	50.1 50.1
	28	27.5	05.3	88	86.4	16.8		145.3			204.2	39.7		263.1	
	29	28.5	05.5	89	87.4	17.0		146.3			205.2			264.1	51.
	30	29.4	05.7	90	88.3	17.2	50	147.2	28.6	10	206.1	40.1	70	265.0	51.
H	31	30.4	05.9	91	89.3	17.4	151	148.2	28.8	211	207.1	40.3	271	266.0	51.
Н	32	31.4	06.1	92	90.3	17.6		149.2			208.1	40.5		267.0	
П	33	32.4	06.3	93	91.3	17.7		150.2			209.1	40.6		268.0	
	34	33 .4	06.5 06.7	94	92.3	17.9 18.1		151.2			210.1 211.0			269.0 269.9	
١	35 36	34 .4 35 .3	06.9	95 96	93.3 94.2	18.3		152.2 153.1	29.8		212.0			270.9	
۱	37	36.3	07.1	97	95.2	18.5		154.1	30.0		213.0	41.4		271.9	
l	38	37.3	07.3	98	96.2	18.7		155.1	30.1	18	214.0	41.6		272.9	
П	39	38.3	07.4	99	97.2	18.9	59	156.1	30.3	19	215.0	41.8		273.9	53.
Н	40	39.3	07.6	100	98.2	19.1	60	157.1	30.5	20	216.0	42.0	80	274.9	53.
П	41	40.2	07.8	101	99.1	18.3		158.0			216.9	42.2	281	275.8	53.1
ı	42	41.2	08.0		100.1	19.5		159.0			217.9	42.4		276.8	
Н	43	42.2	08.2		101.1	19.7		160.0			218.9	42.6		277.8	
П	44 45	43.2 44.2	08.4 08.6		102.1 103.1	19.8 20.0		161.0 16 2 .0			219.9 2 2 0.9	42.7 42.9		278.8 279.8	
Н	46	45.2	08.8		104.1	20.2		163.0		96	221.8	43.1		280.7	
Н	47	46.1	09.0		105.0	20.4		163.9	31.9		222.8			281.7	
П	48	47.1	09.2		106.0	20.6		164.9	32.1	28	223.8	43.5		282.7	
H	49	48.1	09.3		107.0	20.8	69	165.9			224.8		89	283.7	55.
	50	49.1	09.5	10	108.0	21.0	70	166.9	32.4		225.8	43.9	90	284.7	55
	51	50.1	09.7		109.0	21.2		167.9	32.6		226.8	44.1		285.7	56.
H	52	51.0	09.9		109.9	21.4		168.8	32.8	32	2 27 .7	44.3		286.6	55.
Ш	53	52.0	10.1		110.9	21.6		169.8			228.7	44.5		287.6	
	54 55	53.0 54.0	10.3 10.5		111.9 112.9	21.8 21.9		170.8 171.8			2 29 .7 2 3 0.7	44.6		288.6 289.6	56
	56	55.0	10.7		113.9	23.1		172.8	33 .6		231.7	45.0		290.6	56
ı	57	56.0	10.9		114.9	22.3		173.7	33.8		232.6	45.2		291.5	56.
	58	56.9	11.1		115.8	22.5		174.7	34.0	38	233.6	45.4	98	292.5	
۱	59	57.9	11.3		116.8	22.7		175.7	34.2		234.6			293.5	
	_60	58.9	11.4		117.8	22.9		176.7	34.3		235.6			294.5	
-	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.		Dep.	
ŀ					•								[Fo	r 79 D	egrees

Difference of Latitude and Departure for 12 Degrees.

Dist.	Lat.	Dep.	Dist	Lat.	Dep.	Dist	Lat.	Dep.	Dist.	Lat.	Dep	Dist.	Lat.	Dep.
ī	01.0	00.2	61	59.7			118.4			177.0		241		50.1
2	02.0						119.3			178.0			236.7	
3		00.6					120.3			179.0 180.0			237.7 238.7	
4		00.8					121.5 122.5			181.0			239.6	
5 6	04.9 05.9						123.2			181.9	38.7		240.6	
7	06.8						124.9			182.9	38.9		241.6	
8	07.8		68	66.5	14.1		125.2			183.9	39.1		242.6	
9	08.8						126.2			184.9			243.6	
_10	09.8		70				127.2			185.8	39.5	11	244.5	 _
11	10.8						128.1			186.8	39.7 39.9		245.5	
12	11.7 12.7						129.1 130.1			187.8 188.8			246.5 247.5	
13 14	13.7						131.1			189.8	40.3		248.4	
15	14.7							28.1		190.7	40.5		249.4	53.0
16	15.7		. 76	74.3	15.8		133.0			191.7	40.8		250.4	53.2
17	16.6						134.0	28.5		192.7	41.0		251.4	
18	17.6		78 79				135.0 136.0			193.7 194.7	41.2 41.4		252.4 253.3	
19 20	18.6 19.6						136.9			195.6	41.6		254 .3	
	20.5		H				137.9			196.6	41.8		255.3	
21 22	21.5						138.9			197.6			256.3	
23	22.5						139.9			198.6			257.3	
24	23.5				17.5	44	140.9	29.9		199.5		64	258.2	54.9
2 5	24.5						141.8			200:5			259.2	
26	25.4						142.8			201.5			260.2	
27 28	26.4 27.4		87 88		18.1 18.3		143.8 144.8			20 2 . 5 203 . 5	43.0 43.2		261.2 262.1	55.5 55.7
29	28.4						145.7			204.4			263.1	
30	29.3		90				146.7			205.4			264.1	56.1
31	30.3	06.4	91	89.0	18.9		147.7		211	206.4	43.9	271	265.1	56.3
32	31.3		92		19.1	52	148.7	31.6		207.4	44.1	72	266.1	56.6
33	32.3		93				149.7			208.3			267.0	
34	33.3						150.6			209.3			268.0	57.0
35 36	34.2 35.2	07.3 07.5	95 96		19.8 20.0		151.6 152.6			210.3 211.3	44.7		269.0 27 0.0	57. 2 57. 4
37	36.2	07.7	97				153.6			212.3	45.1		270.9	57.6
38	37.2	07.9	98				154.5			213.2	45.3		271.9	57.8
39	38-1	08.1	99	96.8	20.6		155.5	33.1		214.2	45.5		272.9	58.0
40	39.1	08.3	100	97.8	20.8		156.5	33.3		215.2	45.7		273.9	58.2
41	40.1	08.5	101	98.8	21.0	161	157.5	33.5		216.2	45.9		274.9	58.4
42	41.1	08.7	02		21.2		158.5	33.7		217.1	46.2		275.8	58.6
43 44	42.1	08.9 09.1		100.7 101.7	21.4 21.6		159.4 160.4	33.9 34.1		218.1 219.1	46.4 46.6		276 .8 277 .8	58.8 -59.0
45	44.0	09.4		102.7	21.8		161.4	34.3		220.1	46.8		278.8	59.3
46	45.0	09.6		103.7	22.0		162.4	34.5		221.1	47.0		279.8	59.5
47	46.0	09.8		104.7	22.2		163.4	34.7		222.0	47.2	87	280.7	59.7
48	47.0	10.0		105.7	22.5		164.3	34.9		223.0	47.4		281 . 7	59.9
49 50	47.9 48.9	10.2 10.4		106.6 107.6	22.7 22.9		165.3 166.3	35.1 35.3		224.0	47.6		282 · 7 283 · 7	60.1 60.3
51 52	49.9 50.9	10.6 10.8		108.6 109.6	23.I 23.3		167.3 168.2	35.6 35.8		26.0 26.9	48.0 48.2		284.6 285.6	60.5 60.7
53	51.8	11.0		110.5	23.5		169.2	36.0		27.9	48.4		286.6	60.9
54	52.8	11.2	14	111.5	23.7		170.2	36.2		28.9	48.7		287.6	61.1
55	<i>5</i> 3.8	11.4		112.5 13.5	23.9	75	171.2	36.4	35 2	29.9	48.9	95	288.6	61.3
	54.8	11.6			24.1		172.2	36.6		30.8	49.1		289.5	61.5
57 58	55.8 56.7	11.9 12.1		114.4	24.3 24.5		173.1 174.1	36.8		31.8	49.3		290.5	61.7
		12.3		116.4	24.7		175.1	37.0 37.2			49.5		291.5 292.5	62.0 62.2
		12.5	20	117.4	24.9		176:1	37.4			49.9		293.4	62.4
								Lat.						
													78 De	

Difference of Latitude and Departure for 13 Degrees.

								<u>.</u>			_			
Dist.			Dist.				Lat.			Lat.				Dep.
1	0.10	00.2		59.4			117.9	27.2		176.4	40.7		234.8	54.2
2		00.4		60.4			118.9			177.3	40.9		235.8	54.4
3		00.7		61.4			119.8			178.3	41.2		236.8	54.7
4	03.9	00.9		62.4			120.8			179.3			237.7	54.9
5	04.9	01.1		63.3			121.8			180.3	41.6		238.7	55.1
6	05.8	01.3		64.3	14.8		122.8	28.3		181.2	41.8 42.1		239.7 240.7	55.3
7	06.8	01.6		65.3			123.7	28.6		182.2				55.6
8	07.8	01.8		66.3			124.7	28.8		183.2	42.3		241.6	55.8
10	08.8 09.7	02.0		67.2			125.7 126.7	29.0 29.2		184.2 185.1	42.5 42.7		242.6 243.6	
			1	68.2										56.2
11	10.7	02.5		69.2			127.6	29.5		186.1	43.0		244.6	56.5
12	11.7	02.7		70.2	16.2		128.6			187.1	43.2		245.5	56.7
13	12.7	02.9	R	71.1	16.4		129.6			188.1	43.4		246.5	56.9
1 14	13.6			72.1	16.6		130.6	30.1		189.0	43.6		247.5	57.1
15		03.4		78.1	16.9		131.5	30.4		190.0	43.9		248.5	57.4
16	16.6	03.6		74.1	17.1		132.5	30.6		191.0	44.1		249.4	57.6
17	16.6			75.0			133.5	30.8		192.0	44.3		250.4	57.8
18 19				76.0	17.5		134.5 135.4			192.9	44.5		251.4 252.4	
20			01	77.0 77.9	17.8 18.0		136.4	31.3 31.5		19 3 .9	45.0		253.3	
			<u> </u>			1								
31	20.5		81	78.9	18.2		137.4			195.8	45.2		254.3	58.7
22	21.4			79.9	18.4		138.4	31.9		196.8	45.4		255.3	
23	22.4			80.9	18.7		139.3			197.8	45.7		256.3	59.2
24	23.4			81.8	18.9		140.3			198.8 199.7	45.9		257.2	59.4
25 26	24.4 25.3	05.6 05.8		82.8 83.8	19.1 19.3		141.3 142.3	32.6 32.8		200.7	46.1 46.3	80	258.2 259.2	59.6 59.8
27	26.8	06.1		84.8	19.6		143.2	33.1		201.7	46.6		2 60.2	60.1
28	27.3	06.3		85.7	19.8		144.2	33.3		202.7	46.8		261.1	60.3
29	28.3	06.5		86.7	20.0		145.2	33.5		203.6	47.0		262.1	60.5
30	29.2	06.7	90	87.7	20.2		146.2	33.7		204.6	47.2		263.1	60.7
31	30.2	07.0	ilI	88.7	20.5		147.1	34.0		205.6	47.5		264.1	
32	31.2	07.2		89.6	20.5		148.1	34.2		206.6	47.7		265.0	61.0 61.2
33	32.2	07.4	93	90.6	20.9		149.1	34.4		207.5	47.9		266.0	61.4
34	33.1	07.6		91.6	21.1		150.1	34.6		208.5	48.1		267.0	61.6
35	34.1	07.9	95	92.6	21.4		151.0	34.9		209.5	48.4		268.0	61.9
36	35.1	08.1	96	93.5	21.6		152.0	35.1		210.5	48.6		268.9	62.1
37	36.1	08.3	97	94.5	21.8		153.0	35.3		211.4	48.8		269.9	
38	37.0	08.5	98	95.5	22.0		154.0	35.5		212.4	49.0		270.9	
39	38.0	08.8	99	96.5	22.3		154.9	35.8	19	213.4	49.3		271.8	62.8
40	39.0	09.0	100	97.4	22.5	60	155.9	36.0	20	214.4	49.5	80	272.8	63.0
41	39.9	09.2	101	98.4	22.7	161	156.9	36.2	221	215.3	49.7	281	273.8	63.2
42	40.9	09.4	02	99.4	22.9		157.8	36.4		216.3	49.9		274.8	63.4
43	41.9	09.7		100.4	23.2		158.8	36.7		217.3	50.2		275.7	63.7
44	42.9	09.9		101.3	23:4		159.8	36.9		218.3	50.4		276.7	63.9
45	43.8	10.1		102.3	23.6		160.8	37.1	25	219.2	50.6	85	277.7	64.1
46	44.8	10.3		103.3	23.8	66	161.7	37.3		220.2	50.8		278.7	64.3
47	45.8	10.6		104.3	24.1	67	162.7	37.6		221.2	51.1		279.6	64.6
48	46.8	10.8		105.2	24.3		163.7	37.8		222.2	51.3		280.6	64.8
49	47.7	11.0		106.2	24.5		164.7	38.0		223.1	51.5		281.6	65.0
50	48.7	11.2	_10	107,2	24.7	70	165.6	38.2	30	224.1	51.7	90	282.6	65.2
51	49.7	11.5		108.2	25.0	171	166.6	38.5		225.1	52.0		283.5	65.5
52	50.7	11.7		109.1	25.2	72	167.6	38.7		226.1	52.2	92	284.5	65.7
63	51.6			110.1	25.4		168.6	38.9		227.0	52.4		285.5	65.9
54	52.6			111.1	25.6		169.5	39.1		228.0	52.6		286 . 5	66.1
55	53.6	12.4		112.1	25.9		170.5	39.4		229.0	52 .9		287.4	66.4
56	54.6	12.6		113.0	26.1		171.5	39.6		230.0	53.1		288 . 4	66.6
57	55.5	12.8		14.0	26.3		172.5	39.8		230.9	53.3		289 . 4	66.8
58	56.5	13.0			26.5		173.4	40.0		231.9	53.5		290 . 4	67.0
59	57.5	13.3		116.0	26.8		174.4	40.3		232.9	53.8		291 . 3	67.3
	58.5	13.5		116.9			175.4	40.5		233.8	54.0		292.3	67.5
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep. 1	Lat.			
												For	77 De	grees.
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Difference of Latitude and Departure for 14 Degrees.

Dist	. Lat.	Dep.	Dist	Lat.	Dep.	Dist	Lat.	Dep.	Dist. Lat.	Dep.	Dist.	Lat	Dep.
	01.0			-			117.4		· · · · · · · · · · · · · · · · · · ·	-		233.8	58.3
	01.9						118.4	29.5	82 176 .6		42	234.8	58.5
	02.9						119.3					235.8	58.8
	03.9 04.9						120.3 121.3					236 . 8 237 . 7	
i							122.3					238.7	59.5
7							123.2					239.7	
8							124.2					240.6	
			11		1 - 1		125.2					41.6	60.2
10	,					L	126.1					242.6	60.5
11 12			71	68.9			127.1	31.7	191 185.3 92 186.3			43.5	60.7
13							128.1 129.0		93 187.3			44.5 45.5	61.0
14							130.0		94 188.2		54 9	46.5	61.4
15		1 1				35	131.0	32.7	95 189.2	47.2		47.4	61.7
16				73.7			132.0		96[190.2	47.4		48.4	61.9
17							132.9 133.9		97 191.1 98 192.1		572	49.4	62.2 62.4
18 19			78 79				134.9	33.6	99 193.1		599	51.3	62.7
20	19.4		80		19.4		135.8	33.9	200 194.1		60	52.3	62.9
21	20.4		81	78.6	19.6		136.8	34.1	201 195.0	·		53.2	63.1
22	21.3		82	79.6	19.8		137.8	34.4	02 196.0			54.2	63.4
23		05.6	83				138.8	34.6	03 197.0	49.1		55.2	63.6
24			84		20.3		139.7	34.8	04 197.9			56.2	63.9
2 5 26	24.3 25.2		85 86	82.5 83.4	20.6 20.8		140.7 141.7	35.1 35.3	05 198.9 06 199.9			57.1 58.1	64.4
27	26.2		87	84.4	21.0		142.6	35.6	07 200.9			59.1	64.6
28	27.2	06.8	88	85.4	21.3		143.6	35.8	08 201.8	50.3		60.0	64.8
2 9	28.1	07.0	89	86.4	21.5		144.6	36.0	09 202.8			61.0	65.1
30	29.1	07.3	90	87.3	21.8		145.5	36.3	10 203.8	50.8		62.0	65.3
31	30.1	07.5	91	88.3	22.0		146.5	36.5	211 204.7	51.0	271 2		65.6
32	31.0		92	89.3	22.3 22.5		147.5	36.8	12 205.7	51.3		63.9 64.9	65.8
33 34	32.0 33.0	08.0 08.2	93 94	90.2 91.2	22.7		148.5 149.4	37.0 37.3	13 206.7 14 207.6	51.5 51.8	74.7	65.9	66.3
35	34.0		95	92.2	23.0		150.4	37.5	15 208.6		75 2	66.8	66.5
36	34.9	08.7	96	93.1	23.2	56	151.4	37.7	16 209.6	52.3	762	67.8	66.8
37	35.9	09.0	97	94.1	23.5		152.3	38.0	17 210.6	52.5		68.8	67.0
38 39	36.9	09.2	98 99	95.1 96.1	23.7 24.0		153.3	38.2	18 211.5			69.7 70.7	67.5
40	37.8 38.8	09.4 09.7	100	97.0	24.2		154.3 155.2	38.5 38.7	19212.5 20213.5	53.Q 53.2		71.7	67.7
41	39.8	09.9	101	98.0	24.4		156.2	38.9	221 214.4		2812		68.0
42	40.8	10.2	02	99.0	24.7		157.2	39.2	22 215.4	53.5 53.7		73.6	68.2
43	41.7	10.4	03	99.9	24.9		158.2	39.4	23 216.4	53.9	83 2	74.6	68.5
44	42.7	10.6		100.9	25.2		159.1	39.7	24 217.3	54.2			.68.7
45	43.7	10.9		101.9	25.4		160.1	39.9	25 218.3	54.4		76.5	68.9 69.2
46 47	44.6 45.6	11.1		102.9 103.8	25.6 25.9		161.1 162.0	40.2	26 219.3 27 220.3	54.7 51.9		77.5 78.5	69.4
48	46.6	11.6		104.8	26.1		163.0	40.6	28 221.2	55.2		79.4	69.7
49	47.5	11.9	09	105.8	26.4	69	164.0	40.9	29 222.2	55.4		80.4	69.9
50	48.5	12.1	10	106.7	26.6		165.0	41.1	30 223.2	55.6		81.4	70.2
51	49.5	12.3		107.7	26.9		165.9	41.4	231 224.1	55.9	291 2		70.4
52	50.5	12.6 12.8		108.7	27.1 27.3		166.9	41.6	32 225.1	56.1		83.3 84.3	70.6
53 54	51.4 52.4	13.1	- 1	109.6 110.6	27.6		167.9 168.8	41.9	33 226.1 34 227.0	56.4 56.6		85.3	71.1
55	53.4	13.6		111.6	27.8		169.8	42.3	35 228.0	56.9		86.2	71.4
56	54.3	13.5		112.6	28.1	76	170.8	42.6	36 229.0	57.1	96 2	87.2	71.6
57	55.3	13.8		113.5	28.3		171.7	42.8	37 230.0	57.3		88.2	71.9
58	56.3	14.0		114.5	28.5		172.7	43.1	38 230.9	57.6		89.1 90.1	72.1 72.3
59 60	57.2 58.2	14.3 14.5		115.5	28.8		173.7	43.5	39 231.9 40 232.9	57.8 58.1	300 2		72.6
Dist.	Dep.				Lat.				Dist. Dep.		Dist. 1		Lat.
	Эср.	-Jul. "	.,	Jep. 1	-Lat. (*)	J. 60 L.	2-6p. 1	wat fi	Dan Deh	mail. N	For 7		
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TABLE II.

Difference of Latitude and Departure for 15 Degrees.

1.				100 y				- cpa-				. 000.		
Di	st; Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep
$I \vdash$	1 01.0			58.9	15.8		116.9	31.3		174.8	46.8		232 .8	62.
П	2 01.9						117.8	31.6		175.8	47.1		233.8	62.
11	3 02.9 4 05.9		63 64	60.9 61.8	16.3 16.6	1	118.8 119.8	31.8 32.1		176.8 177.7	47.4		2 34 .7 2 35.7	62. 63.
Ш	4 05.9 5 04.8			62.8	16.8		120.7	32.4		178.7	47.9		236.7	63.
Н	6 05.8			63.8	17.1		121.7			179.7	48.1		237.6	63.
П	7 06.8			64.7	17.3		122.7			180.6	48.4		238 . 6	63.
Ш	87.07.7	02.1	68	65.7	17.6		123.6	33.1		181.6	48.7		2 39 · 5	64.
П	9 08.7			66,6			124.6	33.4		182.6	48.9		240 . 5	64.
1 I	10 09.7		70	67.6		1	125.6	33.6		183.5	49.2	i	241.5	64.
	11 10.6		71	68.6	18.4		126.5	33.9		184.5	40.4		242.4	
	2 11.6		72	69.5			127.5	34.2		185.5	49.7		243.4	
	3 12.6 4 13.5		73 74	70.5 71.5	18.9 19.2		128.5 129.4	34.4		186.4 187.4	50.0 50.2		244.4 245.3	65. 65.
	5 14.5		75	72.4			130.4			188.4			246.3	
	6 15.5			73.4	19.7		131.4	35.2		189.3	50.7		247.3	
	7 16.4		77	74.4	19.9		132.3	35.5		190.3	51.0	57	248 · 2	66.
	8 17.4		78	75.3	20.2		133.3	35.7		191.3	51.2		249 . 2	
	9 18.4		79	76.3	20.4		134.3	36.0		192.2			250.2	
1	0 19.5		80	77.3	20.7		135.2	36.2		193.2	51.8		251.1	
	1 20.3		81	78.2	21.0		136.2	36.5		194.2	52.0		252.1	67.
	2 21 3		82	79.2	21.2		137.2 138.1	36.8		195.1	52.3		253.1	
	3 22.2 4 23.2		83 84	80.2 81.1	21.5 21.7		139.1	37.0 37.3		196.1 197.0	52.5 52.8		254.0 255.0	
	5 24.1			82.1	22.0		140.1	37.5		198.0	53.1		256.0	
	6 25.1		86	83.1			141.0	37.8		199.0	53.3		256.9	68.
	7 26.1		87	84.0			142.0	38.0		199.9	53.6		257.9	69.
2			88	85.0	22.8		143.0	38.3		200.9	53.8		2 58.9	69.
2			89	86-0	23.0		143.9	38.6		201.9	54.1		259 .8	69.
3	_		90	86.9	28.3		144.9	38.8		202.8	54.4	·	260.8	69.
3			91	87.9	23.6		145.9	39.1		203.8	54.6		261.8	70.
32			92 93	88.9	23.8		146.8 147.8	89.3 39.6		204.8 205.7	54.9		262.7	70. 70.
33			94	89.8 90.8	24.1 24.3		148.8	39.9		206.7	55.1 55.4		263.7 264.7	70.
35			95	91.8	24.6		149.7	40.1		207.7	55.6		2 65 · 6	71.
36		09.3	96	92.7	24.8		150.7	40.4		208.6	55.9		266.6	71.
37		09.6	97	93.7	25.1		151.7	40.6		209.6	56.2		267 .6	71.
38			98	94.7	25.4		152.6	40.9		210.6	56.4		268.5	72.
39		10.1	99	95.6	25.6 25.9		153.6 154.5	41.2		211.5 212.5	56.7		269.5	72. 72.
40				96.6			155.5	41.7			56.9		270.5	
41		10.6 10.9	101	97.6 98.5	26.1 26.4		156.5	41.9		213.5 214.4	57.2 57.5		271.4 272.4	72. 73.
45		11.1		99.5	26.7		157.4	42.2		215.4	57.7		273.4	73.
44				100.5	26.9		158.4	42.4		216.4	58.0		274.3	73
40	43.5	11.6	05	101.4	27.2		159.4	42.7		217.3	58.2	85	275.3	73.
46	44.4	11.9		102.4			160.3	43.0		218.3			276.3	74.
47		12.2		103.4	27.7	67	161.3	43.2		219.3			277.2	74
48		12.4 12.7		104.3	28.0	50	162.3 163.2	43.5 43.7		220.2 221.2			278.2	74.
49 50		12.7		105.3 106.3	28.2 28.5		164.2	44.0		221.2 222.2	59.3 59.5		279.2 280.1	75
		13.2		107.2	28:7	_	165.2	44.3		223.1	59.8		281.1	75.
51 52	11	13.2		107.2	29.0		166.1	44.5		223.1 224.1	60.0		282.1	75
53		13.7		109.1			167.1	44.8		225.1	60.3		283 .0	
54		14.0		110.1		74	168.1	45.0	34	226.0	60.6	94	284.0	76.
55	53.1	14.2		111.1	29.8		169.0	45.3		227.0	60.8		284.9	76.
56		14.5		112.0	30.0		170.0	45.6		228.0	61.1		285.9	76.
57		14.8 15.0		113.0 114.0	30.5 30.5		171.0 171.9	45.8 46.1		228.9 229.9	61.3 61.6		286.9 287.8	76.1 77.
58		15.3		114.9	30.8		172.9	46.3		230.9	61.9		288.8	77.
5 9		15.5		115.9	31.1		173.9	46.6		231.8	62.1		289.8	77.
Dist				Dep.			Dep.			Dep.		1	Dep.	
DIST	Dep.			P-			Pr.						75 De	
								•				f - 01		Prec:

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TABLE II.

Difference of Latitude and Departure for 16 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
1	01.0	00.3	61	58.6	16.8		116.3			174.0	49.9		231.7	66.4
2	01.9	00.6	62	59.6	17.1	22	117.3	33.6		174.9	50.2	42	232.6	66.7
3	02.9	00.8	63		17.4		118.2			175.9	50.4	43	233.6	67.0
4	03.8		64		17.6		119.2	34.2		176.9	50.7		234.5	67.3
5		01.4	65	62.5			120.2	34.5		177.8	51.0		235.5	67.5
6	05.8		66	63.4			121.1	34.7		178.8	51.3		236.5	67.8
7 8	06.7	01.9 02.2	67	64.4			122.1 123.0	35.0 35.3		179.8 180.7	51.5 51.8		237.4	68.1
9	07.7 08.7	02.5	68 69	65.4 66.3			124.0			181.7	52.1	40	238.4 239.4	58.4 68.6
10	09.6	02.8	70	67.3	19.3		125.0	35.8		182.6	52.4	50	240.3	68.9
11	10.6	03.0	71	68.2	19.6		125.9	36.1		183.6			241.3	69.2
12	11.5			69.2	19.8		126.9			184.6	52.9	201	242.2	
13	12.5	03.6	73	70.2	20.1		127.8			185.5	53.2		243.2	69.7
14	13.5	03.9		71.1	20.4		128.8			186.5	53.5	54	244.2	70.0
15	14.4	04.1		72.1	20.7		129.8	37.2		187.4	53.7	55	245.1	70.3
16	15.4	04.4		73.1	20.9		130.7	37.5	96	188.4	54.0	56	246 . 1	70.6
17	16.3			74.0			131.7			189.4	54.3		247.0	
18	17.3	05.0		75.0			132.7	38.0		190.3	54.6		248.0	
19	18.3		79				133.6			191.3	54.9		249.0	
20	19.2	05.5	80		22.1		134.6		1	192.3	55.1		249.9	
21	20.2	05.8	81	77.9	22.3		135.5			193.2	55.4	261	250.9	71.9
22	21.1	06.1	82	78.8			136.5		02	194.2	55.7		251.9	
23 24	22.1 23.1	06.3	83 84	79.8 80.7			137.5 138.4			195.1	56.0 56.2		252.8	72.5 72.8
25	24.0	06.6 06.9	85				139.4			196 . 1 197 . 1			254.7	
26	25.0	07.2			23.7		140.3			198.0		66	255.7	73.3
27	26.0	07.4	87	83.6			141.3			199.0		67	256.7	73.6
28	26.9	07.7	88	84.6			142.3			199.9				73.9
29	27.9	08.0	89	85.6	24.5	49	143.2	41.1		200.9	57.6	69	258.6	74.1
30	28.8	08.3	90	86.5	24.8	50	144.2	41.3	10	201.9	57.9	70	259.5	74.4
31	29.8	08.5	91	87.5	25.1	151	145.2	41.6	211	202.8	58.2	271	260.5	74.7
32	30.8		92	88.4			146.1			203.8			261.5	
33	31.7		93			53	147.1	42.2	13	204.7	58.7	73	262.4	75.2
34	32.7			90.4		54	148.0	42.4		205.7		74	263.4	75.5
35	33.6	09.6	95				149.0			206.7	59.3		264.3	
36	34.6 35.6	09.9 10.2			26.5 26.7		150.0 150.9			207.6 208.6	59.5 59.8		265.3 266.3	
38	36.5	10.5		94.2	27.0		151.9			209.6	60.1		267.2	
39	37.5	10.7		95.2	27.3		152.8			210.5			268.2	
40	38.5	11.0		96.1	27.6		153.8			211.5			269.2	
41	39.4	11.3	101	97.1	27.8	161	154.8	44.4		212.4	60.9	T	270.1	77.5
42	40.4	11.6		98.0	28.1	62	155.7	44.7		213.4		82	271.1	77.7
43	41.3	11.9		99.0			156.7			214.4			272.0	
44	42.3	12.1	04	100.0	28.7	64	157.6	45.2	24	215.3	61.7	84	273.0	78.3
45	43.3			100.9	28.9		158.6	45.5	25	216.3	62.0		274.0	
46	44.2	12.7		101.9	29.2		159.6		26	217.2	62.3		274.9	
47	45.2	13.0		102.9			160.5	46.0		218.2	62.6		275.9	
48 49	46.1 47.1	13. 2 13.5		103.8 104.8	29.8 30.0		161.5 162.5			219.2 220.1	62.8 63.1		276.8 277.8	
50	48.1	13.8		105.7	30.0		163.4	46.9		220.1 221.1	63.4		278.8	78.9
51	49.0	14.1		106.7			164.4		II				279.7	80.2
52	50.0	14.3		106.7	30.6 30.9	79	165.3	47.1 47.4		222.1 223.0	63.7 63.9		219.1 280.7	80.2
53	50.9	14.6		108.6	31.1		166.3			223.0 224.0	64.2		281.6	
54	51.9			109.6	31.4		167.3			224.9	64.5		282.6	
55	52.9	15.2		110.5	31.7		168.2			225.9	64.8		283.6	
56	53.8	15.4	16	111.5	32.0	76	169.2	48.5	36	226.9	65.1	96	284.5	81.6
57	54.8	15.7		112.5			170.1	48.8	37	227.8	65.3		285.5	
58	55.8	16.0		113.4	32.5		171.1	49.1	38	228.8	65.6		286.5	
59	56.7	16.3		114.4	32.8		172.1			229.7	65.9		287.4	
60	57.7	16.5		115.4			173.0			230.7	66.2		288.4	
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.		Dep.	Lat.
_						-						Fo	F 74 I	erross.

[For 74 Degrees

Difference of Latitude and Departure for 17 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep. _b	Dist., Lat.	Dep.
1	01.0	00.3	61	58.3	17.8		115.7	35.4	181 1		52.9	241 230.5	70.2
2		00.6	62	59.3	18.1		116.7	35.7		74.0	53.2	42 231 .4	70.8
3			63	60.2	18.4		117.6	36.0		75.0	53.5	43 232 .4	71.0
5			64 65	$61.2 \\ 62.2$	18.7 19.0		118.6 119.5	36.3 36.5		76.0 76.9	53.8 54.1	44 233.3 45 234.3	71.
6			66	63.1	19.3		120.5	36.8		77.9	54.4	46 235.3	71.
7	06.7	02.0	67	64.1	19.6		121.5	37.1		78.8	54.7	47.236.2	72.
8			68	65.0	19.9		122.4	37.4		79.8	55.0	48 237 . 2	72.
9			69	66.0	20.2		123.4	37.7		80.7	55.3	49 238.1	72.
10	09.6		70	66.9	20.5	30	124.3	38.0	90 1	81.7	55.6	50 239 1	73.
11	10.5		71	67.9	20.8		125.3	38.3		82.7	55.8	251 240.0	73.
12	11.5		72	68.9	21.1		126.2	38.6		83.6	56.1	52 241.0	73.
13 14	13.4	03.8 04.1	73 74	69.8 70.8	21.3		127.2	38.9		84.6 85.5	56.4	53 241.9 54 242.9	74.
15	14.3	04.4	75	71.7	21.9		128.1 129.1	39.2 39.5		86.5	56.7 57.0	55 243.9	
16	15.3	04.7	76	72.7	22.2		130.1	39.8		87.4	57.3	56 244 .8	
17	16.3	05.0	77	73.6	22.5		131.0	40.1	1	88.4	57.6	57 245 .8	
18	17.2	05.3	78	74.6	22.8		132.0	40.3		89.3	57.9	58 246 .7	
19	18.2	05.6	79	75.5	23.1		132.9	40.6		90.3	58.2	59 247.7	
20	19.1	05.8	80	76.5	23.4		133.9	40.9		91.3	58.5	60 248 . 6	·
21	20.1	06.1	81	77.5	23.7		134.8	41.2		92.2	5র.৪	261 249.6	
22	21.0	06.4	82	78.4	24.0		135.8	41.5		193.2	59.1	62 250.6	
23 24	22.0 23.0	06.7 07.0	83 84	79.4 80.3	24.3 24.6		136.8 137.7	41.8 42.1		194.1 195.1	59.4 59.6	63 251.5 64 252.5	
25	23.9		85	81.3	24.9		133.7	42.4		196.0	59.9	65 253.4	
26	24.9	07.6	86	82.2	25.1		139.6	42.7		197.0	60.2	66 254.4	
27	25.8	07.9	87	83.2	25.4		140.6	43.0		198.0	60.5	67 255.3	
23	26.8	08.2	83	84.2	25.7		141.5	43.3		198.9	60.8	68 256.3	
29	27.7	08.5	89	85.1	26.0.		142.5	43.6		199.9	61.1	69 257.2	
30	28.7	08.8	90	86.1	26.3		143.4	43.9	l'	300-8	61.4	70 258.2	
31	29.6	09.1	91	87.0	26.6		144.4	44.1		3, 105	61.7	271 259.2	
32	30.6 31.6	09.4 09.6	92 93	88.0 88.9	26.9 27.2		145.4 146.3	44.4		202.7 203.7	62.0 62.3		
34	32.5	09.9	94	89.9	27.5		147.3	45.0		204.6	62.6		
35	33 - 5	10.2	95	90.8	27.8		143.2	45.3		205.6	62.9		
36	34.4	10.5	96	91.8	28.1		149.2			206.6	63.2		
37	35.4	10.8	97	92.8	28.4		150.1	45.9		207.5	63.4		
38	36.3	11.1	98	93.7	28.7		151.1	46.2		208.5	63.7	78 265.5	
39 40	37.3 38.3	11.4	99 100	94.7 95.6	28.9 29.2		152.1 153.0	46.5 46.8		209.4 210.4	64.0 64.3		
i	39.2	12.0	101	96.6	29.5		154.0	47.1		211.3	64.6	1,	-!
41	40.2	12.3	02	97.5	29.8		154.9			212.3	64.9		
43	41.1		03	98.5	30.1		155.9			213.3	65.2		
44	42.1	12.9	04	99.5	30.4	64	156.8	47.9	24	214.2	65.5	84 271.6	
45	43.0			100.4			157.8			215.2			
46	44.0	13.4		101.4	31.0		158.7			216.1	66.1		
47	44.9 45.9	13.7 14.0		102.3 103.3	31.3 31.6		159.7 160.7	48.8 49.1		217.1 218.0	66.4		
48 49	46.9			104.2	31.9		161.6			219.0			
50	47.8	14.6		105.2	32.2		162.6			220.0			
51	48.8	14.9		106.1	32.5		163.5	50.0	l:i-	220.9	67.5		
52	49.7	15.2		107.1	32.7		164.5		32	221.9	67.8	92 279.9	85
53	50.7	15.5	13	108.1	33.0		165.4			222.8	68.1		
54		15.8		109.0			166.4			223.8	68.4		
55				110.0			167.4			224.7 225.7	68.7		
56, 57		16.4 16.7		110.9 111.9			168.3 169.3			226.6 226.6		11	
58	55.5	17.0		112.8			170.2			227.6		11 - 1	
59	56.4	17.2		113.8			171.2	52 .3	39	223.6	69.9	99 285	
60	57.4		!	114.8	35.1	80	172.1		11	229.5		-11	
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.		
							•					[For 73 I	e gre

Difference of Latitude and Departure for 18 Degrees.

1	I	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
3 02.9 0 0.9 63 59.9 19.5 . 29117.0 38.0 8\$3174.0 56.6 6 43231.1 75.1 4 03.6 6 05.7 01.9 66 60.9 19.8 24117.9 38.3 84175.0 56.9 44232.1 75.1 6 05.7 01.9 66 62.8 20.4 25119.8 38.9 86176.9 57.2 45233.0 75.7 6 05.7 06.7 02.2 67 63.7 20.7 27120.8 39.9 85175.9 57.2 45233.0 75.7 6 09.8 60 2.8 69 65.6 21.3 29122.7 39.9 89179.7 55.4 49235.8 76.9 10 09.5 03.1 70 66.6 21.6 30123.6 40.2 90180.7 55.7 50257.8 77.3 11 10.5 34.7 77 26 8.5 22.2 32122.7 39.9 89179.7 55.4 49235.8 76.9 11 00 9.5 03.1 70 66.6 21.6 30123.6 40.2 90180.7 55.7 50257.8 77.3 13 12.4 04.0 73 69.4 22.6 33 126.5 41.1 933183.6 59.6 5324.6 77.9 21.1 14 03.7 72 68.5 22.2 32 125.5 40.8 92182.6 59.3 5229.2 77.9 13 12.4 04.0 73 69.4 22.6 33 126.5 41.1 93183.6 59.6 5324.6 78.2 11 16.5 0.4 9 76 71.8 23.2 35 128.4 41.7 95185.6 65.3 6223.5 77.7 3.2 23.2 35 128.4 41.7 95185.6 60.3 55242.5 78.8 11 17 16.2 05.3 77 73.2 23.2 35 128.4 41.7 95185.6 60.6 5243.5 79.8 11 17 16.2 05.3 77 73.2 23.5 37130.3 42.3 97187.4 60.9 5724.4 79.7 19 18.1 05.6 76.7 12.2 41.1 38131.1 42.6 99188.3 61.2 88245.4 79.7 19 18.1 05.9 79 75.1 24.4 99132.2 43.0 99189.3 61.5 69245.8 90.0 20 19.0 66.2 80 76.1 24.7 40133.1 43.3 200199.2 61.8 60.2 43.5 79.1 19 18.1 05.9 79 75.1 24.4 91331.4 13.3 200199.2 61.8 60.2 43.5 79.1 22.2 20.9 06.8 82 78.0 23.5 42155.1 43.9 02192.1 62.4 6224.9 28.0 0.9 2192.0 06.2 80 76.1 24.7 40133.1 13.4 4.8 0199.9 26.1 8 60247.3 80.0 29.2 12.0 06.5 81 77.0 23.0 44157.0 44.5 99189.3 62.4 60.6 6243.5 99.2 22.2 20.9 06.8 82 78.0 23.5 42155.1 43.9 02192.1 62.7 6224.2 62.9 18.0 32.3 12.9 18.0 32.3 12.3 12.4 4 10.9 18.3 12.4 4 10.9	1	ī	01.0	00.3	61	58.0	18.9	121	115.1	37.4	181	172.1	55.9		229.2	74.5
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44 41.8 13.6 01 98.9 32.1 64.156.0 50.7 24.213.0 69.2 84.270.1 87.8 45 42.8 13.9 05 99.9 32.4 651.56.9 51.0 25.214.0 69.5 85.271.1 88.1 46 43.7 14.2 06.100.8 32.8 66.157.9 51.3 26.214.9 69.8 86.272.0 88.4 47 44.5 07.101.8 33.1 66.157.9 51.3 26.214.9 69.8 86.272.0 88.7 48 45.7 14.8 08.102.7 33.4 68.159.8 51.6 27.215.9 70.1 87.273.0 88.7 49 46.6 15.1 09.103.7 33.7 69.160.7 52.2 29.217.8 70.8 89.274.9 89.3 50 47.6 15.5 10.104.6 34.0 70.161.7 52.5 30.218.7 71.1 90.275.8 89.5 51 48.5 16.1 12.106.5 34.6 72.163.6 53.2 32.20.6 77.1 91.277.7 90.2	1															
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46 43.7 14.2 06 100.8 32.8 66 157.9 51.3 26 214.9 69.8 86 272.0 88.4 47 44.7 14.5 07 101.8 33.1 67 158.8 51.6 27 215.9 70.1 87273.0 88.7 49 46.6 15.1 09 103.7 33.7 69 160.7 52.2 29 217.8 70.8 89 274.9 89.3 50 47.6 15.5 10 104.6 34.0 70 161.7 52.5 30 218.7 71.1 90 275.8 89.6 51 48.5 15.8 111 105.6 34.3 171 162.6 52.3 231 219.7 71.4 291 276.8 89.9 52 49.5 16.1 12 106.5 34.6 72 163.6 53.2 32 220.6 71.7 92 277.7 90.2 53 50.4 16.7 14 108.4 35.2 74 165.5 53.8 33 221.6 72.0 93 278.7 90.5 54 51.4 16.7 14 108.4 35.2 74 165.5 53.8 34 222.5 72.0 93 278.7 90.9 <tr< td=""><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td>156.0</td><td>50.7</td><td></td><td></td><td></td><td></td><td></td><td></td></tr<>									156.0	50.7						
47 44.7 14.5 07 101.8 33.1 67 158.8 51.6 27 215.9 70.1 87 273.0 88.77 48 45.7 14.8 08 102.7 33.4 68 159.8 51.9 20 216.8 70.5 88 273.9 89.0 89.74 99.3 50.4 76 15.5 10 104.6 34.0 70 161.7 52.2 29 217.8 70.8 89.274.9 89.3 51 48.5 15.8 111 105.6 34.3 171 162.6 52.3 231 219.7 71.4 291 276.8 89.9 52 49.5 16.1 12 106.5 34.6 72 163.6 53.2 32 220.6 71.7 92 277.7 90.2 53 50.4 16.4 13 107.5 34.9 73 164.5 53.5 33 221.6 72.0 92 277.7 90.2 55 52.3 17.0 15 109.4 35.5 75 166.4 <td>1</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td>	1															
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51 48.5 15.8 111 105.6 34.3 171 162.6 52.3 231 219.7 71.4 291 276.8 89.9 52 49.5 16.1 12 106.5 34.6 72 163.6 53.2 32 220.6 71.7 92 277.7 90.2 53 50.4 16.4 13 107.5 34.9 73 164.5 53.3 32 21.6 72.0 93 278.7 79.5 54 51.4 16.7 14 108.4 35.2 74 166.5 53.8 34 222.5 72.3 94 279.6 90.9 55 52.3 17.0 15 109.4 35.5 75 166.4 54.1 35 223.5 72.6 95 280.6 91.2 56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.8 57 54.2 17.6 17 111.3 36.2 78 36.2 37 225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5<	1	50	47.6	15.5										90	275.8	89.6
52 49.5 16.1 12 106.5 53.4.6 72 163.6 53.2 32 220.6 71.7 92.277.7 90.2 53 50.4 16.4 13 107.5 34.9 73 164.5 53.5 33 221.6 72.0 93 278.7 90.5 54 51.4 16.7 14 108.4 35.2 74 165.5 53.8 34 222.5 72.3 94 279.6 90.9 55 52.3 17.0 15 109.4 35.5 75 166.4 54.1 35 223.5 72.6 95 280.6 91.2 56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.8 57 54.2 17.6 17 111.3 36.2 78 169.3 54.7 37 7225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 1226.4 73.5 98 1283.4 92.1		51	48.5	15.8	111	105.6	34.3	171	162.6	52.3	231	219.7	71.4			
53 50.4 16.4 13 107.5 34.9 73 164.5 53.5 33 221.6 72.0 93 278.7 90.5 54 51.4 16.7 14 108.4 35.2 74 165.5 53.8 34 222.5 72.3 94 279.6 90.9 55 52.3 17.0 15 109.4 35.5 75 166.4 54.1 35 223.5 72.6 95 280.6 91.2 56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.5 57 54.2 17.6 17 111.3 36.2 77 168.3 54.7 37 1225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 226.4 73.5 98 283.4 92.1	1				12	106.5	34.6	72	163.6		32	220.6				
54 51.4 16.7 14 108.4 35.2 74 165.5 53.8 34 222.5 72.3 94 279.6 90.9 55 52.3 17.0 15 109.4 35.5 75 166.4 54.1 35 223.5 72.6 95 280.6 91.2 56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.5 57 54.2 17.6 17 111.3 36.2 77 168.3 54.7 37 225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 226.4 73.5 98 283.4 92.1	1			16.4	13	107.5	34.9	73	164.5		3 3	221.6	72.0	93	278.7	90.5
56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.5 57 54.2 17.6 17 111.3 36.2 77 168.3 54.7 37 225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 226.4 73.5 98 283.4 92.1	1	- 1		16.7	14	108.4	35.2	74	165.5	53.8	34	222.5	72.3	94	279.6	90.9
56 53.3 17.3 16 110.3 35.8 76 167.4 54.4 36 224.4 72.9 96 281.5 91.5 57 54.2 17.6 17 111.3 36.2 77 168.3 54.7 37 225.4 73.2 97 282.5 91.8 58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 226.4 73.5 98 283.4 92.1	1				15	109.4										
58 55.2 17.9 18 112.2 36.5 78 169.3 55.0 38 226.4 73.5 98 283.4 92.1					16	110.3								96	281.5	91.5
I to a late of the second seco								77	168.3					97	282.5	91.8
-1 1								78	169.3							
60 57.1 18.5 20 114.1 37.1 80 171.2 55.6 40 228.3 74.2 300 285.3 92.7					13	110.2		79	170.2	55.3						
	1	'.						80	1/1.2	00.6	40	228.3	14.2	300	285.3	92.1
The process of the pr	127	IDL.i	<i>⊃</i> ep. (Lat.	DIST.	Deb.	Lat.	11/15t.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.
[For 72 Degrees.														[Fo	r 72 D	egrees.

TABLE II.

Difference of Latitude and Departure for 19 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Del
1	00.9	00.3	61	57.7	19.9	121	114.4	39.1		171.1	58.9		227.9	78
2	01.9	00.7	62	58.6	20.2		115.4	39.7		172.1	59.3		228.8	78.
3	02.8	01.0	63	59.6	20.5	23		40.0		173.0	59.6		229.8	79.
4	03.8	01.3	64	60.5	20.8		117.2	40.4		174.0	59.9		230.7	79.
5	04.7	01.6	65	61.5	21.2		118.2	40.7		174.9	60.2	45	231.7	79.
6	05.7		66	62.4	21.5		119.1	41.0		175.9	60.6		232.6	80.
7	06.6	02.3	67	63.3	21.8		120.1	41.3	87	176.8	60.9	47	233.5	80.
8	07.6	02.6	68	64.3	22.1	28	121.0	41.7	88	177.8	61.2	48	234.5	80.
9	08.5	02.9	69	65.2	22.5		122.0	42.0		178.7	61.5		235.4	81.
10	09.5	03.3	70	66.2	22.8	30	122.9	42.3	90	179.6	61.9	50	236.4	81.
11	10.4	03.6	71	67.1	23.1	131	123.9	42.6	191	180.6	62.2		237.3	81.
12	11.3	03.9	. 72	68.1	23.4	32	124.8	43.0	92	181.5	62.5	52	238.3	82.
13	12.3	04.2	73	69.0	23.8		125.8	43.3		182.5	62.8		239.2	
14	13.2	04.6	74	70.0	24.1		126.7	43.6		183.4	63.2		240.2	82.
15	14.2	04.9	75	70.9	24.4		127.6	44.0		184.4	63.5		241.1	83.
16	15.1	05.2	76	71.9	24.7		128.6	44.3		185.3	63.8		242.1	83.
17	16.1	05.5	77	72.8	25.1		129.5	44.6		186.3	64.1		243.0	
18	17.0	05.9	78	73.8	25.4		130.5	44.9		187.2	64.5		243.9	
19	18.0	06.2	79 80	74.7	25.7		131.4	45.3		188.2	64.8		244.9	
20	18.9	06.5	-	75.6	26.0		132.4	45.6	1	189.1	65.1		215.8	1
21	19.9	06.8	81	76.6	26.4		133.3	45.9		190.0	65.4		246.8	
22	20.8	07.2	82	77.5	26.7		134.3	46.2		191.0	65.8		247.7	
23	21.7	07.5	83	78.5	27.0		135.2	46.6		191.9	66.1		248.7	
24	22.7	07.8	84	79.4	27.3		136.2	46.9		192.9	66.4		249.6	
2.5	23.6	08.1	85	80.4	27.7 28.0		137.1 138.0	47.2 47.5		193.8 194.8	66.7		250.6	
26 27	24.6 25.5	08.5 08.8	86 87	81.3 82.3	28.0		139.0	47.9		195.7	67.1 67.4		251.5 252.5	86. 86.
28	26.5	09.1	88	83.2	28.7		139.9	48.2		196.7	67.7		253.4	87.
29	27.4	09.4	89	84.2	29.0		140.9	48.5		197.6	68.0		254.3	87.
30	28.4	09.8	90	85.1	29.3		141.8	48.8		198.6	68.4		255.3	87.
1	29.3				29.6		142.8	49.2		199.5	68.7		256.2	
31	29.3 30.3	10.1 10.4	91 92	86.0 87.0	30.0		143.7	49.2		200.4	69.0		257.2	88.
33	31.2	10.4	93	87.9	30.3		144.7	49.8		201.4	69.3		258.1	88
34	32 . 1	11.1	94	88.9	30.6		145.6	50.1		202.3	69.7		259.1	89
35	33 .1	11.4	95	89.8	30.9		146.6	50.5		203.3	70.0		260.0	
36	34 .0	11.7	96	90.8	31.3		147.5	50.8		204.2	70.3		261.0	
37	35.0	12.0	97	91.7	31.6		148.4	51.1	17	205.2	70.6	77	261.9	90.
38	35.9	12.4	98	92.7	31.9		149.4	51.4	18	206.1	71.0	78	262.9	90.
39	36.9	12.7	99	93.6	32.2	59	150.3	51.8	19	207.1	71.3	79	263.8	90.
40	37.8	13.0	100	94.6	32.6	60	151.3	52.1	20	208.0	71.6	80	264.7	91.
41	38.8	13.3	101	95.5	32.9	161	152.2	52.4	221	209.0	72.0	281	265.7	91.
42	39.7	13.7	02	96.4	33.2	62	153.2	52.7	22	209.9	72.3	82	266.6	91.
43	40.7	14.0	03	97.4	33.5		154.1	53.1		210.9	72.6		267.6	
44	41.6	14.3	04	98.3	33.9		155.1	53.4		211.8	72.9		268.5	
45	42.5	14.7	05	99.3	34.2		156.0	53.7		212.7	73.3		269.5	
46	43.5	15.0		100.2	34.5		157.0	54.0		213.7	73.6		270.4	
47	44.4	15.3		101.2	34.8		157.9	54.4		214.6	73.9		271.4	
48	45.4	15.6		102.1	35.2		158.8	54.7		215.6 216.5	74.2		272.3 273.3	
49	46.3	16.0		103.1	35.5		159.8	55.0	30		74.6 74.9		274.2	
50	47.3	16.3		104.0	35.8		160.7	55.3	i				i	
51	48.2	16.6		105.0	36.1		161.7	55.7	231		75.2		275.1	94
52	49.2	16.9		105.9	36.5		162.6	56.0		219.4	75.5		276 . 1	95
53	50.1	17.3		106.8	36.8		163.6	56.3		220.3	75.9		277.0 278.0	
54	51.1	17.6		107.8	37.1		164.5	56.6		221.3 222.2	76.2		278.0	
55	52.0			108.7	37.4 37.8		165.5 166.4	57.0		222.2	76.5 76.8		279.9	
56	52.9 53.9	18.2 18.6		109.7 110.6	38.1		167.4	57.3 57.6		223.1	77.2		280.8	
57 58	54.8	18.9		111.6	38.4		168.3	58.0		225.0			281.8	
59	55.8	19.2		112.5	38.7		169.2	58.3					282.7	
60	56.7	19.5		113.5	39.1		170.2	58.6		226.9			283.7	
Dist.	Dep.		ll	Dep.		Dist.			li	Dep.	·		Dep.	10
Dist	Dep.	Dat.	.1.131.	Dep	ı,aı.	12/15(.	Dep.	Lat.	- I/IAL	Dep.			r 71 L	
							1					FLO		re ore

TABLE II.

Difference of Latitude and Departure for 20 Degrees.

Di-t	Lat.	Dep	. Dist	. Lat.	Dep.	Dist	. Lat.	Dep.	Dis	Lat.	Dep.	Dist	Lat.	Dep.	
	00.9		6	57.5	20.9	12	1113.7			1 170.1		24	1 226	5 82.4	-
9			6	2 58.3	21.2	29	2 114.6			2 171.0	62.9	2 4:	2 227.4	4 82.8	3
	02.8						3 115.6			3 172.0			3 228.		
4							1116.5			4 172.9			1 229 .: 5 2 30 .:		
							5 117.5 5 118.4			5 173.8 6 174.8			231.		
ì				- 1			119.3			7 175.7			232		
8			61			21	3 120 . 5	43.8		8 176.7		48	233.0	0 84.8	
.:							121.5			9 177.6			234.0		
_10		-	-11		-1	J1	122.2		· i ———	0 178.5	·	-"	234.	-1	_1
11							123.1	1	!!	1 179.5			235.		
12 13							124.0 125.0			2 130.4 3 181.4			2 236.8 3 237.7		
14							125.9			4 182.3			238.7		
15							126.9			183.2			239.6		
16				5 71.4	26.0		127.8	46.5	90	184.2		56	240.6	87.6	5
17							128.7			185.1			241.5		
18 19							129.7 130.6			3 186.1 9 187.0			242.4 243.4		
20							131.6						244.		
21		-			.'		132.0		i	188.9		-1	245		-1
22							133.4			189.8			246		
23							134.4			190.8			247.		
24			84	78.9	28.7		135.3	49.3	04	191.7	69.8	64	248.	90.3	
25							136.3			192.6			249.0		
26 27							137.2			193.6			250.0		
23		09.6			30.1		138.1 139.1	50.3 50.6		194.5 195.5	70.8 71.1		250.9 251.8		
29	27.3					1	140.0			196.4			252.8		
30	28.2						141.0			197.3			253.7		
31	29.1	10.6	91	85.5	31.1	151	141.9	51.6	211	198.3	72.2	271	254.7	92.7	7
32	30.1	10.9		86.5	31.5		142.8	52.0		199.2			255.6		
33							143.8	52.3		200.2	72.9		256.5		
34 35	31.9 32.9	11.6 12.0	94 95				144.7 145.7	52.7		201.1	73.2		257.5		1
36	33.8				32.8		146.6	53.0 53.4		202.0 203.0	73.5 73.9		258.4 259.4		
37	34.8	12.7	97	91.2	33.2		147.5	53.7		203.9	74.2		260.3		
38	35.7	13.0	98	92.1	33.5		148.5	54.0		204.9	74.6		261.2		
39	36.6	13.3	99	93.0	33.9		149.4	54.4		205.8	74.9		262.2		
40	37.6	13.7	100		34.2		150.4	54.7		206.7	75.2	!!	263.1		_
41 42	38.5 39.5	14.0 14.4	101	94.9	34.5		151.3	55.1	221		75.6				
43	40.4	14.7	02	95.8 96.8	34.9 35.2		152.2 1 53.2	55.4 55.7	22	208.6 209.6	75.9 76.3		265.0 265.9		1
44	41.3	15.0	04		35.6		154.1	56.1		210.5	76.6		266.9		
45	42.3	15.4	05	98.7	35.9	65	155.0	56.4		211.4	77.0		267.8		
46	43.2	15.7	06	99.6	36.3		156.0	56.8		212.4	77 .3	86	268.8	97.8	П
47 48	44.2	16.1		100.5	36.6		156.9	57.1		213.3	77.6		269.7		П
49	46.0	16.4 16.8		101.5 102.4	36.9 37.3		157.9 158.8	57.5 57.8		214.2 215.2	78.0		270.6		П
50	47.0	17.1		103.4	37.6		159.7	58.1		216.1	78.3 78.7		271.6 272.5		
51	47.9	17.4		104.3	38.0		160.7	58.5		217.1	79.0		273.5	J	-11
52	48.9	17.8		105.2	38.3		161.6	58.8		218.0	79.3		274.4		П
53	49.8	18.1	13	106.2	38.6		162.6	59.2		218.9	79.7	93	275.3	100.2	Н
54	50.7	18.5		107.1	39.0		163.5	59.5		219.9	80.0	94	276.3	100.6	П
55 56	51.7 52.6	18.8 19.2		108.1 109.0	39.3		164.4	59.9		220.8	80.4			100.9	$\ \ $
57	53.6	19.2		109.0	39.7		165.4 166.3	60.2		221.8 222.7	80.7 81.1		278. I	101.2	П
58	51.5	19.8		110.9	40.4		167.3	60.9		223.6	81.4			101.9	
59	55.4	20.2	19	111.8	40.7		68.2	61.2		224.6	81.7			102.3	
50	56.4	20.5		12.8	41.0	80	69.1	61.6	40	225.5	82.1	300	281.9	102.6	
ist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	11
														egrees.	

Difference of Latitude and Departure for 21 Degrees.

Dist	L. Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
1 .	1 00.9			56.9			113.0	43.4	1	169.0	64.9	241	225.0	
	2 01.9			57.9	22.2		113.9	43.7		169.9	65.2		225.9	86.7
	3 02.8			58.8	22.6		114.8	44.1		170.8	65.6		226.9	
	4 03.7			59.7	22.9		115.8			171.8	65.9		227.8	
	5 04.7	1		60.7			116.7	44.8		172.7	66.3		228.7	87.8
	6 05.6 7 06.5			62.5	23.7 24.0		117.6 118.6	45.2 45.5		173.6 174.6	66.7 67.0		229.7 230.6	
	8 07.3			63.5	24.4		119.5	45.9		175.5	67.4		231.5	
	9 08.4			64.4			120.4	46.2		176.4	67.7		232.5	89.2
1 4 20	0 09.5	03 .6	70	65.4	25.1	30	121.4	46.6		177.4	68.1	50	233.4	89.6
1	1 10.5	03.9	71	66.3	25.4	131	122.3	46.9	191	178.3	68.4	251	234.3	90.0
1:				67.2	25.8		123.2	47.3		179.2	68.8		235.3	90.3
1:				68.2			124.2			180.2	69.2		236.2	
1 14				69.1			125.1	48.0		181.1	69.5		237.1	91.0
1 13				70.0			126.0	48.4		182.0	69.9		238.1	91.4
111			76 77	71.0 71.9	27.2 27.6		127.0 1 2 7.9	48.7 49.1		183.0 183.9	70.2 70.6		239.0 239.9	
is				72.8	28.0		128.8	49.5		184.8	71.0		240.9	
1				73.8	28.3		129.8	49.8		185.8	71.3		241.8	
20	18.7	07.2	80	74.7	28.7		130.7	50.2		186.7	71.7		242.7	
2	19.6	07.5	81	75.6	29.0	141	131.6	50.5	201	187.6	72.0	261	243.7	93.5
25	20.5	07.9	82	76.6	29.4	42	132.6	50.9	02	188.6	72.4	62	244.6	93.9
2:				77.5			133.5			189.5	72.7		245.5	
24				78.4			134.4			190.5	73.1		246.5	
2.5				79.4 80.3			135.4			191.4	73.5		247.4	
27				81.2	30.8 31.2		136.3 137.2	52.3 52.7		192.3 193.3	73.8 74.2		248.3 249.3	
28				82.2	31.5		138.2	53.0		194.2	74.5		250.2	96.0
29				83.1	31.9		139.1	53.4		195.1	74.9		251.1	96.4
30	28.0	10.8	90	84.0	32.3		140.0	53.8		196.1	75.3		252.1	96.8
31	28.9	11.1	91	85.0	32.6	151	141.0	54.1	211	197.0	75.6	271	253.0	97.1
32	29.9		92	85.9	33.0		141.9	54.5	12	197.9	76.0		253.9	
33				86.8	33.3		142.8	54 .8		198.9	76.3		254.9	
34	31 .7	12.2	94	87.8	33.7		143.8	55.2		199.8	76.7		255.8	98.2
35 36	32.7 33.6	12.5 12.9	95 96	88.7 89.6	34.0 34.4		144.7 145.6	55.5 55.9		200.7 201.7	77.0 77.4		256.7 257.7	98.6 98.9
37	34 .5	13.3	97	90.6	34.8		146.6	56.3		202.6	77.8		258.6	
38	35 .5		98	91.5	35.1		147.5	56.6		203.5	78.1		259.5	99.6
39	36 .4	14.0	99	92.4	35.5	59	148.4	57.0		204.5	78.5	79	260.5	100.0
40	37.3	14.3	100	93.4	35 .8	60	149.4	5 7 · 3	20	205.4	78 .8	80	261.4	100.3
41	38.3	14.7	101	94.3	36.2		150.3	57.7		206.3	79.2		262.3	
42	39.2	15.1	02	95.2	36.6		151.2	58.1		207.3	79.6		263.3	
43	40.1 41.1	15.4 15.8	03 04	96.2 97.1	36.9 37.3		152.2 153.1	58.4 58.8		208.2 209.1	79.9 80.3		264 . 2 265 . 1	
45	42.0	16.1	05	98.0	37.6		154.0	59.1		210.1	80.6		266.1	
46	42.9	16.5	06	99.0	38.0		155.0	59.5		211.0	81.0		267.0	
47	43.9	16.8	07	99.9	38.3		155.9	59.8	27	211.9	81.3	87	26 7 .9	102.9
48	44.8	17.2		100.8	38.7		156.8	60.2		212.9	81.7		268.9	
49	45.7	17.6		101 .8	39.1		157.8	60.6		213.8	82.1		269.8	
50	46.7	17.9		102.7	39.4		158.7	60.9		214.7	82.4		270.7	
51	47.6	18.3		103.6	39.8		159.6	61.3		215.7	82.8		271.7	
52	48.5 49.5	18.6 19.0		104.6	40.1		160.6 161.5	61.6		216.6	83.1 83.5		272.6 273.5	
53 54	50.4	19.4		106.4	40.9		162.4	62.0 62.4		218.5	83.9		274.5	
55	51.3	19.7		07.4	41.2		163.4	62.7		219.4	84.2		275.4	
56	52.3	20.1	16	108.3	41.6	76	164.3	63.1	36	220.3	84.6	96	276.3	106.1
57	53.2	20.4		109.2	41.9		165.2	63.4		221.3	84.9		277.3	
58	54.1	20.8		10.2	42.3		66.2	63.8		222.2	85.3		278.2	
59	55.1 56.0	21.1 21.5		11.1	42.6 43.0		67.1	64.1		223.1	85.6 86.0		279 . 1 280 . 1	
60							Des	64.5					Dep.	
Dist.	Dep. !	Lat.	D18(1	ъер. 1	Lat.	Dist.	Dep.	Lat.	DISL.	Dep.			69 De	
												r or	שעניי	Preca.

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TABLE II.

Difference of Latitude and Departure for 22 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.	Dist. Lat.	Dep. :	Dist. Lat. D	ep.
1	00.9	00.4	61	56.6	22.9	121 112.2	45.3	181 167.8	67.8	241 223.5 9	0.3
2	01.9	00.7	62	57.5	23.2	22 113.1	45.7	82 168.7	68.2		0.7
3	02.8	01.1	63	58.4	23.6	23 114.0		83 169.7			1.0
4	03.7	01.5	64	59.3	24.0	24 115.0		84 170.6	68.9		1.4
5 6	05.6	01.9 02.2	65 66	60.3 61.2	24.3 24.7	25 115.9 26 116.8		85 171.5 86 172.5	69.3 69.7		1.8
7	06.5	02.6	67	62.1	25.1	27 117.8		87 173.4	70.1		2.5
8	07.4	03.0	68	63.0	25.5	28 118.7	47.9	88 174.3	70.4		2.9
9	08.3	03.4	69	64.0	25.8	29 119.6	48.3	89 175.2	70.8		3.3
10	09.3	03.7	70	64.9	26.2	30 120 . 5	48.7	90 176.2			3.7
11	10.2	04.1	71	65.8	26.6	131 121.5	49.1	191 177.1	71.5		4.0
12	11.1	04.5		66.8	27.0	32 122 .4		92 178.0	71.9		4.4
13	12.1	04.9	73	67.7	27.3	33 123.3		93 178.9	72.3	53 234.6 9	4.8
14 15	13.0 13.9	05.2 05.6	74 75	68.6	27.7	34 124.2 35 125.2		94 179.9 95 180.8	72.7 73.0		5.2
16	14.8	06.0		69.5 70.5	28.1 28.5	36 126.1			73.4		5.9
17	15.8	06.4	77	71.4	28.8	37 127.0		97 182.7	73.8		6.3
18	16.7	06.7	78	72.3	29.2	38 128.0		98 183.6	74.2		6.6
19	17.6		79	73.2	29.6	39 128.9	52.1		74.5		7.0
20	18.5	07.5	80	74.2	30 .0	40 129.8	52.4	200 185.4	74.9	60 241.1 9	7.4
21	19.5	07.9	81	75.1	30.3	141 130.7	52.8	201 186 . 4	75.3		7.8
22	20.4	08.2	82	76.0	30.7	42 131.7	53.2	02 187.3	75.7		8.1
23 24	21.3	08.6	83	77.0	31.1	43 132.6		03 188.2	76.0		98.5 98.9
23	22.3 23.2	09.0 09.4	84 85	77.9 78.8	31.5 31.8	44 133.5 45 134.4		04 189.1 05 190.1	76.4 76.8		9.3
26	24.1	09.7	86	79.7	32.2	46 135.4		06 191.0	77.2	66 246.6	
27	25.0			80.7	32.6	47 136.3			77.5	67 247.6 10	
28	26.0	10.5		81.6	33.0	48 137.2	55.4		77.9	68 248.5 10	00.4 📋
29	26.9	10.9	89	82.5	33.3	49 138.2	55.8	09193.8	78.3	69 249 . 4 10	
30	27.8	11.2	90	83.4	33.7	50 139.1	56.2	10 194.7	78.7	70 250 . 3 10	
31	28.7	11.6	91	84.4	34.1	151 140.0		211 195.6	79.0	271 251 . 3 10	
32	29.7	12.0	92	85.3	34.5	52 140.9		12 196.6	79.4	72 252 . 2 10	
33 34	30.6 31.5		93 94	86.2 87.2	34.8 35.2	53 141.9 54 142.8		13 197.5 14 198.4	79.8 80.2	73 253 . 1 10 74 254 . 0 10	
35	32.5		95		35.6	55 143.7	58.1	15 199.3	80.5	75 255.0 10	3.0
36	33.4		96	89.0	36.0	56 144.6	58.4		80.9	76 255.9 10	
37	34.3	13.9	97	89.9	36.3	57 145.6	58.8	17 201.2	81.3	77 256.8 10	3.8
38	35.2	14.2	98	90.9	36.7	58 146.5	59.2	18 202.1	81.7	78 257 .8 10	
39	36.2	14.6	99	91.8	37.1	59 147.4			82.0	79 258.7 10	
40	37.1	15.0	100	92.7	37.5	60 148.3		1	82.4	80 259 . 6 10	
41	38.0	15.4	101	93.6	37.8	161 149.3		221 204.9	82.8	281 260 . 5 10	
42 43	38.9 39.9	15.7 16.1	02	94.6 95.5	38.2 38.6	62 150.2 63 151.1		22 205.8 23 206.8	83.2 83.5	82 261 .5 10 83 262 .4 10	
44	40.8		04	96.4	39.0	64 152.1	61.1 61.4		83.9	84 263 . 3 10	
45	41.7	16.9	05	97.4	39.3	65 153.0		25 208.6	84.3	85 264.2 10	
46	42.7	17.2	06	98.3	39.7	66 153.9	62.2	26 209.5	84.7	86 265.2 10	07.1
47	43.6			99.2	40.1	67 154.8			85.0	87 266 . 1 10	
48	44.5	18.0		100.1	40.5	68 155.8		28 211.4	85.4	88 267 .0 10	
49 50	45.4 46.4	18.4 18.7		101.1 102.0	40.8 41.2	69 156.7 70 157.6	63.3	29 212.3 30 213.3	85.8	90 268 . 9 10	
								I			
51 52	47.3 48.2	19.1 19.5		102.9 103.8	41.6 42.0	171 158.5 72 159.5		231 214.2 32 215.1	86.5 86.9	291 269 . 8 10 92 270 . 7 10	
53	49.1	19.5		103.8	42.0	73 160.4				93 271.7	09.8
54	50.1	20.2		105.7	42.7	74 161.3				94 272 .6 1	10.1
55	51.0	20.6		106.6	43.1	75 162.3	65.6	35 217.9		95 273.5 1	10.5
56	51.9	21.0	16	107.6	43.5	76 163.2	65.9	36 218.8	88.4	96 274.41	10.9
57	52.8	21.4		108.5	43.8	77 164.1		37 219.7	88.8	97 275.41	
58	53.8	21.7 22.1		109.4	44.2	78 165.0		38 220.7	89.2	98 276 .3 1	
5 9 60	54.7 55.6	22.1		110.3 111.3	44.6 45.0	79 166.0 80 166.9		39221.6 40222.5	89.5 89.9	99 277 .2 1 300 278 .2 1	12.4
Dist.										Dist. Dep.	
2131.	Dep.	Latt.	·DIST.i	Dep.	Lett.	Dist. Dep.	LMI.	Dist. Dep.	LAL.	(For 69 Der	

[For 68 Degrees.

Difference of Latitude and Departure for 23 Degrees.

Dist.	. Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.
1	00.9	00.4	61	56.2	23.8	121	111.4	47.3		166.6	70.7	241 221.8	94.2
, 2				57.1	24.2		112.3	47.7		167.5	71.1	42 222.8	94.6
3	02.8			58.0	24.6		113.2	48.1		168.5	71.5	43 223.7	94.9
4 5	03.7 04.6			58.9 59.8	25.0 25.4		114.1 115.1	48.5 48.8		169.4 170.3	71.9 72.3	44 224.6 45 225.5	95.3 95.7
6	05.5			60.8	25.8		116.0	49.2		171.2	72.7	46 226.4	96.1
7	06.4	1		61.7	26.2		116.9	49.6		172.1	73.1	47 227 .4	96.5
8	07.4			62.6	26.6		117.8	50.0		173.1	73.5	48 228.3	96.9
10	08.3 09.2	03.5 03.9		63.5 64.4	27.0		118.7 119.7	50.4 50.8		174.0 174.9	73.8 74.2	49 229.2	97.3 97.7
					27.4		120.6					50 230.1	
11 12	10.1	04.3 04.7		65.4 66.3	27.7 28.1		121.5	51.2 51.6		175.8 176.7	74.6 75.0	251 231.0 52 232.0	98.1 98.5
13	12.0			67.2	28.5		122.4	52.0		177.7	75.4	53 232.9	98.9
14	12.9			68.1	28.9		123.3	52.4		178.6	75.8	54 233.8	99.2
15	13.8			69.0	29.3		124.3	52.7		179.5	76.2	55 234.7	99.6
16 17	14.7			70.0	29.7		125.2 126.1	53.1		180.4 181.3	76.6 77.0	56 235.6 57 236.6	
18	15.6 16.6			70.9	30.1 30.5		127.0	53.5 53.9		182.3	77.4	58 237.5	
19				72.7	30.9		128.0	54.3		183.2	77.8	59 238.4	
20		07.8	80	73.6	31.3		128.9	54.7		184.1	78.1	60 239.3	
21	19.3	08.2	81	74.6	31.6	141	129.8	55.1	201	185.0	78.5	261 240.3	102.0
22	20.3	08.6	82	75.5	32.0		130.7	55.5		185.9	78.9	62 241.2	
23	21.2	09.0		76.4	32.4		131.6	55.9		186.9	79.3	63 242.1	
24 25	22.1 23.0	09.4 09.8	84 85	77.3 78.2	32.8 33.2		132.6 133.5	56.3 56.7		187.8 188.7	79.7 80.1	64 243.0 65 243.9	
26	23.9	10.2	86	79.2	33.6		134.4	57.0		189.6	80.5	66 244.9	
27	24.9	10.5		80.1	34.0		135.3	57.4		190.5	80.9	67 245.8	
23	25.8	10.9	88	81.0	34.4	48	136.2	57 .8		191.5	81.3	68 246 . 7	104.7
29	26.7	11.3	89	81.9	34.8		137.2	58.2		192.4	81.7	69 247.6	
30	27.6	11.7	90	82.8	35.2		138.1	58.6		193.3	82.1	70 248.5	
31	28.5	12.1	91	83.8	35.6		139.0	59.0		194.2	82.4	271 249.5	
32 33	29.5 30.4	12.5 12.9	92 93	84.7 85.6	35.9 36.3		139.9 140.8	59.4 59.8		195.1 196.1	82.8 83.2	72 250.4 73 251.3	
34	31.3	13.3	94	86.5	36.7		141.8	60.2		197.0	83.6	74 252.2	
35	32.2	13.7	95	87.4	37.1	55	142.7	60.6		197.9	84.0	75 253.1	
36	33.1	14.1	96	88.4	37.5		143.6	61.0		198.8	84.4	76 254 . 1	
37	34.1 35.0	14.5 14.8	97 98	89.3 90.2	37.9 38.3		144.5 145.4	61.3		199.7 200.7	84.8 85.2	77 255.0 78 255.9	
39	35.9	15.2	99	91.1	38.7		46.4	62.1		201.6	85.6	792.6.8	
40	36.8	15.6	100	92.1	39.1		147.3	62.5		202.5	86.0	80 257.7	
41	37.7	16.0	101	93.0	39.5	161	148.2	62.9	221	203.4	86.4	281 258.7	109.8
42	38.7	16.4	02	93.9	39.9		149.1	63.3		204.4	86.7	82 259.6	
43	39.6	16.8	03	94.8	40.2		150.0	63.7		205.3	87.1	83 260.5	
44	40.5	17.2 17.6	04	95.7 96.7	40.6		151.0 151.9	64.1 64.5		206.2 207.1	87.5 87.9	84 261.4 85 262.3	
46	42.3	18.0	06	97.6	41.4		152.8	64.9		208.0	88.3	86 263.3	
47	43.3	18.4	07	98.5	41.8		153.7	65.3		209.0	88.7	87 264.2	112.1
48	44.2,	18.8	08	99.4	42.2		154.6	65.6		209.9	89.1	88 265 . 1	
49	45.1	19.1		100.3	42.6		155.6	66.0		210.8 211.7	89.5 89.9	89 266.0	
50	46.0	19.5		101.3	43.0		156.5	66.4				90 266.9	
51 52	46.9 47.9	19.9 20.3		102.2 103.1	43.4 43.8		157.4 158.3	66.8 67.2		212. 6 213.6	90.3 90.6	291 267.9 92 268.8	
53	48.8	20.7		104.0	44.2		159.2	67.6		214.5	91.0	93 269.7	
54	49.7	21.1		104.9	44.5		160.2	68.0	34	215.4	91.4	94 270 . 6	114.9
55	50.6	21.5		105.9	44.9		161.1	68.4		216.3	91.8	95 271.5	
56	51.5	21.9		106.8	45.3		162.0	68.8		217.2 218.2	92.2 92.6	96 272.5 97 273.4	
57 58	52.5 53.4	22.3		107.7	45.7		162.9 163.8	69.2 69.6		218.2	93.0	98 274.3	
59	54.3	23.1		109.5	46.5		64.8	69.9		220.0	93.4	99 275.2	116.8
60	55.2	23.4	20	110.5	46.9		65.7	70.3		220.9	93.8	300 276 . 2	
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep.	
												(For 67 D	0.00000

[For 67 Degrees.

TABLE II.

Difference of Latitude and Departure for 24 Degrees.

ĵĨ	Dist.	Lat.	Dep.	Dist	Lat.	Den. I	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat	Dep.
F	1	00.9		61	55.7	24.8	11	110.5	49.2		165.4	73.6		220.2	98.0
1	2	01.8	00.8	62		25.2	22	111.5	49.6	82	166.3	74.0	42	221.1	98.4
1	3			63		25.6	23	112.4			167.2	74.4	43	222.0	98.8
1	4 5							113.3 114.2	50.4 50.8		168.1 169.0	74.8 75.2		222.9 223.8	
	6			66	60.3			115.1	51.2		169.0	75.7		223.8 224.7	
İ	7	06.4	02.8	67	61.2	27.3	27	116.0	51.7	87	170.8	76.1	47	225.6	100.5
	8							116.9	52.1		171.7	76.5		226.6	
1	9 10		03.7 04.1	69 70				117.8 118.8	52.5 52.9		172.7 173.6	76.9 77.3		227.5 228.4	
+	11	10.0		71	64.9	28.9	·	119.7	53.3		174.5	77.7		229.3	
	12	1			65.8			120.6	53.7		175.4	78.1		230.2	
1	13	11.9	05.3	73	66.7	29.7	33	121.5	54.1	93	176.3	78.5	53	231.1	102.9
1	14 15							122.4 123.3	54.5 54.9		177.2 178.1	78.9 79.3		232.0 233.0	
1	16			75 76				123.3 124.2	55.3		178.1	79.7		233.0 233.9	
1	17	15.5	06.9	77	70.3	31.3	37	125.2	55.7	97	180.0	80.1	57	234.8	104.5
1	18	16.4	07.3	78	71.3	31.7	38	126.1	56.1		180.9	80.5	58	235.7	104.9
1	19 20	17.4 18.3		79 80		32.1 32.5		127.0 127.9	56.5 56.9		181.8 182.7	80.9 81.3		236.6 2 37.5	
1	21	19.2		81	74.0			128.8	57.3	!!!	183.6	81.8		238.4	
1	22	20.1	08.9	82	74.9	33.4	42	129.7	57.8	02	184.5	82.2	62	239.3	106.6
	23	21.0	09.4	83	75.8	33.8	43	130.6	58.2	03	185.4	82.6	63	240.3	107.0
1	24 25	21.9 22.8		84 85		34.2 34.6		131.6 132.5	58.6 59.0		186.4 187.3	83.0 83.4		241.2 242.1	
1	25 26	22.8 23.8		85 86				133.4			188.2	83.4		242.1 243.0	
1	27	24.7	11.0	87	79.5	35.4	47	134.3	59.8	07	189.1	84.2	67	243.9	108.6
1	28 90	25.6						135.2			190.0	84.6		244.8 945.7	
	29 30	26.5 27.4		89 90				136.1 137.0			190.9 191.8	85.0 85.4		245.7 246.7	
-	31	28.3		91	83.1	37.0		137.9	61.4	ll	192.8	85.8	i——	247.6	
	32	29.2		92			52	138.9	61.8	12	193.7	86.2		243.5	
1	33	30.1	13.4	93	85.0	37.8	53	139.8	62.2	13	194.6	86.6	73	249.4	111.0
1	34 35	31.1 32.0		94 95	85.9 86.8			140.7 141.6	62.6 63.0		195.5 196.4	87.0 87.4		250.3 251.2	
1	36	32.9		96	87.7	39.0		142.5	63.5		197.3	87.9		252.1	
	37	33.8	15.0	97	83.6	39.5	57	143.4	63.9	17	198.2	88.3	77	253.1	112.7
1	38 39	34.7 35.6	15.5 15.9	98 99	89.5 90.4			144.3 145.3	64.3 64.7		199.2 200.1	88.7 89.1		254.0 254.0	113.1 113.5
1	40	36.5		100	91.4			146.2	65.1		200.1 201.0	89.1		255.8	
	41	37.5		101	92.3	41.1	1	147.1	65.5	!!	201.9	89.9		256.7	
1	42	38.4	17.1	02	93.2	41.5	62	148.0	65.9	22	202.8	90.3	82	257.6	114.7
	43	39.3		03	94.1	41.9		148.9	66.3		203.7	90.7		258.5	
1	44	40.2 41.1	17.9 18.3	04 05	95.0 95.9			149.8 150.7	66.7 67.1		204.6 205.5	91.1 91.5			115.5 115.9
	46	42.0	18.7	06	96.8	43.1	66	151.6	67.5	26	206.5	91.9	86	261.3	116.3
	47	42.9	19.1	07	97.7	43.5	67	152.6	67.9	27	207.4	92.3	87	262.2	116.7
	48 49	43.9 44.8	19.5 19.9	08 09	98.7 99.6	43.9 44.3		153.5 154.4	68.3 68.7		208.3 209.2	92.7 93.1		263.1 264 0	117.1 117.5
	50	45.7	20.3		100.5	44.3		155.3	69.1		210.1	93.1			118.0
r	51	46.6	20.7		101.4	45.1		156.2	69.6	1	211.0	94.0	1		118.4
	52	47.5	21.2	12	102.3	45.6	72	157.1	70.0	32	211.9	94.4	92	266.8	118.8
	53	48.4	21.6		103.2	46.0		158.0	70.4		212.9	94.8			119.2
1	54 55	49.3 50.2	22.0 22.4		104.1 105.1	46.4 46.8		159.0 159.9	70.8 71.2		213.8 214.7	95.2 95.6			119.6 120.0
1	56	51.2	22.8	16	106.0	47.2	76	160.8	71.6	36	215.6	96.0	96	270.4	120.4
1	57	52.1	23.2	17	106.9	47.6	77	161.7	72.0	37	216.5	96.4	97	271.3	120.8
1	58 59	53.0 53.9	23.6 24.0		107.8 108.7	48.0 48.4		162.6 163.5	72.4 72.8		217.4 218.3	96.8 97.2			121.2 121.6
ĺ	60	54.8	24.4		109.6	48.8		164.4	73.2		219.3	97.6			122.0
D	ist.			Dist.					Lat.					Dep.	
_													TF or		

[For 66 Degrees.

Difference of Latitude and Departure for 25 Degrees.

	•								•					
Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
ī	00.9	00.4	61	55.3	25.8	121	109.7	51.1	181	164.0	76.5	241	218.4	
1 2		00.8	. 62	56.2	26.2		110.6	51.6		164.9	76.9	42	219.3	102.3
3	02.7	01.3	63	57.1	26.6		111.5	52.0		165.9	77.3		220.2	
4	03.6		64	58.0	27.0	24	112.4	52.4	84	166.8	77.8		221.1	
5	04.5		65	58.9	27.5		113.3	52.8		167.7	78.2		222.0	
6	05.4	02.5	66	59.8		26	114.2	53.2		168.6	78.6	46	223.0	104.0
7	06.3	03.0	67	60.7	28.3		115.1	53.7		169.5	79.0		223.9	
8	07.3	03.4	68	61.6			116.0	54.1		170.4	79.5		224.8	
9	08.2	03.8	69	62.5	29.2		116.9	54.5		171.3	79.9		225.7	
10	09.1	04.2	70	63.4	29.6		117.8	54.9		172.2	80.3		226.6	
11	10.0	04.6	71	64.3	30.0	131	118.7	55.4	101	173.1	80.7	251	227.5	
12			72	65.3	30.4		119.6	55.8		174.0	81.1		228.4	
13			73	66.2	30.9		120.5	56.2		174.9	81.6		229.3	
14			74	67.1	31.3		121.4	56.6		175.8	82.0		230.2	
15	1		75	68.0	31.7		122.4	57.1		176.7	82.4		231.1	
16			76	68.9	32.1		123.3	57.5		177.6	82.8		232.0	
17			77	69.8	32.5		124.2	57.9		178.5	83.3		232.9	
18			78	70.7	33.0		125.1	58.3		179.4	83.7		233.8	
19			79	71.6	33.4		126.0	58.7		180.4	84.1		234.7	
20			80	72.5	33.8		126.9	59.2		181.3	84.5		235.6	
	·													
21			81	73.4	34.2	141	127.8	59.6		182.2	84.9	261		110.3
22			82	74.3	34.7		128.7	60.0		183.1	85.4		237.5	
23			83	75.2	35.1		129.6	60.4		184.0	85.8		238.4	
24		10.1	84	76.1	35.5		130.5	60.9		184.9	86.2		239.3	
25		10.6	85	77.0	35.9		131.4	61.3		185.8	86.6		240.2	
		11.0	86	77.9	36.3		132.3	61.7		186.7	87.1		241.1	
27	24.5		87	78.8	36.8		133.2	62.1		187.6	87.5		242.0	
28			88	79.8	37.2		134.1	62.5		188.5	87.9			113.3
			89	80.7	37.6		135.0			189.4	88.3		243.8	
30		12.7	90	81.6	38.0		135.9	63.4	10	190.3	88.7	70	244.7	114.1
31		13.1	91	82.5	38.5	151	136.9	63.8		191.2	89.2	271	245.6	114.5
32			92	83.4	38.9	52	137.8	64.2	12	192.1	89.6	72	246.5	115.0
33			93	84.3	39.3	53	138.7	64.7	13	193.0	90.0			115.4
34			94	85.2	39.7	54	139.6	65.1	14	193.9	90.4			115.8
35			95	86.1	40.1	55	140.5	65.5		194.9	90.9	75	249.2	116.2
36		15.2	96	87.0	40.6	56	141.4	65.9		195.8	91.3	76	250.1	116.6
37			97	87.9	41.0	57	142.3	66.4	17	196.7	91.7	77	251.0	117.1
38			98	88.8	41.4	58	143.2	66.8	18	197.6	92.1	78	252.0	117.5
39			99	89.7	41.8	59	144.1	67.2		198.5	92.6	79	252.9	117.9
40	36.3	16.9	100	90.6	42.3	60	145.0	67.6	20	199.4	93.0	80	253.B	118.3
41	37.2	17.3	101	91.5	42.7	161	145.9	68.0	221	200.3	93.4	281	254.7	118.8
42	38.1		02	92.4	43.1		146.8	68.5	22	201.2	93.8	82	255.6	119.2
43	39.0	18.2	03	93.3	43.5		147.7	68.9	23	202.1	94.2			119.6
44	39.9	18.6	04	94.3	44.0		148.6	69.3	24	203.0	94.7			120.0
45	40.8	19.0	05	95.2	44.4		149.5	69.7	25	203.9	95.1			120.4
46	41.7	19.4	06	96.1	44.8		150.4	70.2	26	204.8	95.5	86	259.2	120.9
47	42.6	19.9	07	97.0	45.2		151.4	70.6	27	205.7	95.9	87	260.1	121.3
48	43.5	20.3	08	97.9	45.6	68	152.3	71.0	28	206.6	96.4	88	261.0	121.7
49	44.4	20.7	09	98.8	46.1		153.2	71.4	29	207.5	96.8	89	261.9	122.1
50	45.3	21.1	10	99.7	46.5		154.1	71.8	30	208.5	97.2	90	262.8	122.6
51	46.2	21.6	111	100.6	46.9	171	155.0	72.3	931	209.4	97.6	991	968 7	123.0
52	47.1	22.0		101.5	47.3		155.9	72.7		210.3	98.0			123.4
53		22.4		102.4	47.8		156.8	73.1		211.2	98.5			123.8
54	48.9	22.8		103.3	48.2		157.7	73.5		212.1	98.9			124.2
55	49.8	23.2		104.2	48.6		158.6	74.0		213.0	99.3			124.7
56	50.8	23.7		105.1	49.0		159.5	74.4		213.9	99.7			125.1
57	51.7	24.1		106.0	49.4		160.4	74.3			100.2			125.5
58	52.6	24.5		106.9	49.9		161.3	75.2		215.7				125.9
59	53.5	24.9		107.9	50.3		162.2	75.6			101.0			126.4
60	54.4	25.4		108.8	50.7		163.1	76.1			101.4			126.8
					1									h
Dist.	Dep.	Lat.	DISE.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.			
					1							IF o.	r 65 D	egrees.

TABLE II.

Difference of Latitude and Departure for 24 Degrees.

		Dil	teren	ce of	Laun	tude and Departure for 24 Degrees.
st.	Lat.	Dep.	Dist	Lat.	Dep. h	Dist, Lat. Dep. Dist. Lat. Dep. Dist. Lat. Dep
1	00.9	00.4	61	55.7	24 8	101 110 6 AU 9 181 165 A 73 6 241 2261 -
2	01.8	00.8	62	56.6	25.2	22 111.5 49.6 32 166.3 74.0 42
3	02.7	01.2	63	57.6	25.6	23112.4 50.0 83167.2 74.4 43222 0 24113.3 50.4 84168.1 74.8 44222
4	03.7	01.6	64	58.5 59.4	26.0	24 113.3 50.4 84 168.1 74.8 44 223.4 4
5	04.6	02.0	65	59.4	26.4	25 114.2 50.8 85 169.0 75.2 46 224 7 115.1 51.2 86 169.9 75.7 46 224 7 115.1
6	06.4		67	60.3 61.2 62.1	27.3	OF THE OF STATE OF ST
8	07.3	03.3	68	62.1	27.7	28 116.9 52.1 88 171.7 76.5 48 2.5
9	08.2	03.7	69	63.0	28.1	29 117.8 52.5 89 172 7 76.9 49
10	09.1	04.1	70			30 118.8 52.9 90 173.6 77.3 DO
11	10.0	04.5	71	64.9	28.9	131 119.7 53.3 191 174 5 77.7
12	11.0	04-9		65.8	29.3	30 120 6 53 7 00 tes 4 78 1
13	11.9				29.7	33 121.5 54.1 93 176.3 78.6
14	The second second	05.7		68.5	30.5	34 122.4 54.5 94177.2 73.9 58- 35 123.3 54.9 96178.1 79.3
16	14.6					35 123 .3 54 .9 96 178 .1 79 .3 36 124 .2 55 .3 96 179 .1 79 .1
17	15.5	06.9	77	70.3	31.3	
10	16.4	07.3 07.7	77 78	71.3	31.7	30 136 1 56 1 00 190 9 00 m
19	17.4	07-7	79	72.2	32.1	39 127 0 56 5 99181 8 90 9 40 127 9 56 9 200182 7 81 3
20	18.3	08.1	80	73.1		39 127 0 56.5 99181 8 80.9 40 127.9 56.9 200 82.7 81.3
21	19.2	08.5	81	74.0	32.9	141128 0 0 0 0 001103 6 81
22	20.1	08.9	82	74.9	33.4 33.8	42 129.7 57.8 02 184.5 82.2 43 130.6 58.2 05 185.4 82.6
24	21.9	09.8	84	76.7		
25	20.8	10.2	85	76.7	34.6	
25 26 27	23.8	10.6	86	78.6	35.0	46 133 4 50 4 06 188 2 33
27	24.7	11.0	87	79.5		47 134.3 59.8 67 189 1 84
28	25.6	11.4	88	80.4	35.8	48 135.2 60.2 08 190.0 04.5
30	27.4	12.2	89 90	81.3	36.2	
21	28.3	12.6	_	82.2		50 137.0 61.0 10 191.8 85
	29 2	13.0	91		37 4	A CONTRACTOR MALE AND ADDRESS OF THE PARTY O
014	10.1	13.4	93	84.0	37.4 37.8	4 52 138 9 61 8 12 193 7 15 8 53 139 6 52 2 15 194 5
64	31.1	13.8	94	350 5	38.2	2 64 140.7 62.6 14 195.0
30	32.0	14.2	95	86.	38.6	6 55 141.6 63.0 18 196
	53.8	15.0	96	87	7 39 0 6 39 0	2 04 140.7 65.6 14 195.0 6 25 141.6 65.0 14 196.1 0 36 142.5 65.5 16 197 5 87 143.4 63.9 17 195
	34.7	15.4	97	88.		0 171141.4 63.9 141100
10	35 6	15.5	99	90	A 400-	2 20 145 3 64 7 10
14.1		16.3	3.530	100	40.7	7 60(146,1 65-1) 50
10		16.7	10		U) 61	1 16) 147 1 63 50
84	(B) (b)	17.4	10		10.	5 62 140 AL 60 91.
		17 4	0	114	3 80	3) PO 144 15 Mars
					10000	T 2003
						THE RESERVE OF THE PERSON NAMED IN COLUMN TWO IS NOT THE PERSON NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TWO IS NAMED IN COLUMN TW
						2

·r	27	Degrees.
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		. ~1	Deg	iccs.	
		Lat.	Dep.	Dist. Lat.	Dep.
	-	1.1.3	82.2	241 214.7	109.4
		9.501	82.6	42 215.6	
,		163.1	83.1	43 216.5	
•		163.9	83.5	44 217.4	
:		164.8		45 218.3	
		1165.7		46 219.2 47 220.1	
•		166.6		1 1 1	
- -		39 168.4			
		90 169.3		r 1	
<u> </u>		191 170.2	86.7	251 223.6	
		92 171.1	87.2		
- · ·		93 172.0	87.6	53 225 4	114 (
		94 172.9	88.1	54 226.3	115.
<u> </u>		95 173.7		55 227.2	115.8
	4.	96 174 . 6			
·	- 1	97 175.5	89.4	57 229.0 58 229.9	
The state of the s	í	98 176.4 99 177.3	$89.9 \\ 90.3$		
The state of the s	6	200 178.2		60 231.7	
LINE WILLIAM TO THE REAL PROPERTY AND THE PROPERTY AND THE PROPERTY AND THE PROPERTY AND	0	201 179.1	91.3	261 232.6	
FRE COLUMN		02 130.0	91.7	62 233.4	
1 8 2 dia 12	9	03 130 .9	92.2	63 234.3	
THE STATE OF THE S	, 1	04 181 .8	92.6	64 235.2	
-1.82 and -	5.8	05 132.7	93.1	65 236.1	
THE STATE OF THE PARTY OF THE P	6.3			66 237.6	
THE RESERVE OF THE PARTY OF THE	6.7		94.0	67 237.9	
THE RESERVE OF THE PARTY OF THE	67.2	08 185.3	94.4	68 238 .8	
I down the same of	67.6 68.1	09 186.2 10 187.1	94.9	69 239.7 70 240.6	
TARREST MANAGEMENT OF THE PARTY NAMED IN COLUMN			95.3	i	
STORY OF STREET	68.6	211 188.0	95.8	271 241.5 72 242.4	123.0
3	69.0 69.5	12 188.9 13 189.8	96.2 96.7	73 243.2	
7.2	69.9	14 190.7	97.2	74 244.1	
1.83.1	70.4		97.6	75 245 0	
139.0	70.8	16 192.5	98.1	76 245.9	
17,139.9	71.3	17 193.3	98.5	77 246.8	
3 140.8 30 141.7	71.7	18 194.2	99.0	78 247.7	
	72.2 72.6	19 195.1	99.4	79 248.6 80 249.5	
(0)142.6		20 196.0	99.9		
0 141 143.5	73.1	221 196.9		281 250.4 82 251.3	
63 144.3 63 145.2	73.5 74.0	22 197.8 23 198.7	100.8	83 252.9	
51146.1	74.5	24 199 .6			
7 65 147.0	74.9	25 200.5		85 253.9	
6 1 66 147.9	75.4			86 254.8	
57.0 17 148.8	75.8	27 202.3		87 255.7	
49.0 68 149.7	76.3	28 203.1		88 256 . 6	
19.5 69150.6	76.7	29 204.0		89 25 7 .5	
0 49.9 70 151.5	77.2	30 204.9			
50.4 171 152.4	77.6	231 205.8		291 259.3 92 260.2	
59.8 50 8 72 153.3 100.7 51.3 73 154.1	78.1 78.5	32 206.7 33 207.6			
100 7 51.3 73 154.1 14101.6 51.8 74 155.0	79.0	34 208 . 5		94 262.0	
15 102.5 52.2 75 155.9	79.4	35 209 . 4		95 262.8	133.9
16 103.4 52.7 76 156.8	79.9	36 210.3	107.1	96 263.7	
9 17 104.2 53.1 77 157.7	80.4	37 211.2		97 264.6	
5.3 18 105.1 53.6 78 158.6	80.8				
26.8 19 106.0 54.0 79 159.5	81.3				
27.2 20 106.9 54.5 80 160.4	81.7	40 213.8			
Lat. Dist. Dep. Lat. Dist. Dep.	Lat.	Dist. Dep.	rgf.	Dist. Dep.	
			-	[For 63 De	egrees

TABLE II.

Difference of Latitude and Departure for 26 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.
-1	00.9	00.4	61	54.8	26.7	121	108.8	53.0	181	162.7	79.3	241	216.6	105.6
2			62		27.2		109.7	53.5		163.6			217.5	
3		01.3	63				110.6			164.5			218.4	
4 5		01.8 02.2	64 65				111.5 112.3	54.4 54.8		165.4 166.3			219.3	107.4
6		02.6	66		28.9		113.2			160.3 167.2				107.8
7	06.3	03.1	67				114.1			168.1				108.3
8	07.2	03.5	68	61.1	29.8		115.0			169.0	82.4	48	222.9	108.7
9	08.1	03.9	69		30.2		115.9			169.9				109.2
10	09.0	04.4	70		30.7		116.8			170.8				109.6
11	09.9	04.8	71	63.8	31.1		117.7	57.4		171.7				110.0
12 13	10.8 11.7	05.3 05.7	72 73				118.6			172.6		52	226.5	110.5 110.9
14	12.6	06.1	74				119.5 120.4			173.5 174.4				111.3
15	13.5		75				121.3			175.3				111.8
16	14.4	07.0	76		33.3		122.2		96	176.2	85.9	56	230.1	112.2
17	15.3	07.5	77	69.2	33.8		123.1		97	177.1	86.4			112.7
18 19	16.2 17.1	07.9	78		34.2		124.0			178.0				113.1
20	18.0	08.3 08.8	79 80				124.9			178.9				113.5 114.0
21	18.9				!		125.8	61.4	I	179.8				
22	19.8	09.2	81 82	72.8 73.7	35.5 35.9		126.7 127.6	61.8 62.2		180.7 181.6				114.4 114.9
23	20.7	10.1	83				128.5			182.5		63	936 4	115.3
2.1		10.5	84				129.4			183.4		64	237.3	115.7
25	22.5	11.0	85	76.4	37.3		130.3			184.3	89.9	65	238.2	116.2
26	23.4	11.4	86	77.3	37.7		131.2			185.2				116.6
27 28	24.3	11.8 12.3	87	78.2	38.1		132.1			186 . 1				117.0
29	26.1	12.7	88 89	79.1 80.0	38.6 39.0		133:0 133.9			186.9 187.8				117.5 117.9
30	27.0		90		39.5		134.8	65.3 65.8		188.7		70	242.7	118.4
31	27.9	13.6	91	81.8	39.9		135.7			189.6				118.8
32	28.8	14.0	92	82.7	40.3		136.6			190.5				119.2
33			93	83.6	40.8	53	137.5	67.1		191.4		72	945 A	1110 7
34	30.6	14.9	94	84.5	41.2	54	138.4	67.5		192.3		74	246.3	120.1
35	31.5	15.3	95		41.6		139.3			193.2		75	247.2	120.6
36 37	32.4 33.3	15.8 16.2	96 97	86.3 87.2	42.1	57	140.2 141.1	68.4 68.8		194.1 195.0				121.0 121.4
38	34.2	16.7	98	88.1	43.0	58	142.0				95.6	78	249.9	121.9
39	35.1	17.1	99	89.0	43.4		142.9			196.8		79	250.8	122.3
40	36.0	17.5	100	89.9	43.8	60	143.8			197.7				122.7
41	36.9	18.0	101	90.8	44.3	161	144.7	70.6	221	198.6	96.9	281	25 2 .6	123.2
42	37.7	18.4	02		44.7		145.6	71.0		199.5	97.3			123.6
43 44	38.6	18.8	03		45.2		146.5			200.4				124.1
45	39.5 40.4	19.3 19.7	04 05		45.6 46.0		147.4 148.3			201.3 202.2				124.5 124.9
46	41.3	20.2	06	95.3	46.5		149.2	72.8		203.1				125.4
47	42.2	20.6	07		46.9		150.1	73.2		204.0				125.8
48	43.1	21.0	08	97.1	47.3	68	151.0	73.6		204.9	99.9	88	258.9	126.3
49	44.0	21.5	09		47.8		151.9				100.4	89	259.8	126.7
5 0	44.9	21.9	10		48.2		152.8	74.5			100.8			127.1
51	45.8	22.4		99.8	48.7		153.7	75.0			101.3			127.6
52 53	46.7 47.6	22.8 23.2		100.7 101.6	49.1 49.5		154.6		32	208.5	101.7	92	262.4	128.0
54	48.5	23.7		102.5	50.0	74	155.5 156.4	75.8 76.3	33	203.4 910 3	102.1 102.6	93	203.3	128.4 128.9
55	49.4	24.1		103.4	50.4		157.3		35	211.2	103.0	9.5	265.1	129.3
56	50.3	24.5	16	104.3	50.9	76	158.2	77.2	36	212.1	103.5	96	266.0	129.8
57	51.2	25.0		105.2	51.3	77	159.1	77.6	37	213.0	103.9	97	266.9	130.2
58	52.1	25.4		106.1	51.7		160.0	78.0			104.3			130.6
59 60	53.0 53.9	25.9 26.3		107.0 107.9			160.9 161.8		39	214.8	104.8	300	268.7	131.1
	Dep.	7 nt	13:20	Dan	72.0	D:-	101.0	70.9	Dist.	De-	105.2 Lat.	Dia:	Day	131.5
23(.)	₽ep.	LIUI.	יואוטי	Dep.	Lat.	1)15[.]	Dep.	Lal.	DIST.	Dep.	Lat.	DIST.	· Dep.	Lat.

[For 64 Degrees

Difference of Latitude and Departure for 27 Degrees.

								- opui						
l —	Lat.	·	-		Dep.	Dist.		Dep.	·					Dep.
	1 00.9 2 01.8		61				107.8			161.3 162.2		241	214.7	109.4
	02.7						109.6			163.1	83.1			110.3
	03.6						110.5			163.9	83.5			110.8
	04.5						111.4			164.8				111.2
			66				112.3 113.2			165.7 166.6		40	219.2	1112.1
	07.1						114.0			167.5				112.6
9	08.0	04.1	69	61.5	31.3	29	114.9	58.6	89	168.4	85.8	49		113.0
10			70			·	115.8			169.3		:	1	113.5
11			71				116.7			170.2				114.0
12		05.4 05.9	72				117.6 118.5			171.1 172.0	87.2 87.6			114.4 114.9
14			74				119.4			172.9				115.3
15	13.4		75		34.0		120.3			173.7	88.5			115.8
16		07.3 07.7	76 77		34.5 35.0		121.2 122.1	61.7 62.2		174.6		41		116.2
17			78		35.4		123.0			175.5 176.4				117.1
19	16.9	08.6	79	70.4	35.9	39	123.8	63.1		177.3	90.3	59	230.8	117.6
20	17.8	09.1	80		36.3		124.7	63.6		178.2	90.8	60	231.7	118.0
21	18.7	09.5	81	72.2	36.8		125.6	64.0	201	179.1	91.3	261		118.5
22		10.0 10.4	82 83		37.2 37.7		126.5 127.4	64.5 64.9		180.0 180.9				118.9
23 24		10.9	84		38.1		128.3			181.8	92.2 92.6			119.4 119.9
25	22.3	11.3	85	75.7	38.6	45	129.2	65.8		182.7	93.1			120.3
26		11.8	86		39.0		130.1	66.3		133.5	93 .5			120.8
27	24.1 24.9	12.3 12.7	87 88	77.5 78.4	39.5 40.0		131.0 131.9	66.7 67.2		184.4 185.3	94.0 94.4			121.2 121.7
29		13.2	89	79.3	40.4		132.8			186.2	94.9		239.7	
30		13.6	90	80.2	40.9	50	133.7	68.1		187.1	95.3			122.6
31	27.6	14.1	91	81.1	41.3		134.5	68.6		188.0	95.8			123.0
32	28.5	14.5	92	82.0	41.8		135.4	69.0		188.9	96.2			123.5
33	29.4 30.3	15.0 15.4	93 94	82.9 83.8	42.2 42.7		136.3 137.2	69.5 69.9		189.8 190.7	96.7 97.2			123.9 124.4
35	31.2	15.9	95	84.6	43.1		138.1	70.4		191.6	97.6			124.8
36	32.1	16.3	96	85.5	43.6	56	139.0	70.8		192.5	98.1	76	245.9	125.3
37	33.0 33.9	16.8 17.3	97	86.4 87.3	44.0 44.5		139.9 140.8	71.3 71.7		193.3	98.5		246.8 247.7	
38		17.7	98 99	88.2	44.9		141.7	72.2		194.2 195.1	99.0 99.4		248.6	
40	35.6	18.2	100	89.1	45.4	60	142.6	72.6		196.0	99.9		249.5	
41	36.5	18.6	101	90.0	45.9		143.5	73.1		196.9		281	250.4	127.6
42	37.4	19.1	02	90.9	46.3		144.3	73.5		197.8			251.3	
43	38.3 39.2	19.5 20.0	03 04	91.8 92.7	46.8 47.2		145.2 146.1	74.0 74.5		198.7 199.6			252.2 253.0	
45	40.1	20.4	05	93.6	47.7		147.0	74.9		200.5			253.9	
46	41.0	20.9	06	94.4	48.1		147.9	75.4		201.4			254.8	
47	41.9	21.3 21.8	07	95.3	48.6 49.0		148.8 149.7	75.8 76.3		202.3			255.7	
48 49	42.8 43.7	22.2	08 09	96.2 97.1	49.5		150.6	76.7		203.1 204.0			256.6 25 7 .5	
50	44.6	22.7	10	98.0	49.9		151.5	77.2		204.9			258.4	
51	45.4	23.2	111	98.9	50.4	171	152.4	77.6		205.8			259.3	
52	46.3	23.6	12		50 .8		153.3	78.1	32	206.7	105.3	92	260.2	132.6
53	47.2	24.1		100.7	51.3		154.1	78.5		207.6			261.1	
54 55	48.1 49.0	24.5 25.0		101.6 102.5	51.8 52.2		155.0 15 5 .9	79.0 79.4		208. <i>5</i> 209.4			2 62 .0 2 62 .8	
56	49.9	25.4	16	103.4	52.7	76	156.8	79.9	36	210.3	107.1	96	263.7	134.4
57	50.8	25.9	17	104.2	53.1	77	157.7	80.4	37	211.2	107.6		264.6	
58 59	51.7 52.6	26.3 26.8		105.1 106.0	53.6 54.0		158.6 159.5	80.8 81.3		212.1 213.0			265.5 266.4	
60	53.5	27.2		106.0	54.5		160.4	81.7		213.8			267.3	
Dist.		Lat.				Dist.					Lat.			
	P						P						63 De	
														J

Difference of Latitude and Departure for 28 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.	Dist. Lat.	Dep.	Dist. Lat. I	
1	00.9	00.5	61	53.9	28.6	121 106.8	56.8	181 159.8	85.0	241 212.8 11	
.5	01.8		62	54.7	29.1	22 107.7	57 .3	82 160.7	85.4	42 213.7 11	
3	02.6	01.4	63	55.6	29.6	23 108.6	57.7	83 161 .6		43 214 6 11	
4				56.5	30.0	24 109 .5	58.2	84 162.5	86.4	44 215.411	
5	01.4	02.3		57.4	30.5	25 110.4	58.7	85 163.3	86.9	45 216 .3 1	
• 07	05.3		66	58.3	31.0	26 111.3	59.2	86 164.2	87.3 87.8	46 217.211	
7	06.2	03.3	67	59.2	31.5	27 112.1 28 113.0	59.6 60.1	87 165.1 88 166.0	88.3	48 219.01	
8	07.1	03.8	63	60.0 60.9	31.9 32.4	29 113.9	60.6	89 166.9	88.7	49 219.91	
10	07.9 08.8	01.2 01.7	69 70		32.9	30 114.8	61.0	90 167.8	89.2	50 220.71	
1			I					191 168.6	89.7	251 221 .6 1	
111	09.7	05.2	71	62.7		131 115.7 32 116.5	61.5 62.0	92 169.6	90.1	52 222.51	18.3
12	10.6	05.6	72	63.6 64.5	33.8 34.3	33 117.4	62.4	93 170.4	90.6	53 223.4 11	
13	11.5 12.4		73 74	65.3	34.7	34 118.3	62.9	94 171 .3	91.1	54 224.3 11	
15	13.2	07.0		66.2	35.2	35 119.2	63.4	95 172.2	91.5	55 225.2 11	
16	14.1	07.5		67.1	35.7	36 120.1	63.8	96 173.1	92.0	56 226.0 12	0.2
17	15.0			68.0		37 121.0	64.3	97 173.9	92.5	57 226 . 9 12	20.7
18	15.9	08.5	78	68.9	36.6	38 121 .8	64.8	98 174.8	93.0	58 227.8 19	21.1
19	16.8	08.9		69.8	37.1	39 122.7	65.3	99 175.7	93.4	59 228.7 19	
20	17.7	09.4		70.6	37.6	40 123.6	65.7	200 176.6	93.9	60 229 . 6 1	
21	13.5	09.9	81	71.5	38.0	141 121.5	66.2	201 177.5	94.4	261 230.4 19	22.5
22	19.4	10.3	82	72.1	38.5	42 125.4	66.7	02 178.4	94.8	62 231 . 3 15	23.0
23	20.3	10.8	83		39.0	43 126.3	67.1	03 179.2	95.3	63 232.2 19	23.5
24	21.2	11.3			39.4	44 127.1	67.6	04 180.1	95.8	64 233.1 12	23.9
25	22.1	11 7	85		39.9		68.1	05 181.0	96.2	65 234.0 12	
26	23.0	12.2	86	75.9	40.4	46 128.9	63.5	06 181.9	96.7	66 234.9 1	
27	23.8	12.7	87	76.8	40.8	47 129.8	69.0	07 182.8	97.2	67 235.7 19	
28	24.7	13.1	88	77.7	41.3	48 130.7	69.5	08 183.7	97.7	68 236 . 6 1	
29	25.6	13.6	89	78.6		49 131.6	70.0	09 184.5		69 237.5 1	
30	26.5	14.1	90	79.5	42.3	50 132.4	70.4	10 185.4	98.6	70 238 . 4 19	
31	27.4	14.6	91	80.3	42.7	151 133.3	70.9	211 186.3		271 239.3 15	27.2
32	28.3	15.0	92	81.2	43.2	52 134.2	71.4	12 187.2	99.5	72 240 .2 1	
33	29.1	15.5	93	82.1	43.7	53 135.1	71.8	13 188.1		73 241 .0 15	
34	30.0	16.0	94	83.0	44.1	54 136.0	72.3	14 189.0		74 241 .9 15	0.83
35	30.9	16.4	.95	83.9	41.6	55 136.9	72.8	15 189.8		75 242 .8 15	
36	31.8	16.9	96	84.8	45.1	56 137.7	73.2	16 190.7 17 191.6		76 243.7 15	
37	32.7	17.4	97	85.6	45.5	57 138.6 58 139.5	73.7 74.2	18 192.5		78 245 .51	
33	33.6	17.8	98	86.5	46.0 46.5	59 140.4	74.6	19 193.4		79 246 .3 13	
39 40	34.4 35.3	13.3 18.8	99 1 0 0	87.4 88.3	46.9	60 141 3	75.1	20 194.2		80 247.21	
1			i —			I		221 195.1	;	281 248.1 1:	
41	36.2	19.2	101	89.9	47.4 47.9	161 142.2 62 143.0	75.6 76.1	22 196.0		82 249 .0 1	
42	37.1	19.7	02	90.1		63 143.9	76.5	23 196.9		83 249 .9 13	
43	38.0	20.2 20.7	03	90.9 91.8		64 144 .8	77.0	24 197.8		84 250.813	
44 45	38.8 39.7	21.1	04 05	92.7	49.3	65 145.7	77.5	25 198.7		85 251 .6 13	
46	40.6	21.6	06	93.6	49.8	66 146.6	77.9	26 199.5		86 252.5 13	34.3
47	41.5	22.1	07	94.5	50.2	67 147.5	78.4	27 200.4		87 253.4 15	
48	42.4	22.5	08	95.4	50.7	68 148.3	78.9	28 201.3	107.0	88 254.3 13	35.2
49	43.3	23.0	09	96.2	51.2	69 149.2	79.3	29 202.2	107.5	89 255.2 13	
50	44.1	23.5	10	97.1	51.6	70 150.1	79.8	30 203.1	108.0	90 256 . 1 13	36.1
51	45.0	23.9	111	98.0	52.1	171 151.0	80.3	231 204.0	108.4	291 256.913	36.6
52	45.9	24.4	12	98.9	52.6	72 151.9	80.7	32 204.8		92 257.8 13	
53	46.8	24.9	13	99.8	53.1	73 152.7	81.2	33 205.7	109.4	93 258.7 15	
54	47.7	25.4		100.7	53.5	74 153.6	81.7	34 206.6	109.9	94 259 .6 13	
55	48.6	25.8		101.5	54.0	75 154.5	82.2	35 207.5	110.3	95 260 . 5 13	
56	49.4	26.3	16	102.4	54.5	76 155.4	82.6	36 208.4		96 261 .4 13	
57	50.3	26.8		103.3	54.9	77 156.3	83.1	37 209 . 3		97 262.2 13	9.4
58	51.2	27.2		104.2	55.4	78 157.2	83.6	38 210.1		98 263 . 1 13	
59	52.1	27.7		105.1	55.9	79 158.0	84.0	39 211.0		99 264.0 14	
60		28.2		106.0	56.3	80 158.9	84.5	40 211.9		300 264.9 14	
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep.	Lat.	Dist. Dep.	Lat.	Dist. Dep. I	
										(For 62 Dec	reet

[For 62 Degrees.

Difference of Latitude and Departure for 29 Degrees.

1	Dist	. Lat.	1 Dan	Dia	Tat	(Dan	(INC.	I T at	There	·YN:-4	. Y -4	i Dan	· 105	-4 .	T
1	-		-	-1	-1		Dist.		Dep.			Dep.	.	_	Dep.
١	2	1						105.8 106.7	58.7 59.1		158. 159.				117.3
1	3							107.6			160				117.8
1	4					31.0	24	108.5		84	160.	89.2	44 21	3.4	118.3
1	5							109.3	60.6		161				118.8
1	6							110.2 111.1	61.1		162.		45 21	5 · Z	119. 3 119. 7
	8							112.0	61.6 62.1		163 . 164 .				120.2
i	9							112.8			165				120.7
١	10	08.1	7 04.8	70	61.2	33.9	30	113.7	63.0	90	166.5	92.1	50 21	8.7	121.2
1	11	09.6	05.3	71	62.1	34.4	131	114.6	63.5	191	167.	92.6	251 21	9.5	121.7
1	12					34.9		115.4	64.0		167				122.2
	13					35.4		116.3	64.5		168.				122.7
1	14 15					35.9 36.4		117.2 118.1	65.0 65.4		169.1 170.6				123.1 123.6
1	16							118.9	65.9		171.4		11 - 1		124.1
	17		08.2	77	67.3	37.3		119.8	66.4		172.5			4.8	124.6
П	18					37.8		120.7	66.9		173.2				125.1
П	19 2 0			79 80	69.1 70.0	38.3 38.8		121.6 122.4	67.4 67.9		174.0 174.9				125.6 126.1
	21	18.4	_	11							175.8		l		
	21 22	19.2		81 82	70.8 71.7	39.3 39.8		123.3 124.2	68.4 68.8		176.7		261 22 62 22	0.3	120.5 127 A
	23	20.1			72.6	40.2		125.1	69.3		177.5		63 23		
П	24	21.0		84	73.5	40.7		125.9	69.8		178.4		64 23		
H	25	21.9		85	74.3	41.2		126.8	70.3		179.5		65 23		
11	26	22.7		86	75.2	41.7		127.7	70.8		180.5		66 23		
11	27 28	23.6 24.5		87 88	76.1 77.0	42.7		128.6 129.4	71.3 71.8			100.4 100.8	67 23 68 23		
11	29	25.4	14.1	89	77.8	43.1		130.3	72.2			101.3	69 23		
П	30	26.2	14.5	90	78.7	43.6	50	131.2	72.7	10	183.7	101.8	70 23	6.1	130.9
П	31	27 1	15.0	91	79.6	44.1	151	132.1	73.2			102.3	271 23		
П	32	28.0	15.5	92	80.5	44.6		132.9	73.7			102.8	72 23		
H	33 34	28.9 29.7	16.0 16.5	93 94	81.3 82.2	45.1 45.6		13 3 .8 134.7	74.2 74.7			103.3 103.7	73 23 74 23		
П	36	30.6	17.0	95	83.1	46.1		135.6	75.1			104.2	75 24		
П	36	31.5	17.5	96	84.0	46.5		36.4	75.6			104.7	76 24	1 . 4 1	33.8
П	37	32.4	17.9	97	84.8	47.0		137.3	76.1			105.2	77 249		
П	38 39	33.2 34.1	18.4 18.9	98 99	85.7	47.5 48.0		138.2 139.1	76.6			105.7 106.2	78 243 79 24		
П	40	35.0	19.4	100	87.5	48.5		39.9	77.1			106.7	80 244		
lŀ	41	35.9	19.9	101	88.3	49.0		140.8	78.1			107.1	281 245	/_	
П	42	36.7	20.4	02	89.2	49.5		41.7	78.5	22	194.2	107.6	82 240		
П	43	37.6	20.8	03	90.1	49.9		142.6	79.0	23	195.0	108.1	83 247		
П	44	38.5	21.3	04	91.0	50.4		43.4	79.5			108.6	84 248		
П	45 46	39.4 40.2	21.8 22.3	05 06	91.8 92.7	50.9 51.4		44.3	80.0 80.5			109.1 109.6	85 249 86 250		
	47	41.1	22.8	07	93.6	51.9		46.1	81.0			110.1	87 251		
11	48	42.0	23.3	08	94.5	52.4		46.9	81.4	28	199.4	110.5	88 251		
11	49	42.9	23.8	09	95.3	52.8		47.8	81.9			111.0	89 255		
_	50	43.7	24.2	10	96.2	53.3		48.7	82.4			111.5	90 253		
	51	41.6	24.7	111	97.1	53.8		49.6	82.9			112.0	291 254		
	52 53	45.5	25.2 25.7	12 13	98.0 98.8	54.3 54.8		50.4 51.3	83.4 83.9			112.5 113.0	92 255 93 256		
П	54	47.2	26.2	14	99.7	55.3		52.2	84.4			113.4	94 257	1.1 1	42.5.
		48.1	26.7		100.6	55.8		53.1	84.8	35	205.5	113.9	95 258	10.1	43.0
	56	49.0	27.1		101.5	56.2		53.9	85.3			114.4	96 258		
1	57 58	49.9	27.6 28.1		102.3	56.7 57.2		54.8 55.7	85.8 86.3			114.9 115.4	97 259 98 260		
		50.7 51.6	28.1		103.2	57.7		56.6	86.8			115.9	99 261		
		52.5	29.1		05.0	58.2		57.4	87.3			116.4	300 262		
ī	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. 1	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. De	p. 1	Lat.
_													For 61	Deg	rees.

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TABLE II.

Difference of Latitude and Departure for 30 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	-
. 1	00.9	00.5	61	52.8	30.5		104.8	60.5		156.8	90.5			120.5	П
2	01.7		62	53.7	31.0		105.7	61.0		157.6	91.0			121.0	П
3	02.6 03.5		63 64	54.6 55.4	31.5 32.0		106.5 107.4	61.5 62.0		158.5 159.3	91.5 92.0			121.5	
5	04.3		65	56.3	32.5		108.3			160.2	92.5			122.5	П
6	05.2	03.0	66	57.2	33.0		109.1	63.0	86	161.1	93.0			123.0	П
7	06.1	03.5	67	58.0	33.5		110.0	63.5		161.9	93.5			123.5	11
8	06.9 07.8	04.0 04.5	68 69	58.9 59.8	34.0 34.5		110.9 111.7	64.0 64.5		162.8 163.7	94.0 94.5			124.0 124.5	П
10	08.7	05.0	70	60.6	35.0		112.6	65.0		164.5	95.0			125.0	Н
11	09.5	05.5	71	61.5	35.5		113.4	65.5		165.4	95.5	·		125.5	1
12	10.4	06.0	72	62.4	36.0	32	114.3	66.0	92	166.3	96.0	52	218.2	126.0	П
13	11.3	06.5	73	63.2	36.5		115.2	66.5		167.1	96.5			126.5	
14 15	12.1 13.0	07.0 07.5	74 75	64.1 65.0	37.0 37.5		116.0 116.9	67.0 67.5		168.0 168.9	97.0 97.5			127.0 127.5	
16	13.9	08.0	76	65.8	38.0	36		68.0		169.7	98.0			128.0	b
17	14.7	08.5	77	66.7	38.5	37	118.6	68.5		170.6				128.5	
18	15.6		78	67.5			119.5	69.0		171.5				129.0	П
19 20	16.5 17.3	09.5 10.0	79 80	68.4 69.3	39.5 40.0		120.4 121.2	69.5 70.0		172.3 173.2				129.5 130.0	1
21	18.2	10.5	81	70.1	40.5	141	122.1	70.5		174.1		l\		130.5	1
22	19.1	11.0	82	71.0			123.0	71.0			101.0			131.0	
23	19.9		83	71.9	41.5	43	123.8	71.5	03		101.5			131.5	1
24	20.8	12.0	84	72.7	42.0		124.7	72.0			102.0			132.0	ı
2.5 26	21.7 22.5		85 86	73.6 74.5	42.5 43.0		125.6 126.4			177.5	102.5 103.0			132.5 133.0	ı
27	23.4		87	75.3			127.3				103.5			133.5	1
23	24.2	14.0	88	76.2			128.2				104.0			134.0	ı
29	25.1		89	77.1	44.5		129.0			181.0		69	233.0	134.5	П
30	26.0	15.0	90	77.9	45.0	-	129.9		U		105.0			135.0	1
31	26.8 27.7	15.5	91	78.8	45.5		130.8	75.5			105.5			135.5	Ł
32 33	28.6	16.0 16.5	92 93	79.7 80.5	46.0 46.5		131.6 1 32 .5				106.0 106.5			136.0 136.5	ł
34	29.4		94	81.4			133.4				107.0			137.0	ı
35	30.3		95	82.3	47.5	55	134.2	77.5	15	186.2	107.5			137.5	ı
36 37	31.2		96	83.1 84.0	48.0		135.1	78.0			108.0			138.0	1
38	32.0 32.9		98	84.9	48.5 49.0		136.0 136.8				108.5 109.0			138.5 139.0	1
39	33.8		99	85.7	49.5		137.7	79.5			109.5			139.5	i
40	34.6	20.0	100	86.6	50.0	60	138.6	80.0			110.0	80	242.	140.0	
41	35.5		101	87.5	50.5		139.4	80.5		191.4				140.5	7
42 43	36.4		02	88.3	51.0		140.3				111.0			141.0	١
43	37.2 38.1	21.5 22.0	03 04		51.5 52.0		141.2 142.0			193.1 194.0				141.5 142.0	1
45	39.0	22.5	05	90.9	52.5		142.9				112.5			142.5	1
46	39.8	23.0	06	91.8	53.0	66	143.8	83.0	26	195.7	113.0	86	247.7	143.0	
47 48	40.7 41.6	23.5 24.0	07 08	92.7 93.5	53.5		144.6				113.5			143.5	
49	42.4		09	94.4	54.0 54.5		145.5 146.4				114.0 114.5			144.0 144.5	
50	43.3	25.0	10	95.3	55.0		147.2	85.0			115.0			145.0	
51	44.2	25.5	111	96.1	55.5	171	148.1	85.5	n		115.5	!!		143.5	_,
52	45.0		12	97.0	56.0		149.0	86.0	32	200.9	116.0	92	252.9	146.0	1
53 54	45.9 46.8	26.5	13	97.9	56.5		149.8				116.5			146.5	
55	47.6		14 15	98.7 99.6	57.0 57.5		150.7 151.6	87.0 87.5			117.0 117.5			5 147.0 5 147.5	
56	48.5			100.5	58.0		152.4				118.0			148.0	
57	49.4	28.5	17	101.3	58.5	77	153.3	88.5	37	205. 2	118.5	97	257 .9	148.5	,
58 59	50.2 51.1			102.2	59.0		154.2				119.0			1149.0	
60	52.0	29.5 30.0		103.1 103.9	60.0		155.0 155.9				119.5 120.0			9 149.5 3 150.0	
Dist.					Lat.				l!			11	·	Lat.	-¦
				···						<u> </u>	- 2746	<u> </u>		egrees	<u>.</u>

TABLE II.

Difference of Latitude and Departure for 31 Degrees.

1	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat. Dep.
ŀ	1	00.9	00.5	61	52.3	31.4	121	103.7	62.3	181	155.1	93.2	241 206 . 6 124 . 1
П	2	01.7	01.0		53.1	31.9		104.6	62 .8		156.0		
Н	3	02.6		63	54.0	32.4		105.4	63.3		156.9		43 208.3 125.2
Н	4	03.4			54.9	33.0		106.3	63.9		157.7 158.6	94.8 95.3	45 210.0 126.2
П	5	04.3	02.6 03.1		55.7	33.5 34.0		107.1 108.0	64.4 64.9		159.4		
Н	6	05.1 06.0		66 67	56.6 57.4	34.5		108.9	65.4		160.3		47 211.7 127.2
Н	8	06.9	04.1		58.3	35.0		109.7	65.9		161.1	96.8	48 212 .6 127 .7
Н	9	07.7		69	59.1	35.5		110.6	66.4		162.0	1	49 213.4 128.2
П	10	08.6		70	60.0		30	111.4	67.0	90	162.9	97.9	50 214.3 128.8
i t	11	09.4	05.7	71	60.9	36.6	131	112.3	67.5	191	163.7	98.4	251 215.1 129.3
Н	12	10.3	06.2	72	61.7	37.1		113.1	68.0		164.6		52216.0129.8
Н	13	11.1	06.7	73	62.6	37.6	33	114.0	68.5	93	165.4	99.4	53 216.9 130.3
Н	14	12.0		74	63.4	38.1	34	114.9	69.0		166.3		54 217.7 130.8
П	15	12.9	07.7	75	64.3	38.6		115.7	69.5			100.4	55218.6 131.3
П	16	13.7	08.2	76	65.1	39.1		116.6	70.0			100.9	56 219.4 131.8 57 220.3 132.4
Н	17	14.6			66.0	39.7		117.4	70.6	97	168.9	101.5	58 221 . 1 132 . 9
	18	15.4			66.9	40.2		118.3	71.1 71.6			102.0 102.5	59 222.0 133.4
	19	16.3			67.7 68.6	40.7		119.1 120.0	72.1			103.0	60 222 . 9 133 . 9
١L	20	17.1	10.3	1									261 223.7 134.4
	21	18.0	10.8	81	69.4	41.7		120.9 121.7	72.6 73.1			103.5 104.0	62 224.6 134.9
1	22	18.9		82	70.3 71.1	42.7		122.6	73.7			104.6	63 225.4 135.5
	23 24	19.7 20.6			72.0	43.5		123.4	74.2		174:9		64 226 . 3 136 . 0
1	25	21.4	12.9	85	72.9	43.8		124.3	74.7			105.6	65 227.1 136.5
Н	26	22.3		86	73.7	44.3		125.1	75.2		176.6		66 228.0 137.0
Н	27	23.1			74.6	44.8		126.0	75.7		177.4		67 228.9 137.5
П	28	24.0			75.4	45.3		126.9	76.2		178.3		68 229.7 138.0
H	29	24.9		89	76.3	45.8		127.7	76.7		179.1		69 230 . 6 138 . 5
П	30	25.7	15.5	90	77.1	46.4	50	128.6	77 .3	10	180.0	108.2	70 231.4 139.1
ŀ	31	26.6	16.0	91	78.0	46.9		129.4	77.8		180.9		271 232.3 139.6
н	32	27.4	16.5	92	78.9	47.4		130.3	78.3			109.2	72 233.1 140.1
П	33	28.3	17.0	93	79.7	47.9		131.1	78.8		182.6		73 234.0 140.6
Н	34	29.1	17.5	94	80.6	48.4		132.0	79.3			110.2	74 234 . 9 141 . 1 75 235 . 7 141 . 6
11	35	30.0		95	81.4	48.9		132.9	79.8		184.3	111.2	76 236 . 6 142 . 2
Н	36	30.9		96	82.3	49.4		13 3.7 134.6	80.3 80.9		186.0		77 237 .4 142 .7
П	37 38	31.7 32.6		97 98	83.1 84.0	50.0 50.5		135.4	81.4			112.3	78 238 . 3 143 . 2
Н	39	33.4		99	84.9	51.0		136.3	81.9		187.7		79 239 . 1 143 . 7
Н	40	34.3	20.6	100	85.7	51.5		137.1	82.4		188.6		80 240.0 144.2
-	41	85.1	21.1	101	86.6	52.0	161	138.0	82.9	221	189.4	113.8	281 240.9 144.7
Н	42	36.0		02	87.4	52.5		138.9	83.4		190.3		82 241 .7 145 .2
H	43	36.9		03	88.3			139.7	84.0	23	191.1	114.9	83 242 .6 145 .8
П	44	37.7	22.7	04	89.1	53.6	64	140.6	84.5			115.4	84 243.4 146.3
П	45	38.6			90.0	54.1		141.4	85.0			115.9	85 244 . 3 146 . 8
П	46	39.4	23.7	96	90.9	54.6		142.3	85.5	1		116.4	86 245.1 147.3
Н	47	40.3		07	91.7	55.1		143.1	86.0			116.9	87 246 . 0 147 . 8 88 246 . 9 148 . 3
П	48	41.1		08	92.6	55.6		144.0	86.5		196.3	117.4	89 247 .7 148 .8
1	49	42.0		09	93.4	56.1 56.7		144.9 145.7	87.0 87.6		197.1		90 248 .6 149 .4
1	50	42.9		10	94.3					1	198.0		291 249.4 149.9
	51	43.7		111	95.1	57.2		146.6	88.1 88.6			119.0	92 250 .3 150 .4
H	52	44.6		12	96.0 96.9	57.7 58.2		147.4 148.3				120.0	93 251.2 150.9
1	53	45.4		13 14		58.7		149.1	89.6			120.5	94 252.0 151.4
Н	54 55	46.3				59.2		150.0	90.1		201.4		95 252.9 151.9
П	56	48.0				59.7		150.9	90.6	36	202.3	121.5	96 253.7 152.5
Н	57	48.9			100.3	60.3		151.7	91.2	37	203.1	122.1	97 254.6 153.0
Н	58	49.7			101.1	60.8	78	152.6	91.7			122.6	98 255 . 4 153 . 5
П	59	50.6			102.0			153.4	92.2		204.9		99 256 .3 154 .0
		4	30.9	li en	102.9	61.8	80	154.3	92.7	40	Z U5.7	123.6	300 257 . 1 154 . 5
П	60	61.4	(30.3	20	102.3								
j	60 Dist.				Dep.			Dep.			Dep.	Lat.	Dist. Dep. Lat.

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Difference of Latitude and Departure for 32 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	-
1	00.8	00.5	61	51.7	32.3	121	102.6	64.1	181	153.5	95.9		204.4	127.7	11
2	01.7	01.1	62	52.6	32.9	22	103.5	64.7	82	154.3				128.2	П
3	02.5	01.6	63	53.4	234		104.3	65.2		155.2	97.0			128.8	П
4	03.4	02.1	64	54.3	33.9		105.2	65.7		156.0				129.3	П
5	04.2	02.6		55.1 56.0	34.4 35.0		106.0 106.9	66.2		156.9 157.7				129.8 130.4	П
6	05.1 05.9	03.2 03.7	66 67	56.8	35.5		107.7	66.8 67.3		158.6		47	209.5	130.9	П
8	06.8		68	57.7	36.0		108.6	67.8		159.4		48	210.3	131.4	П
9	07.6		69	58.5	36.6		109.4	68.4		160.3		49	211.2	131.9	П
10	08.5	05.3	70	59.4	37.1	30	110.2	68.9	90	161.1	100.7	50	212.0	132.5	П
11	09.3	05.8	71	60.2	37.6		111.1	69.4		162.0				133.0	П
12	10.2	06.4	72	61.1	38.2		111.9	69.9		162.8				133.5	П
13	11.0		73	61.9	38.7		112.8	70.5		163.7				134.1	П
14	11.9 12.7	07.4 07.9	74 75	62.8 63.6	39.2 39.7		113.6 114.5	71.0 71.5		164.5 165.4			216.3	134.6	П
15 16	13.6		76	64.5	40.3		115.3	72.1		166.2				135.7	
17	14.4		77	65.3	40.8		116.2	72.6		167.1				136.2	1
18	15.3	09.5	78	66.1	41.3	38	117.0	73.1		167.9		58		136.7	Н
19	16.1	10.1	79	67.0	41.9		117.9	73.7		168.8				137.2	П
20	17.0	10.6	80	67.8	42.4		118.7	74.2		169.6	i			137.8	
21	17.8	11.1	81	68.7	42.9		119.6	74.7		170.5				138.3	
22	18.7	11.7	82	69.5	43.5		120.4		02	171.3	107.0	62		138.8	
23 24	19.5 20.4	12.2 12.7	83 84	70.4 71.2	44.0 44.5		121.3 122.1	75.8 76.3	03	172.2 173.0	107.6	64		139.4 139.9	
25	21.2	13.2	85	72.1	45.0		123.0			173.8				140.4	Н
26	22.0	13.8	86	72.9	45.6		123.8	77.4		174.7				141.0	
27	22.9	14.3	87	73.8	46.1		124.7	77.9		175.5		67	226.4	141.5	1
28	23.7	14.8	88	74.6	46.6		125.5	78.4		176.4				142.0	П
29	24.6	15.4	89	75.5	47.2		126.4	79.0		177.2				142.5	١
30	25.4	15.9	90	76.3	47.7		127.2	79.5		178.1	·			143.1	-
31	26.3	16.4	91	77.2	48.2		128.1	80.0		178.9				143.6	1
32 33	27.1 28.0	17.0 17.5	92 93	78.0 78.9	48.8 49.3		128.9 129.8	80.5 81.1		179.8 180.6		72	200.	144.1	
34	28.8	18.0	94	79.7	49.8		130.6			181.5				145.2	
35	29.7	18.5	95		50.3		131.4	82.1		182.3				145.7	1
36	30.5	19.1	96	81.4	50.9	56	132.3		16	183.2	114.5			146.3	1
37	31.4	19.6	97	82.3	51.4		133.1	83.2		184.0				146.8	1
38	32.2 33.1	20.1	98	83.1	51.9 52.5		134.0			184.9 185.7				147.3	1
39 40	33.9	20.7 21.2	99 100	84.0 84.8	53.0		134.8 135.7	84.3 84.8		186.6				147.8 148.4	İ
41	34.8	21.7	101	85.7	53.5		136.5			187.4					-
42	35.6	22.3	02	86.5	54.1		137.4	85.3 85.8		188.3				148.9 149.4	
43	36.5	22.8	03	87.3			138.2			189.1				150.0	1
44	37.3	23.3	04	88.2	55.1		139.1	86.9		190.0				150.5	1.
45	38.2	23.8	05		55.6		139.9	87.4		190.8		85	241.	151.0	1
46	39.0	24.4	06		56.2		140.8			191.7				151.6	1
47 48	39.9 40.7	24.9 25.4	07	90.7 91.6	56.7 57.2		141.6 142.5			192.5 193.4				152.1 152.6	
49	41.6	26.0	09	92.4	57.8		143.3			193.4				153.1	
50	42.4	26.5	10	93.3	58.3		144.2			195.1				153.7	i
51	43.3	27.0	111	94.1	58.8	I	145.0			195.9				154.2	-
52	44.1	27.6	12	95.0	59.4		145.9	91.1		196.7				154.7	1
53	44.9	28.1	13	95.8	59.9		146.7	91.7	33	197.6	123.5	93		155.3	1
54	45.8	28.6	14	96.7	60.4		147.6	92.2		198.4				155.8	1
55 EC	46.6	29.1	15	97.5	60.9		148.4	92.7		199.3				156.3	1
56 57	47.5 48.3	29.7 30.2	16 17	98.4 99.2	61.5 62.0		149.3 150.1	93.3 93.8		200.1				156.9	-
58	49.2	30.Z 30.7		100.1	62.5		151.0			201.0 201.8				157.4 157.9	1
59	50.0	31.3		100.9	63.1		151.8	94.9		202.7			253.	158.4	1
60	50.9	31.8		101.8	63.6		152.6	95.4		203.5		300	254.	159.0	ĺ
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.			Lat.	1
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[For 58 Degrees.

TABLE II.

Difference of Latitude and Departure for 33 Degrees.

								- opur							
Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.		Lat.	Dep.	Dist.	Lat	$\cdot 1^{10}$)ep.
1 1		00.5	61	51.2	33.2		101.5	65.9		151.8			202.	- 1	
2			62		33.8		102.3	66.4		152.6			203.		
3			63 64	52.8 53.7	34.3 34.9		103.2 104.0	67.0 67.5	83	153.5 154.3			203. 204.		
3		02.7	65	54.5	35.4		104.8	68.1		155.2			205.		
6	05.0	03.3		55.4	35.9		105.7	68.6			101.3		206		
7	05.9	03.8	67	56.2	36.5	27	106.5	69.2			101.8	47	207.	2 13	34.
8		04.4	68	57.0	37.0		107.3	69.7	88	157.7	102.4		208		
9		04.9	69	57.9	37.6		108.2	70.3		158.5			208		
10		05.4	70	58.7	38.1	l'	109.0	70.8		159.3		_	209.		
1 11		06.0	71	59.5	38.7		109.9	71.3		160.2			210.		
12 13		06.5 07.1	72 73	60.4	39.2 39.8		110.7	71.9 72.4		161.0 161.9			211. 212.		
14		07.6	74	61.2 62.1	40.3		111.5 112.4	73.0		162.7			213.		
15		08.2	75	62.9	40.8		113.2	73.5		163.5			213.		
16	13.4	08.7	76	63.7	41.4		114.1	74.1		164.4			214.		
17		09.3	77	64.6	41.9		114.9	74.6			107.3		215.		
18		09.8	78	65.4	42.5		115.7	75.2		166.1			216		
19 20		10.3 10.9	79 80	66.3	43.0		116.6	75.7		166.9			217. 218.		
21	17.6	11.4		67.1	43.6		117.4	76.2		167.7		·			
21		12.0	81 82	67.9 68.8	44.1 44.7		118.3 119.1	76.8 77.3		168.6	109.5		218. 219.		
23			83		45.2		119.9	77.9			110.6		220.		
24		13.1	84				120.8	78.4		171.1			221.		
25				71.3	46.3		121.6	79.0		171.9		65	222.	2 14	44.
26				72.1	46.8		122.4	79.5		172.8			223.		
27			87	73.0			123.3	80.1		173.6			223.		
28 29							124.1 125.0	80.6			113.3		224 .		
30		16.3		75.5	49.0		125.8	81.2 81.7			113.8 114.4		225. 226.		
31	26.0	16.9	il	76.3		1	126.6	82.2	'I		114.9		227		
32		17.4		77.2	50.1		127.5	82.8			115.5		228		
33		18.0	93		50.7		128.3	83.3	11		116.0	16 -	229		
34					51.2		129.2	83.9	14		116.6	74	229.		
35				79.7	51.7		130.0	84.4		180.3		75	230 .	61.	49.
36 37	30.2 31.0	19.6 20.2	96 97	80.5			130.8 131.7	85.0			117.6		231.		
38		20.7	98	81.4 82.2	53.4		132.5	85.5 86.1		182.0 182.8			232 . 233 .		
. 39			99	83.0			133.3	86.6			119.3		234		
40	33.5	21.8	100	83.9	54.5	60	134.2	87.1		184.5			234.		
41	34.4	22.3	101	84.7	55.0	161	135.0	87.7	221	185.3	120.4	281	235.	71:	53.
42		22 .9	02	85.5			135.9	88.2		186.2	120.9	82	236.	5 1.	эЗ .
43		23.4		86.4			136.7	88.8			121.5		237 .		
44	36.9	24.0	04	87.2			137.5	89.3			122.0		238.		
45	37.7 38.6	24.5 25.1		88.1 88.9	57.2 57.7		138.4 139.2	89.9 90.4		188.7 189.5	122.5		239. 239 .		
47	39.4	25.6	07	89.7	58.3		140.1	91.0			123.6		240.		
48	40.3	26.1	08	90.6	58.8	68	140.9	91.5			124.2		241.		
49	41.1	26.7	09	91.4	59.4	69	141.7	92.0	29	192.1	124.7	89	242.		
_50	41.9		10	92.3			142.6	92.6		192.9			243 .		
51	42.8	27.8	111	93.1	60.5		143.4	93.1		193.7			244.		
52	43.6	28.3	12	93.9	61.0		144.3	93.7		194.6			244		
53 54	44.4	28.9 29.4	13 14	94.8 95.6	61.5 62.1		145.1 145.9	94.2 94.8		195.4 196.2			245. 246.		
55	46.1	30.0		96.4	62.6		146.8	95.3		196.Z 197.1			240 . 247 .		
56	47.0	30.5	16	97.3	63.2		147.6	95.9		197.9			248		
57	47.8	31.0	17	98.1	63.7		148.4	96.4		198.8		97	249.	1 16	1.1
58	48.6	31.6	18	99.0	64.3		149.3	96.9		19 9 .6			249		
59	49.5	32.1	19	99.8	64.8		150.1	97.5		200.4			250.1		
60	50.3	32.7		100.6	65.4		151.0	98.0		201.3			251.0		
Dist.	Dep.	Lat.	Uist.	Вер.	Lat.	Dist.	Dep.	Lat.	Dist.)	Dep.	Lat.				
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TABLE II.

Difference of Latitude and Departure for 34 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat	Dep.	Dist.	Lat.	Dep.
1	00.8	00.6	61	50.6	34.1		100.3				1 101 .			134.8
2	01.7	01.1	62	51.4	34.7		101.1		82	150.	9 101 . 7 102 .	42	200.6	135.3 135.9
3	02.5 03.3	01.7 02.2	63 64	52.2 53.1	35.2 35.8		102.0 102.8		84	152	5 102		202.3	136.4
5	04.1	02.8	65	53.9	36.3		103.6				4 103.	45	203.1	137.0
6	05.0	03.4		54.7	36.9		104.5		86	154.	2 104.0	46	203.9	
7	05.8	03.9		55.5	37.5		105.3	71.0			0 104.4		204.8	
8	06.6	04.5	68	56.4	38.0		106.1 106.9	71.6 72.1			9 105. 7 105.	7 48	205.6 206.4	139.2
9 10	07.5 08.3	05.0 05.6	69 70	57.2 58.0	38.6 39.1		107.8	72.7			5 106	50	207.3	139.8
11	09.1	06.2	71	58.9	39.7		108.6	73.3			3:106		208.1	
12	09.9	06.7	72	59.7	40.3		109.4				2 107.	52	208.9	140.9
13	10.8	07.3	73	60.5	40.8		110.3	74.4			0 107.		209.7	
14		07.8		61.3	41.4		111.1	74.9			8108.		210.6 211.4	
15	12.4			62.2	41.9		111.9	75.5			7 109 . 5 109 .	111 7.11	211.4 212.2	
16 17	13.3 14.1	08.9 09.5	76 77	63.0 63.8	42.5 43.1	37	112.7 113.6	76.1			3 110		213.1	
18	14.9		78	64.7	43.6		114.4				1 110.	7 58	213.9	144.3
19	15.8		79	65.5	44.2		115.2				0 111		214.7	144.8
20	16.6	11.2	80	66.3	44.7	40	116.1	78.3			8 111.		215.5	
21	17.4	11.7	81	67.2	45.3		116.9	78.8			6 112.		216.4	
22	18.2			68.0	45.9		117.7				5113.0	- 11	217.2 218.0	
23	19.1	12.9	83	68.8	46.4		118.6 119.4		1		3 113.1 1114.1		218.9	
24 25	19.9 20.7	13.4 14.0	84 85	69.6 70 .5	47.0 47.5		120.2				0,114.0		219.7	
26	21.6	14.5		71.3	48.1		121.0		06	170.	8 115.9	66	220.5	148.7
27	22.4		87	72.1	48.6	47	121.9	82.2			6 115.		221.4	
28	23.2		88	73.0			122.7				4:116.		222.2 223.0	
29	24.0		- 89	73.8	49.8	49	123.5	83.5 83.9			3116.9 1117.4		223.0 223.8	
30	24.9		90	74.6	50.3	<u> </u>	124.4				9 118.4		224.7	
31	25.7	17.3	91	75.4	50.9		125.2 126.0	81.4 85.0			8,118.		225.5	
32 3 3	26.5 27.4	17.9 18.5	92	76.3 77.1	51.4 52.0		126.8				6119.		226.3	
34	28.2			77.9	52.6		127.7				4 119.		227.2	
35	29.0	19.6	95	78.8	53.1		128.5				2 120 .		228.0	
36	29.8		96	79.6	53.7		129.3		16	179.	1 120.1 9 121.		228.8 229.6	
37 38	30.7 31.5	20.7 21.2	97 98	80.4 81.2	54.2 54.8	1	130.2 131.0				7 121		230.5	
39	32.3		11	82.1	55.4		131.8				6 122.	79	231.3	156.0
40	33.2			82.9	85.9	60	132.6	89.5			4 123.	80	232.1	156.6
41	34.0	22.9	101	83.7	56.5	161	133.5	90.0	221	183.	2 123.		233.0	
42	34.8	23.5	02				134.3				0 124		233.8	
43	35.6		03	85.4	57.6		135.1				9 124.			158.3 158.8
44 45	36.5 37.3		04 05	86.2 87.0	58.2 58.7		136.0 136.8				7 125. 5 125.		236 3	159.4
46	38.1	25.7		87.9	59.3		137.6				4 126	J 86	237.1	159.9
47	39.0			88.7	59.8	- 67	138.4	93.4	27	188.	2 126 .	87		160.5
48	39.8	26.8		89.5	60.4	68	139.3				0 127.			161.0
49	40.6			90.4	61.0		140.1				8 128.		239.6	161.6 1 62.2
50	41.5	28.0	10	91.2	61.5		140.9	95.1			7 128.	Ji		
51	42.3	28.5	111	92.0	62.1		141.8 142.6				5 129. 3 129.			162.7 163.3
52 53	43.1 43.9	29.1 29.6	12 13	92.9 93.7	62.6 63.2		143.4		33	193	2 130 .:	93		163.8
54	44.8				63.7		144.3				0 130			164.4
55	45.6	30.8	15	95.3	64.3	75	145.1	97.9	35	194.	8 131.4	95		165.0
56	46.4	31.3	16	96.2	64.9		145.9	98.4			7 132.		245.4	165.5
57			17	97.0	65.4		146.7				5 132.4 3 133.			166.1 166.6
58 59	48.1 48.9	32.4 33.0	18 19	97.8 98.7	66.0 66.5	78	147.6 148.4				1 133.			167.2
60	49.7	33.6	20	99.5	67.1		149.2				0 134			167.8
Dist.		Lat.	!		Lat			Lat.	Dist	Den	. Lat.			
-				yp			P.			F		F	or 56 I	Degrees.

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Difference of Latitude and Departure for 35 Degrees.

Dist.	Lat.	Dan	Diat	1-4	D	10!-4	W-4	_	- XI		_		···	
l I———	00.8	Dep. 00.6			Dep.			Dep.				Dist.		Dep.
. 2	01.6	01.1	61 62	50.0 50.8	35.0 35.6	121	99.1 99.9	69.4 70.0		148.3		241		138.2
l ŝ	02.5	01.7	63	51.6	36.1		100.8			149.1 149.9			198.2 199.1	
1 4	03.3	02.3		52.4			101.6			150.7			199.9	
5	04.1	02.9		53.2	37.3	25	102.4			151.5			200.7	
6	04.9	03.4	-66	54.1	37.9		103.2			152.4		46	201.5	141.1
7	05.7 06.6	04.0 04.6	67 68	54.9 55.7	38.4		104.0	72.8		153.2			202.3	
9	07.4		69	56.5	39.0 39.6		104.9 105.7	73.4 74.0		154.0 154.8			203.1	
10	08.2		70	57.3			106.5	74.6		155.6			204.0 204.8	
11	09.0	06.3	71	58.2	40.7		107.3	75.1		156.5				
12	09.8			59.0			108.1	75.7		157.3			205.6 206.4	
13	10.6	07.5		59.8	41.9		108.9	76.3		158.1			207.2	
14	11.5			60.6	42.4		109.8	76.9			111.3	54	208.1	
15 16	12.3 13.1			61.4			110.6	77.4		159.7			208.9	
17	13.9		77	62.3 63.1	43.6 44.2		111.4 112.2	78.0 78.6		160.6			209.7	
18	14.7	10.3	78	63.9			113.0			161.4 16 2 .2			210.5 211.3	
19			79	64.7	45.3		113.9	79.7		163.0			212.2	
20	16.4	11.5	80	65.5	45.9		114.7	80.3		163.8			213.0	
21	17.2	12.0	81	66.4	46.5	141	115.5	80.9	201	164.6	115.3		213.8	
22	18.0		82	67.2	47.0	42	116.3	81.4		165.5			214.6	
23	18.8			68.0	47.5		117.1	82.0		166.3			215.4	
24 25	19.7 20.5		84 85	68.8 69.6	48.2 48.8		118.0			167.1			216.3	
26	21.3	14.9	86	70.4	49.3		118.8 119.6	83.2 83.7		167.9 168.7			217.1	
27	22.1	15.5	87	71.3	49.9		120.4			169.6			217.9 218.7	
28	22.9		88	72.1	50.5		121.2			170.4			219.5	
29	23.8	16.6	89	72.9	51.0		122.1	85.5	09	171.2	119.9		220.4	
30	24.6	17.2	90	73.7	51.6		122.9	86.0	10	172.0	120.5	70	221.2	154.9
31	25.4	17.8	91	74.5	52.2		123.7	86.6		172.8		271	222.0	155.4
32	26.2 27.0	18.4	92	75.4	52.8		124.5	87.2		173.7			222.8	
33 34	27.9	18.9 19.5	93 94	76.2 77.0	53.3 53.9		125.3 126.1	87.8 88.3		174.5 175.3		73	223.6	156.6
35	28.7	20.1	95	77.8	54.5		127.0			176.1			224.4 225.3	
36	29.5	20.6	96	78.6	55.1		127.8	89.5		176.9			226.1	
37	30.3		97	79.5	55.6		128.6	90.1		177.8	124.5		226.9	
38	31.1	21.8	98	80.3	56.2		129.4	90.6		178.6			2 27 .7	
39 40	31.9 32.8	22.4 22.9	99 100	81.1 81.9	56.8 57.4		130.2	91.2		179.4			228.5	
41	33.6	23.5		82.7			131.1	91.8		180.2			229.4	
42	34.4	24.1	101 02	83.6	57.9 58.5		131.9 132.7	92.3 92.9		181.0	126.8 127.3		230.2 231.0	
45	35.2	24.7	03	84.4	59.1		133.5			182.7			231.8 231.8	
44	36.0	25.2	04	85.2	59.7		134.3	94.1		183.5			232.6	
45	36.9	25.8		86.0	60.2	65	135.2	94.6	25	184.3	129.1	85	233.5	163.5
46	37.7	26.4	06	86.8	60.8		136.0			185.1			234.3	
47 48	38.5 39.3	27.0 27.5	07 08	87.6 88.5	61.4 61.9		1 36 .8	95.8 96.4		185.9			235.1	
49	40.1	28.1	09	89.3			138.4			186.8 187.6			235.9 236.7	
50	41.0	28.7	10	90.1	63.1		139.3			188.4			237.6	
51	41.8	29.3	111	90.9	63.7		140.1	98.1		189.2			238.4	
52	42.6	29.8	12	91.7	64.2		140.9			190.0			239.2	
53	43.4	30.4	13	92.6	64.8	73	141.7	99.2	33	190.9	133.6	93	240.0	168.1
54	44.2	31.0	14	93.4				99.8		191.7			240.8	
55 56	45.1 45.9	31.5	15	94.2	66.0		143.4			192.5			241.6	
57	46.7	32.1 32.7	16 17	95.0 95.8	66.5 67.1		144.2 145.0			193.3 194.1			242.5 243.3	
58	47.5	33.3	18	96.7	67.7		145.8			195.0			244.1	
59	48.3	33.8	19	97.5	68.3		146.6			195.8				171.5
60	49.1	34.4	20	98.3	68.8	80	147.4	103.2		196.6		300	245.7	172.1
Dist.	Dep.	Lat.	Dit	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.
												[For	55 D	egrees.

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Difference of Latitude and Departure for 36 Degrees.

-	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Т
	1			[35.9	121	97.9		181	-	106.4		195.0		1
	2			62			22	98.7				107.0		195.8		ı
	3	02.4	01.8				23	99.5	72.3			107.6		196.6	142.8	ı
	4				51.8			100.3				108.2		197.4		ł
	5			65	52.6			101.1				108.7			144.0	ı
	6			66	53.4			101.9				109.		199.0		ı
	7			67 68	54.2 55.0			102.7 103.6				3 109 . 9 1 110 . 5		199.8 200.6		1
	9	07.3		69	55.8			104.4				111.1		201.4		l
	10		05.9	70	56.6			105.2				111.7		202.3		1
	11	08.9	06.5	71	57.4	41.7	-	106.0		ľ.——		112.3		203.1		1
	12		07.1		58.2	42.3		106.8				112.9		203.9		١
	13			73		42.9		107.6				173.4		204.7		
	14	11.3		74	59.9	43.5		108.4				114.0		205.5		ı
	15	12.1	08.8	75	60.7	44.1		109.2				114.6		206.3	149.9	1
•	16	12.9			61.5	44.7	36	110.0	79.9			115.2		207.1		ı
	17				62.3	45.3		110.8				115.8		207.9		1
	18			78	63.1	45.8		111.6				116.4		208.7		ı
	19	15.4		79	63.9	46.4		112.5				117.0		209.5		1
	20	16.2	11.8	80	64.7	47.0		113.3				117.6		210.3		1
	21	17.0	12.3	81	65.5	47.6		114.1	82.9			118.1		211.2		1
	22	17.8			66.3	48.2		114.9				118.7 119.3		212.0		1
1	23 24	18.6 19.4		83 84	67.1 68.0	48.8 49.4		115.7 116.5				19.9 و		212.8 213.6		1
-	25	20.2		85	68.8	50.0		117.3				120.5		214.4		1
1	26	21.0			69.6	50.5		118.1				121.1		215.2		Г
١	27	21.8		87	70.4	51.1		118.9				121.7		216.0		
1	28	22.7	16.5	88	71.2	51.7	48	119.7	87.0	08	168.3	122.5	68	216.8		Г
1	29	23.5		89	72.0	52.3		120.5		09	169.1	122.8		217.6		L
ı	30	24.3	17.6	90	72.8	52 .9	''	121.4				123.4		218.4	158.7	. [
ł	31	25.1	18.2	91	73.6	53.5		122.2				124.0		219.2		1
ı	32	25.9	18.8	92	74.4	54.1		123.0				124.6		220.1		1
Ì	- 33	26.7	19.4	93	75.2			1 23 .8				125.2		220.9		ı
I	34	27.5	20.0	94	76.0	55.3		124.6	90.5			125.8		221.7		1
Į	35 36	28.3 29.1	20.6 21.2	95 96	76.9 77.7	55.8 56.4		125.4 126.2				126.4 127.0		222.5 223.3		1
ı	37	29.9	21.7	97	78.5	57.0		127.0				127.5		224.1		ı
١	38	30.7	22.3	98	79.3	57.6		127.8				128.1			163.4	ı
١	39	31.6	22.9	99	80.1	58.2	59	128.6	93.5	19	177.5	128.7			164.0	ŀ
1	40	32.4	23 .5	100	80.9	58.8	60	129.4	94.0	20	178.0	129.5	80	226.5	164.6	ı
1	41	33.2	24.1	101	81.7	59.4	161	130.3	94.6	221	178.8	129.9	281	227.3	165.2	7
I	42	34.0	24.7	02	82.5	60.0		131.1	95.2			130.5		22 8.1		1
ı	43	34.8	25.3	03	83.3	60.5		131.9				131.1			166.3	П
١	44	35.6	25.9	04	84.1	61.1		132.7	96.4			131.7			166.9	1
١	45 46	36.4	26.5	05	84.9	61.7		13 3 .5 134.5	97.0			132.3			167.5	1
١	47	37.2 38.0	27.0 27.6	06	85.8 86.6	62.3 62.9		135.1	97.6 98.2			132.8 133.4		231.4 232.2		
١	48	38.8	28.2	08	87.4	63.5		135.9	98.7			134.0			169.3	1
ı	49	39.6	28.8	09	88.2	64.1		136.7	99.3			134.6			169.9	ı
١	50	40.5	29.4	10	89.0	64.7		137.5	99.9			135.2			170.5	1
ı	51	41.3	30.0	111	89.8	65.2	171	138.3	100.5	231	186.9	135.8			171.0	1
١	52	42.1	30.6	12	90.6	65.8			101.1			136.4			171.6	1
ł	53	42.9	31.2	13	91.4	66.4			101.7			137.0			172.2	1
١	54	43.7	31.7	14	92.2	67.0			102.3			137.5			172.8	1
١	55	44.5	32.3	15	93.0	67.6			102.9			138.1			173.4	1
1	56	45.3	32.9	16	93.8	68.2			103.5			138.7			174.0	Ţ
1	57 58	46.1 46.9	33.5 34.1	17 18	94.7 95.5	68.8 69.4			104.0 104.6			139.3 1 3 9.9			174.6 175.2	1
1	59	47.7	34.7	19	96.3	69.9			105.2			140.5			175.7	1
	60	48.5	35.3	20	97.1	70.5			105.8			141 1			176.3	١
1:	Dist.	Dep.		Dist.	Dep.	11	Dist.					Lat.	III	Dep.		1
_		- JP:	~~~	2 .00.1	÷ 26.1			<u></u>		LOTE -1	<i></i>	- 2244			egrees.	÷
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Difference of Latitude and Departure for 37 Degrees.

	Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat. Dep.
		00.8	-00.6	61	48.7	36.7	121	96.6			144.6		241 192.5 145.
	1			62	49.5	37 .3	22	97.4				109.5	
	3		01.8	-63			23 24				146.2 146.9		43 194 . 1 146 . 44 194 . 9 146 .
	4			64 65	51.1 51.9	38.5 39.1		99.8				111.3	45 195 .7 147.
	6			66	52.7			100.6				111.9	46 196 . 5 148 .
	1 7			67				101.4				112.5	
	8			68				102.2			150.1		48 198 . 1 149 .
	10			69 70				103.0 103.8			150.9	114.3	491198.9149. 501199.7150.
				71			li	104.6				114.9	{
- 1	11			72	56.7 57.5	42.7 43.3		105.4				115.5	
- 1	13							106.2				116.2	
ı	14	11.2	08.4	74	59.1		34	107.0	80.6			116.8	
1	15			75		45.1		107.8				117.4	
1	16			76	60.7 61.5	45.7		108.6 109.4				118.0 118.6	
1	18				62.3	46.3 46.9		110.2				119.2	
1	19			79		47.5		111.0				119.8	
-1	20	16.0	12.0	80	63.9	48.1	40	111.8	84.3	200	159.7	120.4	60 207.6 156.
1	21	16.8		81				112.6		201	160.5	121.0	
1	22			82				113.4				121.6	
1	23 24				66.3			114.2 115.0				122.2	
1	25				67.1 67.9			115.8				122.8 123.4	
İ	26	20.8		86				116.6				124.0	
ı	27	21.6	16.2	87				117.4		07	165.3	124.6	67 213.2 160.
1	28	22.4			70.3			118.2				125.2	
ı	29 30	23.2 24.0	17.5 18.1	89 90		53.6 54.2		119.0 119.8				125.8 126.4	
1	31				72.7			120.6					l.—————————
1	32	24.8 25.6	19.3	91 92	73.5	54.8 55.4		121.4				127.0 127.6	
	33	26.4	19.9	93	74.3	56.0		122.2				128.2	
	34	27.2	20.5	94	75.1	56.6	54	123.0	92.7	14	170.9	128.8	74 218.8 164.
П	35	28.0	21.1	95	75.9	57.2		123.8				129.4	
Н	36 37	28.8 29.5	21.7	96 97	76.7	57.8 58.4		1 24. 6 125.4				130.0 130.6	
	38	30.3	22.9	98	78.3	59.0		126.2				131.2	78 222 .0 167
ı	3 9	31.1	23.5	99	79.1	59.6		127.0			174.9		79 222.8 167.
1	40	31.9	24.1	100	79.9	60.2	60	127.8	96.3	20	175.7	132.4	80223.6168.
1	41	32.7	24.7	101	80.7	60.8		128.6				133.0	281 224.4 169.
1	42	33.5	25.3	02	81.5	61.4		129.4			177.3		82 225 . 2 169 .
1	43 44	34.3 35.1	25.9 26.5	03 04	82.3 83.1	62.0 62.6		130.2 131.0			178.1	134.2 134.8	83 226 .0 170 . 84 226 .8 170 .
1	45	35.9	27.1	05	83.9	63.2		131.8				135.4	85 227 .6 171
ł	46	36.7	27.7	06	84.7	63.8	66	132.6	99.9	26	180.5	136.0	86 228.4 172.
1	47	37.5	28.3	07	85.5	64.4			100.5			136.6	
1	48 49	38.3 39.1	28.9 29.5	08 09	86.3 87.1	65.6			101.1 101.7		182.1 182.9	137.2	88 230.0 173. 89 230.8 173.
١	50	39.9	30.1	10	87.8	66.2			102.3		183.7		90 231 .6 174
ŀ	51	40.7	30,7	111	88.6	66.8			102.9		184.5		291 232.4 175.
1	52	41.5	31.3	12	89.4	67.4			103.5		185.3		92 233 .2 175.
Ł	53	42.3	31.9	13	90.2	68.0	73	138.2	104.1	33	186.1	140.2	93 234.0 176.
l	54	43.1	32.5	14	91.0	68.6			104.7		186.9		94 234 . 8 176 .
ı	55 56	43.9	33.1 33.7	15 16	91.8 92.6	69.2 69.3			105.3 105.9		187.7 188.5		95 235.6 177.1 96 236.4 178.1
1	57	45.5	34.3	17	93.4	70.4			106.5		189.3		97 237.2 178.
1	58	46.3	34.9	18	94.2	71.0	78	142.2	107.1		190.1		98238.0179.
l	59	47.1	35.5	19	95.0	71.6			107.7		190.9		99 238 . 8 179 . 9
Ļ	60	47.9	36.1	20	95.8	72.2			108.3			144.4	300 239 . 6 180 . 8
I	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist. Dep. Lat.
			•										[For 53 Degree

Difference of Latitude and Departure for 38 Degrees.

Dist.	Lat.	Dep.	Dist	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist .	Let	Dep.
1	00.8		61	48.1	37.6	121	95.3	74.5		142.6	_			148.4
2			62	48.9	38.2	22	96.1	75.1		143.4				149.0
3	02.4		63	49.6	38.8	23	96.9	75.7		144.2				149.6
4	03.2		64	50.4	39.4		97.7	76.3		145.0				150.2
5	03.9			51.2	40.0	25	98.5 99.3	77.0			113.9			150.8
6	04.7			52.0 52.8	40.6 41.2		99.3 100.1	77.6 78.2		146.6 147.4				151.5
8	05.5 06.3			53.6	41.9		100.9	78.8		148.1				152.1 152.7
9	07.1	05.5		54.4	42.5		101.7	79.4		148.9				153.3
10	07.9	06.2	70	55.2	43.1		102.4	80.0		149.7				153.9
11	08.7	06.8	71	55.9	43.7		103.2	80.7		150.5				154.5
12	09.5			56.7	44.3		104.0	81.3		151.3				155.1
13				57.5	44.9		104.8	81.9		152.1				155.8
14	11.0	08.6	74	58.3			105.6	82.5	94	152.9	119.4	54	200.2	156.4
15	11.8			59.1	46.2		106.4	83.1		153.7				157.0
16	12.6			59.9	46.8		107.2	83.7		154.5				157.6
17	13.4	10.5		60.7			108.0 108.7	84.3	97	155.2	121.3			158.2
18 19	14.2 15.0		78 79	61.5 62.3	48.0 48.6		109.7 109.5			156.0 156.8				158.8
20	15.8		80	63.0	49.3		110.3	86.2		157.6				159.5 160.1
-					49.9	1	111.1			158.4				
21	16.5 17.3	12.9 13.5	81	63.8	50.5		111.1 111.9	86.8 87.4			124.4			160.7 161.3
23	18.1	14.2		64.6 65.4	51.1		112.7	88.0			125.0			161.9
24	18.9		84	66.2	51.7		113.5	88.7			125.6			162.5
25	19.7	15.4			52.3		114.3				126.2			163.2
26	20.5	16.0		67.8	52.9	46	115.0	89.9	06	162.3	126.8	66	209.6	163.8
27	21.3	16.6	87	68.6	53.6		115.8	90.5	07	163.1	127.4	67		164.4
28	22.1	17.2	88	69.3	54.2		116.6	91.1		163.9				165.0
29	22.9	17.9	89	70.1	54.8		117.4	91.7		164.7				165.6
30	23.6	18.5	90	70.9	55.4	_	118.2	92.3		165.5		ļ		166.2
31	24.4	19.1	91	71.7	56.0		119.0	93.0		166.3				166.8
32	25.2 26.0		92	72.5	56.6 57.3		119.8 1 2 0.6	93.6 94.2			130.5 131.1			167 . 5 168 . 1
33	26.8			73.3 74.1	57.9		120.6	94.8			131.8			168.7
35	27.6			74.9	58.5		122.1	95.4			132.4			169.3
36	28.4	22.2	96	75.6	59.1		122.9	96.0			133.0			169.9
37	29.2	22.8	97	76.4	59.7		123.7	96.7			133.6			170.5
38	29.9	23.4	98	77.2	60.3		124.5		18	171.8	134.2	78	219.1	171.2
39	30.7	24.0	99	78.0	61.0		125.3				134.8	79	219.9	171.8
40	31.5	24.6	100	78.8	61.6	60	126.1	98.5			135.4	80	220.6	172.4
41	32.3	25.2	101	79.6	62.2	161	126.9	99.1		174.2				173.0
42	33.1	25.9		80.4	62.8		127.7			174.9				173.6
43	33.9	26.5	03	81.2	63.4	63	128.4	100.4			137.3			174.2
44	34.7	27.1	04	82.0	64.0	64	129.2	101.0			137.9 138.5	84	スズン、8 994 c	174.8 175.5
45 46	35.5 36.2	27.7 28.3	05 06	82.7 83.5	64.6 65.3			101.6 102.2		177.3 178.1				176.1
47	37.0	28.9		84.3	65.9			102.2		178.9				176.7
48	37.8	29.6	08	85.1	66.5			103.4		179.7				177.3
49	38.6	30.2	09	85.9	67.1			104.0	29	180.5	141.0			177.9
50	39.4	30.8	10	86.7	67.7	70	134.0	104.7	30	181.2	141.6	90	228 .5	178.5
51	40.2	31.4	111	87.5	68.3	171	134.7	105.3		182.0		291	229.5	179.2
52	41.0	32.0	12	88.3	69.0	72	135.5	105.9	32	182.8		92	230.1	179.B
53	41.8	32.6	13	89.0	69.6			106.5			143.4			180.4
54	42.6	33.2	14	89.8	70.2		137.1			184.4				181 0
55	43.3	33 .9	15	90.6	70.8		137.9			185.2				181 -6
56 57	44.1 44.9	34.5	16	91.4	71.4			108.4		186.0 186.8				182.2 183.9
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59	46.5	36.3	19	93.8	73.3		141.1			188.3		99	235.6	184.1
60	47.3	36.9	20	94.6	73.9		141.8			189.1		300	236 . 4	184.7
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51 39.6 32.1 111 86.3 69.9 171132.9107.6 231179.5145.4 291226.118 52 40.4 32.7 12 87.0 70.5 72133.7108.2 32180.3146.0 92226.918 53 41.2 53.4 13 87.8 71.1 73134.4108.9 33181.1146.6 93227.718 54 42.0 34.0 14 88.6 71.7 74135.2109.5 34181.9147.3 94.228.518 55 42.7 34.6 16 89.4 72.4 75135.0110.1 35182.6147.9 95229.318 56 43.5 35.2 16 90.1 73.0 76136.8110.8 36183.4148.5 96230.018 57 44.3 36.9 17 90.9 73.6 77137.6111.4 57184.2149.1 97230.818 58 45.1 36.5 18 91.7 74.3 78138.3112.0 38185.0149.8 99232.418 59 45.9 37.1 19 92.5 74.9 79139.1112.6 39185.7150.4 99232.418 60 46.6 37.8 20 93.3 75.5 80139.913.3 40186.5151.0 300233.118 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
52 40.4 32.7 12 87.0 70.5 72 133.7 108.2 32 180.3 146.0 92 226.9 18 53 41.2 53.4 13 87.8 71.1 73 134.4 108.9 33 181.1 146.6 93 227.7 18 54 42.0 34.0 14 88.6 71.7 74 135.2 109.5 34 181.9 147.3 94 228.5 18 55 42.7 34.6 16 89.4 72.4 75 135.0 110.1 35 182.6 147.9 95 229.3 18 56 43.5 35.2 16 90.1 73.0 76 136.8 110.8 36 183.4 148.5 96 230.0 18 57 44.3 36.9 17 90.9 73.6 77 137.6 111.4 57 184.2 149.1 97 230.8 18 58 45.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 99 231.6 18 59 45.9 37.1 19 92.5 74.9 79 139.1 112.6 39 185.7 150.4 99 231.4 18 60 46.6 37.8 20 93.3 75.5 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.			_										i !		
53 41.2 33.4 13 87.8 71.1 73 134.4 108.9 33 181.1 146.6 93 227.7 18 54 42.0 34.0 14 88.6 71.7 74 135.2 109.5 34 181.9 147.3 94.228.5 18 55 42.7 34.6 15 89.4 72.4 75 136.0 110.1 35 182.6 147.9 94.228.5 18 56 43.5 35.2 16 90.1 73.0 76 136.8 110.8 36 183.4 148.5 96 230.0 18 57 44.3 36.9 17 90.9 73.6 77 137.6 111.4 57 184.2 149.1 97 230.8 18 58 45.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 98.231.6 18 59 46.9 37.1 19 92.5 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.5 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
54 42.0 34.0 14 88.6 71.7 74 135.2 109.5 34 181.9 147.3 94 228.5 18 55 42.7 34.6 15 89.4 72.4 75 135.0 110.1 35 182.6 147.9 95 229.3 18 56 43.5 35.2 16 90.1 73.0 76 136.8 110.8 36 183.4 148.5 96 230.0 18 57 44.3 35.9 17 90.9 73.6 77 137.6 111.4 57 184.2 149.1 97 230.8 18 58 45.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 98 231.6 18 59 45.9 37.1 19 92.5 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.6 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
55 42.7 34.6 15 89.4 72.4 75 136.0 110.1 35 182.6 147.9 95 229.3 18 56 43.5 35.2 16 90.1 73.0 76 136.8 110.8 36 183.4 148.5 96 230.0 18 57 44.3 35.9 17 90.9 73.6 77 137.6 111.4 57 184.2 149.1 97 1230.8 185.0 149.8 98 231.6 18 58 46.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 98 231.6 18 59 45.9 37.1 19 92.5 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.5 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.		42.0	34.0												
57 44.3 35.9 17 90.9 73.6 77 137.6 111.4 57 184.2 149.1 97 230.8 18 58 45.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 98 231.6 18 59 45.9 37.1 19 92.6 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.5 80 139.9 113.3 40 186.5 151.0 300233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.						72.4	75	136.0	110.1	35	182.6	147.9	95		
58 45.1 36.5 18 91.7 74.3 78 138.3 112.0 38 185.0 149.8 98 231.6 18 59 45.9 37.1 19 92.6 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.5 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
59 45.9 37.1 19 92.5 74.9 79 139.1 112.6 39 185.7 150.4 99 232.4 18 60 46.6 37.8 20 93.3 75.6 80 139.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
60 46.6 37.8 20 93.3 75.5 80 39.9 113.3 40 186.5 151.0 300 233.1 18 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.													u		
Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.															
FP- 21 Day	Dist.														
I KOP DI LIPPI		- Jp.			cp.		-27436-	<i>∠</i> -cp.	, artil.	,2/124	cp.	, 4,44.	[Fo	- 51 T	egres

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Difference of Latitude and Departure for 40 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist	Lat.	Dep.	Dist.I	Lat.	Dep.	Dist.	Lat.	Dep.
1	00.8	00.6	61	46.7	39.2	121	92.7	77.8	181	138.7	116.3	241	184.6	154.9
2	01.5			47.5	39.9	22	93.5	78.4		139.4			185.4	
3	02.3 03.1	01.9 02.6		48.3 49.0	40.5 41.1	23 24	94.2 95.0	79.1 79.7		140.2 141.0			186.1 186.9	
5	03.8	03.2	65	49.8	41.8	25	95.8	80.3		141.7			187.7	
6	04.6	03.9	66	50.6	42.4	26	96.5	81.0	86	142.5	119.6	46	188.4	158.
7	05.4 06.1	04.5 05.1	67 68	51.3 52.1	43.1 43.7	27 28	97.3 98.1	81.6 82.3		143.3 144.0			189.2	
8	06.9	05.8		52.9	44.4	29	98.8	82.9		144.8			190.0 190.7	
10	07.7	06.4	70	53.6	45.0	30	99.6	83.6		145.5			191.5	
11	08.4	07.1	71	54.4	45.6	131	100.4	84.2	191	146.3	122.8	251	192.3	
12	09.2	07.7	72	55.2	46.3		101.1	84.8			123.4		193.0	
13	10. 0	08.4 09.0	73 74	55.9 56.7	46.9 47.6		101.9 102.6	85.5 86.1		147.8 148.6			193.8 194.6	
14	11.5	09.6	75	57.5	48.2		103.4	86.8			125.3			
16	12.3	10.3	76	58.2	48.9	36	104.2	87.4	96	150.1	126.0	56	195.3 196.1	164.
17	13.0	10.9	77	59.0	49.5		104.9	88.1		150.9			196.9	
18 19	13.8 14.6	11.6 12.2	78 79	59.8 60.5	50.1 50.8		105.7 106.5	88.7 89.3		151.7 152.4	127.3		197.6 198.4	
20	15.3	12.9	80	61.3	51.4		107.2	90.0		153.2			199.2	
21	16.1	13.5	81	62.0	52.1	141	108.0	90.6		154.0			199.9	
22	16.9	14.1	82	62.8	52.7	42	108.8	91.3		154.7		62	200.7	168.
23	17.6	14.8	83	63.6	53.4		109.5	91.9		155.5			201.5	
24 25	18.4 19.2	15.4 16.1	84 85	64.3 65.1	54.0 54.6		110.3 111.1	92.6 93.2		156.3	131.1 131.8		202.2 203.0	
26	19.9	16.7	86	65.9	55.3		111.8	93.8		157.8			203.8	
27	20.7	17.4	87	66 6	55.9	47	112.6	94.5	07	158.6	133.1	67	204.5	171.
28	21.4	18.0	88	67.4	56.6		113.4	95.1		169.3			205.3	
29 30	22.2 23.0	18.6 19.3	89 90	68.2 68.9	57.2 57.9		114.1 114.9	95.8 96.4		160.1	134.3		206 . 1 206 . 8	
31	23.7	19.9	91	69.7	58.5		115.7	97.1		161.6			207.6	
32	24.5	20.6	92	70.5	59.1		116.4	97.7		162.4			208.4	
33	25.3	21.2	93	71.2	59.8		117.2	98.3	13	163.2	136.9	73	209.1	175 .
34	26.0	21.9	94	72.0	60.4		118.0	99.0			137.6		209.9 210.7	
35 36	26.8 27.6	22.5 23.1	95 96	72.8 73.5	61.1		118.7 119.5	99.6 100.3		164.7 165.5			211.4	
37	28.3	23.8	97	74.3	62.4	57	120.3	100.9			139.5		212.2	
38	29.1	24.4	98	75.1	63.0		121.0				140.1		213.D	
39 40	29.9 30.6		99 100	75.8 76.6	63.6 64.3		121.8 122.6			167.8 168.5	140.8		213.7 214.5	
41	31.4	26.4	101	77.4			123.3		-	169.3			215.3	
42	32.2	27.0	02	78.1	64.9 65.6		124.1			170.1			216.0	
43	32.9	27.6	03	78.9	66.2	63	124.9	104.8	23	170.8	143.3	83	216.8	181 .
44	33.7	28.3	04	79.7	66.8		125.6				144.0		217.6	
45	34.5 35.2	28 .9 29 .6	05 06	80.4 81.2	67.5 68.1		126.4 127.2				144.6 145.3		218.3 219.1	
47	36.0	30.2	07	82.0	68.8		127.9			173.9		87	219.9	184
48	36.8	30.9	08	82.7	69.4	68	128.7	108.0		174.7			220.6	
49 50	37.5 38.3	31.5 32.1	09	83.5	70.1		129.5			175.4			221.4 222.2	
	39.1		10	84.3	70.7		130.2 131.0			176.2		_	222.9	
51	39.1 39.8	32.8 33.4	111 12	85.0 85.8	71.3 72.0		131.8			177.0 177.7			223.7	
53	40.6	34.1	13	86.6	72.6	73	132.5	111.2	33	178.5	149.8	93	224.5	188 .
54	41.4	34.7	14	87.3	73.3		133.3				150.4		225.2	
55 56	42.1	35.4 36.0	15	88.1 88.9	73.9 74.6		134.1 134.8			180.0 180.8	151.1		226.0 226.7	
57	43.7	36.6	16 17	89.6	75.2		135.6				152.3		227.5	
58	44.4	37.3	18	90.4	75.8		136.4		38	182.3	153.0	98	228.3	191.
59	45.2	37.9	19	91.2	76.5		137.1				153.6		229.0	
60	46.0	38.6	20	91.9	77.1		137.9 Dep.				154.3		229.8	1
13(.)	Dep.	Lat.	DISt.	Dep.	Lat.	Digt.	illen i	Lat.	I Digt.	1 len	laf.	·LUST.	Deb.	· LBI

Difference of Latitude and Departure for 41 Degrees.

1	,Di	st. La	ŧ	Dep.	i Dis	ti La	e. i O	en.	RDLst.	Lei		Dep.	l:Dist.	La	•	Den	Dist.	I T.el	-	De	
ļ	F	'	.8	00.	-11			0.0				79.4		_	_	118.	_	181			
1		2 01	.6	01.5	3 6	2 46	.8 4	Ø.7	22	92	.1	80.0) 82 }	137	.4	119.	42	182			
1		3 02		02.0		3 47		1.3				80.7				120.		183			
١	1	4 03 5 03		02.0		4 48 5 49		2.0 2.6				81.4 82.0				120. 121.		184			
١	1	6 04		03.9				3.3				82.7				122	: HL	185			
١	1	7 05		04.6				4.0				83.5				122		186			
į	1	8 06 9 06		05.9 05.9		8 51 9 52		4.6 5.3	28 29			84.6 84.6		142		123. 124.		187			
1	1 1	0 07		06.6				5.9	30			85.5				124.		188			
ı	-	1 08	_	07.5	J			6.6	131	98		85.9	_	144	1	125.	<u> </u>	189			
1	1	2 09	. 1	07.5	7	2 54	.3 4	7.2	32	99	6	86.6	92	144	.9	126.	52	190	.설	165	j.;
١		3, 09		08.8				7.9		100	4	87.5				126. 127.	. n	190			
١		4 10 5 11		09.2 09.8				8.5 9.2		101 . 101 .		87.9 88.6				127.		191 192			
İ		6 12		10.5				9.9		102		89.1				128.		193			
ı		7 17		11.5				0.5		103.		89.9				129.		194			
ı		8 13 9 14		11.8 12.5				1.2 1.8		104. 104.		90.5 91.2				129.		194 195			
1			.1	13.1				2.5,		105		91.8				131.		196			
1	2		_	13.8	H	_'		3.1		106.		92.5	1	151			1	197	_		_
1	2	2 16	.6	14.4	8:	2 61	9 5	3.8	42	107.	2	93.2	02	152	. 5	132.	62	197			
ł	2			15.1				4.5		107.		93.8				133.		198			
١	2			15.7 16.4				5.1		108 . 109 .		94.8 95.1				1 33 . (1 34 . (199 200			
t	2			17.1				5.4		110.		95.8				155.		200			
1	2	7 20.	4	17.7	84	65.	7 5	7.1		110.		96.4				135.8	67	201.	5	175	. 2
1	2			18.4 19.0				7.7 3.4		111. 112.		97.1 97.8				136 . : 1 37 . 1		202 . 203 .			
1	3			19.7				.0		113.		8.4				137.8		203.			
	3	23.	귛-	20.3	<u> </u>			7.7		114.	ساء	9.1	211					204.			_
	35	24.	획	21 0	92	69.	4 60	.4	52	114.	7 9	9.7	12	160	0	139.1	72	205.	3 1	178	.4
	33			21.6				.0		115.		10.4				139.7		206.			
П	34 34			22.3 23.0	94 95			.7		116.9 117.0						140.4 141.1		206 . 207 .			
П	36	27.	2	23.6	96			.0		17.						41.7	76	20 8.	3 1	81	. 1
П	37		9 :	24.3	97					18.						42.4		209.			
П	38 39			24.9 25.6	98 99					119.5 120.(143.0 143.7		209. 210.			
IJ	40			16.2	100				60	20.8	10	5.0				44.3		211.			
H	41	30.	9	6.9	101	76.	2, 66	3	161	21.8	īō	5.6	221	166.	8 1	45.0		212.			
il	49			7.6	02		66			22.5						45.6	82	212.	B	85.	0
П	43 44			28.2 18.9	03 04		7i 67 5i 68			23 .0 23 .8						46.3 47.0		213.0 214.3			
Н	45				05	79				24.5						47.6		215.			
	46	34.7		9.5 0.2	06	80.0			66.1	25.3	10	B.9				48.3		115.			
1	47 48	35.8 36.2		10.8 11.5	07 08					26 .0 26.8						48.9 49.6		216.(217.			
1	49	37.0		2.1	09	82.5				20.0 27.5						50.2		218.1			
	50	37.7		2.8	10	83.0				28.3						50.9		218.9			
Ī	51	38.5		3.5	111	83.8			171,1				231				291				
1	52 53	39.2 40.0		4.1 4.8	12	84 5				29 .8						52.2		20.4 21.1			
1	54	40.8		5.4	13 14	85.5 86.0				30.6 31.3						52 .9! 53.5		21.			
ı	55	41.5	3	6.10	15	86.8	75.	4		32.1			35 1	77.	4 1	54.2	95	22.6	5 1	93.	ĕ
	56 57	42.3	3	6.7	16	87.5				32.8			36	70.	1 1	54.8		23.4 24.1			
l	58	43.0 43.8		7.4 B.1	17 18	88.3 89.1			77 13	33.6 34.3						55.5 56.1		24.9			
ľ	59	44.5	31	8.7	19	89.8	7 8 ·	1	79 13	35 . 1	117	- 1	39	80.4	6 1	56.8	99	25.7	71:	96.	2
L	60	45.3	_	9.4	20	90.6	78.		80 13				واستند	-	-1-	57.5	300	_	-	_	_
g D	ist.	Dep.	L	at. II	Dist.	Dep.	Lat	. ľD	ist. L	ep.	L	<u>i I</u>	Dist	Dep	1	Lat.	Dist.		_		-
											-						[For	59 L	reg	ree	4

Difference of Latitude and Departure for 42 Degrees.

Dist.	Lat	Dep.	Dist.	Lat.	Dep.	Dist	Lat	Dep.	Dist J	Tat. I	Dep.	.Diet	Let	Dep.
						121	89.9			34.5		_		161.3
1 2		00.7 01.3	61 62	45.3 46.1	41.5		90.7	81.6			121.8	49	170 8	161.9
3					42.2	23				136.0		43	180.6	162.6
4				47.6		24		83.0		136.7				163.3
5				48.3		25	92.9	83.6	85	137.5	123.8			163.9
6			66	49.0	44.2	26	93.6	84.3	86	138.2	124.5			164.6
7	05.2	04.7	67	49.8	44.8	27	94.4	85.0	87	139.0	125.1	47	183.6	165.3
8	05.9	05.4	68	50.5	45.5	28	95.1	85.6		139.7				165.9
9	06.7	06.0	69	51.3	46.2	29	95.9	86.3		140.5				166.6
10	07:4	06.7	70	52.0	46.8	30	96.6	87.0	90	141.2	127.1			167.3
11	08.2	07.4	71	52.8	47.5	131	97.4	87.7		141.9		251	186.5	168.0
12	08.9	08.0	72	53.5	48.2	32	98.1	88.3		142.7		52	187.3	168.6
13	09.7	08.7	73	54.2	48.8	33	98.8	89.0		143.4				169.3
14	10.4	09.4	74	55-0	49.5	34					129.8			170.0
15	11.1	10.0	75	55.7	50.2		100.3			144.9				170.6
16	11.9	10.7		56.5	50.9		101.1	91.0		145.7				171.3
17	12.6	11.4 12.0		57.2	51.5 52.2		101 .8 10 2 .6	91.7 92.3		146.4				172.0
18 19	13.4 14.1	12.7	79	58.0 58.7	52.9		103.3	9 3 .0		147.9				172.6 173.3
20	14.9	13.4	80	59.5	53.5		104.0	93.7		148.6		60	193 9	174.0
		14.1	!											
21	15.6		81	60.2	54.2		104.8 105.5	94.3		149.4				174.6
22	16.3	14.7 15.4	82 83	60.9 61.7	54.9	42	106.3	95.0 95.7		150.1 150.9				175.3 176.0
24	17.1 17.8	16.1		62.4	55.5 56.2	44	107.0	96.4		151.6				176.7
25	18.6	16.7	85	63.2	56.9		107.8	97.0		152.3				177.3
26	19.3	17.4	86	63.9	57.5		108.2	97.7		153.1				178.0
27	20.1	18.1	87	64.7	58.2		109.2			153.8				178.7
28	20.8	18.7	88	65.4	58.9		110.0			154.6				179.3
29	21.6	19.4	89	66.1	59.6		110.7			155.3				180.0
30	22.3	20.1	_ 90	66.9	60.2	50	111.5	100.4	10	156.1	140.5	70	200.6	180.7
31	23.0	20.7	91	67.6	60.9	151	112.2	101.0	211	156.8	141.2	271	201.4	181.3
32	23.8	21.4	92	68.4			113.0			157.5				182.0
33				69.1		53	113.7	102.4			142.5			182.7
34	25.3	22.8	94	69.9				103.0			143.2			183.3
35	26.0	23.4	95	70.6			115.2				143.9			184.0
36	26.8	24.1		71.3	64.2			104.4			144.5			184.7
37	27.5			72.1	64.9			105.1			145.2			185.3
38 39		25.4 26.1	98 99	72.8 73.6	65.6 66.2		117.4	106.4			145.9 146.5			186.0 186.7
40	29.7	26.8	100	74.3	66.9		118.9				147.2			187.4
_			J									!!		
41	30.5	27.4	101	75.1	67.6		119.6				147.9			188.0 5188.7
43	31.2 32.0			75.8 76.5			120.4	108.4			148.5 149.2			189.4
44	32.7	29.4						109.7			149.2	84	211	190.0
45								110.4			150.6			190.7
46	34.2							iii.i			151.2			5 191.4
47	34.9							111.7			151.9		213.	192.0
48		32.1	08	80.3				112.4			152.6		214.0	192.7
49	36.4	32.8	09	81.0	72.9	69	125.6	113.1	29	170.2	153.2	89		8 193.4
.50	37.2	33.5	10	81.7	73.6	70	126.3	113.8	30	170.9	153.9	90	215.	5 194.0
51	37.9	34.1	111	82.5	74.3	171	127.1	114.4	231	171.7	154.6			3194.7
52	38.6	34.8		83.2	74.9			115.1			155.2			0 195.4
53			13		75.6			115.8			155.9			7 196.1
54	40.1	36.1	14	84.7	76.3			116.4			156.6			5 196.7
55	40.9	36.8	15	85.5	77.0			117.1			157.2			2 197.4
56 57	41.6	37.5	16	86.2	77.6			117.8			157.9			0198.1
58	42.4	38.1 38.8	17 18	86.9 87.7	78.3	33	131.6	118.4	37	176 0	158.6 159.3	9		7 198.7 5 199.4
59	43.8	39.5	19	88.4	79.0	7 8	132.3	119.1			159.9			2 200.1
60	44.6	40.1	20	89.2	79.6 80.3			119.8 1 20 .4	16		160.6	11		9 200.7
	Dep.	Lat.			_			7 24						. Let.
,		Dat.		Dep.	Lat.	na i	Dep.	, will.	II DISC.	Deb.	- Dat-			Degrees
												Fr. o	. 40 1	~eBices

Difference of Latitude and Departure for 43 Degrees.

Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat. Dep.
1	00.7	00.7	61	44.6	41.6	121	88.5	82.5		132.4		241 176.3 164.4
2	01.5	01.4	62	45.3	42.3	22	89.2	83.2		133.1		42 177.0 165.0 43 177.7 165.7
3	02.2 02.9	02.0	63 64	46.1 46.8	43.0	23 24	90.0 90.7	83.9 84.6		133.8 134.6		44 178.5 166.4
5	03.7	03.4	65	47.5	44.3	25	91.4			135.3		45 179.2 167.1
: 6	04.4	04.1	66	48.3	45.0	26	92.2	85.9	86	136.0	126.9	46 179.9 167.8
7	05.1	04.8	67	49.0	45.7	27	92.9			136.8		47 180 .6 168 .5
8	05.9 06.6	05.5 06.1	68 69	49.7 50.5		28 29	93.6 94.3	87.3 88.0		137.5 138.2		48 181.4 169.1 49 182.1 169.8
10	07.3	06.8	70	51.2	47.7	30	95.1	88.7		139.0		50182.8170.5
11	08.0	07.5	71	51.9	48.4	131	95.8	89.3		139.7		251 183.6 171.2
12	08.8		72	52.7	49.1	32	96.5			140.4		52 184.3 171.9
13	09.5		73	53.4		33	97.8			141.2		53 185.0 172.5
14			74	54.1	50.5	34	98.0 98.7			141.9		54185.8173.2 55186.5173.9
15			75 76	54.9 55.6		35 36	99.5			142.6 143.3		56 187.2 174.6
17	12.4	11.6	77	56.3		37	100.2	93.4		144.1		57 188.0 175.3
18	13.2	12.3	78	57.0	53.2	38	100.9	94.1	98	144.8	135.0	58 188.7 176.0
19	13.9		79	57.8		39	101.7	94.8		145.5		59 129 4 176 .6
20	14.6		80	58.5	54.6		102.4				136.4	60 190.2 177.3
21	15.4	14.3 15.0	81	59.2	55.2		103.1 103.9	96.2 96.8		147.0 147.7		261 190.9 178.0 62 191.6 178.7
22 23	16.1 16.8	15.7	82 83	60.0	55.9 56.6	43	104.6	97.5			138.4	63 192.3 179.4
24	17.6		84	61.4		44	105.3			149.2		64 193 . 1 180 . 0
25	18.3		85	62.2	58.0	45	106.0	98.9	05	149.9	139.8	65 193.8 180.7
26	19.0		86	62.9	58.7			99.6		150.7		66 194.5 181.4
27	19.7	18.4	87	63.6	59.3	47	107.5	100.3		151.4		67 195.3 182.1 68 196.0 182.8
28 29	20.5 21.2	19.1 19.8	88 89	64.4 65.1	60.0 60.7	48	100.2	100.9 101.6		152.1 152.9		69 196.7 183.5
30	21.9	20.5	90	65.8	61.4	50	109.7	102.3		153.6		70 197.5 184.1
31	22.7	21.1	91	66.6	62.1			103.0	211	154.3	143.9	271 198.2 184.8
32	23.4		92	67.3	62.7			103.7	12	155.0	144.6	72 198.9 185.5
33	24.1	22.5	93	68.0				104.3			145.3	73 199.7 186.2
34	24.9		94	68.7	64. J			105.0		156.5 1 57 .2		74 200.4 186.9 75 201.1 187.5
35 36	25.6 26.3	23.9 24.6	95 96	69.5 70.2	64.8 65.5			105.7 106.4			147.3	76 201.9 188.2
37	27.1	25.2	.97	70.9	66.2		114.8			158.7		77 202 .6 188 .9
38	27.8		98	71.7	66.8			107.8		159.4		78 203.3 189.6
39	28.5		99	72.4	67.5			108.4		160.2		79 204.0 190.3
40	29.8		100	73.1	68.2		117.0			160.9		80 204.8 191.0
41	30.0	28.0 28.6	101	73.9 74.6	68.9 69.6			109.8 110.5		161.6 162.4		281 205.5 191.6 82 206.2 192.3
42	30.7 31.4	29.3	02 03	75.3	70.2			111.2		163.1		83 207.0 193.0
44	32.2		04	76.1	70.9			111.8	24	163.8	152.8	84 207.7 193.7
4.5	32.9	30.7	05	76.8	71.6			112.5			153.4	
46	33.6		06	77.5	72.3			113.2	26	165.3	154.1	86 209.2 195.1 87 209.9 195.7
47	34.4	32.1 32.7	07	78.3 79.0	73.0 73.7			113.9 114.6		166.0	155.5	88 210.6 196.4
48 49	35.1 35.8	33.4	08 09	79.7	74.3			115.3		167.5		89211.4197.1
50	36.6	34.1	10	80.4	75.0			115.9		168.2		90212.1197.8
51	37.3	34.8	111	81.2	75.7	171	125.1	116.6	231	168.9		291 212.8 198.5
52	38.0	35.5	12	81.9	76.4			117.3	32	169.7	158.2	92213.6 199.1
53	38.8		13	82.6	77.1			118.0		170.4		93214.3 199.8 94215.0 200.5
54 55	39.5 40.2		14 15	83.4 84.1	77.7 78.4			118.7 119.3		171.1 171.9	160.3	
56	41.0		16	84.8				120.0	36	172.6	161.0	96 216.5 201.9
67	41.7	38.9	17	85.6	79.8	77	129.4	120.7	37	173.3	161.6	97 217.2 202.6
58		39.6	18	86.3				121.4	38	174.1	162.3	98 217.9 203.2
59	43.1	40.3	19 20	87.0 87.8	81.2 81.8			122 . 1 122 . 8	39	174.8 175.5	163.7	99 218.7 203.9 300 219.4 204.6
Dist.	43.9	40.9					Dep.		Diet	Den	Let	Dist. Dep. Lat.
ואועוי	Deb.	LEL	. احال	Dep.	, Let.	ا <i>حالا</i> زا		· Let.	-2-4-1	- Jup.	2000	[For 47 Degrees.
										•		[: -: -: -: -: -: -: -: -: -: -: -: -:

TABLE II.

Difference of Latitude and Departure for 44 Degrees.

1 00 7 00 7 6 1 43 9 42 4 121 87 0 84.1 181 90.2125 7 241173.4167.4 2 01 4 01 4 63 44.6 43.1 22 87.8 84.1 181 90.2125 4 42174.1168.1 24 4 02 9 02 81 64 46.0 44.0 24 82 82 86.8 6 84.6 13 13.6 127.1 43174.8168.8 84.6 6 6 6 6 6 6 6 7 5 45.8 26 90.6 87.5 86 133.1 128.5 44.177.0 170.9 8 6 05.6 05.6 06.4 66 47.5 45.8 26 90.6 87.5 86 133.1 128.5 44.177.0 170.2 8 05.6 05.6 05.6 66 47.5 45.8 26 90.6 87.5 86 133.1 128.5 44.177.0 170.2 8 05.6 05.6 05.6 66 47.5 45.8 26 90.6 87.5 86 133.1 128.5 44.177.0 170.5 180.5 180.5 10 07.2 06.0 10 07.2 06.0 10 07.2 06.0 10 07.2 06.0 10 07.2 06.0 10 11.4 93.1 19.4 2 91.0 19.1 137.4 137.5 16.5 1.1 10 07.2 07.6 7 10 11.4 93.3 13.9 42.9 10.0 19.1 137.4 133.7 251 120.6 1417.3 4175.5 150.0 12.0 10 07.2 06.0 10 07.2 07.6 10 11.4 93.3 19.4 2.9 10.0 19.1 137.4 133.7 251 120.6 174.4 12.0 10 09.7 7 74 53.2 5.5 50.7 39.5 95.7 19.1 13.0 15.7 133.0 50.1 120.0 177.4 135.0 10 07.2 06.0 10 07.2 07.2 13.0 07.2 07.2 13.0 10 07.2 07.2	Dis	t. Lat.	Dep.	Dist.	Lat.	Dep.	Dist.	Lat.	Dep.	Dist. Lat.	Dep.	Dist. Lat. Dep.
2 01.4 01.4 62 44.6 43.1 22 37.8 84.77 83131.6127.1 43114.81688.1 3 02.2 02.1 65 45.5 45.8 42.8 92.8 66.1 93131.6127.1 431175.6169.5 63.6 03.5 65 46.8 45.2 24 99.9 86.8 85.133.1128.5 441175.6169.5 60 45.3 04.9 66 47.5 45.8 25 99.9 86.8 85.133.1128.5 441175.6169.5 80 56.0 05.6 66 48.9 47.2 28 92.8 86.8 85.133.1128.5 441177.0170.9 97.6 60 50 60.9 70 80.0 4 86.6 30 37.2 19.4 88.8 130.9.130.5 541177.0170.9 10 07.2 06.9 70 80.4 48.6 30 33.5 90.3 90 136.7 133.0 560 1378.1177.1171.6 11.0 10 70.9 07.6 71 61.1 49.3 131 94.2 91.0 19.137.4 133.7 251180.6 174.4 140.1 09.7 74 63.2 56.0 6.2 1.5 50.7 33 95.7 92.4 93.138.8 133.4 85118.3 175.1 13.0 94.4 09.0 73 62.5 50.7 33 95.7 92.4 93.138.8 133.4 85118.3 175.1 15.0 8 10.4 75 64.0 52.1 55.9 71.9 3.8 9140.5135.5 55182.0 175.7 151.1 12.2 11.8 77 55.4 53.5 57.9 94.9 96.6 99.143.1 133.2 55182.0 175.7 12.2 11.8 177 55.4 53.5 57.9 94.9 99.143.1 133.2 55182.0 175.7 12.2 11.8 17.1 15.1 16.6 11.5 11.5										181 130.5	125.7	241 173.4 167.4
4 02.9 02.8 6.4 4.0 4.1 2.4 89.2 66.1 84.1 32.4 127.8 44177.5169.5 80.5 03.6 03.6 03.6 62.5 65.4 46.8 46.2 25.8 69.6 677.5 68.8 81533.3 1128.5 45176.2 170.170.9 70.5 04.2 66.7 47.5 45.8 26.8 9.9 68.8 81533.3 1128.5 45176.2 170.170.9 90.6 5.0 6.5 68.8 48.9 47.2 28.9 2.1 88.9 8153.2 130.6 44174.177.171.6 69.0 6.5 66.3 69.4 46.5 47.9 29.9 23.8 89.8 8153.2 130.6 44178.4 172.3 10.0 07.2 66.9 70.50.4 48.6 63.0 93.5 90.3 90.136.7 133.4 9179.1 173.0 10.0 72.0 6.9 70.50.4 48.6 63.0 93.5 90.3 90.136.7 133.4 9179.1 173.0 112.0 8.6 68.3 71.1 49.3 131.1 44.2 11.1 49.3 131.1 44.2 11.1 49.3 131.1 44.2 11.1 49.3 131.1 44.2 11.1 49.3 131.1 44.2 11.1 49.3 131.1 44.2 11.1 12.3 130.6 4.2 11.1 49.3 131.1 44.2 11.1 12.3 130.6 4.2 11.1 12.3 130.6 4.2 11.1 12.3 130.6 4.2 11.1 12.3 130.6 4.2 11.1 12.3 130.6 4.2 11.1 12.3 130.6 4.2 11.1 12.3 130.9 4.0 19.7 74.6 53.2 51.4 4.5 64.3 13.1 14.5 13.1 13.4 12.3 13.1 14.5 13.2 11.1 14.5 13.2 11.1 14.5 13.2 11.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 13.1 14.5 13.2 14.1 14.5 14.5 14.5 14.5 14.5 14.5 14.5	Н	2 01.4	01.4	62	44.6	43.1						
5 03.6 03.5 63 46.8 47.2 25 89.9 66.8 85 133.3 1128.5 45176.2170.2 9 66.4 04.5 66 47.5 45.8 26 90.6 87.5 86135.3 8129.2 46177.0 170.2 8 05.8 05.6 05.6 68 48.9 47.2 28 92.1 88.9 8133.2 130.6 48178.4 172.6 3 10 07.2 06.9 70 50.4 48.6 30 93.5 90.5 90.5 97.6 76 71 10 07.9 06.9 70 50.4 48.6 30 93.5 90.3 90.8 9135.0 133.0 50179.8 173.7 11 07.9 06.9 70 50.4 48.6 30 93.5 90.3 90.1 30.5 133.0 50179.8 173.7 11 07.9 06.9 70 50.4 48.6 30 93.5 90.3 90.1 30.5 133.0 50179.8 173.7 11 07.9 06.9 70 50.4 48.6 30 93.5 90.3 91.7 92138.8 133.3 4 52181.3 175.1 13 09.4 09.0 73 52.5 50.7 33 95.7 92.4 49.1 133.7 4 52182.0 174.4 10.1 99.7 74.5 52.0 52.0 12.0 12.1 12.1												
6 04.5 04.2 66 47.9 22 80.6 87.5 86 133.8 192.9 2.4 46177.0 170.9 6 70.5 0 04.2 67 48.2 46.5 27 91.4 88.2 87 134.5 192.9 2.4 47177.7171.9 6 8 05.8 05.6 68 48.9 47.2 28 92.1 88.9 88 136.0 151.3 49179.1 173.7 110 07.9 07.6 9 70.6 9 70.5 04.4 46.6 30.9 35.5 90.3 90.130.7 132.0 6 49179.1 173.7 112 07.9 07.6 9 70.5 04.4 46.6 30.9 35.5 90.3 90.130.7 132.0 6 149179.1 173.7 112 07.9 07.6 9 70.5 04.4 46.5 0.0 13.2 91.0 91.7 92.1 132.7 251 120.6 174.4 120.1 09.7 74 85.2 61.4 34.9 64.9 35.7 193.8 91.5 193.8 1133.4 155 122.0 175.7 131.1 12.9 12.5 78 56.1 64.2 38 99.3 95.9 95.1 99.1 14.1 156.8 57 184.9 177.5 18 12.9 12.5 78 56.1 64.2 38 99.3 95.9 98.1 42.5 113.2 79.9 114.4 13.9 80 67.5 55.6 40 100.7 97.3 200 145.9 139.8 96.1 13.6 18.8 15.3 82.5 0.5 55.6 40 100.7 97.3 200 145.9 139.8 96.1 13.6 18.8 15.3 82.5 0.5 55.6 40 100.7 97.3 200 145.9 139.9 13.7 18.1 16.7 18.6 61.7 60.4 55.5 6.6 40 100.7 97.3 200 145.9 139.9 60.1 14.4 13.9 80 67.5 55.6 40 100.7 97.3 200 145.9 139.9 60 147.0 180.6 18.2 177.1 18.1 18.1 18.1 18.6 18.5 83.5 55.7 57.7 4.5 102.9 99.3 00.1 44.6 133.6 6 18.8 15.8 15.3 82.5 0.4 55.0 6.1 45.0 14.0 14.0 14.0 14.0 15.0 14.0 14.0 14.0 14.0 15.0 14.0 14.0 14.0 14.0 14.0 14.0 14.0 14	11											
7 05.0 04.9 67 48.2 47.2 28 92.1 88.2 87134.5 129.9 47117.7171.6 8 05.8 05.8 05.6 06.5 68 48.9 47.2 28 92.1 88.9 88 136.3 2130.6 48178.4172.5 10 07.2 06.9 70 .50.4 48.6 30 93.5 90.3 90.3 05.7 132.0 50 1279.8 173.0 11 07.9 07.6 71 13 09.4 09.0 73. 51.1 49.3 131 94.2 19.1 0 19 1137.4 132.7 251 120.6 174.4 13.1 14 10.1 09.7 74 63.2 51.4 24 96.4 93.1 34 10.1 10.8 10.4 75 54.0 52.1 35 97.1 93.8 94.5 19.1 0 19 1137.4 132.7 251 120.6 174.4 10.1 10.8 10.4 75 54.0 52.1 35 97.1 93.8 94.5 19.1 0 19 1137.4 132.7 13.2 11.8 77 55.4 05.5 57 98.5 95.2 9714.0 136.2 66184.2 177.8 18 12.9 12.5 78 66.1 64.2 38 99.3 95.9 93 140.5 136.2 66184.2 177.8 19 13.7 13.2 79 56.8 54.9 39 100.0 96.6 93 143.1 136.2 66184.2 177.8 19 13.7 13.2 79 56.8 54.9 39 100.0 96.6 93 143.1 139.8 60187.0 180.6 12.2 18.8 18.2 18.3 18.3 18.3 18.3 18.3 18.3 18.3 18.3												
8 80 5.8 0 5.6 68 48.9 47.2 28 92.1 88.9 88 133.3 2150.6 48178.4172.3 9 0.6 5.6 0.5 16 94 5.4 77.9 29.24.8 99.6 6.0 135.0 135.1 34 9179.1173.7 110 77.9 0 77.6 71 51.1 49.3 151 94.2 91.0 17. 110 77.9 0 77.6 71 51.1 49.3 151 94.2 91.0 17. 121 18.1 18.1 19.5 18.2 51.8 50.0 32 95.0 91.7 92.138.1 133.4 52.113.7 11. 19. 19. 19. 19. 19. 19. 19. 19. 19.										87 134	129.9	
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60 43.2 41.7 20 66.3 83.4 80 129.5 125.0 40 172.6 166.7 300 215.8 208.4 Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.	5	8 41.7	40.3	18	84.9	82.0						
Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat. Dist. Dep. Lat.					85.6							997210.17207.7
									120.0			Died Dor 7 5
f.r.os. 40 Deficer	Dis	Li Dep.	Lat.	Dist	Dep.	Lat.	Dist.	Dep.	Lat.	INSL. Dep.	LOL	
	1											fror so refrage

TABLE II.

Difference of Latitude and Departure for 45 Degrees.

-			5 1.	•	S	131.a		1	T		D 1	DI A	¥ -4	
Dist.			Dist.	Lat.		Dist.	Lat.			Let.				Bep.
1 2			61 62	43.1 43.8	43.1 43.8	121 22	85.6 8 6.3	85.6 86.3	181	128. 0 128.7	128.7			170.4 171.1
3			63	44.5		23	87.0		83	129.4	129.4			171.8
4	02.8	02.8	64	45.3		24	87.7	87.7	84	130.1	130.1			172.5
5	03.5		65	46.0			88.4		85	130.8	150.8			173.2
6	04.2		66	46.7	46.7		89.1			131.5				173.9
8	04.9 05.7		67 68	47.4	47.4 48.1	27 28	89.8 90.5			132.2 132.9				174.7
1 9	06.4		69	48.8		29	91.2		189	133.6	133.6			176.1
10	07.1		70	49.5		80	91.9			134.4				176.8
111	07.8	07.8	71	50.2	50.2	131	92.6	92.6	191	135.1	135.1	251	177.1	177.5
12			72	50.9		82				135.B				178.2
13	08.5 09.2		73	51.6	51.6	83	93.3 94.0		93	136.5	136.5			178.9
14	09.9		74	52.3		34	94.8		94	137.2	157.2			179.6
15	10.6			53.0		35	95.5			137.9				180.3
16	11.3 12.0		76	53.7 54.4	53.7 54.4	36 37	96.2 96.9	96.2 96.9			138.6 139.3			0 181.0 7 181.7
18	12.7		78	55.2	55.2	38	97.6	97.6		140.0				182.4
19	13.4		79	55.9	55.9	39	98.3				140.7			183.1
20	14.1	14.1	80	56.6	56.6	40	99.0	99.0		141.4			183.	B 183.8
21	14.8		81	57.3	57.3	141	99.7	99.7	201	142.1	142.1	261	184.0	6 184.6
22	15.6		82	58.0	58.0	n		100.4		142.8				185.3
23	16.3		83	58.7	58.7			101.1			143.5			0 986
24 25	17.0 17.7		84	59.4				101.8		144.2				1186.7
26	18.4		85 86	60.1 60.8				102.5 103.2		145.7	145.0	00	100	1 187 .4 1 188 . 1
27	19.1		87	61.5	61.5	47		103.9			146.4			188.8
28	19.8	19.8	88	62.2	62.2			104.7			147.1			189.5
29	20.5		89	62.9	62.9			105.4			147.8			190.2
30	21.2		90	63.6	63 .6	50	106.1		1	148.5				190.9
31	21.9		91	64.3	64.3	151		106.8		149.2				6 191 .6
32	22.6		92	65.1	65.1			107.5		149.9				3 192.3
33 34	23.3 24.0		93 94	65.8 66.5	65.8 66.5			108.2 108.9			150.6 151.3			0 193.0 7 193.7
35	24.7	24.7	95	67.2	67.2			109.6			152.0			194.5
36	25.5			67.9	67.9			110.3			152.7	76	195	195.2
37	26.2			68.6	68.6			111.0			153.4	77	195	9 195.9
38	26.9			69.3				111.7			154.1			6 196.6
39 40	27.6 28.3		99 100	70.0 70.7	70.0 70.7		112.4	112.4		155.6	154.9			3 197.3 D 198.0
1	29.0			_	71.4	il	_	113.8				L	_	
41	29.0		101	71.4 72.1	72.1			114.6			156.3 157.0			198.7 199.4
43	30.4			72.8		63	115.3	115.3		157.7				1 200 . 1
44	31.1			73.5				116.0	•		158.4			8 200 .8
45	31.8			74.2				116.7			159.1			5 201.5
46	32.5			75.0	75.0			117.4			159.8			2 202.2
47	33.2			75.7	75.7 76.4			118.1 118.8			160.5			9 202.9
48	33.9 34.6		88	76.4 77.1	77.1			119.5			161.2 161.9		204	6 203.6 4 204.4
50	35.4		10	77.8	77.8			120.2			162.6			205.1
51	36.1		111	78.5	78.5	تمنحا		120.9			168.3			206.8
52	36.8			79.2	79.2			121.6			164.0			206.5
53	37.5	37.5	13	79.9	79.9	73	122.3	122.3	33	164.8	164.8	93	207.	207.2
54	38.2			80.6	80.6	74	123.0	123.0	34	165.5				207.9
55	38.9		15	81.3				123.7		166.2				208.6
56 57	39.6 40.3			82.0 82.7	82.0 82.7			124.5 125.2		166.9 167.6				3209.3 210.0
58	41.0		17 18	83.4	83.4	78		125.9		168.3				210.7
59	41.7		19	84.1	84.1	79		126.6		169.0		99	211.4	211.4
60	42.4		20	84.9	84.9	80	127.3	127.3	40	169.7	169.7			212.1
Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.	Dist.	Dep.	Lat.
														egrees.

TABLE III.

· MERIDIONAL PARTS.

M	<u>.</u>	0d. j	1d.	2d.	3d.	4d.	5d.	6d.	7d.	8d.	9d.	10d.	ild.	12d.	13d.	M.	ī
	<u>, </u>	0	60	120	180	240	300	361	421	482	542	603	664	725	787	-0-	l
	1	1	61	121	181	241	301	362	422	483		604	665		788	1	ĺ
	2	2	62	122	182	242	302	363	423	484	544	605	666	727	789	2	l
. 1	3	3	63	123	183	243	303	364	424	485	545	606	667	728	790	3	
. 4	•	4	64	124	184	244	304	365	425	486		607	668	729	791	4	ļ
- (5	5	65	125	185	. 245	305	366	426	487	547	608	669	730	792	5	:
6		6	66	126	186	246	306	367	427	488	548	609	670	731	793	6	1
7		7	67	127	187	247	307	368	428	489	549	610	671	732	794	. 7	l
8		8	68	128	188	248	308	369	429	490	550	611	672	734	795	8	ı
9	<u>_</u>	9	69	129	189	249	309	370	430	491	551	612	673	735	796	_9	ı
10)	10	70	130	190	250	310	371	431	492	552	613	674	736	797	10	ı
11		11	71	131	191	251	311	372	432	493	553	614	675	737	798	11	
15		12	72	132	192	252	312	373	433	494	554		676	738	799	12	ı
13		13	73	133	193	253	313	374	434	495	555		677	739	800	13	l
14	_	14	74	134	194	254	314	375	435	496	556	617		740	801	14	ĺ
18	5	15	75	135	195	255	315	376	436	497	557	618	679	741	802	15	l
16		16	76	136	196	256	316	377	437	498	558		680	742	803	16	Į
17		17	77	137	197	257	317	378	438	499	559		681	743	804	17	i
18		. 18	78	138	198	258	318	379	439	500			682		805 806	18 19	l
19		19	79	139	199	259	319	380	440	501	561	622	683	745			ļ
20		20	80	140	200	260	320	381	441	502	562	623	684	746	807	20	ı
21		21	81	141	201	261	321	382	442	503			685	747	808	21	ı
22		22	82	142	202	262	322	383		504			687	748 749	809 810	22 23	1
23		23	83	143	203		323	384	444	505	566		688	750		23 24	١
24	_	24	84	144	204	264	324	385	445	506	567	627	689		811		١
. 25	- 1	25	85	145	205	265	325	386	446	507	568	628	690	751	812	25	١
26		26	86	146	206	266	326	387	447	508	569		691	752	813		١
27		27	87	147	207		327	388	448	509			692		815 816	27 28	ļ
28		28	88	148	208 209		328	389	449 450	510	571 572	632 633			817	28	١
25		29	89	149			330	390		511							ŀ
30		30	90	150	210	270	331	391	451	512					818	30	1
31		31	91	151	211 212	271	332	392	452 453					757	819 820	31 32	١
39		32	92	152 153	213	272 273	333 334	393 394	454	514 515		636 637	697 698		821	33	I
33		33 34	93 94	154	214	274	335	395	455	516		638	699			34 34	1
34		;										!		761	823		ŀ
31		35	95	155	215	275	336	396	456 457	517	578		700	761	1	35 36	١
36		36 37	96	156 157	216 217	276 277	337 338	397 398	.458	518 519	579 580		701 702			37	I
37 38		38	97 98	158	218	278	339	399	459	519 520		642		764	826	38	I
39		39	99	159	219	279	340	400	460	521			704	765	827	39	١
					220				461			644	705		828	40	H
40		40	100	160 161	221	280 281	341 342	401	462	522 523					829	41	١
41		41	101	162	222	282	343	403	463	524			707	768		42	١
45		43	103	163	223	283	344	404	464	525			708		831	43	١
44		44	104	164	224	284	345	405	465	526	587	648	709	770	832	44	١
	_	45		165	225	285	346	406	466	527	588	649	710	771	833	45	4
45 46		45 46	105 106	166	225 226	285 286	347	407	467	528			711	772	834	46	Į
47		47	107	167	227	287	348	408	468				712		835	47	١
48		48	108	168	228		349	409	469	530	591	652	713		836	48	1
49		49	109	169	229	289	350	410	470	531	592		714		837	49	Į
- 50	_	50	110	170	230	290	351	411	471	532	593		715		838	50	1
51		51	111	171	231	290 291	352	412	472	533			716		839	51	1
59		52	112	172	232		353	413		534			717		840	52	j
53		53	113	173	233	293	354	414					718			53	1
54		54	114	174	234	294	355	415	476	536	597				842	54	ı
	_		115	175	235	295	356	416	477	537	598		720	782	843	55	1
56 56		55 56	116		236	296 296	357	417	478					783	844	56	1
57		57	117	177	237	290	358	418	479	539			722	784	845	57	Į
58		58	118	178	238			419	480			662			846	58	1
		591	119	179	239		360	420	481		602				847	59	Į
59	9 1																_

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TABLE III.

MERIDIONAL PARTS.

	M. I	14d.	15d.	16d.	17d.	18d.	19d.	20d.;	21d.j	22d.	234	24d.	25d.	26d.	27d.i	М.	_
ŀ	0	848	910	973	1035	1098	1161		1289	1354					1684	0	ŀ
1	1	850	911	974	36	99	63	26	90	55	20	85	51	18	85	1	1
- 1	2	851	913	975	37	1100	64	27	91	56	21	86	52	19	86	2	ļ
- 1	3	852 853	914	976	3 8 3 9	01	65	28 29	92 93	57	22 23	87 88	53	20 21	87	3	ı
-].	4		915	977		02	66			58			54		88	4	
١ ١	5	854 855	916 917	978 979	1041 42	1103	1167 68	1230 32	1295 96	1359 60	1424 25	1 49 0 91	1556 57	1622 23	1689 90	5 6	ı
- 1	6 7	856	918	980	43	06	69	33	97	61	26	92	58	24	91	7	
1	8	857	919	981	44	07	70	34	98	62	27	93	59	25	93	8	ı
1	9	858	920	982	45	08	71	35	99	63	28	94	60	26	94	9	l
ľ	10	859	921	983	1046	1109	1172	1236	1300	1364	1430	1495	1561	1628	1695	10	ı
ı	11	860	922	984	47	10	73	37	01	66	31	96	62	29	96	11	ı
- 1	12	861	923	985	48	11	74	38	02	67	32	97	63	30	97	12	ļ
- 1	13 14	862 863	924 925	986 987	49 50	12 13	75 76	39 40	03 04	68 69	33 34	98 99	64 65	31 32	98 99	13 14	1
ŀ		864	926	988		1114	1177		1305	1370	1435	1500		1633	1700	15	ł
I	15 16	865	927	989:	52	15	78	1241	06	71	36	02	68	34	01	16	l
ı	17	866	923	990	53	16	79	43	07	72	37	03	69	36	03	17	١
]	18	867	929	991	54	17	81	44	08	73	38	4 04	70	37	04	18	١
l	19	868	930	993	55	18	82	45	10	74	39	05		38	05	19	1
ſ	20	869	931	994		1119	1183	1246	1311	1375	1440	1506		1639	1706	20	1
- 1	21	870	932	995	57	20	84	48	12	76	41	07		40	07	21	1
- 1	22 23	871 872	933 934	996 997	58 59	21 22	85 86	49 50	13 14	77 79	43 44	08 09		41 42	08	22 23	l
ı	24	873	935	998	60	23	87	51	15		45	10		43	11	24	ļ
1	25	874	936	999	1061	1125	1188		1516		1446	1511		1644	1712	25	i
- 1	26	875	937	1000	63	26	89	53		82	47	13		45	13		l
- 1	27	876	938	01	64	27	90	54	18	83	48	14		47	14	27	1
- 1	28	877	939	02	65	28	91	55	19	84	49	15		48	15	28	I
- 1	29	878	941	03	66	29	92	56	20	85	_ 50	16	82	49			l
- 1	30	879	942	1004	1067	1130		1257		1386	1451			1650			ı
- 1	31	880		05	68	31	94	58	22	87	52			51	18		١
- 1	32 33	882 883		06 07	69 70	32 33	95 96	59 60	24 25	88 89	53 55			51 53			١
- 1	34	884	946	08	71	34	98	61		90	56			54	22		١
ŀ	35	885	947	1009	1072	1135	1199			1392	1457	1522		1656	1723	35	1
- 1	36	886	948	10	73			64	28	93				57		36	ı
ı	37	887	949	11	74	37	01	65		94	59			58			ı
- 1	38	888	950		75		02	66		,	60			59			ı
- 1	39	889	951	13	76			67	31		61			60		39	1
1	40	890												1661			۱
ı	41 42	891 892	953 954	15 16				69 70						62			١
l	43	893												1 -			1
ł	44	894	956							1							
	45	895	957	1020			1209	1273	1338	1402	1466	1533	1600	1667	1734	45	1
l	46	896	958		84	47	10							68			Ì
ļ	47	897	959					75	40								١
	48	898			86												١
ļ	49	899		24					-						┺		4
	50 51	900															١
	52	902															ı
1	53	903													7	53	1
	54	904											10	7	44	54	J
	55	905		1030	1093	1156	1220	1281	1134	1413	1479	154					1
	56	906															١
	<i>5</i> 7	907									81						١
- 1	58 59	908															1
	M.	-				1	1			220		-			_		-
!	<i>5</i> 71.	: 1 5 0.	. 15d	·l rog	. ı/d	1 12Q	. 19d	المالاد ر.	'i 21 a	.: 220	. z30	. 290	-1 ×00	1 40Q	.1210	≠ M.	J

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TABLE III.

MERIDIONAL PARTS.

	M.	28d	. 1 290	. 50d	. 31d.	52d.	33d.	34d.	35d	. 36d	37d.	38d.	39d.	40d.	41d.	M.	ī
	0	175	1/181	1888	1958	2028	2100	2171	2244	2318	2393	2468	2545	2623	2702	0	1
	li	5											46	24	03		Ì
	2	5			60	31	01	74	47	20	95	71	48	25	04	2	ı
	3	5												27	06	3	l
	4	. 50	5 24	H 93	63	33	04	76	49	23	98	73	50	28	07	4	I
	5	176	182	1894	1964	2034	2105	2178	2250	2324		2475	2551	2629	2708	5.	۱
	6	54		98	65	35	07			25	2400	76	53	31	10	6	ı
	7	5									01		54	32	11	7	ı
	8	. 60												33	12		l
	9	6			69	39	10	89	5.5					34	14	9	ļ
	.10	176			1970	2040	2111				2405			2636	2715	10	١
	11	64													16	11	۱
	12	64													18	12	l
	13	60													19	13	l
	14	67							ú					-	20	14	l
	15	1768		1-1	1	2046		2190				2487		2642	2722	15	۱
	16	69											1		23	16	I
	17	70													24	17	į
	18 19	73												46 48	. 26 27	18	l
			1				·									19	١
	20	1774					2123				24 18 19			2649		20	١
	21 22	76				53 54	25								29 31	21 22	ı
	23	1 77				56	26 27						75	53	31	23	i
	24	78				57		2200				99			33	24	ı
		1780			·												l
	25 26	81	,		1987 88	205 8 59	2129				2424 25		2577 78	2655 57	2736 36	25 26	ı
i	27	82				60	31 32	03 04			27	03		58	37	27	ł
Ì	28	83				61	33				28	04	81	59	39	28	ı
	29	84				63	34	07			29	05	-82	61	40	29	l
	30	1785	1854	1923		2064	2135		2281					2662	2742	30	l
- 1	31	86		,		65	2133 37	09		56	32	08	85	63	43	31	İ
- 1	32	87				66	38	10		58	33	09	86	65	44	32	İ
- 1	33	89		27	97	67	39	11	85	59	34	10	88	66	46	53	l
-	34	90		28	98	69	40	13		60	35	12	89	67	47	34	ĺ
- 1	35	1791	1860	1929	1999	20 70	2141	2214	2287	2361	2437	2613	2590	2669	2748	35	ĺ
ı	36	92	61		2000	71	43	15		63	38	14	91	70	50	36	ı
- [37	93	62	31	01	72	44	16	90	64	39	15	93	71	51	37	l
- {	38	94	63	32	02	73	45	17	91	65	40	17	94	73	52	38	ĺ
- 1	39	95	64	34	04	75	46	19	92	66	42	18	95	74	54	39	ĺ
ŀ	40	1797	1865	1935	2005	2076	2147	2220	2293	2368	2443	2519	2597	2675	2765	40	l
- 1	41 .	98	66	36	106	77	49	21	95	69	44	21	98	76	56	41	ı
- 1	42	99	68	37	07	78	50	22	96	70	45	22	99	78	58	42	ı
- {	43	1800	69	38	98	79	51	24	97	71	47	23	2601	79	59	43	ĺ
J.	44	01	70	39	10	80	52	25	98	75	48	24	02	80	60	44	l
Ţ	45	1802	1871	1941	2011	2082	2153			2374	2449	2526	2603	2682	2762	45	
Į	46	03	72	42	12	83	55	27	23 01	75	51	27	04	83	63	46	i
1	47	05	73	43	13	84	56	28	02	76	52	28	06	84	64	47	
- [48	06	75	44	!4	85	57	30	03	78	53	30	07	86	66	48	
1	49	07	76	45	15	86	58	31	04	79	54	31	08	87	67	49	
- 1	50	1808			2017						456				2768	50	
- 1	51	09	78	48	18	89	61	33	07	81	57	33	11	90	70	51	
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-	54	13	81	51	21	92	64	37	11	85	61	37	15	94	74	54	
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TABLE III.

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2 84 66 49 33 18 06 95 85 78 72 62 67 68 3 86 67 50 34 20 07 96 87 79 74 70 69 70 69 88 81 75 72 70 71 71 5 2783(2870) 2953 3037 3122 3210 3299 3390 3482 3577 3673 3772 3873 397 6 90 71 54 38 24 12 3301 91 84 78 75 74 75 74 75 74 75 74 75 74 75 77 78 89 74 57 41 27 14 03 94 87 82 78 77 77 77 77 78 79 94 75 53 43 29 16 05 96 88 83 80 79 80 10 2795 2877 2960 3044	1 2 3 3 4 5 5 4 6 7 7 5 8 6 7 8 2 8 9 9 15 10 12 14 15 16 16 16 17 19 18 19 19 19 19 19 19 19 19 19 19 19 19 19
3 86 67 50 34 20 07 96 87 79 74 70 69 70 4 87 69 51 36 21 09 98 88 81 75 72 70 71 5 2783 (2870) (2953) (3037) (3123) (3210) (3299) (3390) (3482) (3577) (3673) (3772) (3873)	3 3 4 7 5 8 8 7 7 5 8 8 9 5 10 7 11 12 14 15 16 16 16 17 19 18 11 19
4 87 69 51 36 21 09 98 88 81 75 72 70 71 5 2783 2870 2953 3037 3123 3210 3299 3390 3482 3577 3673 3772 3873 39 6 90 71 54 38 24 12 3301 91 84 78 75 74 75 7 91 73 56 40 26 13 30 94 87 80 77 75 74 75 8 92 74 57 41 27 14 03 94 87 82 78 77 75 77 78 77 78 77 78 77 78 77 78 77 78 78 79 80 10 2795 2877 2960 3044 5130 3217 3306 3397 3490 3585 3681 3780 3882 39 11 97 78 61 46	5 4 7 5 8 6 0 7 2 8 9 5 10 7 11 19 12 11 13 12 14 14 15 16 16 17 19 18
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TABLE III.

MERIDIONAL PARTS.

M.	156d.	57d.	58d.	59d.	60d.	61d.	62d.	63d.	64d.	65d.	66d.	67d.	68d.	69 d .	M.	7
0	4074	4183			4527	4649			5039		5324		5631	5795	0	1
ĭ	76	84	96	11	29	51	77	07	42	81	26	77		97	ĭ	İ
2	77	86	98	13	31	53	79	09	44	84	28	79		5800	2	l
3	79		4300	15	33	55	81	12		86	31	82		03	3	l
4	81	90	02	17	35	_ 57	84	14	49	88	33	84	42	06	4	
- 5	4083	4192	4304	4419	4537	4660	4786	4916		5191	5336	5487		5809	5	ı
6	85	94	06	21	39	62	88	18	53	93	38	89		11	6	ı
7	86	95	08	23	41	64	90	20	55	95	41	92		14	7	ı
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13	97	07	19	34	53	76	03	34	69	10		07	66	31	13	ł
14	99		21	36	55	78	05	36	71	12		10		34	14	ł
15	4101	4210	4323	4438	4557	4680	4807	4938	5074	5214	5361	5513	5671	5837	15	١.
16	03		25	40	59	82	09	40	76	17		15		39	16	١
17	04		27	42	62	84	11	43	78	19		18	76	42	17	1
18	06	16	28	44	64	87	14	45	81	22	68	20	79	45	18	ı
19	08	18	30	46	66	89	16	47	83	24	71	23	82	48	19	ı
20	4110	4220	4332	4448	4568	4691	4818	4949	5085	5226	5373	5526	5685	5851	20	ı
21	12	21	34	50	70	93		51	88	29	76	28		54	21	1
22	13	23	36	52		95	22	54	90	31	78	31	90	56	22	ı
23	15	25	38	54	74	97	24	56	92	34	80	33	93	59	23	ı
24	17	27	40	56	76	99	26	58		36		3 6	95	62	24	
25	4119	4229		4458		4701	4829	4960		5238	5385	5539			25	l
26	21	31	44	60		03	31	63	99	41	88		5701	68	26	ł
27	22	32	46	62		05	33	65		43	90	44	04	71	27	ı
28	24 26	34 36	47 49	64		07 10	35 37	67 69	04	46 48	93 95	46 49	96 09	74 76	28 29	ı
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30	4128		4351	4468		4712 14		4972 74			5398				30	ı
31 3 2	30 32	40 42	53 55	70 72		16	42 44	76		55	5401 03	54 57	15 17	82 85	31 32	l
33	33	44	57	74	94	18	46	78		58	06	59	20	88	33	1
34	35	46	59	76	96	20	48	81		60	08	62	23	91	34	l
35	4137	4247	4361	4478	4598	4722	4850	4983	5120	5263	5411	5868	5725	5894	35	1
36	39	49	63		4600	24	52	85	22	65	13	67	28	96	36	l
37	41	51	65	82	02	26	55	87	25	67	16	70	31	99	57	ĺ
38	42	53	67	84	04	28	57	90	27	70	18	73	34	5902	38	ı
39	44	55	69	86	06	31	59	92	29	72	21	75	36	05	39	ł
40	4146	4257	4370	4488	4608	4733	4861	4994	5132	5275	5423	5578	5739	5908	40	1
41	48	59	72	90	10	35	63	96	34	77	26	80	42	11	41	١
42	50	60	74	92	12	37	65	99	36	80	28	83	45	14	42	ı
43	52	62	76	94	14	39	68	5001	39	83	31	86	47	17	43	ı
44	53	64	78	95	16	41	70	03	41	84	33	88	50	19	44	ı
45	4155	4266	4380	4497	4618	4743	4872	5005	5143	5287	5436	5591	5753	5922	45	1
46	57	68	82	99	20	45	74	08	46	89	38	94	56	25	46	1
47 48	59 61	70 72	84	4501 03	23 25	47 50	76 79	10 12	48	92	41 43	96 99	58	28 31	47 48	ı
49	62	74	86 88	05	25 27	52	81	14	51 53	94 97	46	560 2	61 64	34	49	1
50	4164	4275	4390	4507	4629	4754	4883	5017	_	5299	5448	5604	5767		50	
50 51	66	77	4390 92	09	31	56	9663 85	19	5155 58	5301	51	07	70	59 37 40	51	
52	68	79	94	11	33	58	87	21	60	04	54	10	70	43	52	ı
53	70	81	96	13	35	60	90	23	62	06	56	12	75	46	53	ı
54	72	83	98	15	37	62	92	26	65	09	59	15	78	48	54	ı
55	4173	4285	4399	4517	4639	4764	4894	5028	5167	5311	5461	5617	5781	5951	55	ŀ
56	75		4401	19	41	66	96	30	69	14	64	20	83	54	56	ĺ
57	77	89	03	21	43	69	98	33	72	16	66	23	86	57	57	l
58	79	91	05	23	45	71	4901	35	74	19	69	25	89	60	58	ı
59	81	92	07	25	47	73	03	37	76	21	71	28	92	63	59	
M.	56d.	57d.	58d.	59d.	60d.	61d.	62d.	63d.	64d.	65d.	66d.	67d.	68d.	69d.	M.	<u>l</u>
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11	M.	70d.	71d.	72d.	73d.	74d.	75d.	76d.	77d.	78d.	79d.	80d.	81d.	8 2 d.	83d.	M.
2 72 52 41 41 41 53 78 18 76 54 56 87 52 60 22 3 75 55 44 54 55 78 92 22 81 69 61 93 58 67 31 4 73, 58 48 48 48 60 86 27 85 64 67 98 55 74 39 5 5981,61616531 6552 6764 6990 72317490 7769 8072 84048 771 91382 9647 7 86 67 58 58 71 97 39 98 78 83 16 84 96 64 8 89 70 61 62 75 701 43 7503 83 88 22 91 9205 72 9 92 73 64 65 79 05 47 07 88 93 27 97 11 80 10 5995,6177 6567 6569 6782 7009 7257 712 7793 8099 8433 8948 9818 9689 11 1 93 80 71 72 86 13 66 16 981804 39 10 25 97 11 60 183 74 76 90 17 60 21 7803 09 45 17 33 9706 11 98 80 71 72 86 13 66 16 981804 39 10 25 97 11 60 183 74 76 90 17 60 21 7803 09 45 17 33 9706 11 98 80 71 72 86 13 66 16 981804 39 10 25 97 11 60 183 74 76 90 17 60 21 7803 09 45 17 33 9706 11 98 80 71 72 86 13 66 16 80 81804 39 10 25 97 11 1 90 11 93 80 71 72 86 13 64 25 08 16 11 23 40 14 14 07 89 80 83 97 25 68 30 13 20 57 30 48 23 11 66 13 95 87 90 04 33 77 39 22 31 64 25 08 16 12 23 40 14 14 14 07 89 80 88 83 97 25 68 30 13 20 57 30 48 23 11 66 13 95 87 90 04 33 77 39 22 31 68 26 45 80 85 85 85 85 85 85 85 85 85 85 85 85 85								7210			8046	8375	8739	9145	9606	0
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51 18. 06 04 13 36 73 27 7702 99 24 82 81 33 10052 45 52 21 09 07 17 40 77 32 06 8004 29 88 88 41 10061 43 53 24 12 11 20 43 81 36 11 09 30 95 96 49 10071 40 54 27; 15 14 24 78 41 16 14 187019103 57 10080 57 55 61306319651767286951789744577219020334787079110956510089 56 56 33 22 21 31 55 94 49 25 55 52 14 17 7310099 57 57 36 25 24 35 59 98 54 30 30 58 20 24 8110108 58 40 28 28 38 63 7202 58 35 35 64 26 31 89 10118																
52 21 09 07 17 40 77 32 06 8004 29 88 88 41 10061 45 53 24 12 11 20 43 81 36 11 09 36 95 96 49 10071 4 54 27 16 14 24 47 85 41 16 14 41 8701 9103 57 10080 3 55 6130/6319/6517/6728/6951 7189/7445 7721 8020/8347/8707 9110/9565 10089 4 25 55 52 14 17 75/10080 57 50089 56 50 75/10080 57 50089 4 8110108 4 49 25 55 52 14 17 75/10080 58 56 24 8110108 4 8 58 40 28 28 38 63/7202 58 35 35 54<						,								1		50 K1
53																51 52
54 27; 15 14 24 47 85 41 16 14 4187019103 5710080 5 55 6130 6319 6517 6728 6951 7189 7445 7721 9020 834787079110956510089 6 53 22 21 31 55 94 49 25 25 52 14 17 73 10099 6 57 36 25 24 35 59 98 54 30 30 58 20 24 81 110108 8 58 40 28 28 38 63 7202 58 35 35 64 26 31 89 10118								1						1		53
55 6130 6319 6517 6728 6951 7189 7445 7721 9020 3347 8707 9110 9565 10089 656 33 22 21 31 55 94 49 25 25 52 14 17 7310099 657 36 25 24 35 59 98 54 30 30 58 20 24 81 10108 658 40 28 28 38 63 7202 58 35 35 64 26 31 89 10118							•									54
56 33 22 21 31 55 94 49 25 25 52 14 17 73 10099 6 57 36 25 24 35 59 98 54 30 30 58 20 24 81 10108 6 58 40 28 28 38 63 7202 58 35 35 64 26 31 89 10118																
57 36 25 24 35 59 98 54 30 30 58 20 24 81 10108 4 58 40 28 28 38 63 7202 58 35 35 64 26 31 89 10118																55 56
58 40 28 28 38 63 7202 58 35 35 64 26 31 89 101 18																57
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M. 70d. 71d. 72d. 73d. 74d. 75d. 76d. 77d. 78d. 79d. 80d. 81d. 82d. 83d. 1			-													M.

FOR THE YEAR 1824,

W	HICH										328, 1832,	1836.
ú	Jan		March							Oot.	Nev. D	ec. (0)
DAYS.	South.	South.	South.			North-			North.	South.	Bouth Son	uthJ≫
ā	C /	0 /	0 /	0 /	0 /	0 '	0 /	0. 1	0 /	0 /	0 / 0	7 5
	23. 5						23. 7				14.3121.	
	23.00						23. 3				14.50 22.	
	22.55						22.58				15. 9 22.	
	22.49						22.53				15.2822	
5	22.42	16. 8	5.57	6.10	16.19	22.35	22.48	16.57	6.45	4.49	15.4622	. 26 5
	22.36						22.42				16. 422.	
7	22.29	15.32	5.10				22.36				16.22 22.	
8	22.21	15.13	4.47				22.29				16.4022.	
	22.13						22.22				16.57 22.	
10	22. 5	14.35	4.00	8. 2	17.41	23. 3	22.14	15.32	4.53	6.44	17.14 22.	.58 10
11	21.56	14.15	3.36	8.24	17.56	23. 7	22. 7	15.15	4.30	7. 6	17.31 23.	311
12	21.46	13.56	3.13	8.46	19.12	23.11	21.58	14.57	4. 7	7.29	17.47 23.	712
13	21.37	13.36	2.49	9. 8	18.27	23.14	21.50	14.38	3.44	7.52	19. 323.	11 13
	21.26			9.30	18.41	23.18	21.41	14.20	3.21	8.14	18.1923.	1514
15	21.16	12.55	2. 2	9.51	19.55	23.20	21 . 32	14. 1	2.58	8.36	18.3423.	18 15
16	21. 5	12.35	1.39	10.12	19. 9	23.23	21.22	13.42	2.35	8.58	18.4923.	2116
17	20.54	12.14	1.15	10.33	19.23	23.24	21.12	13.23	2.11	9.21	19. 423.	23 17
15	20.42	11.53	0.51	10.54	19.36	23.26	21. 1	13. 4	1.48	9.42	19.1823.	25 18
	20.30						20.51				19.3323	
20	20.17	11.11	0. 4S	11.36	20. 2	23.28	20.39	12.25	1. 1	10.26	19.46 23.	. 27 20
21	20. 4	10.49	0.20N	11.56	20.14	23.28	20.28	12. 5	0.38	10.47	20.0023	2921
22	19.51	10.27	0.44	12.16	20.26	23.29	20.16	11.44	0.15N	11. 9	20.13 23	. 28 22
23	19.37	10. 6	1. 8							11.30	20.2523.	. 27 23
24	19.23	9.44	1.31	12.56	20.49	23.26	19.52	11. 4	0.32	11.51	20.37 23	. 26 24
2:	19. 9	9.21	1.55	13.16	21.00	23.25	19.39	10.43	0.56	12.12	20.4923	. 25 25
20	18.54	8.59	2.18	13.35	21.10	23.23	19.26	10.22	1.19		21. 123	
27	18.39	8.37	2.42				19.12				21.1223	
	15.23		3. 5				3 18.58				21 . 23 23	
	19. 8		3.29	14.32	21.40	23.1	18.44	9.19	2.29		21.3323	
30	17.51	1	3.52	14.51	21.49	23.11	18.30	8.57	2.53	13.59	21 . 43 23	.10 30
3	17.35		4.15		21.57	1	18.15	9.35		14.15	23	. 631

TABLE IV. A.—THE EQUATION OF TIME

FOR THE YEAR 1894,

WHICH WILL ANSWER NEARLY FOR THE YEARS 1832, 1832, 1836.

	AUCI	J 44 1	THE WIL			ARLI	run	1111	ILA	LPD 103	80, IO:) z, 100	∙0.
	Jan.	Fch.	March.	April.	May.	June.	July.	Aug.	Sept.	Oct.	Nov.	Dec.	,
9î ≥	Add to	Add to	Add to	Aild to	Sub. fr.	oub fr.	Add to	Add to	Sub. fr.	Sub. fr.	Sub.fr.	Sub.fr.	e X
DA	ιρ. tim	ap. tim	ap. tim	ap. ilm	ւթ. նա	ap. tim	ap. tim	ap. tim	ap, tim	ap. tim	ap. tim	ap. tim	₹
Ω.	M. S.	M. S.	M. S.	M. 8	M. 5.	M.S.	M.S.	M. S.	M. 8.	M. B.	M. S.	M. S.	4
	3.35	13.52	12.36	8.56	3, 5	2.33	3.25	5.58	0.12	10.28	16.15	10.57	7
2	4 1	14. 1		3.37	3.12	2.24	3.37			10.41			
3	4.32			3.19	3.19	2.14	8.48						3
4	5.0				3.25		3.59			11 18			
5	5.27	14.20	11.44	2.44	8.31	1.54	4. 9	5.39	1.30	11.36	16.14	9. 2	5
c	5.54	14.25	11.30		3.36	1.44	4.19		1.50	11.54	16.11	8.36	6
7	6.21	14.29			5.41	1.33	4.29						7
a	6.47	14.52			3.45	1.22	4,39			12.28			
9	7.18		10.45		3.48		4.48			12.44			
10	7.38			1.18	3,51	0.59	4.56	5. 2	3.12	12.00	15.53		
11	8. 9	14.36			3-58	0.47	5. 4		3.32	13.15	15.46		111
12	9,26					0.85	5.12						
18							5.19						
14		14.34		0.14			5.26						
15	9.38		9. 6	sub0. 2	3.57	a:ld0.02	5.32		4.56	14,14	15.10		
11	9.54		8.48		3.56	0.15	5.34		5.17	14.24	14.59		
17	10.15				3.55	0.28	5.11						
18					3.54	0.40	5.49						
19	10.54	14.15			3.51	95	5.59						
21	11 12			1.12	3.48	1.6			6.41			·	
21	11.30			1.26	3 45	1.19	6.00					1.29	
21 25	11 46				3.41	1.82						0.58	22
25	12. 3				3.36								23
21	12.18				3.31	1.38							
25		ا			3.26	I	6. 8						
26	12.46		5.46				6. 8					1. 2	
27					3.18		6. 3						
27			5. 9				8. 7					2 1	
29 80	13.23		4.50 4.32										
31					2.50		I	0.44			المصحفات	1.00	J-1-1-1
31	134	,	4.13	4	2.42	4	6. 1	0. G	•	16.14	9	3.20	131 C

WHICH WILL ANSWER	NEARLY	FOR THE	YEARS	1329.	1833,	1937.

<u>w</u>	HICH			SWEE									
cċ	Jan.		Harch.				July.		Sept.	Oct.	Nov.	Bouth.	ø.
DAYS.	South.			North.	North.	North.				South.	South.	South.	2
2	0 /	0 /	, ,	0 /	0/	0 /	0,	0 1	3 /	0,	0 /	0 /	M
	32 1	17. 5	7 34	1 22	15 4	99 4	23. 8	18 4	9 10	2 10	14 97	21.50	
		16.48					23. 4					21.59	
		16.31					23.00					22. 8	
		16.13					22.55					22.16	_
		15.55					22.49					22.10	
									-				
		15.36					22.43					22.31	
		15.18					22.37					22.38	
		14.59					22.31					22.45	
		14.40					22.24					22.51	
10	21.58	14.20	4. 6	7.57	17.37	23. 1	22.16	15.37	4.58	6.38	17.10	22.56	10
11	21.49	14. 1	3.42	8.19	17.53	23. 6	22. 9	15.19	4.35	7. 1	17.27	23. 1	Πī
12	21.39	13.41	3.19	9.41	18. 8	23.10	22.00	15. 1	4.12	7.24	17.43	23. 6	12
13	21.29	13.21	2.55	9.3	18.23	23.14	21.52	14.43	3.49	7.46	17.59	23.10	13
14	21.18	13.00	2.31	9.24	18.37	23.17	21.43	14.24	3.26	8. 9	18.15	23.14	14
15	21. 8	12.40	2. 8				21.34			8.31	18.31	23.18	15
16	90 56	12.19	1 44	10 7	10 6	92 99	21.24	12 47	9 40	9 52	19 46	23.20	16
		11.58		10.28								23.23	
		11.37		10.49								23.25	
		11.16		11.10								23.26	
				11.31								23.27	
											1		
				11.51								23.28	
		10.11		12.12									
		9.49										23.27	
		9.27					19.55					23.26	
25	18.57	9. 5	1.49	13.11	20.57	23.2	19.42	10.4	0.50	12. 7	20.46	23.25	25
		8.49		13.81	21. 8	33.23	19.29	10.27	1.13	12.2	20.58	23.23	26
27	18.27	8.90		13.50	21.18	33.21	19.15	10. 6	1.37	12.4	21. 9	23.21	27
28	18.11	7.57	3.00	14. 9	21.2	23.18	19. 2	9.4	2.00	13. 8	21.90	23.18	28
29	17.55		3.23	14.28	21.37	23.1	18.48	9:24	2.24	13.2	21.30	23.15	129
30	17.39	i	3.46				18.33					23.11	
31	17.22	1	4. 9		21.55		18.19	8.4	1	14.		23.	' -
3.1	11.44	L	120)	<u>'</u>				1 2.41	<u>'L</u>	17-40		120.	1

TABLE IV. A.—THE EQUATION OF TIME FOR THE YEAR 1835, WHICH WILL ANSWER NEARLY FOR THE YEARS 1829, 1933, 1837.

	Jan.	Feb.	March.	April.	May ,	June.	July.	Aug.	Sept.	Og.	Non	Dec. 16
DAYS.	Add to	Add to	Add to	Add to	Sub. fr.	Sub. fr.						Sub. Sr
₹	ap. tim	ap. tim	ap. tim	np. tim	ap. tim	ap. tim	ep. cim	ap. tim	ap. tim	ap. time	ap. tim	at. tim 5
-	M. 8.	M. S.	M. S.	M. 8.	M. S.	M. S.	M. S.	M S.	M. S.	M. 8.	M. S.	
	3.57	15,58	12.50	8.59	3, 4	2.37	3.21	5.56	0. 9	10.18	16.15	10.43
				5,41	3.12	2.28	3.32	5.54	0 28	10.37		
2 3 4	4.56	14,12		3.23	3.19		3.48		0.47	10,56		9 56 3
4	5.90			3. 5		2.9						
- 6	6,13				8.34							
7	6.39						4.24					
9 10	7. 8						4.3				16, 4	7.49 8
9	7,30				8.48				2.46			7.52 9
10	7.55			1.21								
11	8,18		10.16		3.54	0.51	5. 1					
12 15	8.42	14.86				0.39			5.45			
15	9.5		9.46									
14	9.21	14.31		0.17								
15				. 4			-					4.83 15
16			8.52	sub0.15	3.5							
17	10.29		8.36									
18	10.49											
18	11.	14.10										
20	11.2				3.4	1. 5						I
21	11.4			1.2	3.4	1.10	5.5		6.5			
22	11.5		7. 4			2 1.25	6.	2.41			13.81	1. 5 22
25 24	12.1		6.4					2.2				
24	12,5	13.3	2 6.3				"	2,10				0. 524
25	1							7 1.5				
26	12.5						71	8 1.5				
27	13.						J	7 1.2				
28	13.2						vi :-	? 1.				1.54 28
29	15.5		4.5				'1 -	5 0.4				
90		4	4.9	8 2.5	8 2.5	5 8.			LHOUIZE	20 DV 3	10.23.23	
81	13.5	·	41	97	2.4	4	6.	1 0.1	01	16.1	, C	3.21 3 1

81 13.47

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FOR THE YEAR 1826,

77	нісн	wil	L AN	SWEI	FOR NEA	THE	FOR	1826, THE	YEAR	S 183	0. 183	4, 183	18.
_	Jen			April.			July.		Sept.	Oct.	Nov.	Dec.	<u> </u>
82	South.	South.	South.	North.	North.	North.	North.		North.	South.	South.	South.	2
DA	0 /	0 1	0 1	0 /	0 /	c /	0 /	0 /	0 /	0 1	0 7	0 /	4
1	23. 2	17.10	7.40	4.27	15.00	22. 2	23. 9	18. 7	8.24	3. 5	14.22	21.47	1
	22.57			4.50	15.18	22.10	23. 5	17.52	8. 3			21.57	
3	22.51	16.35	6.54				23. 1			3.51	15.00	22. 5	3
4	22.45	16.17	6.31				22.56					22.14	
' 5	22.39	15.59	6.8	5.59	16.11	22.32	22.50	17. 5	6.56	4.38	15.37	22.22	5
6	22.32	15.41	5.45	6.22	16.28	22.38	22.45	16.49	6.34	5. 1	15.56	22.29	.6
	22.25						22.39					22.36	
	22.17						22.32					22.43	
	22. 9						22.25					22.49	
	22.00						22.18			6.33	17. 6	22.55	10
	21.51						22.10					23.00	
	21,41						22. 2					23.5	
	21.31						21.54					23. 9	
	21.21						21.45					23.13	
	21.10						21.36				·1	23.17	-
	20.59						21.27					23.20	
	20.47						21.17					23.22	
	20.35						21. 6					23.24	
	20:23						20.56					23.26	
	20.10											23.27	
				11.46								23.27	
	19.44			12. 7								23.28	
	19.30											23.27	
	19.16											23.27	
	19. 1						19.45					23.25	
	18.46						19.32					23.24	
	18.31			13.45								23.21	
	18.15			14. 4								23.19	
	17.59			14.23								23.15	
_	17.43		3.40				18.37					23.12	
31	17.26		4. 4)	21.53		18.22	8.46		14. 3		23. 8	31

TABLE IV. A.—THE EQUATION OF TIME

POR THE YEAR 1826, WHICH WILL ANSWER NEARLY FOR THE YEARS 1830, 1834. Jon. Feb. March April. May. June. July. Aug. Sept. Oct. Nov. Dec. Add to Add to Add to Add to Sub. fr. Sub. fr. Add to Add to Sub. fr. Sub DATS. M. 8 M. 8 M. 8. M. S. M. 8. M. 8 M. S. M. 8 M. 8 M. S. M. S. 3. 3 3.11 5.18 5.24 5.30 5.58 5.55 5.51 5.46 5.41 0. 5 0.25 0.42 4. S 5.44 5.26 10.48 10.25 3.50 13.56 12.41 8.17 10.14 16.14 16.15 12.29 12.16 12. 3 4.18 4.46 5.14 14. 4 14.10 3,2 10.33 2.30 3.40 2.21 16.16 10. 1 3 9.37 14.16 3. 8 2.11 3.51 1. 1 16.15 16.14 2.51 1.21 8.41 14.22 11.50 2 1 4. 2 11.27 9.13 2.33 2.16 1.59 1.42 1.25 6. 8 6.34 7.00 7.25 14.26 14.30 14.32 14.34 8.36 4.18 5.35 1.40 2. 0 2.20 11.36 1.51 67 11.45 16.11 8.47 6 3.40 3.44 3.48 5.28 5.21 5.14 11.21 1.40 12. 2 16. 8 16. 4 16. 0 8.22 7.56 4.23 4.33 8 11. 7 10.51 12.19 12.35 4.42 2.41 3. 1 7.29 1.17 10 7.50 14.36 10.38 8.51 1. 6 5. b 12.51 15.54 7. 2 10 14.56 14.56 14.55 14.53 14.50 6.34 11 15.48 15.41 15.33 11112 8.14 10.20 1. 8 3.53 0.54 4.59 4.57 3.22 13, 7 10. 4 9.48 9.31 9.14 4.47 4.87 4.27 4.15 3.55 3.56 3.57 13.22 13.38 13.51 8,31 0.52 0.42 5. 8 5.15 3.43 6. 6 12 13 9. 1 9.23 9.44 5.22 5.29 14 0.31 0.17 4.25 15.24 5.1014 3.57 4.46 15 0. 6 0. 5 14. 4 15.14 4.41 15 14.26 14.22 14.17 14.12 14. 6 3.57 5.35 5.40 5.46 5.50 5.54 4. 4 3.52 3.59 3.26 5. 7 5.28 5.49 6.10 14,17 14,30 14,41 14,63 15, 3 8.56 8.39 10. č 16 sab0.10 10. 8 15. 4 4.12 16 3.57 3.56 3.55 3.53 3.50 10.26 10.45 14.40 14.40 14.37 3.43 17 3.13 18 0.24 0.21 8.21 0.38 18 0.34 8. S 7.45 0.52 2.48 19 2.14 20 11. 4 11.22 0.59 6.31 20 1. 6 3.12 14.13 13.58 13.51 13.43 13.34 13.24 7.27 7. 8 6.50 6.31 5.47 5.44 5.40 5.55 6.52 7.13 7.84 7.55 15.13 15.22 15.31 15.39 13.49 1.44 21 1.14 22 0.44 23 0.15 24 21 11.39 1.19 1.12 8.57 2.58 11.55 6. 0 6. 3 2.43 1.31 1.44 1.55 22 23 1.38 13 27 13. 9 24 25 12.26 1.51 2.13 12.40 6.18 2. 7 5.30 2. 3 8.18 15.46 12.51 1.57 40.17 26 27 28 5.54 5.35 5.17 4.58 1.41 1.24 1. 7 0.50 0.32 0.47 26 1.17 .7 1.46 28 12.53 13.14 2.17 3.24 2.16 6. 6 1233 8.36 15.52 2.28 2.37 2.46 6. 6 6. 6 13. 5 13. 4 12.52 S.18 S.11 2.29 8.56 15.58 12.13 13.17 13.28 9.16 9.36 9.56 2.41 2.58 16. S 16. 7 3. 4 2.16 29 2.45 30 11.32 30 13.38 4.39 2.86 3. 6 6. 8 16.10 11.10

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FOR THE TEAR 1827.

WHICH	WILL A	NSWER	NEARLY	FOR	ŤHE	YEARS	1931,	1825.
Jan Feb.	[March-]	April. May	June Ju	ly. 1 42	e. 1 S	not. Ont.	Non.	Den

_	Jan.	Feb.	March-	April.	May.	June	July.		Sept.		Nov.	Dea	_
DAYB.	South.		South.	North.	North.		North.		North.	South.		Honth	X E.
₹0	0 /	0 /	0,	0 1	0 /	0 /	0 1	0 7	0 /	0 /	0 /	0 /	γď
	-	10 14	F 45	4 31	14 56	20 00	22 10		2 20				
	23. 3						23.10					21.45	
	22.58						23. 6					21.54	
	22.53 22.47						23. 2 22.57					22.3 22.12	
	22.41						22.57					22.20	
	22.34						22.46					22.27	6
	22.26						22.40					22.35	
	22.19						22.34					23.41	
	22.11						22.27					22 .48	
	22. 2						23.20					22.53	
	21.53						22.12					22.59	
	21.44						22. 4					23. 4	
	21.34						21.56					23. 8	
	21.23						21.47					23.12	
-	81.13			9.35	18.45	23.18	21.38	14.15	3.15			23.16	_
16	21. 2	12.29	1.56	9.57	18.59	23.21	21.29	13.56	2.51	8.49	18.38	23.19	16
	20.50						21.19					23.22	
	20.38						21. 9					23.24	
	20.26						20.58					23.25	
20	20.13	11. 5	0.218	11.21	19.53	23.27	20.47	12.39	1.18	10.10	19.36	23.27	20
21	20.00	10.43	0. 3N	11.41	20. 5	23.29	20.36	12.19	0.55	10.32	19.50	23.27	21
	19.47						20.25					33.29	
	19.33						20.13						
	19.19						90. 00						
25	19. 5	9.15	1.38	13. 2	20.52	23.25	19.48	10.58	0.39	11.56	20.41	23.26	25
26	18.50	8.53	2. 1	13.21	91. 3	23.24	19.35	10.37	1. 2	12.17	20.52	23.24	26
27	18.34	8.31					19.22			12.38	21. 4	23.22	97
28	18.19	8. 8					19. 8			12.58	21.15	23.19	28
	18. 3			14.18	21.33	23.17	18.54	9.34	2.12			23.16	
30	17.47	1	3.35	14.37	21 . 42	23.14	18.40	9.13	2.36	13.38	21.35	23.13	30
31	17.30		3.58		21.51		18.26	8.51		13.58		23. 9	31

TABLE IV. A.—THE EQUATION OF TIME
FOR THE TEAR 1937,
WHICH WILL ANSWER NEARLY FOR THE YEARS 1831, 1835.

	WIL			VIIOA	A LT I) N 1 I		CARO		1039.	
-	Jan.		March.	April.		June.		Aug.	Sept.	Oct.	Nev.	Dec.	Ŀ
뼕	Add to	Add to	Add to	Add to	Suh. fr.	Sub. fr.	Add to	Add to	Add to	Sub. fr.	Bab. ir.	Sub. fr.	2
DA	ap. tim	ep. tim	up. tim	ap. tim	ap. tim	ap. tim	ap. tim	ap. tim	ap. tinı	ap. tim	ap. tim	ap. tim	
-	M. 8	M. S.	M. S.	M. 8.	M. B.	M. B.	M. S.	M. 8.	M. S.	M. 8.	M. 6.	M. S.	P
	3.43		12.44	4. 9			3.17	6. 1	0. 1	10. 9			<u> -</u>
÷	411				3. 7		3.29		emb0.18	10.28	16.16		1 2
2	4.40		12.20	3.32	314				0.37				ŝ
4	5.7				3.20		3.51	5.49	0.56				14
8	5.34	14.21	11.54	2.56	8.26	2.1	4. 2	5.43					15
6	6. 1	14.25	11.40	2,39	5.32	1.57	4.12	5.38	1.35	11.41	16.13	8.56	8
7	6.28				3.37	1.40	4.22	5.31	1.58	11.59	16.11	8.30	7
8	6.53					1.29	4.32						
9	7.19						4.41	5.16				7.37	9
10	7-43							5. 8			15.57	7.10	
711	8. 8					0.55	4.59		3.17		15.51	6.42	11
12	8.31						5.7		5.38	18.19			12
13	8.54		9.51	0.41			8.14		3.69		15.36	5.46	
14							5.21	4.30	4.20	18.48		5.18	
15		14.29	9.18	0. 9	3.57	0. 6	5.28	4.19		14. 1	15.18		
16	9.50			smtO. 6		add0. 6		4.7			18. 7	4.20	16
17						0.19							17
18	10.50			0.35	3.54	0.32	5.46					3.20	18
19	10.58				8.52				6. 5			2.51	19
20				1. 2	3.50	0.58					14.17		
21	11.55											1.51	21
22 23 24	11.50				3.43						13.47	1.20	
25	12. 6				3.88		6. 4		7.29		13.30	0.80	
24	12.21		6.36				6. 8 6. 8	2.18			18.13	0.20	34
25	12.30				-					15.44			_
36	12.4									15.50		0.40	
27 28 29 30	13. 2				3.17							1.10	127
28	13.14						6. 9		9.10				2
23	13.96	Ī	4.4		8. 9							2. 9	229
			4.45										12
31	13.46	31	4.27	-	2.47	Γ	б. 4	0.20		16.12	1	3. 7	31

For reducing the Sun's Declination as given in the Nautical Almanac for noon at Garran
wich, to any other Time under any other Meridian.

							mer 1									
			Sub. aft. Add bef.													
	Add in Sub. in		Sub. in		5 Deg	10 Deg	15 Deg.	20 Deg	25 Deg	30 Deg	35 Deg	40 Deg		ub. in W. dd in E.		d in W.
ŀ	Days		Days				M.8							Days.		Days.
١,	-		Decemb				_			_			-	June	·]	June
)CC61MO	20 20		29	0 4	10. 1	0. i	lo. i	0. 2	0. 2	0. 5	do. 3	\bar{b}_2	June	20	June
1		19		23	0. 0	lő. i	0. 2	0. 2	0. a	0. 4	U. 5	0. 6	28		19	- 1
1		18	i	24	0. 1	0. 2	0. a	0. 4	0. 6	0. 2	7 O. E	Ø. 9	24		18	ł
1		17	j	25	lO. 1	10. 3	10.4	10.6	10.7	I O. S	N O.11	0.12	125		17	ŀ
1		16		26	0. 9	0. 4	0. 6	0. 7	0. 9	0.11	0.18	0.15	26		16	ł
1		15 14					0. 6								15	
Т		13	1	29	0. 8	10. 0	0. 7 0. 9	0.10	0.12 0.15	0.10	0.10	0.21	20		14 13	- 1
1	-	12												June	12	
1-		11	Decemb.	_			0.11					0,30	_	July	11	
1			January				0.12								io	ł
1		9		2	0. 4	0. 8	0.18	0.17	0.21	0.26	0.30	0.35	3		9	
1		8	l	3			0.14								8	i
ł		7	l	4			0.15								7	1
1		6	1	5 6			0.16								6	•
		5	1	7			0.17 0.18								6	1
1		3		8			0.19								3	• 1
1		2		9			0.20								8	
ĬĔ)ecemb	. 1		10	0. 2	0.14	0.21	0.29	0.36	0.43	0.54	0.57	11		1.	June
N	ovemb	. 30		11	0. 7	10.15	0.22	19 .3 0	0.37	0.46	0.59	1. 0	12			May
1		29		12	0. 8	0.16	0.23	0.81	0.39	0.47	0.56	1. 8	18		200	· 1
1		28					0.24								29	l
		27 26					0.25								26	ŀ
ı		25		16	0. 8	0.18	0.26 0.27	0.37	0.46	0.55	1: 3	1113	17		27 26	- 1
1		94					0.28								66	- 1
1		28			0.10	0.20	0.29	0.39	0.49	0.58	1. 9	1.19	19		94	I
		92		19			0.80					L			28	
Г		21		200	0.10	0.21	0.81	0.41	0.51	1. 2	1.12	1.22	21		20	
1		20		21	0.11	0.22	0.82	0.48	0.58	1. 4	1.14	1.25	22		ÞΙ	- 1
		19		92	0.11	0.22	0.88	0.44	0.55	1. 6	1.17	1.28	28		20	- 1
1		18 17		28 24	0.11	0.28	0.34 0.34	0.46	0.50	1. 7	1.19	1.80	27 05		19 18	
1		10		25			0.85								17	1
ı		16		26	0.12	0.24	0.36	0.48	1. 0	1.12	1.24	1.36	97		li6	I
1		14		27	0.12	0.25	0.37	0.49	1. 2	1.14	1.26	1.39	28		16	
1		18	_	28	0.18	0.96	0.38	0.51	1. 4	1.16	1.28	1.41	20		14	•
L			January	_	_		0.89								13	
		9	February			0.27								August	10	
ı		4		3			0.42 0.43								8	
1		ä		7			0.46								4	
N	ovemb.	. 1		ġ			0.46					2. 8			9 1	May
0	ctober				0.16	0.32	0.47	1. 8	1.19	1.36	1.50	2. 6	12		BO 1	
1		28					0.48					2. 9			28	1
1		26 24					0.49			1.39	1.56	2.12			96	1
1		21	•	17 90	0.17	0.84	0.50 0.52	1. 4	1.29	1.41	9 1	9 10	01 10		24 21	ı
1-		18					0.53								18	
1			February	2	0.16	0.24	0.54	1.12	1.22	1.40	2. 4	2.22	27		16 16	- 1
1		12	March	1	0.18	0.37	0.55	1.14	1.32	1.51	2. 9	2.28	30	August	12	- 1
1		9		4	0.19	0.88	0.56	1.15	1.34	1.53	2.12	2.30	2	Sept.	9	- 1
		đ		7	0.19	0.38	0.57 0.57	1.16	1.85	1.54	2.18	2.82	5	•	6	_ [
	ctober	3		10	0.19	0.38	0.57	1.17	1.36	1.55	2.14	2.84	8	1		pril
٦	eptemb	껅		10	0.19 0.18	0.30	0.58	1.17	1.87	1.00	2.10	2.30	1 L 1 A		28 21 1	March
		24					0.58								25	- 1
	After	þ	Before				0.59							ore	After	ogle
1 1	-quino:	E. []	Equinox.	Į.	,	ł	- 1	- 1	- 1	1	l	j	Equ			BOX

TABLE V.

For reducing the Sun's DECLINATION as given in the Nautical Almanac for noon at (wich, to any other Time under any other Meridian.

. 122 - A N	Sub. aft. N.	H M	H M	и м	нм	UM	X X XX	U M	٠e	- AN		
	Add bef. N.											
Add in W.		45	50	55	60	65	70			b. in W.		ld in W
Sub. in E.									A	id in E.		b. in E
		MS	MS	MS	MS	MS	MS	MS	-	Days.		
Days.	Days.										-	Days.
	Decemb. 21	0. 0	0. 0	0. 0	0. 0	0. 0	0. 0	0. 0	21	June		June
20	22	0. 8 0. 6	0. 0	0. 4	0. 4	0. 4	0. 0	0. 0	62		20	
18		0.10									19 18	
1 17	25	0.13	0.15	0.16	0.18	0.19	0.20	0.22	25		17	
16	26	0.16	0.18	0.20	0.22	0.24	0.26	0.27	26		16	
16	27	0.20	0.22	0.24	0.26	0.29	0.31	0.33	27		15	
14	28	0.23	0.25	0.28	0.31	0.34	9:3 6	0.38	28		14	
13	29					0.88				_	13	
12	30									June	12	
	Decemb. 31									July	11	
		0.36									10	
9	2	0.39	0.44	0.48	0.53	0.57	1. 2	1. 6	3		9	
1 5	3	0.43	0.48	0.58	0.57	1. 2	1. 7	1.11	5		8	
8 7 6		0.46									7 6	
	6	0.49 0.52	0.68	1. 4	1.10	1.18	1.22	1.27	1 7		6	
1 4	ž		1. 1	1. 7	1.14	1.20	1.26	1.32	8		4	
8		0.58									8	
] 2	9	1. 1	1. 8	1.15	1.22	1.29	1.36	1.43	10		2	
Decemb. 1	10	1. 4	1.12	1.19	1.26	1.33	1.41	1.48	11		Ti	June
Novemb. 30		1. 7										May
229		1.10									30	
28	13	1.18	1.22	1.30	1.38	1.46	1.54	2. 2	14		29	
27	14	1.16	1.25	1.34	1.42	1.50	1.58	2. 7	15		28	
26	16	1.19	1.28	1.37	1.46	1.55	2. 3	2.12	10		27	
25 24	10	1.22	1.31	1.40	1.49	1.09	2. 8	2.17	17		26 25	
23	10	1.25 1.28	1.30	1 47	1.57	2. 3	2.12	2.21	10		24 24	
200	19	1.30	1.41	1 51	2. 1	2.11	2.21	2.81	20		23	
21		1.33									22	
20	21					2.19					21	
19	22	1.39									20	
18	23	1.41	1.58	2. 4	2.15	2.26	2.37	2.48	24		<u> </u>	
17	24	1.43	1.56	2. 7	2.18	2.30	2.41	2.52	25		18	
16	25	1.46	1.58	2.10	2.21	2.33	2.45	2.56	26		17	
15	26		2. 1	2.18	2.25	2.37	2.49	3. 1	27		16	
14	.27	1.51 1.54	2. 4	2.10	0.28	2.40	2.62	3. 5	28		15	
	January 30	1.59	2 11	2.15	2.37	2.44	2.50	3. Y	31	July	14 12	
7	February 1	2. 8								August	8	
á		2.11									6	
8		2.14									4	
Novemb. 1	ģ	2.19	2.83	2.49	3. 4	3.19	3.35	3.50	10			May
October 30		2.22										April
28		2.25									28	-
26	15	2.29 2.32	2.45	3. 2	3.18	3.85	3.51	4. 8	16		26	
24	17	2.32	2.49	3. 5	3.22	3.39	3.56	4.18	18		24	
91	20	2.36	Z.08	0.11	3.28	5.40	4. 8	4.20	Z.		21	
18	23	2.40	2.68	3.16	3.33	3.51	4. 8	4.26	24		18	
	February26	2.46	ö. [320	3.38	3.50	4.14	4.32	27 20	A	15	
9	March 1	2.40	g. 9	2 04	0.42 9 AE	4 1	4.19	4.08	<u>ت</u>	August Sept.	12 9	
1 8	7	2.51	3 17	2 20	3.40	7.7	4.20	4.41	5	ocpt.	6	
October 8	10	2.53	8.18	3.32	8.51	4.10	4.29	4.49	8			April .
Septemb 30	13	2.55	3.14	3.33	3.53	4.18	4.32	4.51	11			March
27	16	2.56	3.15	8.34	8.54	4.14	4.33	4.52	14		28	
24	19	2.56	3.15	3.35	3.55	4.15	4.83	4.52	17	_	25	
	Before	2.56	3.15	3.35	3.55	4.15	4.34	4.53	Be	ore		er _e
! Fauinox.	Equinex.	ł	ŧ	1	- 1	1		- vigit	ĸq	′ . xoسنن	i E.q	ninox.

K TAB.

For reducing the Sun's Declination as given in the Nautical Almanac for noon at Greenwich, to any other Time under any other Meridian.

		wich,												
Add af Sub. be	t. N.	Sub. aft. I Add bef.	N.	H.M 6. 20	5.40	6. "	6. 20	6.40	7. 0	7. 20	Ad	d bef. N.	Sul	b. bef. N.
Add in		Sub. in W	7.	80 Dag	85 Dec	90 Dec	95 Dec					ıb. in W. Id`in E.		id in W. ib. in E.
Sub. in		Add in E							M.S.		-	Days.	-	Days.
Day	s.	Days. Decemb.											-	June
Decemi	0. 21 20	Decemo.	22	0. 5	0. 6	0. 6	0. 7	0. 8	0. 8	0. 8	22	June	20	June
1	19	5	23	0.11	0.12	0.13	0.14	0.15	0.15	0.16	23		19	
	18		24	0.17	0.19	0.20	0.21	0.22	0.23	0.24	24 05		18	
	17 16		26 26	0.23	0.20	0.20	0.25	0.37	0.31 0.38	0.32	26 26		17 16	•
ĺ	15		27	0.35	0.38	0.40	0.42	0.44	0.46	0.49	27		15	
ł	14		28	0.41	0443	0.46	0.49	0.51	0.54	0.57	28		14	
İ	13 12		29 30	0.47	0.50	0.53	1.30	1 6	1. 2	1.0	20	June	13 12	
		Decemb. 3							1.17		ī	July	11	
l		January	"	1. 5	1. 9	1.13	1.17	1.21	1.25	1.29		July	10	
	9	,	2	1.11	1.15	1.19	1.24	1.28	1.32	1.37	3	•	9	
1	8		3						1.40				8	
	7		4	1.22	1.23	1.32	1.44	1.49	1.47 1.54	2. 0	6		7	
i	5		6	1.33	1.39	1.45	1.51	1.57	2. 2	2. 8	7		8	j
į	4		7	1.39	1.45	1 61	1.57	2. 3	2. 9	2.16	8		4	
1	3 2	•	8	1.44	1.50	2 2	2. 4 2.10	2.10	2.16 2.23	2.30	חו		3 2	,
Decemb									2.30				<u>. </u>	June
Novemb			11	2. 0	2. 7	2.15	2.22	2.30	2.87	2.45	12			May
	29		10	2. 5	2.13	2.21	2.29	2.37	2.44	2.52	18		30	
l	28		13	2.10	2.19	2.27	2.35	2.43	2.51 2.58	8. 0	14		29 28	
	27 26	, :	14	2.10	2.30	2.38	2.47	2.56	3. 5	3.13	16		97	
1	25		ıal	2 26	2.35	2.44	2.53	13. 2	3.11	3.21	17		26	
	24		17	2.31	2.40	2.50	2.59	3. 9	3.18 3.24	3.28	18		25	
1	23 22	;	18 19	2.30	2.51	3. I	3.11	3.21	3.31	3.41	20		24 23	
	21		20	2.46	2.56	3. 6	3.17	3.27	3.37	3.48	21		00	
	20	•9	21	2.50	3. 2	3.12	3.23	3.33	3.44	3.55	22		51	
	19								3.50 3.56				20 19	
l	18 17		24	3. 4	3.16	3.27	3.39	3.50	4. 1	4.13	25		18	•
l	16		25	3, 8	3.20	3.32	3.44	3.56	4. 7	4.19	26		17	
Ì	15			3.13	3.25	3.37	3.49	4. I	4.13 4.19	4.20	27		16	
	14 18	9	27 28	3.22	3.34	3.47	4. 0	4.12	4.25	4.38	29		15 14	
	. 11	January 3	30	3.30	3.43	3.56	4. 9	4.22	4.36	4.49	31		12	
		February	1	3.38	3.51	4. 5	4.18	4.32	4.46	4.59	2	August	10	
1	7		3	3.46	4. 0	4.14	4.28	4.42	4.56 5. 5	5.10	6		8	
1	5 3		7	3.59	4,14	4.29	4.44	4.59	5.14	5.29	8		1 4	
Novemi	b. 1		ø	4. 5	4.21	4.36	4.52	5. 7	5.23	5.38	10		2	May
October			11	4.12	4.28	4.44	5. 0	5.16	5.31 5.40	5.47	12			April
l	28 26		16 15	4.19	4.60	4.57	5.14	5.30	5.40 5.47	6. 3	16		28 26	
	24	1	17	4.30	4.47	5. 8	5.21	5.38	5.55	6.12	18		24	
	21								6. 4				21	
	18	F-1-	23	4.44	5. 2	5.19	5.37	5.55	6.13	6.31	24		18	
l	10	February: March	20	4.50	6. 8 6.15	5.20	5.59	6.10	6.20	6.47	27 90	August	15 12	
	9		4	5. O	5.19	5.38	5.57	6.16	6.34	6.53	2	Sept.	9	
	6		7	5. 4	5.23	5.42	6. 1	6.20	6,39	6.58	5	•	6	
October Septem			10	5. 8	5.27	5.46	ნ. ნ ჩ. ₽	6.25	6.44 6.47	7. 3	8		3	April March
-chrem	27		16	5.12	5.31	6.51	6.11	6.31	6.50	7. 9	14		28	
١	24]_ :	19	5.12	5.32	5.52	6.12	6.32	6,51	7.11	17		25	
After		Before Equinox.	Į	5.13	5.88	6.53	6.13	6.33	6.52	7.11	Bei	o re Digitized b	Αñ	
Estimo	UX.	Equinox.							<u> </u>		rq	pigitized b	rLq	O'xoun

For reducing the Sun's DECLINATION as given in the Nautical Almanac for noon at GREENwice, to any other Time under any other Meridian.

	WICH, to	_					-			
Sub. bef. N	Add bef. N.	7.40	8. 0	8.20	8.40	9. 0	9. 20	9.40	Add bef. N	Add aft. N. Sub. bef. N.
Add in W Sub. in E	Sub. in W.	115 Deg.	120 Deg.	125 Deg.	130 Deg.	135 Deg.	140 Deg.	145 Deg.	Sub. in W.	Add in W. Sub. in E.
Days.	Days.	M.S.								Days.
	1 Decemb. 21									21 June
2	D 22	0. 9	0. 9	0. 9	0.10	0.10	0.10	0.10	22	20
]		0.17	0.18	0.18	0.19	0.19	0.20	0.21	23	19
1		0.34			0.28					18 17
) i	6 26	0.42	0.44	0.46	0.48	0.49	0.51	0.53	26	16
1	5 27	0.51	0.53	0.55	0.57 1. 7	0.59	1. 1	1. 3	27	15
1 1		1 8	1. 2	1. 0	1.17	1.9	1.12	1.14	28 20	14 13
i		1.16	1.19	1.23	1.26	1.29	1.32	1.35	30 June	12
i	Decemb. 31	1.24	1.28	1.32	1.35	1.39	1.43	1.46	1 July	11
1	Oµanuary 1	1.33	1.37	1.41	1.45	1.49	1.53	1.57	2	10
		1.42 1.49								8
	8 . 8 7 4	1.58	2. 3	2. 8	2.13	2.19	2.23	2.28	5	17
	6) 5	2. 5	2.11	2.16	2.22	2.28	2.33	2.39	16	6
		2.14								6
ï		2.22 2.29								8
		2.37								2
Decemb.	1 10	2.45	2.52	2.59	8. 6	8.14	3.21	3.28	11	1 June
Novemb. 3	ol 11	2.52	8. 0	3. 7	3.15	8.23	3.30	8.38	12	31 May
2	9 12	3. 0	3. 8	8.16	3.24	8.32	3.39	3A7	13	30
2	5 18 7 14	3. 8 3.15	8.24	3.32	3.41	8.49	8.58	4. 6	15	29 28
2	6) 15	3.22	3.81	8.40	l 3.49	8.58	4. 7	4.16	116	27
2	5 10	3.30	3.39	8.48	8.57	4. 7	4.16	4.25	17	26
2	4 17 2 19	3.87 3.44	3.46 9.54	3.50	4.0	4.10	4.24	4.34	10	25 24
2		3.51								23
2	1 20	3.58	4. 8	4.19	4.29	4.39	4.50	5. U	21	22
2	21	4. 5 4.12	4.16	4.27	4.37	4.48	4.59	5, 9	22	हिं
1:	B 92	4.12	4.23	4.54	4.40	5. 4	5.15	5.15	24	20 19
i	7 24	4.25	4.36	4.48	5. 0	5.12	5.23	5.34	25	18
1	6 25	4.31	4.43	4.55	5. 7	5.19	5.30	5.42	26	17
1		4.38	4.50	5 8	5.21	5.20 6.23	5.46	5.50	28	16 15
1:	28	4.50	5. 3	5.16	5.28	5.40	5.54	6.6	29	14
	l January 30	5. 2	5.15	5.28	5.41	5.54	6. 8	6.21	31 July	12
	February 1	6.13	5.27	5.40	5.54	6. 8	6.22	6.35	2 August	
	7) 8 5. 5	5.24 5.34	5.38	6.02	6.19	0.20	6.47	7. 9	6	8
	R 7	5.44	5.59	6.14	6.29	6.44	6.59	7.14	18	4
Novemb.	1! 9	5.53	6. 9	6.24	6.40	6.55	7.11	7.26	10	2 May
October 3) 11 10	6. 3 6.12	6.18	6,34	6.50	7. 6	7.21	7.87	14	30 April 28
2	el 1 <i>e</i>	1000	000	A EO	P 10	P CA	~ AO	7 60	11R	26
2	4 17	6.29	6.45	7. 2	7.19	7.36	7.52	8. 9	18	24
2	II 22U	יענט.ט וי	0.00	1 4.10	1.01	17.40	10. U	0.22	21	21
11	23	6.48	7. 6	7.24	7.42	8. 0	8.17	8.34	24	18 15
19	February26 March 1	7. 6	7.10	7.42	8. 1	8.20	8.38	8.57	27 30 August	
	4	7.12	7.31	7.50	8. 9	8 28	8.46	9. 6	2 Sept.	9 °
0-4-1-	5 7	7.17	7.36	7.56	8.14	8.33	8.53	9.12	5	6 3 April
October 3 Septemb,8	10	7.23 7.26	7.42 7.45	8.4	8.94	8.42	9. 2	9.18	11	31 March
2°	7) 16	7.29	7.48	8. 7	8.27	8.47	9. 6	9.25	14	28
2	4 19	7.30	7.50	8.10	8.29	8.49	9. 8	9.27	17	26
After Equinox.	Before Equinox.	7.81	7.50	8.10	8.30	8.54)	y. 9	y.28	Before Equinox.	After C Equinox.
Zquillox.	Jaquidox.			<u> </u>	·	<u> </u>	<u> </u>		Equitor.	LEQUINUE.

TABLE V.

For reducing the Sun's Declination as given in the Nautical Almanac for moon at Granswich, to any other Time under any other Meridian.

	<u> </u>	7. 6 - 7. 6 .	****	****		****	** **	17 50			
Add alt.	Ŋ.	Add bef N	H.M	H. M.	H.M.	H. M.	11.01.	11 40	19 O	DOD. SEL. I	Add aft N. Sub. bef. N.
	_		150	156			170	175	180		
Add in V		Sub. in W. Add in E.	Deg.		160 Deg.	165 Deg.	Deg.	Deg.	Deg.	Sub. in W Add in E	
	<u>.</u>		M.S.					M. S.	M. S.		
i)avs.		Days.		M. S.						Days.	Days.
Decemb.		Decemb. 21 22					1	0. 0 0.13		21 June	21 June
1	20 19	23									20 19
	18		1					0.38			18
1	17	25	1 .		0.47	0.49			0.53		17
	16	26	I						1. 6		16
1	15 14	27 28				1.13 1.25		1.17 1.80	1.19 1.32		16
	13	29			1.84		1.40	1.43	1.46		14
1	12	30				1.49	1.52	1.55		30 June	19
	11	Decemb. 31	1.50	1.54	1.57	2. 1	2. 5	2. 8	2.12	1 July	11
1	10	January 1	2. 1	2. 5	2.9	2.13		2.21	2.25	9	10
1	9	2					2.30	2.34	2.38		9
1	8	3			2.32 2.44	2.37 2.49	2.42 2.54	2.47 2.59	2.51 3. 4	4	18
	7	5	1	2.50	2.55	3. 0		8.12	8.17		6
1	5	6			86	3.12	3.18	3.24	3.30		5
1		7			3.17	8.23				-	4
1	8	8						3.48			18
	2	9									9
Decemb.	1	10 11	3.35 3.45	3.42 3.52				4.11 4.22	4.18 4.30		1 June 21 May
Sovemb.	29	12	I								31 May 30
1	28	13		4.18		4.29					20
	27	14		4.23	4.31	4.40					28
1	26	15				4.50					1 27
į	25 24	16 17		4.43 4.53				5.19 5.30			26
	23		4.62	5. 2	5.12						25 24
-	22	19		5.12	5.22	5.32		5.52			28
	21	20	5.10	5.21	5.81	5.42	5.58	6. 3	6.13	21	29
	20	21	5.20		5.41	5.52	6. 3	6.14	6.24		31
	19	22	5.29	5.40	6.51	6. 2	6.13	6.24			20
	18 17	23 24		5.49 5.57	6. 0 6. 9	6.11 6.20	6.23 6.32	6.34 6.43	6.44 6.54		19 18
1	16	25		6. 6	6.17	6.29	6.41	6.53	7. 4		17
	15	26		6.14	6.26	6,38	6.51	7. 3	7.14		16
1	14	27	1		6.34	6.47	7. 0	7.12	7.94		1.5
	13	28 Tanuara 20	•	6.81 6.47	6.43	6.56 7.13	7. 9 7.26	7.22 7.40	7.34	29 31 July	14
		January 30			7. 0						12
1	7	February 1 8		7. 3 7.17	7.16 7.81	7.30 7.45	7.43 7.59	7.57 8.13	8.11 8.28	2 August 4	10
1 .	5	. 5	4	7.31	7.45	8. 0	8.14	8.28	8.43	6	16
L	3	7	7.29	7.44	7.59	8.14	8.28	8.43	8.68		4
Novemb.	1	9		7.56	8.12	8.27	8.42	8.58	9.13		2 May
	30 28	11 13	7.53 8. 4	8. 8 8.20	8.24 8.36	8.40 8.53	8.56 9. 9	9.12 9.25	9.28 9.42		30 April 28
	26	15		8.82	8.48	9. 5	9.21	9,38	9.54		26
	24	17	8.26	8.43	9. 0	9.17	9.84		10. 7		24
	21	20	8.40	8.57	9.14	9.22	9.49	10. 6	10.24	21	21
	18	23		9.10	9.28		10. 3				18
		February26	9. 4	9.22			10.16				15
1	12 9		9.15 9.24			10.10	10.90	10.47	11. 6	30 August 2 Sept.	12
I	2		9.24	9.50	10.	10.20	10.47	11. 6	11.24	2 жерт. 5	6
October	8		9.37	9.56	10.16	10.35	10.54	ii.13	11.32	8	8 April
Septemb.	80	13	9.41	10. O	10.21	10.40	10,59	11.18	11.38	11	31 March
	27	16	9.45	10. 4	10.24	10.44	11. 8	11.22	11.42	14	28
After	24	Befor e	9.47	10. 6	10.26	10.46	11. 5 11. 6	11.24	11.44	17 Refore	26 After_
	. li	Equinox.	J. 70	.u. /	10.01	-0.7/	*** 4			Equinox.	Equinos
	_	,				'		·			

1

TABLE VI.—SUN'S RIGHT ASCENSION.

1	Jan.	Feb.	March.	April.	May.	June.	July.	Aug.	Sept.	Oct.	Nev.	Dec.	1
Days.	H. M.	11. M	H. M.	н. м.	H. M.	H. M.	11. M.	II. M.	H. M.	H. M.	н. м.	H. M.	Days.
1	18.45	20.58	22.48	0.42	2.33	4.35	6.40	3.44	10.41	12.29	14.25	16.29	1
2			22.52	0.45	2.37	4.39	6.44	8.48	10.44	12.32	14.29	16.33	2
3	19.54	21.6	22.55	0.49	2.40	4.44	6.48	8.52	10.48	12.36	14.33	16.37	3
			22.59	0.52		4.49						16.42	
5	19. 3	21.14	23. 3	0.56	2.48	4.52	6.56	9. 0	10.55	12.43	14.41	16.46	3
6	19. 7	21.18	23. 7	1. 0	2.52	4.56	7. 0	9. 4	10.59	12.47	14.45	16.50	6
7	19.12	21.22	23.10	1.3	2.56	5. 0	7. 4	9. 8	11. 2	12.51	14.49	16.55	7
8	19.16	21.26	23.14	1. 7	3. 0	5. 4	7. 8					16.59	S
9	19.21	21.30	23.18		3 . 3	5. 8				12.58	14.57	17. 4	9
10	19.25	21.34	23.21	1.14	3. 7	5.12	7.17	9.19	11.13	13. 2	15. 1	17. 8	10
11	19.29	21.38	23.25	1.18	3.11	5.17	7.21	9.23	11.17	13. 5	15. 5	17.12	11
12	19.34	21.42	23.29	1.22	3.15	5.21	7.25					17.17	
13	19.38	21.46	23.32	1.25	3.19	5.25						17.21	
			23.36	1.29	3.23	5.29	7.33					17.26	
15	19.47	21.54	23.40	1.33	3.27	5.33	7.37	9.38	11.31	13.20	15 21	17.30	15
16	19.51	21.59	23.43	1.36	3.31	5.37	7.41	9.42	11.35	13.24	15.25	17.34	16
17	19.55	22. 1	23.47	1.40	3.35	5.41	7.45					17.39	
18	19.59	22. 5	23.51		3.39	5.46	7.49					17.43	
11			23.54	1.48	3.43	5.50	7.53					17.48	
20	20.8	22.13	23.58	1.51	3.47	5.54	7.57	9.57	11.49	13.39	15.42	17.52	20
21	20.12	22.17	0. 2	1.55	3.51	5.58	8. 1	10. 0	11.53	13.43	15.46	17.57	21
22	20.16	22.21	0. 5	1.59	3.55	6. 2		10. 4					
23	20.21	22.24	0. 9	2. 2	3.59	6.6		10. 8					
	20.25		0.12	2.6	4. 3	6.11		10.11					
25	20.29	22.32	0.16	2.10	4. 7	6.15	8.17	10.15	12. 7	13.58	16. 3	18.14	25
26	20.33	22.36	0.20	2.14	4.11	6.19	3.21	10.19	12.11	14. 2	16. 7	18.19	26
	20.37			2.18	4.15	6.23		10.22				i8.23	
	20.41			2.21	4.19	6.27		10.26				18.28	
29	20.46	22.46		2.25	4.23	6.31		10.30					
30	20.50		0.34	2.29	4.27	6.35	8 37	10. 3 3	12.25	14.17	16.24		I— I
31	20.54		0.38	2.33	4.31		8.41	10.37		14.21	1	18.41	31

This Table gives nearly the mean of the Sun's Right Ascension for the years 1817, 1818, 1839, and 1820, and i sufficiently exact for finding when any Star comes to the meridian. But in all calculations for determining the longitude by celestial observations, the Sun's Right Ascension must be taken from the Nautical Almanac, wher it is calculated to a greater degree of accuracy.

TABLE VI. A.

Correction for the daily variation of the Equation of Time found in Table IV. A. Find the daily variation of Equation of Time at the top, the hour at Greenwich at the side.

1 2 3 4 5 6 1 0 0 0 0 0 0 0 2 0 0 0 0 0 1 3 0 0 0 1 1 1	7 8 9	10 11	12 13	141	5 16	17 18	19	50 3	1 22	23	24 2:	26	27	28	29	30	De.g.
1000000	0 0 0	0 0	-1-	1 1-	7 7	1	-	-			-	17	-1	-	-,	1	15
1 0 0 0 0 0 0	0 0 0	0 0	11 1		1 i	1	9 9	2	1 6		ءُ اه	1	2	4	2	3	30
2 0 0 0 0 0 1 3 0 0 0 1 1 1	1 1 1	1 1	2 2	1 3	2 2	0	2 2 2	3	2 2 3		2 3	2	3	-	2	4	45
	1 1 1	1 1	2 2	1 -1	2 2			3	3 3	3	3	7 3	3	4	2	5	60
4 0 0 1 1 1 1	1 1 2	2 2 2	2 2	2	3 3	.0	3 3	3	4 4	4	-4 -	-	의	-0	5		
5 0 0 1 1 1 1 6 0 1 1 1 1 2 7 0 1 1 1 1 2	1 2 2	2 2 2	3 3	3	3 3	4	1 4	4	4 5	5	5 !	5	6	6	6	6	75
6 0 1 1 1 1 2	2 2 2	2 3 3	3 3	4	4 4	4	5 5	5	5 6	6	6 (7	7	7	7	8	9 0
7 0 1 1 1 1 2	2 2 3	3 3 3	4 4	4	4 5	5	5 6	6	6 6	7	7	8	8	8		9	105
8 0 1 1 1 2 2	2 3 3	3 3 4	4 4	5	5 5	6	6 6	7	7 7	8	8 8	9	9	9	10	10	120
0 0 1 1 0 0 0	9 2 2	2 1 4	5 5	5	6 6	6	7	8	8 8	9	9 0	10	10	11	11	11	135
9 0 1 1 2 2 2 10 0 1 1 2 2 3	9 9	4 4	5 5		6 7		8 8		9 9		10 1						
10 0 1 1 2 2 3	3 3	4 4 5	- 1 -			0	8 9		010	1.5	1011						165
11 0 1 1 2 2 3 12 1 1 2 2 3 3	3 4	4 5 5	-1 -		7 7 8 8		9 10	1 -1-	010	12	12 1	13	1	14	15		180
12 1 1 2 2 3 3	4 4 3	5 5 6	6 7		_ _	-	-1-		1111	1	_				13		
13 1 1 2 2 3 3	4 4 ;	5 5 6	7 7		8 9	9 1	0 10				13 1						
14 1 1 2 2 3 4	4 5 8	5 6 6	7 5		9 9	101	1 11	121	2 13	13	14 1	5 15	16				210
15 1 1 2 3 3 4	4 5 (6 6 7	8 8	3 9	9 10	111	1 12	13 1	314	14	15 1	5 16	17	18	18	19	225
16 1 1 2 3 3 4	5 5	6 7 7	8 9	9 1	0 11	111	2 13	131	4 15	15	161	117	18	19	19	20	240
17 1 1 2 3 4 4	5 6	6 7 8	9 3	أأما	TIT.	121	3 13	14	5 16	16	17 1	118	19	20	21	21	255
19 1 0 0 2 4 5	5 6	7 8 8	910		1119	121	4 14		617	17	181	9 20	20	21	22	23	270
18 1 2 2 3 4 5 19 1 2 2 3 4 5	6 6	7 8 9	1011		9 12	131	4 15		7 17								28.
20 1 0 2 2 3 4 5	6 7	-	10 11	12	3 13		5 16	177	2 12	10	20 2						
20 1 2 3 3 4 5	0 /		10 11	11		- 1	_										
21 1 2 3 4 4 5 22 1 2 3 4 5 6		8 9 10	11 11		3 14		6 17	181	18 19	50	21 2	2 23	24	25	25	120	51.
22 1 2 3 4 5 6	6 7	8 9 10	11 12	2 13 1	14 15	161	7 17	18 1	19 20	21	22 2	3 24	25	30	27	35	330
23 1 2 3 4 5 6 24 1 2 3 4 5 6	7 8	9 10 11	121	2 13 1	14 15	161	7 19	19	21	55	23 2	4 25	150	27	28	139	3 :
24 1 2 3 4 5 6	7 8	9 10 11	121	3 [1] 1	15/16	171	3 19	20	21/22	23	24 2	5 26	27	120	129	30	360

TABLE VII.

AMPLITUDES.

												_	_	_	_	_	_				_	٠					_							_
	L	at.	<u> </u>	4 65	4	9	6	7	&	6	2	Ξ	12	13	7	12	92	11	18	19	50	25	7	£ 1	E	22	\$ 6	N 6	X) G	38	100	3	33	
	28	Σ	80 0	3 8	8	8	36	န	3	47	2	26	=	1	1	2	28	36	3	2	4	2	56		5	4	200	3	£,	2 0	14	0	2 2	1
	35	Ω	83 8	3 8	8	8		83	<u>ස</u>	23	53	23	24	24	25	2 7	₹.	4	\$		35	35	35	200	3			2 2	9	3 6	18	8	8	
	23	Σ	٥-	- •	. 4	9	8	F	14	. 18	83	27	.33	88.			.59		. 15			45	.55	_			å .	٦;		2. 6	-	9	3	1
1		0	23.	3 %	33	23	23	123	23	23	23	23	21	8					57				<u> </u>	25		35				25	6	2		
	22	Σ					<u>.</u>	1	7.	=	2	.26	85	8	<u> </u>		38		<u></u>	-:	8	•	S. 28	٠.'						3 5		3 2	•	ı
		0	0 52	16	8	5	72	102				25 22	0 22	35 22	<u>55</u>	722		123	823			34 23		55 24	2	~~	30.54	<u> </u>	67 20	1 20			8 26	1
	2	DM	21.	: -	: _:	_;	<u>.</u> ;		Ξ	•	•	25 25	<u>ਦ</u>				21.53	65 65						•	ان	Τ.		-		24.9			25.1	ı
	_	Σ	0 -		8						<u>중</u>	<u>83</u>			38				50				3		26				2 2	162	16.			ı
	50	2	20°	S	8	8	8	8	8	20.	20			8					21.	21	•	•		2			N 6	S	, s		2	8		
	61	Σ	0 -	- 6	ಣ	20	-	6	랄	15	8	25	26	31	36	42	\$	\$	-	∞	9	22			2			200	2	5 0	16	35		1
	_	0	5 5	9	6	6	_	_	<u> </u>	<u> </u>	<u>6</u>	-	<u>.</u>	6	<u> 3</u>	<u> </u>	<u> </u>	<u>6</u>	8	8	<u>양</u>			ရွ	<u>ຊ</u>	21.	; ;	7	7	61	6	67	8	١
	≊	Σ	0 -	• 6		₹.	9	. 8	Ξ	٠.	•	•	.25	.29					. 58			- 4			3	99.	- :			4 2	ľ	9	8	1
	_	10	9 18	2 2	28	8	8 5	8 28		3 18	318	=	=				_	8 8	Ę	<u>=</u>	_	<u>=</u>	=	5	=	2	<u>8</u>			2 2				
	2	D M	7.				Ŀ	١.	Ξ.	~	7.16		.2	7.28	7.32	۲. ي	7.42	17.4	2.02	_		_		8.3	Ž.	3.49		٠.		2 4			2.2	
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TABLE VII.

AMPLITUDES.

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Right Ascensions and Declinations of some of the principal fixed Stars, adapted to the beginning of the year 1920, with their annual variations.

Names and situations of the STARS.	Characters.	Magnitudes.	Right Ascension.	Aun. Var. R. A. add after. 1820.	Decimation.	Annual Variation.
			H.M.S.	"	0,	,,
Extremity of the wing of Pe-	1					
gasus, Algenib	7	2.3	0. 3.59	3.08	14.11 N.	+20.1
In the head of the Phenix	•	2.5	0.17.22	2.99	43.16 S.	-20.0
Bright Star in the tail of the Whale	β	2.3	0.34.32	5.00	18.59 S.	-19.9
Polar Star, tail of the Little	"	2.0	0.01.02	0.00	10.00 5.	-15.5
Bear	a	2.3	0.57. 1	14.26	88.21 N.	+19.4
In the girdle of Andromeda	β	e	0.59.40	3.80	34.40 N.	+19.5
The spring of the River Erida,				l	1	,
Achernar	a	1	1.81. 0	2.25	58. 9 8.	-18.5
Almach in the foot of Andro-	į			l		
meda	γ	2	1.52.58	3.62	41.28 N.	+17.7
*The following horn of the Ram,			1 5 7 0		00 00 N	
a ARIETIS		2	1.57. 3 2.10.15	3.35 3.02	22.36 N. 3.48 S.	+17.4
In the neck of the Whale		2	2.52.52	3.12	3.48 S. 3.23 N.	-17.0 +14.8
-In the jaw of the vy hair			2.02.02	02	0.2011.	714.0
In the head of Medusa, Algol	β	2	2.56.28	8.85	40.15 N.	+14.6
The bright Star in Perseus	a	2	3.11.31	4.20	49.15 N.	+13.6
The bright Star of the Pleiades,	l		ł	1	•	·
or Seven Stars	7	3	3.38.47	3.54	23.33 N.	+11.8
*The southern eye of the Bull,	l			1		
ALDEBARAN	a	1	4.25.36	3.43	16. 8 N.	+ 8.0 ,
In the left shoulder of Auriga,	a	1	5. 3.24	4.41	45.48 N.	
Capella The bright foot of Orion, Rigel	β	i	5. 5.58	2.88	8.25 S.	+ 4.6 - 4.9
The northern horn of the Bull	B	2	5.14.55	3.78	28.27 N.	+ 3.8
The western shoulder of Orion	-	2	5.15.29	3.21	6.11 N	+ 4.0
	8	2	5.22.49	3.06	0.27 S.	- 8.5
In the belt of Orion	ε	2	5.27. 4	3.03	1.19 S.	— 5.0
Į.	ζ	2	5.81.41	3.02	2. 3 S.	- 2.6
Drinks Stan in the Dans			F 00 0	2.17	24 10 5	
Bright Star in the Dove The eastern shoulder of Orion	a	2	5.33. 9 5.45.26	2.17 3.25	34.10 S. 7.22 N.	- 2.4 + 1.4
In the foot of the Great Dog	B	2.3	6.14.46	2.64	17.52 S.	+ 1.4 + 1.2
In the poop of the ship Argo		2.0	0.17.70	~.04	17.02 5.	T 1.2
Canopus	4	1	6.19.57	1.33	52.36 S.	+ 1.7
In the ankle of Pollux	7	2.3	6.27.19	3.46	16.33 N.	- 2.3
In the mouth of the Greater	•					
Dog, Sirius	a	1	6.37.13	2.64	16.29 S.	+ 4.4
In the thigh of the Greater Dog		2.5	6.51.32	2.35	28.44 S.	+ 4.4
In the back of the Greater Dog	δ	2.3	7. 1. 3	2.44	26. 7 S.	
In the tail of the Greater Dog In the head of the northern	η	2	7.16.59	2.58	28.57 S.	+ 6.5
Twin, Castor	a	1.2	7 .25 . 6	3.85	32.16 N.	— 7.1

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Right Ascensions and Declinations of some of the principal fixed Stars, adapted to the beginning of the year 1820, with their annual variations.

Names & situations of the $STARS$	Characters.	Magnitudes	Right Ascension.	Anu. Var. R. A. add after 1820.	Declination.	Annual Variation.
·			H. M. S.	"	0 /	"
The Lesser Dog <i>Procyon</i> The head of the southern	a	1.2	7.29.52	8.15	5.41 N,	- 8.5
Twin, Pollux In the row lock of the ship	β	1.2	7.84.17	3.69	28.27 N.	8.0
Argo	ξ	2	7.57.16	2.12	59.50 S.	⊥ 9 7
In the poop of the ship Argo	7	2	8. 4. 1	1.86	46.48 S.	110.3
In the middle of the ship Argo	ð	2.3	8.39.45	1,66	1 24. 3 3.	- 12.9
In the oars of the ship Argo The heart of the female Hydra,	β	2.5	9.11.15	0,75	68.59 5.	+14.9
Alphard	a	2	9.18.44	2.95	7.53 S.	+15.2
The Lion's heart REGULUS South pointer in the sq. of the	•	1.2	9.58.46	5.21	12.51 N.	-17.3
Great Bear North pointer in the sq. of the	β	2	10.50.54	8.71	57.21 N.	19.1
Great Bear	a .	1.2	10.52.32	3.85	62.45 N.	—19.3
The Lion's tail—Denebola S. E. Star of of the Great	β	1.2	11.59.52	3.07	15.35 N.	-20.0
Bear N. E. Star of O of the Great	7	2	11.44.19	5.20	54.42 N.	20.0
Bear	ð	3	12. 6.27	3.02	58. 2 N.	-20.1
In the foot of the Cress	•	1	12.16.41	3.24	62. 6 S.	+ 20.0
In the top of the Cross In the following arm of the	7	~	12.21.15	5.24	56. 6 S.	+20.0
Cross Alioth, first star in the tail of the	β	~	12.37.18	8.41	58.41 S.	-
Great Bear	6	1	12.48. 8	2.75	56.56 N.	-19.7
The Virgin's spike—Spica The second Star in the tail of the	•	1	15.15.45	5.14	10.13 S.	
Great Bear	5	2.3	18.16.59	2.45	55.5₹ N.	 19.0
Last Star in the tail of the Great						
Bear	77		13.40.28	2.58	50.13 N.	18.2
The western foot of the Centaur	β		13.51.14	4.10	59.30 8.	+17.8
In the tail of the Dragon The bright Star in Bootes—Arc-	•	1	13.59.51	1.63	65.14 Ŋ.	17.4
turus	a		14. 7.26	2.73	20. 8 N.	-19.0
The eastern foot of the Centaur The southern scale of the Ba-	4		14.28. 0	4.44	60. 7 S.	
lance The northern scale of the Ba-	a	2.3	14.40.56	3.29	15.17 S.	-
lance	β	2.3	15. 7.20	3.22	8.43 S.	+15.8
Bright Star in the crown Gemma	•	2	15.27. 4	2.53	27.20 N.	-14.5

Right Ascensions and Declinations of some of the principal fixed Stars, adapted to the beginning of the year 1820, with their annual variations.

Names & situations of the STARS.	Obaracters.	Kagnitudes.	Right Ascension.	Ann. Var. R. A. add. after 1820.	Declination.	Annual Variation.
In the neck of the Serpent The northernmost Star of the		2	H. M. S. 15.55.24	2.94	7. 0 N.	
Scorpion's forehead	β		15.55. 0 16.18.23	5.47 3.66	19.18 \$.	
*The Scorpion's heart, Antares In the eastern knee of Ophiuchus	a	-	17. 0. 3	5.42	26. 1 S. 15.50 S.	
In the head of Hercules	a		17. 6.27	2.73	14.36 N.	
In the head of Ophiuchus	a	2	17.26.35	2.77	12.42 N.	
In the head of the Dragon	7		17.52.25	1.58	51.31 N.	
In the bow of Sagittarius	8	2.3	18.12.14	5.98	54.28 S.	- 0.9
The bright Star in the Harp, We- ga, LYRA * Bright Star in the Eagle, Atair,	a	1	18.30.51	2.08	58.57 N.	+ 5.0
a AQUILE		1	19.42. 0	2.95	8.24 N.	+ 9.1
The eye of the Peacock	4	-	20.11.20	4.85	57.18 S.	10.8
The tail of the Swan Deneb	4		20.35.18	2.04	44.39 N.	+12.6
The western wing of the Crane	a	2	21.56.49	5.85	47.49 S.	-17.4
*In the mouth of the southern fish, FOMALHAUT In the shoulder of Pegasus	β	-	22.47.41 22.55. 3	8.34 2.87	30.84 S. 27. 7 N.	
*In the wing of Pegasus, Markab	۱,	9	22.55.48	2.98	14.14 N	1194
In the head of Andromeda	a		23.59. 6	5.08	28. 6 N.	
Near the shoulder of Cassiopea	β		23.59.35	3.05	58. 9 N.	

Note.-If the places of these stars are wanted for any time before the be ginning of the year 1820, multiply the annual variation, in right ascension, by the number of years before 1820, and subtract the product from the right ascension standing in the table; but the product of the annual variation in declination by the number of years before 1820 must be added to, or subtracted from the declination, according as the sign — or + is marked in the Table; but for any years after 1820, the annual variation in right ascension multiplied by the number of years after 1820 must be added to the right ascen sion in the Table, and the annual variation in declination multiplied by the num ber of years after 1820 must be either added to, or subtracted from the declination, according to the signs in the Table.-The annual variation is set down for seconds and decimals of a second. An asterisk is prefixed to the stars whose distances from the moon are given in the Nautical Almanacvery great accuracy is required, the corrections found in Tables XLII. an XLIII. for aberration and nutation, are to be applied to the numbers deduced from Table VIII. but these corrections are generally not of much importance in nautical calculations. The corrected values are however given in the Nautical Almanac for 24 of the bright stars of this catalogue for every ten days in the year, and these values are always to be preferred.

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TABLE X.

For finding the Distance of Terrestrial Objects at Sea, in Statute miles.

	Height in	Distance.	Height in	Distance.	Height in	Distance.	Height in	Distance.
	feet.	Mil. Dec.	feet.	Mil. Dec.	feet.	Mil. Dec.	feet.	M. Tenths
	1	1.32	44	8.78	320	23.67	1000	41.8
	2	1.87	45	8.87	330	24.03	1100	43.9
	3	2.29	46	8.97	340	24.39	1200	45.8
	4	2.65	47	9.07	350	24.75	1300	47.7
	5	2.96	4.8	9.17	360	25.10	1400	49.5
	6	3.24	49	9.26	370	25.45	1500	51.2
	7	3.50	50	9.35	380	25.79	1600	52.9
,	8	3.74	55	9.81	390	26.13	1700	54.5
	9	3.97	60	10.25	400	26.46	1800	56.1
	10	4.18	65	10.67	410	26.79	1900	57.7
	11	4.39	70	11.07	420	27.11	2000	59.2
	12	4.58	75	11.46	430	27.43	2100	60.6
į	13	4.77	80 ′	11.83	440	27.75	2200	62.1
	14	4.95	85	12.20	450	28.06	2300	63.4
	15 16	5.12	30	12.55	460 470	28.37	2400 2500	64.8 66.1
		5.29	95	12.89		28.68		
	17	5.45	100	13.23	480	28.98	2600	67.5
ļ	18 19	5.61	105	13.56	490	. 29.29	2700	68.7
	20	5.77 5.92	110 115	13.88 14.19	500 520	29.58 30.17	2800 2900	70.0 71.2
	21	6.06	120	14.49	540	30.74	3000	72.5
	22	6.21	125	14.79	560	31.31	3100	73.7
	23	6.34	130	15.08	580	31.86	3200	74.8
	.24	6.48	135	15.37	600	32.41	3300	76.0
	25	6.61	140	15.65	620	32.94	3400	77.1
~	26	6.75	145	15.93	640	33.47	3500	78.3
	27	6.87	160	16.20	660	33.99	3600	79.4
	28	7.00	160 .	16.73	680	34.50	3700	80.5
	29-	7.12	170	17.25	700	35.00	3800	81.6
1	30	7.25	180	17.75	720	85.50	3900	82.6
	31	7.37	190	18.24	740	35.99	4000	83.7
	32	7.48	200	18.71	760	36.47	4100	84.7
1	33	7.60	210	19.17	780	36.95	4200	85.7
	34	7.71	220	19.62	800	37.42	4300	86.8
	36	7.83	230	20.06	820	37 .88	4400	87.8
	36	7.94	240	20.50	840	38.34	4500	88.7
	37	8.05	250	20.92	860	38.80	4600	89.7
	38	8.16	260	21.33	880	39.25	4700	90.7
	39	8.26	270 280	21.74	900 920	39.69	4800	91.7
	40 41	8.37	290	22.14	940	40.13	4900 5000	92.6 93.5
	42	8.47 8.57	300	22.53 23.91	960	40.56 40.99	1 mile.	96.1
	45	8.68	310	23.29	980	41.42	r nare.	30.1
		T.00	U.V	~~.~		770		

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sek the nearest number to the reduced time in the top column, and the difference of parallax, proportional logarithm, or semi-diameter for 12 hours in the side column; under the former and opposite the latter, is the correction to be applied to the number, marked first in the Nautical Almanac, additive if increasing, subtractive if decreasing.

<u> </u>										19	EDU	CE) TI	Œ.										
ar. ir	Ь	h 1	h 14	h 2	h	h 3	h	h 4	h	h 5	h	6	6 h	h 7	h 74	Ь 8	h 84	h 9) h	h 10	h 103	h	h	h 12
in12	<u>ь</u>	h	h	h h	23 h	h h	34 h	h	44 h	b h	5 <u>1</u>	b	h	h	h	h	h	Б	h	h	h	11 h	114 h	h
hrs.	12 <u>4</u>	13	13 <u>}</u>	14	144	15	154	16	161	17	17}	18	184	19		20	204	21	213	22	221	23	23 <u>1</u>	24
1	0	0	0	0	0	0	0	0	0	0	0	0	Ţ.	1	Ī	1	1	1	I	1	1	1	1	1
3	0	0	0	0	0	0	1	1	1	1 1	1	1	1 2	1 2		1 2	1 2	1 2	2 2	2 2	2	3		2 3
4	0	0	0	1	1	1	1	1	1	2	2	2	2	2	2	3	3	3	3	3	3	4	4	4
5	_	9	1	1	1	1	1	녆		2		2	3	3		3	4	4	4	4	4	5		5
6	0	ĭ	1	1	1 1	1 2	2 2	2 2	2	2 3	3	3	3	3	4	5	4 5	4	5 6	5	5 6	5 6	6	7
8	Ð	1	1	1	2	2	2	3	3	3	4	4	4	5	5	5	6	6	6	7	7	7	8	8
10	0	1	1	1 2	2	2	3	3	3	4	5	4 5	5	5 6	6	6	6	7	7 8	8	8	8	9	9 10
11	0	ī	1	2	2	3	3	4	4	5	5	5	6	6	7	7	8	8	9	9	10		11	11
12	0	1	1	2	2	3		4	4	5	5	6	6	7	7	8	8	9	9	10			11	12
13	1	1	2 2	2	3	3		5	5	6	6	6	8	8	8	9	9 · (10	10 10	10 11	11		12 13	12	13 14
15	1	Ĭ.	2	2	3	4		5		6	7	7	8	9	9	ro	11	11	1	12		14	14	15
16	1	ı	2	3	3	4		5		7	7	8	9		10	11	11	12	13	13	14	15	15	16
17 18	1	1	2 2	3	4	4		6		7	8	8	9 10	10 10		11 12	12 13	13 13	13 14	14 15			16 17	17 18
19	1	2	2	3	4	5	6	6	7	8	9	9	10	11	12	13	13	14	15	16	17	17	18	19
20	1	2	2	3	4	5	[7	7	8		10	11		12	_	14	15	16	17		18		20
21 22	1	2	3	3	5	5		7	8		10 10	10 11	11' 12	12 13		14 15	15 16	15 16	17 17	17 18			20 21	21
23	1	2	3	4	5	6	7	8	9	10	11	11	12	13	14	15	16	17	18	19	20	21	22	23
24	1	2	3	4	5	6	7	8	9	10 10		12 12	13 14	14 15			17 18	18 19	19 20	20 21			23 24	24 25
26	<u>:</u>	101	3	4	5	6	8	9	10		12	13	14	15		17	18	19	21				25	26
27	1	2	3	4	6	7	8	9	10	11	12	13	15	16	17	18	19	20	21	22	24	25	26	27
28 29	1	2	3	5 5	6	7	8	9 10	10	12		14 14	15 16	16 17		19 19		21 22	22 23	23			27 28	28 29
30	î	2	4	5	6	7	9	10		12		15	16	17		20		22	24	25			29	30
31	1	3	4	5	6	8	9			13		15	17		19	21	22	23		26			30	31
32 33		3	4	5	7 7	8	9 10	11		13 14		16 16	17 18	19 19		21 22	23 23	24 25		27 27		2 9 30	31 32	32
34	1	3	4	6	7	8	10	11	13	14	16	17	18	20	21	23	24	25	27	28	30	31	33	34
35		3	4	6	7	9		12			16	17		20		23	25	26	1	_	_	_	34	35
36 37		3	5	6	7 8	9	10 11	12 12		15 15		18 18		21 22		24 25	25 26	27 28	28 29	30 31			34 35	36 37
38	2	3	5	6	8	9	11	13	14		17	19		22		25	27	28	30	32			36	38
39 40		3	5 5	6		10 10	11	13 13	15	16 17		19 20		23 23		26 27	28	2 9		32 33			37 58	39 40
41		3	5	7		10						20		_	26 26	_	20 29	31	32 32	_		38	3° 39	41
42	2	3	5	7	9	10	12	14	16	17	19	21	23	24	26 26	28	30	31	33	35	37	38	33 40	42
43 44		4	5 5	7 7		11 11									27		30	32		36	38		41	43
45		4		7		11 11		15 15		18 19		22 22		26 26	27 28	29 30	31 32	33 34		37 37		40 41	42 43	44 45
				_		_				_						_	-			-		-		

TAB	LES	XI	, XIII,	XI	V, X	(V,	AND	XV	ĭ.		
TABLE XII. The Refra Heavenly Bodies in App. Ref. App. Ref.		ide.	Depressi Horis	on o		of the	1		a'aus	: XIV Parsi ude.	lax in
Ait. Ait.	Alt.	1661.	Height the eye		Dip o		Sto	n's A	Jt.		un's allax.
D. M. M. S. D. M. M.S. 0. 033. 0 6.307.52	D. 30	M. S. 1.38	Feet.	+		. S.	1-	D.		8	
0. 5 32.11 6.407.41	31	1.35	1			59	1	0		9	
0.1031.22 6.507.31		1.51 1.28	2 3	1		34 42	H	10 20		9	
0.20 29.50 7.107.12	34	1.24	4	- 1	1.	58	1	30	Ì	8	
0.25 29 . 6 7.207 . 3 0.30 28 . 23 7 . 30 6 . 54		1.21	5 6		2. 2.			40 50		7	
0.35 27.41 7.40 6.46		1.16	7	-	2.			55		5	
0.40 27. 0 7.50 6.38 0.45 26.20 8. 0 6.30		1.13		-	2. 2.			60 65		4	
0.50 25.42 8.10 6.22		1. 8	10	4	3.		.]]	70	1	3	
0.55 25. 5 8.20 6.15 1. 0 24.29 8.30 6. 8		1.5	11		3, 3.		1	75 80	- 1	2 2	
1. 5 23.54 8.40 6. 1	43	1. 1	18	- 1	3.	33	ll	85	- 1	1	
1.10 23.20 8.50 5.55 1.15 22.47 9. 0 5.49		0.59 0.57	14	-	3.4 3.4			90		0	
1.20 22.15 9.10 5.43	46	0.55	16	1	3.	56		T Augma		XV. ion o	f the
1.25 21.44 9.20 5.37 1.30 21.15 9.30 5.31		0.53 0.51	17 18	1	4.		M	oon's	Sem	i-dia	neter.
1.35 20.46 9.40 5.26	49	0.50	19	ì	4.	17	Mo		Ale.		entation
1.40 20.18 9.50 5.20 1.45 19.51 10. 05.15		0.48	20	-	4.			D. 0		S.	
1.50 19.25 10.15 5. 8		0.46 0.45	21	.	4.		l	5	- {	1	
1.55 18.59 10.50 5. 0		0.43	23	!	4.			10 15		3 4	
2. 018.3510.45M.54 2. 518.1111. 04.47		0.41 0.40	24 26		5 .4		1	20	Ì	5	
2.1017.4811.154.41		0.38	28	-	5.			35		7 8	
2.15 17.26 11.304.35	58	0·37 0. 3 6	. 50 35	Ì	5.1 5.4			35	- }	ŝ	
2.25 16.44 12. 0/4.23	59	0.54	40		6.			40	<u> </u>	10	
2.30 16.23 12.20 4.16	60 61	0.33 0. 32	4.5 50		6.		1	45 50	- 1	11 12	
2.40 15.45 13. 04. 3	62	0.30	60	-	7.	37		55	- }	13	
2.45 15.27 13.20 3.57 2.50 15. 9 13.40 3.51		0. 2 9 0. 2 8	70- 80	- [8.1 8.4		:1 	60 70	- 1	14 15	
2.55 14.52 14. 03.46	65	0.27	90	1	9.5	0		80	- 1	15	
3. 014.3514.203.40 3. 514.1914.403.35	66	0.25 0.24	100		9.1	51	1	90		16	
8.10 14. 3 15. 0 3.30	68	0.23	Table X	VI.	Dip	of th	e Sei	a at d	iffere	nt D	istances
3.15 13.48 15.30 3.23 3.20 13.33 16. 0 3.17		0.22	1.			om tl					
3.25 13.19 16.30 3.11		0.21 0.20	- ਭਰ	He	ght o	the	Eye (pove	the	Sea i	a Feet.
3.30 13. 5 17. 0 3. 5 3.40 12.39 17.30 2.59	72	0.19	B die	5	10	15	20	25	30	35	40
3.50 12.14 18. 0 2.54	74	0.17 0.16	miles.		Dip.	_	Dip.	Dip.		Dip.	
4. 0 11.50 18.30 2.49 4.10 11.28 19. 0 2.44	75	0.15	3 5	M.	M.	M.	M.	M. 57	M.	M.	M. 91
4.20 11. 7 19.30 2.40	77	0.14 0.13	1	11	25 12	34 17	45 23	28	68	79 40	45
4.30 10.47 20. 0 2.36 4.40 10.28 20.30 2.32	78 79	0.12 0.11	, 3	4	8	12	15	19 15	23 17	27	30 23
4.50 10.1021. 02.28		0.10		3	8	9	12	12	14	16	19
5. 0 9.53 21.30 2.24	81	D. 9 '	12 13	3	4	6	8	10	12	14	16
5.10 9.3722. 02.20 5.20 9.2123. 02.14		0. 8 0. 7	2	2 2	3	5 4	7	8	9	11	12 10
5.30 9. 724. 02. 7 5.40 8.53 25. 02. 2	84	0. 6	2 <u>1</u>	1	3			6	7	8	9
5.50 8.39 26. 0 1.56). 5). 4	34	2	3	4	5	6	6	. 7	8
6. 0 8.2727, 01.51	87 (). 3	4 5	2	3	4	5	5	6	6	7
6.10 8.15 28. 0 1.47 6.20 8. 3 29. 0 1.43). 2	6	2	3	÷	न	5	5	6	6
NOTE TO TABLE XVI.	—Th	e nun	bers of	this	table	belo	w the	bla	ck li	nes,	are the
same as are given in Ta not_being so far distant a	oie : s the	AUI. t land.	ne visible	р	rieod,		espos tized b		00		restapt#
							, CUILL	_	~~	200	

TABLE XVII.

WHEN A STAR IS USED.

Ap.	Cor.	Log.	Alt.	Cor.	Log.	Ap.	Cor.	Log.	*Ap. Alt.	Cor.	Log.	Alt.	Cor.	Log.
) M	MS	100	D M	MS		D M	M S	100	D M	MS		D M	M S	
5. 0	50. 8	0.9581	10. 0	54.45	1.2277	13.15	56. 2	1.3433	19. 0	57.16	1,4925	36.30	58.43	1.7517
10		0.9700			1,2297			1.3459		57.17	1.4960	37. 0	58.45	1.7568
20	50.39	0.9817	6	54.48	1,2317	25	56. 5	1.3485	20		1,4996			1.7618
30		0.9930			1,2337			1.3511			1.5031			1.7668
40		1.0041			1.2357			1.3537			1.5066			1.7717
50		1.0150			1,2377	_		1.3562		-	1.5101			1.7766
3. 0					1.2397						1.5136		58.51	1.7810
10		1.0360			1.2417			1.3612			1.5170			1.7854
20		1.0462			1.2437			1.3637			$\frac{1.5204}{1.5238}$			1.7900 1.7946
30 40	50 10	1 0562 1.0660	30	55 0	1.2457	14. 0	56.16	1.3687			1.5271			1.7287
50	52.29	1.0755	33		1.2490		56.17	1.3711			1.5304			1.8028
7. 0	_	1.0849	-	-	1.2515			1.3735	-	_	1.5338	-		1.8070
10		1.0941			1.2534			1.3759			1.5370			1.8112
20		1.1032	42	55. 5	1.2553			1.3783			1.5401			1.8152
30		1.1120	45	55. 7	1.2572	30	56.22	1.3807	30	57.36	1.5432	44. 0	59. 1	1.8192
35		1.1164	48	55. 8	1,2591			1.3831			1.5463		59. 2	1.8230
40	53.15	1.1207	51	55. 9	1.2610		-	1.3855		57.39	1.5494	45. 0	59, 3	1.8268
7.45	53.19	1.1250			1.2629			1.3878			1.5525			1.8338
48		1.1975			1.2648			1.3901	10	57.41	1.5556	47.	59. 7	1,8411
51		1.1301			1.2667			1.3924	20	57.42	1.5586	48.		1.8478
54		1.1326			1.2686			1.3947	30	57.43	1.5616	50		1.8547 1.8611
57 8. 0		1.1351 1.1376			1.2705 1.2724			1.3970 1.3993	50	57.45	1.5646 1.5676	51.		1.8676
_	-	The second second	-	_				_					-	1.8734
8. 3		1.1401			1.2742			1.4016	23. 0	57.46	1.5706	59		1.8794
6		1.1425 1.1450			1.2760 1.2778			1.4039 1.4061			1.5736 1.5765			1.8846
12		1.1474	21	55.22	1.2796			1.4083	30	57.50	1.5794	55.		1.8900
15		1.1499			1.2814			1.4105			1.5822			1.8956
18	53.44	1.1523	27	55.24	1.2832	40	56,39	14127			1.5850		59.23	1.9003
3.21	53.46	1.1547			1.2850	15.45	56.40	1.4149	24. 0	57.53	1.5879	58.	59.24	1.9050
24		1.1571	33	55.27	1,2868			14171			1.5907		59.26	1.9102
27		1.1595			1.2886			1.4193			1.5935			1.9142
30		1.1619	39	55.29	1.2904						1.5963			1.9183
33		1.1642	42	55.30	1.2922			1.4236	40	57.56	1.5990			1.9226
36		1.1666			1.2940	_	-	1.4258	_	-	1.6017	-		1.9270
3.39		1.1689			1.2957			1.4279			1.6044			1.9302
42	54. 0	1.1712 1.1735	51	55 25	1.2974 1.2991			1.4300			1.6097			1.9335
45		1.1758	57	55.36	1.3008			1.4321			1.6149 1.6201	E.E.		1.9404
51	54. 6		12. 0	55.37	1.3025			1.4363			1.6251	2000		1.9438
54	54.08		3	55.38	1.3042			1.4384			1.6301			1.9471
3.57		1.1826									1.6350		50.30	1.9501
0.0	54.12		9	55.40	1.3076	50	56.53	14425			1.6400			1.9528
3	54.13	1.1871	12	55.41	1.3093	55	56.54	1.4445			1.6449		59.42	1.9553
6	54.15	1.1893	15	55.42	1.3110	17. 0	56.55	1.4465	28. 0	58.13	1.6498	73.		1.9578
	54.17		18	55.43	1.3127	5	06.56	1.4486			1.6545			1.9603
-	54.19	-	-	_	1.3144	_	-	1.4506		_	1.6591		-	1.9625
	54.21		12,24	55.45	1.3161	17.15	56.58	1.4526			1.6635			1.9843
	54.22		27	55.46	1.3178	20	56.59	1.4546			1.6702			1.9660
	54.24				1.3194			1.4566			1.6769			1.9676
	54.26 I 54.28 I				1.3211 1.3227			1.4586			1.6833			1.9706
	54.29		39	55.50	1.3243	40	57. 3	1.4626	30	58 97	1.6957	81.		1.9714
-	54.31 1		-	-	1.3259	_	_	-	-		1.7018	-	_	1.9722
	54.33 I				1.3275			1.4646			1.7018			1.9729
	54.34 1				1.3291	55	57. 5	1.4684			1.7140			1.9734
	54.36 1							1.4708			1.7202			1.9737
	54.37 1				1,3323			1.4741			1.7262	86.	59.56	1.9739
	54.39 1	.2194	57	5.56	1.3339			1.4778				87.	59.57	1.9741
51	54.41 1	2215	13. 0	55.57	.3355	18,30	57,11	1.4815	35, 0	58,39	1.7862	88.	59.58	1.9742
	54,42 1	.2236	5 5	55.59	.3381	40	57.13	1.4852	30	58.40	1.7414	89.		1.9742
		2257		A 15 4	.3407	60	T 14	1 4000	BC O	50 40	1.7466	100	60 0	1.9742

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WHEN THE SUN IS USED.

,⊙Ap.	1_	I.O.A.	Cor.	;	ı () An	Cor.		(OAn	-		ı⊙An		
Alt	Cor. Lo	Alt.	Cor.	Log.	An.	Cor.	Log.	Alt.	Cor.	Log.	OAp Alt.	Cor.	Log.
D M	M S	D N				M S			M S			M S	
5. 0 10	50.16 0.96 50.32 0.97			1. 23 97 1. 24 18			1.3592	19. 0	57.24 57.95	1.5149 1.5187			1.7934
20	50.48 0.98			1,2139			1.3646			1.5225	30	58.53	1.8045
30 40	51. 3 1.00 51.16 1.01			1.2460 1.2481			1.3672 1.3699			1.5262 1.5299			1.8100 1.8154
. 50	51.30 1.02			1.2501			1.3725			1.5336			1.8206
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M 54' 55' 56' 57' 58' 59' 10 12.21 11.22 10.22 9.22 8.23 7.23 10 12.13 11.13 10.149 148.157.15 20 12.13 11.13 10.149 148.157.15 20 12.13 11.50 10.6 9.68.7 7.7 30 11.58 10.58 9.59 8.59 8.07.0 140 11.50 10.51 9.51 8.52 7.53 6.53 TABLE X Apparent Altitude of TABLE X Apparent Altitude of 224 225 225 221 2218 2218 2218 227 228 23 23 19 28 10 2225 2321 28 16 221 2200 220 28 10 2805 2800 2795 2792 2788 278 270 2784 2790 2785 2751 2777 2773 2766 2764 2759 2755 2751 2747 2743 273 0 2748 2744 2739 2735 2731 2727 272 20 2718 2714 2709 2705 2701 2697 269 2718 2714 2709 2705 2701 2697 269 2718 2714 2709 2705 2701 2697 269 2718 2714 2709 2705 2701 2697 269 2686 2684 2660 2665 2665 2665 2664 2660 2665 2665 2665 2664 2660 2665 2665 2665 2664 2660 2665 2665 2665 2664 2660 2665 2665 2665 2664 2660 2665 2665 2665 2664 2660 2665 2664 2660 2665 2664 2660 2665 2665 2664 2660 2665 2665 2664 2660 2665 2665 2664 2660 2665 2665 2664 2660 2665 2665 2664 2660 2665 2665 2664 2660 2665 2665 2665 2665 2665 2665 2665	12.21 11.22 10.22 9.22 8.23 7.23 6.2 10 12.13 11.13 10.14 9.14 8.15 7.15 6.16 11.50 10.51 9.59 8.59 8.07 0 6.4 11.50 10.51 9.51 8.52 7.53 6.53 5.5 TABLE XIX.	M 54' 55' 56' 57' 58' 59' 60' 10 12.21 11.22 10.22 9.22 8.23 7.23 6.24 5 10 12.13 11.13 10.14 9.14 8.15 7.15 6.16 5 20 12. 51 15.50 6.9 .68 .77 .7 6.8 .15 30 11.58 10.58 9.59 8.59 8.07 .06 .15 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4 11.50 10.51 9.51 8.52 7.53 8.52 9.53 5.53 1.52 1.52 1.52 1.52 1.52 1.52 1.52 1.52	M 54' 55' 56' 57' 58' 59' 60' 61' 0 12.21 11.22 10.22 9.22 8.23 7.23 6.24 5.24 10 12.13 11.13 10.14 9.14 8.15 7.15 6.16 5.16 20 12.5 11.5 10.6 9.5 8.8 07.06.15.16 30 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.54 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.54 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.54 Apparent Altitude of)'s centre Apparent Altitude of)'s centre 8 7 3 7 6 7 9 7 12 7 15 7 18 7 21 7 24 7 27 0 2841 2836 2831 2827 2823 2319 2815 2811 2807 10 2825 2821 2816 2812 2808 2804 2800 2796 2791 20 2810 2805 2800 2796 2792 2788 2734 2780 2776 20 2784 2742 4770 2766 2762 2758 2754 2750 2765 20 2764 2759 2755 2751 2747 2743 2739 2755 2731 0 2748 2744 2799 2735 2731 2727 2723 2719 2716 0 2738 2729 2724 2720 2716 2712 2708 2704 2706 20 2718 2714 2709 2705 2701 2697 2638 2639 2639 20 2818 2714 2709 2705 2701 2697 2638 2674 2671 20 2688 2684 2699 2664 2660 2656 2653 2649 2645 2641 0 2638 2684 2639 2645 2671 2671 2678 2634 2659 2656 0 2673 2669 2664 2660 2656 2653 2649 2645 2651 0 2638 2644 2639 2646 2641 2638 2634 2630 2622 20 2840 2824 2620 2616 2612 2608 2604 2600 2597 20 2540 2512 2508 2576 2572 2588 2554 2550 2546 2556 2558 0 2570 2566 2562 2558 2554 2550 2546 2560 2556 2553 0 2526 2522 2518 2514 2510 2506 2503 2499 2495 240 2599 2695 2591 2587 2588 2554 2550 2546 2556 2558 0 2570 2566 2562 2558 2554 2550 2534 2639 2645 2641 0 2628 2644 2649 2460 2467 2463 2460 2456 2455 2461 2467 2463 2460 2456 2452 2588 2554 2550 2544 2470 2476 2470 2475 2471 2467 2463 2460 2456 2452 2568 2558 2554 2550 2531 2527 2524 250 240 2492 2488 2481 2447 2440 2426 2422 2418 2411 2410 2407 2403 2399 2396 2398 2394 2390 2366 2332 2374 2331 2327 2333 2320 2316 2313 240 2332 2334 2330 2376 2372 2368 2354 2351 2347 2343 2340 2342 2338 2376 2372 2368 2354 2351 2347 2343 2340 2342 2338 2376 2372 2368 2354 2351 2347 2343 2340 2342 2338 2376 2372 2368 2352 2351 2357 2354 2330 2376 2372 2368 2352 2351 2357 2354 2330 2376 2372 2368 2352 2354 2351 2347 2347 2345 2340 2345 2341 2337 2347 2349 2345 2351 2347 2349 2345 2351 2347 2349 2345 2351 2347	M 54' 55' 56' 57' 58' 59' 60' 61' 5.24 10 12.13 11.13 10.14 9.14 8.15 7.15 6.16 5.16 10 12.15 11.5 10.6 9.6 8.77 7.6 8.5 8. 82 10.58 9.59 8.59 8. 9.7 0.6 15.1 1 30 11.58 10.58 9.59 8.59 8. 9.7 0.6 15.1 1 30 11.59 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 11.50 10.51 9.51 8.52 7.53 6.53 5.54 4.64 40 40 40 40 40 40 40	M 54' 55' 56' 57' 58' 59' 60' 61' 5.0' 10 12 .21 11 .22 10 .22 9 .22 8 .23 7 .23 6 .24 5 .24 10 42 10 12 11 .13 10 .14 9 .14 8 .15 7 .15 6 .16 5 .16 10 45 10 .51 9 .51 8 .52 7 .53 6 .53 5 .54 4 .54 40 13 30 11 .58 10 .58 9 .59 8 .59 8 .0 7 .0 6 .1 15 .1 30 22 40 11 .50 10 .51 9 .51 8 .52 7 .53 6 .53 5 .54 4 .54 40 13 50 10 .51 9 .51 8 .52 7 .53 6 .53 5 .54 4 .54 40 13 50 20 22 22 22 22 22 22	M	M	M	Of Park M 54' 55' 56' 57' 58' 59' 60' 61'	C	Cor.	Table Stat	M 54' 55' 56' 57' 58' 39' 60' 61' S. 0" 17' 2" 3" 4" 4" 4" 4" 4" 4" 4

CORRECTION

correction.	
s riorisontal Parallax.	TABLE A. tional part for Seconds of Parallax. Add. Tab.B For M of alt. Add. Add.
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Apparent Altitude of)'s centre.	TABLE C. Cor. for Seconds of Parallax. Add.
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0 2700 2696 2692 2639 2686 2682 2679 2676 2672 2669 26 10 2685 2681 2677 2674 2671 2667 2664 2661 2657 2654 261 20 2670 2666 2662 2659 2656 2652 2649 2646 2646 2659 2659 2650 2655 2651 2657 2654 2650 2655 2651 2657 2653 2650 2655 2650 2657 2653 2650 2657 2653 2650 2657 2653 2650 2650 2650 2650 2650 2650 2650 2650	651 2648 7 3 636 2653 8 1 621 2618 9 0 606 2603 592 2589 S. Cor.
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Explanation of Table XIX.

This table consists of two parts, for finding a correction of the moon's distance and logarithm corresponding: they are both in the same page from the beginning of the table to the altitude of 21 degrees, after which the correction is on the left hand page, and the logarithm on the right, both being found at the same opening of the book, in the following

To find the correction of Table XIX.

1. Enter the table marked Correction, and find in the side column the moon's apparent altitude, or the altitude next less if there be any units of miles in the altitude; opposite to this and under the minutes of the moon's horizontal parallax will be the approximate correction.

2. Enter table A, abreast of the approximate correction, and find the seconds of the moon's horizontal parallax, vis. the tens of seconds at the side, and the units at the top, under the latter, and opposite the former will be the correction of table A.

3. Enter table B, abreast of the approximate correction, and find the units of miles in the moon's apparent altitude (neglected above) opposite to which will be a number of seconds, which being added to the corrections found from table XIX. and from table A, will give the sought correction.

To find the Logarithm of Table XIX. Enter the table marked Logarithms, in the column titled at the top with the degrees and minutes nearest to the moon's apparent altitude, and find the logarithm corresponding to the moon's horizontal parallax in the side column, or the next less parallax if there be units of seconds in it. Abreast of this in the table C, opposite the units of seconds of parallax neglected, will be a correction, to be added to the former logarithm, to obtain the

logarithm sought.

It was observed in a former part of this work that in fixing these tables so as to render the corrections of the tables A, B, C, additive, it had been found necessary to make the greatest corrections correspond to 0" of parallax and 0' of altitude, so that when you find the exact parallax and altitude in the side and top columns of table XIX. it will still be necessary to refer to the tables A, B, or C, to take out the corrections corresponding to 0" of parallax or 0" of altitude. This is evident from the inspection of the tables, but it was proper to make this remark as a caution to prevent mistakes. To illustrate these rules, the following taken less than the properties are given, in which all the corrections are not down and added together, but after examples are given, in which all the corrections are put down and added together, but after a little practice it will be very easy to take the numbers from the table by inspection, and add them together without the trouble of writing them down separately.

LOGARITHMS.

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TABLE A. Proportional part for Seconds of Parallax. Add. D M 54' 55' 56' 57' 58' 59' 60' 61' S. 0' 1'' 2'' 3'' 4'' 5'' 6'' 7'' 8'' 9''	of a Add

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For the Correction.

In Tab. xix. to alt. 16° 20′ and par. 58′ is

Tab. A. 45″ pardilax

Tab. A. 45″ pardilax

Tab. B. 5′ altitude

Sought correction

For the Logarithm.

For the Logarithm.

Tab.xix.to nearest alt. 16° 20′ and par. 58′ 40″ is 2099

Tab. C. 5″ parallax

Sought logarithm

2105

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LOGARITHMS.

	Parallax									entre				Cor. for	ELE C. r Second arallax. add.
M	S	30 °	303°	310	3150	320	3210	330	3310	340	3440	350	3510	S.	Cor.
54	0	2360	2358	2356	2354	2352	2349	2347	2345	2344	2342	2340	2338	0	12
													2324	1	11
													2311	2	9
													2297	3	8
													2283	4	7
_	_												2270	5	5
55													2256	6	2
													2243	7 8	1
													2229	9	Ö
													2216	1	"
													2202	<u> </u>	- C
1													2189	<u>S.</u>	Cor.
56													2175	0	12
													2162	.1	11
													2149	2	9
													2136	3	8
													2122	4	5
									·		2113			5 6	4
57											2100			7	3
											2087			8	ĭ
											2074			ğ	l ō
													2057	1	· ·
											2048 2035			8.	Cor.
!	_														
58											2022			0	12 11
											2009			2	9
											1996 1983			3	8
											1971			4	7
											1958			5	6
-														6	4
											1945			7	3
											1933 1920			8	2
											1908			9	1
											1895			1	1
											1883			S.	Cor.
_	_										1870			0	12
											1858			ĭ	ii
											1845			2	10
											1833			8	8
											1821			4	7
											1809			5	6
61	ᇹ	1813	1811	1809	1807	1805	1803	1802	1800	1798	1796	1795	1793	6	5
											1784			7	3
											1772			8	2
											1760			9	1

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CORRECTION.

	s cen.				*	1)	s F	lo	riz	on	tal	Pa	ral	ax.	à				-			1	7	for	A	ec dd.	•	lso		ara		Tal For	r M
D	M	5	1/	1	55		- 5	64	1	57	1	5	8/	1 5	9	1	60	1	61	1	S	0	11	12	13	11/4	1"	5"	6"	7//	8"	9"	M	S
36	0	16	.31	1	5.4	13	14	. 5	4	14.	6	13	.1	7 12	.2	91	1.	40	10	51	(4	41	4	5 4	5	14	43	42	41	41	40	0	0
0	10	16	.36	1	5.4	18	15	. (0	14.	11	13	.2	3 12	.3	4 1	1.	46	10	57	10	39	3	3	7 3	7	36	35	34	33	33	32	1 2	1 7
100	90	16	.42	1	5.1	53	15		5	14.	17	13	.2	8 12	.4	0 1	1.	51	11	. 3	20	3	3	2	92	8	28	27	26	25	24	24	4	2
-4	30	16	.47	1	5.4	8	15	.1	0	14.	22	13	.3	1 12	.4	5 1	1.	57	11	. 9	30	2:	25	2 2	1 2	0	20	19	18	17	16	16	01000000000000	3
	40	16	.52	1	6.	4	15	.1	6	14.	27	13	.3	9 13	5	1 1	2.	3	11	.15	40	1.	1	1 1	3 1	2	12	11		9		8	7	4
	50	16	.57	1	6.	9	15	.2	1	14.	33	13	.4	5 12	5	7 1	2.	9	11	.21	50		1	5	5	4	4	3	2	1	0	0	100	5
37														0 15																	41		0	0
														6 13																			-04074507-000	i
-	20	17	.13	1	6.5	25	15	.3	7	14	.50	14		2 13	3.1	4 1	2.	26	11	.39	20	3	3	0 3	0 2	9	28	27	26	26	25	24	4	1 2
- 1	30	17	.18	1	6.5	30	15	.4	3	14.	55	14		8 13	3.2	0 1	2.	32	11	45	30	2	3 2	22	2 2	1	20	19	18	18	17		6	15
														3 13															10	10	9	8	8	5
_	00					_	No. 1		_ 1		250	-	_	9 1:	-	-	-		_	_		1				<u>□</u> ,	4	3	9	=	1	-	A SECTION	0
														6 15																			0	1
														2 13																			3	13
														7 1.																			4	2
														3 13 9 14																			6	13
		18												5 14				22								4	1	11	9	3	0	0	-NO4-001-00	5
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														1 14																			0	1 i
														7 14																			10000	12
														9 14																			CONTACTOR OF	13
														5 14																			7	14
														11												5	4	3	3	9	1	0	8	1.5
-	00	-	200	1-		35	_	-						9 1			_				1	-	-		9 7	9	10	11	40	40	20	20	-	10
														4 1																			ĭ	1 i
																																23	3	13
														7 1																			3	15
														3 18																			7	14
	50	19	.11	1	8.	25	17	.4	0	16	.55	16		9 14	5.2	4 1	4.	38	13	.53	50)	7	6	5	5	4	3	2	2	1	0	9	36
41	0	19	. 19	1	8.	32	17	.4	7	17	9	16	.1	6 1	5.3	1 1	4	46	14	. 0	1	14	14	3 4	2 4	2	41	40	39	39	38	37	0	10
	10	19	. 23	1	8.	38	17	.5	3	17	. 8	16	.2	3 1	5.3	7 1	4.	52	14	. 7	110	3	6 3	63	5 3	4	33	33	32	31	30	30	1 2	11
																																20		12
																																15	1 5	13
														2 15																			1 7	12
_1	50	19	.47	1	9.	2	13	.1	71	17.	33	16	.4	8 16	· .	3 1	5.	19	14	.34	50)	5	6	5	4	3	3	2	1	0	0	1 9	16

EXAMPLE VI.

Given the moon's apparent altitude 11° 20'	and horizontal	parallax 60 45%	Required the cor-
rection and logarithm?	*		- '
To find the Correction.	1	To find the Logarit	than.

To find the	Correction.	,			To find the l	Logari	ihm.		
In Tab. xix. to alt. 110	20' and pa	r. is	60′ 4 /	30″	Tab. xix. to nearest alt. 110	200∕an	d par	.00	40′′ 9052
Tab. A. 43" parallax				16	Tab. C. 3" parallax		•		9
Tab. B. O' altitude			•	2	Sought logarithm	_	_	_	2061
Sought correction			4	48"		•	•	•	300.

EXAMPLE VII.

Given the moon's apparent altitude 8° 40' and herizontal parallax 56' 20". Required the correction and logarithm?

To find the Corre	cuon.		1	10 pma ine	: Logar	unm.	_	
n Tab. xix. to alt. 80 40' and	par.	56' is 9' 1	18″ Tab.	xix. to nearest alt. 80	ວ 397au	d par. 5	<i>i6</i> / 2 0	" 2 518
Tab. A. 20" parallax .	•			C. 0" parallax .		٠.	•	13
. Tab. B. 0' altitude .			5	Sought logarithm		_		2531
Cha	×	10		need no sed as some	. •	. •	•	

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LOGARITHMS.

N HOL	Parallax.				••)'s c					Cer. for of Pa A	rallax. dd.
M	S	36°	36}°	37 °	3710	38°	3810	390	3910	40°	401°	410	4110	S.	Cor.
54	0	2337	2335	2334	2332	2331	2329	2328	2327	2326	2324	2323	2322	0	12
	110	9999	2321	2320	2318	2317	2315	2314	12313	12312	2310	2309	2308	1	11
	20	2309	2307	2306	2304	2303	2302	2301	2300	2298	2297	22 96	2294	. 2	9
	an.	2296	2294	2293	2291	2290	2288	2287	12286	2285	2283	2282	2281	3	8
	40	2282	2280	2279	2277	2276	2274	2273	2272	2271	2270	226 8	2267	4	7
													2254	5	5
55	·o	2255	2253	2252	2250	2249	2247	2246	2245	2244	2242	2241	2240	6	4
	10	2241	2239	2238	2236	2235	2234	2233	2232	2230	2229	2228	2227	7 8	2
	len	9998	2226	2225	2223	2222	2220	2219	2218	2217	2215	2214	2213		Ö
	30	2214	2212	22 11	22 10	2209	2207	2206	2204	2203	2202	2201	2200	1	ľ
	40	2201	2199	2198	2196	2 195	2193	2192	2191	2190	2189	2188	2186	<u> </u>	
	50	2188	2186	2185	2183	2182	2180	2179	2178	2177	2175	2174	2173	S.	Cor.
56	0	2174	2173	2171	2169	2168	2167	2166	2164	2163	2162	2161	2160	0	12
	l 10	9161	9150	2158	2156	2155	2153	2152	2151	12150	2149	2148	2147 (1	
	hei	9149	9146	2145	9149	2142	2140	2139	2138	2137	2135	2134	2133 l	2	9
	30	2135	2133	2132	2130	2129	2127	2126	2125	2124	2122	2121	2120	3	8
	M	9191	9119	2118	2117	2116	2114	2113	2112	2111	2109	2108	2107	4	7
	50	2108	2106	2105	2104	2103	2101	2100	2099	2097	2096	2095	2094	5	5
57	7	9005	9003	9009	9000	2080	2088	2087	2086	2084	2083	2082	2081	6	4
"	10	9089	9090	2079	9077	2076	2075	2074	2073	2071	2070	2069	2068	7	3
	90	2002	9067	2066	9064	2063	2062	2061	2060	2058	2057	2056	2055	8	. 2
	30	9056	9054	2053	9051	2050	2049	2048	2047	2046	2044	2043	2042	9	. 0
	40	9048	9041	2040	9039	2038	2036	2035	2034	2033	2031	2030	2029		
	50	2030	2028	2027	2026	2025	2023	2022	2021	2020	2018	2017	2016	S.	Cor.
_	<u></u>	8015	9016	9015	8019	0010	9010	9000	2008	9007	2006	2005	2003	0	12
28		2017	2010	2010	8000	1000	1007	1005	1995	1004	1993	1009	1991	i	lii
	10	1000	1000	1000	1007	1006	1085	1094	1982	1981	1980	1979	1978	2	9
	20	1070	1077	1076	1075	1074	1973	1971	1970	1969	1967	1966	1965	3	8
	40	1066	1065	1964	1069	1961	1960	1958	1957	1956	1955	1954	1953	4	7
	~	1054	1959	1951	1040	1948	1947	1946	1944	1943	1942	1941	1940	5	6
_														6	4
59	0	1941	1939	1938	1937	1936	1934	1933	1932	1010	1017	1016	1015	7	3
	10	1929	1921	1010	1924	1923	1921	1000	1919 1907	1006	1004	1004	1000	8	2
	20	1910	1914	1001	1911	1910	1909	1906	1894	1000	1909	1901	1800	9	1
	30	1903	1000	1000	1000	1005	1004	1000	1882	1881	1879	1979	1877	, ,	l
	20	1870	1877	1876	1874	1873	1871	1870	1869	1868	1867	1866	1865	S.	Cor.
_	20	1073	1000	1000	1000	1000	1011	1050	1057	1056	1055	1054	1059	0	12
50	0	1866	1804	1803	1802	1861	1809	1808	1857	1044	1949	1004	1840	ĭ	iĩ
	10	1854	1802	1001	1099	1098	1004	1000	1030	1091	1830	1990	1828	2	liö
	ZU	1661	1090	1005	1002	1004	1000	1003	1820	1810	1818	1817	1816	3	8
	30	1029	1012	1020	1020	1024	1022	1000	1808	1807	1805	1804	1803	4	7
	20	100%	1000	1019	1900	1700	1700	1707	1796	1795	1793	1709	1791	5	6
-														6	5
51	0	1792	1791	1790	1788	1787	1785	1784	1783	1782	1761	1780	1779	7	3
- 1	10	1780	1778	1777	1776	1775	1773	1772	1771	1770	1757	1768	1767	8	2
-	20	1768	1766	1765	1764	1763	1761	1760	1759	1758	1745	1720	1700	9	ī
ı	30	1756	1754	1753	1752	11761	1749	1748	1747	1 /40	140	1/44	1743	H	•

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		Sir.)'s H	forizon	tal Pa	rallax.			1	Prop	To. pa	rt f	or S	ec.	of I	Par.	- 1	Tat For of a	M.
٠Ī	54/	55'	56'	57'	58/	59'	60'	61	S	0" 1	11/2/1	3"	4/15	1/16	7 7"	8"	9"	M	S
ō	19.53			17.39	The same of the sa		1000			7.0	3 43	1			-	100	100	0	9
				17.45							6 35						30	3	1
0				17.52 17.58							1 20						23 15	3	3
-0				18. 4						2 /	4 13	100	-50			10.7	8	156780	3
-0				18.11						7			4		3 0	1	0	š	8
0	20.30	19.46	19. 2	18.18	17.34	16.50	16. 6	15.23	0	43 4	2 42	41	40	39 3	9 38	37	36	0	-
		19.52		18.25							5 34						29	2	1
				18.31							8 27						22	3	3
				18.38							1 20	100		300			7	ğ	1
				18.44				15.58	50		3 13 6 5		4	3	9 9	1	0		5
		-		18.58		_		-		_	1 41	1000	2.7	-	832	-	35	5	6
_				19. 5							435						28	X.	{
				19.11							7 26	100	1	200			21	345	3
				19.18					1		20 19		The state of	3-0165				8	3
				19.25							3 19	15.5	700	10	9 8	8		6788	5
_	_	-	-	19.31					50		6	100		3	2		0	9	-1
	21.46			19.39				16.49			0 40						35	Ŷ	0
				19.46				16.57			33 33 26 26						28	i	1
				19.59							9 19				6 18		14	3	1
				20. 6				17.18	11	10000	2 1	100	1000	9		3 7	7	4 3.27 8	3
50	22.18	21.36	20.50	20.13	19.31	18.49	18. 7	17.26	50	6	5	4	3	2	2	1 0	0	9	5
0	22.25	21.43	21. 1	20.20	19.38	18.56	18.15	17.33			10 40							Q	0
		21.50		20.27							33 33							4	į
				20.33							27 21							I	5
	22.44			20.40							20 15			10	9 5			Ž	13
				20.54						1000	100	5 4	100	3	2 3		ó	- Constant	6
	_	22.24	_		_	-	-	18.18	11-	-	39 39	1	25		16 3	115	34	0	0
		22.31						18.26			33 3								11
				21.16							26 2						20	1	3
		22.44		21.23							19 18				6 13		27.7		1
40	23.31	22.51	22.1	21.30	20.50	20. 9	19.29	18.48	40						9 8		7	7	9.6
	_	-	_	21.37	-	-	-	-			5		3	3			0	-	0
		23. 6				20.25					38 38 32 31						33 26	-0404-01-0-	ĭ
		23.12		21.52				19.12			25 24				9 21		20	3	3
		23.26				20.47		19.28			8 18	1000	70 L 10 lb	16.1	5 14	14	100	5	4
40	24.13	23.33	22.5	22.14							2 11			9	8 8		6		5
50	24.20	23.40	23.	22.21	21.42	21. 2	20.23	19.43	50	6	5 4	4	3	2	2 1	0	0	3	Ľ
-	24.27			22.29			20.31				7 37						100	Q	0
10				22.36							1 30							3	į
				22.44							4 24				1 20		1	Ž	1 4
	24.48			22.51											5 14		100		13
														2	- A	1 0		8	9
		24.16	23.37	22.58 23. 6	22.19	21.41	21. 2	20.23	40	12 1	1 11	10			-24-10	1 0	6 0	9	

LOGARITHMS.

-	Parallax.	Ja.	100		ppar									Cor. for of Pa	Second rallax. dd.
M	S	420	4210	430	4310	440	4440	45°	4510	460	470	480	490	S.	Cor.
54	0	2321	2320	2319	2318	2317	2316	2315	2314	2313	2311	2309	2308	0	12
	10	2307	2306	2305	2304	2303	2302	2301	2300	2299	2297	2296	2294	1	11
	20										2284			2	9
	30										2270			3	8
3	40										2256			4	7
	50					-	100	Contract of	100	A STATE OF	2243		111111111111111111111111111111111111111	5	5
55	0										2229			6	4
2	10												2213	7	3
	20												2199	8	1
3	30										2189			9	0
774	40												2173	10.00	100
	50	2172	2171	2170	2169	2168	2167	2166	2165	2164	2163	2161	2159	S.	Cor.
56	0	2159	2158	2157	2156	2155	2154	2153	2152	2151	2149	2148	2146	0	12
	10	2146	2144	2143	2142	2141	2140	2139	2138	2138	2136	2134	2133	1	11
	20	2132	2131	2130	2129	2128	2127	2126	2125	2125	2123	2121	2120	2	9
	30												2107	3	8
	40												2093	4	7
	50	2093	2092	2091	2090	2089	2088	2087	2086	208	2084	2082	2080	5	5
57	0										2071			6	4
3	10												2054	7	3
-	20												2042	8 9	0
	30												2029	9	0
	40										2019			-	-
	50							1	-		100000		2003	S.	Cor.
58	0												1990	0	12
6	10										1981			1	11
	20												1965	2	9
1	30												1952	3 4	8 7
	40 50	100000				1000				1000			1939 1927	5	6
			200		12 1000	-	-		_	-			1000	6	4
59	0												1914	7	3
3	10												1902	8	2
	20												1889	9	1
0	40												1864	4	1
0,0	50												1852	8.	Cor.
-				200										0	12
60	0												1840	1	11
1	10 20												1815	2	10
K	30												1803	3	8
	40												1791	4	7
	50												1779	5	6
61			2200	1000			-					-	_	6	5
01	10												1766	7	3
	20												1742	8	2
															1

-	_	-	-	-	_	-	-	_	-	-	_	_	_	_	_	-	-	EC.	_		ni.	-	-	-	-	_	_	_	_	÷	_	_	
It	cent.	1																			1					LE				100		Œ	
3. B	Ce	1)	s I	Ior	riz	ont	al	Pa	rai	lax					1	Pre	por	rt. I				ec.	of	P	ir.	Fo of	
Ap.	-																				1		è		112	Add	1.						å
D	M	1-	54	1	1	55	1	5	6/	T	57	1	5	8/	1	59	1	60'	I	61/	S	10/	01//	12//	377	14//	15//	611	17//	18"	900	M	_
50	-	2	_	_	-		0 0	_		1 9	_		_		-	10		1.1	-	-	-11-	-	37			-	-	-	-	33	-	9	1:
W	-	1																			110											1 .	11
- 1																					20												11
																					30											3	13
																					40									100		7	11
																					50								1 -	1	0	š	1
51	_	_	_				_ -	_				_	_	_	-	_	-	2. :	-			-	36		-			-	20	30	91	-	╁
																					10												H
																					20											1 3	14
																					36												
	40	26	, 9	0	25	.4	3 2	25.	. 6	2	1.9	8	23	.51	23	.14	1 23	2.3	7 21	.59	40	12	11	11	10	10	9	8	8	7	6	3	1
	50	26	5.9	7	25	5.5	0 2	25	. 13	3 24	1.5	6	23	. 59	23	. 25	2 20	2.4	5 29	. 8	50	6	5	5	4	3	3		1	1	0	Š	11
2	-	-	-	_	-	5.5	-			-	_	-	-	_	-	_	-	_	-	_	(-	-	1	-	-		-	90	31	31	_	١,
																					10											\ X	1
																					20											1 \$	1
						. 2															36											3	13
																					46							1	7	7	6	0-29456780	1
																					50								1	1	O	ĕ	1
	-	-	9	-	_	_	3 2	_			_	_	_	_	-	_		3.45	1		-	-	34	-	-	-	-	31	31	30	30	0	17
~1				-	-																10											0-0004000	П
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	Parallax.						•	`) 's				Cor. for of Pa	LE C. Seconds rallax. ld.
M	S	_i 50°	510	52°	53°	54°	,5 5 0	56°	57°	58°	59°	S.	Cor.
54	0	2306	2305	2303	2302	2301	2300	2298	2297	2296	2295	0	12
1	10	2293	2291	2290	2288	2287	2286	2285	2284	2283	2282	i	ii
1	20	2279	2277	2276	2275	2274	2272	2271	2270	2269	2268		9
1	30	2265	2264	2262	2261	2260	2259	2258	2257	2256	2254	3	8
1	40	2252	2250	2249	2248	2246	2245	2244	2243	2242	2241	4	7
	50	2238	2237	2235	2234	2233	2232	2230	2230	2229	2227	5	5
55	0	2925	2223	2222	2221	2219	2918	9917	9916	9915	2214	6	4
1	10	2911	2210	2208	2207	2206	9905	9904	9903	2210	2201	7	3
.1	20	2198	2196	2195	2194	2193	2191	2190	9180	9188	2187	8	1
1	30	2184	2183	2182	2180	2179	2178	2177	9176	9175	2174	9	0
	40	2171	2170	2168	2167	2166	2165	2164	2163	2169	2161	ł	
1 1	50	2158	2156	2155	2154	2153	2159	2150	2140	2148	2147	S.	Cor.
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1 1	20	9110	9117	9116	9114	9119	0110	0174	2123	2122	2108		
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1 1	50	9070	9078	9076	9075	207.1	9072	2000	2004	2083	2062		5
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57	0	2066	2005	2063	2062	2061	2060	2059	2053	2057	2056	7	3
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1 1	20	2040	2039	2037	2030	2035	2034	2033	2032	2031	2030	9	ő
1 1	30	2027	2020	2025	2023	2022	2021	2020	2019	2018	2017	9	"
1 1	40	2014	2013	2012	2010	2009	2008	2007	2006	2005	2004	ļ	
	50										1991	S.	Cor.
58	0	1989	1987	1986	1985	1984	1983	1982	1981	1980	1979	0	12
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1 1	20	1963	1962	1961	1960	1958	1957	1956	1955	1954	1953	. 2	9
1 1	30	1951	1949	1948	1947	1946	1945	1944	1943	1942	1941	3	8
1 1	40	1938	1937	1936	1934	1933	1932	1931	1930	1929	1928	4	7
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59	0	1913	1912	1910	1909	1908	1907	1906	1905	1904	1903	6	4
1 3	10	1901	1899	1898	1897	1896	1895	1893	1892	1892	1891	7	3
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1 1	30	1876	1874	1873	1872	1871	1870	1869	1868	1867	1866	9	1
1 1	40	1863	1862	1861	1859	1858	1857	1856	1855	1854	1853		
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I's Hor.	Parallax.		A	ppare	ent A	ltitud	le of					Cor. fo	BLE C. or Seconds Parallax. Add.
M	S	60°	610	620	630	640	65°	66°	670	680	690	S.	Cor.
54	0	2295	2294	2293	2292	2291	2291	2290	2289	2289	2288	0	12
-	10										2274	1	11
THE											2261	2	9
1	30										2247	3	8
1	40										2234	4	7
: 14	50										2220		5
	1175		-	_		-				-	2207	6	4
55											2194	7	3
	10											8	1
	20										2180	9	0
-3	30										2167		
											2154	S.	Cor.
1	50										2140		
56		2133	2133	2132	2131	2130	2130	2129	2128	2128	2127	0	12
	10	2120	2119	2119	2118	2117	2116	2116	2115	2114	2114	1	11
	20	2107	2106	2105	2105	2104	2103	2103	2102	2101	2101	2	9
	30										2088	3	8
	40										2075	4	7
	50	2068	2067	2066	2065	2065	2064	2064	2063	2069	2062	5	5
57	0	2055	2054	2053	2053	2059	2051	2051	2050	2049	2049	6	4
٠.	10	2042	2041	2040	2040	2039	2038	2038	2037	2036	2036	7	3
	20	2029	2028	2028	2027	2026	2025	2025	2024	2024	2023	- 8	2
	30	2016	2016	2015	2014	2013	2013	2012	2011	2011	2010	9	0
	40	2004	2003	2002	2001	2000	2000	1999	1999	1998	1998		- 4
E.	50										1985	S.	Cor.
58	0	1978	1977	1976	1976	1975	1974	1974	1973	1979	1972	0	12
-	10										1959	1	11
	20	1953	1959	1951	1950	1950	1949	1948	1948	1947	1947	2	9
	30	1940	1939	1938	1938	1937	1936	1936	1935	1934	1934	3	8
	40	1927	1927	1926	1925	1924	1924	1923	1923	1929	1921	4	7
	50	1915	1914	1913	1912	1919	1911	1911	1910	1909	1909	5	6
-	-										1896	6	4
59												7	3
19	10										1884	8	2
	20											9	1
	30										1859		475
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60											1822	0	12
00	10	1816	1815	1814	1813	1813	1812	1812	1811	1810	1810	1	11
	20	1803	1803	1809	1801	1801	1800	1799	1799	1798	1798	2	10
	30	1701	1701	1790	1789	1788	1789	1787	1787	1786	1786	3	8
	40	1770	1779	1770	1777	1776	1776	1775	1774	1774	1773	4	7
	50										1761	5	6
-	1000	A COLUMN TO SERVICE AND ADDRESS OF THE PARTY	200			_	_	-	_	_		6	5
61											1749	7	4
45	10										1737	8	2
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	30	1719	1718	1717	1716	1716	1715	1715	1714	1713	1713	9	

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App. ait.	D's Horizontal Parallax.		Tah.B ForM.
	- 100 (A)	Add.	of alt.
D M	54' 55' 56' 57' 58' 59' 60' 61'		M S
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20	41.23 41. 3 40.42 40.22 40. 2 39.41 39.21 39. 1 41.32 41.12 40.51 40.31 40.11 39.51 39.31 39.10	10 17 16 16 16 15 15 15 15 14 14 14	10074-X01-00
	41.4041.2041. 040.4040.2040. 039.4039.20	20 13 13 13 12 12 12 11 11 11 10	3 3
	41.49 41.29 41. 9 40.50 40.30 40.10 39.50 39.30	30 10 10 9 9 9 8 8 8 7 7 40 7 6 6 6 5 5 5 4 4 4	6 6
	41.58 41.38 41.18 40.59 40.39 40.19 40. 0 39.40	50 3 3 3 2 2 2 1 1 1 0	576.078 576.078
_	42. 8 41.48 41.28 41. 9 40.49 40.30 40.10 39.51		
	42.16 41.57 41.38 41.18 40.59 40.39 40.20 40. 1	10 16 15 15 15 15 14 14 14 13 18	010000000000000000000000000000000000000
20	42.25 42. 6 41.47 41.27 41. 8 40.49 40.30 40.11	20 13 12 12 12 11 11 11 10 10 10	\$ 5
	42.34 42.15 41.56 41.37 41.18 40.59 40.39 40.20	30 9 9 9 8 8 8 8 7 7 7	3 3
	12.43 42.24 42. 5 41.46 41.27 41. 8 40.49 40.30	40 6 6 6 5 5 5 4 4 4 3	9 9
50	42.51 42.33 42.14 41.55 41.36 41.18 40.59 40.40	50 3 3 2 2 2 1 1 1 1 0	9 8
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	43 .27 43 . 9 42 .51 42 .33 42 .15 41 .57 41 .39 41 .21	30 9 9 8 8 8 7 7 7 7 6	3 3
	43.36 43.18 43. 0 42.48 42.25 42. 7 41.49 41.31 43.45 43.27 43.10 42.52 42.34 42.17 41.59 41.41	40 6 6 5 5 5 4 4 4 3 3 5 50 3 3 2 2 2 1 1 1 0 0	8 8
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	14.21 44. 4 43.47 43.30 43.13 42.56 42.39 42.22	30 8 8 8 8 7 7 7 6 6 6	3 3
	4.30 44.13 43.57 43.40 43.23 43. 6 42.49 42.32	40 6 5 5 5 4 4 4 4 3 3	9 5
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	6.10 45 .55 45 .40 45 .25 45 .10 44 .55 44 .40 44 .25	30 7 7 7 7 6 6 6 6 5 5 5	134567
	6.19 46. 4 45.50 45.35 45.20 45. 5 44.50 44.35		1 8
-11-	6.28 46.14 45.59 45.44 45.29 45.15 45. 0 44.45		1
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	6.56 46.42 46.28 46.14 45.59 45.45 45.31 45.17 37 46.51 46.37 46.23 46.9 45.55 45.41 45.27 3	20 9 9 9 9 8 8 8 8 7 7 1 30 7 7 7 6 6 6 6 6 5 5 5	131
	7.14 47. 0 46.46 46.33 46.19 46. 5 45.51 45.37	00 7 7 7 6 6 6 6 6 5 5 5 5 8 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9	19
	7.23 47. 946.56 46.42 46.28 46.15 46. 1 45.47 5		8
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1 1	7.51 47.38 47.25 47.12 46.59 46.45 46.32 46.19 2	0 9 8 8 8 8 8 7 7 7 7 7	3
	8. 0 47.47 47.34 47.21 47. 8 46.55 46.42 46.29 3	0 6 6 6 6 6 5 5 5 5 5	3
	8. 9 47. 56 47. 44 47. 31 47. 18 47. 5 46. 52 46. 39 4		171
	8.18 48. 6 47.53 47.40 47.28 47.15 47. 2 46.50 5		ŝ
04	8. 28 48 . 16 48 . 3 47 . 51 47 . 38 47 . 26 47 . 13 47 . 1	0 12 12 12 11 11 11 11 11 10 10 0	0
104	R. 37 48. 25 48. 13 48. 0 47. 48 47. 36 47. 24 47. 11 1	0 10 10 10 9 9 9 9 9 8 8 2	
20 1	B.46 48.34 48.22 48.10 47.58 47.46 47.34 47.21 2	0 0 0 0 0 0 0 0 0 0 0 0	3
30 4	3.55 48.44 48.32 48.20 48. 8 47.56 47.44 47.32 3		acokso.
	9. 5 48.53 48.41 48.29 48.17 48. 6 47.54 47.42 4	0 4 4 4 3 3 3 3 8 3 2 2 7	121
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10 49	.33 49 . 22 49 . 10 48 . 59 48 . 48 48 . 36 48 . 25 48 . 14 10	0 9 9 9 9 8 8 8 8 8 8 8 8 8 8 8 8	13
20 49	.42 49 .31 49 .20 49 . 9 48 .57 48 .46 48 .35 48 .24 24	0 7 7 7 7 6 6 6 6 6 8	131
304	.51 49 .40 49 .29 49 .18 49 . 7 48 .56 48 .45 48 .34 30	0 11 11 11 10 10 10 10 10 10 10 9 0 0 9 9 9 9	0-00741761-00
50150	9 49 59 49 49 49 37 49 97 49 16 48 55 48 45 40	0 4 3 3 3 3 3 3 2 2 2 7	á
190190	. 949.5949.4849.3749.2749.1649. 648.55 5	0 2 2 1 1 1 1 1 1 1 10 0 5	19

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N's Hor	Parallax.			Appa	rent	Altitu	ide of) 's	centr	re.			r. for	Seconds rallax.
M	-	1700	710	720	730	740	750	760	770	780	790	-	S.	Cor.
54	0	2287	2287	2286	2286	2285	2285	2285	2284	2284	2284	200	0	12
-	10										2271		1	11
	20										2257		2	9
	30										2243		3	8
	40										2230		4	7
	50										2216		5	5
-		-	-	1		_		-	1		2203		6	4
55											2190		7	3
	10										2176		8	1
	20												9	0
	30										2163			100
	40										2150 2137	-	S.	Cor.
	50	16.00	200	_	-	-				100	_			
56	. 0	2126	2126	2126	2125	2125	2125	2124	2123	2123	2123		0	12
	10	2113	2113	2112	2112	2112	2111	2111	2110	2110	2110		1	11
	20										2097		2	9
	30										2084		3	8
	40			2073									4	7
	50	2061	2061	2060	2060	2059	2059	2059	2058	2058	2058		5	5
57	0	2048	2048	2047	2047	2046	2046	2046	2045	2045	2045		6	4
	10			2034									7	3
	20	2022	2022	2022	2021	2021	2021	2020	2019	2019	2019		8	2
	30	2010	2009	2009	2008	2008	2008	2007	2007	2007	2007		9	0
	40	1997	1996	1996	1995	1995	1995	1994	1994	1994	1994			200
	50			1983									S.	Cor.
58	0	1971	1971	1970	1970	1970	1970	1969	1968	1968	1968	1	0	12
00	10			1958									1	11
- 1	20			1945									2	9
- 1	30			1933									3	8
- 1	40			1920									4	7
1	50			1907									5	6
-	-	-	ALC: N	-	-	-		-	1	-	-	1	6	4
59	0			1895									7	3
	10			1882									8	2
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60	0			1821									0	12
1	10			1808									1	11
	20			1796									2	10
	30			1784									3	8
i	40			1772									4	7
	50	1761	1760	1760	1759	1759	1759	1758	1758	1758	1758		5	6
51	0	1749	1748	1748	1747	1747	1747	1746	1746	1746	1746		6	5
	10			1736									7	3
	20	1724	1724	1724	1723	1723	1723	1722	1722	1722	1722		8	2
-	30	1712											9	1

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)'s cent)'s Horizontal Parallax.	Propor. part for Sec. of Par.	Tab.B For M of alt. Add.
M	54 55 56 57 58 59 60 61	8 0"1" 2" 3" 4" 5" 6" 7" 8" 9"	MIS
0	50.19 50. 9 49.58 49.48 49.38 49.27 49.17 49.	010101010 9 9 9 9 9 9	0 0
	50.28 50.18 50. 8 49.58 49.47 49.37 49.27 49.1		2 2
	50.38 50.27 50.17 50. 7 49.57 49.47 49.37 49.2		4 4
	50.47 50.37 50.27 50.17 50. 7 49.57 49.47 49.3		9 8
	50.56 50.46 50.36 50.27 50.17 50. 7 49.57 49.41		0-00000000000
-	51. 5 50.55 50.46 50.36 50.27 50.17 50. 8 49.5		9 9
	51.15 51. 6 50.56 50.47 50.37 50.28 50.19 50.		0 0
	51.24 51.15 51. 6 50.57 50.47 50.38 50.29 50.20		3 3
	51.33 51.24 51.15 51. 6 50.57 50.48 50.39 50.30 51.42 51.34 51.25 51.16 51. 7 50.58 50.49 50.40		5 5
	51.52 51.43 51.34 51.26 51.17 51. 8 50.59 50.5		7 7
	52. 151.5251.4451.3551.2751.1851.1051.		9 9
-	52.11 52. 3 51.54 51.46 51.38 51.29 51.21 51.13		0 0
	52.20 52.12 52. 4 51.56 51.47 51.39 51.31 51.23		1 1
	52.29 52.21 52.13 52. 5 51.57 51.49 51.41 51.3		3 3
	52.39 52.31 52.23 52.15 52. 7 51.59 51.52 51.4		5 5
	52.48 52.40 52.32 52.25 52.17 52. 9 52. 2 51.54	40 3 3 2 2 2 2 2 2 2 2 2	405700
50	52.57 52.49 52.42 52.34 52.27 52.19 52.12 52.	50 1 1 1 1 1 1 1 0 0 0	9 9
0	53. 7 53. 0 52.52 52.45 52.38 52.30 52.23 52.10	0 7 7 7 7 6 6 6 6 6	9 0
	53.16 53. 9 53. 2 52.55 52.48 52.41 52.33 52.20		2 2
	53.25 53.18 53.11 53. 5 52.58 52.51 52.44 52.3	20 5 5 4 4 4 4 4 4 4	1 1
	53.35 53.28 53.21 53.14 53. 7 53. 1 52.54 52.47	30 4 3 3 3 3 3 3 3 3	200
	53.44 53.37 53.31 53.24 53.17 53.11 53. 4 52.5		8 8 9
-	53.53 53.47 53.40 53.34 53.27 53.21 53.14 53.		2 3
	54. 3 53.57 53.51 53.44 53.38 53.32 53.26 53.19		0 0
	54.12 54. 6 54. 0 53.54 53.48 53.42 53.36 53.30		2 3
	54.22 54.16 54.10 54. 4 53.58 53.52 53.46 53.40	20 4 4 4 4 4 4 3 3 3 3 3 3 3 3 3 3 3 3 3	3 2
	54.3154.2554.1954.1454.854.253.5653.51 54.4054.3554.2954.2354.1854.1254.754.1		6 6
	54.49 54.44 54.39 54.33 54.28 54.22 54.17 54.11		8 8
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	55.18 55.13 55. 8 55. 3 54.58 54.54 54.49 54.49		5 3
	55.27 55.23 55.18 55.13 55. 8 55. 4 54.59 54.59		3 3
	55.36 55.32 55.27 55.23 55.18 55.14 55. 9 55.	40 2 2 2 2 1 1 1 1 1 1	217
	55.46 55.41 55.37 55.33 55.28 55.24 55.20 55.18		9 9
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	56. 5 56. 1 55.57 55.53 55.49 55.45 55.41 55.37		1 1
20	56.14 56.11 56. 7 56. 3 55.59 55.55 55.52 55.48	20 3 3 3 3 3 2 2 2 2 2	34 4
30	56.24 56.20 56.16 56.13 56. 9 56. 5 56. 2 55.58	30 2 2 2 2 2 2 2 2 2 2	56
	56.33 56.29 56.26 56.22 56.19 56.15 56.12 56. 8	40 2 1 1 1 1 1 1 1 1 1	8 8
1-1-	56.42 56.39 56.36 56.32 56.29 56.26 56.22 56.19		9 9
	56 . 52 56 . 49 56 . 46 56 . 43 56 . 40 56 . 37 56 . 34 56 . 30		0 0
	57. 2 56.59 56.56 56.53 56.50 56.47 56.44 56.41		3 3
	57.11 57. 8 57. 5 57. 3 57. 0 56.57 56.54 56.51		1 1
	57.20 57.18 57.15 57.12 57.10 57. 7 57. 4 57. 2	30 2 2 2 2 1 1 1 1 1 1	6 6
	57.29 57.27 57.26 57.22 57.20 57.17 57.15 57.12 57.39 57.36 57.34 57.32 57.30 57.27 57.25 57.23		600
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10	57 . 49 57 . 47 57 . 45 57 . 43 57 . 41 57 . 38 57 . 36 57 . 34	0 2 2 2 2 2 2 2 2 2 2 2 2	1 1
20	57. 58 57. 56 57. 54 57. 52 57. 50 57. 49 57. 47 57. 45 58. 7 58. 6 58. 4 58. 2 58. 0 57. 59 57. 57 57. 55	10 2 2 2 2 2 2 2 2 2 1 1 20 1 1 1 1 1 1	5 3
30	58. 17 58. 15 58. 14 58. 12 58. 10 58. 9 58. 7 58. 6	30 1 1 1 1 1 1 1 1 1 1	100
40	58.26 58.25 58.23 58.22 58.20 58.19 58.18 58.16	40 1 1 1 1 1 1 1 1 1	7 7
50 5	58.35 58.34 58.33 58.32 58.30 58.29 58.28 58.27	50 1 1 1 1 1 0 0 0 0	9 9
	58.45 58.44 58.43 58.42 58.41 58.40 58.39 58.38		0 0
105	58.55 58.54 58.53 58.52 58.51 58.50 58.49 58.49	10 1 1 1 1 1 1 1 1 1	1 2
20 5	59. 4 59. 3 59. 3 59. 2 59. 1 59. 0 58.59 58.59	20 1 1 1 1 1 1 1 1 1	4 4
30 5	59.13 59.13 59.12 59.12 59.11 59.11 59.10 59.10	30 1 1 1 1 1 1 1 1 1 1	00 0
40 5	59.22 59.22 59.22 59.21 59.21 59.21 59.20 59.20	40 1 1 1 1 1 1 1 1 1 1	R 2
0015	59.32 59.32 59.31 59.31 59.31 59.31 59.31	50 1 1 1 1 1 0 0 0 0 0	318

LOGARITHMS.

)'s Hor.	Farallax.	Apparent Altitude of)'s centre.	Table C. Correction for Second of Parallax. Add.
M	S	800 810 820 830 840 850 860 870 880 89	S. Cor.
54	10 20 30 40	2283 2283 2283 2283 2283 2283 2283 2283	1 11 66 2 9 12 3 8 29 4 7
55	0 10 20 30 40	$\begin{array}{c} 2216 & 2216 & 2216 & 2215 & 2215 & 2215 & 2215 & 2215 & 22\\ 2202 & 2202 & 2202 & 2202 & 2202 & 2202 & 2202 & 2202 & 22\\ 2199 & 2189 & 2189 & 2189 & 2189 & 2189 & 2189 & 2189 & 2189 & 2189 & 2189 & 2175 & 2$	02 08 08 08 08 08 08 08 08 08 08 08 08 08
	50	2136 2136 2136 2136 2136 2135 2135 2135 2135 21	S. Cor.
56	10 20 30 40	$\begin{array}{c} 2123 \ 2123 \ 2122 \ $	9 1 11 66 2 9 3 3 8 70 4 7
57	10 20 30 40	2044 2044 2044 2044 2044 2044 2044 2044	7 3 8 8 9 0
58	0 10 20 30 40	1968 1968 1967 1967 1967 1967 1967 1967 1967 1967	77 0 12 15 1 11 22 9 3 8 7
59	0 10 20 30 40	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6 4 3 9 7 8 2 1 1
60	0 10	1830 1830 1830 1830 1830 1830 1830 1830	8 0 12 5 1 11
,	30 40	1793 1793 1793 1793 1793 1793 1793 1793 1793 1794 1791 1791 1791 1791 1791 1791 1791 1791 1791 1769 1769 1769 1769 1769 1769 1769 1769 1769 1769 1767 1757 17	3 8 7
61	0 10 20	1745 1745 1745 1745 1745 1745 1745 1745	5 6 5 4 3 1 8 2 1
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Corr	.						t Dist	ance.			40		50
Tab.	19	 -	00		10		20	<u> </u>	30				- 1
Cort		Tab.19	Tab.19. +2Cor	Tab.19.	Tab.19. +2Cor	Tab.19	Tab. 19. +2Cor	Tab.19	Tab.19. +2Cor	Tab.19	Tab.19. +2Cor	.b.19	Tab.19. +2Cor
+20		E	<u> </u>		<u>E</u> +		احستا		4+		4	Tab.	<u>+</u>
M	M.	"				"	"	"	"	"	"	100	117
0 1	120 119	196 190	178 172	179 174	161 156	166 161	148 143	154 149	136 131	144 140	126 122	135 131	117
2	118	184	166	169	151	156	138	145	127	136	118	127	109
3	117	178 173	160 155	164 159	146 141	151 147	133 129	141 136	123 118	132 128	114 110	124 120	106 102
5	115	167	149	154	136	142	124	132	114	124	106	116	98
6	114	162	144	149	131	138	120	128	110	120	102	113	95
7 8	113 112	157 152	139 134	144 139	126 121	133 129	115 111	124 120	106 102	116	98 95	109 106	91 88
9	iii	146	128	135	117	125	107	116	98	109	91	103	85
10	110	142	124	130	112	121	103	112	94	105	87 84	99 96	81
11 12	109 108	137 132	119 114	126 121	108 103	116 112	98 94	109 105	91 87	102 99	81	93	78 75
13	107	127	109	117	99	109	91	101	83	95	77	90	72
14	106	123	105	113	95	105	87 83	98	76	92 89	74	87	69
15 16	105 104	118 114	100 96	109 108	91 87	97	79	91	73	86	68	81	63
17	103	109	91	101	83	94	76 72	88 85	70 67	83 80	65 62	78 75	60 57
18 19	102 101	105 101	87 83	97 93	79 75	90 87	69	81	63	77	59	73	55
20	100	97	79	90	72	84	66	78	60	74	56	70	52
21 22	99 98	93 89	75 71	86 83	68 65	80 77	62 59	75 73	57 55	71 69	53 51	68 65	50 47
23	97	86	68	79	61	74	56	70	52	66	48	63	45
24	96	82	64	76	58	71	53	67	49	63	45	60	42
25 26	95 94	79 75	61 57	73 70	55 52	68 65	50 47	64 62	46 44	61 58	43	58 56	40 38
27	93	72	54	67	49	63	45	59	41	56	38	53	35
28 29	9 2 91	69 66	51 48	64 61	46 43	60 57	42 39	57 54	39 36	54 52	36	51 49	33 31
30	90	63	45	58	40	55	37	52	34	49	31	47	29
31	89	60	42	56	38	53	35	50	32	47	29	45	27
32 33	88 87	57 54	39 36	53 51	35 33	50 48	32 30	48 46	30 28	45 44	27 26	44	26 24
34	86	51	33	48	30	46	28	44	26	42	24	40	22
35	85	49	31	46	28	44	26	42	24	40	22	38	20
36 37	84 83	46	28 26	44	26 24	42 40	24 22	40 38	22 20	38 37	20 19	37 35	19 17
38	82	42	24	40	22	38	20	36	18	35	17	34	16
39 40	81	38	22	38	20 18	36	18	35 33	17	33	15	32	14
41	79	36	18	34	16	33	15	32	14	31	13	30	13
42	78	34	16	33	15	31	13	30	12	29	11	29	11
43 44	77 76	32 31	14 13	31 29	13 11	30 29	12 11	29 28	11 10	28 27	10 9	27 26	9 8
45	75	29	11	28	10	27	9	27	9	26	8	25	7
46 47	74 73	28 26	10 8	27 26	9 8	26 25	8 7	25 94	7	25 24	7	24 24	6
48	72 71	25	7 6	24	6 5	24	6 5	23	6 5 5	23 22	6 5 4	23	5
49		24		23		23		23					
50 52	70 68	23 21	5 3 1 0	22 21	4 3 1 0	22 21	4 5	22	4 2	21 20	3 2 1	21 20	3 2 1
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60	-60	18	0	18		18	0	18	0	18	0	18	0
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	l	T	2	E E	44	Te		1	44	Te I	44	F	
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li	Tab.		10	5 0	1	70		80		90	2	00	2	10	2	2 0
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	M	M.	"	"	"	"	7	"	"	"	"	"	-	"	"	
	0	120 119	127 124		121 117	103 99	115	97 93	109 106		104 101	83	100 97	82 79	96 93	78 75
	2 3 4	118 117 116	120 117 113	102 99 95	114 111 107	96 93 8 9	108 105 102	90 87 84	103 100 97	85 82 79	99 96 93	81 78 75	94 92 89	76 74 71	91 88 86	73 70 68
ŀ	5	115	110	92	104	86	99	81	95	77	90	72	87	69	83	65
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	8 9	112 111	100 97	8 2 79	95 92	77 74	91 88	73 70	86 84	68 66	83 80		79 77	61 59	76 74	58 56
	10 11	110 109	94 91	76 73	89 86	71 68	85 82	67 64	81 79	63 61	78 76	60 58	75 73	57 55	72 70	54 52
	12	108	88	70	84	66	80	62	76	58	73	55	70	52	68	50
	13 14	107 106	85 8 2	67 64	81 78	63 60	77	59 57	74 72	56 54	71 69	53 51	68 66	50 48	66	48 46
	15	105	80	62	76	58	72	54	69	51	67	49	64	46	62	44
ı	16 17	104 103	77 74	59 56	73 71	55 53	70 68	52 50	67 65	49 47	64 62	46	62 60		60 58	42
l	18	102	72	54	68	50	65	47	63	45	60	42	58	40	56	38
	19	101	69 67	51 49	66	48	63	45	61 59	43	58 56	40 38	56	38 36	53	35
	21	99	64	46	61	43	59	41	57	39	54	36	53	35	51	33
١.	22 23	98 97	62 60	44	59 57	41 39	57 55	39 37	55 53	37 35	53 51	35 33	51 49	33 31	49 48	31 30
	24	96	57	39	55	37	53	35	51	33	49	31	47	29	46	28
	25 26	95 94	55 53	37 35	53 51	35 33	51 49	33 31	49 47	31 29	47	29 28	46 44	28 26	44	26 25
	27	93	51	33	49	31	47	29	46	28	44	26	43	25	42	24
	28 29	92 91	49 47	31 29	47 45	29 27	45 44	27 26	44 42	26 24	43	25 23	40	23 22	40 39	22 21
	30	90	45	27 26	44	26 24	42 41	24 23	41 39	23 21	40 38	22 20	38 37	20 19	37 36	19
	31 32	89 88	44	24	40	22	39	21	38	20	37	19	36	18	35	18 17
	33 34	87 86	40 39	22 21	39 37	21 19	38 36	20 18	36 35	18 17	35 34	17 16	35 33	17 15	34 33	16 15
	35 36	85 81	37 36	19 18	36 34	18 16	35 33	17 15	34 33	16 15	33 32	15 14	32 31	14	31 30	13 12
	37	83	34	16	33	15	32	14	31	13	31	13	30	12	29	11
	38 39	82 81	33 31	15 13	32 31	14 13	31 30	13 12	30 29	12 11	30 29	12 11	29 28	11 10	28 28	10 10
	40 41	80 79	30 29	12 11	2 9 28	11 10	29 28	11 10	28 27	10	28 27	10 9	27 26	9	27 26	9
	42	78	28	10	27	9	27	9	26	8	26	8	25	7	25	7
	43 44	77 76	27 26	9 8	26 25	8 7	26 25	8 7	25 24	7 6	25 24	7 6	25 24	6	24 24	6 6
	45 46	75 74	25 24	7	24 24	6	24 23	6	24 23	6 5	23 23	5 5	23 22	5 4	23	5 4
-	47	73	23	5	23	5	23	5	22	4	22	4	22	4	22	4
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7 8	113	35 35	17 17	35 34	17 16	34 33	16 15	33 33	15 15	32	15 14	32		29	ii	27	9
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19	101	28	10	28	10	28	10	27	9	27	9	26	8	25	7	23	5
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21	99	27	9	27	9	27	9	26	8	26	8	26	8	24	6	23 23	5
22 23	98 97	27	9	26 26	8	26 26	8 8	26 25	8	26 25	8 7	25 25	7	24 24	6	22	4
23 24	96	26 26	8	26	8 8	25	7	25	7	25	7	25	7	23	5	22	4
25	95	25	7	25	7	25	7	25	7	24	6	24	6	23	5	22	4
26	94	25	7	25	7	25	7	24	6	24	6	24	6	23	5	22	4
27	93	25	7	24	6	24	6	24	6	24	6	23		22	4	21	3
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29	91	24	6	24	6	23	5	23	-5	23	5	23	5	22	4	21	3
30 31	90 89	23 23	5	23 23	5	23 23	5 5	23 23	5	23	5	22	4	21	3	21	3
31 32	88	23	5	23	5	22	4	22	4	22	4	22	4	21	3	20	2
33	87	22	4	22	4	22	4	22	4	22	4	22	4	21	3	20	2
34	86	22	4	22	4	22	4	22	4	22	4	21	3	21	3	20	2
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37 38	83 82	21 21	3 3	21	3	21 21	3	21 21	3	21	3	20		20	2	20	2 1
39	81	21	3	21	3	20	2	20	2	20	2	20	2	20	2	19	ī
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41	79	20	2	20	2	20	2	20	2	20	2	20	2	19	1	19	1
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5 115	25	7	23	5	20	2	18		2	20	5	23	7	25	10	28	12	30	15	33
6114 7113	25 25	7	22 22	4	20 20	2 2	18 18	0	2 2	20 20	4	22 22	7 7	25 25	9	27 27	12 11	30 29	15 14	33 32
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11 109	24 24	6	22	4	20	2	18	0'	2	20	4	22	6	24	8	26	10	28 28	12	30
12 108 13 107	11 1	5 5	22 21	3	20 20	2 2	18 18	0	2 2	20 20	4 3	22 21	5 5	23 23	7	25 25	9	27 27	12 11	30 29
14 106	23 23	5	21	3	20	2	18		2	20	3	21	5	23	7	25 25	9	27	11	29
15 105	23	8	21	3	20	2	18	ō	2	20	3	.21	5	23	6	24	8	26	10	28
16 104 17 103	23 22	5	21 21	3	19 19	1	18 18	0	1	19 19	3	21 21	5 4	23 22	6	24	8 8	26 26	10 9	28 27
18 102 19 101	22	4	21 21	3	19	1	18 18	0	1	19	3	21 21	4	22 22	6	24	7	25	9	27 26
20 100	22	4	21 20	2	19 19	1	18	0	-	19	2	20	4	22	5	23	7	25	8	26
21 99	22	4	20	2	19	1	18	0	1	19	2	20	4	22	5	23	6	24	8	26
22 98 23 97	21 21	3	20 20	2 2	19 19	1	18 18	0	1	19 19	2 2	20 20	3	21 21	5	23 22	6	24	7	25 25
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34 80	400	2	19	1	19		18	0	1	19	1	19	2	20	2	20	3		3	
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36 84 37 83 38 85	19	1	19	1	18	0	18	0	O	18	1	19	1	19	2 2 2 1	20	2	20	3	21
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TABLE XXI.

For turning Degrees and Minutes into Time, and the contrary.

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M	MS	M	мэ	M	MS	M	MS	M	MS	M	MS
1	0. 4	61	4. 4	121	8.4	181	12. 4	241	16. 4	301	20. 4
2	0.8	62	4. 8	122	8.8	182	12. 8	242	16. 8	302 303	20. 8
3	0.12	63 64	4.12	123 124	8.12 8.16	183 184	12.12 12.16	243 244	16.12 16.16	303	20.12
5	0.20	65	4.20	125	8.20	185	12.20	245	16.20	305	20.20
6	0.24	66	4.24	136	8.24	186	12.24	246	16.24	306	20.24
7 8	0.28 0.32	67 68	4.28	127 128	8.28 8.32	187	12.28 12.32	247 248	16.28 16.32	307 308	20.28 20.32
9	0.32	69	4.36	129	8.36	189	12.36	249	16.36	309	20.36
10	. 0.40	70	4.40	130	8.40	190	12.40	250	16.40	310	20.40
11	0.44	7]	4.44	131	8.44	191	12.44	251	16.44	311	20 44
12	0.48	72	4.48	132	8.48	192	12.48 12.52	252	16.48	312 313	20 48 20.52
13 14	0.52 0.56	73 74	4.52	133	8.52 8.56	193 194	12.52	253 254	16.52 16.56	313	20.56
15	1. 0	75	5. 0	135	9. 0	195	13. 0	255	17. 0	315	21.0
16	1. 4	76	5. 4	136	9. 4	196	13. 4	256	17. 4	316	21. 4
17 18	1.8 1.12	77 78	5. 8 5.12	137 138	9.8 9.12	197 198	13. 8 13.12	257 258	17.8 17.12	317 318	21.8
19	1.16	79	5.16	139	9.16	199	13.16	259	17.16	319	21.16
20	1.20	80	5.20	140	9.20	200	13.20	260	17.20	320	21.20
21	1.24	81	5.24	141	9.24	201	13.24	261	17.24	321	21.24
22	1.28	82	5.28	142	9.28	202	13.28 13.32	262	17.28	322 323	21.28
23	1.32	83 84	5.32 5.36	143 144	9.32 9.36	203	13.32	263 264	17.32 17.36	323 324	21.32 21.36
25	1.40	85	5.40	145	9.40	205	13.40	265	17.40	325	21.40
26	1.44	86	5.44	146	9.44	206	13.44	266	17.44	326	21.44
27	1.48	87	5:48	147	9.48	207	13.48 13.52	267 268	17.48 17.52	327 328	21.48 21.52
28 29	1.52 1.56	88 89	5.52	148 149	9.52 9.56	200	13.56	269	17.56	329	21.56
30	2. 0	90	6. 0	150	10. 0	210	14. 0	270	18. 0	330	22. 0
31	2. 4	91	6. 4	151	10. 4	211	14. 4	271	18. 4	531	22. 4
32	2.8	. 92	6. 8	152	8 .01	212	14. 8	272	18. 8	332	22. 8
33 34	2.12 2.16	93 94	6.12	153 154	10.12 10.16	213 214	14.12 14.16	273 274	18.12 18.16	553 534	22.12 22.16
35	2.20	95	6.20	155	10.20	215	14.20	275	18.20	335	22.30
36	2.24	96	6.24	156	10.24	216	14.24	276	18.24	336	22.24
37	2.28 2.32	97 98	6.28	157 158	10.28 10.32	217 218	14.28 14.32	277 278	18.28 18.32	337 338	22.28 22.32
39	2.36	99	6.36	159	10.32	219	14.36	279	18.36	339	22.36
40	2.40	100	6.40	160	10.40	220	14.40	280	18.40	340	22.40
41	2.44	101	6.44	161	10.44	221	14.44	281	18.44	341	22.44
42	2.48	102	6.48	162	10.48	222 223	14.48	282	18.48	342 343	22.48 22.52
43 44	2.52 2.56	103 104	6.52 6.56	163 164	10.52 10.56	223	14.52 14.56	283 284	18.52 18.56	344	22.56
45	3. 0	105	7. 0	165	11.0	225	15. 0	285	19. 0	345	23. 0
46	3. 4	106	7. 4	166	11. 4	226	15. 4	286	19. 4	346	23. 4
47	3. 8 3.12	107 108	7. 8 7.12	167 168	11.8 11.12	227 228	15. 8 15.12	287 288	19.8 19.12	347 348	23. 8 23.12
49	3.16	108	7.12	169	11.12	229	15.12	289	19.12	349	23.16
50	3.20	110	7.20	170	11.20	230	15.20	290	19.20	350	23.20
51	3.24	111	7.24	171	11.24	231	15.24	291	19.24	351	23.24
52	3.28	112	7.28	172	11.28	232	15.28	292	19.28	352	23.28
53 54	3.32 3.36	113 114	7.32 7.36	173 174	11.32 11.36	233 234	15.32 15.36	293 294	19.32 19.36	353 354	23.32 23.36
55	3.40	115	7.40	175	11.40	235	15.40	295	19.40	355	23.40
56	3.44	116	7.44	176	11.44	236	15.44	296	19.44	356	23.44
57	3.48	117	7.48	177	11.48	237	15.48	297	19.48	357	23.48
58 59	3.52 3.56	118 119	7.52	178 179	11.52 11.56	238 239	15.52 15.56	298 299	19:52 19:56	358 359	23.52 23.56
60	4. 0	120	8, 0	180	12. 0	240	16. 0	300	20. 0	360	24. 0

TABLE XXII.

PROPORTIONAL LOGARITHMS.

S.	, h 00	m O	h no	m	ъ Р	m 2/	þ 0°	12h 3/	h O	m 4	р 00	на 5/	h 00	m 6':	h ص	m 7/	00 h	m 8/	S.
0	*	-		2553	_	.9542	_	.7782		6532	_	.5563	_	.4771	-	4102	<u>'</u>	3522	0
ĭ	4.0	334		2481		.9506		7757		6514		.5549		.4759		.4091	_	. 3513	1
2		324		2410		.9471		.7734		6496		. 5534		.4747		.4081		3504	2
. 3		663		2341		.9435		.7710		.6478		. 5520		.4735		.4071		.3495	5
4	3.4	314	2.	2272	ł.	. 9400	1	.7686	1	. 64 60	1	. 5506	1	.4723	1	.4061	1	.3486	. 4
-5	3.3	345	2.	2205	1	.9365	1	.7663	1	.6443	1	.5491	1	.4711	1	.4050	1	.3477	
6		553		2139		.9331		.7639		6425		.5477	1 7	.4699		.4040		.3468	ě
7		833	2.	2073		. 9296		.7616		.6407		. 5463		.4688	1	.4030	1	. 3459	,
. 8		303		2009		.9262		.7593	1	.6390	1	. 5449		.4676		. 4020		. 3450	8
9	3.0	792	2.	. 1946	1	.9228	1	.7570	1	.6372	1	.5435	1	.4664	1	.4010	1	.3441	_ 9
10	3.0	331	2.	1883	1	.9195	1	.7547	1	.6355	1	.5121	1	.4652	1	.4000	1	.3432	10
11	2.9	920		1822	1	.9162	1	.7524	1	. 6338	1	.5407		.4640	ı	.3989	1	. 3423	11
12	2.9	542	2.	1761		.9128	1	.7501	1	. 6320	1	. 5393	1	.4629	1	.3979	1	.3415	15
13,		195		1701		.9096		.7479		. 63 03		. 5 379		.4617		.3969		.3406	1:
14	2.8	873	2.	1642	1	. 9063	1	.7456	. 1	. 6286	1	.5366	1	. 4606	1	. 3 959	1	. 3397	14
15	2.8	573	2.	1584	1	.9031	1	.7434	1	6269	ī	.5351	1	4.594	1	.3949	1	.3388	1.
16		293		1526		. 8999		.7412		.6252		. 5337		. 45 82		. 3939		.3379	10
17		030		1469		. 8967		.7390		.6235		.5324		.4571		. 3929	1	.3371	1'
18		782		. 1413		.8935		.7358		.6218		.5310		.4559		.3919		.3362	11
19		547	2	. 1358	1	8904	1	.7346	1	.6201	1	. b 296	_	. 454 8		.8910	1	.3353	1
20		324		. 1303		.8873		. 7324		.6185		. 5283		.4536		.3900		.3345	2
21		112		. 1249		.8842	1	.7302		.6168		.5269		.4525		.3390		.3386	2
22		910		. 1196		.8811		.7281		.6151		.5256		.4514		.3880		.3327	2
23		717		.1143		.8781		.7259		.6135		.5242		.4502		.3870		.3319	
24		532		. 1091	;	.8751		.7238		.6118		. 5 229		.4491	-	.3860	-	.3310	2
25	4	355		. 1040		.8721		.7217		.6102		.5215		.4480		.3851		.3301	2
26		185		.0989		.8691		.7196		.6085		.5202		.4468		.3841		.3293	20
. 27		021		.0939		.8661		.7175		.6069		.5189		.4457		.3831		.3284	2
28 29		863		.0889		.8632		.7154		. 6053		,5175		.4446		.3821		.3276	2:
		710		0:40	-	.8602		.7133	I	6037	-	.5162	_	.4435	-	.8812	-	.3267	-
30		563		0792		.8573		.7112		.6021		.5149		.4424		. 3802		.3259	30
31	2.5			0744		8544		.7091				.5136		.4412		.3792		.3250	8
32 33		283		0696		8516		.7071		.5989		.5128		4300		.3783		.3242 .3233	3:
34		149 019		0649 0603		. 84 87 . 8459		. 7 050 . 70 30		. 5 973 . 5957		.5110 .5097		. 4 39 0 . 4 379		. 3 773 . 3 764		.3225	3.
							_				-		_						
35		894		0557	-	.8431		.7010		.5941		.5084		.4368		.3754		.3216	30
36 37	2.4	652		0512 0467		· 8403 · 8375		.6990 .6970		-59 2 5 - 5 909		.5071 .5058		.4357 .4346		.3745 .3786		.3208 .3199	3
38		536		0422		. 834 8		. 6 950		. 5 894		.5045		4335		. 372 6		.3191	31
39		424		0378		. 8320		.6930		. 5 878		.5032		.4325	i	3716	١i	.3183	3
40		314	_	0334	-	. 8293		.6910	_	.5863	ļ	.5019	_	.4314				.3174	4
40 .		206		.0334 .0291		. 8293 . 8266		.6890		. 5663 . 5847		.5007		.4303		.3707 .3697		.3166	4
42		102		.0248		. 8239		.6871		. 5 832		.4994		.4292		. 3 688		.3156	4
43		000		0206		.8212		.6851		. 5816		.4981		.4281		.3678		.3149	
44		900		0164		.8186		.6852		. 5801		4969		.4270		.3669		.3141	4
45		802		0122		.8159	_	.6812	١	.5786		.4956		.4260	I	.3660	1	.3133	4
46		707		0081		.8133		.6793		. 577 1		.4943		4249		.3650		.3124	44
47		613		0040		.8107		.6774		.5755		.4931		.4238		.3641		.3116	. 4
48		522		0000		.8081		.6755		. 5740		.4918		. 4228		.3632		.3108	4
49	2.9	432	1.	9960	1	. 8056		.67 3 6		. 5725	1	.4906	1	.4217	1	. 3623	1	.3100	49
50	2.1	345	1	9920	1	.8030	ī	.6717	1	.5710	1	.4894	ī	.4206	1	.3613	1	.3091	50
51		259		9881		.8004		.6698		. 5695		.4881		.4196		.3604	1 .	.3063	5
52		174		9842		.7979		.6679		5680		.4869		.4185		. 3595	1	3075	55
53		091	1.	. 9803	1	.7954	1	.6661		. 5 666		4856		.4175		. 3586		.3067	53
54	2.5	010	1.	.9765	1	.7929	1	.6642	1	. 5651	1	. 4 844	1	.4164	1	. 3 576	1	. 3059	54
55	2.2	931	1	9727	1	.7904	1	.6624	1	. 5636	1	.4832	1	.4154	1	.3567	1	.3051	54
56		852	_	9690		.7879		.6605		. 5621		.4820	1	.4143		. 3558		.3043	56
57		775	1.	9652	1	.7855	1	.6587	1	. 5607		.4808		.4133		.3549		.3034	5
58		70u		. 9615		.7830		.6568		. 5592		.4795		.4122		.3540		.3026	58
59	2,2	626	1.	9679		.7806		.6550	<u>. </u>	. 5578		.4783		.4112	_	. 3531	1	.3018	59
			00		<u>₀</u>		00		00		00	- 1	00	CI	00	7	00	- 8/	S.

Digitized by GOOQL

TABLE XXII.

PROPORTIONAL LOGARITHMS.

ī	S.	b b	m 9/	00		h 0°.	m 11'		m 12	0о р		h 0°	. m	o p		o°	m 16'	h 0°	m 17	S.
-	0		3010		. 2553		2139	1.	1761	1	. 1413		1091		0792		0512	1	.0248	0
	1		3002 2994		. 2545 . 2538		2132 2126		1755 1749		. 1408 . 1402		1086 1081		.0787 .078 2		.0507 .0502		. 0244 . 0240	1 2
1	. 2 3		2 936		.2531		2119		1748		.1397		1076		0777		0498		.0235	3
	4		2978		. 2524		2113		1737		. 1391		1071	1.	.0773	Ť.	0493	1	.0231	4
1	5		2970		.2517		2106		1731		. 1386		1066		0768		0489	_	.0227	5
1	6 7		2962 2954		. 2510 . 2502		2099 2093		1725 1719		. 1380 . 1374		1061 1055		.076 3 .0758		0484 0480		.0223 .0219	6
1	8		2946		2495		2086		1713		. 1369		1050		0758		0475		.0214	8
1	9		2939		.2488		2080		1707		. 1363		1045		0749		0471		.0210	9
1	10		2 931		.2481		2073		1701		. 1358		1040	-	0744		0467	_	. 0206	10
1	11 12		2923 2915		. 2474 . 2467		2067 2061		.1695 .1689		. 1352 . 1347		. 1035 . 1036		.0739 .0734		.0462 .0458		.0202 .0197	11
1	13		2907		.2460		2054		1683		. 1342		1025		0730		0453		0193	13
1	14		2899		.2453		2048		1677		. 1386		1020		0725		0449	1.	0189	14
Γ	15		2891		.2445		2041		1671		. 1331		1015		0720		0444		.0185	15
	16 17		2883 2876		. 243 8 . 243 1		2035 2028		1665 1660		. 1325 . 1320		. 1009 . 1004		.0715 .0711		.0440 .0435		.0181 .0176	16 17
1	18		2 868		2424		2022		1654		.1314		0999		0706		.0431		.0172	18
١.	19	1.	2860	1	.2417		2016		1648		. 1309		0994		07 01	1	.0426	1	.0168	19
ľ	20		2852		.2410		2009		1642		.1303		0989		0696		0422		.0164	20
1	21 22		2845 2837		.2403 .2396		2003 1996		1636 1630		. 1298 . 1292		.0984 .0979		.0692 .0687		.0418 .0413		.0160 .0156	21 22
1	23		2829		.2389		1990		1624		.1287		.0974		.0682		.0409		.0151	23
1	24		2821		.2382		1984		1619		. 1282		0969		0678		.0404		.0147	24
	25		2814		.2375		1977		1613		.1276		0964		0673		.0400		.0143	25
	26 27		2806 279 8		. 2368 . 2362		1971		1607 1601		.1271 .1266		.0959 .0954		.0668 .06 63		. 0395 . 03 91		.0139 .0135	26 27
1	28		2791		.2355		1965 1958		1595		. 1260		0934		.0659		.0387		.0131	28
1	29		2783		.2348		1952		1589		. 1255		0944		0654		.038%		.0126	29
-	30		2775		.2341		1946		1684		.1249		0939		0649		0378		.0122	30
1	31 32		2768 2760		. 2334 . 2327		19 3 9		1578 1572		. 1244 . 1239		.0934 .0929		.0645 .0640		.0374 .0369		.0118 .0114	31 32
1	32 33		2753		.2320		1933		1566		. 1239 . 1233		0929		. 06 3 5		-0365		.0110	33
	34		2745	1	. 2313		1921		1561		. 1 22 8		0919		0631		.0360		.0106	34
	35		2738		.2307		1914		1555		. 1223		0914		0626		.0356		.0102	35
1	36 37 .		2730 2722		. 2 300 . 22 93		1908 1902		1549 1543		. 1217 . 1212		.0909 .0904		.0621 .0617		. 0352 . 0347		.0098 .0093	36 37
	38		2715		.2286		1896		1538		.1207		.0899		0612		.0343		.0089	38
	39	1.	2707	1	. 227 9	1.	1889	1.	1532	1	. 1201		0894	1.	0608	1	.0339	1	.0085	39
1	40		2700		.2272		1883		1526		.1196		0889		0603		.0334		.0081	40
	41 42		2 692 26 85		. 22 66 . 22 59		1877 1871		1520 1515		.1191 .1186		.0884 .0880		0598 0594		.03 30 .03 26		.0077 .0073	41
1	43		2678		. 2252		1865		1509		.1180		0875		0589		.0321		.0069	43
	44	1.	2670	1	. 2245	1.	1859	1.	1503	1	.1175	1.	.0870	1.	0585	1	0317	1	.0065	44
Γ	45		2663		. 2239		1852		1498		.1170		0865		0580		0313	_	.0061	45
	46 47		2 655 2 648		. 2232 . 2225		1846 1840		1492 1486		. 1164 . 1159		0860 0855		0575 0571		.0308 .0304		.0057 .0053	46
	48		2640		.2218		1834	1.	1481		.1154		0850		0566		0300		.0049	48
	49		2633		.2212	1.	1828	_	1475	-	.1149	1.	0845	1.	0562		0295		.0044	49
Γ	50		2626		. 2205		1822		1469		.1143		0840		0557		0291		.0040	50
l	51 52		2618 2611		.2198 .2192		1816 1809		1464 1458		. 11 3 8 . 11 3 3		0835 0831		0552 0548		0287 0282		.0036 .003£	51 52
	53		2604		.2185		1803		1452		.1128		0826		0543		0278	_	.0028	53
	54		2596	-	.2178	1.	1797		1447	-	.1123	_	0821	_	0539	1.	0274	1	.0024	54
	55		2589		.2172		1791		1441		.1117		0816		0534		0270		.0020	55
	56 57		2582 2574		.2165 .2159		1785 1779		1436 1430		.1112 .1107		0811 0806		0530 0525		0265 0261		.0016 .0012	56 57
1	58	1.	2567	1	2152	1.	1773	1.	1424	1	. 1102	1.	0801		0521		0257		.0008	58
_	59		2560		.2145		1767		1419		. 1097		0797		0516		0252		.0004	59
_	S.	0 0	9'	00	10	00	11/	0°	12/	<u>0</u> 0	13/	<u>00</u>	14'	<u>o</u> °_	15/	0 °	16'	<u>0</u> 0	17	S .

TABLE XXII.

PROPORTIONAL LOGARITHMS.

1-	i h m	h m	h m	h m	h m	h m	h m	h m l	h m	h m	h m	h m /	
8.			0° 20′	0° 21′	00 22/				00 26		0 [ୁ] 28∕		8.
0	10000		9542	9331	9128	8935	8751	8573	8403	8239	8081	7929	0
1 2	9996 9992	9761 9758	9539 9535	9327	9125	8932	8748	8570	8400	8236 8234	8079	7926	1
3	9992	9754	9532	93 24 93 2 0	91 22 9119	8929 8926	8745 8742	8568 8565	8397 8395	8234 8231	8076 8073	7924 7921	2
4	9984	9750	9528	9317	9115	8923	8739	8562	8392	8228	8071	7919	4
5	9980	9746	9524	9313	9112	8920	8736	8559	8389	8226	8068	7916	5
6	9976	9742	9521	9310	9109	8917	8733	8556	8386	8223	8066	7914	6
8	9972 9968	9739 9735	9517 9514	9306 9303	9106 9102	8913 8910	8730 8727	8553 8550	8384 8381	8220 8218	8063 8061	7911 7909	8
9	9964	9731	9510	9300	9099	8907	8724	8547	8378	8215	8058	7906	9,
10	9960	9727	9506	9296	9096	8904	8721	8544	8375	8212	8055	7904	10
111	9956	9723	9503	9993	9092	8901	8718	8542	8372	8210	8053	7901	11
12	995 2 9 94 8	9720 9716	9499 9496	9 2 89 928 6	9089	8898	8715	85 3 9 8536	8 37 0 8 36 7	8207 8204	8050 8048		12 13
14	9944	9712	9482	9283	9086 9083	8895 8892	8712 8709	8533	8364	8202	8045		14
15	9940	9708	9488	9279	9079	8888	8706	8530	8361	8199	8043	7891	15
16	9936	9795	9485	9276	9076	8885	8703	8527	8359	8196	8040		16
17	9982	9701	9481	9272	9073	8682	8700	8524	8356	8194	8037	7887	17
18	9928 9924	9697 9693	9478 9474	9 269 9 26 6	9070 9066	8879 8876	8697 8694	85 22 8519	8353 8350	8191 8188	3035 8032		18 19
20	9920	9690	9471	9262	9063	8873	8691	8516	8348	8186	8030		20
21	9916	9636	9467	9269	9060	8870	8688	8513	8345	8183	8027	7877	21
22	9912	9682	9464	9255	9057	8867	8685	8510	8342	8181	8025	7874	22
23	9908 9905	9678 9675	9460 9456	92 <i>5</i> 2 9249	9053 9050	8864 8861	8682 8679	8507 8504	8339 8 337	8178 817 <i>5</i>	8022		23
											8020		24
25 26	9901 9897	9671	9453 9449	9245 92 42	9047 9044	8867 8864	8676 8678	8502 8499	8334 8331	8173 8170	8017 8014	7867 7864	25 26
27	9893	9664	9446	9238	9041	8851	8670	8496	8528	8167	8012	7862	27
28	9889	9660	9442	9235	9037	8848	8667		8326	8165	8009	7859	28
29	9885	9656	9439	9232	9034	8845	8664	8490	8323	8162	8007	7857	
30	9881 9877	9652 9649	9435 9432	9228	9031 9028	8842	8661	8487 8484	8320 8318	8159	8004 8002	7855	30
31 32	9873	9645	9428	9225 9222	9024	8839 8836	8658 8655	8482	8315	8157. 8154	7999	7852 7850	31 32
33	9869	9641	9425	9218	9021	8833	8652	8479	8312	8152	7997	7847	33
34	9865	9638	9421	9215	9018	8880	8649		8309	8149	7994	7845	.34
35	9861	9634	9418	9212	9015	8827	8646	8473	8307	8146	7992	7842	35
36 37	9858 9854		9414 9411	9208 9205	9012 9008	. 8 824 8821	8643 8640	8470 8467	8304 8301	8144 8141	7989 7987	7840	36 37
38	9850		9407	9901	9005	8817	8637	8465	8298	8138	7984		38
39	9846	9619	9404	9198	9002	8814	8635	8462	8296	8136	7981	7532	39
40	. 9842	1	9400	9195	8999		8632	8459	8293	8133	7979	7830	40
41	9838		9397	9191	8996	8808	8629		8290	8131	7976		41
42	9834		9 5 93		8992 8989	8805 880 2	8 62 6	8451	8288 8285	8128 8125	7974 7971	7823	42
14	9827	9601	9386	9181	8986	8799	8620	8448	8282	8123	7969		44
45	9823		9383		8983	8796	8617	8445	8279	8120	7966	7818	45
46	9819		9379	9175	8980		8614	8442	8277	8117	7964	7815	46
47	9815	9590 9586	9376 9372	9171 9168	8977 8973	8790 8787	8611 8608	8439 8437	8274 8271	8115 811 2	7961 7959	7813 7811	47 48
49	9807		9369		8970		8605	8434	8269	8110	7956		49
50	9803	9679	9365	9162	8967	8781	8602	8431	8266	8107	7954	7806	50
51	9800	9575	9362	9158	8964	8778	8599	8428	8263	8104	7951	7803	51
52 53	9796		9358 9355		8961 8958	8775 8772	8597 8594	8425 8423		810 2 8099	7949 7946		52 53
54	9792		9351	9152 9148	8954		8591	8423	8255	8097	7946		54
55	9784	9561	9348		8951	8766	8588	8417	8253	8094	7941	7794	55
56	9780		9344		8948	8763	8585	8414	8250	8091	7939	7791	<i>5</i> 6
57	9777		9541	9138	8945			8411	8347	8089	7936		57
58 59	9773		9337 9334		8942 8959		8579 8576	8409 8406	8 244 8 242	8086 8084	7934 7931		58 59
8.		00 19				1							- S .
	10 10			·	J . A.A.		15 NA		Digitized	by G	ÖÖĞ	te "	

TABLE XXIL

PROPORTIONAL LOGARITHMS.

	TE	h m	rk	h m	F	h m	h	h == 1	h m	h m	b m	1		-
S.	h m	00 31/				0° 35/	00.36/	00 37/	0 38/	00 39/	00 40	00 414	8.	ı
0	7782	7639	750)	7368	7238	7112	6990		6755		6532	6425	0	١
1	7779	7637	7499	7365	7236	7110	6988		6753		6530	6423	1	١
2 3	17777 17774	7634 7632	7497 7494	7363 7361	7234 7232	7108 7106	6986 6984	6867 6865	6751 6749	6638 6637	65 2 9 65 2 7	6421 6420	2 3	١
4	7772	7630	7492	7359	7229	7104	6982	6863	6747		6525	6418	4	l
5	7769	7627	7490	7357	7227	7102	6980	6861	6745		6523	6416	5	l
6	7767	7625	7488	7354	7225	7100	6978	6859	6743		6521	6414	6	l
7 8	7765 7762	7623 7620	7485 7483	7352 7350	7223 7221	7098 7096	6976 6974	6857 6855	6742 6740		6519 6518		7	l
9	7760	7618	7481	7348	7219	7093	6972	6853	6738		6516	6409	9	l
10	7757	7616	7479	7346	7217	7091	6970	6851	6736	6624	6514	6407	ю	1
11	7755	7613	7476	7344	7215	7089	6968	6849	6734	6622		6406	11	l
12 13	7763 7750	7611 7609	7474 7472	7341 7339	7212 7210	7087 7085	6966 6964	6847 6845	6732 6730	6620 6618		6404	12 13	l
14	7748	7607	7470	7837	7208	7083	6962	6843	6798	6616		6400	14	Į
15	7745	7604	7467	7335	7206	7081	6960	6841	6726	6614	65 05	6398	15	1
16	7743	7602	7465	7333	7204	7079	6958	6840	6725	6612	6503	6397	16	l
17 18	7741 7738	7600 7597	7463 7461	7330 7328	7202 7200	7077 70 75	6956 6954	6838 6836	6723 6721	6611 6609	6501 6500	6395 6393	17 18	1
19	7736	7595	7458	7326	7198	7073	6952	6834	6719	6607		6391	19	l
20	7734	7593	7456	7324	7196	7071	6950	6832	6717	6605	6496	6390	20	ı
21	7731	7590	7454	7322	7193	7069	694 8	6830	6715	6603	6494	6388	21	l
22	7729		7452	7320	7191	7067	6946	6828 6826	6713 6711	6601 6600	6492 6491	6386 6384	22	1
23 [.] 24	7726	7586 7583	7450 7447	7317 7315	7189 7187	7065 7063	6944 6942	6824	6709	6598	6489	6383	24	ı
25	7722	7581	7445	7313	7185	7061	6940	6822	6708	6596	6487	6381	25	ł
26	7719	7579	7443	7311	7183	7059	693 8	6820	6706	6594	6485	6379	26	١
27	7717	7577	7441	7309	7181	7057	6936	6818	6704	6692	6484	6377	27	l
28 29	7714	7574 7572	7438 7436	7307 7304	7179 7177	7055 7052	6934 6932	6816 6814	6702 6700	6590 6589	6482 6480	6376 6374	28 29	ı
30	7710	7570	7434	7302	7175	7050	6930	6812	6698	6587	6478	6372	30	1
31	7707	7567	7432	7300	7172	7048	6928	6810	6696	65 85	6476	6371	31	1
32	7705	7565	7429	7298	7170	7046	6926	6809		. 6583	6475	6369	32	l
33 34	7703 7700	7563 7560	7427 7425	7296 7294	7168 71 6 6	7044 7042	6924 6922	6807 6805	6692 6691	6581 6579	6473 6471	6367 6365	33 34	ı
35	7698	7558	7423	7291	7164	7040	6920	6803	6689	6578	6469	6364	35	ı
36	7696	7556	7421	7289	7162	7038	6918	6801	6687	6576	6467	6362	36	ı
37	7693	7.554	7418	7287	7160	7036	6916	6799	6685	6574		6360	37	ı
3 8 39	7691	7551	7416	7285 7283	7158	7034 7032	6914 6912	6797 6795	6683 6681	6572 6570	6464 6462	6358 6367	38 39	
40	7688	7549 7547	7414	7281	7156	7030	6910	6793	6679	6568	6460	6355	40	ĺ
41	7686 7684	7544	7409	7279	7152	7028	6908	6791	6677	6567	6459	6353	44	١
42	7681	7542	7407	7276	7149	7026	6906	6789	6676	6565	6457	6351	42	ı
43	7679	7540	7405	7274	7147	7024	6904	6787	6674 6672	6563 6561	6455 6453	6350 6348	43	l
45	7677	7638	7403	7272	7145	7022	6902	6785 6784	6670	6559	6451	6346	44	۱
46 46	7674 7672	7535 7533	7401 7398	7270 7268	7143 7141	7020 7018		6784 6782	6668	6558	6450	6344	46	
47	7670	7631	7396	7266	7139	7016	6896	6780	6666	6556	6448	6343	47	
48	7667	7528	7394	7964	7137	7014	6894	6778	6664	6554	6446	6341	48	I
49	7665	7526	7392	7261	7135	7012	6892	6776	6663	6562	6444	6339	49	1
50 51	7663 7660	7524 7522	7390 7387	7259 7257	7133 7131	7010 7008	6890 6888	6774 6772	6661 6659	6550 6648	6443 6441	6338	50 51	1
52	7658	7519	7385	7255	7129		6886	6770	6657	6547	6439	6334	52	1
53	7655	7517	7383	7253	7127		6884	6768	6655	6545	6437	6332	53	l
-54	7653	7515	7381	7251	7124	7002	6882	6766	6653	6543	6435	6331	56	Į
55 56	7651	7513 7510	7379 7376	7249 7246	7122 7120	7000 6998	6881 6879	6764 6763	6651 6650	6541 6539	6434 6432	6829 6327	55 56	ļ
50 57	7648 7646	7508	7374	7244	7118	6996	6877	6761	6648	6538	6430		57	i
58	7644	7506	7372	7242	7116	6994	6875	6759	6646	6536	6428	6324	58	
59	7641	7503	7370	7240	7114			6787		6534	6427	6322	<i>5</i> 9	
S.	100 20	0° 31′	0° 32′	0° 33/	00 34∕	00 35	0º 36 /	00 37	8 ° 38/	0° 39	0° 40′	0° 41/1	8.	١.

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PROPORTIONAL LOGARITHMS.

	8.	h m 0° 42′	h m 0° 43′	k m (° 44/	h m 0° 45′	h.m. 0°46′	h in 0° 47'	b m 0° 484	h m 0°49′	n m 0° 50'	na ma 9°51′	h m 0° 52/	h m 0° 53′	8.
F	0	6320		6118	6u21	5920	5832	5740	5651	5563		5393	5310	0
	1 2	6319 6317	6216	6117	6019	5924 5922	5830 5829	5739	5649	5562 5560	5476 5474	5391 5390	5309 5307	1 2
1	3	6815	6215 6218	6115 6113	6017 6016	5922	5829	5737 5736	5648 5646	5559	5473	5389	5306	3
	4	6313		6112	6014	5919	5826	5734	5645	5557	5471	5387	53 05	4
۲	5	6312		6110	6013	5917	5824	ό733	5643	5556	5470	5386	53 03	5
1	6	6810	Q	6108	6011	5916	5823	5731	5642	5554	5469	5384	5302	6
	7	6308	6206	6107	6009	5914		5730	5640		5467	5383	5300	7
1	8	6306	6205	6105		5913		5728	5639	5051	5466	5382	5299	8
L	9	6305	6203	6103	6006	5911	5818	5727	5637	5550	5464	5380	5298	9
:	10	6303 6801	6201 6200	6102 6100	6005	5909	5816	5725	5636	5549 5547	5463 5461	5379 5377	5296 5295	10 11
	11. 12	6300	6198	6099	6003 6001	5908 5906	5815 5813	5724 5722	5635 5633		5460			12
1	13	6298	6196	6097	6000		5812	5721	5632		5459			13
i	14	6296	6195		5998	5903	5810	5719	5630		8457	5373	5291	14
-	15	6294	6193	6094	5997	5902	5809	5718	5629	5541	5456	5372	5290	15
	16	6293	6191	6092	5995	5900	5807	5716	5697		5454	5370		16
	17	6291	6190	6090	5993	5898	5806	5715	5626		5453			17
1	18	6289 6288	6188 6186	6089 6087	599 2 5990	5897	5804 5803				5452 5450			18
-	19								5623	<u></u>				20
-	20	6286 6284	6185 6183	6085 6084			5801 5800	5710 5709			5449 5447		1	20
1	22	6282		6082			5798		5618		5446			22
-	23	6281	6179				5796					1		23
-1	24	6279	6178	6079	5989	5888	5795	.5704	5615		5443	5389	5277	24
-	25	6277	6176	1			5793	5703	5614	5527	5442	5358		25
- 1	26	6276							5615					26
	27	6274												27
	2 8 2 9	6272	6171 6169	6072			5789 5787		5608					28 29
ŀ			6168						I				5269	30
- 1	30 31	6269 6267	6166				5786 5784						1	31
	32	6265												32
	33	6264	6163	6064			5781	5691	5602		5439			33
	34	6262	6161	6063	5966	5872	5780	- 5689	5601	5514	5429	5346	5264	34
	35	6260			5965					5515	5428		5262	- 35
	36	6259					5777				5426			36
	· 57 38	6257	6156								5425 5423			37
1	39	6254									5422			39
ŀ	40	6252		-							5421	5387	5256	40
- 1	-41	6250					5769	,			5419		1	41
	42	6248	6148	6050	5954	5860	5768	6677	5589	5508	5418	5385	5253	42
- 1	43	6247									5416			43
L	41	6245												44
1	45	6245									8414		5249	45
- 1	46	6242						5671 5670	5535 5582		5412 5411			46 47
- 1	47 48	6238									5409			48
ł	49	6237						5667	5579		5408			49
ŀ	50	6236	6135	-		5847		5666	5578		8407	5324		50
1	51	6233	6133	6035	5939	5846			5576			5322	5241	51
	52	6232					5752		5575		5404			52
- 1	53	6230 6228							5573		5402	5320		53
	54				-		5749		5579		5401	5318		54
- 1	55	6226 6225							5570		5400		5235 5234	55
	56 57	6225							5569 5567		6398 6397			56 57
	58	6221							5566		5395			58
li	59	6220					5742				5394			. 59
	Š.	00 42	00 43	00 44/	0° 45	00 46	00 47	0° 48/	0° 49	00 50	00 51	00.52	0° 53/	S.
-					·							77.77		

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PROPORTIONAL LOGARITHMS.

1 -	h m	b == 1	p 200 1	h m	h m	h m	h == '	· 10 1	h mil	}	h m	<u> </u>		Г
S.						00 594			n m 10 2/				S.	1
-0-	5229	5149	5071	i.		4844		4699		4559	4491	4424	-	1
ĭ	5227	5148	5070	4993		4643				4558	4490		1	Í
2	5226	5146	5068	4991	4916	4842	4769	4697	4626	4557	4489	4421	2	1
3	6225	5145	5067	4990	4915	4841	4768	4696	4625	4556	4488	4420	3	ł
4	5223	5144	5066	4989	4913	4839	4766			4555		4419	4	1
5	5222	5143	5064		4912	4838	4765			4554		4418	5	ţ
6	5221	5141	5063	4986	4911	4837	4764	4692	4622	4552	4484	4417	6	
7	5219	5140	5062	4985	4910	4836	4763	4691	4621	4551	4483	4416	7	1
8	5218	5139	5061	4984	4908		4762	4690	4619	4550	4482	4415		1
9	5217	5137	5059	4983	4907	4833	4760			4549		4414	9	ĺ
10	5215	5136	5058		4906		4759			4548			10	i
11	5214	5135	5057	4980	4905	4831	4758	4686	4616	4547	4479	4411	īį '	
12	5213				4903							4410	12	
13	5211	5132			4902		4756						13	
14	5210	5131	5053		4901	4827	4754	<u> </u>		. —		4408	14	:
15	5209	5129	5051	4975			4753			4542			15	1
16	5207	5128	5050				4752						16	1
1 17	5206		5049 5048			4823	4751			4540			17	i
18 19	5205 5203		5048 5046	1 1			4750			4539			18	!
	·	5124					4748			4538			19	1
20	5202	5123	5045		4894		4747	1		4536			20	
21	5201	5122			4892								31 22	•
22	5199 5198		5043 5041	4966 4965						4534 4533			22	
24	5198	5118			4889 4889					4533 4532			23	
	-				<u> </u>									•
25 26	5195 5194	5116 5115		4962	4887	4814 4812		1		4531 4530			25	
26	5194				4886 4885					4530 4528			26 27	
28	5191	5112			4884							4393	28	
29	5190				4882					4526		4391	28	
30	5189									4525		4390	30	1
31	5189				4880		4735 4734			4524		4389	30	ı
32	5186		5030	4954						4523			31	1
33	5185	5106	502 8	4952	4877							4387	32 33	1
34	5183				4876								34	ĺ
36	5182		5026	4950	4875		4729			4619		4385	35	1
36	5181	5102	5025	4949	4874	4800	4728	4657		4518		4384	36	1
37	5179	5101	5023	4947	4873	4799	4727	4656	. 4586	4517		4383	37 37	i
38	5178	5099	5022	4946	4971	4798	4726	4655	4585	4516	4448	4381	38	İ
39	5177	5098			4870		4724			4515		4380	39	1
40	5175									4514	lI	4579	40	1
41	5174	5095	5018		4868	4794	4722	4651	4581	4512	4445	4378	41	1
42	5173	5094	5017		4866	4793	4721	4650	4580	4511	4444	4377	43	1
43	5172				1 -000		4720	4649	4579	4510	4443	4376	43	1
44	5170		.'						4578	4609	4441	4375	44	j
45	5169				1					4508		4574	45	1
46	5168						4716	4645	4575	4507	4439	4373	46	1
47	5166				1				4574	4506	4438	4372	47	1
48 49	5165											4370	48	1
	5164									4503		4369	49	1
50	5162				4856					4502		4368	50	1
51 59	5161		1		1							4367	51	1
52 53	5160 5158				4854 4863					4500		4366	52	1
54	5157									4499 4498		4365	53 54	ı
													54	1
55	5156		1							4497	4429	4363	55	1
56 57	5154 5153									4495			56	1
58	5152							4632		4494		4361 4359	57 50	1
59	5150												58 59	I
\$.		1				00 59					i			1
1 2.	10- 04/	0- 00	10° 50',	יייי 27'	u~ 58′	U~ 59/	<u>v 0</u> ′.	1/1	10 2/1	10 3	10 4/	10 %	S,	٠

PROPORTIONAL LOGARITHMS.

٠.														
1	. s .	hm 1º6/	h m 10 7/	h ma 1° 8′	hm 1°9⁄	10 104 pr no	h m	h m		h m 1º 14/	h m 10 15/	h m 10164		8.
ľ	0	4357	4292		4164	4102			3919		3802	3745		
1	1	4356	4291	4227	4163	4101	4039		3919	3859	3801	3744		1
	2	4355	4290	4226	4162	4100	4038		3918		5800	3743	3666	2
H	3	4354	4289	4224	4161	4099	4037	3976	3917		3799	3742	3685	
	4	4353	4288	4223	4160	4098	4036	3975	3916		3798	3741	3684	4
	5	4352	4287	4222	4159	4097	4035	3974	3915	3856	3797	3740	3683	-5
1	6	4351	. 4285	4221	4158			3973		3855	3796	3739	3682	6
1	7	4350	4284		4157			3972	3913	3854	37 95	3738	3681	7
Н	8	4349	4283	4219	4156	4093		3971	3912			3737	3680	8
1	9	4347	42 82	4218	4155	4092	4031	3970	3911	3852	3793	3736	3679	9
	10	4346	4281	4217	4154	4091	4030	3969	3910		3792	3735	3678	
	11	4345	4280		4153			3968		3850	3792	3734	3677	10
ı	12	4344	4279	4215	4152	4089		3967	3908	3849	3791	3733		11 12
	13	4343	4278	4214	4151	4088		3966		384 8	3790	3732	3676	13
	14	4342	4277	4213	4150		4026	3965		3847	3789	3731	3675	14
ı	15	4341	4276	4212	4149	4086	4025	3964		3846	3788	3730	3674	
- 1	16	4340	4275	4211	4147						3787	3729	3673	15
. 1	17	4339	4274	4210							3786	3728	3672	16 17
١,	18	4338	4273	4209	4145		4022	3961		3843	3785	3727	3671	18
- 1	19	4336	4271	4207	4144			3960		3842	3784	3727	3670	19
ı	20	4335	4270	4206	4143		4020		3900		3783	3726	3669	
- 1	21	4334	4269	4205	4142		4019		3899		3782	3725		20
1	22	4333	4268	4204	4141		4018		3898		3781	3724	3669 3667	21 22
- 1	23	4332	4267	4203	4140		4017	3956	3897		3780	3723		23
- 1	24	4331	4266	4202	4139		4016	3955	3896		3779	3722		24
ŀ	25	4330	4265	4201	4138	4076	4015	3954	3895	3 836	3778	3721		
- 1	26	4329	4264	4200	4137		4014	3963	3894	3835	3777	3720	3664	25
- 1	27	4328	4263	4199	4136		4013	3952	3893	3834	3776	3719	3663 3663	26 27
- 1	28	4327	4262	4198	4135	4073		3951	3892	3833	3775	3718	3662	28
-1	29	4326	4261	4197	4134		4011	3950	3891	3832	3774		3661	29 29
r	30	4325	4260	4196	4133		4010		3390	3831	3773	3716		
- [31	4323	4259	4195	4132				3889	3830	3772	3715	3660	30
-	32	4322	4258	4194	4131				3888		3771	3714	3659 3658	31 32
-	33	4321	4256	4193			4007	3946	3887	3828	3770	3713	3657	33
- [34	4320	4255	4192	4129	4067	4006	3945	3886		3769	3712	3656	34
-	35	4319	4254	4191	4128	4066	4005	3944	3885		3768	3711	i	
- 1	36	4318	4253	4189	4127	4065	4004		3884		3768	3710	3655 3654	35
-1	37	4317	4252	4188	4126		4003		3883			3709		36 37
1	39	4316	4251	4187	4125	4063	4002	3941	3882	3823		3709	3652	38
1	39	4315	4250	4186	4124	4062	4001	3940	3881	3822	3765	3708	3651	39
ŀ	40	4314	4249	4185	4122	4061	4000	3939	3880	3821	3764	3707	3650	
	41	4313	424 9	4184	4121	4060		3938 3938		3820	3763	3706	3649	40 41
İ	42	4311	4247	4183	4120	4059	3998	3937		3820	3762	3705	3649	42
-	43	4310	4246	4182	4119	4058	3997	3936	3877	3819	3761	3704		43
İ	44	4309	4245	4181	4118	4056	3996	3935	3876	3818	3760			44
ŀ	45	4308	4241	4180	4117	4055	3995		3875	3817	3759	3702	3646	45
I	46	4307	4243	4179	4116	4054	3993			3816	3758	3701	3645	46
-	47	4306	4241	4173	4115	4053	3992		3873	3815	3757	3700	3644	47
ı	48	4305	4240	4177	4114	4052	3991	3931	3872	3814	3756	3699	3643	48
-	49	4304	4239	4176	4113	4051	3990	5930	3871	3813	3755	3698	3642	49
r	50	4303	4238	4175	4112	4050	3989		3870	3812	3754	3697	3641	50
٠	51	4302	4237	4174	4111	4049			3869	3811	3753	3696		51
1	52	4301	4236	4173	4110	4048		3927	3868	3810	3752	3695		52
1	53	4300	4235	4172	4109	4047			3867	3809	3751	3694	3638	53
ı	54	4298	4234	4171	4108	4046	3985		3866	3808	3750	3693	3637	54
1	55	4297	4233	4169	4107	4045	3984	3924	3865	3807	3749	3693		55
1	56	4296	4232	4168	4106	4044		3923			3748	3692	3635	56
1	57	4295	4231	4167	4105			3922	3863	3805	3747	3691	3635	57
1	58	4294	4230	4166	4104			3921	3862	3804	3746	3690	3634	58
i	59	4293	4229	4165	4103	4041			3861	3803	3746	3689		59
t		10 6'						10 12/						S
- <u>'</u> -	-~.	V.	1		- 31	10			- 10	Digitiza	🗴 پرط افد	4 (.) (.)	SIC.	
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PROPORTIONAL LOGARITHMS.

				- T	τ	<u> </u>	<u> </u>	h =	h = ;	h =	h = 7	<u> </u>		٣
8.	h m 1° 18′	h m 10 19	h m 1° 20′	h m 1° 21′	h m 1° 22′	n m 1° 23′	n m 1° 24/	h m 1° 25′	10 26	10 27/	10 28	10 29	S.	١
-0	3632		3522		3415	3362	3310	3259	3208	3158	3108	3059	0	Į
1	3631	3576	3521	3467	3414	3361	3309	3258	3207	3157	3107	3058	1	١
2	3630	3575	3520	3466	3413	3360		3257	3206	3156	3106	3057	2	١
3	3629	3574	3519	3465	3412	3359 3358	3307 3306	3256 3255	3205 3204	3155 3154	3105 3105	3056 3056	3 4	I
4	3628	3573	3518		3411	3358	3306		3204					1
5	3627	3572	3517	3463	3410	3358 3357	3306 3305	3254 3253	3204 3203	3153 3153	3104 3103	3055 3054	5 6	1
6	3626 3625	3571 3570	3516 3515	3463	3409 3408	3357 3356	3305 3304	3253 3253	3203 3202	3163 3162	3103	3054 3053	7	1
7 8	3625 3624	3570 3569	3515 3515		3408 3403	3355	3303	3252	3201	3151	3101	3052	8	١
9	3623	3568	3515 3514		3407	3354	3302	3251	3200	3150	3101	3052	9	1
10	3623		3513	-	3406	3353	3301	3250	3199	3149	3100	3051	10	1
10	3622	3566	3513 3512		3405	3352	3300	3249	3198	3148	3099	3050	11	١
12	3621	3565	3511	3457	3404	3351	3300	3248	3198	3148	3098	3049	12	١
13	3620	3565	3510	3456	3403	3351	3299	3247	3197	3147	3097	3048	13	1
14	3619	3564	3509	3455	3402	3350	3298	3247	3196	3146	3096	3047	14	١
15	3618		3508	3454	3401	3349	3297	3246	3195	3145	3096	3047	15	١
16	3617	3562	3507	3454	3400	3348 3347	3296 3295	3245 3244	3194		3095 3094	3046 3045	16 17	١
17	3616		3506 3506	3453 3452	3400 3399	3347 3346	3295 3294	3244 3243	3193 3193			3045 3044	17	1
18 19	3615 3614		3506 3505		3399 3398	3346 3345	3294 3294	3243	3193 3192	3143 3142		3044 3043	18	1
					3398	3345	3293	3242	3192	3141	3091	3043	20	1
20 21	3613 3612		3504 3503	1	3397 3396	3345 3344	3293 3292	3241	3190	3140	3091	3043 3042	21	١
21 22	3612		3503 3502		3396 3395	3343	3291	3240	3189	3139	5090	3041	22	١
23	3610	3555	3501	3447	3394	3342	3290	3239	3188	3138	3089	3040	23	١
24	3610		3500		3393	3341	3289	3238	3188	3138	3088	3039	24	ļ
25	3609	3554	3499	3446	3393	3340	3288	3237	3187	3137	3087	3039	25	1
26	.3608	3553	3498	3445	3392	3339	3288	3236	3186	3136	3087	3038	26	1
27	3607	3552	3497	3444	3391	3338	3287	3236	3185	3135	3086	3037	27	
28	3606	3551	3497	3443	3390	3338 3337	3286	3235 3234	3184		3085 3084	3036 3035	28 29	1
29	3605	3550	3496		3389	3337	3285		3183			3035		1
30	3604	3549	3495		3388	3336 3335	3284 3283	3233 3232	3183	3133		3034 3034	30 31	1
31 32	3603 3602		3494 3493		3387 3386	3335 3334	3283 3282	3232 3231	3182 3184		3082 3082	3034 3033	31 32	1
32 33	3602 3601	3547 3546	3493 3492		3386 3386	3334 3333	3282 3282	3231	3181 3180			3033 3032	32 33	1
33 34	3600	3545	3492 3491	3438	3385	3332	3281	3230			(34	-
35	3599	3545	3490	3437	3384	3332	3280	3229	3178		3079	3030	35	-
35 36	3598	3544	3489	3436	3383	3331	3279	3228	3178	3128	3078	3030	36	1
37	3598	3543	3488	3435	3382	3330	3278	3227	3177	3127	3078	3029	37	1
38	3597	3542	3488	3434	3381	3329	3277	3226	3176	3126	3077	3028	38	1
39	3596	3541	3487	3433	3380	3328	3276	3225	3175	3125	3076	3027	39	1
40	3595	3540	3486	3432	3379	3327	3276	3225	3174		3075	3026	40	1
41	3594		3485		3379	3326	3275	3224	3173			3026		1
42	3593	3538	3484 3483	3431	3378 3377	3325 3325	3274 3273	3223 3222	3173		3073 3073	3025 3024	42	I
43 44	3592 3591	3537 3536	3483 3482	3430 3439	3377 3376	332 <i>5</i> 3324	3273 3272	3222 3221	3172 3171	3122 3121	3073 3072	3024 3023	43	1
						3323	3272	3221	3171		3071	3023	45	1
45 46	3590 3589	3535 3535	3481 3480	3428 3427	3375 3374,		3271 3270	3220 3220				3022 3022	45 46	1
46	3589 3588	3535 3534	3480 3480		3374	3322	3270	3219				3022	47	1
48	3587	3533	3479	3425	3372	3320	3269	3218	3168	3118	3069	3020	48	1
49	3587	3532	3478		3372	3319	3268	3217	3167	3117	3068	3019	49	J
50	3586	3631	3477	3423	3371	3319	3267	3216	3166	3116	3067	3018	50	1
51	3585	3530	3476	3423	3370	3318	3266	3215	3165	3115	3066	3018	51	-
52	3584	3529	3475	3422	3369	3317	3265	3214	3164	3114	3065	3017	52	1
53	3583	3528	3474	3421	3368	3316	3265	3214	3163	3114	3065	3016	5 3	J
54	3582	3527	3473	3420	3367	3315	3264	3213	3163			3015		ļ
55	3581	3526	3472	3419	3366	3314	3263	3212	3162	3112	3063	3014	55 56	1
56	3580	3525	3471	3418	3365	3313	3262	3211	3161	3111	3062	3014		-
57 58	3579 3578	3525	3471	3417	3365	3313	3261	3210 3209		3110		3013 3012		I
58 59	3578 3577	3524	3470 3469	3416 3415	3364 3363	3312 3311	3260 3259	3209 3209	3159 3158		3060 3060	3012	58 59	1
	10 18												- 8.	1
5. 1	1~ 18"	1-19/	1~ 20/	1. 514	- 22/l	1 23	1-24.	1 25	1- 56/	17 27	1~ 28/	L Z9	· 5.	1

TABLE XXIL

PROPORTIONAL LOGARITHMS.

										,			` .
8.	h m 1° 30′	h m 1031/		h m 1°33/	h m 1° 34/	h m 1° 35′	h m 1° 36′	h m 1° 37′	bm 1° 38′	h m 1039/	h m 1°40′	h m 1° 41′	S.
0	3010	2962	2915	2008	2021	2/10	2/30	2000	2040	2596	2553	2510	0
1	3009	2962		2867	2821	2775		2684	2640	2596	2552	2509	1
2	3009	2961	2913		2820			2684	2639	2595	2551	2508	2
3	3008 3007	2960 2959	2912 2912	2866 2865	2819		2728 2727	2683	2638	2594	2551	2507	3
4					2818	2772		2682	2638	2593	2550		4
5	3006	2958	2911	2864	2818	2772	2726	2681	2637	2593	2549	2506	5
6	3005 3005	2958 2957	2910 2909	2863 2862	2817 2816	2771	2725	2681 2680	2636 2635	2592 2591	2548		6
8	3003	2956	2909	2862	2815		2725 2724	2679	2635	2591	2548 2547	2504 2504	7
9	3003	2955	2908	2861	2815		2723	2678	2634	2590	2546	2503	8 9
	3002	2954	2907	2860	2814		2722	·2678	2633	2589			
10 11	3001	2954	2907	2859	2813	2768 2767	2722	2677	2632	2589 2588	2545 2545	2502 2502	10 11
12	3001	2953	2905	2859	2812	2766	2721	2676	2632	2588	2544		12
13'	3000	2952	2905	2858	2811	2766	2720	2675	2631	2587	2543		.13
14	2999	2951	2904	2857	2811	2765	2719	2675	2630	2586	2543	2499	14
15	2998	2950	2903	2856	2810	2764	2719	2674	2629	2585	2542	2499	15
16	2997	2950	2902	2855	2809	2763	2718	2673	2629	2585	2541	2498	16
17	2997	2949	2901	2855	2808	2763	2717	2672	2628	2584	2540		17
18	2996	2948	2901	2854	2808	2762	2716	2672	2627	2583	2540		18
19	2995	2947	2900	2853	2807	2761	2716	2671	2626	2583	2539	2496	19
20	2994	2946	2899	2852	2806	2760	2715	2670	2626	2582	2538	2495	20
21	2993	2946	2898	2852	2805	2760	2714	2669	2625	2581	2538		21
22	2993	2945	2898	2851	2805	2759	2713	2669	2624	2580	2537	2494	22
23	2992	2944	2897	2850	2804	2758	2713	2668	2624	2580	2536	2493	23
24	2991	2943	2896	2849	2803	2757	2712	2667	2623	2579	2535	2492	24
25	2990	2942	2895	2848	2802	2756	2711	2666	2622	2578	2535	2492	25
26	2989	2942	2894	2848	2801	2756	2710		2621	2577	2534	2491	26
27	2989	2941	2894	2847	2801	2755	2710		2621	2577	2533		27
28	2988	2940	2893	2846	2800	2754	2709	2664	2620	2576	2533		28
29	2987	2939	2892	2845	2799	2753	2708	2663	2619	2575	2532	2489	29
30	2986	2939	2891	2845	2798	2753	2707	2663	2618	2574	2531	2488	30
31	2985	2938	2891	2844	2798	2752	2707	2662	2618	2574	2530		31
32	2985	2937	2890	2843	2797	2751	2706	2661	2617	2573	2530		32
33	2984 2983	2936 2935	2889 2888	2842 2842	2796 2795	2750	2705	2660 2660	2616 2615	2572 2572	2529 2528	2486	33
34						2750	2704						34
35	2982	2935	2887	2841	2795	2749	2704	2659	2615	2571	2527	2485	35
36 37	2981 2981	2934 2933	2887 2886	2840 2839	2794 2793	2748 2747	2703 2702	2658 2657	2614 2613	2570 2569	2527 2526	2484	36
38	2980	2932	2885	2838	2792	2747	2701	2657	2612	2569	2525		37 3 8
39	2979	2931	2884	2838	2792	2746	2701	2656	2612	2568	2525		39
40	2978	2931	2883	2837	2791	2745	2700	2655	2611	2567	2524	2481	40
41	2977	2930	2883	2836	2790	2744	2699	2655	2610	2566			41
42	2977	2929	2882	2835	2789	2744	2698	2654	2610	2566			42
43	2976	2928	2881	2835	2788	2743	2698	2653	2609	2565		2479	43
44	2975	2927	2880	2834	2788	2742	2697	2652	2608	2564		2478	. 44
45	2974	2927	2880	2833	2787	2741	2696	2652	2607	2564	2520	2477	45
46	2973	2926	2879	2832	2786	2741	2695	2651	2607	2563	2520		46
47	2973	2925	2878	2831	2785	2740	2695	2650	2606	2562	2519		47
48	2972	2924	2877	2831	2785	2739	2694	2649	2605	2561	2518	2475	48
49	2971	2924	2876	2830	2784	2738	2693	2649	2604	2 561	2517	2475	49
50	2970	2923	2876	2829	2783	2738	2692	2648	2604	2560	2517	2474	50
51	2969	2922	2875	2828	2782	2737	2692	2647	2603	2559	2516		51
52	2969	2921	2874	2828	2782	2736	2691	2646	2602	2559	2515		52
53	2968	2920	2873	2827		2735		2646		2558	2515		53
54	2967	2920	2873	2826	2780	2735	2689	264 5	2601	2557		2471	54
55	2966	2919	2872	2825	2779	2734	2689	2644	2600			2470	55
56	2965		2871	2825	2779		2688	2643	2599				56
57	2965		2870	2824	2778		2687	2643	2599			2469	67
58	2964		2869		2777		2687	2642	2598				58
59	2963		2869		2776		2686	2641	2597			2467 10 41	59
S.		10 444	10 00/	10 007					110 00/	10 00/	110 AM	110 41/	S.

PROPORTIONAL LOGARITHMS.

S																_
	į	3.								10 49/	h m	h m			S.	ł
1																1
2																į .
Section Sect																Ĺ
A																1
5 2483 2421 2379 2347 2296 2206 2115 2175 2106 2096 2097 2019 5 6 2462 2420 2378 2337 2295 2234 2174 2134 2095 2006 2017 2118 6 7 2460 2419 2377 2333 2291 2253 2213 2173 2134 2094 2055 2017 8 10 2460 2417 2376 2333 2292 2251 2211 2173 2133 2094 2055 2017 8 11 2450 2416 2374 2333 2291 2250 2210 2170 2131 2093 2054 2161 11 12 2450 2441 2373 2331 2291 2200 2210 2170 2131 2092 2003 2014 13 15 2460 2441 2371 2330 </td <td></td> <td>-</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>ı</td>		-														ı
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-1	27	1568	1534	1499	1465	1432	1398	1365	1332	1300	1267	1235	27
- 1	28	1567	1533	1499	1465	1431	1398	1365	1332	1299	1267	1234	28
1.	29	1567	1532	1498		1431	1397	1364	1331	1298	1266	1234	29
- 1	30	1566 1566	1532 1531	1498 1497	1464 1463	1430 1429	1397 1396	1363 1363	1331 1330	1298 1297	1266 1265	1233 1233	50 - 31
- 1	31 32	1565	1531	1496	1463	1429	1396	1362	1329	1297	1264	1232	32
-	33	1565	1530	1496	1462	1428	1395	1362	1329	1296	1264	1232	33
1	34	1564	1530	1495	1461	1428	1394	1361	1328	1296	1263	1231	34
ſ	35	1563	1529	1495	1461	1427	1394	1361	1328	1295	1263	1231	35
1	36	1563	1528	1494 1494	1460	1427 1426	1393 1393	1360 1360	1327	1295	1262 1262	1230 1230	36 37
١	37 38	1562 1562	1528 1527	1493	1460 1459	1426	1392	1359	1327 1326	1294 1294	1261	1229	31 38
ı	39	1561	1527	1493	1459	1425	1392	1359	1326	1293	1261	1229	39
ŀ	40 .	1561	1526	1492	1458	1424	1391	1358	1325	1292	1260	1228	40
1	41	1560	1526	1491	1458	1424	1391	1357	1325	1292	1260	1227	41
1	42	1559	1525	1491	1457	1423	1390 1389	1357 1356	1324	1291	1259	1227 1226	42 43
-	43 44	1559 1558	1524 15 24	1490 1490	1456 1456	1423 1422	1389	1356	1323 1323	1291 1290	1259 1258	1226	44 44
}	45	1558	1523	1489	1455	1422		1355	1322	1290	1257	1225	45
	45 46	1557	1523	1489	1455	1421	1388	1355	1322	1289	1257	1225	46
I	47	1556	1522	1488	1454	1421	1387	1354	1321	1289	1256	1224	47
Į	48	1556	1522	1487	1454	1420		1354	1321	1288	1256	1224	48
	4.9	1555	1521	1487	1453	1419		1353	1520	1288	1255	1223	49
	. 50	1555	1520 1520	1486 1486	1452 1452	1419 1418		1352 1352	1320 1319	1287 1287	1255 1254	1223	50 51
	· 51 52	1554 1554	1519	1485	1451	1418		1351	1519	1286			52
Ì	53	1553		1485	1451	1417	1384	1351	1318	1285	1253	1221	53 .
	54	1552	1518	1484	1450	1417		1350	1317	1285		1221	54
-	55	1552	1518	1483	1450			1350	1317	1284	1252	1220	56
ł	56	1551	1517	1483	1449	1416		1349 1349	1316	1284	1252 1251	1219 1219	56 87
1	57 58	1551 1550	1516 1516		1449 1448			1348	1316 1315				57 58
	59	1550				1414			1315			1218	59
_	S.	I		20 7/							20 14	20/15	8.

lightized by C1009

PROPORTIONAL LOGARITHMS.

8.	h m 2º 16/	h m 20 171	h m 2° 18′	h m 20 19/	h m 2° 20′	h m 2° 21′	h m 2° 22′	h m 2° 23/			h m 2° 26′	8.
0	1217	1186	1154	1123	1091	1061	1030	0999	0969	0939	0909	(
1	1217	1185	1153	1122	1091	1060	1029	0999	0969	0939 0938	0909	1
2	1216	1184	1153	1122 1121	1090 1090	1060 1059	1029 1028	0998 0998	0968 0968	0938	0908	- 3
3 4	1216	1184 1183	1152 1152	1120	1089	1058	1028	0997	0967	0937	0907	
5	1215	1183	1151	1120	1089	1058	1027	0997	0967	0937	0907	
6	1214	1182	1151	1119	1088	1057	1027	0996	0966	0936	0906	(
7	1214		1150	1119	1088	1057	1026	0996	0966	0936	0906	
8	1213		1150 1149	1118	1087 1087	1056 1056	1026 1025	0995 0995	0965 0965	0935 0935	0905	;
9	1213	1181	1149	1117	1086	1055	1025	0994	0964	0934	0904	1
10 11	1212	1180 1180		1117	1086	1055	1024	0994	0964	0934	0904	î
12	1211	1179	1148	1116	1085	1054	1024	0993	0963	0933	0903	1
13	1210	1179	1147	1116	1085	1054	1023	0993	0963	0933	0903	1
14	1210	1178	1147	1115	1084	1053	1023	0992	0962	0932	0902	1
15	1209	1178	1146	1115	1084	1053	1022	0992	0962	0932	0902	1
16	1209	1177	1146	1114	1083	1052	1022	0991	0961	0931	0901	1 1
17	1208	1177	1145 1145	1114	1083 1082	1052	1021	0991 0990	0961 0960	0931 0930	0901 0900	i
18 19	1208	1176 1175	1144	1113 1113		1051 1051	1021 1020	0990	0960	0930	0900	i
20	1207	1175	1143	1112	1081	1050	1020	0989	0959	0929	0899	2
21	1206	1174		1112		1050		0989	0959	0929	0899	2
22	1206	1174		1111	1080		1019	0988	0958	0928	0898	2
23	1205	1173	1142	1111	1080	1049	1018	0988	0958	0928	0898	2
24	1205	1173		1110		1048	1018	0987	0957	0927	0897	2
25	1204	1172	1141	1110	1079	1048	1017	0987	0957	0927	0897	2
2 6 27	1204		1140 1140		1078 1078		1017	0986	0956	0926	0896	2 2
28	1203	1171	1139	1109 1108	1077		1016 1016	0986 0985	0956 0955	0925	0896 0895	2
29	1202	1170	1139	1108	1076	1046	1015	0985	0955	0925	0895	2
30	1201	1170	1138	1107	1076	1045		0984	0954	0924	0894	3
31	1201	1169	1138	1106	1075	1045		0984	0954	0924	0894	3
32	1200		1137	1106	1075			0983	0953	0923	0893	3
33 34	1200	1168	1137 1136	1105	1074	1044	1013	0983	0953	0923 0922	0893 0892	3
35	1199	1168		1105	1073		1013	0982	0952 0952	0922		3
36	1199 1198			1104 1104		1043 1042		0981	0951	0921	0892 0891	3
37	1198	1166	1135	1103	1072	1042		0981	0951	0921	0891	3
38	1197	1165	1134	1103	1072	1041	1011	0980	0950	0920	0890	3
39	1197	1165	1134	1102	1071	1041	1010	0980	0956	0920	0890	3
40	1196	1164	1133	1102	1071	1040		0979	0949	0919	6380	4
41	1196				1070			0979 0978		0919	0889	4
42 43	1195			1101 1100				0978		0918 0918	0888 0888	4
44	1194	1162		1100		1038	1007	0977	0947	0917	0887	4
45	1193	1162	1130	1099	1068	1037	1007	0977	0947	0917	0887	4
46	1193	1161	1130	1099	1068	1037	1006	0976	0946	0916	0886	4
47	1192		1129	1098				0976	0946	0916	0886	4
48 49	1192	1160 1160	1129 1128	1098 1097		1036 1035		0975 0975	0945 0945	0915 0915	088 <i>5</i> 088 <i>5</i>	4
50	1191	1159	1128		1066	1035	1004	0974	0944	0914	0884	5
51	1190		1127					0974	0944	0914	0884	5
52	1190		1127	1096				0973		0913	0883	
53	1189	1158	1126	1095	1064	1033	1003	0973	0943		0883	5
54			1126				1002				0833	5
55	1188							0972			0882	5
56 57	1188 1187		1125 1124			1032 1031					0882 0881	5
58	1187										0881	5
59	1186	1154	1123	1092	1061	1030	1000	0970	0940	0910	0880	5
S.	20 16	90 17/	90 19/	90 10/	90 90/	90 91/	00 00/	90 93/	90 94/	90 95/	90 96/	.5

PROPORTIONAL LOGARITHMS.

S.	h m 2° 27′	h m 20 28/	h m 2° 29′	h m 2° 30′	h m 2° 31′	h m 2° 32′	h m 2° 33′	h m 2° 34/	h m 2° 35′	h m 2º 36'	h m 20 37/	S.
0	0380	0850	0821	0792	0763		0706		0649		0594	0
1	0879	0850	0820		0762	0734		0677	0649	0621	0593	ì
2	0879	0849	0820		0762	0733		0677	0648		0593	2
3 4	0878	0849	0819	0790	0762	0733		0676	0648	0620	0592	3
	0878	0848	0819	0790	0761	0732	0704	0676	0648	0620	0592	4
5	0877	0848	0818	0789	0761	0732	0703	0675	0647	0619	0591	5
6	0877	0847	0818	0789	0760		0703	0675		0619		6
8	0876 0876	0847 0846	0817	0788	0760		0703					7
9	0875	0846	0817 0816	0788 0787	0759 0759		0702					8
10										0617		9
11	0875 0874	0845	0816	0787	0758	0730	0701	0673		0617	0589	10
12	0874	0845 0844	0816 0815		0758			0672				11
13	0873	0844			0757 0757		0700 0700					12
14	0873	0843	0814		0756		0699	0671 0671	0643			13 14
15	0872	0843	0814		0756							
16	0872	0842	0814		0755		0699 0698	0670				15
17	0871	0842										16
18	0871	0841	0812									17 18
19	0870	0841	0812					0669		0613		18
20	0870	0840	0811	0782	0753							
21	0869	0840	0811	0782								20 21
22 23	0869	0839	0810		0752							22
23	0868	0839	0810		0752							23
24	0868	0838	0809			0723						24
25	0867	0838	0 809	0780	0751	0722					·	25
26	0867	0837	0808									26
27	0866	0837	0808		0750		0693					27
28	0866	0836	0807	0778	0750		0693					28
29	0865	0836	0807	0778	0749	0721	0692	0664				29
30	0865	0835	0806	0777	0749	0720	0692	0663	0635	0608	0580	30
31	0864	0835	0 306	0777	0748	0720						31
32	0864	0834							0634			32
33	0863	0834		0776	0747		0690					33
34	0863	0834	0804	0775	0747		0690	0662	0634	0606	0578	34
35	0862	0833	0804				0689	0661	0633	0605	0578	35
36	0862	0833									0577	36
37	0861	0832										37
3 8 3 9	0861	0832			0745							
	0860	0831	0802	0773			0687	!			.'	
40	.0860	0831	Q801	0773	0744							40
41 42	0859	0830		0772								
43	0859 0858	0830 0829	0801 0800									42
44	0858	0829	0800		0742							43
. 45	0857	0828					i —					44
46	0857	0828	0799 0799	0770 0770	0741	0713	0685					45
47	0856	0827	0798								0573	46
48	0356	0827			0740		0683				0572 0572	47 48
	0855	0826	0797	0768	0740		0683					49
49		0826	0797	0768	0739		0682				0571	
	0855				0739							50 51
49 50 51	0855 0855	0825	0796	0767			, ,,,,,					
5Ó			0796 0796	0767 0767			0681	1 0653	0625	0.507	0.670	59
50 51	0855	0825	0796	0767	.0738	0710						
50 51 52	0855 0854	0825 0825	0796		.0738	0710		0653	0625	0597	0569	53
50 51 52 53	0855 0854 0854	0825 0825 0824	0796 0795	0767 0766 0766	0738 0738 0737	0710 0709 0709	0681 0680	0653 0652	0625 0624	0597 0596	0569 0569	53 54
50 51 52 53 54 55 56	0855 0854 0854 0853	0825 0825 0824 0824	0796 0795 0795	0767 0766 0766 0765	0738 0738 0737	0710 0709 0709 0708	0681 0680 0680	0653 0652 0652	0625 0624 0624	0597 0596	0569 0569 0568	53 54 55
50 51 52 53 54	0855 0854 0854 0853	0825 0825 0824 0824 0823	0796 0795 0795 0794	0767 0766 0766 0765 0765	0738 0738 0737 0737 0736	0710 0709 0709 0708 0708	0681 0680 0680 0679	0653 0652 0652 0651	0625 0624 0624 0623	0597 0596 0596 0596	0569 0569 0568 0568	53 54 55 56
50 51 52 53 54 55 56 57 58	0855 0854 0854 0853 0853 0852 0852 0851	0825 0824 0824 0823 0823 0823 0822 0822	0796 0795 0795 0794 0794 0793 0793	0767 0766 0766 0765	0738 0738 0737 0737 0736 0736	0710 0709 0709 0708 0708 0708	0681 0680 0680 0679 0679	0653 0652 0652 0651 0651	0625 0624 0624 0623 0623	0597 0596 0596 0596	0569 0569 0568 0568 0568	53 54 55 56 57
50 51 52 53 54 56 56 57 58 59	0855 0854 0853 0853 0852 0852 0851 0851	0825 0825 0824 0824 0823 0823 0822 0822 0821	0796 0795 0795 0794 0794 0793 0793 0792	0767 0766 0765 0765 0764 0764 0763	0738 0737 0737 0736 0736 0735 0735	0710 0709 0709 0708 0708 0707 0707	0681 0680 0679 0679 0678 0678	0653 0652 0651 0651 0650 0650	0625 0624 0623 0623 0623 0622	0597 0596 0596 0596 0595 0595	0569 0569 0568 0568 0568 0567	53 54 55 56 57 58 59

PROPORTIONAL LOGARITHMS.

S. pb m h m b m h m h m h m h m h m h m h m															_
0 0.666 0.539 0.512 0.484 0.456 0.431 0.404 0.776 0.352 0.396 0.300 0 1 0.666 0.587 0.511 0.484 0.457 0.430 0.403 0.377 0.351 0.325 0.299 1 2 0.666 0.587 0.510 0.483 0.457 0.450 0.403 0.377 0.351 0.325 0.299 1 3 0.666 0.537 0.510 0.483 0.456 0.450 0.403 0.377 0.351 0.325 0.299 1 5 0.666 0.566 0.537 0.510 0.483 0.456 0.450 0.403 0.377 0.351 0.325 0.299 1 5 0.666 0.566 0.597 0.510 0.483 0.456 0.450 0.403 0.377 0.350 0.324 0.298 4 0.456 0.456 0.429 0.403 0.376 0.360 0.334 0.298 4 0.456 0.456 0.429 0.402 0.376 0.349 0.332 0.297 6 0.666 0.665 0.556 0.609 0.482 0.455 0.422 0.402 0.376 0.349 0.332 0.297 7 0.663 0.364 0.009 0.481 0.454 0.427 0.401 0.374 0.348 0.332 0.297 7 0.663 0.364 0.009 0.481 0.454 0.427 0.401 0.374 0.348 0.332 0.296 9 0.662 0.353 0.608 0.491 0.454 0.427 0.401 0.374 0.348 0.332 0.296 9 0.662 0.353 0.607 0.480 0.453 0.426 0.399 0.373 0.346 0.330 0.296 0.479 0.453 0.426 0.399 0.373 0.347 0.331 0.295 10 11 0.0661 0.534 0.607 0.430 0.453 0.426 0.399 0.373 0.346 0.330 0.294 12 13 0.661 0.533 0.606 0.479 0.452 0.426 0.399 0.373 0.346 0.330 0.294 12 13 0.660 0.332 0.606 0.479 0.452 0.426 0.399 0.373 0.346 0.330 0.294 12 13 0.660 0.332 0.606 0.479 0.452 0.426 0.399 0.373 0.346 0.330 0.294 12 13 0.660 0.332 0.606 0.477 0.450 0.424 0.397 0.371 0.345 0.319 0.295 11 1 0.650 0.531 0.600 0.477 0.450 0.423 0.399 0.372 0.346 0.319 0.294 14 15 0.550 0.531 0.600 0.477 0.460 0.423 0.399 0.372 0.346 0.319 0.294 13 13 0.650 0.351 0.600 0.477 0.460 0.423 0.397 0.371 0.345 0.319 0.295 11 1 0.657 0.500 0.600 0.476 0.449 0.422 0.399 0.372 0.346 0.319 0.295 11 1 0.657 0.500 0.600 0.476 0.449 0.422 0.399 0.372 0.346 0.319 0.295 11 1 0.657 0.500 0.600 0.476 0.449 0.422 0.399 0.309 0.340 0.311 0.295 11 1 0.657 0.550 0.600 0.600 0.476 0.449 0.422 0.399 0.309 0.342 0.316 0.291 1 1 0.550 0.552 0.600 0.600 0.476 0.449 0.422 0.399 0.309 0.342 0.316 0.291 1 1 0.550 0.552 0.600 0.600 0.476 0.449 0.422 0.399 0.309 0.342 0.316 0.291 1 1 0.550 0.552 0.600 0.600 0.476 0.449 0.423 0.399 0.309 0.342 0.316 0.291 1 1 0.550 0.552 0.550 0.600		S.											h m	S.	1
1 0.666 0538 0511 0484 0457 0450 0405 0377 0351 0325 0299 2 3 0565 0537 0510 0483 0456 0430 0403 0377 0350 0324 0298 3 4 0564 0537 0510 0483 0456 0430 0403 0377 0350 0324 0298 3 5 0664 0537 0510 0483 0456 0430 0403 0376 0350 0324 0298 4 5 0664 0535 0509 0482 0455 0429 0402 0376 0349 0333 0297 6 6 0663 0356 0509 0482 0455 0428 0402 0375 0349 0333 0297 6 8 0562 0353 0508 0491 0454 0427 0401 0375 0349 0333 0297 6 8 0562 0353 0508 0491 0454 0427 0401 0376 0349 0333 0297 7 8 0562 0353 0507 0480 0454 0427 0401 0376 0349 0333 0297 7 10 0456 052 0534 0507 0480 0454 0427 0401 0376 0349 0333 0297 7 11 0561 0353 0506 0491 0454 0427 0401 0376 0349 0332 0296 8 11 0561 0353 0506 0479 0453 0426 0399 0373 0346 0320 0294 12 13 0561 0533 0506 0479 0452 0426 0399 0373 0346 0320 0294 12 13 0560 0533 0506 0479 0452 0426 0399 0373 0346 0320 0294 12 15 0561 0535 0506 0479 0452 0426 0399 0373 0346 0320 0294 14 16 0569 0331 0504 0477 0450 0423 0398 0371 0345 0319 0294 14 16 0569 0331 0504 0477 0450 0423 0397 0371 0345 0319 0294 14 17 0568 0331 0504 0477 0450 0423 0397 0371 0345 0319 0293 15 18 0569 0331 0504 0477 0450 0423 0397 0371 0345 0319 0293 15 18 0569 0351 0503 0476 0449 0422 0396 0370 0344 0318 0392 17 20 0557 0530 0503 0476 0449 0422 0396 0370 0344 0318 0392 17 21 0557 0530 0503 0476 0449 0422 0396 0370 0344 0318 0392 17 22 0556 0529 0500 0473 0449 0422 0396 0370 0344 0318 0392 17 23 0556 0529 0500 0473 0449 0422 0396 0370 0344 0318 0392 18 24 0556 0529 0500 0473 0449 0422 0396 0370 0345 0317 0291 19 25 0557 0550 0509 0500 0473 0449 0422 0396 0370 0345 0317 0291 19 25 0557 0550 0509 0500 0473 0440 0419 0393 0366 0340 0314 0289 22 25 0556 0529 0500 0476 0449 0422 0396 0370 0345 0317 0291 19 25 0550 0529 0500 0476 0449 0422 0396 0370 0345 0317 0291 21 22 0556 0529 0500 0476 0449 0422 0396 0370 0345 0318 0392 0318 0392 0318 0392 0318 0392 0318 0392 0318 0392 0318 0392 0392 0392 0392 0392 0392 0392 0392				. — —											1
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56 0541 0513 0486 0459 0433 0406 0380 0353 0327 0301 0276 56 57 0540 0513 0485 0459 0432 0406 0379 0353 0327 0301 0275 57 58 0540 0512 0485 0432 0405 0379 0353 0356 0300 0275 58 59 0539 0512 0485 0431 0405 0379 0352 0326 0300 0274 59	ŀ		-												ı
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59 0539 0512 0485 0458 0431 0405 0378 0352 0326 0300 0274 59	- 1	57	0540	0513	0486		0432	0406	0379	0353	0327	0801	0275	57	ı
1 22 222	!					0458		0405	0379	0353					ĺ
S. 20 38/ 20 39/ 20 40/'20 41/ 20 42/ 20 43/ 20 44/ 20 45/ 20 46/ 20 47/ 20 48/ S.	Ļ														į
	ᆚ	S.	2° 38/	20 39/	20 40/	41	20 42/	20 43/	0 44	20 45/	20 46/	20 471	20 48/1	S.	L

7

PROPORTIONAL LOGARITHMS.

	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m	h m		
S.		20 50	20 51/	20 52/	2° 53⁄	20 54/	2º 55/	2º 56′	2 ° 57′	2 ° 58′	2 0 59′	8.	
0	0274			0197		0147	0122		0073		0024	0	l
1	0273		0222				0122		0073	0048	0024	1	l
2 3	0273		0222 0221	019 7 0196	0171 0171	0146	0122 0121	0097	0072 0072	0048 0047	0023	2	ł
4	0273		0221	0196		0146 0146	0121	0096	0072	0047	0023 0023	3 4	İ
5	0272		0221	0195	0170		0120	0096	0071	0046	0023		
6	0271		0220				0120		0071		0022	5	l
7	0271		0220			0144	0119	0095	0070		0021	7	
8	0270		0219	0194	0169	0144	0119	0094	0070		0021	8	l
9	0270		0219	0194	0169	0143	0119	0094	0069	0045	0021	9	İ
10	0270		0219	0193		0143			0069		0020	10	
111	0269		0218								0020		ı
12 13	0269		0218 0217	0192 0192							0019	12	1
14	0268		0217	0192	0166	0142 0141	0117 0117	0092	0068 0067		0019 0019	13 14	1
15	0267	0242	0216	0191	0166		0116	0091	0067				ı
16	0267		0216	0191	0166		0116		0067		0018 0018	15 16	l
17	0267		0216						0066		0017	17	ı
18	0266		0215		0165	0140	0115	0090	0066	0041	0017	18	1
19	0266		0215	0189	0164	0139	0114	0090	0065	0041	0017	19	1
20	0265		0214	0189	0164	0139	0114	0089	0065		0016	20	1
21	0265		0214						0064		0016		
22	0264		0213 0213						0064		0015		ŀ
24	0264		0213		0163 0162		0113 0112		0064		0015 0015	23 24	1
25	0263		0212	0187	l					1			l
26	0263		0212	0187	0162 0161	0137 0136	0112		0063		0014	25 26	l
27	0262		0211	0186					0062		0013		1
28	0262		0211	0186					0062		0013	28	ı
29	0261		0211	0185	0160	0135	0110	0086	0061	0037	0012	29	l
30	0261		0210	0185			0110		0061	0036	0012	30	ı
31	0261		0210	0184					0060		0012	31	1
32	0260 0260		0209 0209	0184 0184	0159				0060		0011	32	ı
34	0259		0203	0183	0158 0158		0109 0108		0060 0059		0011	33 34	ł
35	0259		0208	0183	0158	0133	0108		0059		0010		1
36	0258		0208	0182	0157	0132	0108	0083	0059		0010	35 36	ı
37	0258	0233	0207	0182	0157		0107		0058		0009	57	l
38	0258	0232	0207	0181	0156		0107	0082	0057	0033	0009	38	1
39	0257	0232	0206	0181		0131	0106	0082		0033	0008	39 .	•
40	0257	0231	0206	0181	0156		0106	0081	0057	0032	0008	40	1
41	0256 0256		0205 0205	0180			0105		0056		0008	41	1
43	0255		0205	0180 0179	0155 0154	0130 0129	0105 0105	0080	0056 0055	0031 0031	0007	42 43	
44	0255	0230	0204	0179	0154	0129	0104		0055	0031	0006	44	ı
45	0255	0229	0204	0179		0129	0104	0079	0055	0030	0006	45	
46	0254	0229	0203	0178		0128	0103	0079	0054	0030	0006	46	Ĭ
47	0254		0203	0178	0153	0128	0103	0078	0054	0029	0005	47	1
48 49	0253 0253		0202	0177	0152	0127	0103	0078	0053	0029	0005	48	1
50	0252	0227	0202	0177	0152	0127	0102	0077	0053	0029	0004	49	1
51	0252	0227 0227	0202 0201	0176 0176	0151 0151	0126 0126		0077	0053		0004	50	
52	0252		0201	0176	0151	0126	0101 0101	0077 0076	0052 0052	0028	0004	51	l
53	0251	0226	0200	0175	0150	0125	0100	0076	0051	0027	0003	52 53	ı
54	0251	0225	0200	0175	0150	0125	0100	0075	0051	0027	0002	54	l
55	0250		0200	0174	0149	0124	0100	0075	0051	0026	0002	55	
56	0250	0224	0199	0174	0149	0124	0099	0075	0050	0026	0002	56	
57 58	0250	0224	0199	0174	0148	0124	0099	0074	0050	0025	0001	57	
59	0249	0223	0198 0198	0173 0173	0148	0123	0098 0098	0074	0049	0025	0001	58	l
S.	20 49		90 51/	0110	01.90	0123	0038	0073	0049	0025	0000	59.	!
<u></u>	F 33	~00	- 014	0Z'	2 03/	z 54/	z~ 65/°	z~ 56/	z~ 57/	z ^い 58年	zo 59/	S .	1

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<u> </u>	TABL			To ED T		the La	titud	le by	tv tv					n. 147
	- 117		HOU		IME.		├				HOU	TIME.	<u>'</u>	
M	0"	10"	20/	30″	40″	507	MI	0/		10"	201	30"	40″	50"
	Infinite.			-			<u> </u>	Inf. I				-	70	
	3.	13833	1			1	0	2.	Ĭ	16270	46373	63982	76476	86167
	2.					43936		2.94	085			l '		
	2.36018 2.05916			18409	13834	09695		3.	197	97669	3000	11694	16269	20408 39313
اء		02990		96225	93422	90790	3	3.41	7 96	44144	46371	48490	50510	52440
	1.88307		83732	81613	7959 3	77663	4	3.54	2 89	56061	57764	59403	60982	62506
·	1.75814						5	3.63	978	65402	66781	68117	69413	70672
	1.66125													77542
	1.58208 1.51515									79609 85980				83471 88686
	1.45718									90294				
9	1.40605	39809	39027	38258	37503	36762			_					97545
	1.36032						11	3.98		93860		i i	l	1
	1.31896							4.	000		00170			01376
	1.28120 1. 24 647													04896 08151
	1.21432													11178
15	1.18440	17961	17487	17018	16554	16096					·			14007
	1.15642						16	4.14	461	14911	15355	15796	16231	16663
	1.13013						17	4.17	090	17513	17932	18346	18757	19164
	1 . 105 3 6 1 . 08193						10	4.19	007 910	19907	99664	20755	21140	21529 23770
I	1.05970													25901
	1.03857									26588				
	1:01843						22	4.28	2 60	28587	28911	29233	29553	29870
	0.99918													31725
·	0.98077			1			11							33503
	0.96310									34080				
	0.94614 0.92982													36853 38434
	0.91411													39960
29	0.89894	89647	89401	89156	88913	88671	29	4.40	209	40456	40702	40947	41190	41432
	0.88430													42856
	0.87015 0.85644													44233 45568
	0.8 4317									46003				
	0.83030									47284				
35	0.81780	81576	81372	81169	80967	80767	35	4.48	323	48527	48731	48934	49136	49336
	0.80567						36	4.49	53 6	49735	49933	50130	50326	50522
	0.79387													51675
	0.78 23 9 0.771 22									52052 53165				
	0.76033										_			54956
	0.74972									55306				
42	0.73937	73767	73597	73429	73261	73093								57010
	0.72927									57343				
	0.71940								-	58325				
	0.70976 0.70034									59285 6 022 3				59913
	0.69113									61141				
1 48	0.68212	68064	67916	67769	67622	67476	48	4.61	891	62039	62187	62334	6 24 81	62627
49	0.67330	67185	67040	66896	66752	66609				62918				
	0.66466						50	4.63	637	63779	63921	64062	64203	64343
	0.65620									6462 2 65448				
	0.64791 0.63978									66 25 8				
	0.63181									67053				
	0.62400						55	4.67	703	67832	67961	68089	68216	68344
56	0.61632	61506	61380	61254	61129	61004	56	4.68	171	68597	68723	68849	68974	69099
	0.60879						57	4.69	224	69348	69472 7000	70995	69718 20 440	70860
).60140).59414						50	. છક. 4. જ∩	500 580	70085 70809	709 2 8	71047	71166	71285
1030	,.03717	UJ4J9	UJ 1 10	J2000	00301	00010	. 03	7.70		.0003				

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H	ALF E			IME.			_				TIME		
	1	HOU						0"		HOU		HA //	
M 0"	10"	20"	30"		50"	M	-			20"		40″	
0.0.58700 1.0.57999	58583	58465	58348	58232	58115	ľ	4.	72104	79919	,71038 79335	79450	79564	71988
10.57999 20.57310	157884 157196	57083	56970	56857	56745								73358
30.56633	156521	56409	56298	56187	56076		4.	73470	73582	73694	73805	73916	74027
40.55966	55856	55747	55637	55528	55419	4	_						74684
50.55311	55203	55095	54987	54880	54773	и .	4.	74792	74900	75008	75116	75223	75330
60.54666	54559	54453	54347	54212	54136	6	4.	75437	75544	75650	75756	75861	75967 76593
7 0.54031 8 0.53406	53920	53822 53900	53098	59995	52893		4	76697	76800	76903	77005	77108	77210
90.52791	52690	52589	52487	52387	32286	9		77312					
10 0.52186						10	4.	77917	78017	78117	78217	78316	78415
110.51589	51491	51393	51294	51197	51099	14		78514					
120.51002						12							79584
130.50423 140.49852													80156 8 0720
150.49290						1			i				81275
160.49290								81367					
170.48189						17	4.	81914	82004	82094	82184	82274	82363
180.47650	47561	47473	47384	47295	47207			82453					
190.47119						U		82984					
200.46595	46508	46422	46335	46249	46163			83508					
21 0.46078 22 0.45567								. 84025 . 84536					
23 0 . 45064								85039					
24 0 . 44567								85536					
25 0.44077	43995	43915	43834	43753	13673	25	4.	86026	86108	86188	86269	86350	86430
26 0 . 43592	43512	43432	43353	43273	43194								86909
27 0 . 43114						27 28		. 86989 . 87461					87382
28 0 . 42642 29 0 . 42176								. 879 2 7					
30 0 . 41716						H		88387					
310.41261						31	4.	88842	88917	88992	89067	89142	89216
320.40819	40738	40664	40590	40516	40442			89291					
330.40368	40295	40222	40149	40076	40003			89735					
34 0.39930	4							90173					
35 0 . 39497 36 0 . 39069								90606 910 3 4					
37 0.38646								91457					
380.38227								91876					
390.37814	1					39	4.	92289	92358	92426	9 2494	92562	92630
40 0 . 37405								92698					
410.37001 420.36602								9310 2 93501					
43 0 . 36206								93897					
44 0 . 35816								94287					
45 0.35429								94674					
46 0 . 35047								95056					
47 0 . 34669 48 0 . 34295								. 9 5434 . 95808					
490.34295								.96178					
500.33559						II——	_	96544					
510.33197								96906					
520.32839	32780	32720	32661	32602	32543	52	4.	97264	97323	973 83	97442	97501	97560
530.32485								97618					
54 0.32134						0	_	97969					·
550.31787	31729	31672	31614	31557	31500			98316					
560.31443 570.31103								98660 99000					
48 0.30766						58	4.	99337	99393	99448	99504	99559	99615
59 0 . 30433						59	4.	99670	99725	99780	99835	99890	99945
	-3-					-	•						000

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HALF ELAPSED TIME.	MIDDLE TIME.
2 HOURS.	2 HQURS.
M 0" 10" 20" 30" 40" 50"	M 0" 10" 20" 30" 40" 50"
0 0 . 30103 30048 29994 29959 29885 29831	95.000000005500109001640021800272
10.29776 29722 29668 29614 29561 29507 20.29453 29400 29346 29293 29239 29186	1 5.00327 00381 00435 00489 00542 00596 2 5.00650 00703,00757 00810 00864 00917
30.291332908029027289742892128869	\$5.00970 01023 01076 01129 01182 01234
40.28816 28764 28711 28659 28607 28554	45.01287 01339 01392 01444 01496 01549
5 0 . 28502 28450 28398 28346 28295 28243	55.01601 01653 01705 01757 01808 01860
60.28191 28140 28089 28037 27986 27935	6 5.01912 01963 02014 02066 02117 02168
70.278842783327782277312768027630 80.275792752927478274282737827327	7 5.02219 02270 02321 02372 02423 02473 8 5.02524 02574 02625 02675 02725 02776
90.27277 27227 27177 27127 27078 27028	95.028260287602926029760302503075
10 0.26978 26929 26879 26830 26781 26731	10 5.03125 03174 03224 03273 03322 03372
110.26682 26633 26584 26535 26487 26438	11 5.03421 03470 03519 03568 03616 03665
12 0 . 26389 26341 26292 26244 26195 26147	125.037140376203811038590390803956 135.040040406204100041480419604244
150.26099 26051 26003 25955 25907 25859 140.25811 25763 25716 25668 25621 25573	145.042920434004387044350448204530
15 0 . 25526 25479 25432 25385 25338 25291	155.04577 04624 04671 04718 04765 04812
160.25244 25197 25150 25104 25057 25011	165.048590496604953049990504605092
17 0 . 24964 24918 24872 24825 24779 24738	175.051390518505231052780532405370
18:0.24687 24641 24595 24550 24504 24458 19:0.24413 24367 24322 24276 24231 24186	18 5.05416 05462 05508 05553 05599 05645 19 5.05690 05736 05781 05827 05872 05917
200.241412409624051240062396123916	205.05962 06007 06052 06097 06142 06187
210.23871 23827 23782 23738 23693 23649	215.0623206276063210636506410106454
22 0 . 23605 23560 23516 23472 23428 23384	21 5.06232 06276 06321 06365 06410 06454 22 5.06498 06543 06587 06631 06675 06719
230.238402329623253232092316523122	
24 0 . 23078 23035 22991 22948 22905 22862	24 5 .07025 07068 07112 07155 07198 07241
250.228192277522732226902264722604 260.225612251922476224332239122349	25 5.07284 07528 07371 07413 07456 07499 26 5.07542 07584 07627 07670 07712 07754
27 0 . 22306 22264 22222 22180 22138 22096	2715 07797107839107881107923107965108007
28,0.22054 22012 21970 21928 21887 21845	28 5 . 08049 08091 08133 08175 08216 08258 29 5 . 08300 08341 08383 08424 08465 08507
29 0 . 21803 21762 21720 21679 21638 21596	29 5.08300 08341 08383 08424 08465 08507
30 0.21555 21514 21473 21432 21391 21350	305.08548 08589 08630 08671 08712 08753
310.21309 21269 21228 21187 21147 21106 320.21066 21025 20985 20945 20905 20864	315.08794 08834 08875 08916 08966 08997 32 5.09037 09078 09118 09158 09198 09239
33 0 . 20824 20784 20744 20704 20665 20625	33 5.09279 09319 09359 09399 09438 09478
34 0 . 20585 20545 20506 20466 20427 20387	84 5.09518 09558 09597 09637 09676 09716
35 0 . 20348 20309 20269 20230 20191 20152	355.097550979409834098730991209951
36 0.20113 20074 20035 19996 19957 19919 37 0.19880 19841 19803 19764 19726 19687	36 5.09990 10029 10068 10107 10146 10184 37 5.10223 10262 10300 10339 10377 10416
380:196491961119572195341949619458	385.104541049210531105691060710645
39 0 . 19420 19382 19344 19306 19269 19231	395.10683 10721 10759 10797 10834 10872
40 0 . 19193 19156 19118 19081 19043 19006	40 5 . 10910 10947 10985 11022 11060 11097
410.18968 18931 18894 18857 18820 18783	415.11135/11172/11209/11246/11283/11320 42/5.11357/11394/11431/11468/11505/11542
420.18746 18709 18672 18635 18598 18561 430.18525 18488 18451 18415 18378 18342	435.115781161511652116881172511761
440.18306 18269 18233 18197 18161 18125	44 5 . 11797 11834 11870 11906 11942 11978
45 0 . 18089 18063 18017 17981 17945 17909	40 5 . 12014 12050 12086 12122 12158 12194
460.178741783817802177671773117696	46 5.12229 12265 12301 12336 12372 12407
470.176601762517590175541751917484 48.0.174491741417379173441730917274	475.12443 12478 12513 12549 12584 12619 48 5.12654 12689 12724 12759 12794 12829
49 0 . 17239 17205 17170 17135 17101 17066	49 5 . 12864 12898 12933 12968 13002 13037
500.17032 16997 16963 16928 16894 16860	50 5.13071 13106 13140 13175 13209 13243
510.16826 16792 16758 16723 16690 16656	51 5.13277 [331] 13345 13380 13413 13447
52 0.16622 16588 16554 16520 16487 16453 53 0.16419 16386 16352 16319 16285 16262	52 5.13481 13515 13549 13583 13616 13650 53 5.13684 13717 13751 13784 13818 13851
540.162191618616152161191608616055	54 5.13884 13917 13951 13984 14017 14050
55 0 . 16020 15987 15964 15921 15888 15856	55 5.14083 14116 14149 14182 14215 14247
560.15823 15790 15758 15725 15692 15660	56 5.14280 14313 14345 14378 14411 14443
57 0.15627 15595 15563 15530 15498 15466	57 5 . 14476 14508 14540 14573 14605 14637
580.154341540215370153381530615274 590.152421521015178151461511515083	58 5.14669 14701 14733 14765 14797 14829 59 5.14861 14893 14925 14957 14988 15020
1 0910 - 10922(10210[10110]10190[10110]10083	030.1900111905011934E119300119300110020

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TABLE XXIII.

To find the Latitude by two Altitudes of the Sun.

HALF ELAPSED TIME.	MIDDLE TIME.
\$ HOURS.	3 HOURS.
M 0" 10" 20" 30" 40" 50"	M 0" 10" 20" 30" 40" 50"
0 0 . 15051 15020 14988 14957 14926 1489	
1 0 . 14863 14832 14800 14769 14738 1470 2 0 . 14676 14645 14614 14583 14552 1452	
30.14490144601442914398143681433	38.15613 15643 15674 15705 15735 15766
40.14307 14276 14246 14215 14185 1415	45.15796 15827 15857 15888 15918 15948
50.14124 14094 14064 14034 14004 1397	
60.13944 13914 13884 13854 13824 1379 70.13765 13735 13705 13676 13646 1361	
80.13587135581352813499134701344	
90.13411 13382 13353 13324 13295 1326	
100.13237 13208 13179 13150 13121 1309	105.16866 16895 16924 16953 16982 17010
110.13064 13035 13007 12978 12950 1292	
120.12893 12864 12836 12808 12779 1275 130.12723 12695 12666 12638 12610 1258	
140.12554 12526 12499 12471 12443 1241	
150.12387 12360 12332 12305 12277 1224	9 155.17716 17743 17771 17798 17826 17854
160.12222121951216712140121131208	5 16 5 . 17881 17908 17936 17963 17990 18018
170.12058 12031 12004 11977 11949 1192	
180.11895 11868 11842 11815 11788 1176 190.11734 11708 11681 11654 11628 1160	
200.11575 11548 11522 11495 11469 1144	
210.11416 11390 11364 11338 11312 1128	5 21 5.18687 18713 18739 18765 18791 18818
220.11259112331120711181111561113	0 225.18844 18870 18896 18922 18947 18973
230.11104 11078 11052 11027 11001 1097 240.10950 10924 10899 10873 10848 1082	
25 0 . 10797 10772 10746 10721 10696 1067	_
260.10646106201059510570105451062	
270.10496 10471 10446 10421 10396 1037	1 27 5.19607 19632 19657 19682 19707 19732
280.10347 10322 10298 10273 10248 1022	
290.10199101751015110126101021007	
30 0.10053 10029 10005 09981 09957 0993 31 0.09909 09885 09861 09837 09813 0978	
32 0.09765 09741 09718 09694 09670 0964	
33 0.09623 09599 09576 09552 09529 0950	5 33 5.20480 20504 20527 20551 20574 20597
34 0 . 09482 09459 09435 09412 09389 0936	
35 0.09343 09319 09296 09273 09250 0922 36 0.09204 09181 09158 09136 09113 0909	
370.09067 09044 09022 08999 08977 0895	
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3 0.05924 05906 05888 05871 05853 05836 4 0.05818 05801 05783 05766 05748 05731	35.24179 24197 24215 24232 24250 24267 45.24285 24302 24320 24337 24355 24372
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37 0.02913 02901 02889 02877 02865 02853	37 5.27190 27202 27214 27226 27238 27250
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28 0.00425 00420 00416 00412 00407 00403 28 5.29678 29683 29687 29691 29696 29700 29 5.29704 29709 29713 29711 29721 29786 30 0.00349 00345 00350 00356 00357 00353 00329 31 0.00349 00345 00341 00337 00333 00329 32 0.00325 00321 00317 00313 00310 00306 32 5.29736 29738 29738 29738 29736 29776 29774 32 0.00325 00321 00317 00313 00310 00306 32 5.29778 29782 29786 29790 29793 29774 33 0.00329 00295 00229 00226 00262 35 5.2981 29805 29808 29812 29816 29819 35 0.00329 00255 00259 00259 00259 00226 35 5.29844 29848 29851 29854 29858 29851 35 0.00329 00215 00213 00210 0027 00203 36 0.00329 00215 00213 00210 0037 00203 38 0.00329 00215 00213 00210 0037 00203 38 0.00329 00215 00213 00210 0037 00203 39 0.00183 00480 00177 00174 00171 00168 40 0.00166 00163 00160 00157 00155 00152 41 0.00149 00147 00144 00142 00139 00137 44 0.00169 00140 0012 00099 00097 00095 45 0.00038 00039 00037 00035 00035 00035 00035 00043 00043 00043 00043 00045 0	26	0.00480	00475	00470	00466	00461	00456								
29 9. 00399 00394 00390 00386 00382 00377 29 5.29704 29709 29713 29717 29721 29736 29750 200. 00373 00345 00341 00337 00333 00332 00328 00341 00313 00310 00383 00321 00317 00313 00310 00383 00321 00317 00313 00310 00385 00321 00317 00313 00310 00385 00321 00317 00313 00310 00385 00321 00317 00313 00310 00385 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00387 00388															
30 0 .00373 00369 00365 00361 00357 00353 31 0 .00349 00345 00341 00337 00333 00329 32 0 .00325 00321 00317 00313 00310 00306 33 0 .00302 00299 00295 00291 00287 00284 34 0 .00280 00276 00273 00269 00266 00262 35 0 .00259 00255 00252 00249 00245 00222 35 0 .00259 00255 00252 00249 00245 00222 35 0 .00259 00255 00252 00249 00245 00222 35 0 .00219 00216 00213 00210 00207 00203 38 0 .00300 00187 00194 00191 00188 00185 39 0 .00183 00180 00177 00174 00171 00168 39 0 .00183 00180 00177 00174 00171 00168 39 0 .00183 00180 00177 00174 00171 00168 40 0 .00166 00163 00160 00157 00155 00152 41 0 .00149 00447 00144 00142 00199 00137 42 0 .00134 00152 00129 00127 00124 00112 43 0 .00120 00117 00115 00113 00110 00108 44 0 .00016 00104 00102 00099 00097 00095 44 0 .00016 00104 00102 00099 00097 00095 45 0 .00008 00001 00009 00087 00085 00085 46 0 .00001 000040 00099 00097 00095 47 0 .000070 00066 00066 00065 00065 00065 00065 48 0 .00001 000079 00077 00075 00074 00072 49 0 .00050 00049 00047 00046 00044 00043 50 0 .00040 00059 00087 00035 00055 000	29	0.00425 A.00399	00330	00316 00390	00412	00407 00382	00403								
31 0.09349 00345 00317 00313 00310 00305 003052 00321 00317 00313 00310 00305 003021 00317 00313 00310 00305 003021 00317 00313 00310 00305 003021 00317 00221 00221 00221 00221 00222 00225 00222 00225									_						
328.0.0032800292800295 00291 00286 00284 335 5.29801 29808 2981229616 29919 340 00288000275 00291 00289 00256 00262 355 5.29801 29808 2981229616 29919 350 0.00280 00255 00252 00229 00225 00222 365 5.29864 29868 29871 29874 29878 29881 370 0.00280 00215 00213 00210 00807 00203 380 0.00290 00216 00213 00210 00807 00203 380 0.0020 00180 00177 00174 00191 00188 00185 399 0.00183 00180 00177 00174 00171 00168 00168 395 2.29905 29905 29912 29912 29918 399 0.00183 00180 00177 00174 00171 00168 00168 00165 00165 00165 00157 00155 00152 410 0.00169 00144 00142 00199 00137 42 0.00134 00152 00129 00127 00124 00132 00129 00127 00124 00132 00129 00127 00124 00132 00129 00127 00124 00132 00129 00127 00124 00132 00129 00129 00137 440 0.00160 00104 00102 00099 00097 00095 445 0.00008 00008 00005 00065 0	31	0.00349	00345	00341	00337	00333	00329								
34 0.00280 00276 00273 00269 00266 00262 34 5.29823 29827 29830 29834 29837 29841 350 0.00259 00225 00225 00222 00225 00222 00225 00222 00225								32	5.	2977 8	29782	29786	29790	29793	29797
35 0.0025 00255 00252 00249 00245 00242 35 5.29844 29848 29851 29854 29858 29861 29900 2025 00.00219 00216 00213 00210 00807 (0.0023 300.00219 00216 00213 00210 00807 (0.0023 300.00219 00216 00213 00210 00807 (0.0023 300.00219 00216 00213 00210 00807 (0.0023 300.00219 00216 00213 00210 00807 (0.0023 300.00219 00180 00177 00174 00171 (0.0168 30185 395.29905 29905 29905 29905 29912 29915 29918 399 0.00183 00180 00177 00174 00171 (0.0168 30165								33	5.	29801	29805	29808	29812	29816	29819
36 0.00339 00235 00232 00229 00225 00222 36 5.29864 29868 29871 29874 29878 29881 37 0.00219 00216 00213 00210 00807 00203 38 0.00209 00191 00194 00191 00188 00185 38 0.00209 00191 00194 00171 00168 38 5.29903 29905 29902 29912 29915 29918 39 0.00183 00180 00177 00174 00171 00168 40 0.00165 00163 00160 00157 00155 00152 4012 00122 00127 00124 00132 40123 00127 00124 00132 40123 00127 00124 00132 40123 00127 00124 00132 40123 00127 00124 00132 40123 00127 00124 00132 40123 00127 00124 00132 40123 00123 00123 00132 00123 0013 0013 0															
37 5.2984 2987 2989 29898 29990 29913 29915 29918 39 0.00183 00183 00183 00180 00177 00174 00171 00183 00183 00183 00180 00177 00174 00171 00183 00183 00183 00180 00160 00157 00174 00171 00168 39 5.29920 29923 29920 29929 29932 29935 29915 29918 3990 29932 29935 29915 29918 3990 29932 29935 29918 3990 29932 29935 29918 3990 29932 29935 2993	36	D. 0023 9	00235	00232	00229	00225	00222	36	5. 5.	29864 29864	29868	29871	29834	29878	29861
39 0.00183 00189 00177 00174 00171 00168 40 0.00165 00165 00167 00155 00152 41 0.00149 00147 00144 00142 00139 00137 42 0.00134 00132 00129 00127 00124 00132 43 0.00120 00117 00115 00113 00110 00108 44 0.00106 00104 00102 00099 00097 00095 45 0.00098 00091 00089 0087 00085 0085 45 0.00098 00091 00089 0087 00085 0085 45 0.00098 00091 00089 0087 00085 0085 46 0.00081 00079 00077 0075 00074 00072 47 0.00070 00066 00066 00065 00065 00065 48 0.00060 00068 00056 00065 00065 00065 48 0.00060 00068 00056 00065 00065 00065 49 0.00050 00049 00047 00046 00044 00043 48 0.00060 00088 00056 00065 00065 00065 50 0.00041 00040 00039 0037 00036 00052 49 0.00050 00049 00047 00046 00044 00043 50 0.00050 00049 00047 00046 00044 00043 50 0.00050 00049 00047 00046 00044 00043 50 0.00050 00049 00047 00046 00044 00045 50 0.00050 00049 00047 00046 00044 00045 50 0.00050 00040 00039 00037 00036 00055 50 0.00050 00040 00039 00037 00036 00055 50 0.00050 00040 00039 00039 00039 00039 50 0.00050 00040 00039 00039 00039 00039 50 0.00050 00040 00039 00039 00039 00039 50 0.00050 00040 00039 00039 00039 00039 50 0.00050 00040 00039 00039 00039 00039 50 0.00050 00050 00050 00050 00050 00050 50 0.00050 00050 00050 00050 00050 00050 50 0.00050 000	37	0.00219	00216	00213	00210	00207	00203								
40 0.09166 00163 00160 00157 00155 00152 41 0.00149 00147 00144 00142 00199 00137 42 0.00134 00132 00129 00127 00124 00122 43 0.00120 00117 00115 00113 00113 00110 00108 44 0.00106 00104 00102 00099 00097 00095 45 0.00098 00091 00089 00087 00085 00083 46 0.00081 00079 00077 00075 00074 00072 47 0.0007000066 00066 00065 00065 00065 48 0.00060 00068 00055 00055 00055 00055 00055 48 0.00060 00068 00055 00055 00055 00055 49 0.00060 00068 00061 00065 00065 50 0.00041 00040 00099 00037 00036 00035 51 0.00030 00040 00099 00037 00036 00035 52 0.00026 00062 00062 00065 00065 53 0.00020 00019 00018 00017 00017 00016 54 0.00016 00014 00013 00013 00012 00011 55 0.00016 00014 00013 00013 00012 00011 55 0.00016 00014 00013 00013 00012 00011 55 0.00016 00014 00013 00013 00012 00011 55 0.00016 00014 00013 00013 00012 00011 55 0.00016 00014 00013 00013 00012 00011 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00007 56 0.00007 00006 00006 00008 00008 00008 57 0.00008 00001 00001 00009 00008 00008 00008 58 5.3010 30100 30100 30100 30100 30101 301	38	0.00209	00197	00194	00191	00188	00185								
41 0.00149 00147 00144 00142 00139 00137 42 0.00134 00132 00139 00137 00124 00132 42 0.20134 00132 00139 00137 00134 00113 00110 00108 43 0.20104 00117 00115 00113 00110 00108 44 0.20106 00104 00102 00039 00037 00035 00035 45 0.20038 00031 00039 00037 00035 00035 45 0.20038 00031 00039 00037 00035 00035 45 0.20038 00031 00039 00037 00035 00035 45 0.20038 00038 00043 00043 00045 00054 45 0.20038 00038 00035 00035 00035 00035 46 0.20038 00039 00037 00035 00035 47 0.20038 00039 00035 00035 00035 00035 48 0.20038 00039 00035 00035 00035 00035 49 0.20038 00039 00037 00036 00035 49 0.20038 00039 00037 00036 00035 50 0.20038 00039 00037 00036 00035 50 0.20038 00038 00039 00037 00036 00035 50 0.20038 00039 00039 00037 00036 00035 50 0.20038 00039 00039 00037 00036 00035 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00039 00039 00039 00039 00039 50 0.20038 00038 00039 00039 00039 00039 50 0.20038 00038 00039 00039 00039 00039 50 0.20038 00038 00039 00039 00039 00039 00039 50 0.20038 00038 00039 00									_						
42 0.00134 00132 00129 00127 00124 00122															
43 0.0012000117 00115 00113 00110 00108	42	لمع 0.001 م	00132	00100	00197	00194	00100								
45 0.00098 00097 00077 00075 00074 00092 46 5.30010 30012 30014 30016 30018 30020 46 0.00081 00079 00077 00075 00074 00072 47 5.30032 30025 30025 30028 30029 30031 47 5.30033 30035 30035 30047 30048 30040 30044 48 0.00060 00068 00065 0006	43	0.00120	00117	00115	00118	00110	00108	43	5.	29983	29986	29988	29990	29993	29995
46 5.30022 30024 \$0026 \$0028 \$30029 \$30031 47 0.00070 00068 \$00066 \$00065 \$00065 \$00065 \$00061 48 0.00060 \$00068 \$00056 \$00055 \$00053 \$00052 \$45 5.30033 \$30035 \$30037 \$30038 \$30040 \$30042 \$45 5.30043 \$30045 \$30047 \$30048 \$30055 \$30055 \$49 5.30053 \$30045 \$30047 \$30048 \$30055 \$30055 \$45 5.30068 \$30067 \$30068 \$30066 \$30067 \$30068 \$51 0.00038 \$00032 \$00031 00030 00029 \$00028 \$52 0.00036 \$00024 \$00024 \$00023 \$00021 \$53 0.00030 \$00018 \$00017 \$00017 \$55 5.30073 \$30074 \$30075 \$3007	4 =	0.00000	0000	00102	00099	00097	00095		_						
48 0.00060 00068 00055 00055 00053 00052 49 0.00050 00049 00047 00046 00044 00043 50 0.00041 00040 00039 00037 00036 00035 51 0.00038 00038 00031 00030 00029 00028 52 0.00026 00025 00024 00023 00022 00021 53 0.00020 00019 00018 00017 00016 54 0.00015 00014 00013 00013 00012 00011 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 0008 00008 00007 55 0.00010 00010 00010 0009 0008 00008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00097 00098 0098 00998	46	0.000R1	00079	00089 00077	00087 0007=	00085	00085								
48 0.00060 00068 00055 00055 00053 00052 49 0.00050 00049 00047 00046 00044 00043 50 0.00041 00040 00039 00037 00036 00035 51 0.00038 00038 00031 00030 00029 00028 52 0.00026 00025 00024 00023 00022 00021 53 0.00020 00019 00018 00017 00016 54 0.00015 00014 00013 00013 00012 00011 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 00008 00008 00007 55 0.00010 00010 00009 0008 00008 00007 55 0.00010 00010 00010 0009 0008 00008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00007 55 0.00010 00010 00010 0009 0008 0008 00097 00098 0098 00998	47	0.00070	00068	00066	00065	00063	00061								
50 0.00041 00040 00039 00037 00036 00035 510 .0003300032 00031 00030 00029 00028 52 0.00026 00025 00024 00023 00022 00021 53 0.00026 00025 00024 00023 00021 00011 00013 00013 00013 00012 00011 55 0.00010 00010 00009 00008 00008 00007 56 0.00007 00066 00066 00066 00068 00008 00007 56 0.00007 00066 00066 00066 00068 00008 00007 56 0.00007 00066 00066 00068 00008 00007 56 0.00007 00066 00066 00068 0000	48	0.00060	00058	00056	00055	00053	00052	48	5.	30043	30045	30047	30048	30050	30051
51 0.00038/00082/00031 00030 00029 00028 52 0.00026 00022/00023 00022 00021 53 0.00020 00019 00018 00017 00011 55 0.00020 00019 00018 00017 00011 55 0.00010 00010 00009 00008 00008 56 0.00010 00010 00009 00008 00008 56 0.00017 00066 00066 00006 00008 00007 56 0.00007 00066 00066 00006 00008 00000 57 0.00004 00003 00003 00003 00002 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 58 0.00002 00008 00001 00001 00001 00001 00001 58 0.00002 00008 00001 00															
55 0.00026 00025 00024 00023 00022 00021 55 0.00020 00019 00018 00017 00016 55 0.00020 00019 00018 00013 00012 00011 55 0.00010 00010 00009 00008 000007 55 0.00010 00009 00008 00008 000007 55 0.00007 00006 00006 00008 000004 55 0.00007 00006 00006 00008 000004 55 0.00007 00006 00008 00008 000004 55 0.00008 00008 00008 00008 00008 57 0.00008 00008 00008 00008 00008 57 0.00008 00008 00008 00008 00008 58 0.00008 00008 00008 00008 00008 58 0.00008 00008 00008 00008 00008 58 0.00008 00008 00008 00008 00008 58 0.00008 00008 00008 00008 00008 58 0.0008 0008 0008 0008 0008 0008 58 0.0008 0008 0008 0008 0008 0008 0008	50 51	0.0004) 0.00049	00040	00039	00037	00036	00035								
55 0.00020 00019 00018 00017 00016 54 0.00015 00014 00013 00017 00016 55 0.00010 00010 00019 00018 00007 56 0.00010 00010 000009 00008 00005 56 0.00010 00010 00003 00003 00002 57 0.00004 00003 00003 00003 00002 58 0.00002 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 00001 00001 58 0.00002 00001 000001 000001 00001 00001 00001 00001 00001 00001 00001 00001 00001 00001 0000	52	0.00026	00025	00024	00023	00022	00021	51	5. 5	30070 30077	30071	30072	30073	30074	30075
55 0.00016 00010 00003 00013 00012 00011 54 5.30088 30089 30090 30090 30091 30092 55 0.00010 00009 00008 00008 00007 55 5.30093 300	53	0.00020	00019	00018	00017	00017	00016								
57 0.00004 00003 00003 00003 00002 00002 58 0.00002 00003 00001 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 00001 00001 58 0.00002 00001 00001 00001 00001 00001 00001	54	0.00015	00014	00013	00013	00012	00011								
58[0.00002]00001[00001]00001[00001]00001 57[5.30099]30100[30100]30100[30101]30101 58[0.00002]00001[00001]00001[00001]00001 58[5.30101]30102[30102]30102[30102]30102	55 EC	0.00010	00010	00009	80000	00008	00007								
580.00002 00001 00001 00001 00001 00001 58 5.30101 30102 30102 30102 30102 30102	57	0.00007	000006	00000	00005	00008	00004								
	58	U . U00002	00001	00001	00001	nana i	00001								
	59	0.00000	00000	00000	00000	00000	00000								

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TABLE XXIII:

To find the Latitude by two Altitudes of the Sun. LOG. RISING OR VERSED SINE.

		•			. RIS	ING OF	VF	RSED					
	·		HOUL							1 HO	UR.		
M	0"	10"	20"	30"	40"	50"	M	0"	10″	20"	30"	40"	50" .
U	Inf. Neg.						0	3.53243	53482	53721	53 959	54197	54134
	8. 9.	12230	00196	37654	COC 46	00004	1	9 84070	E 400 E	E E 1 40	E E 057 E		
1		1	02.130	31034	020-12	0.004	1	3.54670	34903	22140	00010	22008	55841
11	0.	11250	22846	33079	42230	50509	2	3.56074	56306	56537	56767	56997	57226
2		65619	71455	77448	83054	88319				i			
3	0.93284	97960		06672	10714	14575	3	3.57455	57683	57910	58137	58363	58589
4		21817	25221	23502	31660	34708	4	3.58814	59038	59262	59486	59708	59930
1 -							5	3.60152					
6	1.53480	55860	58184	60440	62639	64784	6	3.61469	61686	6190:	62120	69326	69551
3		68920	70917	72869	74778	76646	7	3.62766	62980	63194	63407	63620	63832
8		90297	91969	43739	0.1000	0630A	8 9	3.64043 3.65302	65510	104463 1168713	04075	64885	6.091
1 10				- COACO	1000	20037	<u> </u>		·	· '	1	·	·
10	2.	3.1203	00701	02091	03458	04805	10	3.66542					
11			08723	09991	11240	12479	11 12	3.67765 3.68969					
19	2.13687	14885	16066	17232	18382	19517	iŝ	3.70158					
13		28100	22836	23915 30190	31119	32093	14	3.71329					
1 1/1/1/1/1/1/1/1/1/1/1/1/1/1/1/1/1/1/1							15	3.7248	79076	7900	730:0	7 10 15	73136
1 1 16	2.33667	39 6 67	40457	41339	42211	13075	16	3.7362	73813	7400	74189	7437	74563
l lif	2.43930	14777	45616	16117	17270	18085	17	3.73624 3.74750 3.75860	74936	7512	75307	75491	75676
18	31 2.48893	349693	50486	551271	52050	152821	18	3.75860	7604	7622	76409	76592	76774
19							19			-1			
20							20						
21	2.62274 2.66312						21	3.79108 3.80159					
2							23		8137	8154	81717	81888	82059
2							24	3.62230	82400	82570	82739	82908	83077
2/	2.77403	77982	78555	79124	79689	80251	25	3.83240	83414	8358	8374	83917	84083
26	2.80809	81365	81914	82461	83005	83546	26						
2		84617	85148	85675	86199	86720	27 28	3.85242	285400	85570	185734	85897	786060
99		90779	91273	91765	92454	92739	29	3.86223 3.87192	8735	28 751 :	87679	28783	287991
30						·	-		 	 	-		
31							30						
33		99270	99719	1		l	31 32	3.89097 3.90034	90189	1903L	90498	9065	190507 190507
1 1 2	3. 3.01488	01095	00200			01049	33	3.90960	91114	H91267	91420	9157.	91724
33							34	3.91876	9202	92179	92331	92482	92632
3							35	3.92782	99035	0300	93939	03301	03530
36	3.09032	09432	09831	10227	10622	11015	36						
3		11796	12184	12570	12954	13337	37	3.94566					
33	3.13718	14097	114175	14850	15225	15597	38						
39	3.15969	10338	10/00	140/2	143/	11000		3.96311		al			
40	3.18162	18522	18881	19238	19594	19949	40 41	3.97170 3.98021	9816	9745	1984A4	19773	1927550 192793
41	3.20301	20663	21003	21351	21699	22044	42		9900	9914	99280	9941	99557
45		22132	95005	23114 95199	95750	26080	43	3.99696			4	1	1
1 4		26745	27071	27398	27720	23042		4.	000				00384
		.					44	4.00521					
45							45	4.01337					
46							47	4.02947	0.3080	03219	02317	02081	03608
4	3.33950	34250	34549	34847	35144	35439	48	4.03740					
49	3.35734	36028	36321	36613	36903	37193	49	4.04526					
50	3.37482	37770	38057	38343	38628	38912	50	4.05304					
51	3.39195	39477	39759	40039	40319	40597	51	4.06074	06202	06330	06457	05584	06711
52		41152	41427	41702	43604	43971	52 53	4.06838 4.07595	07790	1091 107842	07970	01343 08095	08220
54		14405	44670	41935	45199	45462	54	4.08344	08468	08592	08716	03840	08964
55	3 45/794	45006	16947	46500	46786	47094	55	4.09087	09210	09333	09456	09578	09701
56		47539	47795	48050	48305	48558	56	4.00823	0994	10067	10188	10310	10431
57	3.48811	40064	19315	49566	49816	50066	57	4 10559	1067.	110794	110915	411035	111155
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13 4 . 21409 14 4 . 22041							4.52390 4.52812					
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33 4.33180 34 4.33724							4.60388 4.60765					
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To find the Latitude by two Altitudes of the Sun.

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0	ō.17609	17627	17646	17664	17682	17700			.23226					
	5.17718 5.17827								.23304 .23382					
3	5.17935	17953	17971	17989	18007	18024			.23459					
4									.23536					
5	5.18149	15167	18185	18203	18220	18238	5	5	.23612	23625	23638	23650	23663	23676
6	5 18956	18273	18291	18309	18326	18344			. 23 688					
7	5.18362	18379	18397	18414	18432	18449	7	5	.23764	23776	23789	23801	23814	23826
8	5.18467 5.18572	18484	18502	18519	18631	18554			.23839 .23913					
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10	5.18070	18093	18214	18831	18848	18866			.24060					
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	5.19087						14	5	. 24278	24290	24302	24314	24 326	24338
15	5.19189	19206	19223	19240	19256	19273			.24349					
	5.19290	19307	19324	19340	19357	19374			.24421					
17	5.19390	19407	19424	19441	19457	19474			.24491 .24561					
18	5.19490 5.19590	19507	19623	19639	19656	19672			.24631					
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20	5.19787	19804	19820	19836	19852	19869	21	b	.24769	24780	24792	24803	24814	24326
22	5.19885	19901	19918	19934	19950	19966	22	5	.24837	24849	24860	24871	24882	24894
23	5.19982	19999	20015	20031	20047	20063			. 24905					
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25	5.20175	20191	20207	20223	20239	20255			. 25039					
26	5.20271	20287	20303	20319	20335	20351			.25106					
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31	5.20742	20758	20773	20789	20804	20820			.25431					
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40	5.21558	21573	21588	21602	21617	21632	40	5	.25987	25997	26007	26017	26027	26037
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5	99752	07034	99860	05292	99937	03548	99984	01803	100000		2
5		07063	99858	05321	99936	03577	99983	01832	100000		3
5	99748	07092	99857	05350	99935	03606	99983	01862	100000	00116	4
5	99746	07121	99855	05379	99934	03635	99982	01891	100000	00145	. 5
5	99744	07150	99854	05408	99933	03664	99982	01920	100000	00175	6
5	99742	07179	99852	05437	99932	03693	99981	01949	100000	00904	7
5	99740	07208	99851	05466	99931	03723	99980	01978	100000		8
5	99738	07237	99849	05495	99930	03752	99980	02007	100000		9
5	99736	07266	99847	05524	99929	03781	99979	02036	100000		
4	99734	07295	99846	05553	99927	03781	99979	02065		00320	10
	99731	07324	99844			03839					11
4				05582	99926		99978	02094		00349	12
4	99729	07353	99842	05611	99925	03868	99977	02123	99999	00378	13
4	99727	07382	99841	05640	99924	03897	99977	02152	99999	00407	14
4	99725	07411	99839	05669	99923	03926	99976	02181	99999	00436	15
4		07440	99838	05698	99922	03955	99976	02211		00465	16
4	99721	07469	99836	05727	99921	03984	99975	02240		00495	17
4	99719	07498	99834	05756	99919	04013	99974	02269	99999	00524	18
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4	99716	07527	9983	05785	99918	04042	99974	02298	99998	00553	19
4	99714	07556	99831	05814	99917	04071	99973	02327		00582	20
3	99712	07585	99829	05844	99916	04100	99972	02356		00611	21
3	99710	07614	99827	05873	99915	04129	99972	02385	99998	00640	22
3	99708	07643	99826	05902	99913	04159	99971	02414	99998	90669	23
3		07672	99824	05931	99912		99970	02443	99998	00698	24
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3	99703		99822	05960							2 5
3		07730	99821	05989	99910		99969	02501		00756	26
3			99819	96018	99909	04275	99968	02530		00785	27
3	99696	07788	99817	06047		04304	99967	02560		00814	2 8
3		07817	99815	06076	99906	04333	99966	02589		00844	2 9
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2	99689	07875	99812	06134	99904	04391	99965	02647	99996	00902	31
9			99810	06163	99902	04420	99964	02676		00931	32
3		07983	99808	06192	99901	04449	99963	02705		00960	
				06221	99900	04478	99963	02734		00989	33
3		07962	99806								34
2			99804	06250	99898	04507	99962	02763	99995	01018	35
_ 2	99678	08020	99803	06279	99897	04536	99961	02792	99995	01047	36
2	99676	08049	99801	06308	99896	04565	99960	02821	99994	01076	37
9	99673	08078	99799	06337	99894	04594	99959	02850	99994	01105	38
9	99671	08107	99797	96366	99893	04623	99959	02879	99994	01134	3 9
9			99795	06395	99892	04658	99958	02908	99993	01164	40
î			99793	06424	99890	04682	99957			01193	41
1					99889	04711	99956			01222	42
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1			99790	06482	99888	04740	99955		99992	01251	43
1	1 1		99788		99886	04769	99954			01280	44
1				06540	99885			03054		01309	45
1			99784		99883	04827		03083		01338	46
3	99652		99782	06598	99882	04856	99952			01367	47
1	99649	08368	99780	06627	99881	04885	99951	03141	99990	01396	48
1		08397	99778	06656	99879	04914	99950	03170	99990	01425	49
			99776	06685	.99878	04943		03170		01454	50
1	1			06714	99876	04972	99948	03133		01483	51
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	99635	08542	99768	06802	99872	05059	99945	03316		01571	54_
	99632	08571	99766	06831	99870	05088	99944	03345	99987	01600	55
										01629	56
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						05175	99941			01687	58
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										01745	60
1	N.sine.	IN. cos.	N.sine.	N. cos.	X sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	
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1	7	08918	99602		99431	12389	99230		98998		98737	53	ł
	8	08947	99599				99236				98732	52	١.
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1	35	09729	99526			13197	99125				98604	25	
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()	52	10221	99476				99059	15414	98805			8	
1	53	10250	99473		99279	13716	99055	15442	98800		98516	7	
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H	55	10308	99467	12043	99272	13773	99047	15500	98791	17222	98506	5	
{	56	10337	99464	12071	99269	13802	99043	15529	98787	17250	98501	4	
1	57	10366	99461	12100	99265	13831	99039	15557	98782		98496	3	
	<i>5</i> 8	10395	99458	12129	99262	13860	99035	15586			98491	2	
	59	10424	99455	12158	99258	13889		15615			98486	1	
	60	10453	99452	12187	99255	13917		15643				0	
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M.	N.sine.								N.sinc.		
0	17365	98481	19081	98163	20791		22495		24192		60
1	17393		19109	98157	20820						59
2	17422	98471	19138	98152			22552			97015	58
3	17451 17479	98466 98461	19167 19195	98146 98140		97797 97791	22580 22608	97417 97411		97008 97001	57 56
5	17508		19224	98135				97404		96994	55 ·
6	17537	98450	19252	98129			22665	97398		96987	54
7	17565		19281	98124			22693	97391		96980	
8	17594	1		98118			22722	97384		96973	53 52
9	17623					97760	22750				51
10	17651	98430				97754		97371			50
11	17680				21104	97748					49
12	17708	98420	19423	98096	21132	97742	22835	97358	24531	96945	48
13	17737	98414	19452	98090	21161	97735	22863	97351	21559	96937	47
14	17766			98084							46
15	17794			98079	21218		22920	97338			45
16	17823	98399				97717	22948				44
17	17852			98067	21275	97711	22977	97325			43
18	17880	98339	19595	98061	21303		23005	97318	24700	96902	42
19	17909	98333	19623	98056		97698	23033	97311	24728	96894	41
20	17937	98378									40
21	17966									1 - 1	39
22	17995										
23 24	18023 18052									96866	37
				l							
25 86	18031	98352			21502		23203	97271	24897		
2 6	18109							97264			
27 28	18138 18166					97648 97642					
20 29	18195							97251 97244			32 31
3 0	18224			97992							30
31	18252			97987	21672			l		ļ	
31 32	18281	98315				97623		97230 97223			29 28
33	18309										
34	18538			97969							
35	18367	98299	20079	97963							
36	18395	98294	20108	97958							24
37	18421	98288	20136	97952	21843	97525	23542	97189		96764	23
38	18452	98283	20165	97946		97579		97182			22
39	12481	98277	20193	97940	21899	97573					21
40	18509					97566	23627	97169			20
41	18538										19
42	18567	, 									18
43	18595										17
44	18624										16
45	18652										15
46	18681 18710									1	14
47 48	18738										13 12
						·				I	
49	18767							97106			11
50 51	18795 18824							9 7 100 97093			10
52	18852										9
53	18881							97079			7
54		98196							25713		
55	18938			·	f				25741		
<i>5</i> 6	18967									96623	
57	18995									96615	
58	19024								25826		
59	19052					97444	24164	97037	25854	96600	1
60	19081				22495	97437	24192	97030	25882	96593	0
	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N. sine.	M.
		90	78		77			50 ,		150	i
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TABLE XXIV. OF NATURAL SINES.

	1:	50	10	50	17		18		19			Γ
M.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos	N.sine.	N. cos.	N.sine.	N. cos.		Γ
0	25832	96593	27564	96126	29237	95630	30902	95106	32557	94552	60	1
1	25910	96585	27592	96118		95622	30929	95097			59 58	l
2	25938	96578	27620 27648			95613 95605	30957 30985	95088 95079			57	l
3 4	25966 25994	96570 96562	27676			95596	31012	95070		94514	56	١
5	26022	96555				95588	31040		32694	94504	55	ı
6	26050		.27731	96078	29404	95579	31068	95052	32722	94495	54	
7	26079	96540	27759	96070	29432	95571	31095	95043		94465	53	l
8	26107	96532		96062		95562	31123	95033			52	i
9	26135		27815			95554	31151	95024 95015	32804 32832		51 50	Į
10	26163 26191	96517	27843 27871	96046 96037		95545 95536	31178 31206	95006			49	١
11 12	26219	96502		96029		95528	31233	94997	32887	94438	48	ı
13	26247	96494		96021	29599	95519	31261	94988	32914	94128	47	l
14	26275	96486		96013		95511	31289		32942	94418	46	ł
15	26303	96479		96005		95502	31316	94970		94109	45	l
16	26331	96471		95997		95493	31344	94961	32997	94399	44	ı
17	26359	96463				95485	31372			94390 94380	43 42	ĺ
18	26387	96456	28067	95981	29737	95476	31399	94943				ĺ
19	26415	964-18	28095	95972 95964	29765 29793	95467 95459	31427 31454	94933 94924		94370 94361	41 40	ı
20	26443 26471	96440 96433	28123 28150			95459	31482	94915			39	ı
21 22	26500	96425	28178			95441	31510			94342	38	ı
23	265 28	96417		95940	29876	95433	31537	94897	33189	94332	37	ı
24	26 556	96410	28234	95931	29904	95424	31565				36	ı
25	26581	96402	28262	95923		95415	31593		33244	94313	35	ı
26	26612	96394				95407	31620			94303	34	l
27	26640	96386	28313		29987	95398	31648 31675	94860 94851	33 2 98 33326	94293 94284	33 32	ı
28	26668 26696	96379 96371		95898 95890		95389 95380	31703	94842	33353	94274	31	ı
29 30	26724	96363				95372	31730		33381	94264	30	į
$\frac{30}{31}$	26752	96355	28429	95874	30098	95363	31758	94823	33408	94254	29	
32	26780	96347		95865		95354	31786			94245	28	ĺ
33	26308	96340			30154	95345	31813	94805	33463		27	ı
34	26836	96332	28513	95849		95337	31841	94795			26	ı
35	26864	96324	28541	95841	30209	95328	31868	94786 94777	33518 33545	94215 942 0 6	25 24	
36	26892	96316	23569	95832	30237	95319	31896					
37	26920 26948	96308 96301	28597 28625	95824 95816	30265 30292	95310 95301	31923 31951	94768 94758	33573 33600	94196 94186	23 22	İ
3 8 39	26976	96293	28652	95807	30320	95293	31979			94176	21	
40	27004	96285	28680		30348	95284	32006	94740	33655	94167	20	
41	27032		28708	95791	30376	95275	32034	94730	33682	94157	19	ı
42	270G0	96269	28736	95782	30403	95266	32061	94721	33710	94147	18	
43	27038	96261	28764	95774	30431	95257	32089	94712	33737	94137	17	
44	27116	96253				95248	32116	94702			16	ĺ
45 40	27144	96246 96238				95240	32144 32171	94693 94684	33792 33819	94118 94108	15 14	l
46 47	27172 27200	96238		95749 957 4 0		95231 95222	32171	94674		94098	13	ı
48	27228	96222	28903	95732	30570	95213	32227	94665		94088	12	ĺ
49	27256	96214		95724	30597	95204	32254	94656	33901	94078	11	ı
50	27284	96206	28959			95195	32282	94646	33929		10	į
51	27312	96198		95707	30653	95186	32309	94637	33956	94058	9	ı
52	27340	96190				95177	32337	94627		94049	8	i
53	27368 27396						323 64 3239 2				6	ı
54	27396		29098								5	ı
55 56	27424 27452						32419				4	ı
57			29154			95133	32474	94580			3	ĺ
58		96142					32502		34147	93589	2	l
59	27536	96134	29209	95639		95115	32529	94561				ı
60												l
											M.	· _
	74	lo ,	7:	30	7			10			I	Ĺ
	60 27564 96126 29237 95630 30902 95106 32557 94552 34202 93909											

TABLE XXIV

1_							AL SI					
L			00		lo	22			30	. 24		
ſ		N.sine.										
1	0	34202 34229	939 6 9 93959							,	91355	60
1	2	34257	93949	35891	93337				92039			59 58
ı	3	34284	93939	35918	93327							57
1	4	34311	93929		93316					40780	91307	56
1	5	34339 34366	93919 93909	35973	93306							55
ł				36000	93295	37622	92653			40833	91283	5.1
ł	8	34393 34421	93899 93889	36027 36054	93285 93274	37649 37676		39260 39287	91971 91959	40860		53
1	9	34448	93879	36081	93264		92620			40886 40913		52 51
1	10	34475	93869	36108			92609	39341	91936	40939	91236	50
1	11	34503	93859	36135	93243	37757	92598		91925	40966		49
-	12	34530	93849	36162	93232	37784	92537	39394	91914	40992	91212	48
1	13	34557	93839	36190	93222	37811	92576	39421	91962	41019		47
	14 15	34584 34612	93829 93819	36217	93211	37838	92565		91891	41045		46
1	16	34639	93809	36244 36271	93201 93190	37865 37892	92554 92543	39474 39501	91879 91863	41072	91176 91164	45
	17	34666	93799	36298	93180	37919	92532	39528	91856	41125	91152	44 43
1	18	34694	93789	36325	93169	37916	92521		91845		91140	42
1	19	34721	93779	36352	93159	37973			91833	41178	91128	41 .
1	20	34748	93769		93148	37999	92499	39608	91822	41204	91116	40
1	21	34775	93759	36406		38026			91810	41231	91104	39
1	22 23	34803 34830	93748 93738	36434 36461	931 27 93116	38053 38080	92477 92466		91799	41257	91092	38
1	24	34857	93728	36488	93106		92455	39715	91787 91775	41284 41310	91080 91068	37 36
	25	34884	93718	36515	93095	38134	92444	39741	91764	41337		35
ļ	26	34912	93708		93084		92432		91752	41363	91044	34
	27	34939	93698	36569	93074		92421	39795	91741		91032	33
Į į	28	34966	93688	36596	93063		92410		91729	41416		32
1	29 30	34993 35021	9 3677 9 3 667	36623	93052 93042	38241	92399	39848	91718	41443	91008	31
Ĺ				36650		38268	92388	39875	91706	41469	90996	30
1	31 32	35048 35075	93657 93647	36677 36704	930 3 1 93020	38295 38322	92377. 92366	39902 39928	91694 91683	41496 41522	90984 90972	29 28
1	33	35102	93637	36731	93010		92355		91671	41549	90960	27 27
	34	35130	93626	36758	92999						90943	26
1	35	35157	93616	36785	92988		92332		91648	41602	90936	25
H	36	35184	93606		92978	38430			91636	41628	90924	24
ì	57	35211	93596	36839	92967	38456	92310	40062	91625	41655	90911	23
1	38 39	35239 35266	93585 93575	36867 36894	9 2 956 9 2 945	38483 38510		40088 40115	91613 91601	41681 41707	90899 90887	22 21
	40	35293	93565	36921	92935			40141	91590		90875	20
	41	35320	93555	36948	92924	38564		40168	91578	41760	90863	19
	42	35347	93544	36975	92913	38591	92254	40195	91566	41787	90851	18
1	43	35375	93534	37002	92902	38617	92243	40221	91555	41813	90839	17
1	44	35402	93524	37029	92892	38644		40248		41840	90826	16
ı	45 46	35429 35456	93514 93503	37056 37083	92881 92870	38671 38698	92220 92209	40275 40301	91531 91519	41866 41892	90814 90802	15 14
П	47	35484	93493	37110	92859		92198				90790	13
1	48	35511	93483	37137	92849	38752			91496	41945	90778	12
1	49	35538	93472	37164	92838	38778	92175	40381	91484	41972	90766	11
1	50	35565	93462	37191	92827	38805	92164		91472	41998	90753	10
	51	35592	93452	37218	92816						90741	9
1	52 53	35619 35647	93441 93431	37245	92805 92794			40461 40488	91449	42051 42077	90729	8
	54	35674				38912	92119	40514				6
	55	35701	93410	37326	92773				91414		90692	5
ļ.	56	35728	93400	37353	92762				91402	42156	90680	4
	57	35755	93389	37380	92751	38993	92085	40594	91390		90668	3
	58	35782								42209 42235	90655	2
	59 60	35810 35837	93368 93358	37434 37461	92729 92718		97007	40647 40674			90643 90631	1 0
		N. cos.	Nain	N 555	32110							М.
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OF NATURAL SINES.

	1 2/	50	20		NATUR	70	2	80 (290			_
M.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.		Т
0	42262	90631	43837	89879	45399	89101			(87462	60	7
1	42288		43863	89867	45425	89087				87448	59	ı
2	42315			89854	45451	89074 89061		88267		87434 87420	58 57	ı
3	42341 42367	90594 90582	43916 43942	89841 89828	45477 45503					87406	56	1
5	42394	90569	43968	89816	45529	39035				87391	55	1
6	42420	90557	43994	89803	1 .	89021	1	88213		87377	54	ı
7	42446	90545	44020	89790	45580	89008	47127	88199	48659	87363	53	1
8	42473		44046	89777	45606	88995				87349	52	ı
9	42499			89764						87335	51	ł
10	42525			89752						87321	50	1
111	42552 42578			89739 89726						87306 87292	49 48	1
12										87278	47	-
13	42601 42631	90470 90458	44177	89713 89700		88928 88915			48811 48837	87264	46	1
14	42657		44229	89687						87250	45	1
16	42633	1		89674			1			87235	44	1
17	42709	90421	44281	89662	45839	88875				87221	43	1
18	42736	96403	44307	89649	1	88862	·			87207	42	1
19	42762	9 0396								87193	41	1
20	42788	90383		89623						87178	40	1
21 22	42815 42341	90371		89610 89597				88006 87993		87164 87150	39 38	l
23	42867	90346				88795		87979		87136	37	ı
24	42894	90334				88782		87965	49090	87121	36	1
25	42920	90321	44490	89558	46046	88768	47588	87951	49116	87107	35	1
26	42946	90309		89545					49141	87093		1
27	42972	90296		89532						87079	33	1
28	42999				46123					87064	32	ı
29 30	43025 43051	90271 90259	44594 44620	89506 89493	46149 46175					87050 87036	31 30	1
												1
31 32	43077 43104	90246 90233		89480 89467		88688 88674				87021 87007	29 28	1
33	43130		44698	89454						86993	27	ı
34	43156			89441						86978	26	1
35	43182	90196		89428					49369	86964	25	1
36	43209	90183	44776	89415	46330	88620	47869	87798	49394	86949	24	1
37	43235	90171	44802	89402		88607		87784		86935	25]
38	43261	90158	44828	89389		88593				86921	23	Ł
39 40	43287 43313	90146 9013 3	44854 44880	89376 89363				87756 87743		86906 86892	21 20	l
41	43340	90120	44906	89350					49521	86878	19	I
42	43366	90108	44932	89337						86863	18	ŀ
43	43392	90095	44958	89324		88526		87701	49571	86849	17	1
44	43418	90082	44984	89311	46536	88512		87687		86834	16	1
45	43445	90070	45010	89298		88499				86820	15	1
46	43471	90057	45036	89285		88485		87659	49647	86805	14	1
47	43497 43523	90045 90032	45062 45088	89272 89259	46613 46639	88 4 72 88 4 58		87645 87631	49672 49697	86791 86777	13	L
49	43549											{
50	43575	90019 90007	45114 45140	89245 89232	46664 46690	88445 88431		87617 87603	49723 49748	86762 86748	11 10	
51	43602	89994		89219	46716	88417		87589	49773	86733	9	
52	43628	89981	45192	89206	46742	88404	48277	87575		86719	8	
53	43654		45218					87561	49821	86704	7	1
54	43680			89180	46793	88377	48328	87546	49849	86690	6]
55	43706							87532			5	1
56	43733		45295	89153							4	1
57 58	43759 43785			89140 89127		88336 88322					3	ı
59	43811			89114		88308			49950 49975	8663 3 8661 7	1	١
60	43837									86603	ò	1
1							N. cos.				M.	4.
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Ī	M.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sinc.	N. cos.	N.sine.	N. cos.	
ľ	0	50000	86603	51504	85717	52992	84805			55919	82904	60
ı	1	50025	86588				84789			55943	82887	59
١	2	50050	86573	51554	85687	53041	84774		83835	55968	82871	58
ı	.3	50076	86559	51579		53066	84759	54537	83819		82855	57
1	4	50101	86544			53091	84743		83804		82839	56
l	5	50126	86530	51628	85642						82842	5
Ł	6	50151	86515	51653	85627	53140	84712	54610	83772	56064	82306	54
ľ	7	50176	86501	51678	85612	53164	84697	54 635	83756	56088	82790	5
I	8	50201	86486	51703	85597	53189	84681	54659	83740		82773	5
l	9	50227	86471	51728	85582		84666	54683	83724		82757	5
ı	10	50252	86457	51753	85567	53238	84650	54703	837 08		82741	5
١	11	50277	86442	51778	85551	53263	84635			56184	82724	4
L	12	50302	86427	51803	85536	53288	84619	54756	83676	56208	82708	4
I	13	50327	86413	51828	85521	53312	84604	54781	83660	56232	82692	4
١	14	50352	86398	51852	85506	53337	84588	54805	83645	56256	82675	4
İ	15	50377	86384	51877	85491	53361	84573	54829	83629	56280	82659	4.
ı	16	50403	86369	51902	85476	53386	84557	54854	83613	56305	82643	4
l	17	50428	86354	51927	85461	53411	84542	5487 8	83597	56329	82626	4:
١	18	50453	86340	51952	85446	53435	81526	54902	83581	56353	82610	4
ľ	19	50478	86325	51977	85431	53460	84511	54927	83565	56377	82593	4
l	20	50503	86310	52002	85416		84495		83549	56401	82577	4
ı	21	50528	86295	52026	85401	53509	84480		83533		82561	3
1	22	50553	86281	52051	85385	53534	84464	54999	83517	56449	82544	3
l	23	50578	*86266	52076	85370	53558	84448	55024	83501		82528	3
l	24	50603	86251	52101	85355	53583	84433		83485	56497	82511	3
r	25	50628	86237	52126	85340	53607	84417	55072	83469	56521	82495	3.
l	26	50654	86222	52151	85325		84402	55097	83453		82478	3
Ì	27	50679	86207	52175	85310		84386		83437	56569	82462	3
ı	28	50704	86192	52200	85294		84370	55145	83421	56593	82446	3
l	29	50729	86178	52225	85279	53705	84355	55169	83405		82429	3
l	30	50754	86163	52250	85264		84339	55194	83389	56641	82413	3
ŀ												
ı	31	50779	86148	52275	85249	53754 53779	84324	55218	83373		82396	2
ı	32	50804	86133	52299	85234		84308	55242	83356	56689	82330	2
ı	33	50829	86119	52324 52349	85218 85203		84292	55266	83340		82363	2
l	34	50854 50879	86104	52374			84277	55291	83324		82347	2
1	35 3 6	50904	86089 86074	52399	85188 85173	53853 53877	84261 84245	55315 55339	83308 83292	56760	82330	2
ŀ										56784	82314	2
l	37	50929	86059	52423	85157	53902	84230	55363	83276	56808	82297	2
I	38	50954	86045	52448	85142			55388	83260		82281	2
ı	39	50979	86030	52473	85127	53951	84198		83244	56856	82264	2
l	40	51004	86015	52498	85112		84182	55436			82248	2
١	41	51029	86000	52522	85096	54000	84167	55460	83212	56904	82231	1
L	42	51054	85985	52547	85081	54024	84151	55484	83195		82214	1
ı	43	51079	85970	52572	85066	54049	84135	55509		56952	82198	1
١	44	51104	85956	52597	85051	54073	84120				82181	1
ı	45	51129	85941	52621	85035		84104	55557	83147		82165	1
l	46	51154	85926	52646	85020		84088		83131		82148	1
l	47	51179	85911	52671	85005	54146	84072	55605	83115		82132	1
-	48	51204	85896	52696	84989	54171	84057	55630		57071	82115	1
ĺ	49	51229	85881	52720	84974	54195	84041	55654	83082	57095	82098	1
I	50	51254	85866			54220	84025	55678	83066		82082	1
ŀ	-61	51279	85851	52770				55702			82065	
l	52	51304	85836		84928						82048	
l	53	51329	85821		84913	54293	83978	55750	83017	57191	82032	
L	54	51354	85806			54317				57215		
l	55	51379	85792	5 2869		54342				57238	81999	
ı	56	51404	85777	52 893								
ı	57	51429	85762	52918	84851	54391	83915					;
ı	58	51454	85747	52943	84836	54415	83899	55871				
ĺ	59	51479	85732			54140				57334	81932	
ı	60	51504	85717	52992		54464					81915	(
		N cos	Naine	N cos	N sine	N cos	Neina	N cos	Neine	N. cos.	Neine	N.
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OF NATURAL SINES.

		A .				AL SIN						
<u></u>	35		36		37		38		39			匚
M.		N. cos.				N. cos.			N.sine.			
0	57358	81915	58779	80902	60182	79864	61566	78801	62932	77715	60	
1	57381	81899 8188 2	58802	80885	60205	79846	61589	78783	62955	77696	59	
2 3	57405 57429	81865	58826 58849	80867 80850	60228 60251	79829 79811	61612 61635	78765 78747	62977	77678	58	l
4	57453	81848	58873	80833	60274	79793	61658	78729	63000 63022	77660 77641	57	ı
5	57477	81832	58896	80816	60298	79776	61681	78711	63045	77623	56 55	
6	57501	81815	58920	80799	60321	797 58	61704	78694	63068	77605	54	
7	57524	81798	58943	80782	60344	79741	61726	78676	63090	77586	53	i
8	57548	81782	58967	80765	60367	79723	61749	78658	63113	77568	52	1
9	57572	81765	58990	80748	60390	79706	61772	78640	63135	77550	51	ŀ
10	57596	81748	59014	80730	60414	79688	61795	78622	63158	77531	50	
11	57619	81731	59037	80713	60437	79671	61818	78604	63180	77513	49	
12	57643	81714	59061	80696	60460	79653	61841	78586	63203	77494	48	!
13	57667	81698	59084	80679	60483	79635	61864	78568	63225	77476	47	•
14	57691	81681	59108	80662	60506	79618	61887	78550	63248	77458	46	
15	57715	81664	59131	80644	60529	79600	61909	78532	63271	77439	45	
16 17	57738 57762	81647 81631	59154 59178	80627 80610	60553 60576	79583	61932	78514	63293	77421	44	1
18	57786	81614	59201	80593	60599	79565 79547	61955 61978	78496 78478	63316 63338	77402	43	
			59225							77384	42	
19 20	57810 57833	81597 81580	59225 59248	80576 80558	60622 60645	79530 79512	62001 62024	78460	63361	77366	41	
21	57857	81563	59272	80541	60663	79512 79494	62046	78442 78424	63383 63406	77347 77329	40	
22	57881	81546	59295	80524	60691	79177	62069	78405	63428	77310	39 38	
23	57904	81530	59318	80507	60714	79459	62092	78387	63451	77292	37	
24	57928	81513	59342	80489		79441	62115	78369	63473	77273	36	
25	57952	81496	59365	80472	60761	79424	62138	78351	63496	77255	35	
26	57976	81479	59389	80455	60784		62160	78333	63518	77236	34	
27	579 99	81462	59412	80438	60807		62183	78315	63540	77218	33	
28	58023	81445	59436	80420	60830	79371	62206	78297	63563	77199	32	
29	58047	81428	59459	80403	60853	79353	62229	78279	63585	77181	31	
30	58070	81412	59482	80386	60876	79335	62251	78261	63608	77162	30	
31	58094	81395	59506	80368	60899	79318	62274	78243	63630	77144	29	
32	58118	81378	59529	80351	60922	79300	62297	78225	63 653	77125	28	
33	58141	81361	59552	80334	60945	79282	62320	78206	63675	77107	27	١
34 35	58165 58189	81344 81327	59576 59599	80316 80299	60968 60991	79264 79247	62342 62365	78188	63698	77088	26	
36	58212	81310	59622	80282	61015	79229	62383	78170 78152	63720 63742	77070 77051	25 24	١.
37	58236	81293										ĺ
38	58260	81276	59646 59669	80264 80 2 47	61038 61061	79211 7 9193	62411 62433	78134 78116	63765	77033	23	
39	58283	81259	5 9693	80230	61084	79176	62456	78098	63787 63810	77014 76996	22 21	
40	58307	81242	59716	80212	61107	79158	62479	78079	63832	76977	20	
41	5 83 3 0	81225	5973 9	80195	61130	79140	62502	78061	63854	76959	19	
42	58354	81208	597 63	80178	61153	79122	62524	78043	63877	76940	18	İ
43	58378	81191	59786	80160	61176	79105	62547	78025	63899	76921	17	i
41	58401	81174	59809	80143	61199	79087	62570	78007	63922	76903	16	
45	58425	81157	59832	80125	61222	79069	62592	77988		76884	15	
46	58449	81140	59356	80108	61245	79051	62615	77970	63966	76866	14	
47	58472	81123	5987 9	80091	61268	79033	62638	77952	63989		13	!
48	58496	81106	59902	80073	61291	79016	62660	77934	64011	76828	12	1
49	58519	81089	59926	80056	61314	78998	62683	77916	64033	76810	11	
50 51	58543	81072	59949	80038	61337	78980	62706	77897	64056	76791	10	H
51 52	58567 58590	81055 81038	59972 59995	80021	61360	78962	62728	77879	64078	76772	9	
53	58614	81021	60019	80003 79986	61383 61406	78944 78926	62751	77861	64100	76754		
54	58637	81004		79968	.61429			77843 77824			6	
55	58661	80987	60065			78891						
56	58684				61474		62819 62842			76698	5	
57	58708		60112		61497						3	l
58	58731		60135						64234			
59	68755		60158		61543				64256			:
60	58779		60182	79864	61566	78801	62932	77715	64279	76604		l
	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine.	N. cos.	N.sine	М.	•
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OF NATURAL SINES.

		1 4	00	41	0	NATUR	O SIN		30		10		_
+	M.					N.sine.							+
-	0		76604	65606	75471	66913	74314		73135	69466	71934	60	1
-	ĭ	,	76586				74295		73116			59	1
1	2	64323		65650			74276					<i>5</i> 8	1
-	3	64316					74256					57	1
-	4	64368					74237	68285	73056		71853	56	1
. [5 6	64390 64412		65716 65 73 8	75375 75356	67021 67043	74217 74198	68306 68327	73036 73016	69 57 0 69591	71833 71813	55 54	1
\vdash				65759	75337	67064	74178		72996		71792		-
	7	61435 64457			75318			68349 68370		69612 69633		53 52	
1	9	64479			75299		74139		72957	69654		51	
-)	10	64501			75280							50	1
1	11	64524			75261	67151	74100	68434		69696		49	ł
L	12	64546		65869	75241	67172	74080	68455	72897		71691	48	1
	13	64568	76361	65891	75222	67194	74061	68476	72877	69737	71671	47	ı
1	14	64590	76342	65913	75203		74041	68497				46	ı
	15	64612	76323 76304	65935 65956	75184 75165		74022 74002	68518 68539	72837 72817			45 44	
-	16 17	64635 64657	76286	65978	75146	67280	73983	68561	72797	69821	71590	43	1
	18	64679	76267	66000	75126	67301	73963		72777		71569	42	
-	19	64701	76248	66022	75107	67323		68603		69862	71549	41	1
	20	64723		66044	75088	67341		68624				40	1
1	21	64746	76210	66066	75069	67366	73904	68645	72717	69904	71508	39	1
	22	64768	76192		75050							38	
	23	64790	76173		75030		73865	68688	72677	69946		37	1
_	24	64812	76154	66131	75011	67430	73846	68709	72657	69966	71447	36	-
-	25	64834	76135	66153 661 7 5	74992 74973	67452 67473	73826	68730		69987 70008	71427 71407	35 34	١
1	26 27	64856 64878	76116 76097	66197	74953		73806 73787	68751 68772	72617 72597	70029		33	
.	23	64901	76078	66218	74934	67516	73767	68793		70049	71366	32	
1	29	64923	76059	66240	74915	67538	73747	68814	72557	70070		31	
1	30	64945	76041	66262	74896	67559	73728	68835	72537	70091	71325	30	
-	31	64967	76022	66284	74876	67580	73708	68857	72517	70112	71305	29	1
1	32	64989	76003	66306	74857	67602	73688	68878	72497	70132		28	1
	33	65011	75984		74838		73669	68899	72477 72457	70153 70174		27 26	ì
	34 35	65033 65055	75965 75946	66349 66371	74818 74799	67645 67666	73649 73629	68920 68941	72437			25	1
1	3 6	65077	75927	66393	74780	67688	73610	68962	72417	70215		24	1
-	37	65100	75908	66414	74760	67709	73590	68983	72397	70236	71182	23	1
1	38	65122	75889	66436	74741	67730	73570	69004	72377		71162	22	
1	39	65144	75870	66458	74722	67752	73551	69025	72357	70277	71141	21	;
ł	40	65166	75851	66480	74703		73531	69046	72337			20	:
ł	41	65188	75832	66501 66523	74683 74664		73511 73491	69067 69088	72317 72297	70319 70339		19 18	•
-	42	65210	75813										1
ì	43	65232	75794 75775	66545 66566	74644 74625		73472 73452		72277 72257	70360 70381		17 16	1
1	44 45	65254 65276	75756		74606		73432					15	
1	46	65298	75738	66610	74586		73413	69172	72216	70422	70998	14	
1	47	65320	75719	6 6632	74567		73393					13	1
	48	65342	75700	. 66653	74548		73373	69214	72176			12	_
Г	49	65364	75680	66675	74528		73353	69235	72156			11	
1	50	65386		66697	74509		73333					10 9	
j	51 - 52	65408 65430	75642 75623	66718 66740	74489 74470		73314 73294	69 277 69 2 98	72116 72095			8	
1	- 52 - 53	65452							72075		70855	7	
	54	65474			74431		73254					6	1
1	55	65496	75566	66805	74412	68093	73234		72035	70608	70813	5	1
ł	56	65518		66827	74392	68115	73215	69382	72015	70628	70793	4	1
1	57	65540	75528		74373		73195					3	1
1	58	65562					73175	69424				2	
1	59 60	65684 65606		66891 66913	74334 74314		73155 73135		71954 71934			1	1
; -		N. cos.										M.	·
+-			1 .51ne.		30		70		50		50	172.	÷
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Of Logarithmic Sines, Tangents, and Secants to every Point and Quarter Point of the Compass.

	Co. sine.	Sine.	Co. tang.	Tangent.	Co. secunt.	Secant.	Point
4	9.84949	9.84949	10.00000	10.00000	10.15051	10.15051	4
3 💈	9.82708	9.86979	9.95729	10.04271	10.13021	10.17292	4 1
3 I	9.80236	9.88819	9.91417	10.08583	10.11181	10.19764	4 🛔
3 <u>1</u>	9.77503	9.90483	9.87020	10.12980	10.09517	10.22497	4 }
3	9.74474	9.91985	9.82489	10.17511	10.08015	10.25526	5
2 🖣	9.71105	9.93335	9.77770	10.22230	10.06665	10.28895	5 🖠
2 4	9.67339	9.94543	9.72796	10.27204	10.05457	10.32661	5 🛓
2 1	9.63099	9.95616	9.67483	10.32517	10.04384	10.36901	5
2	9.58284	9.96562	9.61722	10.38278	10.03438	10.41716	6
1 1 1	9.52749	9.97384	9.55365	10.44635	10.02616	10.47251	6 🛔
	9.46282	9.98088	9.48194	10.51806	10.01912	10.53713	6 🛔
ī #	9.38557	9.98679	9.39879	10.60121	10.01321	10.61443	6 1
1	9.29024	9.99157	9.29866	10.70134	10.00843	10.70976	7
0 1	9.16652	9.99527	9.17125	10.82875	10.00173	10.83348	7 1
0 1	8.99130	9.99790	8.99340	11.00660	10.00210	11.00870	7 🛔
0.1	8.69080	9.99948	8.69132	11.30868	10.00052	11.30920	7 1
0	Inf. neg.	10.00000	Inf. neg.	Infinite.	10.00000	Infinite.	8
Points.	Sine.	Co. sine.	Tangent.	Co. tang.	Secant.	Co. secant.	

TABLE XXVI.

LOGARITHMS OF NUMBERS.

No. 1—100. Log. 0.00000—2.00000.									
N.	Log.	N.	Log.	N.	Log.	N.	Log.	N.	Log.
1	0.00000	21	1.32222	41	1.61278	61	1.78533	81	1.90549
2	0.30103	22	1.34242	42	1.62325	62	1.79239	82	1.91381
3	0.47712	23	1.36173	43	1.63347	63	1.79934	83	1.91908
٠ 4	0.60206	24	1.38021	44	1.64345	64	1.80618	84	1.92428
5	0.69897	25	1.39794	45	1.65321	65	1.81291	85	1.92942
6	0.77815	26	1.41497	46	1.66276	66	1.81954	86	1.93450
7	0.84510	27	1.43136	47	1.67210	67	1.82607	87	1.93952
8	0.90809	28	1.44716	48	1.68124	68	1.83251	88	1.94413
9	0:95424	29	1.46240	49	1.69020	69	1.83885	89	1.94959
10	1.00000	30	1,47712	50	1.69897	70	1.84510	90	1.95124
11	1.04139	31	1.49136	51	1.70757	71	1.85126	91	1.95904
12	1,07918	32	1.50515	52	1.71600	72	1.85733	92	1.96379
13	1.11394	33	1.51851	53	1.72428	73	1.86332	93	1.96848
14	1.14613	34	1.53148	54	1.73239	74	1.86923	94	1.97313
15	1.17609	35	1.54407	55	1.74036	75	1.87506	95	1.97772
16	1.20412	36	1.55630	56	1.74819	76	1.88081	96	1.98227
17	1.23045	37	1.56820	57	1.75587	77	1.88649	97	1.98677
18 ,	1.25527	38	1.57978	58	1.76343	78	1.89209	98	1.99123
19	1.27875	39	1.59106	59	1.77085	79	1.89763	99	1.99564
20	1.30103	40	1.60206	60	1.77815	80	1.90309	100	2.00000

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LOGARITHMS OF NUMBERS.

N.	o. 100-	1600.							T.o	g. 00000	204	19
	No.	1 0	1	2	3	4	5	6	7	1 8	9	1
1	100	00000	00043	00087	00130	00173	00217	00260	00303	00346	00389	-
	101	00432										i
	102	00860	00903	00945	00988	01030	01072	01115	01157		01242	
	103	01284	01326	01368	01410	01452			01578		01662	
!	104	01703	01745	01787	01828	01870		01953	01995	02036	02078	_[
	105	02119	02160	02202	02243	02284	02325	02366	02407	02449	02490	1
	106	02531	02572	02612	02653	02694	02735	02776	02816		02898	
	107	02938				03100 03503	03141 03543	03181 0358 3	03222 03623		03302 03703	
	100	03743				03902	03941	03981	03023		04100	1
į	110	04139	04179	04218		04297	04336		04415		04493	-
	111	04532	04571	04610		04689	04727		04805		04883	1
1	112	1 04922	04961	04999	05038	05077	05115	05154	05192		05269	1
	113	05308	05346	05385	05423	05461	05500	05538	05576		05652	ł
	114	05690	05729	05767	05805	05843	05881	05918	05956	05994	06032	1
- 1	115	06070	06108	06145	06183	06221	06258	06296	06333	06371	06408	1
- 1	116	06446	06483		06558	06595	06633		06707		06781	
	117	06819	06856	06893	06930	06967	07004	07041	07078		07151	1
	118 119	07188	07225 07591	07962 07628	07298 07664	07335 07700	07372 07737	07408 07773	07445 07809	0748 2 07346	07518 07882	1
												┨.
	120 121	07918 08279	07954 08314		08027 08386	08063 08422	08099 08458	08135 08493	08171 08529	08207 08565	08243 08600	
- 1	122	08636	08672	08707	08743	08778	08814	08849	08884		089 55	1
- 1	123	08991	09026	09061	09096	09132	09167	09202	09237	09272	09307	1
	124	09342	09377	09412	09447	09482	09517	09552	09587	09621	09656	1
ŀ	125	09691	09726	09760	09795	09830	09864	09899	09934	09968	10003	1
- 1	126	10037	10072	10106	10140	10175	10209	10243	10278	10312	10346	i
- 1	127	10380	10415	10449	10483	10517	10551	10585	10619	10653	10687	1
- 1	128	10721	10755	10789	10823	10857	10890	10924	10958	10992	11025	1
- [-	129	11059	11093	11126	11160	11193	11227	17261	11294	11327	11361	
	130	11394	11428	11461	11494	11528	11561 11893	11594	11628	11661	11694 12024	1
- 1	131 13 2	11727 12057	11760 12090	11793 12123	11826 12156	11860 12189	12222	11926 1 225 4	11959 1 22 87	11992 12320	12352	ł
- 1	133	12385	12418	12450	12483	12516	12548	12581	42 613		12678	1
	134	12710	12743	12775	12808	12840	12872	129	12937	12969	13001	1
 	135	13033	13066	13098	13130	13162	13194	13226	13258	13290	13322	1
	136	13354	13386	13418	13450	13481	13513	13545	13577	13609	13640	l
- 1	137	13672	13704	13735	13767	13799	13830	13862	13893	13925	13956	l
- 1	138	13988 14301	14019 14333	14051 14364	14082	14114	14145 14457	14176 14489	14208	14239	14270	
-	139				14395				14520	14551	14582	
	140	14613 14922	14644 14953	14675 14983	14706 15014	14737 15045	14768 15076	14799 15106	14829 15137	14860	14891	1
- 1	141 142	15229	15259	15290	15320	15351	15381	15412	15442	15168 15473	15198 15503	İ
- 1	143	15534	15564	15594	15625	15655	15685	15715	15746	15776	15806	
	144	15836	15866	15897	15927	15957	15987	16017	16047	16077	16107	ł
	145	16137	16167	16197	16227	16256	16286	16316	16346	16376	16406	1
	146	16435	16465	16495	16524	16554	16584	16613	16643	16673	16702	
	147	16732	16761	16791	16820	16850	16879	16909	16938	16967	16997	1
- 1	148	17026	17056	17085	17114	17143	17173	17202	17231	17260	17289	٠.
L	149	17319	17348	17377	17406	17435	17464	17493	17522	17551	17580	}
- 1	150	17609	17638	17667	17696	17725	17754	17782	17811	17840	17869	1
	151	17898 18184	17926 18213	17955 18 24 1	17984 18 2 70	18013 18298	18041 18 32 7	18070 18355	18099 18 3 84	18127 18412	18156 18441	1
- 1	152 153	18469	18498	18526	18554	18583	18611	18639	18667	18696	18724	
	154	18752	18780	18808	18837	18865	18893	18921	18949	18977	19005	
-	155	19033	19061	19089	19117	19145	19173	19201	19229	19257	19285	
ı	156	19312	19340	19368	19396	19424	19451	19479	19507	19535	19562	
- 1	157	19590	19618	19645	19673	19700	19728	19756	19783	19811	19838	
	158	19866	19893	19921	19948	19976	20003	20030	20058	20085	20112	١.
1	159	20140	20167	20194	20222	20249	20276	20303	20330	20358	20385	l
1-	No.	0	1	2	3	4	5	6	7	8	9	!

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TABLE XXVI.

No.	. 1600-	2200).						Log	. 20412	3424	12.
7	No.	0	1	2	3	4	5	6	7	8	9	$\overline{1}$
i-	160	20412	20439	20466	20493	20520	20548	20575	20602	20629	20656	1
Ì	161	20683	20710	20737	20763	20790	20817	20844	20871	20898	20925	ļ
	162	20952	20978	21005	21032	21059	21085	21112	21139	21165	21192	l
1	163 164	21219 21484	21245 21511	21272 21537	21299 21564	21325 21590	21352 21617	21378 21648	21405 21669	21431 21696	21458 21722	
-		21748	21775	21801	21827	21854	21880	21906	21932	21958	21985	ł
ı	165 166	22011	22037	22063	22089	22115	22141	22167	22194	22220	22246	
	167	22272	22298	22324	22350	22376	22401	22427	22453	22479	22505	ļ
1	168	22531	22557	22583	22608	22634	22660	22686	22712	22737	22763	
-	169	22789	22814	22840	22866	22891	22917	22943	22968	22994	23019	Į
- }	170	23045 23300	23070 23325	23096 23350	23121 23376	23147 23401	23172 23426	23198 23452	23223 23477	23249 23502	23274 23528	i
- 1	171 172	23553	23578	23603	23629	23654	23679	23704	23729	23754	23779	}
	173	23805	23830	23855	23880	23905	23930	23955	23980	24005	24030	l
ŀ	174	24055	24080	24105	24130	24155	24180	24204	24229	24254	24279	l
-	175	24304	24329	24353	24378	24403	24428	24452	24477	24502	24527	
]	176	24551	24576	24601	24625	24650	24674	24699	24724	24748	24773 25018	}
	177	24797 25042	24822 25066	24846 25091	24871 25115	24895 25139	24920 25164	24944 25188	24969 25212	24993 25237	25261	ŀ
- 1	178 179	25285	25310	25334	25358	25382	25406	25431	25455	25479	25503	1
-	180	25527	25551	25575	25600	25624	25648	25672	25696	25720	25744	i
- 1	181	25768	25792	25816	25840	25864	25888	25912	25935	25959	25983	1
- 1	182	26007	26031	26055	26079	26102	26126	26150	26174	26198	26221	1
	183	26245	26269	26293	26316	26340	26364	26337	26411	26435	26458 26694	1
_	184	26482	26505	26529	26553	26576	26600	26623	26647	26670		1
	185 1 86	26717 26951	26741 26975	26764 26998	26788 27021	26811 27045	26834 27068	26858 27091	26831 27114	26905 27138	26928 27161	
	187	27184	27207	27231	27254		27300	27 323	27346	27370	27393	1
	188	27416	27439	27462	27485	27508	27531	27554	27577	27600	27623	ſ
	189	27646	27669	27692	27715	27738	27761	27784	27807	27830	27852]
Γ	190	27875	27898	27921	27944	27967	27989	28012	28035	28058	28081	1
- 1	191	28103	28126	28149	28171	28194 28421	28217	28240 28466	28262 28488	28285	28307 28533	l
- 1	1 92 193	28330 28556		28375 •2 8601	28398 28623	28646	28443 28668	28691	28713	28511 28735	28758	1
	194	28780		28825	28847	28870	28892	28914	28937	28959	28981	ı
-	195	29003	29026	29048	29070	29092	29115	29137	29159	29181	29203	1 '
	196	29226	29248	29270	29292	29314	29336	29358	29380	29403	29425	ł
- 1	197	29447	29469	29491	29513	29535	29557	29579	29601	29623	29645	l I
	198 1 9 9	29667 29885	29688 29907	29710 29929	29732 29951	29754 29973	29776 29994	29798 30016	29820 30038	29842 30060	29863 30081	1
L	200	30103		30146	30168	30190	30211	30233	30255	30276	30298	1
1	201	30390	30341	30363	30384	30406	30428	30233	30235	30492	30514	1
	202	30535	30557	30578	30600	30621	30643	30664	30685	30707	30728	
- 1	203	30750	30771	30792	30814	30835	30856	30878	30899	30920	30942	
L	204 · ·	30963	30984	31006	31027	31048	31069	31091	31112	31133	31154	
	,205 205	31175	31197	31218	31239	31260	31281	31302	31323	31345	31366	
	2 06 207	31387 31597	31408 31618	31429 31639	31450 31660		31492 31702		31534 31744	31555 31765	31576 31785	
	208	31806	31827	31848	31869	31890	31702	31931	31952	31705	31793	
	209	32015	32035	32056	32077	32098	32118	32139	32160	32181	32201	}
	210	32222	32243	32263	32284	32305	32325	32346	32366	32387	32408	1
	211	32428	32449	32469	32490	32510	32531	32552	32572	32593	32613	1
- 1	212 213	32634 32838	32654 32858	32675	32695 32899		32736	32756	32777	32797	32818	
- 1	214	33041	33062	32879 33082	33102	32919 33122	32940 33143	32960 33163	32980 33183	33001 33203	33021 33224	li
-	215	33244	33264	33284	33304	33325	33345	33365	33385	33405	33495	
	216	33445	33465	33486	33506	33526	33546	33566	33586	33606	33626	
i	217	33646	3 366 6	33686	33706	33726	33746	33766	33786	33806	3 3 8 2 6	
- 1	218 219	33846 34044	33866	33885	33905	33925	33945	33965	33985	34005	34025	
-	No.		34064	34034	34104	34124	34143	34163	34183	34203	34223	}
	110.	0 !	1	2 1	_ 3 _	4 1	5	6	7	8 1	9	L., Ì

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TABLE XXVI.

	No. 2	2200-	2800					Log.	34242	4	47 16.
-	No.	0	1	2	3	4	5	6	7	8	9
- (220	34242	34262	34282	34301	34321	34341	34361	34380	34400	34420
1	221 222	34439	34459		34498	34518	34537	34557	34577	34596	34616
	222	34635 34830	34655 34850	34674 34869	34694 34889	34713 34908	34733 34928	34753 34947	34772 34967	34792 34986	34811 35005
- 1	224	35025	35044	35064	35083	35102	35122	35141	35160	35180	35199
	225	35218	35238	85257	35276	35295	35315	35334	35353	35372	35392
	226	35411	35430	35449	35468	35488	35507	35526	35545	35564	35583
1	227	35603	35622	35641	35660	35679	3 5698	35717	35736	35755	35774
	228 229	35793 35984	35813 36003	35832 36021	35851 36040	35870 36059	35989 36078	35908 36097	35927 36116	35946 36135	35965 36154
	230	36173	36192	36211	36229	36248	36267	36286	36305	36324	36342
	231	36361	36380	36399	36418	36436	36455	36474	36493	36511	36530
i	232	36549	36568	36586	36605	36624	36642	36661	36680	36698	36717
	233	36736	36754	36773	36791	36810		36847	36866	36884	36903
	234	36922	36940	36959	36977	36996		37033	37051	37070	37088
	235 236	37107 37291	37125 37310	37144 37328	37162	37181	37199 37383	37218 37401	37236 37420	37254	37273 37457
	237	37475	37493	37511	37346 37530	37365 37548	37566	37585	37603	37438 37621	37639
Į	238	37658	37676	37694	37712	37731	37749	37767	37785	37803	37822
	239	37840	37858	37876	37894	37912	37931	37949	37967	37985	38003
1	240	38021	38039	38057	38075	38093	38112	38130	38148	38166	38134
ŀ	241 242	38202	38220	38238	38256	38274	38292	38310	38328	38346	38364
- 1	243	38382 38561	38399 38578	38417 38596	38435 38614	38453 38632	38471 38650	38489 38668	38507 38686	38525 38703	38543 38721
- 1	244	38739	38757	38775	38792	38810	38828	38846	38863	38881	38899
	245	38917	38934	38952	38970	38987	39005	39023	39041	39058	39076
1	246	39094	39111	39129	39146	39164	39182	39199	39217	39235	39252
- 1	247	39270	39287	39305	39322	39340	39358	39375	39393	39410	39428
ı	24 8 24 9	39445 39620	39463 39637	39480 39655	39498 39672	39515 39690	39533 39707	39550 39724	39568 39742	39585 39759	39602 39777
ł	250	39794	39811	39829	39846	39863	39881	39898	39915	39933	39950
- 1	251	39967	39985	40002	40019	40037	40054	40071	40088	40106	40123
- 1	252	40140	40157	40175	40192	40209	40226	40243	40261	40278	40295
1	253	40312	40329	40346	40364	40381	40398	40415	40432	40449	40466
	254	40483	40500	40518	40535	40552	40569	40586	40603	40620	40637
- 1	255 256	40654 40824	40671 40841	40688 40858	40705 40875	40722 40892	40739 40909	40756 40926	40773 40943	40790 40960	40807 40976
	257	40993	41010	41027	41044	41061	41078	41095	41111	41128	41145
- 1	258	41162	41179	41196	41212	41229	41246	41263	41280	41296	41313
1	259	41330	41347	41363	41380	41397	41414	41430	41447	41464	41481
- 1	260	41497	41514	41531	41547	41564	41581	41597	41614	41631	41647
- 1	261 262	41664 418 3 0	41681 41847	41697 41863	41714 41880	41731 41896	41747 41913	41764 41929	41780 41946	41797 41963	41814 41979
	263	41996	42012	42029	42045	42062	42078	42095	42111	42127	42144
	264	42160	42177	42193	42210	42226	42243	42259	42275	42292	42308
- 1	265	42325	42341	42357	42374	42390	42406	42423	42439	42455	42472
-	266	42488	42504	42521	42537	42553	42570	42586	42602	42619	42635
- 1	267 268	42651 42813	42667 42830	42684 42846	42700 42862	42716 42878	42732 42894	42749 42911	42765 42927	42781 42943	42797 42959
1	269	42975	42991	43008	43024	43040	43056	43072	43088	43104	43120
Ì	270	43136	43152	43169	43185	43201	43217	43233	43249	43265	43281
ł	271	43297	43313	43329	43345	43361	43377	43393	43409	43425	43441
1	272	48457	43473	43489	43505	43521	43537	43553	43569	43584	43600
١	273 274	43616 43775	43632 43791	43648 43807	43664 43823	43680 43838	43696 43854	43712 43870	43727 43886	43743 43902	43759 43917
	275	43933	43949	43965	43981	43996	44012	44028	44044	44059	44075
- 1	276	44091	44107	44122	44138	44154	44170	44185	44201	44217	44232
- 1	277	44248	44264	44279	44295	44311	44326	44342	44358	44373	44389
	278	44404	44420	44436	44451	44467	44483	44498	44514	44529	44545
	279	44560	44576	44592	44607	44623	44638	44654	44669	44685	44700
	No.	0	1	2	3	4	5 1	Digifized	by (7 ₁ (<u>१०श्री</u>	9 .

LOGARITHMS OF NUMBERS.

No. 2800 3409. Log. 44716 53148.												
No.	0	3400	2	3	4	1 5	6	og. 4411	8		÷	
280	44716	44731	44747	44762	·	44793		44824	44840	9	4	
281	44871	44886	44902	44917		44948		44979	44994	44855 45010	;	
232	45025	45040	45056	45071		45102	45117	45133	45148		1	
283	45179	45194	45209	45225	45240	45255	45271	45286	45301	45317		
284	45332	45347	45362	45378	45393	45408	45423	45439	45454	45469	1	
285	45484	45500	45515	45530		45561	45576	45591	45606	45621	1	
286	45637		45667			45712			45758	45773	1	
287	45788			45834					45909	45924	1	
288 239	45939 46090		45969 46120	45984 46135				46045 46195	46060		1	
290	46240		46270						46210		1	
290	46339			46285 46434				46345 46494	46359 46509	46374 46 523		
292	46538									46672	1	
293	46687		46716		46746		46776			46820		
294	46335		46364	46879			46923	46938	46953	46967		
295	46982	46997	47012	47026	47041	47056	47070	47085	47100	47114	1	
296	47129	47144	47159	47173	47188	47202	47217	47232	47246	47261	1	
297	47276		47305	47319	47334	47349	47363			47407		
298		47436		47465	47480	47494		47524			}	
299		47532	47 596	47611	47625	47640	47654	47669			j	
300	47712		47741	47750	47770	47784	47799	47813	47828		1	
301 302	47857 43001		47855	47900		47929	47943				1	
302		48015 48159	48029 48173	48044 48187	43058 48202		48087 48230	48101 48244	48116 48259	48130 48273	1	
304		43302	48316	48330	48344	48359	48373	48387	48401	48416		
305	48430		48458	48473		48501	48515	48530	48544	48558	-	
306	48572	48586	48601	48615		48643	48657	48671	48686	48700	1	
397	1	48728	48742	48756			48799	48813	48827	48841)	
:`08	48855	48869	48883	48897	48911	48926	48940	48954	48968	48982	1 .	
309	48996	49010	49024	49038	49052	49066	49080	49094	49108	49122	1 1	
310	49136	49150	49164	49178	49192	49206	49220	49234	49248	49262	1	
311		49290	49304	49318	49332	49346		49374	49388	49402	1 1	
312		49429	49443	49457	49471	49485	49499	49513	49527	49541		
313 314	49.554	49568 49707	49582 49721	49596 49734	49610 49748	49624	49638	49651	49665	49679	1 1	
						49762	49776	49790	49803	49817	J I	
315	49931 49969	49345 49982	49859 49996	49872 50010	49886 50024	49900 50037	49914	49927	49941	49955		
317	50106	50120	50133	50147	50161	50174	50051 50188	50065 50202	50079 50215	5009% 50229	1	
318	50243	50256	50270	50284	50297	50311	50325	50338	50352	50365	1	
319	50379	50393	50406	50420	50433	50447	50461	50474	50488	50501		
320	50515	5 0529	50542	50556	50569	50583	50596	50610	50623	50637	1	
321	50651	50664	50678	50691	50705	50718	50732	50745	50759	50772		
322	50786	50799	50813	50826	50840	50853	50866	50880	50893	50907		
323	50920	50934	50947	50961	50974	50987	51001	51014	5102 8	51041		
324	51055	51068	51081	51095	51108	51121	51135	51148	51162	51175		
325	51188	51202	51215	51228	51242	61255	51268	51282	51295	51308	l	
326	51322	- 4	51348	51362	51375	51388	51402	51415	51428	51441		
327 328	51435		51481	51495	51508	51521	51534	51548	51561	51.574		
328	51587 51720	51601 51733	51614 51746	51627 51759	51640 51772	51654 51786	51667 51799	51680	51693	51706		
330	51351							51812	51825	51838		
331	51983	51865 51996	51878 52009	51891 52022	51904 52035	51917 52048	51930	51943	51957	51970	1	
332	52114	52127	52140	52153	52166	52179	52061 52192	52075 52205	52088 52218	52101 52231		
333	52244	52257	52270	52284	52297	52310	52323	52336	52349	52362		
334	52375	52383	52401	52414	52427	52440	52453	52466	52479	52492		
335	52504	52517	52530	52543	52556	52569	52582	52595	52608	52621		
336	52634	52647	52660	52673	52686	52699	52711	52724	52737	52750		
337	527 63	52776	52789	52802	52815	52827	52840	52853	52866	52879		
338	52892		52917	52930	52943	52956	52969	52982	52994	53007	!	
339	53020		53046	53058	53071	53084	53097	53110	53122	53135		
No.	0 1	1	2	3	4	5	6	7	8	- 9	L	

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TABLE XXVI.

No. 0 1 2 3 4 5 6 7 7 8 9 340 5314 5316 53173 5316 53194 53276 53389 53314 53326 53333 53332 53336 53377 53390 3342 53320 53415 53428 53441 53453 53466 53479 53491 53594 53513 53643 3345 53526 53565 53567 53580 53593 53505 53561 53631 53643 3344 53566 53661 53651 53661 53631 53643 3344 53566 53661 53651 53661 53631 53673 53795 53746 53777 53769 3446 53906 53920 53932 53936 53937 53936 53937 53936 53936 53936 53937 53936 53936 53936 53937 53936 53936 53936 53937 53936 5393	No. 3	100	—4000					Lo	g. 5314	8	-60206.
341 53275 53388 53301 53314 53326 53336 53346 53477 53390 3343 53495 53446 53473 53491 53454 53454 53454 53454 53454 53454 53455 3456 34579 53491 53504 53516 33474 53451 53455 53567 53580 53593 53605 53618 53631 53643 344 53566 53593 53794 53777 53759 33643 344 53566 53593 53794 53797 53757 53759 346 53990 53990 53930 53933 53945 53936 53937 53870 53877 53759 346 53990 53990 53933 53945 53950 53982 53995 54008 54020 5415 5415 5415 5415 5415 5415 5415 541					3	. 4	5				
341 53275 53388 53301 53314 53326 53336 53346 53477 53390 3343 53495 53446 53473 53491 53454 53454 53454 53454 53454 53454 53455 3456 34579 53491 53504 53516 33474 53451 53455 53567 53580 53593 53605 53618 53631 53643 344 53566 53593 53794 53777 53759 33643 344 53566 53593 53794 53797 53757 53759 346 53990 53990 53930 53933 53945 53936 53937 53870 53877 53759 346 53990 53990 53933 53945 53950 53982 53995 54008 54020 5415 5415 5415 5415 5415 5415 5415 541			53161	53173	53186	53199	53212	53224	53237	53250	
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388 58883 58894 58906 58917 58928 58939 58950 58961 58973 58984 389 58995 59006 59017 59028 59040 59051 59062 5973 59084 59095 390 59168 59129 59140 59151 59162 59173 59184 59195 59207 391 59218 59229 59240 59251 59262 59273 59284 59295 59306 59318 392 59329 59400 59311 59362 59373 59384 59395 59406 59417 59428 393 59439 59450 59461 59472 59483 59494 59506 59517 59528 59539 394 5950 5961 59671 59682 59633 59704 59715 59726 59727 59788 59759 395 5960 59671 59682 59802 59813 5								5 8 838	58850		58872
390 59106 59118 59129 59140 59151 59162 59173 59184 59195 59207 391 59218 59229 59240 59251 59262 59273 59284 59295 59306 59318 392 59329 59340 59351 59362 59373 59384 59395 59406 59417 59428 393 59439 59450 59451 59572 59583 59594 59606 59517 59528 59539 394 59560 59561 59572 59583 59594 59616 59627 59628 59649 395 59660 59671 59682 59693 59716 59715 59726 59737 59748 59759 396 59770 59780 59791 59802 59813 59824 59835 59846 59857 59868 397 5988 59990 60010 60021 60032 60043 <t< th=""><th></th><th>58883</th><th>58894</th><th>58906</th><th></th><th></th><th></th><th></th><th></th><th></th><th></th></t<>		58883	58894	58906							
391 59218 59229 59240 59251 59262 59273 59284 59295 59306 59318 392 59329 59340 59351 59362 59373 59384 59395 59406 59417 59428 393 59450 59450 59461 59472 59483 59494 59506 59517 59528 59539 394 59560 59561 59572 59583 59594 59606 59616 59627 59638 59693 395 59660 59671 59682 59693 59715 59726 59737 59748 59759 396 59770 59780 59791 59802 59813 59824 59835 59846 59877 59868 397 5988 59990 60010 60021 60032 60043 60044 60065 60076 60086 399 60097 60108 60119 60130 60141 60152 <t< th=""><th></th><th>58995</th><th>59006</th><th></th><th></th><th></th><th></th><th></th><th></th><th></th><th></th></t<>		58995	59006								
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394 59560 59561 59572 59583 59594 59605 59616 59627 59638 59649 395 59660 59671 59682 59693 59704 59715 59726 59737 59748 59759 396 59770 59780 59791 59802 59813 59824 59835 59846 59857 59868 397 59879 59890 59901 59912 59923 59934 59945 59956 59966 59977 398 59988 59999 60010 60021 60032 60043 60054 60065 60076 60036 399 60097 60108 60119 60130 60141 60152 60163 60173 60184 60195											
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396 59770 59780 59791 59802 59813 59824 59835 59846 59857 59868 397 59879 59890 59901 59912 59923 59934 59945 59956 59966 59977 398 59988 59999 60010 60021 60032 60043 60054 60054 60065 60076 60086 399 60097 60108 60119 60130 60141 60152 60163 60173 60184 60195											59759
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000 0000											
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LOGARITHMS OF NUMBERS.

. 4000-	46	00.						Log.	60206-	662	76.
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404	60638	60649	60660	60670	60681	60692			60831	60842	ł
405	60746	60756	60767	60778	60788 60895	60799 60906	60810 60917	60821 60927	60938	60949	i
406 407	60853	60863 60970	60874 60981	60885 60991	61002	61013	61023	61034	61045	61055	1
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410	61278	61289	61300	61310	61321	61331	61342	61352	61363	61374	7
411	61384	61395	61405	61416	61426	61437	61448	61458	61469	61479	i
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413	61595	61606	61616	61627 61731	61637 61742	61648 6175 2	61658 61763	61669 617 73	61679 61784	61 690 61 794	i i
414	61700	61711	61721								-
415	61805	61815	61826 61930	61836 61941	61847 61951	61857 6196 2	61868 61972	61878 619 82	61888 61993	61899 62003	1
416 417	62014	61920 62024	62034	62045	62055	62066	62076	62086	62097	62107	1
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421	62428	62439	62449	62459	62469	62480	62490	62500		62521	1
422	62531	62542	62552	62562	62572	62583	62593	62603		62624 62726	l
423	62634	62644	62655	62665	62675 62778	62685 62788	62696 62798	62706 62808	62716 62818	628 29	1
424	62737	62747	62757	62767	62880	62890	62900	62910	62921	62931	1
425 426	62839 62941	62849 62951	62859 62961	62870 62972	62982	6299 2	63002	63012	63022	63033	1
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429	63246	63256	63266	63276	632 86	63296	63306	63317	63327	63337	J
430	63347	63357	63367	63377	63387	63397	63407	63417	63428	63438	1
431	63448	63458	63468	63478	63488	63498	63508	63518	63528	63538	1
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435	63849	63859	63869	63879	63839	63899	63909	63919	63929	63939	†
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437	64048	64058	64068	64078	64088	64098	64108	64118	64128	64137	
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443	64738	64748	64758	64768	64777	64787	64797	64807	64816	64826	1
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451 452	65418 65514	65427 65523	65437 65533	65447 65543	65456 65552	65466 65562	65475 65571	65485 65581	65496 65591	65600	ł
452 453	65610	65619	65629	65639	65648	65658	65667	65677	65686	65696	lΙ
454	65706	65715	65725	65734	65744	65753	65763	65772	65782	65792	
455	65801	65811	65820	65830	65839	65849	65858	65868	65877	65887	[]
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459 No.	66181	66191	66200	66210	66219			7			
No.	1 0	1 1	2	3 1	4	5	6 1	7	8	<u></u>	$\sigma l d$

i	No. 4600 5200. Log. 66276 71600												
			5200						66276-		1600.		
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	460	66276	66285		66304	66314	66323	66332	66342		66361		
	461 462	66370 66464	66380 66474	66389 6648 3	66398 66492	66408 66502	66417 66511	66427 66521	66436 665 3 0	66445 66539	66455		
•	463	66558	66567	66577	66586	66596	66605	66614	66624	66633	66549 66642		
	464	66652	66661	66671	66680	66689	66699	66708	66717	66727	66736		
	465	66745	66755	66764	66773	66783	66792	66801	66811	66820	66829		
	466	66839	66848	66857	66867	66876	66885	66894	66904	66913	66922		
ŀ	467	66932	66941	66950	66960	66969	66978	66987	66997	67006	67015		
'	468 469	67025 67117	67034	67043	67052	67062	67071	67080	67089	67099	67108		
			67127	67136	67145	67154	67164	67173	67182	67191	67201		
1	470 471	67210 67302	67219 67311	67228 67321	67237 67330	67247 67339	67256 67348	67265 67357	67274 67367	67284 67376	67293		
1	472	67394	67403	67413	67422	67431	67440	67449	67459	67468	67385 67477		
	473	67486	67495	67504	67514	67523	67532	67541	67550	67560	67569		
[474	67578	67587	67596	67605	67614	67624	67633	67642	67651	67660		
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	476	67761	67770	67779	67788	67797	67806	67815	67825	67834	67843		
	477	67852	67861	67870	67879	67888	67897	67906	67916	67925	67934		
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										68106	68115		
	480 481	68124 68215	68133 682 24	68142 68233	68151 68 242	68160 68251	68169 68260	68178 68269	68187 68 2 78	68196 68287	68205		
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	483	68395	68404	68413	68422	68431	68440	68449	68458	68467	68476		
	484	68485	68494	68502	68511	68520	68529	68538	68547	68556	68565		
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	488 489	6884 2 68931	68851 68940	68860 68949	68869 68958	68878 68966	68886 68975	68895 68984	68904 68 993	68913	68922		
	490	69020	69028	69037	69046	69055	69064			69002	69011		
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	493	69285	69294	69302	69311	69320	69329	69338	69346	69355	69364		
	494	69373	69381	69390	69399	69408	69417	69425	69434	69443	69452		
	495	69461	69469	69478	69487	69496	69504	69513	69522	69531	695 3 9		
	496	69548	69557	69566	69574	69583	69592	69601	69609	69618	69627		
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	499	69810	69819	69827	69836	69845	69854	69862	69871	69880			
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	502	70070	70079	70088	70096	70105	70114	70122	70131	70140	70143		
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	505	70329	70338	70346	70355	70364	70372	70381	70389	70398	70406		
	506 507	70415 70501	70424 70509	70432 70518	70441 70526	70449 70535	70458 70544	70467 70552	70475 70561	70484 70569	70492 70578		
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	511	70842	70851	70859	70868	70876	7088 5	70 893	70902	70910	70919		
İ	512	70927	70935	70944	70952	70961	70969	70978	70986	70995	71003		
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- 1	515 516	71181 71 2 65	71189 71273	71198 71282	71206 71290	71214 71299	71223 71307	71231 71315	71240 71324	712 td 71332	71257 71341		
ı	517	71349	71357	71366	71374	71383	71391	71399	71408	71416	71425		
	518	71433	71441	71450	71458	71466	71475	71483	71492	71500	71508		
,	519	71517	71525	71533	71542	71550	71559	71567	71575	71584	71592		
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LOGARITHMS OF NUMBERS.

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522	71767	71775	71784	71792	71800	71809	71817	71825	71834	71842 71925	l
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524	71933	71941	71950	71958	71966					72090	ł
525	72016	72024	72032	72041	72049 72 132	72057 72140	72066 72148	72074 72156	72082 72165	72173	1
526 527	72099 72181	72107 72189	72115 72198	72123 72206	72132	72222	72230	72239	72247	72255	i
528	72263	72272	72280	72288	72296	72304	72313	72321	72329	72337	
5 2 9	72346	72354	72362	72370	72378	72387	72395	72403	72411	72419	ĺ
530	72428	72436	72444	72452	72460	72469	72477	72485	72493	72501	l
531	72509	72518	72526	72534	72542	72550	72558	72567	72575	72583	ł
532	72591	72599	72607	72616	72624	72632	72640	72648	72656	72665	İ
533	72673	72681	72689	72697	72705	72713	72722	72730 72811	72738 72819	72746 72827	
534	72754	72762	72770	72779	72787	72795	72803		!-	72908	
535	72835	72843	72852	72860	72868	72876	72884	72892 72973	72900 72981	72908	
536	72916	72925	72933	72941 73022	72949 73030	72957 73038	72965 73046	73054	73062	73070	
537 538	72997 73078	73006 73086	73014 73094	73102	73111	73119	73127	73135	73143	73151	
539	73159	73167	73175	73183	73191	73199	73207	73215	73223	73 2 31	ŀ
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543	73480	73488	73496	73504	73512	73520	73528	73536	73544		
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548 549	73957	73965	73973	73981	73989	73997	74005	74013	74020	74028	
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551	74115	74123	74131	74139	74147	74155	74162	74170	74178	74186	
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561	74896	74904	74912	74920	74927	74935	74943	74950	74958	74966	1 1
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566	75282	75289	75297	75305	75312	75320	75328 75404	75335 75412	75343 75420	75351 75427	1 1
567 568	75358 75 43 5	75366 75442	75374 75450	75381 75458	75389 75465	75397 75473	75481	75488	75496	75504	l i
<i>5</i> 69	75511	75519	75526	75534	75542	75549	75557	75565	75572	75580	1 :
570	75587	75595	75603	75610	75618	75626	75633	75641	75648	75656	1 :
571	75664	75671	75679	75686	75694	75702	75709	75717	75724	75732	1
572	75740	75747	75755	75762	75770	75778	75785	75793	75800	75808	1
573	75815	75823	75831	75838	75846	75853	75861	75868	75876	75884	
574	75891	75899	75906	75914	75921	75929	75937	75944	75952	75959	
575	75967	75974	75982	75989	75997	76005	76012	76020	76027	76035 76110	
576	76042	76050	76057	76065	76072 76148	76080 76155	76087 76163	76095 76170	7610 3 76178	76185	1
57 7 5 78	76118 7619 3	76125 76200	76133 76208	76140 76215	76223	76230	76238	76245	76253]
579	76268	76275	76283	76290	76298	76505	76313	76320	76328	76335	
No.	0	1	2	3	4	5	6	7	8	9	l l
	<u> </u>										 '

LOGARITHMS OF NUMBERS.

	LOGARITHMS OF NUMBERS.													
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	581	76418	76425	76433	76440	76448	76455	76462	76470	76477	76435	1		
	582.	76492	76500	76507	76515	76522	76530	76537	76545	76552 76626	76559 76634	ł		
	583 584	76567 76641	76574 76649	7658 2 76656	76589 76664	76597 76671	76604 76678	76612 76686	76619 76693	76701	76708	1		
			76723	76730	76738	76745	76753	76760	76768	76775	76782	1		
	585 586	76716 76790	76797	76805	76812	76819	76827	76334	76842	76849	76856			
1	58 7	76864	76871	76879	76886	76893	76901	76908	76916	76923	76930			
	588	76939	76945	76953	76960	76967	76975	76982	76989	76997	77004	1		
	589	77012	77019	77026	77934	77041	77048	77056	77063	770 70	77078			
	590	77085	77093	77100	77107	77115	77122	77129	77137	77144	77151			
	591	77159	77166	77173	77181	77188	77195	77203	77210	77217 77291	77225 77298	1		
1	592	77232	77240	77247	77254 77327	77262 77335	77 2 69 77 342	77276 77349	77283 77357	77364	77371	1		
1	593 594	77305 77379	7731 3 77386	77320 77393	77401	77408	77415	77422	77430	77437	77444			
- 1	595	77452	77459	77466	77474	77481	77488	77495	77503	77510	77517	1		
	596	77525	77532	77539	77546	77554	77561	77568	77576	77583	77590	ł		
i	597	77597	77605	77612	77619	77627	77634	77641	77648	77656	77663			
- 1	598	77670	77677	77685	77692	77699	77706	77714	77721	77728				
	599	77743	77750	77757	77764	77772	77779	77786	77793	77801	77808			
	600	77815	77822	77830	77837	77844	77851	77859	77866	77873	77680	1		
	601	77837	77895	77902	77909	77916	77924 77996	77931 78003	77938 78010	77945 78017	77952 78025	1		
	60 2 603	77960 78032	77967 78039	77974 78046	77981 78053	77988 78061	78068	78075	78082	78089	78097	1		
	604	78104	78111	78118	78125	78132	78140	78147	78154	78161	78168			
	605	78176	78183	78190	78197	78204	78211	78219	78226	78233	78240	1		
	606	78247	78254	78262	78269	78276	78283	78290	78297	78305	78312			
	607	78319	78326	78333	78340	78347	78355	78362	78369	78376	78383	l		
	608	78390	78398	78405	78412	78419	78426	78433	78440	78447	78455			
	609	78462	78469	78476	78483	78490	78497	78504	78512	78519	78526	-		
	610	78533	78540	78547	78554	78561	78569	78576	78583	78590	78697 78668	1		
	611	78604	78611	78618	78625	78633 78704	78640 78711	78647 78718	78654 787 2 5	78661 78732	787 3 9	1		
	61 2 613	78575 78746	78682 78753	78689 78760	78696 78767	78774	78781	78789	78796	78803	78810			
	614	78817	78824	78831	78838	78845	78852	78859	78866	78873	78880	1		
	615	78883	78895	78902	78909	78916	78923	78930	78937	78944	78951	1		
	616	78958	78965	78972	78979	78986	78993	79000	79007	79014	79021			
	617	79029	79036	79043	79050	79057	79064	79071	79078	79085	79092	ļ		
	618	79099	79106	79113	79120	79127	79134	79141 79211	79148 79218	79155 79225	791 62 79232	ł		
	619	79169	79176	79183	79190	79197	79204				79302	ł		
	620	79239	79246	79253	79260	79 2 67 79337	79274	79281 79351	79288 79358	79 2 95 79365	79372	l		
- 1	621 622	79309 79379	79316 79386	79323 79393	79330 79400	79407	79414	79421	79428	79435	79442	i		
- 1	623	79449	79456	79463	79470	79477	79484	79491	79498	79505	79511	1		
- 1	624	79518	79525	79532	79539	79546	79553	79560	79567	79574	79581]		
- 1	625	79588	79595	79602	79609	79616	79623	79630	79637	79644	79650	l		
- 1	626	79657	79664	79671	79678	79685	79692	79699	79706	79713	79720	1		
ı	627	79727	79734	79741	79748	79754	79761	79768	79775	7978 2 79851	79789 79858	1		
	628	79796	79803	79810 79879	79817 79886	798 24 79893	79831 79900	79837 79906	79844 79913	79920	79927	1		
- 1	629	79865	79872				79969	79975	79982	79989	79996	1		
- 1	630	79934 80003	79941 80010	79948 80017	79 96 5 800 24	79962 80030	80037	80044	80051	80058	80065			
1	631 632	80003	80079	80085	80092	.80099	80106	80113	80120	80127	80134	1		
]	633	80140	80147	80154	80161	80168	80175	80182	80188	80195	80202	l		
1	634	80209	80216	80223	80229	80236	80243	80250	80257	80264	80271	4		
ľ	636	80277	80284	80291	80298	80305	80312	80318	80325	80332 80400	80339 80407			
ı	636	80346	80353	80359	80366	80373	80380	80387	80393 80462	80468	80407	1		
- 1	637	80414	80421	80428	80434 80502	80441 80509	80448 80516	80455 80523	80530	80536	80543	1		
Į	638 639	80482 80550	80489 80557	80496 80564	80570	80577	80584	80591	80598	80604	80611	1		
L		<u> </u>		2	3	4	5	-6	7	8	9	1		
	No.	1 0										-		

LOGARITHMS OF NUMBERS.

	No. 6400——7000. Log. 80618——84510.												
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	640	80618	80625	80632	80638	80645	80652	80659	80665	80672		1	
	641	80686	80693	80699	80706	80713	80720	80726	80733	80740	80747		
	642	80754 80821	80760 80828	80767 80835	80774 80841	80781 80848	80787 80855	80794 80862	80801	80808	80814	1	
	644	80889	80895	80902	80909	80916	80922	80929	80868 80936	80875 80943	8088 2 80949	1	
	645	80956	80963	80969	80976	80983	80990	80996	81003	81010	81017	1	
	646	81023	81030	81037	81043	81050	81057	81064	81070	81077	81084	ł	
	647 648	81090 81158	81097 81164	81104 81171	81111 81178	81117	81124		81137	81144	81151	l	
	649	81224	81231	81238	81245	81184 81251	81191 81258	81198 81 2 65	81204 81271	31211 81278	81218 81285		
	650	81291	81298	81305	81311	81318	81325	81331	81338	81345	81351	ł	
	651	81358		81371	81378	81385	81391	81398	81405	81411	81418	1	
	652	81425	81431	81438	81445	81451	81458	81465	81471	81478	81485	ı	
	658 654	81491 81558	81498 81 <i>5</i> 64	81505 81571	81511 81578	81518	81525	81531	81538	81544	81551	i	
	655	81624	81631	81637	81644	81584 81651	81591	81598	81604	81611	81617	ļ	
	656	81690		81704	81710	81717	81657 81723	81664 81730	81671 81737	81677 81743	81684 817 50	١.	
	657	81757	81763	81770	81776	81783	81790		81803	81809	81816	1	
	658	81823		81836	81842	81849	81856	81862	81869	81875	81882]	
	659	81889	81895	81902	81908	81915	81921	81928	81935	81941	81948		
	660	81954	81961	81968	31974	81981	81987	81994	82000	82007	82014	1	
	661 662	82020 82086	82027 82092	82033 82099	82040 82105	82046 82112	82053 82119	82060 82125	82066	82073	82079		
	663	82151	82158	82164	82171	82178	82184	82123	82132 82197	82138 82 2 04	82145 82210	ł	
	664	82217	82223	82230	82236	82243	82249	82256	82263	82269	82276	1	
	665	82282	82289	82295	82302	82308	82315	82321	82328	82334	82341	1	
	666	82347	82354	82360	82367	82373	82380	82387	82393	82400	82406		
	667	82413 82478	82419 82484	82426	82432	82439	82445	82452	8 2 458	82465	82471		
	669	82543	82549	82491 82556	82497 82562	82504 82569	82510 8257 <i>5</i>	82517 82582	82523 82588	8 2 530 8 2 595	82536 82601		
	670	82607	82614	82620	82627	82633	82640	82646	82653	82659		1	
	671	82672	82679	82685	82692	82698	82705	82711	82718	82724	8 2666 827 3 0		
	672	82737	82743	82750	82756	82763	82769	82776	82782	82789	82795		
	673	82802 82866	82808 82872	82814	82821	82827	82334	82840	82847	82853	82860	1	
	675	82930		82879	82885	82892	82898		82911	82918	82924	1	
	676	82930	82937 83001	82943 83008	82950 83014	82956 83020	82963	82969	82975	82982	82988	l I	
	677	83059	83065	83072	83078	83085	83027 83091	83033 83097	83040 83104	83046	83052 83117	1	
	678	83123	83129	83136	83142	83149	83155	83161	83168	83110 83174	83181		
	679	83187	83193	83200	83206	83213	83219	83225	83232	83238	83245	1	
	680	83251	83257	83264	83270	83276	83283	83289	83296	83302	83308		
	681 682	83315 83378	83321 83385	83327	83334	83340	83347	83353	83359	83366	83372		
	683	83442		83391 83455	83398 83461	83404 83467	83410		83423	83429	83436		
	684	85506	83512	83518	83525	83531	83474 83537	83480 83544	83487 83550	83493 83556	83499 83563		
	685	83569	83575	83582	83588	83594	83601	83607	83613	83620	83626	1	
	686	83632	83639	83645	83651	83658	83664	83670	83677	83683	83689		
	687 688	83696	83702	83708	83715	83721	83727	83734	83740	83746	83753		
	689	83759 83822	83765 83828	83771 83835	83778 83841	83784 83847	83790 83853	83797	83803	83809	83816]	
	690	83885	83891	83897	83904	83910	83916	83860	83866	83872	83879		
	691	83948	83954	83960	83967	83973	83916	83923 8398 <i>5</i>	83929 83992	83935 83998	83942 84004		
	692	84011	84017	84023	84029	84036	84042	84048	84055	84061	84067		
	693	84073	84080	84086	84092	84098	84105	84111	84117	84123	84130		
	694	84136	84142	84148	84155	84161	84167	84173	84180	84186	84192		
	695 696	84198 84261	84205 84267	84211 84273	84217	84223	84230	84236	84242	84248	84255	ļ	
	697	84323	84330	84336	84280 84342	84286 84348	84292 84354	84298	84305	-84311	84317	1	
l	698	84386	84392	84398	84404	84410	84417	84361 84423	84367 84429	84373 84435	84379 84442		
ļ	699	84448	84454	84460	84466	84473	84479	84485	84491	84497	84504	i	
	No.	0	1	2	3	4	5	6	7	8	9		

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	No. 7000		7600.	·		· •			Log. 84	510	8 80 81.	
	No.	0	1 1	2 1	3	4	5	6	7	8 1	9	_
1	700	84510	84516	84522	84528	84535	84541	84547	84553	84559	84566	1
	701	84572	84578	84584	84590	84597	84603	84609	84615	84621	84628	1
	702 703	84634 84696	84640	84646	84652	34658	84665	84671	84677	84683	84689	l
İ	704	84757	84702 84763	84708 84770	84714 84776	84720 84782	84726 84788	84733 84794	84739 84800	84745 84807	84751 84813	
	705	84819	84825	84831	84837	84844	84850	84856	84862	84868	84874	-
l	706	84880	84887	84893	84899	84905	84911	84917	84924	84930	84936	1
	707	84942	84948	84954	84960	84967	84973	84979	84985	84991	84997	1
	708 709	85003 85065	85009	85016	85022	85028	85034	85040	85046	85052	85058	1
			85071	85077	85083	85089	85095	85101	85107	85114	85120	1
	710 711	85126 85187	85132 85193	85138 85199	85144 85205	85150 85211	85156 85217	85163 85224	85169	85175	85181	1
1	712	85248	85254	85260	85266	85272	85278	85285	85230 85291	85236 85297	85242 85303	-
	713	85309	85315	85321	85327	85333	85339	85345	85352	85358	85364	
	714	85370	85376	85382	85388	85394	85400	85406	85412	85418	85425	}
	715	85431	85437	85443	85449	85455	85461	85467	85473	85479	85485	1
1 1	716	85491	85497	85503	85509	85516	85522	85528	85534	85540	85546	1
i	717 718	85552 85612	85558 85618	85564 85625	85570 85631	85576 85637	85582 85643	85588 85649	85594	85600	85606 95667	1
. 1	719	85673	85679	85685	85691	85697	85703	85709	85655 85715	85661 85721	85667 85727	I
	720	85733	85739	85745	857ŏ1	85757	85763	85769	85775	85781	85788	1
!	721	85794	85800	85806	85812	85818	85824	85830	85836	85842	85848	
ŀ	722	85854	85860	85866	85872	85878	85884	85890	85896	85902	85908	
	723	85914	85920	85926	85932	85938	85944	85950	85956	85962	85968	1
	724	85974	85980	85986	85992	85998	86004	86010	86016	86022	86028	
	725 726	86034 86094	86040 86100	86046 86106	86052 86112	86058 86118	86064 86124	86070 86130	86076	86082 86141	86088 86147	1
	727	86153	86159	86165	86171	86177	86183	86189	86136 86195	86201	86207	
	728	86213	86219	86225	86231	86237	86243	86249	86255	86261	86267	
	729	86273	86279	86285	86291	86299	86303	86308	86314	86320	863 26	
	730	86332	86338	86344	86350	86356	86362	86368	86374	86380	86386	1
	731 73 2	8639 2 86451	86398 86457	86404 86463	86410 86469	8641 <i>5</i> 8647 <i>5</i>	86421 86481	86427 86487	86433	86439	86445 86504	
	733	86510	86516	86522	86528	86534	86540	86546	86493 8655 2	86499 86558	86564	
	734	86570	86576	86581	86587	86593	86599	86605	86611	86617	86623	1
	735	86629	86635	86641	86646	86652	86658	86664	86670	86676	86682	1
	736	86688	86694	86700	86705	86711	86717	86723	86729	86735	86741	1
	737 738	86747 86806	86753	86759 86817	86764 86823	86770 868 2 9	86776 86835	86782	86788	86794	86800	ł
	739	86864	86812 86870	86876	86882	86888	86894	86841 86900	86847 86906	86853 86911	86 859 8691 7	
	740	86923	86929	86935	86941	86947	86953	86958	86964	86970	86976	1
	741	86982	86988	86994	86999	87005	87011	87017	87023	87029	87035	į
	742	87040	87046	87052	87058	87064	87070	87075	87081	87087	87093	
	743 744	87099 871 <i>5</i> 7	87105	87111	87116	87122	87128	87134 87192	87140	87146	87151 87210	1
	745	87216	87163 87221	87169 87227	87175 87233	87181 87239	87186 87245	87251	87198	87204 87262	87268	-
	746	87274	87221 87280	87227 87286	87291	87 23 9	87303	87309	872 <i>5</i> 6 87315	87320	87326	1
	747	87332	87338	87344	87349	87355	87361	87367	87373	87379	87384	1
1	748	87390	87396	87402	87408	87413	87419	87425	87431	87437	87442	i
	749	87448	87454	87460	87466	87471	87477	87483	87489	87495	87500	1
- 1	750	87506	87512	87518	87523	87529	87535	87541	87547	87552	87558	Į
- 1	751 75 2	87564 87622	87570 87628	87576 87633	87581 876 3 9	87587 876 45	87593 87651	87599 87656	87604 87662	87610 87668	87616 87674	1
	753	87679	87685	87691	87697	87703	87708	87714	87720	87726	87731	1
1	754	87737	87743	87749	87754	87760	87766	87772	87777	87783	87789	
l	755	87795	87800	87806	87812	87818	87823	87829	87835	87841	87846	
ı	756	87852	87858	87864	87869	87875	87881	87887	87892	87898	87904	
l	757 758	87910 87967	87915 87973	87921 87978	879 27 87984	87933 87990	879 38 8799 6	87944 88001	87950 88007	87955 88013	87961 88018	1
- 1	759		88030			88047						1

TABLE XXVI.

	No. 7	600	8 20 0					Lo	g. 8809	1	-91381.	
	No.	0	1	2	3	4	5	6	.7	8	9	1
1	760	88081	88087	88093	88098	88104	88110	88116	88121	88127	88133	1
1	761	88138	88144	88150	88156	88161	88167	88173	88178	88184	88190	1
	762	88195	88201	88207	88213	88218	88224	88230	88235	88241	88247	
	763	88252	88258	88264	88270	88275	88281	88287	88292	88298	88304	1 1
. 1	764	88309	88315	88321	88326	88332	88338	88343	88349	88355	88360	1 1
	765	88366	R3372	88377	88383	88389	88395	88400	88406	88412	88417	
1	766	88423	88429	88434	88440	88446	88451	88457	88463	88468	88474	
- 1	767	88480	88485	88491	88497	88502	88508 88564	88513 88570	88519 88576	885 2 5 88581	88530	1 1
	768 769	88536 88593	88 542 8 8 598	88547 88604	88553 88610	88559 88615	88621	88627	88632	88638	88587 886 43	
1						88672	88677	88683				1 1
1	770 771	83649 88705	88655 88711	88660 88717	88666 88722	88728	88734	88739	88689 8 874 5	88694 88750	88700 88756	1 1
į	772	88762	88767	88773	88779	88784	88790	88795	88801	88807	8881 2	1 1
1	773	88818	88824	88829	88835	88840	88846	88852	88857	88863	88868	1 1
	774	88874	88880	88885	88891	88897	88902	88908	88913	88919	88925	1 1
	775	88930	88936	88941	88947	88953	88958	88964	88969	88975	88981	1 1
	776	88986	8899	88997	89003	89009	89014	89020	89025	89031	89037	1 1
1	777	89042	89048	89053	89059	89064	89070	89076	89081	89087	89092	1 1
	778	89098	89104	89109	89115	89120	89126	89131	89137	89143	89148	1 1
I	779	89154	89159	89165	89170	89176	89182	89187	89198	89198	89204	1 1
- 1	780	89209	89215	89221	89226	89232	89237	89243	89248	89254	89260	1
	781	89265	89271	89276	89282	89287	89293	89298	89304	89310	89315	1 1
	782	89321	89326	89332	89337	89343	89348	89354	89360	89365	89371	1 1
-	783	39376	89382	89387	89393	89398	89404	89409	89415	89421	89426	1 1
I	784	89432	89437	89443	89448	89454	89459	89465	89470	89476	89481	
1	785	89487	89492	89498	89504	89509	89515	89520	89526	89531	89537	1 1
ļ	786	89542	89548	89553	89559	89564	89570	89575	89581	89586	89592	1
}	787	89597	89603	89609	89614	89620	89625	89631	89636	89642	89647	1 1
	788 789	8965 3 89708	89658 89713	89664 89719	89669 89724	89675 89730	89680	89686	89691	89697	89702	1 1
							89735	89741	89746	89752	89757	1
j	790 791	89763 89818	89768 898 23	89774 8 982 9	89779	89785	89790	89796	89801	89807	89812	1 1
	792	89873	89878	89883	89834 89889	89840 89894	89845 89900	89851 89905	89856 89911	8986 2 89916	89867 899 22	1
1	793	89927	89933	89938	89944	89949	89955	89960	89966	89971	89977	1
1	794	89982	89988	89993	89998	90004	90009	90015	90020	90026	90031	1 1
ł	795	90037	90042	90048	90053	90059	90064	90069	90075	90080	90086	
	796	90091	90097	90102	90108	90113	90119	90124	90129	90135	90140	
- 1	797	90146	90151	90157	90162	90168	90173	90179	90184	90189	90195	
- 1	798	90200	90206	90211	90217	90222	90227	90233	90238	90244	90249	
	799	90255	90260	90266	90271	90276	90282	90287	90293	90298	90304	
l	800	90309	90314	90320	90325	90331	90336	90342	90347	90352	90358	
ı	801	90363	90369	90374	90380	90385	90390	90396	90401	90407	90412	
1	802	90417	90423	90428	90434	90439	90445	90450	90455	90461	90466	1
1	80 3 804	90472	90477	90482	90488	90493	90499	90504	90509	90515	90590	
		90526	90531	90536	90542	90547	90553	90558	90563	90569	90574	
	805	90580	90.25	90590	90596	90601	90607	90612	90617	90623	90628	Ì
1	806 80 7	90634 9068 7	90639 90693	90644 90698	90650	90655	90660	90666	90671	90677	90682	- 1
l	808	90741	90747	90098	90703 90757	90709 90763	90714 90768	90720 90773	90725 90779	90730 90784	90756 90789	- 1
	809	90795	90800	90806	90811	90816	90822	90827	90832	90838	90843	- 1
	810	90849	90854	90859	90865	90870	90875	90881	90886	90891	90897	Į
1	811	90902	90907	90913	90918	90924	90979	90934	90886	90891	90897	1
ł	812	90956	90961	90966	90972	90977	90982	90988	90993	90998	91004	1
	813	91009	91014	91020	91025	91030	91036	91041	91046	91052	91057	- 1
. [814	91062	91068	91073	91078	91084	91089	91094	91100	91105	91110	
ı	815	91116	91121	91126	91132	91137	91142	91148	91153	91158	91164	
ļ	816	91169	91174	91180	91185	91190	91196	91201	91206	91212	91217	l
1	.817	91222	91228	91233	91238	91243	91249	91254	91259	91265	91270	1
- !	818	91275	91281	91286	91291	91297	91302	91307	91312	91318	91323	1
- 1	819	91328	91334	91339	91344	91350	91355	91360		91371	91376	ı
	No.	0	1 1	2	3	4	5	6	7	8	. 9	اـــ
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LOGARITHMS OF NUMBERS.

No. 820		00.						Log.	01581	94448.
1 0.	. 0	1 1	2	3	4	5 .	6 1	7	8 1	9
820	91381	91387	91392	91397	91403		91413	91418	91424	91429
821	91434		91445	91450	91455	91461	91466	91471	91477	91482
822	91487	91492	91498	91503	91508	91514	91519	91524	915 29	91535
823	91540	91545	91551	91556	91561	91566	91572	91577	91582	91587
824	91593	91598	91603	91609	91614	91619	91624	91630	91635	91640
825 826	91645 91698	91651 91703	91656 91709	91661 91714	91666 91719	91672 91724	91677 91730	91682 91735	91687 91740	91693 91745
827	91751	91756	91761	91766	91772	91777	91782	91787	91793	91798
828	91803	91808	91814	91819	91824	91829	91834	91840	91845	91850
829	91855	91861	91866	91871	91876	91882	91887	91892	91897	91903
830	91908	91913	91918	91924	91929	91934	91939	91944	91950	91955.
831	91960 92012	91965 92018	91971 9 2 023	91976 92028	91981 9 2 033	91986 9 2 038	91991 9 2 044	91997 9 204 9	92002 92054	92007 92059
832 833	92065	92070	92075	92080	92085	92091		92101	92106	92111
834	92117	92122	92127	92132	92137	92143		92153	92158	92163
835	92169	92174	92179	92184	92189	92195	92200	92205	92210	92215
836	92221	92226	92231	92236	92241	92247	92252	92257	92262	92267
837	92273 92324	92278 92330	92283 92835	92288 92340	92293	92298	92304 92355	92309	92314 92366	92319 92371
838 839	92324	92381	92387	92340	92345 92397	92350 92402	92300	92361 92412	92300	92371
840	92428	92433	92438	92443	92449	92454	92459	92464	92469	92474
841	92480	92485	92490	92495	92500	92505	92511	92516	92521	92526
842	92531	92536	92542	92547	92552	92557	92562	92567	92572	92578
843	92583	92588	92593	92598	92603	92609	92614	92619	92624	92629
844	92634	92639	92645	92650	92655	92660	92665	92670	92675	92681
845	92686	92691	92696	92701	92706	92711	92716	92722	92727	92732
846 847	92737 92788	92742 92793	92747 92799	92752 92804	92758 92809	92763 92814	92768 92819	92773 92824	92778 92829	92783 92834
848	92840	92845	92850	92855	92860	92865		92875	92881	92886
849	92891	92896	92901	92906	92911	92 916	92921	92927	92932	92937
850	92942	92947	92952	92957	92962	92967	92973	92978	92983	92988
851	92993	92998	93003	93008	93013	93018	93024	93029	93034	93039
852	93044	9 3049 9 3100	93054 93105	93059 93110	93064 93115	93069 931 20	93075 93125	93080	9308 <i>5</i> 93136	93090 93141
853 854	93095 93146	93151	93156	93161	93166	93171	93176	93131 93181	93186	93192
855	93197	93202	93207	93212	93217	93222	93227	93232	93237	93242
856	93247	93252	93258	93263	93268	93273	93278	93283	93288	93293
857	93298	93303	93308	93513	93318	93323	93328	93334	93339	93344
858	93349 93399	93354	93359 93409	93364 93414	93869	93374	93379	93384	93389	93394
859		93404			93420	93425	93430	93435	93440	93445
860 861	93450 93500	93455 93505	93460 93510	93465 93515	93470 93520	93475 93526	93480 93531	93485 93536	93490 93541	93495 93546
862	93551	93556	93561	93566	93571	93576	93581	93586	93591	93596
863	98601	93606	93611	93616	93621	93626		93636	93641	93646
864	93651	93656	93661	93666	93671	93676	93682	93687	93692	93697
865	93702	93707	93712	93717	93722	93727	93732	93737	93742	93747
866	93752 93802	93757 93807	93762 93812	93767 93817	93772 938 22	93777 93827	9378 2 938 32	93787 93837	93792	93797
867 868	93852	93857	93862	93867	93872	93877	93882	93887	93842 93892	93847 93897
869	93902	93907	93912	93917	93922	93927	93932	93937	93942	93947
870	93952	93957	93962	93967	98972	93977	93982	93987	93992	93997
871	94002	94007	94012	94017	94022	94027	94032	94037	94042	94047
872	94052	94057	94062	94067	94072	94077	94082	94086	94091	94096
873 874	94101 94151	94106 94156	94111 94161	94116 94166	94121 94171	94126 94176	94131 94181	94136 94186	94141 94191	94146 94196
875	94201	94206	94211	94216	94221	94226	94231	94236	94240	94245
875 876	94250	94255	94260	94265	94270	94275	94231	94236	94240	94245
877	94300	94305	94310	94315	94320	94325	94330	94335	94340	94345
878	94349	94354	94359	94364	94369	94374	94379	94384	94389	94394
879	94399	94404	94409	94414	94419	94424	94429	94433	94438	94443
No.	0	1	2	3	4	5	6	7	8	9

TABLE XXVI.

No. 88	9	9400.						.og. 944	48	97313.	
No.	0 1	1 7	2	3	4	5	6	7	8 1	9	
880	94448	94453	94458	94463	94468	94473	94478	94483	94488	94493	
881	94498	94503	94507 94557	94512	94517	94522	94527	94532	94537	94542 94591	ŀ
882 883	94547 94596	94552	94606	94562 94611	94567 94616	94571 94621	94576 94626	94581 94630	94586 94635	94640	ŀ
884	94645	94650	94655	94660	94665	94670	94675	94680	94685	94689	
885	94694	94699	94704	94709	94714	94719	94724	94729	94734	94738	
886	94743	94748	94753	94758	94763	94768	94773	94778	94783	94787	
887	94792	94797	94802	94807	94812	94817	94822	94827	94832	94836	
888	94841	94846	94851	94856 94905	94861	94866	94871	94876 94924	94880 94929	94885	
889	94890	94895	94900		94910	94915	94919			94934	
890 891	94939 94988	94944 94993	94949 94998	94954 9 5002	94959 95007	94963 95012	94968 95017	94973 95022	94978 95027	9498 3 9503 2	
892	95036	95041	95046	95051	95056	95061	95066	95071	95075	95080	
893	95085	95090	95095	95100	95105	95109	95114	95119	95124	95129	l
894	95134	95139	95143	95148	95153	95158	951 63	95168	95173	95177	
895	95182	95187	95192	95197	95202	95207	95211	95216	95221	95226	l
896	95231	95236	95240	95245	95250	95255	95260	95265	95270	95274	1
897 898	95 2 79 95328	95284 95332	95289 95337	95294 95342	95 2 99 95347	9 530 3 9 53 52	95308 95357	95313 95361	95318 95366	953 23 95371	
899	95376	95381	95386	95390	95395	95400	95405	95410	95415	95419	1
900	95424	95429	95434	95439	95444		95453	95458	95463	95468	1
901	95472		95482	95487	95492	95497	95501	95506	95511	95516	ı
902	95521	95525		95535	95540	95545	95550	95554	95559	95564	i
903	95569	95574	95578	95583	95588	95593	95598	95602	95607	95612	١
904	95617	95622	95626	95631	95636	95641	95646	95650	95655	95660	1
905	95665	95670		95679	95684	95689	95694	95698	95703	95708	1
906	95713 95761	95718 95766	95722 95770	95727 95775	95732 95780	95737 95785	957 42 95789	95746 95794	95751 95799	95756 95804	l
908	95809			95823	95828		95837	95842	95847	95852	L
909	95856		95866	95871	95875		95885	95890	95895		1
910	95904	95909	95914	95918	95923	95928	95983	95938	95942	95947	1
911	95952			95966	95971		95980	95985	95990	95995	ł
912	95999				96019		96028	96033	96038	96042	1
913	96047 96095		960 57 961 04	96061 96109	96066 96114	96071 96118	96076		96085	96090	l
914	96142	96147	96152				96123	96128		96137	1
915 916	96190		96199	96156 9 620 4	96161 96209	96166 96213	96171 96218	96175 96223	96180 96227	96185 962 32	١
917	96237	96242	96246	96251	96256	96261	96265		96275		1
918	96284		96294	96298	96303	96308			96322		١
919	96332		96341	96346	96350	96355	96360	96565	96369	96374	1
920	96379		96388	96393	96398		96407	96412	96417	96421	1
921	96426		96435	96440	96445	96450	96454	96459			ì
922 923	96473 96520			96487 96534	96492 96539	96497 96544	96501 96548	96506 96553	9651 F 96538		1
923	96567	96572	96577	96581	96586	96591	96595		96605		1
925	96614			96628	96633	96638	96642	96647	-96652		1
926	96661				96680			96694	96699	96703	1
927	96708	96713	96717	96722	96727	96731	96736	96741	96745		1
928	96755		96764	96769	96774	96778	96783	96788	96792	96797	I
929	96802			96816		96825	96830	96834	96839	96844	-
930	96848		1			96872	96876		96886	96890	
931	96895			96909 96956		96918 96965			96932 96979		1
933	96988						97016		97025		l
934	97035								97072		
935	97081	97086	97090	97095	97100		97109	97114	97118		1
936	97128					97151	97155	97160	97165	97169	1
937	97174						97202			97216	1
938	97220 9 7261	97225 97 2 71	97230 97276							97262 97308	1
No.			<u> </u>			·					4
1 410.	1 0	1 1	1 2	1 3	1 4	5	1 6	7	8	9	•

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TABLE XXVL

<u> </u>					TAME I II						
[<u> </u>	No. 9		1000						97313		99996.
	No.	0		2	3	4	5	6	7	8	9
1	940	97313 97359	97317	97322	97327		97336	97340	973-15	97350	97354
	941 942	97405	97364 97410	97368 97414	97373 97419	97424	97382 97428	97387 97433	97391 97437	97396 97442	97400 97447
	943	97451	97456	97460	97465	97470	97474	97479	97483	97488	97493
	944	97497	97502	97506	97511	97516	97520	97525	97529	97534	97539
1	945	97543	97548	97552	97557	97562	97566	97571	97575	97580	97585
	946	97589	97594	97598	97603	97607	97612	97617	97621	97626	97630
1	947	97635	97640	97644	97649	97653	97658	97663	97667	97672	97676
	948 949	97681 977 2 7	97685 97731	97690 97736	97695 97740	97699 97745	97704 97749	97708 97754	97713 97759	97717 97763	9772 2 97768
1	950	97772	97777	97782	97786	97791	97795	97800	97804	97809	97813
	951	97818	97823	97827	97832	97836	97841	97845	97850	97855	97859
1	952	97864	97868	97873	97877	97882	97886	97891	97896	97900	
	953	97909	97914	97918	97923	97928	97932	97937	97941	97946	97950
	954	97955	97959	97964	97968	97973	97978	97982	97987	97991	97996
	955	98000	98005	98009	98014	98019	98023	98028	98032	98037	98041
	956 957	98046 98091	98050 98096	98055 98100	98059 98105	98064 98109	98068 98114	98073 98118	98078	98082	98087 981 32
	958	98137	98141	98146	98150	98155	98159	98164	98123 98168	981 27 981 7 3	
1	959	98182	98186	98191	98195	98200	98204	98209	98214	98218	
	960	98227	98232	98236	98241	98245	98250	98254	98259	98263	98268
	961	98272	98277	98281	98286	98290	98295	98299	98304	98308	98313 [,]
[962	98318	98322	98327	98331	98336	98340	98345	98349	98354	98358
	963 964	98363 98408	98367 98412	98372 98417	98376 98421	98381 98426	98385 98430	98390 98435	98394	98399	98403
	965	98453	98457	98462					98439	98444	
	966	98498	98502	98507	98466 98511	98471 98516	98475 985 2 0	98480 98525	98484 98529	98489 98534	9849 3 98 53 8
1	967	98543	98547	98552	98556	98561		98570	98574	98579	
l	968	98588	98592	98597	98601	98605		98614	98619	98623	
	969	98632	98637	98641	98646	98650	98655	98659	98664	98668	98673
H	970	98677	98682	98686	98691	98695	98700	98704	98709	98713	
	971 972	987 22 98767	98726 98771	98731 98776	98735 98780	98740 98784	98744 98789	98749	98753	98758	
i	973	98811	98816	98820	98825	98829	98834	98793 988 3 8	98798 98843	988 02 98847	98807 98851
	974	98856	98860	98865	98869	98874	98878	98883	98887	98892	
	975	98900	98905	98909	98914	98918	98923	98327	98932	98936	98941
1	976	98945	98949	98954	98958	98963	98967	98972	98976	98981	98985
	977	98989	98994	98998	99003	99007	99012	99016	99021	99025	
	978 979	99034 99078	99038 99083	99043	99047 9909 2	99052 99096	99056 99100	99061 99105	99065 99109	99069	
1	980	99123	99127	99181	99136	99140	99145	99149		99114	99118
1 1	981	99167	99171	99176	99180	99185	99189	99193	99154 99198	99158 99 202	99162 99 2 07
١,	982	99211	99216	99220	99224	99229	99233	99238	99242	99247	99251
	985	99255	99260	99264	99269	99273	99277	99282	99286	99291	99295
,	984	99300	99304	99308	99313	99317	99322	99326	99330	99335	
	985	99344	99348	99352	99357	99361	99366	99370	99374	99379	1 1
	986 987	99388 99432	99392 99436	99 396 99441	99401 99445	99405 99449	99410 99454	99414 99458	99419 99463	99 423 99467	
	988	99476	99480	99484	99489	99493	99498	99502	99506	99511	99471 99515
	989	99520	99524	99528	99533	99537	99542	99546	99550	99555	99559
	990	99564	99568	99572	99577	99581	99585	99590	99594	99599	99603
1	991	99607	99612	99616	99621	99625	99629	99634	99638	99642	99647
	992	99651	99656	99660	99664	99669	99673	99677	99682	99686	99691
	993 994	99695 99739	99699 99743	99704 99747	99708 99752	99712 99756	99717 99760	99721 99765	99726 99769	99730 99774	99734
	995	99782	99787	99791	99796	99800		_			99778
	996	99826	99850	99835	99839	99843	998 04 99848	998 08 998 52	99813 99856	99017 99861	99822 99865
	997	99870	99874	99878	99883	99887	99891	99896	99900	99904	99909
	998	99913	99917	99922	99926	99930	99935	99939	99944	99948	99952
	999	99957	99961	99965	99970	99974	99978	99983	99987	99991	99996
[No.	0	1	2	3	4	5	6	7	8	9

Log. Sines, Tangents and Secants.

90 Degs.

Degs. 89.

TABLE XXVII. Log. Sines, Tangents and Secants.

1	1 E)eg						•	,	6			Degs. 1
			NIF4	·¥.	He	PUTE	M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant
	0	11	52	0	0	8	ō	8.24186			11.75808		11.75814
	ĭ			52	1	8	8	24903		24910		00007	75097
	2		bl	44		8		25609	99993	25616	74384	00007	74391
1	3		51	36		8	24	26304	99993		73688	00007	73696
	4		51	28		8	32	26988	99992	26996	73004	00008	73012
	5	11	51	30	0	8	40	8.27661	9.99992	8.27669	11.72331	10.00008	11.72339
	6	-	51	12	1	8	48	28324	99992	28332	71668	60008	71676
	7		51	4	1	8	56	28977	99992	28986	71014	80000	71023
	8		50	56		9	4	29621	99992	29629	70371	00008	70379
	9		50	48		9	12	30255	99991	30263	69737	00009	69745
1	0	11	50	40	0	9	20	8.30879	9.99991	8.30888	11.69112	10.00009	11.69121
	i		50	32	1	9	28	31495	99991	31505	68495		68505
	2	ŀ	50	24	l	9	36	52 103	99990	32112	67888	00010	
	3		50	16		9	44	32702	99990	32711	672 89		67298
1 1	4		50	8	ŀ	9	52	33292	99990	33302	66698	00010	66708
+	5	11	50	ō	0	10	0	8.33875	9.99990	8.33886	11.66114	10.00010	11.66125
	6		49	52		10	8	34450	99989		65539		65550
	7		49	44	l	10	16	3501 8	99989	35029	64971	00011	649 82
	8		49	36		10	24	35578	99989	35590	64410		64422
	9		49	28	l	10	32	36131	99989	36143	63 85 7	00011	63 869
9	0	11	49	20	0	10	40	8.36678	9.99988	8.36689	11.63311	10.00012	11.63322
1 2			49	12	ľ	10	48	37217	99988	37229	62771	00012	62783
2			49	4		10	56	37750	99988	37762	62238	00012	62250
2	3		48	56		11	4	38276	99987	38289	61711	00013	61724
2	4		48	48		11	12	38796	99987	38809	61191	00013	61204
2	5	11	48	40	0	11	20	8.39310	9.99987	8.39323	11.60677	10.00013	11.60690
2	6		48	32		11	28	39818	99986	39832	60168	00014	60182
2	7		48	24		11	36	40320	99986	40334	59666		596 80
2	8		48	16		11	44	40816	99986	40830	59170		59184
2	9		48	8		11	52	41307	99985	41321	58679	00015	5 8693
3	0	11	48	Õ	0	12	o	8.41792	9.99985	8.41807	11.58193	10.00015	
3	1		47	52		12	8	42272	99985	42287	57713		57728
3			47	44		12	16	42746	99984	42762	57238		57254
3			47	36	1	12	24	43216	99984	43232	56768	00016	56784
3	4		47	28	L	12	32	43680	99984	43696	56304	00016	56320
3	5	11	47	20	0	12	40	8.44139	9.99983	8.44156	11.55844	10.00017	
3	6		47	12		12	48	44594	99983	44611	55389	00017	55406
3			47	4		12	56	45044	99983	45061	54939	00017	54956
	8		46	56		13	4	45489	99982	45507	54493		54511
3	9		46	43	Ŀ	13	12	45930	99982	45948	54052	00018	. 54070
4	4)	11	46	\$	0	13	20	8.46366	9.99982	8.46385		10•00018	11.53634
4			46	32		13	28	46799	99981	46817	53183		53201
	2		46	24		13	36	47226	99981	47245	52755		52774
4			46	16		13	44	47650	99981	47669	52331	00019 00020	52350 51931
4	-		46	8	_	13	52	48069	99980	48089	51911		
4		11	46	0	0	14	0	8.48455	9.99980			10.00020	
4			45	52		14	8	48896	99979	48917	51083	00021	51104
4			45	44		14	16	49304	99979	49325	50675	00021	50696 50292
4			45	36		14	24	49708	99979	49729	50 27 1 49870		49892
4	_		45	28	_	14	32	<i>5</i> 0108	99978	50130			
5	~ 1	11	45	20	0	14	40	8.50504	9.99978	8.50527		10.00022	
5			45	12		14	48	50897	99977	50920	49080		49103 48713
5			45	4		14	56	51287	99977	51310	48690 48304	90023	48327
5			44	56		15	4	51673 52055	99977 99976	51696 52079	47921	00023	47945
5		_	44	48		15	12						
5		11	44	40	0	15	20	8.52434	9.99976		11.47541	10.00024	47190
5			44	32	l	15	28	52 810		1			46817
5			44	24	l	15	36	53183					1
5			44	16	l	15	44	5855% 53919					
6			44	8		15 16	5 2	54282		54308			45718
-													
M	1.	H	uri	·.M.	Hk	WY.	M.	Co-sine.	Sine.	Co-tang	rangent	Co-secant	De

TABLE XXVII.

2 Deg	79.		_					Degs.	177.
M.		Hourp.m.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M.
0	11 44 (0 16 0	8.54282	9.99974	8.54308	11.45692	10.00026		60
1	43 5		54642	99973	54669	45331	00027	45358	59
2	43 44		54999	99973	55027	44973	00027	45001	58
3	43 36		55354	99972	55382	44618	00028	44646	57
4	43 28		55705	99972	55734	44266	00028	44295	56
5	11 43 20		8.56054	9.99971	8.56083			11.43946	55
6	43 19		56400	99971	56429	43571	00029	43600	54
7	43 4		56743	99970	56773	43227	00030	43257	53
8	42 56		57084	99970	57114	42886	00030	42916	52
9	42 48		57421	99969	57452	42548	00031	42579	51
10	11 42 40		8.57757	9.99969				11.42243	50
11	42 32		58089	99968	58121	41879	00032	41911	49
12	42 24		58419	99968	58451	41549	00032	41581	48
13	42 16 42 8		58747	99967	58779		00033	41253	47
14		4	59072	99967	59105	40895	00033	40928	46
15	11 42 (8.59395	9.99967			10.00033		45
16	41 52		59715	99966	59749		00034	40285	44
17	41 44		60033 60 34 9	99966 99965	60068 60384			39967	43
18 19	41 28		60662	99964	60698		00035 00036	39651 39338	42 41
									
20 21	11 41 20		8.60973 61282	9.99964				11.39027	40
21 22	41 4		61589	99963 99963	61319 61 62 6		00037 00037	38718	39 38
23	40 56		61894	99962	61931	38374 38069	00038	38411 38106	38 37
24	40 4		62196	99962			00038	37804	36
25	11 40 40		8.62497	9.99961					
26	40 32		- 62795	99961	62834			11.37503 37205	35 34
27	40 24		63091	99960		36869		36909	33
28	40 16		63385	99960				36615	32
29	40 8		63678	99959	63718		00041	36322	31
30	11 40		8.63968	9.99959				11.36032	30
31	39 59		64256	99958	64298		00042	35744	29
32	39 44		64543	99958	64585			35457	23
33	39 36		64827	99957	64870			35173	27
34	39 28	20 32	65110	99956	65154		00044	34890	26
35	11 39 20	0 20 40	8.65391	9.99956	8.65435		10.00044		25
36	39 19		65670	99955	65715	34285		34330	24
37	39 4		65947	99955				34053	23
38	38 56		66223	99954	66269		00046	33777	22
39	38 48	21 12	66497	99954	66543	33457	00046	33503	21
40	11 38 40	0 21 20	8.66769	9.99953	8.66816	11.33184	10.00047	11.33231	20
41	38 32		67039	99952		32913		32961	19
42	38 24		67308	99952				32692	18
43	38 16		675 75	99951	67624	32376		32425	17
44	38 8		67841	99951	67890	32110	00049	32159	16
45	11 38 (9.99960	8.68154	11.31846	10.00050	11.31896	15
46	37 59		68367	99949	68417		00051	31633	14
47	37 44		68627	99949	68678	31322		31373	13
48	37 36 37 28		68886	99948	68938			31114	
49			69144	99948	69196			30856	11
50	11 37 20		8.69400	9.99947		11.30547		11.30600	10
51	37 19		69654	99946	69708			30346	9
52 52	37 4	22 56	69907	99946	69962			30093	8
53 54	36 56 36 48		70159	99945	70214			29841	7
			70409	99944	70465			29591	6
55	11 36 40		8.70658	9.99944		11.29286		11.29342	5
56 57	36 32 36 24		70905	99943	70962			29095	4
58	36 24 36 16		71151	99942	71208			28849	3
59	36 8	23 44 23 52	71395 71638	99942 99941	71453			28605	:
60	36 0	24 0	71880	99940	71697 71940	28303 28060		28362	1
M.		Houra.M.						28120	0
12 De		LLUGIA.E.	CO-SIDE.	Sine.	Co-tang.	Langent	Co-secant	Secant.	M.
	•-	-			_		Digitized	Deen.	87

M Houra M Houra M Sine Co-sine Tangent Co-tang Secant Co-secant M	3	B Deg	3.					_		·-, - ·	ones and	ocounts.		Degs.	176
0 11 36 0 0 24 0 8.711840 9.99940 8.711940 11.28060 10.00060 11.28120 60 2 35 44 24 16 72359 99940 73181 73181 73181 73181 3 3 3 3 3 5 6 2 4 2 72359 99939 73400 27760 00061 37641 58 3 3 3 5 5 2 4 40 8.73069 9.99938 72896 27341 00062 27403 57 74761 58 58 58 58 24 52 72334 99938 72896 27104 00062 27163 56 58 58 56 25 4 24 56 73353 99936 73360 26400 00064 26665 53 53 56 56 54 73767 99936 73366 26640 00064 26665 53 58 56 25 4 73767 99934 73802 26168 0.0066 26633 52 58 58 56 25 4 73767 99934 73802 26168 0.0066 26633 52 58 73567 99354 73600 26400 00064 26665 53 53 53 53 53 53 53	1			our	.w.	Ho	uri	P.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.		
3 35 36 44 24 16 7 23.59 999.39 7 24.00 27.580 00061 276.41 58 3.5 36 24 24 725.97 999.38 728.96 271.04 00062 271.66 56 56 13.5 12 24 42 725.97 999.38 728.96 271.04 00062 271.66 56 56 35 12 24 42 73.03 999.36 73.566 265.34 00064 256.67 37.67 9 36 4 24 56 73.05 999.36 73.566 265.34 00064 256.67 37.67 9 36 4 24 56 73.05 999.36 73.566 265.34 00064 256.65 33 34 4.56 25 12 73.05 999.36 73.566 265.34 00064 256.65 33 34 4.56 25 12 73.05 999.38 74.06.8 29.93 1.66 00.06 26.66 35 36 11 1 3.4 40 0 5.2 0 8.742.26 9.999.38 74.06 25 29.93 1.27 1.27 1.27 1.27 1.27 1.27 1.27 1.27		_	11					_			8.71940	11.28060	10.00060	11.28120	60
S	!!														
S	П					1									
6 35 12 24 48 73303 99936 73366 26634 00064 26697 34 8	H			35	28		24	32							
To So 4 24 56 73555 99935 73600 25400 00064 25465 35 35 35 36 35 43 36 25 4 73767 99935 74063 25937 00066 26003 31 1 34 40 0 25 25 74454 99933 74621 2479 00067 25364 36 36 36 36 36 36 36			11			0						11.26368	10.00063	11.26931	55
S	П					1								1	
11 34 40 0 25 20 8.74226 9.9934 8.74292 11.25708 10.00666 26003 51 11 34 32 25 25 74454 99933 744521 25479 00067 2.5344 62 62 62 62 74560 99932 74474 26202 00068 26094 47 47 47 47 47 47 47	Н					ı									
10	H					l									
12	1	10	11	34		0			8.74226	9.99934		11.25708	10.00066	·	
14						Ì									
14	Н					١.									
15						ŀ									
16	iŀ	15	11	34	0	0	26	0	8.75353	9.99930	8.75423	11.24577	10.00070		
18 33 36 26 24 76015 99920 76007 23913 00072 23985 42 19 33 28 26 32 76234 99927 76306 23934 00073 23766 41 20 11 33 20 0 26 40 8.76451 9.99926 76742 23528 00074 11.23549 40 21 33 12 26 48 76667 99926 76742 23528 00074 23333 39 22 33 4 26 56 76883 99925 76958 23042 00075 23117 38 23 32 56 27 4 77097 99924 77173 22827 00076 23913 70077 22690 36 24 32 48 27 12 77310 99923 77387 22613 00077 22690 36 26 32 32 27 28 77733 99922 77811 22897 00076 22267 34 22 32 4 27 36 77945 99921 78022 21978 00079 22057 33 224 27 36 77945 99921 78022 21978 00079 22057 33 22 4 27 36 77945 99920 78232 21768 00680 21848 32 9 32 8 27 52 78360 99920 78241 21559 00680 21848 33 1 31 52 28 8 78774 99918 78845 21559 00680 21640 31 31 31 52 28 8 78774 99918 78855 21145 00682 21226 23 32 32 14 28 16 78979 99917 79061 20939 00683 20817 27 33 34 31 31 32 28 48 79789 99917 79061 20939 00683 20817 27 34 28 28 28 28 28 28 28 28 28 28 28 28 28	Н														44
19						ı						1			
20 11 33 20 0 26 40 8.76451 9.99926 8.76525 11.23475 10.00074 11.23549 40 21 33 12 26 48 76667 99926 76742 23258 00074 23333 99925 76958 23042 00075 23117 38 23 32 56 27 4 77097 99924 77173 22827 00076 22903 37 25 25 11 32 40 0 27 20 8.77522 9.99923 77387 22613 00077 22650 36 25 27 28 77330 99924 77387 22613 00077 22650 36 25 27 28 27 28 77330 99922 77811 22189 00078 22267 34 27 32 24 27 36 77943 99921 78022 21978 00079 22067 33 22 29 32 8 27 52 78360 99920 78252 21768 00080 21640 31 31 31 52 28 8 78774 99918 8.78649 11.21551 10.00081 11.2142 30 31 31 52 28 8 78774 99918 8.78649 11.21551 10.00081 11.2126 30 31 31 52 28 8 78774 99918 78855 21145 00082 21226 29 32 31 34 28 16 78979 99917 79061 20939 00083 21021 28 33 31 34 28 16 78979 99917 79061 20939 00083 21021 28 33 31 35 22 88 8 78774 99918 78855 21145 00082 21226 29 32 31 34 28 16 78979 99917 79061 20939 00083 21021 28 33 31 35 28 28 24 79183 99917 79266 20734 00083 20817 27 33 31 35 28 28 38 78778 99916 79470 20530 00084 20614 26 36 31 12 28 48 79789 99916 79470 20530 00084 20614 26 36 31 12 28 48 79789 99913 80076 19924 00087 19811 22 30 40 11 30 40 0 29 20 8.80585 99918 80277 19725 00086 20211 24 30 30 40 0 29 20 8.80585 99918 80277 19725 00086 20211 24 30 30 40 0 29 20 8.80585 99918 80277 19725 00086 11.20412 25 30 48 29 36 80978 99909 8136 80776 19924 00087 19811 22 21 44 30 32 29 28 80782 99910 80872 19128 00090 19221 12 44 30 8 29 52 81367 99908 81469 18541 00092 18653 16 46 29 52 30 8 81752 99906 81464 18736 00091 11.2440 15 50 12 92 20 0 30 40 8.82513 9.99908 81469 18541 00092 18653 16 50 11 29 20 0 30 40 8.82513 9.99908 82420 17580 00096 17676 11 50 12 92 20 0 30 40 8.82513 9.99908 81469 18541 00092 18653 16 50 12 92 20 8.83630 99898 83504 11.17390 10.00097 11.17487 10 50 12 92 20 0 30 40 8.82513 9.99908 81464 18154 00092 18653 16 50 12 92 20 8.83630 99898 83501 11.17463 10 00099 17112 85 50 12 92 30 30 32 83524 99900 83176 16659 00101 16654 5 50 20 48 31 32 88 56 30 99898 83501 16659 00101 16654 5 50 20 48 31 32 88 56 30 99898 83501 16659 00101 1	Н														
21 33 12 26 48 76667 99926 76742 23258 00074 23333 99 22 33 4 26 56 76883 99925 76958 23042 00075 23117 38 23 32 56 27 4 77097 99924 77173 22827 00076 22903 57 24 32 48 27 12 77310 99923 77387 22613 00077 22690 36 25 11 32 40 0 27 20 8.77522 9.99923 8.776001 11.22400 10.00077 11.22478 35 26 32 32 27 28 77733 99921 78022 21978 00079 22067 34 27 32 24 27 36 77943 99921 78022 21978 00079 22067 33 28 32 16 27 44 78162 99920 78232 21768 00080 21848 32 29 32 8 27 52 78360 99920 78232 21768 00080 21848 32 29 32 8 27 52 78360 99920 78432 21559 00080 21640 31 30 11 32 0 0 28 0 8.78568 9.99919 8.78649 11.21351 10.00081 11.21432 30 31 31 52 28 8 78774 99918 78855 21145 00082 21226 29 32 31 44 28 16 78979 99917 79266 20734 00083 20012 23 33 31 36 28 24 79183 99917 79266 20734 00083 20121 23 34 31 28 28 32 79386 99916 79470 20530 00084 20614 26 35 11 31 20 0 28 40 8.79588 9.99918 8.79673 11.20327 10.00085 11.20412 25 36 31 12 28 48 79789 99914 79875 20125 00086 20211 24 37 31 4 28 56 79990 99913 80076 19924 00087 20010 23 38 30 56 29 4 90189 99913 80076 19924 00087 20010 23 38 30 56 29 4 80189 99913 80076 19924 00087 20010 23 38 30 48 29 12 80388 99912 80476 19524 00088 19612 21 40 11 30 40 0 29 20 8.80585 9.99911 8.80674 11.19326 10.00089 11.20412 25 40 11 30 40 0 29 20 8.80585 9.9991 80872 19128 00090 49218 19 42 30 24 29 36 80978 99909 81068 19932 00091 11.8241 12 42 30 24 29 36 80978 99909 81068 19932 00091 11.8241 12 43 30 16 29 44 81173 99908 81469 18932 00091 18827 17 44 30 32 29 28 80782 99900 81264 18736 00091 18827 17 45 30 16 29 44 81173 99904 82230 17770 00096 17866 12 49 29 28 30 32 82324 99904 82420 17580 00096 17666 11 50 11 29 20 0 30 40 8.82513 9.99903 83561 16639 00102 11.6554 5 51 128 40 0 31 20 8.8346 9.9989 83561 16639 00102 11.6554 5 52 29 4 30 56 83888 99901 82991 11.6648 00102 11.6554 5 54 28 48 31 12 83560 99898 83561 16639 00100 16024 5 55 12 28 48 31 52 84870 99898 83561 16639 00100 16024 5 55 28 48 31 52 84870 99898 83561 16639 00100 16004 2 56 28 0 32 0 48458 99894 84464 15556 00106 16004 2 56 2	t		11			0								<u>'</u>	
23						ľ						232 58	00074		
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93 Dogs. Digitized by COSIDegs. 50	\bot			urr	.N.	Ho)127/	. W.	Co-sine.	Sine.	Co-tang				
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Log. Sines, Tangents and Secants.

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	٠,		7	-	34					11.12894			
15	11	26			34 34	-		9.99880 99879	8.87106 87277	12723		12844	45 44
16 17	1	25			34	-		99879	87447	12553		12675	43
18	Į	25			34			99878	87616	12384		12506	42
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20	īī		20	1_				9.99876		11.12047			40
21	١.,	25	12		34			99875	88120	11880		12005	39
22		25	14	1	34			99874	88287	11713		11839	38
23		24	56	1	35		88326	99873	88453	11547	00127	11674	37
24	1	24	48		35			99872	88618	11382	00128	11510	36
25	11	24	40	0	35	20	8.88654	9.99871	0 00703	11.11217	10.00129		35
26		24	32		35		88817	99870	88948	11052	00130	11183	34
27		24	24]	35		88980	99869	89111	10889	00131	11020	33
28		24	16	1	35		89142	99868	89274	10726	00132	10858	32
29		24	8	1	35		89304	99867	89437	10563	00133	10696	31
50	11	24	0	To	36	0	8.89464	9.99866	8.89598	11.10402		11.10536	30
31		23	52		36	8	89625	99865	89760	10240	00135	10375	29
32		23	44	1	36		89784	99864	89920	10080	00136	10216	28
33		23	36	1	36	24	89943	99863	90080	09920	00137	10057	27
14		23	28		36	32	90102	99862	90240	09760	00138	09898	26
15	11	23	20	0	36	40	8.90260	9.99861	8.90399	11.09601	10.00139	11.09740	25
16		23	12		36	48	90417	99860	90557	09443	00140	09583	24
17		23	4		36	56	90574	99859	90715	09285	00141	09426	23
18		22	56		37	4	90730	998 <i>5</i> 8	90872	09128	00142	09270	22
19		22	48		37	12	90885	99857	91029	08971	00143	09115	21
	11	22	40	0	37	20	8.91040	9.99856	8.91185	11.08815	10.00144	11.08960	20
11		22	32	1	37	28	91195	99855	91340	08660	00145	08805	19
2		22	24		37	36	91349	99854	91495	08505	00146	08651	18
3		22	16		37	44	91502	99853	91650	08350	00147	08498	17
4		22	8	_	37	52	91655	99852	91803	08197	00148	08345	16
	11	22	0	0	38	0	8.91807	9.99851			10.00120	11.08193	15
6		21	52		88	8	91959	99850	92110	07890	00150	08041	14
7		21	44		38	16	92110	99848	92262	07738	00152	07890	13
8		21 21	36 28		38 38	24	92261	99847	92414	07586	00153	07739	12
9		_				32	92411	99846	92565	07435	00154	07589	11
- 1	11	21	20	0	38	40	8.92561	9.99845	8.927161		10.00155	11.07439	10
1		21	12		38	48	92710	99844	92866	07134	00156	07290	9
3		21 20	56		38 39	56 4	92859 93007	99843	93016	06984 06835	00157	07141 06993	8
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9		20 20	10		39	52	93885	99836	94049	05951	00164	06115	il
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Log. Sines, Tangents and Secants.

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	M		our.		13.		-	Sine.	Carina	Tanant	Catana	Bassat	Degs. 1	. M
I I-					<u> </u>				Co-sine.		Co-tang.		Co-secant	
	0	111	20	0	0	40	0				11.05805			60
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11	4		19	28	ŀ	40	32	94603	99830			00170		56
		-	19	20	0	40	40	8.94746	9.99829		11.05083		11.05254	55
	5 6	11	19	12	٧	40	48	94887	9.99829	95060	04940	00172		
1 1	7		19	4		40	56	95029	99827	95202	04798	00173		53
	8	į	18	56		41	4	95170	99825	95344	04656	00175		
11	9		18	48		41	12	95310	99824	95486	04514	00176	04690	
	10	11	18	40	0		20	8.95450	9.59823				11.04550	
1	iĭ		18	32	ľ	41	28	95589	99822	95767	04233	00178		49
Н	12		18	24		41	36	95728	99821	95908		00179		
11	13		18	16		41	44	95867	99820	96047	03953	00180	04133	47
11	14		18	8		41	52	96005	99819	96187	03 813	00181	03995	46
	15	11	18	0	0	42	0	8.96143	9.99817	8.96325	11.03675	10.00183	11.03857	45
	16	_	17	52		42	8	96280	99816	96464	03536	00184	03720	44
!	17		17	44	}	42	16	96417	99815	96602	03398	00185	03583	43
	18		17	36		42	24	96553	99814	96739	03261	00186	03447	42
1 1	19		17	28		42	32	96689	99813	96877	03123	00187	03311	41
	20	11	17	20	0	42	40	8.96825	9.99812	8.97013	11.02987	10.00188	11.03175	40
	21		17	12	_	42	48	96960	99810	97150	02850	00190	03040	39
	22		17	4	•	42	56	97095	99809	97285	02715	00191	02905	38
	23		16	56		43	4	97229	99808	97421	02579	00192	02771	37
<u> </u>	24		16	48		43	12	97363	99807	97556	02444	00193	02637	36
	25	11	16	40	0	43	20	8.97496	9.99806	8.97691		10.00194	11.02504	35
	26_		16	32		43	28	97629	99804	97825	02175	00196	02371	34
	27		16	24	-	43	36	97762	99803	97959	02041	60197	02238	33
	78		16	16		43	44	97894	99802	98092	01908	00198	02106	32
-	29	_	16	8		43	52	98026	99801	98225	01775	00199	01974	31
	30	11	16	0	0	44	0	8.98157	9.99800		11.01642	10.00200	11.01843	30
	31		15	52		44	8	98283	99798	98490	01510	00202	01712	29
	32		15	44		44	16	98419	99797	98622	01378	00203	01581	28.
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	36	11	15	20	0	44	40	8.98808	9.99793		11.00985		11.01192	25
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	38		14	56		45	4	99194	99790	99 27 5 99 4 05	00725 00595	00209	009 3 4 00806	23 22
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	43	ĺ	14	16		45	44	99830	99783		10.99954	00217	00170	17
	44		14	8		45	52	99956	99782	00174	99826	00218	00044	16
I	46	11	14	0	0	46	0	9.00082	9.99781		10.99699		10.99918	15
	46	* *	13	52	ľ	46	8	00207	99780	00427	99573	00220	99798	14
	47		13	44		46	16	00332	99778	00553	99447	00222	99668	13
1	48		13	36		46	24	00456	99777	00679	99321	00223	99544	12
	49		13	28		46	32	00581	99776	00805	99195	00224	99419	11
-	50	11	13	20	0	46	40	9.00704	9.99775	9.00930	10.99070	10.00225	10.99296	10
	51		13	12		46	48	00828	99773	01055	98945	00227	99172	9
	52		13	4		46	56	00951	99772	01179	98821	00228	99049	8
	53	ł	12	56		47	4	01074	99771	01303	98697	00229	98926	7
1	54		12	48		47	12	01196	99769	01427	98573	00231	98804	6
	55	11	12	40	0	47	20	9.01318	9.99768	9.01550	10.98450	10.00232	10.98682	5
	56		12	32		47	28	01440	99767	01673	98327	00233	98560	4
	57		12	24		47	36	01561	99765	01796	98204	00235	98439	3
	58		12	16		47	44	01682	99764	01918	98082	90236	98318	. 2
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Log. Sines, Tangents and Secants.

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Ho	UFA	.w.	Ho	urp	M.	Sine.	Co-sine.	Tangent.			Co-secant	M
11		0	0	48	0	9.01923	9.99761			10.00239		60
	11	52		48	8	02043	99760		97717	00240	- :	59
	11	44		48	16	02163	99759	02404	97596		97837	58
	11	36 28		48 48	24 32	02283 02402	99757	02525	97475	00243		57
	11		<u> </u>				99756	02645	97355	00244		56
	11	20	0	48	40	9.02520	9.99755				10.97480	55
	11	12		48	48	02639	99753	02885	97115	00247	97361	54
	11 10	4 56		48 49	56 4	02757	99752	03005	96995		1	53
	10	48		49	12	02874 02992	99751 99749	03124 03242	96876 9 67 58		97126 97008	52 51
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	10	40	0	49	20	9.03109	9.99748			10.00252 00253		50
	10 10	32 24		49 49	28 36	03226 03342	99747 99745	03479 03597	96521 96403			49 48
	10	16	ļ	49	44	03342	99744	03714	96286		96542	47
	10	8		49	52	03574	99742	03832	96168		96426	46
11	10	-0	0	50	- 0	9.03690	9.99741			10.00259	10.96310	45
11	9	52	U	50	8	03805	9.99740	04065	959 3 5		96195	44
	9	44		50	16	03920	99738	04181	95819		96080	43
	9	36		50	24	04034	99737	04297	95703		95966	42
	9	28		50	32	04149	99736	04413	95587			41
11	9	20	0		40	9.04262	9.99734	9.04528			10.95738	40
4.1	9	12	ľ	50	48		9.99734		95357			
	9	4	1	50	56	04490		04758	95242		95510	
	8	56		51	4	04603	99730		95127			37
	8	48		51	12	04715	99728	04987	95013	1		36
11	8	40	0	51	20	9.04828	9.99727			10.00278	10.95179	35
••	8	32	ľ	51	28	04940	99726	05214	94786		95060	34
	8	24		51	36	05052	99724	05328	94672		94948	33
	8	16		51	44	05164	99723	05441	94559		94836	32
	8	8		51	52	05275	99721	05553	94447	00279	94725	31
11	8	0	0	52	0	9.05386	9.99720	9.05666	10.94334	10.00280	10.94614	30
	7	52	ľ	52	8	05497	99718	05778	94222	00282	94503	29
	7	44		52	16	05607	99717	05890	94110	00283	. 94393	28
	7	36		5 2	24	05717	99716	06002	93998	00284	94283	27
	7	28		5 2	32	05827	99714	06113	93887	00286	94173	26
11	7	20	0	52	40	9.05937	9.99713	9.06224	10.93776	10.00287	10.94063	25
	7	12		52	48	06046	99711	06335	93665	00289	93954	24
	7	4		52	56	06155	99710	06445	935 55	00290	93845	23
	6	56		53	4	06264	99708	06556	93444	00292	93736	22
	6	48		53	12	06372	99707	06666	93334	00293	93628	21.
11	6	40	0	53	20	9.06481	9.99705	9.06775	10.93225	10.00295	10.93519	20
	6	32		53	28	06589	99704	06885	93115	00296	93411	19
	6	24		53	36	06696	99702	06994	93006	00298	93304	18
	6	16		53	44	06804	99701	07103	92897	00299	93196	17
	_6	8	_	53	52	06911	99699	07211	92789	00301	93089	16
11	6	0	0	54	Ü	9.07018	9.99698		10.92680		10.92982	15
	5	52		54	8	07124	99696	07428	92572	00304	92876	14
	5	44		54	16	07231	99695	07536	92464	00305	92769	13
	5	36		54	24	07337	99693	07643	92357	00307	92663	12 11
	5	28		54	32	07442	99692	07751	92249	00308	92558	
11	5	20	0	54	40	9.07548	9.99690			10.00310		10
	5	12		54	48	07653	99689	07964	92036	00311	92347	9
	5	4		54	5 6	07758	99687	08071	91929	00313 00314	9 2242 9 2137	8 7
		56		55 55	12	07863 07968	99686 99684	08177 08 28 3	918 23 91717	00314	92032	6
		48										
11	4	40	0	55	20	9.08072	9.99683				10.91928	5
		32		55	28	08176	99681	08495	91505	00319 00320	918 24 917 2 0	4
		24		55 55	36	08280 08383	99680 99678	08500 08705	91400	00322	91617	2
	4	16		55	44	08486	99677	08810	91190	00323	91514	ĩ
		ol		K.K.								
	4	8		55 56	52	08589	99675	08914	91086	00325	91411	ō

Degs.

Degs. 23.

	7 Des	ps .							,	90000		•	Degs	. 172.
ī	M		ULA	.Mr	Ho	urr	.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
i	0	11	4	0	0	56	-0	9.08589	9.99675		10.91086	10.00325	10.91411	60
١	1		3	52		<i>5</i> 6	8	08692	99674		90981	00326	91308	59
	2		3	44		56	16	08795	99672		90877	00328	91205	58
ı	3		3	36	! L	56	24	08897	99670		90773	00330	91103	57
П	4		3	28		56	32	08999	99669	09330	90670	00331	91001	56
Н	5	11	3	20	(0	66	40	9.09101	9.99667					55
	6	ĺ	3	12	1	56	48	09202	99666		90463 90360	00334 00336	90798 90696	54 53
П	7	1	3	4 56	1	56 57	56	09304 09405	99664				90595	52
Н	9	1	2	48	1	57	12	09506			90155		90494	51
ì		11	- 2	40	0	57	20	9.09606					10.90394	50
П	10 11	14	2	32		57	28	09707				00342	90293	49
Н	12	1	2	24		57	36	09807					90193	48
Н	13	1	2	16		57	44	09907					90093	47
Н	14	l	2	8	1	57	5 2	10006	99653	10353	89647	00347	89994	46
H	15	11	2	0	0	58	0	9.10106	9.99651	9.10454	10.89546	10.00349	10.89894	45
	• 16		ĩ	52	!	58	8	10205	99650	10555	89445	00350	89795	44
	17	1	1	44	ł	58	16	10304					89696	43
	18	1	1	36		58	24	10402					89598	42
l	19	L	1	28	_	58		10501	99645			00355	89499	41
l	20	11	_	20		58	40	9.10599					10.89401	40
П	21	1	1	12		58	48	10697						39
Н	22	l	1	4		58 59	56 4	10795				00360	89205 89107	38 37
П	23 24		0	56 48		59	12	10893				00363	89010	36
Н		١.,	_				20						10.88913	35
H	25 26	11	0	40 32		59 59	28	9.11087 11184	1		10.88548 88449		88816	34
П	27	l	ŏ	24		59	36	11281				00368	88719	3 3
П	28	1	ŏ	16		59	44	11377			88253			32
Н	29	l	Ŏ	8		59	52	11474				00371	88526	31
П	30	11	-0	0	T	0	0	9.11570	9.99627	9.11943	10.88057	10.00373	10.88430	30
H	31	10	59	52		0	8	11666					88334	29
П	32	1	59	44		0		11761	99624				88239	28
П	33		59	36		0	24	11857					88143	27
H	34	L	59	28	_	_0	32	11952					88048	26
П	35	10		20		0	40	9.12047					10.87953	25
П	36	1	59	12	1	0	48	12142					87858	24
H	37 38	l	59 58	4 56	1	0	56 4	12236 12331	9961 <i>5</i> 99613		87379 87283	00385 00387	87764 87669	23 22
Н	39	l	58	48	1	i	12	12425				00388	87575	21
Н	40	10	58	40	ī		20	9.12519						
	41	ľ	58	32	1	1	20 28	12612				10.00390 00392	10.87481 87388	20 19
	42		58	24	1	1	36	12706		1		00393	87294	18
1	43		58	16	1	ī	44	12799					87201	17
۱	44	l	58	8	1	i	52	12892	99603	13289	86711	00397	87108	16
1	45	10	58	0	1	2	0	9.12985	9.99601	9.13384	10.86616	10.00399	10.87015	15
	46		57	52	ł	2	8	13078	99600		86522	00400	86922	14
ı	47	1	57	44	١	2	16	13171	99598		86427	00402	86829	13
1	48	l	57	36		2	24	13263	99596		86333	00404	86737	12
١	49	بيا	57	28	<u> </u>	2	32	13355	99595		86239	00405	86645	11
ı	50	10	57	20	1	2	40	9.13447	9.99593		10.86146		10.86553	10
1	51 52		57 57	12	l	2	48 56	13539 13630	99591	13948 14041	86052	00409	86461 86370	9
1	-53	l	56	56	(3	30	13722	99589 99588	14134	85959 85866	00411 00412	86278	7
١	54		56	48		3	12	13813	99586	14227	85773	90414	86187	6
ł	55	10	56	40	1	-3	20	9.13904	9.99584		10.85680		10.86096	5
-	56			32	-	3	28	13994	99582		85588	00418	86006	4
ı	57		56	24	ŀ	3	36	14085	99581	14504	85496	00419	85915	3
	58		56	16	}	3	44	14175	99579		85403	00421	85825	2
1	59		56	8		3	52	14266	99577	14688	85312	00423	85734	1
ŀ	60		56	0		4	9	14356	99575	14780	85220	00425	85644	0
ı	M	Ho	urp.	.	Ho	ura	. m . l	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secant	Secant.	M

Log. Sines, Tangents and Secants.

8 D	egs	•									Degs. 1	71.
M	Ho	M.A.M.	He)U71	.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	10	<i>5</i> 6 0	1	4	0	9.14356				10.00425	10.85644	60
1		55 52		4	8	14445	99574	14872	85128	00426	85555	59
2		85 44	l	4	16	14535	99572	14963	85037	00428	85465	58
3		<i>55</i> 36	ŀ	4	24	14624	99570	15054	84946	00430	85376	57
4		<i>5</i> 5 2 8		4	32	14714	99568	15145	84 855	00432	85286	56
5	10	55 20	1	4	49	9.14803	9.99566	9.15236	10.84764	10.00434	10.85197	55
. 6		55 12	ŀ	4	48	14891	99565	15327	84673	00435	85109	54
7		55 4		4	56	14980	99563	15417	84583	00437	85020	53 .
8		54 56	ŀ	5	4	15069	99561	15508	84492	00439	84931	52
9	4	54 48	1	5	12	15157	99559	15598	84402	00441	84843	51
10	10	54 40	1	6	20	9.15245	9.99557	9.15688	10.84312	10.00443	10.84755	50
liil		54 32	-	5	28	15333	99556	15777	84223			49
12		54 24		5	36	15421	99554	15867			84579	48
13		54 16	ŀ	5	44	15508	99552	15956			84492	47
14		54 8		5	52	15596	99550	16046	83954		84404	46
1		54 0	1	6	0	9.15683	9.99548			10.00452		45
16		53 52	•	6	8	15770	99546	16224	83776		84230	44
17		63 44	i	6	16	15857	99545	16313	83688	00455	84143	43
18		53 36	l	6	24	15944	99543	16401	83599		84056	42
19		5 3 2 8	l	6	32	16030	99541	16489		00459	83970	41
			┢┰		49						10.83884	
.20			l	6		9.16116	9.99539			10.00461		40
. 21		53 12	I	6	48	16203	99537	16665	83335 83247			39
22 23		53 4 52 56	ľ	6 7	56 4	16 2 89 16 3 74	99535 99533	1675 3 16841	83159	,	8362 6	38 37
23		0% 00 5 % 4 8		7	12	16460	99532			00468	83540	36
			_					16928				
. ~- 1	10		1	7	20	9.16545	9.99530	9.17016		10.00470		35
26		52 32		7	28	16631	99528	17103	82897	00472	83369	34
27		52 24	Ì	7	36	16716	99526	17190			83284	33-
28		5 2 16	ŀ	7	44	16801	99524	17277	82723		83199	32
29		52 8		7	52	16886	99522	17363	82637	00478	83114	31
30	10	5 2 0	1	8	0	9.16970	9.99520	9.17450	10.82550	10.00480		30
31		51 52	1	8	8	17055	99518	17536	82464	00482	82946	29
32		51 44	1	8	16	17139	99517	17622	82378		82861	28
33		51 36		8	24	17223	99515	17708	82292	00485	82777	27
34	- 4	51 28	1	8	32	17307	99513	17794	82206	00487	82693	26
36 4	10	61 20	ī	8	40	9.17391	9.99511	9.17880	10.82120	10.00489	10.82609	25
36		51 12		8	48	17474	99809	17965	82035	00491	825 26	24
37		51 4		8	56	17558	99507	18051	81949	00493	82442	23
38		50 56	٠.	9	4	17641	99505	18136	81864	00495	82359	22
39		50 48		9	12	17724	99503	18221	81779	00497	82276	21
40	10	50 40	1	9	20	9.17807	9.99501	9.18306	10.81694	10.00499	10.82193	20
41		60 32	•	9	28	17890	99499	18391	81609	00501	82110	19
42		50 24		9	36	17973	99497	18475	81525	00503	82027	18
43		50 16		. 9	44	18055	99495	18560	81440	00505	81945	17
44		50 8		9	52	18137	99494	18644	81356	00506	81863	16
		60 0	1	10	-	9.18220	9.99492		10.81272	10.00508	10.81780	15
46		60 0 49 52	'	10	8	18502	99490	18812	81188	00510	81698	14
47		49 02 49 44	l	10	16	18383	99488	18896	81104	00512	81617	13
48		49 36		10	24	18465	99486	18979	81021	00514	81535	19
49		49 2 8		10	32	18547	99484	19063	80937	00516	81453	ii
			-									10
50		49 20	1	10	40	9.18628	9.99462			10.00518 005 2 0	81291	9
51		49 12	ŀ	10	48	18709	99480	19229	80771 80688	00522	81210	8
52		49 4		10	56	18790	99478	19312	80605		81129	7
53		48 56		11	า	- 18871 1895 2	99476 99474	19395 19478	80522	00524	81048	6
54		48 48			12							
55		48 40	1	11	20	9.19033	9.99472		10.80439	10.00528	10.80967	5
56		48 32		11	28	19113	99470	19643	80357	00530	80887	4
57		48 24	1	11	36	19193	99468	19725	80275		80807	3
58		48 16		11	44	19273	99466	19807	80193		80727	2
59		48 8		11	52	19353	99464	19889	80111	00536	80647	1
60		48 0		12	0	19433	99462	19971	, 80029		80567	
M	Hou	IP.M.	Ho	WA	. x .	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secent	Secant.	M

Digitized by G Des 186.

9 D	009	L				-	og. om	o, rang	enus anu	DCCC01113)	Degs. 1	70.
M		NIFA	.м.	Ho	HIP	.м.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.		M
0	'	48	0		12	0	9.19433			10.80029			60
li	10		52	•	12	8	19513	99460		79947	00540	80487	59
2	j	47	44	į		16	19592	99458		79866	00542	80408	58
3	}	47	36	١	12		19672	99456		79784	00544	80328	57
4	1	47	28		12		19751	99454	20297	79703	00546	80249	56
5	10	47	20	1	12	40	9.19830	9.99452	9.20378	10.79622	10.00548	10.80170	55
6		47	12	1	12		19909	99450		79541	00550		54
7	١	47	4		12		19988	99448		79460	00552		53
8	1	46	56	ł	13		20067	99446		79379	00554		52
9		46	48		13	12	20145	99444	20701	79299	00556	79855	51
10	10	46	40	ī	13	20	9.20223	9.99442	9.20782	10.79218	10 00558	10.79777	50
lii	i	46	32	٠	13		20302	99440		79138	00560		49
12		46	24		13		20380				00562	1	48
13		46	16	1	13		20458			78978	00564		47
14	1	46	8		13		20535			78898	00566		46
15	10	46	0	1	14	0	9.20613	9.99432	9 21182	10.78818	10 00569	10.79387	45
16	1.0	45	52	-	14		20691	99429		78739	00571		44
17	1	45	44		14		20768			78659	00573	1	43
18	1	45	36		14		20845			78580	00575	I	42
19	1	45	28	1	14		20922	99423		78501	00577]	41
20	100	45	20	-	14		9.20999					10.79001	40
21	120	45	12	1	14		21076			78343	00581	78924	39
22	١.	45	4	l	14		21153				00585		38
23	1	44	56		15		21229	99415		78186	00585		37
24	ļ	44	48	ł	15		21306	99413			00587		36
25	ro		40	1	15		9.21382	9.99411		10.78029	10.00589		35
26	ï	44	32	١.	15		21458	99409		77951	00591	78542	34
27	1	44	24		15		21534			77873	00593	1	33
28		44	16		15		21610	99404		77795	00596	1	32
29		44	8		15		21685	99402		77717	00598	I —I	31
30	100	44	_ 0	1	16	0	9.21761	9.99400		10.77639		10.78239	30
31	1.0		52		16	8	21836	99398			00602		29
32	ļ	43	44		16		21912				00604		28
33	l	43	36	1	16		21987	99394			09606	1	27
34		43	28		16		22062	99392		77330	00608		26
35	10	43	20	1	16		9.22137	9.99390		10 77953		10.77863	25
36	120	43	12	'	16		22211	99388			00612		24
37		43	4		16		22286	99385		77099	00614		23
38	1	42	56		17	4	22361	99383		77023	00617		22
39		42	48	1	17		22435		23054		00619		21
40	ì	42	40	1	17	20	9.22509			-			20
41	120	42	32	١.	17		22583			76794	00623		19
42	1	42	24		17		22657	99375		76717	00628	1	18
43	١.	42	16		17		22731	99372		76641	00628	1	17
44	1	42	8		Î7		22805			76565	00630		16
45	10	42	_		18		9.22878			10.76490		10.77122	15
46	1		52		18		22952	99366		76414	00634		14
47		41	44		18		23025			76339		1	13
48	1	41	36		18		23098	99362		76263	00638	1	12
49	1	41	28	l	18		23171	99359		76188	00641		11
50	100	41	20	1			9.23244	4				10.76756	10
51	-"	41	12	١.	18		23317	1					9
52	1	41	4	ı		56	23390			75963			8
53	1		56		19		23462						
54			48		19	12	23535	99348	34186	75814	00655	76465	6
55	10	40				20		9.99346				10.76393	5
56	*"		32			28							4
57	1		24			36	23752						3
58	1		16			44							2
59	1	40	8			52						76105	1
60	1	40	ŏ		20		23967	99338					0
	H						Co-sine.	Sine.	Co-tang.	Tangent	Co-secan	Secant.	M
99				-44(<u> </u>	Digitized		Dega.	
33	ve	6ª.						Aa	1	Digitized	-, -0	O. TLERY.	5 0.

10	Degs.					TOE. DIE	cs, I ail	Sciica dire	. Devany	70	Degs. 1	69.
	Hour	.M.	H	our	·M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.		M
-0	10 40	0	1		0	9.23967	9.99335	9.24632	10.75368		10.76033	60
1	39			20	8		99333	24706	75294	00667	75961	
2	39	44 36		20 20	16 24	24110 24181	99331 99328	24779 24853	75 22 1 75147	00669 00672	75890 75819	58
3 4	39 39	28		20	32	24253	99326	24926	75074	00674	75747	57 56
5	10 39	20	ī	20	40	9.24324	9.99324		10.75000		10.75676	55
6	39	12	•	20	48	24395	99322	25073	74927	00678	75605	54
7	59	4		20	56	24466	99319	25146	74854	00681	75534	53
8	38	56		21 21	4	24536	99317	25219	74781	00683	.75464	52
9	38	48	-		12	24607	99315	25292	74708	00685	75393	51
10	10 38 38	40 32	1	21 21	20 28	9.24677 24748	9.99313 99310	9.25365 25437	10.74635 74563	10.00687 00690		50
11 12	38	24		21	36		99308		74490	00692	75252 75182	49 48
13	38	16		21	44	24888	99306		74418	00694	75112	47
14	38	8		21	52	24958	99304	256 55	74345	00696	75042	46
15	10 38	0	1		0		9.99301	9.25727		10.00699	10.74972	45
16	37	5 2		22	8		99299	25799	74201	00701	74902	44
17 18	37 37	44 36		22 22			99 2 97 9 92 94	25871 25943	74129 74057	00703	74832	43
19	37	28		22			99294	26015	73985	00706		42
20	10 37	20	1				9.99290		10.73914	10.00710		40
21	37	12	•	22		25445	99288		73842	00712		39
22	37	4		22	5 6	25514	99285	262 29	73771	00715	74486	
23	36	56		23	4	25583	99283		73699	00717		37
24	36	48	_	23			99281	26372	73628	00719		
25	10 36 36	40 32	1	23 23	20 28	9.25721	9.99278		10.73557		10.74279	35
26 27	36	24		23		25790 25858	99 2 76 99 2 74		73486 73415			34 33
28	36	16		23	44	25927	99271		73345			
29	36	8		23	52	25 995	99269		73274	00731	74005	31
30	10 36	0	1	24	0	9.26063	9.99267	9.26797	10.73203	10.00733	10.73937	30
31	35	52		24			99264		73133	00736	73869	29
32	35	44 36		24 24			99262		73063	00738		28
33 34	35 35			24		26267 26335	99 2 60 99 2 57		7299 2 729 22	00740 00743		27 26
35	10 35	20	1	24	-		9.99255			10.00745		
36	35	12	•	24			99252		72782	00748	10.73597 73530	25 24
37	35	4		24	56	26538	99250		72712			23
38	34	56		25	4	26605	99248		72643	00752	73395	22
39	34	48		25	12	26672	99245		72573			2)
40	10 34	40	1		20		9.99243	9.27496		10.00757		20
41 42	34 34	32 24		25 25		26806 26873	99241 99238		72434 72365			19
43	34	16		25	44	26940	99236		72300	00762 00764		18 17
44	34	8		25	52	27007	99233		72227	00767	72993	16
45	10 34	0	1	26	0	9.27073	9.99231	9.27842	10.72158	10.00769		15
46	33	52		26	8		99229	27911	720 89	00771	72860	14
47	33	44		26	16		99226		72020	00774		13
48 49	33 33	36 28		26 26	24 32	27273 27339	99 22 4 99 22 1	28049 28117	71951 71883	00776		12
50	10 33	20	<u></u>	26	40	9.27405	9.99219		10.71814	00779		11
51	33	12	•	26	48	27471	99219	28254	71746	10.00781 00783		10
52	33	4		26	56	27537	99214		71677	00786	, ,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	8
53		56		27	4	27602	99212	28391	71609	00788	72398	7
- 54		48			12							6
55	10 32		1	27				9.28527			10.72266	5
56 57		32 24			28 36	27799 27864	99204 99202		71405			4
58	32				44	27930	99202		71338 71 27 0			3 2
59	32	8		27		27995	99197	28798	71202			
60	32	<u>oj</u>		28	0	28060	99195	28865	71135	00805	71940	Ö
M	Hour	M.	Ho	WF.A	.M.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secunt	Secent.	M
100 D	egs.									Digitized		<u> </u>
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	11	De	gs.					rog. om	ents and	Secants.	•	Deg. 1	68.	
17	M	_	our	.M.	H	uri	·M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	
H	0	10	32	0	ī	28	0	9.28060	9.99195		10.71135			60
П	1	l		52	\	28	8	28125	99192	28933		00808	71875	59
! !	2		31				16		99190	29000		00810		
Н	3 4	ļ	31 31	36 28	Ì	28 28	24 32	28254 28319	9918 7 99185	29067 29134	70933	00813 00815	71746	57
1 }		1.0		_	 .						70866		71681	56
H	5 6	10	31 31	20 12	1	28 28	40 48	9.28384 28448	9.9918 2 99180	9.29201 29268	10.70799 707 32	00820	71552	55 54
ı	7	•	31	4	ŀ	28	56	28512	99177	29335	70665	00823	71488	53
	8	!	30	56	l	29	4	28577	99175	29402	70598	00825		52
ll	9	1	3 0	48		29	12	28641	99172	29468	7053 2	00828	71359	51
1	10	10	30	40	1	29	20	9.28705	9.99170	9.29535	10.70465	10.00830	10.71295	50
1 1	11	I	3 0	32		29	28	28769	99167	29601	70399	00833	71231	49
11	12		30	24		29	36	28833	99165	29668	70332	00835	71167	48
1 !	13 14	l	30 30	16 8		29 29	44 52	28896	99162	29734	70266 70200	00838 00840	71104	47 46
-		<u> </u>			-			28960	99160	29800			71040	
H	15 16	10	30 29	0 52	1	30 30	0 8	9.29024 29087	9.99157 99155	9.29866 29932	10.70134 70068	10.00843 00845		45 44
11	17		29	44		30		29150	99152	29998	70002	00848	70850	43
П	18	l	29	36	ŀ	30	24	29214	99150	30064	69936	00850	70786	42
П	19]	29	28		30	32	29277	99147	30130	69870	00853	70723	41
ľ	20	10	29	20	1	30	40	9.29340	9.99145	9.30195	10.69805	10.00855	10.70660	40
	21	ł	2 9	12		3 0	48	29403	99142	30261	69739	00858	70597	39
Н	22	1	29	4		30	56	29466	99140	30326	69674	00860	70534	38
Н	23	ł	28	56		31 31	12	29529	99137	30391	69609	00863	70471	37 36
-	24	١	28	48	-		_	29591	99135	30457	69543	00865	70409	
Н	25 26	Ιτο	28 28	40 32	1	31 31	20	9.29654 29716	9.99132 99130	9.30522 30587	10.69478 69413	00870	70284	35 34
П	27		28	24		31	36	29779	99127	30652	69348	00873	70231	33
Н	28		28	16		31	44	29841	99124	30717	69283	00876	70159	32
H	29	1	28	8		31	52	29903	99122	30782	69218	00878	70097	31
	30	10	28	0	1	32	0	9.29966	9.99119	9.30846	10.69154	10.00881	10.70034	30
11	31		27	52		32	8	30028	99117	30911	69089	00883	69972	29
Н	32	1	27	44		32	16	30090	99114	30975	69025	00886	69910	28
Н	33		27	36		32 32	24 32	30151	99112	31040	68960	00888 00891	69849 69787	27 26
-	34	<u> </u>	27	28	-		_	30213	99109	31104	68896			25
	35 36	10	27 27	20 12	1	32 32	.40 48	9.30275 30336	9.99106 99104	9.31168 3123 3	10.68832 68767	00896	69664	24
	37		27	4		32	56	30398	99101	31297	68703	00899	69602	23
	38	l	26	56		33	4	30459	99099	31361	68639	00901	69541	22
	39		26	4 8		33	12	30521	99096	31425	6857 <i>5</i>	00904	6947 9	21
	40	10	26	40	1	33	20	9.30582	9.99093		10.6851 r			20
	41	l		32	l	33	28	30643	99091	31552	68448	00909	69357	19
	42		26	24		33	36	30704	99088	31616	68384 68 32 1	0091 2 00914	69 2 96 69 2 35	18 17
	43 44	1	26 26	16 8		33 33	44 52	30765 308 2 6	99086 9908 3	31679 31743	68257	00917	69174	16
-		10		- °	-		- 52	9.30887	9.99080		10.68194			15
{	45 46	۱,۵	26 25	52	'	34	8	9.30887	99078	31870	68130	00922	69053	14
Н	47	1	25	44		34	16	31008	99075	31933	68067	00925		13
П	48	l	25	36		34	24	31068	99072	31996	68004	60928	68932	12
	49		25	28	L	34	32	31129	99070	32059	67941	00930	68871	11
l ľ	50	10	25	20	1	34	40	9.31189	9.99067		10.67878			10
	51	l	25	12	Ì	34		31250	99064 9906 2	32185 32248		00936 00938	68750 68690	9 8
	52 53	1	25 94	4	ı	34 35		31310 31370	99059				68630	7
	53 54	1		48	l	35	12	31430				: 00944		6
;	55	10	24			35			9.99054		10.67564	10.00946	10.68510	5
	56	٦		32		35	28	31549	99051		67502	00949	68451	4
ĺ	57	1		24		35	36	31609	99048	32561	67439	00952	68391	3
	58	1	24	16	l		44		99046	32623	67377	00954 00957	68331 68272	2 1
-	59	1	24	8		_	52		990 43 990 40	32685 32747	67315 67253	00957	68212	0
Į.	60	<u> _</u>	24	0		36	_0		_ ,	Co. 4077			Secant.	
	M		MLI	· . M.	H	ura	.M.	Co-sine.	ine.	Co-tang.	Digitized by		Ci Degs.	
	101	De	g.								2.9.11200 0		D 1084.	

Log. Sines, Tangents and Secants.

12 De	e 25.					_		,				Degs	. 167.
M.		our	A.M	Н	our	P.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M.
0	10		_			0	9.31788	9.99040		10.67253		10.68212	60
ì	1	23			36		31847	99038	32810				59
2	Ì	23			36		31907	99035	32872				58
3	1	23			36		31966	99032	32933		00968		57
4	_	23		_	36	32	32025	99030	32995	-	00970		56
5	10	23	20	1		40	9.32084	9.99027	9.33057 33119	10.66943 66881	00973		55 54
6	1	23 23	12	ľ	36 36	48 56	32143 32202	99024 99022	33180		00978		53
8	l	22	56		37	4	32261	99019	33242		00981	67739	52
9	l	22	48		37	12	32319	99016	33303		.00984		51
10	10	22	40	1	37	20	9.32378	9.99013	9.33365	10.66635	10.00987	10.67622	50
111		22	32		37	28	32437	99011	33426		00989	67563	49
12	ļ	22	24		37	36	32495	99008	33487	66513	00992	67505	48
13	1	22	16	1	37	44	32553	99005	33548		00995 00998	67 447 67 3 88	47
14	<u> </u>	22	8	-	37	52	32612	99002	33609				
15	10	22	0	1		0	9.32670			10.66330 66269	01003	67272	45
16		21 21	52 44	l	3 8 3 8	8 16	32728 32786	98997 98994	33731 33792		01005	67214	44
18	1	21	36	ĺ	38		32844	98991	33853		01009	67156	42
19	1	21	28		38	32	32902	98989	33913	66087	01011	67098	41
20	10	21	20	1	38	40	9.32960	9.98986	9.33974	10.66026	10.01014	10.67040	40
21	ľ	21	12	ľ	38	48	33018	98983	34034	65966	01017	66982	39
22	ı	21	4		38	56	33075	98980	34095		01020	66925	38
23		20	56		39	4	33133	98978	34155		01022	66867	37
24	L	20	48	L	39	12	33190	98975	34215		01025	66810	36
25	10	20	40	1		20	9.33248	9.98972		10.65724			35
26		20 20	32 24		3 9	28 36	33305 33362	98969 .98967	34336 34396	65664 65604	91031 01033	66695 666 3 8	34 33
27		20 20	16		39	44	33420	98964	34456		01035	66580	32
29		20	.8		39	52	33477	98961	34516	65484	01039	66523	31
30	10	20	0	1		0	9.33534	9.98958		10.65424		10.66466	30
31	1	19	52	^	40	8	33591	98956	34635	65365	01045	66409	29
32	l	19	44		40	16	33647	98953	34695	65305	01047	66353	28
33	l	19	36		40	24	33704	989 <i>5</i> 0	34755	65245	01050	66296	27
34	_	19	28		40	32	33761	98947	34814	65186	01053	66239	26
35	10	19	20	1	40	40	9.33818	9.98944		10.65126		10.66182	25
36		19	12	ı	40 40	48	33874	98941	34933	65067 65008	01059	66126	24
37 38		19 18	56		41	56	33931 33987	98938 98936	34992 35051	64949	01062 01064	66069 66013	23
39		18	48		41	12	34043	98933	35111	64889	01067	65957	21
40	10	18	40	1		20	9.34100	9.98930		10.64830			20
41	- 0	18	32	•	41	28	34156	98927	35229	64771	01073	65844	19
42		18	24		41	36	34212	98924	35288	64712	01076	65788	18
43		18	16		41	44	34268	98921	35347	64653	01079	65732	17
44		18	8		41	52	34324	98919	35405	64595	01081	65676	16
45	10	18	0	1		0	9.34380	9.98916	9.35464		10.01084		15
46		17	52		42	.8	34436	98913	35523	64477	01087	65564	14
47		17 17	36		42 42	16 24	34491 34547	98910 98907	35581 35640	64419 64360	01090 01 09 3	65509 65453	13
49		17	28		42	32	34602	98907	35698	64302	01093	65398	11
استنسا	10	17	20	7	42	40	9.34658	9.98901		10.64243			10
51	-	17	12	•	42	48	34713	98898	35815	64185	01102	65287	9
52		17	4		42	56	34769	98896	35873	64127	01104	65231	8
53		16	56		43	•	34824	98893	35931	64069	01107	65176	7
54	_	16	48		43	12	34879	98890	35989	64011	01110	65121	6
55	10	16	40	1	43	20	9.34984	9.98887	9.36047	10.63953	10.01113	10.65066	5
56		16	32		43	28	34989	98884	36105	63895	01116	65011	4
57			24		43	36	35044	98881	36163	63837	01119	64956	3
58 59		16 16	16		43 43	44 52	35099 35154	98878 9887 <i>5</i>	36221 36279	63779 63721	01122	64901 64846	2
60		16	0		44	3	35209	98872	36336	63664	01128	64791	e i
	Ho			ìγ	OFA.	- T.	Co-sine.					Secant	M
102 D		***			u A			Jule.	CO-MILE.	Tangent.	Co-secant	Secont ,	

102 Degs.

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Log. Sines, Tangents and Secants.

	1	3 I	Deg	3.				- B	,	,			Degs. 166	5.
ī	M	He	our.	M.M	H	ouri	P.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
	0	10	16	0	1	44	0	9.35209	9.98872	9.36336	10.63664	10.01128	10.64791	60
١	1	!	15			44	8			36394	63606	01131	64737	59
I	2	1	15		1	44			93867	36452		01133	64682	58
İ	3 4		15 15			44 44	24 32	35373 35427	- 98864 98861	36509 36566	63491 63434	01136 011 3 9	64627	57
		1.0		20	1								64573	56
	5 6	טנן	15 15	12	1	44 44	40 48	9.35481 35636	9.98858 98855	9.36624 36681	10.63376 63319	01142		55
	7	1	15	4		44	56	35590	98852	36738	63262	01148	64464 64410	54 53
	8	1	14	56		45	4	35644	98849	36795	63205	01151	64356	52
Н	9	l	14	48	Ì	45	12	\$5698	98846	36852	63148	01154	64302	51
	10	10	14	40	1	45	20	9.35752	9.98843	9.36909	10.63091	10.01157	10.64248	50
•	11	1	14	32	i	45	28	35806	98840	36966	63034	01160	64194	49
1	12		14	24		45	36	35860	98837	37023		01163	1	48
l	13		14 14	16	1	45 45	44 52	35914	98834	57080	62920	01166	64086	47
	14	_		-8	Ļ			35968	98831	37137	62863	01169	64032	46
	15 16	10	14 13	0 52	1	46 46	0	9.36022	9.98828		10.62807 62750			45
	17	1	13	52 44	1	46	8 16	36075 36129	98825 98822	37250 37306	62694	01175 01178	63925 63871	44 43
1	18	1	13	36		46	24	36182	98819	37363	62637	01181	6 3 818	42
1	19	l	13	28		46	32	36236	98816	37419	62581	01184	63764	41
!	20	10	13	20	1	46	40	9.36289	9.98813	9.37476	10.62524	10.01187	l	40
1	21		13	12		46	48	36342	98810	37532	62468	01190		39
1	22		13	4		46	56	36395	98807	37588	62412	01193	63605	38
l	23	١.	12	56		47	4	36449	98804	37644	62356	01196	63551	37
	24	_	12	48		47	12	36502	98801	37700	62300	01199	63498	36
İ	25	10	12	40	1	47	20	9.36555	9.98798		10.62244			35
	26 · 27		12 12	32 24	1	47 47	28 36	36608 36660	98 7 95 98 7 92	37812	62188 62132	01205	63392	34
1	23	l	12	16	1	47	44	36713	98739	37868 37924	62076	01 2 08 01211	63340 63287	33 32
1	29		12	8	l	47	52	36766	98786	37980	62020	01214	63234	32 31
1	30	10	12	Ó	1	48	0	9.36819	9.98783	9.38035				30
١	31		11	52	Ī -	48	8	36871	98780	38091	61909	01220		29
}	32	1	11	44	ŀ	4 8	16	36924	98777	38147	61853	01223	63076	28
1	33	l	11	36		48	24	36976	98774	38202	61798	01226	63024	27
	34	_	11	28	_	48	32	37028	98771	38257	61743	01229	62972	26
}	35	10	11	20	1	48	40	9.37081	9.98768	9.38313	10.61687			25
1	36		11	12		48	48	37133	98765	38368	61632	01235	62867	24
1	37		11 10	4 56	ŀ	48 49	56 4	37185 37237	98762 98759	38423	61577 61521	01238	62815	23
1	39		10	48		49	12	37289	98756	38479 38534	61466	01241 01244	62763 62711	22 21
1	40	10	10	40	1	49	20	9.37341	9.98753	9.38589			10.62659	20
١	41	ï	10	32	١.	49	28	37 39 3	98750	38644	61356	01250	62607	20 19
1	42	1	10	24		49	36	37445	98746	38699	61301	01254	62555	18
1	43		10	16		49	44	37497	98743	38754	61246	01257	62503	17
ł	.44	_	10	8	L	49	52	37549	98740	38808	61192	01260	624 51	16
1		10	10	0	1	50	0	9.37600	9.98737		10.61137		10.62400	15
1	46	[9	52	İ	50	8	37652	98734	38918	61082	01266	62348	14
I	47		9	44 36		50 50	16 24	37703 377 5 5	98731 98728	38972 39027	61028 60973	01269	62297	13
١	49		9	28	l	50	32	37806	98725	39027	60918	01272 01275	62245 62194	12 11
	50	10	- <u>5</u>	20	7	50	40	9.37858	9.98722	9.39136			1	
	51	"	9	12	۱ ٔ	50	48	9.37838 37 909	98719	39130 39190	10.60864 60810	01281	62091	10 9
١	52	1	9	4	l	50	56	37960	98715	39245	60755	01285	62040	8
	53	1	8	56	l	51	4	38011	98712	39299	60701	01288	61989	7
	54		8	48	L	51	12	38062	98709	39353	60647	01291	61938	6
1	55	10	8	40	1	51	20	9.38113	9.98706	9.39407	10.60593		10.61887	5
1	56	1	8	32	l	51	28	38164	98703	3961	60539	01297	61836	4
1	57		8	24	l	51	36	38215	98700	39515	60485	01300	61785	3
- 1	58		8	16 8		51 51	44 52	38 2 66 3 8317	98697 98694	39569 39623	60431 60377	01303 01306	61734	2
į	60		8	ő		52	0	38368	98690	39623 39677	60323	01300	61683 61632	1 0
i		Ha			Ha	UFA								_
	1 101	110	W. P	· E · l	***	41.4			Jine.	Co-tang.	- angent.	Co-secant	Secrut.	:71

103 Degs.

Degs. 76.

Log. Sines, Tangents and Secants.

14	Deg						Log. Sine	s, rang	ents and	DCC811112	•	Deg. 1	65.
M			A. W	111	our	P. W	Sine.	Co-sine.	Tangent	Co-tan#	Secant	Co-secant	
0		8	0	-	-		{·			10.60323	10.01310	10 61639	60
1	1.0	7	52		52			98687	39731	60269	01313	61582	59
2	1	7	44		52			98684					
3	1	7	36		52			98681	39838				57
4	_	7	28	L	52			98678			!		
5	10	7	20	1	52	40	9.38620	9.98675				10.61380	
6	ļ	7	12		52			98671	39999	60001			
7 8		7 6	4 56	l	52 53		38721 38771	98668 9866 <i>5</i>	40052 40106	59948 59894			
9		6	48	l	53			98662	40159				
10		6	40	ī			9.38871	9.98659				10.61129	11
11		6	32	*	53		38921	98656	40266	59734			49
12		ĕ	24		53			98652	40319	59681		61029	
13		6	16		53		39021	98649	40372			60979	
14		6	8		53	52	39071	98646	40425	59575	01354	60929	46
15	10	6	Ö	1	54	0	9.39121	9.98643		10.59522	10.01357	10.60879	45
16		5	52		54	8	39170	98640	40531	59469		60830	44
17		5	44		54		39220	98636	40584	59416		60780	43
18		5	36		54	24	39270	98633	40636	59364		60730	42
19		5	28	Ļ	54	32	39319	98630	40689	59311	01370	60681	41
20		5	20	1	_	40	9.39369	9.98627			10.01373		40
21 22	,	5	12		54 54	48 56	39418	98623	40795	59205		60582	39 38
23		5 4	56		55	4	39467 39517	98620 98617	40847 40900	59153 59100		60533 60483	
24		4	48		55	12	3 9566	98614	40952	59048		60434	36
25	_	- -	40	1	55	20	9.39613	9.98610			10.01390		35
26		4	32	•	55	28	39664	98607	41067	58943		60336	34
27		4	24		55	36	39713	98604	41109	58891		60287	33
28		4	16		55	44	39762	98601	41161	58839		60238	32
29		4	8	ı	55	52	39811	98597	41214	58786		60189	31
30	10	4	-0	1	56	0	9.39860	9.98594	9.41266	10.58734	10.01406	10.60140	30
31		3	52		56	8	39909	98591	41318	58682	01409	60091	29
32		3	44		56	16	3 9958	98588	41370	58630		60042	28
33		3	36	!	56	24	40006	98534	41422	58578		59994	27
34		3_	28		5 6	32	40055	98581	41474	58526		59945	26
35		3	20	1	56	40	9.40103	9.98578			10.01422		25
36		3	12		56	48	40152	98574	41578	58422		59848	24
37 38		3 2	4		56 57	56	40200 40249	98571	41629	58371	01429 01432	59800 59751	
39		z 2	56 48		57	4 12	40249	98568 98565	41681 41733	58319 58267	01432	59703	21
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40		2 2	32	I	57 57	20 28	9.40346 40394	9.98561 98558	9.41784 41836	10.58216 58164		10.59654 59606	20 19
42		z 2	3z 24		57	36	40394 40442	98555	41887	58113		59558	18
43		2	16		57	44	40490	98551	41939	58061	01449	59510	17
41		ž	8		57	52	40538	98548	41990	58010		59462	16
		2	-6	ī	58	0	9.40586	9.98545	9.42041	10.57959	10.01455	10.59414	15
46		ĩ	52	-	50	8	40634	98541	42093	57907	01469	59366	14
47		1	44		58	16	40682	98538	42144	57856	01462	59318	13
48		1	36		5 8	24	40730	98535	42195	57805	01465	59 2 70	12
49		1	28		58	32	40778	98531	42246	57754	01469	59222	11
		i	20	1	58	40	9.40825	9.98528			10.01472		10
51		ı	12		58	48	40873	98525	42348	57652	01475	59127	9 1
52		1	4		58	56	40921	98521	42399	57601	01479	59079	8
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		0	40	1	59	20 28	9.41063	9.98511			10.01489 01492	10.58937 58889	5
56 57		0	32 24		59 59	36	41111	98508 98505	42603 42653	57397 57347	01492	58842	3
58		0	16		59 59	44	41205	98501	42704	57 2 96	01499	58795	2
59		Ď	8		59	52	41252	98498	42755	57245	01502	58745	ī
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104 Degs.

Degs. 75

Log. Sines, Tangents and Secants.

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10	_			.M.	Ho	urp.n		Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	·M
1 9 5.69 52 0 8 41347 98491 42856 57144 01509 58653 59 3 9 56 0 24 41441 98484 42957 57043 01516 58566 58 3 9 58 0 32 41488 98481 43907 56939 01619 58656 58 5 9 .59 20 2 0 40 9 .41635 \$98477 9 .45067 10 .56943 10 .01523 10 .58466 56 6 59 12 0 43 41682 98471 43108 56892 01525 58212 56 6 59 12 0 43 41682 98471 43108 56892 01525 58272 53 8 58 56 1 4 41675 98467 43208 56892 01525 58272 53 8 58 46 1 12 41722 98464 43258 56732 01535 58237 53 10 9 .88 40 2 1 20 9 .41768 9 .98467 43208 10 .05692 10 .01501 10 .5823 10 .01523 10 .58468 11 12 58 32 1 128 41816 98457 43208 56692 10 .01501 10 .5823 10 .0151 10 .0151 10 .01523 10 .58468 11 12 58 32 1 128 41816 98457 43208 56692 10 .01501 10 .5823 10 .0151 10 .	H					_		9.41300	9.98494	9.42805	10.57195	10.01506	10.58700	60
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5 9.69 20 2 0 49 9.41530 ⊕84377 9.40071 0.654310 10.5231 10.58456 55 6 9.9 4 0 56 41628 98471 43108 68929 01529 68472 53 8 58 56 56 1 4 41677 98467 43208 56742 01536 582275 53 10 9.88 40 2 120 9.41768 9.98460 9.43308 10.56692 10.1540 10.58232 50 11 58 32 1 52 41816 98450 43336 66592 01.547 58139 48 13 58 16 14 41868 984407 43508 66592 01.547 58139 48 15 9.68 0 2 2 9.94201 9.98440 43507 56933 01.567 10.5799 46 15 9.56	1				ł	0 3	2	41488	98481	43007	56993	01519	58512	56
6			9 59	20	2	0 4	히	9.41535	98477	9.43057	10.56943	10.01523	10.58465	55
8	1				~		- 1			43108	56892	01526		
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9	1	-		56				41675	· 98467		56792			
10	1				1	1 1	2	41722	98464	432 58	56742	01536	58278	51
11	 -			40	2	1 2	ō	9.41768	9.98460	9.43308	10.56692	10.01540	10.58232	50
12	1				~				98457	43358	56642	01543	58185	49
13	ł				ł	1 3	6	41861	98453	43408	56592	01547		
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16 57 52 2 8 42047 98440 43607 56393 01560 57953 44 17 57 44 2 16 42093 98436 43657 56343 01561 57907 43 18 57 36 2 24 42140 98433 43707 56293 01567 57860 42 19 57 28 2 32 42186 98429 43756 56244 01671 57814 41 20 9.57 20 2 2 40 9.42332 98426 9.43966 10.56194 19.01574 10.57768 12 21 57 12 24 42 4278 98422 43855 56145 01678 57722 39 22 57 4 2 56 42324 98419 43905 56095 01681 57676 38 23 56 56 3 4 42370 98415 43955 56095 01681 57676 38 24 56 48 3 12 42416 98412 44004 58996 01588 57584 36 25 9.56 40 2 3 20 9.42461 9.98409 9.44053 10.59947 10.01691 10.5739 35 26 56 56 33 32 842507 98405 44102 55898 01595 57447 33 28 56 16 3 44 42999 99398 44201 55799 01602 57401 32 29 56 8 3 52 42644 98395 44250 55750 01605 57356 31 30 9.56 0 2 4 0 9.42690 9.9391 9.44299 10.55701 10.01609 10.57310 30 31 55 52 4 8 42735 98388 44397 55603 01612 572556 31 32 55 44 4 16 42781 98384 44397 55603 01612 572953 34 55 20 2 4 40 9.42692 98370 44495 55509 01602 57401 32 35 55 52 4 24 4826 98391 44446 55554 01619 57112 27 35 55 52 4 4 4 16 42781 98384 44397 55603 01616 57219 28 35 55 52 4 2 4 48282 98381 44446 55554 01619 57112 27 35 55 52 4 2 4 42826 98391 44495 55603 01616 57219 28 35 55 52 4 2 4 482735 98386 44597 55603 01616 57219 28 36 55 12 4 48 42962 98370 44592 55408 91630 57028 24 37 55 4 4 56 43008 98350 44591 55500 01623 57129 28 38 54 56 5 4 43098 98359 44780 55510 01627 10.7083 25 38 54 56 5 4 43098 98359 44780 55510 01657 56947 12 44 54 8 5 52 8 43188 98352 44384 55652 01641 56902 21 44 54 8 5 56 43333 98344 44981 55019 01664 10568 56902 21 45 54 24 5 56 4 43502 98370 44592 55408 91630 57038 24 45 53 26 6 8 34312 98334 44981 55019 01664 56581 19 45 52 4 7 56 6 83433 98394 44884 55116 01651 56676 18 45 5 52 4 7 4 43969 98391 45505 54681 01669 56643 13 45 5 52 6 7 4 43509 98393 44500 55610 01651 56665 56688 14 45 53 52 6 8 43412 98334 44981 55019 01668 56643 13 45 52 6 7 4 43508 98396 5834 54690 55610 01651 56667 16 45 5 9.52 40 2 7 20 9.43367 9.98373 4559 54681 01669 56643 13 46 53 52 6 7 4 43506 98391 45656 54680 01673 56693 10 50	1		58	8	ĺ	1 5	2	41954	98447	43508	56492	01553	58046	46
16	-	15	9.58	-0	2	2	ō	9.42001	9.98443	9.43558	10.56442	10.01557	10.57999	
17	1		5.57		1				98440	43607	56393			
19	1				l			42093	98436	43 657				
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20	1			28	l	2 3	2	42 186	98429	43756	56244	01571	57 814	
21 57 12 2 48 42278 98422 43865 56145 01678 57722 39 22 57 4 2 56 4224 98419 43905 56095 01581 57676 38 23 56 66 34 42370 98412 44004 58996 01688 57630 37 24 56 48 3 12 42416 98412 44004 58996 01698 57630 37 25 9.66 40 2 3 20 9.42461 9.84402 58988 01695 57493 34 26 56 52 3 28 42507 98405 44102 58898 01695 57493 34 27 56 24 3 36 42559 98398 44201 55898 01695 57493 34 28 56 16 3 44 42599 98398 44201 55799 01602 57401 32 29 56 8 3 52 42644 98395 44250 55750 01605 57356 31 30 9.56 0 2 4 0 9.42690 9.98391 9.44299 10.55701 10.01609 10.57310 30 31 55 52 4 8 42725 98388 44348 55652 01612 57259 38 32 55 54 4 16 42781 98384 44397 55603 01616 57219 28 33 55 56 4 24 42826 98381 44446 55554 01619 57174 27 34 55 28 4 32 42872 98377 44495 55500 10625 57103 26 35 9.55 20 2 4 40 9.42917 9.98373 9.44594 10.55466 10.01637 10.37083 25 36 55 12 4 48 42962 98381 44446 55554 01619 57174 27 38 54 56 6 5 4 4308 98366 44641 55359 01634 57023 26 39 54 48 5 12 43098 98359 44738 55602 01612 57203 26 40 9.54 40 2 5 20 9.43143 9.98356 9.44787 10.50213 10.01644 10.56857 20 41 54 32 5 28 4318 9.8356 9.44787 10.50213 10.01644 10.56857 20 41 54 32 5 28 4318 9.8356 9.44787 10.50213 10.01644 10.56857 20 41 54 32 5 28 4318 9.8356 9.44787 10.50213 10.01644 10.56857 20 41 54 32 5 28 4318 9.8356 9.44787 10.50213 10.01644 10.56857 20 41 54 32 5 28 43182 9.8334 44981 55019 01653 56677 16 45 9.54 0 2 6 0 9.43367 9.98331 44981 55019 01666 56588 14 45 53 56 6 6 4 43680 98313 45126 54874 01669 56581 16 55 9.52 40 7 20 9.43367 9.98331 45367 54681 01669 56583 16 55 9.52 40 7 7 40 43546 98324 44981 55019 01666 56588 14 55 9.52 40 7 7 20 9.43313 9.98302 9.45511 10.54489 10.01662 10.56633 16 56 9.52 40 7 7 20 9.43313 9.98302 9.45511 10.54489 10.01660 10.56187 5656 5658 14 56 9.52 40 7 7 20 9.43367 9.98331 45367 54681 01669 56546 11 56 9.52 40 7 7 20 9.43313 9.98302 9.45511 10.54489 10.01660 10.56187 5656 565 52 32 7 28 43866 98391 45066 45394 01706 56649 10.56187 565 566 52 32 7 28 43866 98391 45654 5459 01698 10.56187 5656 566 52 32 7 28 43890 98388 45702 44366	-		9.57	20	2	2 4	Ιō	9.42232	9.98426	9.43806	10.56194	10.01574	10.57768	
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105 Degs.

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TABLE XXVII.

Logs. Sines, Tangents and Secants.

16	Der	28.				1	ogs. Sin	co, rame	Genra anno	Decanti	•	Degs. 1	63.
			.M.	Ho	urr	.M.	Sine.	Co-sine.	Tangent.	Co-tang-	Secant.	Co-secanti	M
0	9	52	- 0	2	8	0	9.44034	9.98284		10.54250	10.01716	10.55966	60
1		51	52		8	8	44078	98281	45797	54203	01719	55922	59
2	1	51	44	ì		16	44122	98277	45845	54155	01723	55878	58
3	İ	51	36		8	24	44166	98273	45892	54108	01727	55834	57
4	_	51	28		_8	32	44210	98270	45940	54060	01730	55790	56
5	9	51	20	2	8	40	9.44253	9.98266	9.45	10.54013			55
6		51	12		8	48	44297	98262	46035	53965	01738	55703	54
7		51 50	4 56		8	36 4	44341 44385	98259 98255	46082	53918 53870	01741 01745	55659	53 52
8		50 50	48		9	12	44428	98251	46180 46177	53823	01749	55615 55572	51
استسا	_	50	40	2									
10	y	50	32	Z	9	20 28	9.44472 44516	9.98248 98244	9.46224 46271	10.53776 53729	10.01752 01756	55484	50 49
12		50	24		9	36	44559	98240	46319	53681	01760	55441	48
13		50	16		9	44	44602	98237	46366	53634	01763	55398	47
14		50	8		9	52	44646	98233	46413	53587	01767	55354	46
15	9	50	-0	2	10	-0	9.44689	9.98229	9.46460	10.53540	10.01771	10.55311	45
16		49	52	Ī	10	8	44733	98226	46507	53493	01774	55267	44
17	l	49	44		10	16	44776	98222	46554	53446	01778	55224	43
18		49	36	l	10	24	44819	98218	46601	53399	01782	55181	42
19	_	49	28	_	10	32	44862	98215	46648	53352	01785	55138	41
20	9	49	20	2	10	40	9.44905	9.98211	9.46694	10.53306		10.55095	40
21		49	12		10	48	44948	98207	46741	53259	01793	55052	39
22	1	49	4	l	10	56	44992	98204	46788	53212	01796	55008	38
23 24		48	56 48		11	12	45035	98200	46835	53165	01800		37
		48		_			45077	98196	46881	53119	01804	54923	36
25	9	48 48	40 32	2		20	9.45120	9.98192	9.46928			10.54880	35
26 27		48	24		11 11	28 36	45163 45206	98189 98185	46975 47021	53025 52979	01811 0181 <i>5</i>		34 33
28		48	16		11	44	45249	98181	47068	52932	01819		32
29		48	8		ii	52	45292	98177	47114	52886	01823		31
30	9	48	0	2	12	0	9.45334	9.98174	9.47160			10.54666	30
31	1	47	52	~	12	8	45377	98170	47207	52793	01830		29
32	l	47	44	1	12	16	45419	98166	47253	52747	01834		28
33		47	36		12	24	45462	98162	47299	52701	01838		27
34		47	28		12	32	45504	98159	47346	52654	01841	54496	26
35	9	47	20	2	12	40	9.45547	9.98155	9.47392			10.54453	25
36		47	12		12	48	45589	98151	47438		01849	54411	24
37	ŀ	47	4		12	56	45632	98147	47484	52516	01853		23
38 39	l	46	56		13	4 12	45674	98144	47530	52470			22
		46	48	<u> </u>	13		45716	98140	47576	52424	01860		21
40	9	46	40	2	13	20	9.45758	9.98136	9.47622			10.54242	20
41	l	46 46	32 24	l	13 13	28 36	45801 45843	98132 981 2 9	47668 47714	52332 52286	01868 01871	54199 54157	19 18
43	1	46	16	1	13	30 44	45885	98125	47760				17
44	1	46	8		13		45927	98121	47806		01879		16
45	9	46	- ö	2		0	9.45969	9.98117	9.47852				15
46]	45		. ~	14	ě	46011	98113	47897	52103	01887	53989	14
47	ĺ	45		1	14		46053	98110	47943	52057	01890	1	13
48	ĺ	45	36	1	14	24	46095	98106	47989		01894	53905	12
49		45	28	1	14	32	46136	98102	48035	51965	01898	53864	11
50	9	45	20	2	14	40	9.46178	9.98098	9.48080	10.51920	10.01902	10.53822	10
51	1	45		1	14	48	46220	98094	48126	51874	01906		9
52		45	4	ł	14	56	46262	98090	48171	51829	01910		.8
53	1	44		1	15	4	46303	98087	48217	51783	01913		7
54	<u></u>	44		<u> </u>	15	12	46345	98083	48262	51738	01917	53655	6
55	9	44	40	2	15	20	9.46386	9.98079	9.48307			10.53614	5
56		44	32	1	15	28	46428	98075	48353	51647	01925		4
57	L	44	24	1	15 15	36 44	46469	98071	48398		01929 01935		3 2
58	Γ	44	16 8	1	15	52	46511 .4655 2	98067 98063	48443 48489	51557 51511	01933		
60	l	44	Ô		16	0	46594	98060	48534	51466	01940		ò
	H-		·.M.	u.			Co-sine.						M
174	410	HIFF	·= .	ne	MLY	·	CU-sine.	Sine.	CO-tang.	Tangent.	Co-secan	i Secult.	1 571

106 Degs.

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Log. Sines, Tangents and Secants.

	17 I	egs.					•	oog. Ditt	co, raig	ents and	becants	•	Degs.	169
ī	M	Hou	ГА	. M .	Ho	игр	.м.	Sine.	Co-sine.	Tangent	Co-tang	Secant.		M I
ľ	U	9 4	4	0	2	16	Ú	9.46594	9.93060	`			10.53406	60
!	1			52	1	16	8	46635		48579	51421		5 3365	59
1	2			44	l	16	16	46676	98052	48624	51376	01948	5 3324	58
1	3 4		13	36 28	ŀ	16	24	46717	98048	48669	51331	01952	53283	57
ı							32	46758	98044	48714	51286	01956	53242	56
	5		13	20	2	16	40	9.46 00	9.98040		10.51241	10.01960	10.53200	55
1	6 7		13 13	12	ŀ	16 16	48 56	46641	98036	48804	1	01964	53159	54
ll	8			56	1	17	4	4688 2 46923	98032 98029	43849 48894	51151	01968	53118	53
1	9			48		17	12	46964		48939	51106 51061	01971 01975	53077 53036	52 51
1	10	9 4	12	40	2	17	20	9.47005	9.98021		10.51016			50
	11		2	32	_	17	28	47045	98017	49029	50971	01983	52955	50 49
П	12	4	2	24		17	36	47086	98013	49073			52914	48
	13		2	16	1	17	44	47127	98009	49118			52873	47
ł	14		2	8	_	17	52	47168	98005	49 163		01995	52832	46
1	15		2	0	2	18	Q	9.47209			10.50793		10.52791	45
	16			52		18	8	47249	97997	492.52			52751	44
1	17			44		18	16	47290		49296		02007	52710	43
1	18 19			36 28	ŀ	18 18	24 32	47330 47371	97989 97936	49341	50659	02011	52670	42
1	20		_	20	9	18		9.47411		49385		02014	52629	41
1	20		, l , l	12	2	18	40 48	47452	9.97982 9797 8	9.49430 49474	10.50570 50526			40
1	22		,1	4	1	18	56	47492	97974	49519		02022 02026	52548 52508	39
1	23		0	56		19	4	47533	97970	49563	50437	02030	52467	37
1	24		ω	48	l	19	12	47573	97966	49607	50393	02034	52427	36
1	25	9 4	<u>,0</u>	40	2	19	20	9.47613	9.97962	9.49652	10.50348	10 02038	1	35
1	26		ω	32		19	28	47654	97958	49696	50304	02042	52346	34
1	27			24	1	19	3 6	47694	97954	49740		02046	52 306	33
١	28			16	ŀ	19	44	47734	97950	49784		02050	52 266	32
1	29		<u>ю</u>	-8		19	52	47774	97946	49828		02054	52226	31
1	30		Ю	0		20	0	9.47614	9.97942		10.50128			30 '
1	31			52		20 20	8	47854	97938	49916		02062	52146	29
1	32			44 36		20	16 24	47894 47934	97934 97930	49960 50004		02066 02070	52106 52066	28
1	34			28		20	32	47974	97926	50 048			52006 52026	26
1	30	9 3	9	201	2	20	40	9,48014	9.97922				10.51986	25
ł	36		9	12		20	48	480o	97913	501.6	49564	02082	5 1946	24
ì	37		9	4		20	56	48091	97914	50180		0206	51905	23
1	38			56		21	4	43133	97910	50223	49777	02096	51867	22
1	39			48	_	21	12	48173	97906	50267	49733	02094	51627	21
1	40		18	40	2	21	20	9.43213	9.97902	9.50311	10.49659	10.02095		20
1	41		33	32	}	21	28	48252	97898	50355		02102	51746	19
1	42		18 18	24 10	ŀ	21 21	36 44	48292 48332	97894 97890	50398 50442	49602 49558	02106	51700	18 17
	44	-	13	10		21	52	43371	97886	50442 50485		02110 02114	516d8 51629	16
	40		88	$-\frac{0}{0}$	-	22	0	9.48411	9.97882		10.49471			
١	46		17 17	52	۔ ا	22	8	48450		50572		02122	51550	15 14
- 1	47		17	44	l		. 16	48490		50616		02126	51510	13
- 1	48	3	17	36		22	24	48529	97870	5 0659		02130	51171	12
١	49	3	7	28		22	32	48568	97866	50703	49297	02134	51432	11
- 1	50		17	20	2	22	40	9.48607	9.97861	9.50746	10.49254	10.02139	10.51393	10
١	51		7	12		22	48	48647	97857	50789	49211	02143		9
- 1	52		17	4		22	56	48686	97853	50833	49167	02147	51314	8 j
- 1	53 54		16 16	56 48		23 23	12	48725 48764	97849 97845	50876 50919	49124 49081	02151 02155	51275 51236	7 6
- 1			_	40	-	23	20							5
- 1	55 56			32	2	23	20	9.48803 48842	9.97841 97837	9.50962 51005	10.49038 48995	02163	51158	4
- 1	57		-	24		23	36	48881	97833	51003 51048	48952	02167	51119	3
١	58			16		23	44	48920	97829	51092	48908	02171	51060	2
Į	59	3	6	8		23	52	48959	97825	51135	48865	02175	51(141	1
Į	60	3	6	0	_	24	ი	48998	97821	51178	48822	02179	51002	0
1	M	Hou	FP.	M.	Ho	ur.	.M.	Co-sine.	Sine.	Cotang.	Tangent.	Co-secant	Secant.	M

107 Degs.

12 1

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Log. Sines, Tangents and Secants.

18 De	egs .				,	,				s. 161.
M	Houra.m.	Hour	·M·	Sine.	Co-sine.		Co-tang.		Co-secant	
0	9 36 0	2 24	0	9.48998	9.97821				10.51002	
1	35,52	24	. 8		97817	51221			,	
2	35 44		16			51264				58 57
3	35 36 35 28	24 24	24 32	49115 49153	97808 97804	51306 51349	. 48694 48651	02192		56
4		'				1	1			
5	9 35 20		40	1	9.97800		10.48608 48565		10.50808 50769	55 54
6	35 12 35 4	24 24	48 56	49231 49269	97796 97 7 92	51435 51478	48522			53
7 8	34 56		4	49309	97788	51520	48480		50692	
و	34 48	25	12	49347	97784	51563	48437	02216	50653	
10	9 34 40	2 25	20	9.49385	9.97779	9.51606	10.48394	10.02221	10.50615	50
11	34 32	25	28	49424	97775	51648	48352	02225	50576	49
12	34 24	25	36	49462	97771	51691	48309	02229	50538	48
13	34 16	25	44	49500	97767	51734	48266	02233	50500	47
14	34 8	25	52	49539	97 763	51776	48224	02237	50461	46
15	9 34 0	2 26	0	9.49577	9.97759	9.51819	10.48181	10.02241	10.50423	45
16	33 52	26	8	49615	97754	51861	48139		50385	44
17	33 44		16	49654	97750	51903	48097			43
18	33 36		24	49692	97746	51946	48054		50308	42
19	33 28	26		49730	97742	51988	48012	02258		41
20	9 33 20	2 26	40	9.4 9768	9.97738			10.02262		40
21	33 12	26	48	4 98 0 6	97734	52073	47927	02266	50194	39
22	33 4		56	49844	97729	52115	47885		50156	38
23	32 56	27 27	4 12	49882	97725	52157	47843			37
24	32 48			49920	97721	52200	47800			36
25	9 32 40	2 27	20	9.49958	9.97717	9.52242		10.02283		35
26	32 32	27 27	28 36	49996	97713	52284	47716		50004	34
27	32 24 32 16	• 27	44	50034 50072	97708 97704	52326 52368	47674 47632		49966 499 2 8	33
28 2 9	32 8	27	52	50110	97700	52410	47590			31
		2 28						10.02304		
30	9 32 0 31 52	2 28	0 8	9.50148 50185	9.97696 97691	9.52452 52494	47506		49815	30 29
31 32	31 44	28	16	50223	97687	52536	47464			
33	31 36	28	24	50261	97683	52578	47422		49739	27
34	31 28	28	32	50298	97679	52620	47380		49702	
35	9 31 20	2 28	40	9.50336	9.97674	9.52661		10.02326	10 49664	25
36	31 12	28	48	50374	97670	52703	47297			24
37	31 4	28	56	50411	97666	52745	47255			23
38	30 56	29	4	50449	97662	52787	47213		49551	22
39	30 48	29	12	504 86	97657	` 5282 9	47171	02343	49514	21
40	9 30 40	2 29	20	9.50523	9.97653	9.52870	10.47130	10.02347	10.49477	20
41	30 32	29	28	5 0561	97649	52912	47088		49439	19
42	30 24	29	36	50598	97645	52953	47047	02355	49402	18
43	30 16	29	44	50635	97640	52995	47005			17
44	30 8	29	52	50673	97636	53037	46963	02364	49327	16
4.5	9 30 0	2 30	0	9.50710	9.97632	9.53078		10.02368		15
46	29 52	30	8	50747	97628	53120	46880		49253	14
47	29 44	30	16	50784	97623	53161	46839		49216	13
48	29 36	30	24	50821	97619	53202	46798		49179	12
49	29 28	30		50858	97615	53244	46756	02385	49142	11
50	9 29 20	2 30	40	9.50896	9.97610			10.02390		10
51	29 12		48	50933	97606	53327	46673		49067	9
52 53	29 4 28 56	30 31	50	50970 51 00 7	97602 97597	53368	46632		49030	8
53 54	28 48		12	51043	97597 97593	53409 53450	46591 46550	02403		7 6
55	9 28 40	2 31 31		9.51080	9.97589			10.02411		5
56 57	28 32 28 24		36	51117 51154	97584 97580	53533	46467 46426			-
58	28 16	31		51191	97576	53574 53615	46385		48846 48809	3 2
59	28 8	31		51227	97571	53656	46344	02424	48773	î
60	28 0	32	õ	51264	97567	5 3697	46303	02433	48736	i l
M	Hourp.m.				Sine.			Co-secant		Mo
108 D	n ore				~1110.	Jo tang.	+ m12 cut.		December 1	

TABLE XXVII. Long, Sines, Tangents and Secants

												Degs		
M	Н			H.	our	P.M.	Sine.	Co-sine.				Co-secant		
0	9						9.51264	9.97567	9.53697	10.46303	10.02433	10.48736		
1			59		32			97563	63738			48699		
2	1	27			32							4 8662		
3	1	27			32		1					48626		
4	L	27	28	3	32	32	51411	97550	53861	46139	02450	4 8 5 89		
5	5	27	20) 2	32	40	9.51447	9.97545	9.53902	10.46098	10.02455	10.48553		
6	1	27	12	2	32	48	51484	97541						
7	1	27		ŀ)	32	56	51520	97536	53 984	46016	02464	48480		
8	1	20			33					45975	02468	48443		
9		26	48	3	33	12	51593	97528	54065	45935	02472	48407		
10	- 3	26	44	5	33	20	9.51629	9.97523	9.54106	10.45894	10.02477	10.48371		
11	1	26	35	2	33	28	51666	97519				48331		
12		26	24	ı.	33	36	51702	97518						
13	1	26	16	i	33	44	5173 8					48262		
14	1	26	; {	3	33	52	51774	97506	54 269	45731	02494	48226		
15	9	26	. (2	34	0	9.51811			10 45691	10.02499			
16	1	25			34									
17		25			34									
18	1	23			34									
19	1	2.				32								
20	17	25		-'	34				·		10.02521	l		
21	1	25			34				1					
22		25			34									
23	1	24			35									
24	1	21			35									
	-			-	35									
25	۱۶	24				28			9.54714		10.02543	10.47829		
26	ł	24 24			35									
27	1	24			35									
28 29	1	24			35									
	- -			.)					-l	.1				
30	9	L 24			36	0						10.47650		
31	1	23			36	8								
32	1	23			36				1					
33		23			36	24								
34	_		28	'	36	32		97417						
35	9	23				40		9.97412			10 02588			
36			12		36	48	52563	97408	,					
37	1	23	_		36		52598	97403	55195			47 402		
38	1	22			37	4		97399				473 66		
39		22			37	12	52669	97394	1		I	47331		
40	9	22			37	20		9.97390		10.44685	10.02610	10.47295		
41	1		32	1	37			97385	5 5355			47 260		
42	1		24		,37	3 6		97381	55395					
43	1	22			37	44		97376	1			471 89		
44	L	22	- 8	_	37	52	52 846	97372	55474	44526	02 628	47154		
45	9	22	ō	2	38	0	9.52881	9.97367	9.55514	10.44486	10.02633	10.47119		
46	1	21	52		38	8	52916	97363	55554			47084		
47	1	21	44	ł		16	52951	97358	55593	44407		4704 9		
48	1	21	36		3 8	24	5 2986	97353	55633	44367	02647	47014		
49	1	21	2 8		3 8	32	53021	97349	55673	44327	02651	46979		
50	9	21	20	2	38	40	9.53056	9.97344	9.55712	10.44288	10.02656	10.46944		
51	١		12	~	38		53092	97340	55752	44248	02660	46908		
52		21	4	1		56	53126	97335				46874		
53			56	1	39	4	53161	97331	55831	44169	02669	46839		
54	1		48		39		53196	97326	55870	44130	02674	46804		
55	0	20	40	9	39		9.53231	9.97322			10.02678			
56	-		32	~	39		53266	97317	5.55910 55949	44051	02683	46734		
57	1		24		39		53301	97312	55989	44011	02688	4 669 9		
84	1	20			39		53 336	97308	56028	43972	02692	46664		
59		20	8		39		53370	97303	56067	43933	02697	4 663 0		
60	l	20	ŏ	i	40	ď	53405	97299	56107	43893	02761	4659 5		
	u.			HA			Co-sine.							
	no		·	110	UFA.	<u></u>	CO-MING.	Sine.	W-LEUE.	∓ enken(·)	Co-secant	Secant.		

M Hourp.m. Houra.m. Co-sine. 109 Degs.

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Log. Sines, Tangents and Secants.

M I 2 3 4 5 6 7 8 9	9	20 19 19 19	0 52 44	2	игр. 40 40	M. 0	9.53405	Co-sine. 9.97299	Tangent.	Co-tang. 10.43::93	Secant. (M 60
1 2 3 4 5 6 7 8	9	19 19 19	52 44	_				9.97299	9 56107	10.43::93	10.02701	0 46595	60
2 3 4 5 6 7 8	9	19 19	44		40	al		J. J. ~ ~ J.				10.70000	
3 4 5 6 7 8	9	19	-			81	. 53440	97294	56146	43854	02706	46560	59 i
5 6 7 8	9				40	16	53475	97289	561 85	43815	02711	46525	58
5 6 7 8	9	19	36		4 0		53509	97235	56224	43776	02715	46491	57
6 7 8	-		23		4()	J2	53544	97280	56264	43736	02720	46456	56
7 8		19	20	2	40	40	9.53578	9.97276	9.56303	10.43697	10.02724	10.46422	55
8		19	12		40	48	53613	97271	56342	43658	02729	46387	54
		19	4		40	56	53647	97266	56381	43619	02734	46353	53
9			56		41	4	53682	97262	56420	43580	02738	46318	52
		18	4 3		41	12	53716	97257	56459	43541	02743	46284	51
10	9	18	40	2	41	20	9.53751	9.97252	9.56498	10.43502	10.02745	10.46249	50
11		18	32		41	28	53785	97248	56537	43463	02752	46215	49
12		18	24		41	36	53819	97243	56576	43424	02757	46181	48
13		18	16		41	44	53854	97238	56613	43305	02762	46146	47
14		18	8		41	52	53 888	97234	56654	43346	02766	46112	46
15	9	18	0	2	42	0	9.53922	9.97229	9.56693	10.43307	10.02771	10.46073	45
16		17	52	ì	42	8	5 3957	97224	56732	43268	0277 6	46043	44
17		17	44	1	42	16	53991	97220	56771	43229	02780	46009	43
18		17	3 6	1	42	24	54025	97215	56810	43190	02785	45975	42
19	}	17	2 3	١	42	32	540 59	97210	56849	43151	02790	45941	41
20	9	17	20	2	42	40	9.54093	9.97206		10.43113			40
21	1	17	12	1	42	48	54127	97201	ა69 2 6		02799	45873	
22		17	4	ļ	42	5 6	54161	97196			02904	45839	
23		16	56	1	43	4	541 95	97199	57004		02808	45805	
21		16	48	١.	43	12	542 29	97187	57042	I	02813	45771	36
25	9	16	40	2	43	20	9.54263	9.97182		10.429 9			35
26		16	32	l	43	28	54297	97178			02822	45703	
27		16	24		43		5 4331	97173			02827	45669	
28		16	16	1	43	44	54365	97168				45635	32
29	1	16	8		43	52	5439 9	97163				45601	31
30	9	16	-0	2	44	U	9.54453				10.02841		30
31		15	52	1	44		544 66						29
32	,	15			44		54 500						
33	1	15			44		54534				02855		27
34		15	28	_	44		54567	97140					26
კა	9	15				40	9.54601	9.97135			10.02865		25
36		15			44		54635						24
37		15			44		54668				02874		23
3 8		14			45		54702		57581				
39	<u>نــا</u>	14		-1-	45		54735				J. —		21
40	9						9.54769				10.02889		20
41		14			45					1			
42	1	14		٠.	45								
43	!	14			45								
-44	_	14	_	_	45		·						_
45	9											10.45064	
46	1	13			46								
47	1	13		- 1	46								
48	1	13 13			46								
49	-			-1-	46					-			
50	9					-				10.4196		10.4489	
51	1	13			46								
52		13		4	40	-							
53	1	19	-		47								
_54	-	15		_	47		1		-1	-			
ر.ن	9				47					10.4177		10.4473	
56	1	15			47								
57		12		-,	47								
58	1	12	_		47								
59 60	1	12		B) Di	48						0298		
	1.												
M			P.M	.ĮH	our	A.M.	Co-sine.	Sine.	· Co-tang	Tangent	- Co-secan	Secant.	M

110 Degs.

Dan 89

Log. Sines, Tangents and Secants.

21	De	gs.							•			Degs. 1	58.
M		OUFA	.м.	Ho	uri	·M·	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	9	12	0	2	48	0	9.55433	9.97015	9.58418	10.41582	10.02985	10.44567	60
1	1	11	52	ļ	48	8	55466	97010	58455		02990	44534	59
2	1	11	44	1	48	16	55499	97005	58493		02995	44501	58
3		11	36	l	48	24	55532	97001	58531	41469	02999	44468	57
4	_	11	28		48	32	55564	96996	58569	41431	03004	44436	56
5	9	11	20	2	48	40	9.55597	9.96991		10.41394			55
6	l	11	12		48	48	55630	96986	58644	41356	03014 03019	44370 44337	54 53
7 8		11	4 56		48 49	56	55663 55693	96981 96976	58681 58719	41319 41281	03019	44305	52
9		10 10	48		49	12	55728	96971	58757	41243	03029	44272	51
	<u> </u>			2	49	20		9.96966		10.41206			50
10	9	10	40 32	2	49	28	9.55 7 61 55 7 93	96962	58832	41168	03038	44207	49
12		10	24		49	36	55826	96957	58869	41131	03043	44174	48
13	1	10	16		49	44	55858	96952	58907	41093	03048	44142	47
14	1	10	8		49	52	55891	96947	58944	41056	03053	44109	46
15	9	10	0	2	50	0	9.55923	9.96942	9.58981	10.41019	10.03058	10.44077	45
16	٦	.9	52	٦	50	8	55956	96937	59019	40981	03 063	44044	44
17		9	44		50	16	55988	96932	59056	40944	03068	44012	43
18		9	36	ŀ	50	24	56021	96927	59094	40906	03073	43979	42
19	ł	9	28		. 50	32	56053	96922	59131	40869	03078	43947	41
20	9	9	20	2	50	40	9.56085	9.96917	9.59168			10.43915	40
21		9	12	l	50	48	56118	96912	59205		03088	43 882	39
22		9	4		50	56	56150	96907	59243		03093	43850	38
23	}	8	56		51	4	56182	96903	59280	40720	03097	43818	37
24	_	_8	48		51	12	56215	96898	59317	40685	03102	43785	36
25	9	8	40	2	51	20	9.56247	9.96893	9.59354		10.03107	10.43753	35
26	ŀ	8	32		51	28	56279	96888	59391	40609	03112	43721	34
27		8	24		51	36	56311	96883	59429	40571 40534	03117 03122	43689 43657	33 32
28 29		8	16 8		51 51	44 52	56343 56375	96878 96873	59466 59503	40334	03122	43625	31
	_		_	_		_							30
30	9	8	0	2	52	0	9.56408	9.96868 9686 3	9.59540 59577	40423	10.03132 03137	10.43592 43560	29
31 32		7	52 44	l	52 52.	8 16	56440 56472	96858	59614	40386	03142	43528	28
33		7	36	1	52	24	56504	96853	59651	40349	03147	43496	27
34		7	28	ŀ	52	32	56536	96848	5 9688	40312	03152	43464	26
35	9	7	20	2	52	40	9.56568	9.96843	9.59725			10.43432	25
36	3	7	12	1 ~	52	48	56599	96838	59762	40238	03162	43401	24
37		7	4	ŀ	52	56	56631	96833	59799	40201	03167	43 369	23
38		6	56		53	4	56663	96828	59835	40165	03172	43337	22
39		6	48	1	53	12	56695	96823	59872	40128	03177	43305	21
40	9	6	40	2	53	20	9.56727	9.96818	9.59909	10.40091	10.03182	10.43273	20
41	_	6	32	ľ	53	28	56759	96813	59946	40054	03187	43241	19
42		6	24	1	53	36	56790	96808	59983	40017	03192	43210	18
43		6	16	ı	53	44	56822	96803	60019	39981	03197	43178	17
44		6	8		53	52	56854	96798	60056	39944	03202	43146	16
45	9	6	0	2	54	0	9.56886	9.96793		10.39907	10.03207	10.43114	15
46		5	52	ı	54	8	56917	96788	60130	39870	03212	43083	14
47		5	44	Ī	54	16	56949	96783	60166	39834	03217 03222	43051	13
48		5	36 28	l	54	24	56980 57019	96778 96772	60 2 03 60 2 40	39797 39760	03222 032 2 8	43020 42988	12
49	_	_5_		<u> </u>	54	32	57012						
50	9	5	20	2	54	40	9.57044	9.96767 96762	9.60276 60313		10.03233 03238	10.42956 42925	10 · 9
51 52		5 5	12	1	54 54	48 56	57075 57107	96757	60349	39687 39651	03238	42893	8
53		4	56	l	55	4	57138	96752	60386	39614	03248	42 862	7
54		4	48	l	55	12	57169	96747	60422	39578	03253	42831	6
55	9	4	40	2	55	20	9.57201	9.96742	9.60459			10.42799	5
56	-		32	~	55	28	57232	96737	60495	39505	03263	42768	4
57		4	24		55	36	57264	96732	60532	39468	03268	42736	3
58		4	16		55	44	57295	96727	605 68	39432	03273	42705	2
59		4	8	1	55	52	57326	96722	60605	39395	03278	42674	1
60 l		4	0		56	O	573 58	96717	60641	9359	03283	426 42	0

206

TABLE XXVII. Log. Sines, Tangents and Secants.

22 De	egs.	•	Dog. Dilk	, - u ₆				Degs	
M	Houra.m.	Ноигр.м.	Sine	Co-sine.	Tangent.	Co-tang	Secant.	Co-secant	M
0	9 4 0	2 56 0	9.57358	9.96717				10.42642	
1	3 52	56 8	57389	96711	60677			42611	59
2	3 44	56 16	57420	96706	60714		03294 03299	42580	58
3 4	3 36 3 28	56 24 56 32	57451 57482	96701 96696	60750 60786	39250 3 9214	03299	42 549 42 518	57 56
1	9 3 20	2 56 40	9.57514	9.96691			10.03309		55
5	3 12	56 48	57545	96686	60859	39141	03314	42455	54
7	3 4	56 56	57576	96681	60895	39105	03319	42424	53 !
8	2 56	57 4	57607	96676	60931	39069	03324	42393	52
9	2 48	57 12	5763 8	96670	60967	39 033	03330	42 362	51
10	9 2 40	2 57 20	9.57669	9.96665	9.61004	10,38996	10.03335	10.42331	50
11	2 32	57 28	57700	96660	61040	38960	03340	42300	49
12	2 24 2 16	57 36 57 44	57731	96655	61076	38924 38888	03345 03 350	42269 42238	48
13	2 16 2 8	57 44 57 52	577 62 577 93	96650 966 4 5	61112 61148	38852	03355	42207	46
15	9 2 0	2 58 0	9.57824	9.96640			10.03360		45
16	1 52	58 8	57855	96634	61220	38780	0 3366	42145	44
17	1 44		57 885	96629	61256	38744	03371	42115	43
18	1 36	58 24	57916	96624	61292	38708		42084	42
19	1 28	58 32	57947	96619	61328	38672	03381	42053	41
20	9 1 20	2 58 40	9.57978	9.96614			10.03386	10.42022	40
21	1 12	58 48	5 8008	96608	61400	38600	05392	41992	39
22 23	1 4 0 56	58 56 59 4	58039 58070	96603 96598	61436 61472	38564 38528	03397 03402	41961 41930	38
24	0 48	59 4 59 12	58101	96593	61508	38492	03407	41899	36
25	9 0 40	2 59 20	9.58131	9.96588			10.03412		35
26	. 0 32	59 28	58162	96582	61579	38421	03418	41838	34
27	0 24	59 36	58192	96577	61615	38385	03423	41808	33
28	0 16	59 44	58223	96572	61651	38349	03428	41777	32
29	0 8	59 52	58253	96567	61687	38313	03433	41747	31
30	9 0 0	3 0 0	9.58284	9.96562			10.03438		30
31	8 59 52	0 8	58314	96556	61758 61794	38242 38206	03444 03449	41686	29 28
32 33	59 44 59 36	0 16 0 24	58345 58375	96551 96546	61830	38170	03454	41655 41625	27
34	59 28	0 32	58406	96541	61865	38135	03459	41594	26
35	8 59 20	3 0 40	9.58436	9.96535				10.41564	25
36	59 12	0 48	58467	96530	61936	38064	03470	41533	24
37	59 4	0 56	5 849 7	96525	61972	38028	03475	41503	23
38	58 56	1 4	58527	96520	62008	37992	03480	41473	22
39	58 48	1 12	58557	96514	62043	37957	03486	41443	
40	8 58 40	3 1 20	9.58588	9.96509	9.62079 62114		10.03491		20 19
41	58 32 58 24	1 28 1 36	58618 58648	96504 96498	62114	37 886 3785 0	03496 03502	41382 41352	18
43	58 16	1 44	5 8678	96493	62185	37815	03502	41322	17
44	58 8	1 52	58709	96488	62221	37779	03512	41291	16
45	8 58 0	3 2 0	9.58739	9.96483	9.62256	10.37744	10.03517	10.41261	15
46	57 52	2 8	5 8769	96477	62292	37708	03523	41231	14
47	57 44	2 16	58799	96472	62327	37673	03528	41201	13
48	57 36 57 28	2 24 2 32	58829	96467 96461	62362 62398	37638 37602	03533	41171 41141	12
49			58859				03539		
50 51	8 57 20 57 12	3 2 40 2 48	9.58889 58919	9.96456 96451	9.62433 62468	10.37567 375 32	10.03544 03549	10.41111 41081	10
52.	57 4	2 56	58949	96445	62504	37496	03555	41051	8
53	56 56	3 4	58979	96440	62539	37461	03560	41021	7
54	56 48	3 12	59009	96435	62574	37426	03565	40991	6
55	8 56 40	3 3 20	9.59039	9.96429	9.62609	10.37391	10.03571	10.40961	5
56	56 32	3 28	59069	96424	62645	37355	03576	40931	4 1
57	56 24	3 36	59098	96419	62680	37320	03581	40902	3 [
58 59	56 16 56 8	3 44 3 52	59128 59158	96413 96408	62715 62750	37 2 85 372 50	03587 03592	40872 40842	1
60	56 0	4 0	59188	96408	62785	37215	03592	40812	6
	Hourp.m.		Co-sine.				Co-secant		1
		-AUUITA-A	CO-MIE.	DINE.	- Counting.	- and chr	- sociality	100	CLTC

112 Degs.

TABLE XXVII.

9	23 De	23.							,				Degs.	156.,
<u> </u>	M		ura	.M.	Ho	urp	.w.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
1	0	8	56	0	3	4	0	9.59188	9.96403		10.37215			60
ļ	1	ļ	55			4	8	59218	96397	62820	37180		40782	59
- 1	2 3		55 55	44 36			16 24	59247 59277	9639 2 96387	62855 62890	37145 37110	03608 03613	40753 40723	58 57
-	4			28		4	32	593 07	96381	62926	37074	03619	40693	56
-	5		55	20	3	4	40	9.59336	9.96376		10.37039	10.03621	10.40664	55
- 1	6	"	55	12		4	48	59366	96370	62996	37004	03630	40634	54
1	7	ì	55	4		4	56	593 96	96 3 65	63031	3 6969	03635	40604	53
١	8	1		56		5	4	59425	96360	63066	3 6934	03640	40575	52
-	9	L	54	48		5	12	59455	96354	63101	36899	03646	40545	51
İ	10	8	54	40	3	5	20	9.59484		9.63135		03651 03657	10.40516, 40486	50
	11	ı	54 54	32 24		5 5	28 36	59514 59543	963 43 963 3 8	63170 63205	36830 36795		40457	49 43
1	12 13		54	16		5	44	59573		63240	36760	1		47
	14		54	8		5	52	59602		63275	36725			46
	15	8	54	-0	3	6	0	9.59632	9.96322	9.63310	10.36690	10.03678	10.40368	45
	16	Ĭ	53	52	Ī	6	8	59661	96316	63345	3 6655	03684	40339	44
	17	1	53		l	6	16	59690		63379	36621	03689	- 1	43
	18	1		36	1	6	24 32	59720 50740		63414 63449	36586 36551	03695 03700		42 41
ŀ	19	<u> </u> _	53			6		59749						
	20	8	53 53	20 12	3	6 6	40 48	9.59778 59808		9.63484	10.36516 36481	03711	10.40222 40192	40 39
1	21 22	1	53	1Z 4	1	6	56	59837		63553				38
l	23	1	52	56		7	4	5 9866		63588				37
1	24	1	52	48	1	7	12	59895		63623			40105	36
1	25	8	52	40	3	7	20	9.59924	9.96267	9.63657	10.36343		10.40076	35
١	26	ł	52	32		7	28	59954		63692				34
	27	1	52	24		7	36	59983						33
1	28	1	. 52	16		7 7	44 52	60012	1	63761 63796				32 31
	29	-	52	- <u>8</u>	<u> </u>		_	60041					10.39930	30
	30	8	52 51	0 52	3	8	0 8	9.60070			10.36170 36135			29
1	32	1	51	44		8	16	60128		63899				28
	33	1	51	36	ŀ	8	24	60157				03777	39843	27
1	34	١.	51	28		8	32	60186	96218	63968	36032			26
L	35	8	51	20	3	8	40	9.60215			10.35997		10.39785	25
1	36	١.	51	12	1	8	48	60244		64037				24
1	37	İ	51	4		8	56 4	60273		64072				23 22
1	38 39		50 50	_		9	12	6030 2 60331		64106 64140				21
١	40	8		40	3	9	20	9.60359			10.35825		10.39641	20
	41	ľ°	50			9	28	60388	1					19
	42		50			9	36	60417						18
1	43		50	16	1	9	44	60446	96168	64278	35722	0383	39554	17
	44	L	50	8	1_	9	52	60474						16
	45	8	50	0	1 -		0	9.60503			10.35654		10.39497	15
١	46		49	-	1	10	8	60532				1		14
1	47		49 49	44 36		10 10	16 24	60561 60589						13 12
	49		49	28		10	32	60618	1					11
	50	8		20			40	9.60646			10.35485		10.39354	10
-	51	1	49			10	48	60675						9
1	52		49	4		10	56	60704						8
1	53			56			4							7
1	54	-		48			12				J			6
	55	8	48				20	9.60789			10.35315		10.39211	
	56			32 24			28 36	60818 60846						3
١	57	1		16			44	60875						2
	59		48		1		52	60903						ĩ
- [60		48			12		60931						0
	M	He	ouri	·.м.	He	ur	.M.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secan	Secant.	M
	113 1										→ *Digil	zea by 😉	O Deg	a. 66.
		9											•	

Log. Sines, Tangents and Secants.

	_					I	Log. Sine	s, Tang	ents and	Secants	•	Dom 1	e e	1
	Deg						<u> </u>	<u> </u>	T 1	Catanal	Secont	Degs. 1: Co-secant	M	-1
1	-				urr	<u> </u>			rangent.	10.35142	10 00007	10. 900cul	;	1
0 1	8	48	0 52		12 12	9		9.96073 96067	64892		03933		60 59	ı
2	İ	47				16		96062	64926		03938		58	1
3		47	36		12		61016	96056	64 960	35040	03944	38984	57	1
4		47	28		12	32	61045	96050	64994	35006	03950	38955	56	1
5	8	47	20	3	12	40	9.61073			10.34972			55	1
6	ŧ	47	12		12	48	61101	96039	65062 65096	34938 34904	03961 0 3966	38899 38871	54 53	1
7 8	1	47 46	4 56		12 13	56 4	61129 61158	96034 96028	65130	34870	03972	38842	52	1
9		46	48			12	61186	96022	65161	34836	03978	38814	51	1
0	8	46	40	3			9.61214	9.96017	9.65197	10.34803	10.03983	10.38786	50	ı
ì	-		32	-		28	61242	96011	65231	347 69	03989	387 58	49	ŀ
2	l		24			36	61270	96005	65265	34735	03995	38730	48	1
3	1	46				44	61298	96000 95994	65299 65333	34701 34667	04000 04006	38702 38674	47	1
4	<u> _</u>	46	8	Ļ	13		61326			10.34634			45	. 1
5 6	8	46	0 52	3	14 14	0 8	9.61354 61382	9.95988 95982	65400	34600		38618	44	ı
7	1	45	44		14		61411	95977	65434	34566		38589	43	١
8			3 6		14		61438	95971	65467	3 4533			42	١
9	}	45	8 8		14	32	614 66	95965	65501	344 99	04035	38534	41	1
0	3	45	20	3	14	40	9.61494				10.04040	10.38506	40	ı
1	1		12		14		61522	95954	65568	34432	04046		39 38	
2	1	45	4		14 15	56 4	61550 61578	95948 95942	65602 65636	34398 34364	04052 04058	1	37	
3 4		44	56 48	ĺ	15		61606	95942 95937	65669	34331	04063		36	ı
5	Ω	44	40	3		20	9.61634	9.95931		10.34297		10.38366	35	
6	ľ	44	32	١	15		61662	95925	65736	34264			34	
7		44	24			36	61689	95920	65770	34 230			35	
8	Ì	44	16		15	44	61717	95914	65803		04086		32	
9		44	8			52	61745	95908	65837	34163		38255	31	
0	8	44	0	3	16	0	9.61773	9.95902				10.38227 38200	30 29	
1 2	1	43 43	52 44	1	16 16	8 16	61800 61828	9589 7 95891	65904 65937	34096 34063			28	ı
3	l		36	İ	16			95885	65971	34029			27	l
4	1	43	28		16	32		95879	66004	33996		38117	2 6	
5	8	43	20	3	16	40	9.61911	9.95873	9.66038	10.33962	10.04127		25	
5	1		12			48	61939	95368	66071	33929			2:	l
7		43	4		16	56	61966	95862	66104				23 22	; ;
3		42 42	56 48		17 17	12	61994 62021	95856 95850	66138 66171	33862 33829				i
5	-			-	17	20	9.62049	9.95844				10.37951	20	
1	8	42 42	40 32	٦	17	28	62076	95839	66238					
2	1	42			17		62104	95833	1	33729		37896	18	
3	ļ	42			17	44	62131	95827	6 630 4	3 3696	04173			
4		42	8			52	62159	95821	66337				16	!
5	3	42	0	3	18			9.95815				10.37H14	15	1
6	1	41	52		18	8	62214	95810						
7	1	41 41	44 36		18	16 24	62241 62268	95804 95798	66437 66470		1			١
9	l	41	28	٠	18	32	62296	95792	66503					
5	8	41	20	3	18	40	9.62323	9.95786				10.37677	10	ĺ
ì	۱		12	Ĭ		48	62350	95780			04220	37650	9	
2		41	4			56	62377	95775			1			1
3			56			4		95769						1
4			48	_		12		95763				1		ł
5	3	40		3	19		9.62459		9.66702 66735			10.37641 37514	5	
5 7	1		32 24			28 36	62486 62513	95751 95745						1
3			16			44		95739		1		37459	2	İ
3		4 0	8		19	52	62568	95733	66834	33166	04267			1
)	_	40	0		20	'		95728	66867	' 				1
	Ho	urr	м.	Ho	ura	.м.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secant	Secant.	M	1

14 Degs.

Digitized by Per 65glC

25	De			_							. -	Degs. 1	54.
M	Н	oura	.М.	-		.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	8	40	0		20	0		9.95728	9.66867	10.33133	10.04272	10.37405	60
1			521		20	8	62622	95722	66900	33100			59
2	1	39 39	36	1	20 20		62649	95716	66933				58
3		39 39	28)	20 20	24 32	62676 62703	95710 95704	66966		04290		٠.
	 -	39	20	<u>_</u>					66999		04296		56
6	18	39	12	3	20	40 48	9.62750 62757		9.67032 67065	10.32968 32935			55
7		39	4	Ī	20	56	62784						10.0
8		38	56	1	21	4	62811	95680		32869	04314		
9		38	48	1	21	12	62838				04326		
10	8	38	40	3	21	20	9.62865					10.37135	1
11	1	38	32	1	21	28	62892	95663	67229	32771	04337	37108	49
12		38	24	ļ	21	36	62918	95657	67262	32738	04343	37082	48
13		38 38	16	ļ	21		62945		67295				47
14	+	38	8		21	52	62972						
15 16	8	38	0 52		22	0	9.62999			10.32640			45
16	1	57 57	52 44		22 22	8 16	63026 63052		67393 67426				
18	1	37 37	36	١	22	16 24	63052		67426 67458				,
19	1	37	28	I		32	63106		67491	32542 32509			
20	P	37	20	3		40	9.63133			10.32476			40
21	ً ا	37	12		22	48	63159	95603			04397		
22	1	37	4	I	22	56	63186	95597	67589	32411	04403	36814	38
23		36	56	1	23	4	63213	95591	67622	32378	04409	367 87	37
24	-	36	48	ب	23	12	63239					·	1
25 96	8	36	40	3		20	9.63266					10.36734	
26 27	1	36 36	32 24	ļ	23 23	28 36	63292 63319				04427		
27 28	1		16	ļ	23	36 44	63319 63345		6775 2 67785				
29	l	36	8	1	23 23	52	63372						
30	8		0	3	24	0	9.63398			10.32150			
31	۱	35	52	-	24	8	63425						
32	1	35	44	ļ	24	16	63451	95537	67915		04463	36549	
33		35	36	Ī		24	63478	95531	67947	32 053	04469	36522	27
34	<u> </u>		28			32		I					~ -
35	8	35	20			40						10.36469	
36 37	1	35 35	12	ļ	24 24	48					04487		
38		34	56	Ì	24 25	56 4	63583 63610		68077 68109		04493 04500		
39		34	48	I	25	12	63636		68109 68142		04500		
40	Ж	34	40	3		20						10.36338	
41	•	34	32	1	25	28	63689		68206				
42		34	24	1	2 5	36	63715	95476	68239	31761	04524	36285	18
43		34	16	ı	25	44	63741	95470	68271	31729	04530	36259	17
44		34	-8		25	52	63767		68303		04536		
45 46	8	34	0	3		0	9.63794			10.31664			15
46 47		33 33	52 44	ļ	26 26	8 16	63820 63846			1			
48		33 33	36	Ĭ	26 26	16 24	63846 63872	95446 95440					
49				1	26 26	32	63898		68432 68465		04560 04566		
50	8	33	20	3	_	40	9.63924					10.36076	
51	_	33	12			48	63950				04579		
52		33	4	1	26	56	63976	95415	68561	31439	04585	36024	
53		32	56		27	4	64002	95409	68593	31407	04591	35998	7
51		32		_		12	64028	95403	68626	31374		35972	
	8	32			27							10.35946	
56 57		32		ļ	27	28	64080	95591	68690	31310	04609	35920	4
5 7 58		52 32				36							
59		32 32	16		27 27		64132 64158						
50		32 32	ö		28	02	64184		68786 68818				
	Ī		_	_			Co-sine.	Sine.		Tangent.			M
5 De							dilg.	C c	rang.	- Tent.	Co-secant Digitize		
	-	•						~ 6			21911120	Del	AG.

210

TABLE XXVII.

Log. Sines, Tangents and Secants.

1	26	Đe	103.					8	,	,			Degs. 1	
-			our		Ш			Sine.	Co-sine.	Tangent.	Co-tang-	Secant.	Co-secant	
	M	-	_				_		9.95366		10.31182		10.35816	60
П	0	18		0	3	28	0	9.64184	9.95360	68850		04640	35790	
П	1	ı	31	52		28 28	8	64210 64236	95354	68882		04646	35764	58
Н	2	ı	31	44			16 24	64262	95348	68914	31086	04652	35738	57
Н	3		31			28	32	64288	95341	68946	31054	04659	35712	56
Н	4	_	31	28		28								
П	5	8		20	3		40	9.64313	9.95335	9.68978				55
Н	6		31	12		2 8	48	64339	95329	69010	30990	04671	35661	54
ı	7	Į	31	4		28	56	64365	95323	69042	30958	04677	35635	53
	8	1	30	56		29	. 4	64391	95317	69074	30926	04683 04690	35609 35583	52
Н	9	1	30	48		29	12	64417	95310	69106	30894			51
П	10	8	30	40	3	29	20	9.64442	9.95304	9.69138		10.04696		50
	11	-	30	32	1	29	28	64468	95298	69170	30830	04702	35532	49
	12	l	30			29	36	64494	95292	69202	30798	04708	35506	48
	13	1	30	16		29	44	64519	95286	69234	30766	04714	35481	47
	14	ı	30	8	ŀ	29	52	64545	95279	69266	30734	04721	35455	46
1	15	8	30	0	3	30	Ó	9.64571	9.95273	9.69298	10.30702	10.04727	10.35429	45
		۱°	29	52	١,	30	8	64596	95267	69329	30671	04733	35404	44
١	16 17		29	44	l	30	16	64622	95261	69361	30639	04739	35378	43
1		ŀ	29	36		30	24	64647	95254	69393	30607	04746	35353	42
-	18 19	ı	29	28	1	30	32	64673	95248	69425	30575	04752	35327	41
H		<u>ب</u>			_				9.95242	9.69457	10.30543	10.04758	10.35309	40
ı	20	8		20	3	30	40	9.64698	9.95242	69488	30512	04764	35276	39
	21	l	29	12		30	48	64724	95230	69520	30480	04771	35251	38
١	22	1	29	4		30	56	64749	95219 95223	69552	30448	04777	35225	37
H	23	1	28	56		31	4	64775	95223 95217	69584	30416	04783	35200	36
i	24		28	48	_	31	12	64800						
	25	8	28	40	3	31	20	9.64826	9.95211	9.69615		10.04789		35
'	26	1	28	32		31	28	64851	95204	69647	30353	04796	35149	34
1	27	l	28	24	ŀ	31	36	64877	95198	69679	30321	04802	35123	33
1	28	ı	28	16		31	44	64902	95192	69710	30290	04808	35098	32
	29		28	8	1	31	52	64927	95185	69742	302 58	04815	35073	31
	30	8	28	ō	3	32	0	9.64953	9.95179	9.69774	10.30226	10.04821		30
П	31	-	27	52	1	32	8	64978	95173	69805	30195	04827	35022	29
	32	1	27	44		32	16	65003	95167	69837	30163	04833	34997	28
.	33		27	36		32	24	65029	95160	69868	30132	04840	34971	27
	34	1	27	28		32	32	65054	95154	69900	30100	04846	34946	26
	35	8	27	20	3	32	40	9.65079	9.95148	9.69932	10.30068	10.04852	10.34921	25
١	36	١٣	27	12	١	32	48	65104	95141	69963	30037	04859	34896	24
	37	ĺ	27	4		32	56	65130	95135	69995	30005	04865	34870	23
	38	l	26	56		33	4	65155	95129	70026	29974	04871	34845	22
	39	•	26	48		33	12	65180	. 95122	70058	29942	04878	34820	21
		۱-,			-	33	20	9.65205	9.95116	9.70080	10.29911		10.34795	20
	40	8		40	3		28		9.95110	70121	29879	04890	34770	19
١	41	1	26	32	1	33	28 36	65230	95103	70152		04897	34745	18
I	42	1	26	24	1	33		65255	95097	70132		04903		i7
١	43	1	26	16	1	33 33	44	65281	95097 95090	70215	29785	04910		16
. 1	44	1_	26	8	_		52	65306						
	45	8		0	3	34	0	9.65331	9.95084	9.70247				15
,	46	1	25		-	34	8	65356	95078	70278	29722	04922		14
	47	1	25			34	16	65381	95071	70309	29691	04929	34619	13
	48		25	36	1	34	24	65406	95065	70341	29659	04935		12
	49	L	25	28		34	32	65431	95059	70372	29628	04941	34569	11
	50	8	25	20	3		40	9.65456	9.95052		10.29596		10.34544	10
١	51	1	25	12	ı	34	48	65481	95046	70435	29565	04954	34519	9
١	52	1	25	4	1	34	56	65506	95039	70466	29534	04961	34494	8
١	53	1	24	56	1	35	4	65531	95033	70498	29502	04967	34469	7
ļ	54	1	24	48	l	35	12	65556	95027	70529	29471	04973	34444	6_
ı	55	8	24	40	3	35	20	9.65580	9.95020	9.70560	10.29440	10.04980	10.34490	5
ł	56	۱۳	24	32	ľ	35	28	65605	95014	70592	29408	04986	34595	4
- 1	57	1	24	24	l	35	36	65630	95007	70623	29377	04993		3
١	58	1	24	16		35	44	65655	95001	70654	29346	04999	34345	2
•	59	1	24	8	l	35	52	65680	94995	70685	29315	05005	34320	1
١	60	I	24	0		36	Õ	65705	94988	70717	29283	05012	34295	0
ł	M	T.			77									M
_!	WI	PE 7 (DULL	·М.	HO	ours	M.M.	Co-sine.	Sine.	CO-cang.	Tangent.	CO-SECSTION	DECEMB	***

116 Degs.

İ		~ T	.				J	Log. Sine	s, Tang	ents and	Secants	• .	D 1-0	
l-,-,			eg:		¥¥-			0:	· ·	T	(A- 4	04	Degs. 152	
			_	_		urr		Sine.			Co-tang.	Secant.	Co-secant	M
H	0	8	24 23	0 52	3	36 36	8	9.65705 65729	9.94988 94982	70743	10.29283 29252	05018	10.34295 34271	60 59
!	2		23	44		36	16	65754	94975	70779	29221	05025	34246	58
11	3		23	36		36	24	65779	94969	70810	29190	05031	34221	57
11	4		23	28		36	32	65804	94962	70841	2 9159	05038	34196	56
-	5	8	23	20	3	36	40	9.65828	9.94956	9.70873	10.29127	10.05044	10.34172	55
11	6		23	12		36	48	65853	94949	70904	29096	05051	34147	54
11	7		23	4		36	56	65878	94943	70935	29065	05057	34122	53
H	8		22 22	56 48		37 37	4 12	65902 65927	94936 94930	70966 70997	29034 29003	05064 05070	34098	52
l I-	9	_		_	_		20						34073	51
	10 11	8	22 22	40 32	3	37 37	28	9.65952 65976	9.949 23 94917	9.71028 71059	10.28972 28941	10.05077 05083	10.34048 34024	50 49
	12		ē2	24		37	36	66001	94911	71090	28910	05089	33999	48
	13		22	16		37	44	66025	94904	71121	28879	05096	33975	47
П	14		22	8		37	52	66050	94898	71153	28847	05102	33950	46
	15	8	22	0	3	38	0	9.66075	9.94891	9.71184	10.28816	10.05109	10.33925	45
	16		21	52		38	8	66099	94885	71215	287 85	05115	33901	44
	17		21	44		38	16	66124	94878	71246	28754	05122	33876	43
	18 19		21 21	36 28		38 38	24 32	66148 66173	94871 94865	71277 71308	28723 28692	05129 05135	33852 33827	42 41
ı	_	<u> </u>			_		40			9.71339		10.05142		
	20 21	8	21 21	20 12	3	38 38	48	9.66197 66221	9.94858 94852	71370	10.28661 28630	05148	10.33803 33779	40 39
	22		21	4		38	56	66246	94845	71401	28599	05155	33754	38
	23		20	56		39	4	66270	94839	71431	28569	05161	33730	37
11:	24		20	43		39	12	6 6295	94832	71462	285 38	05168	33705	36
	25	8	20	40	3	39	20	9.66319	9.94826	9.71493	10.28507	10.05174	10.33681	35
	26		20	32		39	28	66343	94819	71524	28476	05181	33657	34
	27		20	24		39	36	66368	94813	71555	28445	05187	33632	33
	28 29		20 20	16 8		39 39	44 52	6639 2 66416	94806 94799	71586 71617	28414 28383	05194 05201	33608 33584	32 31
! <u>`</u>	30	_	20	- °	3	40	02		9.94793	9.71648		10.05207		30
	31	8	20 19	52	3	40	8	9.66441 66465	94786	71679	10.28352 28321	05214	10.33559 33535	29
	32		19	44		40	16	66489	94780	71709	28291	05220	33511	28
	33		19	36		40	24	66513	94773	71740	28260	05227	33487	27
	34		19	28		40	32	66537	94767	71771	28229	05233	33463	26
	35	8	19	20	3	40	40	9.66562	9.94760		10.28198			25
	36		19	12		40	48	66586	94753	71833	28167	05247	33414	24
	37 38		19 18	4 56		40 41	56 4	66610 66634	94747 94740	71863 71894	28137 28106	05253 05260	33390 33366	23 22
	39		18	48		41	12	66658	94734	71925	28075	05266	33342	21
	40	8	18	40	3	41	20	9.66682	9.94727		10.28045	10.05273	l	20
	41	Ŭ	18	32	ľ	41	28	66706	94720	71986	28014	05280		19
	42		18	24		41	3 6	66731	94714	72017	27983	05286	33269	18
	43		18	16	l	41	44	66755	94707	72048		05293 05300	33245	17
i	44		18	8	_	41	52	66779	94700	72078	27922			16
	15	8	18	0	3	42 42	0	9.66803	9.94694	9.72109 72140	10.27891	10.05306 05313		15
	16 17		17 17	52 44	ŀ	42	8 16	66827 66851	94687 94680	72170				14 13
	18		17	36		42	24	66875	94674	72201	27799	05326		12
1 4	49		17	28		42	32	66899	94667	72231	27769	05333	3 3101	11
	50	8	17	20	3	42	40	9.66922	9.94660	9.72262	10.27738	10.05340	10.33078	10
	51		17	12	l	42	48	66946	94654	72293				
	52		17	4	1	42	56	66970	94647	72323				
	53 54		16 16	56 48	l	43 43	4 12	66994 67018	94640 94634	72354 72384				7
1	55	8	16	40	3	43	20	9.67042	9.94627	9.72415		10.05373	-	5
	56	0	16	32	٦	43		67066	94620					
	57		16	24	l	43		67090	94614			1		
	8		16	16	l	43	44	67113	94507	72506				2
	59		16	8	•	43		67137	94600					
1 6	io l		16	O	•	44	n	67161	94595	72567	1 27435	05407	71 32839	0

0 8 1 2 3 4 4 5 8 6 7 8 9 9 10 8 11 12 13 13 14 15 16 16 17 18 19 20 8 21 22 23 24 25 8 29	Hou 8 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	IFA.M. 16 0 15 52 15 44 15 36 15 28 15 20 15 12 15 4 4 4 48 4 40	3	44 44 44	0 8 16 24	67185	9.94593 94587 94580 94573	72598	10.27433 27402 27372	10.05407 05413 05420	32815	
0 8 1 2 3 4 4 5 8 6 7 8 9 9 10 8 11 12 13 13 14 15 16 16 17 18 19 20 8 21 22 23 24 25 8 29	8 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	16 0 15 52 15 44 15 36 15 28 5 20 5 12 5 4 4 56 4 48 4 40	3	44 44 44 44 44	0 8 16 24 32	9.67161 67185 67208 67232	9.94593 94587 94580 94573	9.72567 72598 72628	10.27433 27402 27372	10.05407 05413 05420	10.32839 32 815	60 5 9
1 2 3 4 5 8 6 7 7 8 9 10 8 11 12 13 13 14 15 16 17 18 19 20 21 22 23 24 25 8 29	11 11 11 11 12 14 14 14 14 14 14 14 14 14 14 14 14 14	15 52 15 44 15 36 15 28 15 20 15 12 15 4 4 56 4 48		44 44 44 44	8 16 24 32	67185 67208 67232	94587 94580 94573	72598 72628	27402 27372	05413 05420	32815	59
2 3 4 5 8 6 7 8 9 10 8 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 8 29	11 11 8 11 14 14 14 14 14 14 15 16 17 18 11 11 11 11 11 11 11 11 11 11 11 11	15 44 15 36 15 28 15 20 15 12 15 4 4 56 4 48 4 40	3	44 44 44	16 24 32	67208 67232	94580 94573	72628	27372	05420		
3 4 5 8 6 7 8 9 9 10 1 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 8 29	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	5 36 5 28 5 20 5 12 5 4 4 56 4 48 4 40	3	44	24 32	67232	94573				32792	5.0
5 8 8 9 10 8 11 12 13 14 15 16 17 18 19 20 8 21 22 22 23 24 25 8 29	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	5 28 5 20 5 12 5 4 4 56 4 48 4 40	3	44	32			72650				
5 8 8 9 9 10 8 11 12 13 13 14 15 16 16 17 18 19 20 21 22 23 24 25 8 29	8 1.1 1.4 1.4 1.4 1.4 1.4 1.4 1.4 1.4 1.4	5 20 5 12 5 4 4 56 4 48 4 40	3	44		67256	OAKET		27341	05427	32768	57
6 7 8 9 10 8 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 27 28 29	1: 1: 1: 8 1: 1: 1: 8 1: 1:	5 12 5 4 4 56 4 48 4 40	3		40		94567	72689	27311	05433	32744	56
6 7 8 9 10 8 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 27 28 29	1: 1: 1: 8 1: 1: 1: 8 1: 1:	5 12 5 4 4 56 4 48 4 40		44		9.67280	9.94560	9.72720	10.27280	10.05440	10.32720	55
7 8 9 9 10 8 11 12 13 14 15 16 17 18 19 20 8 21 22 23 24 25 8 29	14 14 8 14 14 14 14 14 15 15	5 4 4 56 4 48 4 40			48	67303	94553	72750	27250	05447	32697	54
8 9 10 8 111 12 13 14 15 8 16 17 18 19 20 8 22 22 23 24 25 26 27 28 29	14 14 14 14 14 14 14 15 15	4 48 4 40		44	56	67327	94546	72 780	27220	05454	32673	58
9 10 8 11 12 13 14 15 16 17 18 19 20 8 21 22 22 24 25 8 29 29	8 14 14 14 14 8 14 13	4 40		45	4	67350	94540	72811	27189	05460	32650	52
10 8 11 12 13 14 15 16 17 18 19 20 8 21 22 23 24 25 27 28 29	14 14 14 8 14 13			45	12	67374	94533	72841	27159	05467	32626	51
11 12 13 14 15 16 17 18 19 20 8 21 22 23 24 25 26 28 29	14 14 14 8 14 13		-	45	20	9.67398	9.94526	0 79979	10.27128	10 05474		
12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29	14 14 8 14 1;	T U~	J	45	28	67421	94519	72902	27098	05481	32579	50
13 14 15 16 17 18 19 20 8 21 22 22 23 24 25 8 8 27 28 29	1 4 8 1 4 1 1	4 24		45	36	67445	94513	72932	27068	05487	3255à	49 48
14 8 15 16 17 18 19 20 8 21 22 23 24 25 8 29	8 14 13			45	44	67468	94506	72963	27037	05494	32532	47
15 8 16 17 18 19 20 8 21 22 23 24 25 8 26 27 28 29	8 14 13			45	52	67492	94499	72993	27007	05501	32508	46
16 17 18 19 20 8 21 22 23 24 25 26 27 28 29	1:		-									
17 18 19 20 8 21 22 23 24 25 8 8 27 28 29	13		3	46	0	9.67515	9.94492			10.05508		45
18 19 20 8 21 22 23 24 25 26 27 28 29				46 46	8 16	67539	94485	73054	26946 96016	05515	32461	44
19 20 8 21 22 23 24 25 8 26 27 28 29	1:			46	24	67562 67586	94479 94472	73084 73114		05521 05528	32438	43
20 8 21 22 23 24 25 8 26 27 28 29	1:			46	32	67586 67609	94465	73144	26 866 26 856	05535		42
21 22 23 24 25 8 26 27 28 29					_						32391	41
22 23 24 25 26 27 28 29	B 1		3	46	40	9.67633	9.94458		10.26825			40
23 24 25 8 26 27 28 29	13			46	48	67656	94451	73205	26795	05549	32344	39
24 25 8 26 27 28 29	1:			46	56	67680	94445		26765	055 55	32320	38
25 8 26 27 28 29	15			47	4	67703	94438		26735	05562	32297	37
26 27 28 29	15			47	12	67726	94431	73295	26705	05569	32274	36
27 28 29	3 15		3	47	20	9.67750	9.94424		10.26674		10.32250	35
28 29	12			47	28	67773	94417	73356	26644	05583	32227	34
29	15	-			36	67796	94410		26614	05590	32204	33
	15			47	44	67820	94404		26584	05596	32180	32
90 0	15	2 8		47	52	67843	94397	73446	26554	05603	32157	31
30 8	8 15	2 0	3	48	0	9.67866	9.94390	9.73476	10.26524	10.05610	10.32134	30
31	1	1 52		48	8	67890	943 83		26493	05617	32110	29
32	1			48	16	67913	94376	73537	26463	05624	32087	28
33		1 36		48	24	67936	94369	73567	26433	05631	32064	27
34	1	1 28		48	32	67 959	94362	73597	26403	05638	32041	26
35 8	3 1	1 20	3	48	40	9.67982	9.94355	9.73627	10.26373	10.05645	10.32018	25
36	11	1 12		48	48	68006	94349	73657	26343	05651	31994	24
37	11			48	56	68029	94342	73687	26313	05658	31971	23
38	10			49	4	68052	94335	73717	26283	05665	31948	22
39	10	0 48		49	12	68075	94328	73747	26253	05672	31925	21
40 8	3 10	0 40	3	49	20	9.68098	9.94321	9.73777	10.26223	10.05679	10.31902	20
41	10				28	68121	94314	73807	26193	05686	31879	19
42	10			49	36	68144	94307	73837	26163	05693	31856	18
43.	10			49	44	68167	94300	73867	26133	05700	31833	17
44	10			49	52	68190	94293	73897	26103	05707	31810	
45 8	3 10	0 0	3	50	-0	9.68213	9.94286			10.05714		15
46		9 52	•	50	8	68237	94279	73957	26043	05721	31763	
47		9 44			16	68260	94273	73987	26013	05727	31740	14
48		9 36		50	24	68283	94266	74017	25983	05734	31717	12
49		9 28		50	32	68305	94259	74047	25953	05741	31695	11
50 8		9 20	3	50	40	9.68328	9.94252			10.05748		
51		9 12	J		48	68351	94245	74107	25893 25893	05755		10
52		9 4		50		68374	0400-1					
53		8 56		51		68397	94238	74137 74166	25863 25834			
54		8 48		51		68420	94224		25804 25804			
	_	8 40	-		_							
55 8 56		8 32	J	51			9.94217			10.05783		
57		8 24		51 51		68466	94210	74.256	25744			
58		B 16		51			94203					
59	8					6851 2 68534	94196 94189	743 16				
60	8			n.			291041	743415	256 55	05811	31466	
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*18 D	8 8 8	в о		52	o	68557	94182	7437'5		05818		_
ייט או	8	B 0		52	o		94182	7437'5		05818	Secant) Degs.	0910

Log. Sines,	Tangents	and S	Secants.
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ı	29 D	egs.			_									. 150
17	M	Ho	ur.	M.A	H.	our	P.M.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secure	M
11	0	8	8						9.94182		10.25625			60
Ш	1	!	7			52			94175					59
П	2	1	7			52			94168					58
11	3		7			52		68625	94161	1		058 3 9 05846		57 56
H	4	<u>_</u>	7			52		68648	94154					
П	5	8	7		-	52			9.94147 94140		10.25476 25446		10.313 2 9 31306	55
Ħ	6		7			52 52		68694 68716	94130			05860 05867	31284	54 53
	7 8		6	_	1	53		687 3 9	94126			05874		52
11	9		6	-		53		68762	94119	74643		05881	31238	51
11	10	8	-6	_	-	53	20	9.68784	9.94112	9.74673			10.31216	50
11	11	1 "	6			53		68807	94105			05895		49
11	12	Ī	6			53		68829	94098				31171	48
П	13	1	6	16	1	53	44	68852	94090	74762		05910	31148	47
П	14		6	8	L	53	52	68875	94083	74791	25209	05917	31125	46
H	15	8	6	0	3	54	0	9.68897	9.94076	9.74821	10.25179	10.05924	10.31103	45
11	16	1	5			84	8	68920	94069	74851	25149	05931	31080	44
	17		5	44	ł	54	16	68942	94062	74880	25120		31058	43
!	18	1	5		1	54	24	68965	94055	74910		05945	31035	42
11	19		5		 	54	32	68987	94048	74939	25061	05952	31013	41
11	20	8	5	20	3	54	40	9.69010			10.25031			40
П	21	ı	5		1	54	48	69033	94034 94027	74998 75028	25002 24972	05966 05973	30968 30945	39
П	22 23	1	5 4	56	1	54 55	56 4	69055 69077	94020	75028	24942	059 80		
Н	24	1	4	48	1	55	12	69100	94012	75087	24913	05988	30900	36
H	25	B	4	40	-	55	20	9.69122			10.24883			35
П	26	l°	4	32] 3	55	28	69144	93998	75146	24854	06002	30856	34
Н	27	l	4	24]	55	36	69167	93991	75176	24824	06009	30833	33
Н	28	1	4	16	1	55	44	69189	93984	75205	24795	06016	30811	32
Н	29	l	4	8		55	52	69212	93977	75235	24765	06u23	30788	31
Ħ	30	8	4	0	3	56	0	9.69234	9.93970	9.75264	10.24736	10.06030	10.30766	30
11	31		8	52		56	8	69256	93963	75294	24706	96037	30744	29
11	32	l	3	44		56	16	69279	93955	75323	24677	06045	30721	28
H	33		3	36	ı	66	24	69301	93948	75353	24647	06052	30699	27
S L	34		3	28	_	56	32	69323	93941	75882	24618	06059	30677	26
П	35	8	3	20	3	56	40	9.69345	9.93934		10.24589			25
11	36	l	3	12	1	56	48	69368	93927	75441	24559	06073	30632	24
11	37 38		3	56		56 57	56	69390 69412	93920 93912	75470	24530	06080		23
١!	36 39	l	2	48		57	4 12	69434	93905	75500 75529	24500 24471	06088 0609 <i>5</i>	3 0588 30 566	22 21
lŀ		-			-		_							
П	40 41	8	2 2	40 32	8	57 57	20 28	9.69456 69479	9.93898 93891	9.75558 75 588	10. 24442 2 4412	10.06102 06109	10.30544 30521	20 19
11	42		2	24		57	36	69501	93884	75617	24383	06116	30499	18
П	43		2	16		57	44	69523	93876	75647	24353	06124	30477	17
H	44		2	8	1	57	52	69545	93869	75676	24324	06131	30455	16
۱t	45	8	2	0	3	58	-0	9.69567	9.93862	9.75705	10.24295	10.06138	10.30433	15
H	46	آ	ĩ	52	1	58	8	69589	93855	75735	24265	06145	30411	14
11	47	I	1	44	l	56	16	69611	93847	75764	24236	06153	30389	13
Н	48		1	36	ł	58	24	69633	93840	75793	24207	06160	30367	12
I L	49		1	28		58	32	69655	93833	75822	24178	06167	30345	11
١ſ	50	8	1	20	3	58	40	9.69677	9.93826			10.06174		10
Н	51	l	1	12		58	48	69699	93819	75881	24119	06181	30301	9
П	52	ŀ	1	4		56	56	69721	95811	75910	24090	06189	30279	8
П	55 54			56 48		69 59	.3	69743	93797	75939 75969	24061	06196 06203	30257 30235	7
-		_	_	_	_	_					24031			<u>-6</u>
П	55	8		40	3	59 59		9.69787	9.93789		10.24002	10.06211 06218		5
П	56 57			32 24		59		69809 69831	93782 9377 <i>5</i>	76027 76 056	23973 23944	06218	30191 30169	4 3
H	58			16		59		69853	93768	76086	23914	06232	30147	2
Н	59		ŏ	8		59		69875	93760	76115	23885	06240	30125	ĩ
H	60		Ö	o	4	0	0	69897	93753	76144	23856	06247	30103	ō
! 	M	Hoi	IFP	<u></u>	Ho	UFA	<u>.</u>	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secent	Secant.	M
÷	1191										• Enditzed			

TABLE XXVII.

30 1	Degs.		~b. DI		3			Degs. 1	49.
M		Hourp.m.	Sine.	Co-sine.	Tangent.			Co-secant	M
0	8 0 0		9.69897	9.93753		10.23856			60
1	7 59 52 59 44		69919 69941	93746 93738	76173 7 62 02	23827 23798	06254 06262	30081 30059	59 58
2 3	59 44 59 36		69963	93731	76231	23769	06269	30037	57
4	59 28		69984	93724	76261	23739	06276	30016	56
5	7 59 20	4 0 40	9.70006	9.93717	9.76290	10.23710	10.06283	10.29994	55
6	59 12		70028	93709	76319	23681	06291	29972	54
7	59 4			93702	76348	23652	06298	29950	53
8	58 56 58 48		70072 70093	9369 <i>5</i> 93687	76377 76406	23623 23594	06305 06313	29928 29907	52 51
10	7 58 40		9.70115	9.93680		10.23565			50
11	58 32		70137	93673	76464	23536	06327	29863	49
12	58 24	1 36	70159	93665	76493		06335	29841	48
13	58 16			93658	76522	23478	06342	29820	47
14	58 8			93650	76551	23449	06350	29798	46
15	7 58 0		9.70224	9.93643 93636	9.76580 76609	10.23420 23391	10.06357 0 63 64	10.29776 29755	45
16 17	57 52 57 44	1 2		93628	76639	23361	06372	29733	43
18	57 36			93621	76668		06379	29712	42
19	57 28			93614	76697	23303	06386	29690	41
20	7 57 20	4 2 40	9.70332	9.93606	9.76725	10.23275	10.06394	10.29668	40
21	57 12				76754	23246	06401	29647	39
22	57 4			93591	76783 76812		06409 06416	29625 29604	38
23 24	56 56 56 48			93584 93577	76841	23 188 23 159	06423		36
25	7 56 40	·		9.93569		10.23130			35
26	56 32			93562	76 899	23101	06438	29589	34
27	56 24			93554	76928	23072	06446	29518	33
28	56 16			93547	76957	23043			32
29	56 8			93539	76986	23014	06461	29475	31
30	7 56 (9.93532		10.22985			30
31	55 52 55 44			93525 93517	77044 77073	22956 22927	06475 06483	29432 29410	29 28
32 33	55 36			93510	77101	22899	06490		27
34	55 28		1	93502	77130	22870	06498	29367	26
35	7 55 20	4 4 40	9.70654	9.93495	9.77159	10.22841	10.06505	10.29346	25
36	55 12			93487	77188	22812	06513	29325	24
37	55 4			93480	77217		06520		23
3 8 39	54 56 54 48			93472 93465	77246 77274		06528 06535		23 21
						·			
40 41	7 54 40 54 32			9.93457 93450		10.22697 22668	06550		20 19
42	54 24						06558		18
43	54 16	5 44	70824	93435	77390	22610	06565	29176	17
44	54 8			93427	77418		06573	29154	16
45	7 54 0			9.93420		10.22553			15
46	53 52 53 44				77476 77505	22524	06588 06595		14
47 48	53 36			93405 93397	77533	22495 22467	06603		12
49	53 28			93390			06610		ii
50	7 53 20	4 6 40		9.93382		10.22409	10.06618	10.29027	10
51	53 12	6 48	70994	93375	77619	22381	06625	29006	9
52	53 4				77648				
53	52 56 52 48								
- 54						·			_
55 56	7 52 40 52 32					10.22266 22237			5 4
50 57	52 24								
58	52 16	7 44	71142					28858	2
59	52 8	7 52	71163	93314	77849	22151	06686		
60	52 0			93307	77877	22123	06693		-
IVI I	HOURP.M.	Houra.m.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secant		octe
120	Degs.			_				Degs.	99

TABLE XXVII. Log. Sines, Tangents and Secants.

ı		D	_				I	∠og. Sine	s, Tang	ents and	Secants.	•		
-	M	De	ura		T).			Sine.	Co de la	.			Deg. 1	_
1			_	-	no	_	_		Co-sine.		Co-tang.		Co-secant	M
1	0	7	52 51	0 52	4	8	8	9.71184 71 2 05	9.93307 93299	9.77877 77906	10.22123			60
1	2	1	51	44		8	16	71226	93291	77935	22094 22065	06701 06709	28795 28774	59 58
1	3	l	51	36	l	8	24	71247	93284	77963	22037	06716	28753	5 7
	4		51	28		8	32	71268	93276	77992	22008	06724	28732	56
1	5	7	51	20	4	8	40	9.71289	9.93269		10.21980	10.06731		55
l	6	ı	51	12		8	48	71310	93261	78049	21951	06739	28690	54
	8	ı	51 50	4 56		8	56	71331	93253	78077	21923	06747	28669	53
ı	ŝ	ı	50	48	ŀ	9	12	71352 71373	93 24 6 9 323 8	78106 78135	21894	06754	28648	52
ı	10	7	50	40	4	-	20	9.71393	9.93230		21865	06762	28627	51
1	ii	Ι'	50	32	7	9	28	71414	93223	9.78163 78192	10.21837 21808	06777		50
l	12	ı	50	24	1	9	36	71435	93215	78220	21780	06785	28586 28565	49 48
ı	13	١.	5 0	16		9	44	71456	93207	78249	21751	06793	28544	47
l	14	_	50	8	_	9	52	71477	93200	78277	21723	06800	28523	46
ı	15	7	50	0	4	10	0	9.71498	9.93192		10.21694		10.28502	45
ı	16	1	49	52 44		10	8	71519	93184	78334	21666	06816	28481	44
1	18	1	49 49	36		10 10	16 24	71539 71560	93177 93169	78363 78391	21637	06823	28461	43
ı	19	1	49	28	i	10	32	71581	93161	78419	21609 21581	06831 06839	28440	42 41
1	20	7	49	20	4	10	40	9.71602	9.93154		10.21552		28419	
1	21	[49	12	-	10	48	71622	93146	78476	21524	06854	28378	40 39
1	22	1	49	4		10	56	71643	93138	78505	21495	06862	28357	38
1	23	1	48	56		11	4	71664	93131	78533	21467	06869	28336	37
ł	24	_	48	48	L.,	11	12	71685	93123	78562	21438	06877	2 8315	36
ı	25 26	7	48	40	4	11	20	9.71705	9.93115	9.78590	10.21410	10.06885	10.28295	35
1	27	1	48 48	32 24	l	11 11	28 36	71726	93108	78618	21382	06892	28274	34
1	28	1	48	16	1	11	44	71747 71767	93100 93092	78647 78675	21353 21325	06900 06908	28253	33
ı	29	l	48	. 8	ļ	11	52	71788	93084	78704	21326 21296	06916	28233 28212	32 31
1	30	7	48	0	4	12	<u>_</u> 0	9.71809	9.93077		10.21268			30
ĺ	31	}	47	5 2		12	8	71829	93069	78760	21240	06931	28171	29
Ì	32	ı	47	44		12	16	71850	93061	78789	21211	06939	28150	28
	33 34	1	47 47	36 28		12 12	24 32	71870	93051	78817	21183	06947	28130	27
1	35	7	47	20	<u> </u>	12	40	71891	93046	78845	21155	06954	28109	26
ı	36	۱'	47	12	4	12	48	9.71911 71932	9.93038		10.21126			25
Į	37	1	47	4		12	56	71952	93030 930 2 2	7890 2 78930	21098 21070	06970 06978	28068 28048	24
1	38	Ļ	46	56	ŀ	13	4	71973	93014	78959	21041	06986	28048 28027	23 22
ı	39		46	48		13	12	71994	93001	78987	21013	06993	28006	21
1	40	7	46	40	4	13	20	9.72014	9.92999	9.79015	10.20985	10.07001		20
	41	1	46	32	1	13	28	72034	92991	79043	20957	07009	27966	19
	42	1	46 46	24 16	l	13 13	36 44	7 2 055 72075	92983		20928	07017	27945	18
	44	1	46	10	1	13	52	72075 72096	92976 92968		20900 20872	07024 07032	27925	17
1	45	7	46	- 0	4	14	- 6	9.72116					27904	16
L	46	١.	45	52	"	14	8	72137	9.92960	79185	10.20844 20815	07048		15
ł	47	1	45	44	l	14	16	72157	92944		20787		27863 27843	14 13
l	48	1	45	36	I	14	24	72177	92936	79241	20759	07064	27823	12
i	49	 _	45	28	<u> </u>	14	32	72198	92929	79269	20731	07071	27802	11
ł	50	7	45	20	4	14	40	9.72218	9.92921	9.79297	10.20703	10.07079	10.27782	10
ł	51 52	1	45 45	12		14 14	48 56	72238 72259	92913	79326	20674	07087	27762	9
ł	53	l	44	56	l	15	4	72279	92905 92897	79354 7938 2	20646 20618	07095 07103	27741	8
L	54		44	48		15	12	72299	92889	79410	20518		27721 27701	7 6
1	55	7	44	40	4	15	20	9.72320	9.92881		10.20562			- 5
1	56		44	32		15	28	72340	92874	79466	20534	07126	27660	4
l	57	1	44	24	l	15	36	72360	92866	79495	20505	07134	27640	3
	58 59	l	44	16 8	Ì	15	44	72381	92858		20477	07142	27619	2
1	60	ĺ	44	0		15 16	52	72401 72421	92850 92842	79551 79579	20449 20421	07150 07158	27599	1
1		H			Ha			Co.sine.					27579	0
1_			-		-40	w.	·#•	THE STREET	Sine.	UO-CARE.	Tangent.	Co-secant	Secant.	M

121 Degs.

M	32 De	gs.	•	Dog. Dan					Degs.	147.
1			Hourp.m.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.		
1	0									
** *** *** *** *** *** *** *** *** ***										
A										
5 7 43 20 4 16 40 9.72522 9.92803 9.72719 10.24281 10.07197 10.27147 55 6 43 12 16 49 72552 92795 77474 20.253 07205 27458 44 16 66 72562 92787 79764 20.224 07213 27458 55 8 42 56 17 4 72562 92777 79766 20.224 07213 27458 55 9 42 48 17 12 72562 92777 79804 20.196 07221 27418 51 17 12 72562 92777 79804 20.196 07221 27418 51 17 12 72562 92777 79804 20.196 07221 27418 51 17 12 72562 92777 79804 20.196 07221 27418 51 17 12 72562 92777 79804 20.196 07221 27418 51 17 12 72562 92771 79808 20.112 07245 27357 49 11 1 42 24 24 17 36 72563 92747 79916 20.084 07253 27357 42 11 1 42 8 16 17 44 72563 92737 79944 20.066 07251 27357 44 11 1 42 8 16 72703 92731 79972 20.028 07259 27397 45 16 41 52 18 8 72743 92715 92731 79972 20.028 07259 27397 45 16 41 52 18 8 72743 92715 92707 800.06 10.07271 10.27277 40 16 41 44 18 16 72763 92707 800.06 10.00000 10.00000 10.0000 10.0000 10.0000 10.0000 10.0000 10.0000 10.0000 10.0000 10.0000 10.0000 1										
The color of the			4 16 40	9.72522	9.92803	9.79719	10.20281	10.07197	10.27478	55
Second Part										
9 42 48 17 12 72602 92771 79832 20168 07229 77398 51 10 7 42 40 4 17 20 72652 9.92763 9.7866 10.20140 10.072371 77378 50 11 44 32 17 17 28 72663 927767 79988 20112 07245 27357 40 12 42 24 17 36 72663 92739 79944 20066 07251 27317 47 13 42 16 17 44 72663 92739 79944 20066 07251 27317 47 14 42 8 17 52 72703 92731 79972 20029 07269 27397 46 15 7 42 0 4 18 0 9.72723 9.92723 9.80000 10.2000 10.07217 10.27174 47 16 41 52 18 8 72743 92717 80023 19972 07229 27397 44 17 41 44 18 16 72763 92707 80056 19944 07293 27357 44 18 41 36 18 24 72783 92699 80084 19916 07301 27317 42 19 41 28 18 32 72803 92691 80112 19888 07309 27197 41 20 7 41 20 4 18 40 9.72833 9.9263 80162 19832 07329 27197 41 21 41 12 18 48 72843 92675 80162 19832 07329 27157 39 22 41 4 18 66 72863 92667 80195 19805 07335 27157 39 23 40 56 19 4 72883 92659 80162 19832 07325 27157 39 24 40 46 19 12 72902 92651 80251 19749 07349 27098 36 25 7 40 40 4 19 20 9.72922 9.92643 80251 19749 07349 27098 36 25 6 40 32 1928 72942 92635 80307 19593 07356 27038 34 27 40 24 19 36 72962 92667 80195 19805 07335 27157 83 28 40 16 19 44 72982 92635 80307 19593 07356 27038 34 27 40 0 4 20 0 9.73022 9.92603 80361 19590 07335 27157 818 318 30 7 40 0 4 20 0 9.73022 9.92603 80447 19555 07406 25999 23 31 39 52 20 8 73041 92567 80365 19556 07413 26999 31 30 7 40 0 4 20 0 9.73022 9.92603 80447 19555 07406 25999 23 31 39 52 20 8 73041 92567 80569 19351 0.07357 10.25078 34 33 39 44 20 16 73061 92567 80569 10.9442 10.07397 10.25078 34 35 39 30 44 20 16 73061 92567 80569 10.9449 10.07357 10.25078 34 37 39 4 20 56 73160 92567 80569 10.9449 10.07357 10.25078 34 37 39 4 20 56 73160 92567 80360 19370 07406 25999 25 32 39 44 20 16 73061 92567 80360 19470 07429 26899 35 35 57 30 4 20 40 9.73121 9.92567 80447 19555 07406 25999 25 35 37 47 22 4 2 4 3 73160 92568 80560 19351 0.07357 10.25678 30 36 7 39 20 4 20 40 9.73121 9.92567 80569 10.9349 10.07357 10.25678 30 36 7 30 6 4 22 0 9.73318 9.92449 80947 19053 07551 25660 11 36 7 37 80 4 22 0 9.73318 9.92449 80947 19053 07551 25660 11 36 7 36 40 4 22 0										-
10										- 1
11										
12		1							27357	49
14	12		17 36							
15										
16 41 52 18 8 72743 92715 80028 19972 07825 27257 44 17 41 44 18 16 72763 92707 80056 19944 07393 27237 44 18 41 36 18 24 72783 92599 80084 19916 07301 27217 41 19 41 28 18 32 72803 92691 80112 19888 07309 27197 41 20 7 41 20 4 18 40 9.72823 9.2663 9.80140 10.19860 10.07317 10.27177 42 21 41 12 18 48 72843 92675 80168 19832 07325 27157 39 22 41 4 18 56 72863 92667 80195 19805 07333 27157 39 23 40 56 19 4 72883 92659 80223 19777 07341 27177 42 24 40 48 19 12 72902 92651 80251 19749 07349 27098 36 25 7 40 40 4 19 20 9.72922 9.2643 9.80879 10.19721 10.07387 10.27078 35 26 40 32 19 28 72942 92635 80351 19665 07373 271038 32 27 40 24 19 36 72962 92657 80335 19665 07373 27038 34 28 40 16 19 44 72982 92619 80363 19637 07365 27038 34 29 40 8 19 52 73002 92611 80391 19609 07389 36998 31 30 7 40 0 4 20 0 9.73022 9.92603 9.80419 10.19881 10.07387 10.26978 30 31 39 52 20 8 73041 92595 80447 19553 07406 25998 31 33 39 36 20 24 73081 92579 80502 19498 10.07387 10.26978 33 35 39 44 20 16 73661 92887 80474 19525 07445 25999 25 35 39 44 20 16 73661 92887 80474 19525 07445 25999 25 36 39 12 20 48 73140 92555 80586 19414 07445 25680 44 38 32 21 24 73881 92579 80502 19498 07421 226919 27 36 39 38 48 21 12 73200 92525 80689 19331 0.07487 10.26678 25 38 38 56 21 4 73180 92538 80669 19331 0.07487 10.26678 25 38 38 48 21 12 73200 92525 80689 19331 0.07487 10.26678 25 38 38 48 21 12 73200 92529 80699 19331 0.07470 256800 21 40 7 38 40 4 21 20 9.73318 92538 80669 19331 0.07470 25680 21 41 38 32 21 28 732399 92546 80753 19247 07494 26741 18 43 38 16 21 44 73278 92458 80698 19390 07550 25683 10.26681 10.26										
17										1
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	44			16	74794	91840	82953		08160		58
	36			24	74812	91832	82980	17020	08168		57
27 9	28		32	32	74831	91823	83008	16992	08177	25169	56
7 27 5	20	4	32	40	9.74850	9.91215		10.16965		10.25150	55
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27	4		32	56	74887	91798 91789	83089 83117	16911 16883	08211	25094	52
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	48	_		12	9.74943	9.91772	9.83171	10.16829		10.25057	50
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26	8		33	52	75017	91738	83280	16720	08262	24983	46
7 26	ᇹ	4	34	-0	9.75036	9.91729	9.83307	10.16693	10.08271	10.24964	45
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	10	4	35	20	9.75221	9.91643	9.83578	10.16422	10.08357	10.24779	35
	32	•	35	28	75239	91634	83605	16395	08366	24761	34
	24		35	36	75258	91625	83632	16368	08375		33
	16		35	44	75276	91617	83659	16341	08383	24724	32
24	8		35	52	75294	91608	83686	16314	08392		31
7 24	0	4	36	0	9.75313	9.91599		10.16287		10.24687	30
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7 22	40	4	37	20	9.75496	9.91512		10.16016		10.24504	20
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	16		37	44	75551	91486	84065 84092	15935 15908	08514 08523		17 16
22	8		37	52	75569	91477					
7 22	0	4	38	0	9.75587	9.91469	9.84119 84146	10.15881 15854	10.085 3 1 08540		15 14
	52 44		38 38	8 16	75605 75624	91460 91451	84173		08549		13
	36		38	24	75642	91442	84200		08558		12
	28		38	32	75660	91433	84227	15773	08567		11
	20	4	38	40	9.75678	9.91425	9.84254	10.15746	10.08575	10.24322	10
	12	· -	38	48	75696	91416	84280		08584	24304	9
21				56	75714	91407	84307	15693	08593		8
20			39	4	75733	91398					7
20 4	48		39	12	75751	91389		15639	08611		6
7 20		4	39		9.75769	9.91381				10.24231	5
20 :				28	75787	91372					
20 2				36	75805	91363			08637 08646		3
20 1 20			39 39	44	75823	91J54 913 4 5	84496				ĩ
20 20	8		39 40	5Z	75841 75859	91336	84523		08664		ö
	1	H-			Co-sine.			Tangent.			M
Door	•	110	UF A	·	~0-81∏6' ,	Suic.	CO-rang.	- rangent	Correctant	Dem	

Degs.

35 D	POS						TYPE SIII	es, rang	cura and	i December	•	Degs	. 144.
M			.M.	Ho	uri	. M.	Sine.	Co-siné.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	7	20	Q	4	40	0	9.75859	9.91336	9.84523	10.15477	10.08664	10.24141	60
ľi	i ʻ	19	52	•	40	8	75877	91328	84550	15450	08672		59
2	1	19	44	ĺ	40	16	75895	91319	84576	15424	08681	24105	58
3	1	19	36	ŀ	40	24	75913	91310	84 603	15397	0 8690		57
4	İ	19	2 8	l	40	32	75931	91301	84630	15370	08699	24069	56
5	7	19	20	4	40	40	9.75949	9.91292	9.84657	10.15343			55
6	t	19	12		40	48	75 967	91283	84634	15316	08717	24033	54
7	1	19	4	l	40	56	75985	91274	84711	15289	08726 08734	24015 23997	53 52
8	1	18			41	4	76003	91266	8473 8	15262 15236	08743		51
-9	_	18	48	<u> </u>	41	12	76021	91257	84764				
10	7			4	41	20	9.76039	9.91248	9.84791	10.15209 15182	08761	10.23961 23943	50 49
11	l	18	32 24	l	41 41	28 36	76057 76075	91239 91230	84818 84845	15135	08770		48
12	1	18 18	16		41	44	76093	91221	84872	15128	08779		47
14	į	18	8		41	52	76111	91212	84899	15101	08788		46
	7	18	_ 0	4	42	0	9.76129	9.91203	9.84925	10.15075	10.08797	10.23871	45
15	'	17	52	•	42	8	76146	91194	8495 2	15048	08806		44
17	1	17	44		42	16	76164	91185	84979	15021	08815		43
18	1	17	36	l	42	24	76182	91176	85006	14994	08824		42
19	1	17		ı	42	32	76200	91167	8503 3	14967	08833	23 800	41
20	7	17	20	4	42	40	9.76218	9.91158	9.85059	10.14941	10.08842	10.23782	40
21	Ι.	17	12	آ ا	42	48	76236	91149	85086	14914	08851	23764	39
22	ĺ	17	4	1	42	56	76253	91141	85113	14387	08859		38
23	ł	16	56		43	4	76271	91132	85140		08868		37
24	•	16	48		43	12	76289	91123	85166	14834	08877	23711	36
25	7	16	40	4	43	20	9.76307	9.91114				10.23693	35
26	[16	32		48	28	76324	91105	83220	14780	08895		34
27	ł	16	24		43	36	76342	91096	85247	14753	08904 08913		
28	1	16			43	44	763 60	91087	25275	14727 14700	08922		32 31
29	<u>L</u>	16	8	_	43	52	76378	91078	85300		10.08931		
30	7	16	0	4	44	0	9.76395	9.91069	9.85327	10.14673 14646	08940		30 29
31	1	15	52		44	8	76413	9106 0 91051	85354 85380	14620	08949		28
32	Į	15	44 36		44 44	16 24	76431 76448	91042	85407	14593	08958		27
33	l	15 15	28	1	44	32	76466	91033	85434	14566	08967		26
	7		20	4		40	9.76484	9.91023	9.85460		10.08977	10.23516	25
35 36	۱'	15 15	12	*	44	48	76501	91014	85487	14513	08986		24
37	l	15	12		44	56	76519	91005	85514	14486	08995		23
38	ı	14	56		45	4	76537	90996	85540	14460	09004	23 463	22
39	ı	14	48	1	45	12	76554	90987	85567	14433	09013	23446	21
40	7	14	40	4	45	20	9.76572	9.90978	9.85594	10.14406	10.09022	10.23428	20
41	Ι.	14	32	_	45	28	76590	90969	85620	14380	09031	23410	19
42		14	24		45	36	76607	90960		14358	09040		18
43	l	14	16	l	45	44	76625	90951	85674	14326	09049		17 16
44		14	8	L	45	52	76642	90942	85700	14300	09058		
45	7	14	0	4	46	9	9.76660	9.90933	9.85727		10.09067		15
46	1	13		ľ	46	. 8	76677	90924	85754	14246	09076 09085		14 13
47	t	13	44	l	46	16	76695	90915	85780	14220 14193	09088	23288	12
48	I	13	36		46	24	76712	90906 90896	85807 858 34	14166	09104		11
49	L	13		<u> </u>	46	32	76730			10.14140		10.23253	10
50	7		20	4	46	40	9.76747	9.9088 7 90878	9.85860 85887	14113	09122		9
51		13 13	12 4	l	46 46	48 56	76765 7678 2				09131		8
52 53	ł		56		47								7
54	1		48			12		90851	85967	14033	09149	23183	6
55	7		40		47		9.76835	9.90842			10.09158	10.23165	5
56	۱'		32			28	76852	90832	86020			23148	4
57	1		24			36				13954	09177		
58			16			44		90814	86078	13927	09186		2
59	1	12	8		47	52	76904						1
60		12			48	_		90796	_			I	0
M	He	uri	P.M.	He	our/	L.M.	Co-sine.	Sine.	Co-tang.	Tangent.	Cosecam		
12	5 D									- Digitization	, , , ,	O Degs.	51.
		•											

Log. Sines, Tangents and Secants.

36	Degs											Degs. 1	43.
M	Hou		M.	Ho	urp	.M.	Sinc.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	7 1	2	0	4	48	0	9.76922	9.90796	9.86126	10.13874	10.09204	10.23078	60
ĭ		ĩ	52	-	48	8	76939	90787	86153	13847	09213	23061	59
2		1 4	14		48	16	76957	90777	86179	13821	09223	23043	58
3	1	1 :	36		48	24	76974	90768	86206	13794	09232	23026	57
4	1	1 2	28		48	32	76991	90759	86232	13768	09241	23009	56
5	7 1	1 9	50	4	48	40	9.77009	9.90750	9.86259	10.13741	10.09250	10.22991	55
6			12	-	48	48	77026	90741	86285	13715			
ž	l i		4		48	56	77043	90731	86312	13688		22957	53
8			56		49	4	77061	90722	86 338	13662	09278		52
9			18		49	12	77078	90713	86365	13635	09287	12922	51
10			10	4	49	20	9.77095	9.90704	9.86392	10.13608	10.09296	10.22905	50
11	,		32	•	49	28	77112	90694	86418	13582	.09306		49
12			24		49	36	77130	90685	86445			1	48
13			16		49	44	77147	90676	86471	13529			47
14		ō	8		49	52	77164	90667	86498	13502	09333	22 336	46
15		0	0	4	50	0	9.77182	9.90657	9 86594	10 13476	10.09343	10 92819	45
16	' '		52	7	50	8	77199	90648	86551	13449			44
17	l		14		50	16	77216	90639	86577	13423			43
18	1		36		50	24	77233	. 90630	86603	13397			42
19	į .		28		50	32	77250	90620	86630				
20	-	_	20	-	50	40	9.77263	9.90611	9.86656			10.22732	
20 21	7		12	*	50	40	9.77265 77285	90602					
21 22	1	9	4		50	56	77302	90592					
23	l	-	56		51	4	77319	90583					
23 24	ı		48		51	12	77336	90574					
	7		-4	4	51	20				10.13211			
25 ec	[]		40 32	4	51	28	9.77353 77370	9.90565 90555				i 10. 2264 7 i 2263 0	35
26 27	1		3Z 24		51	36	77387	90555					
2 1	1		16		51	44	77405	90537	86868				
20 29	!	8	8		51	52	77422	90527					
	-			_									
30	7	8	0	*	52	0	9.77439	9.90518				10.22561	
31	1		52		52 52	8		90509					
32	1		44 36		52 52	16 24	77473 77490						
33 34	1		20 28		52 52	32	77507	90480					
	-		_	_									
35	7		20	4	52	40	9.77524	9.90471		10.12947		10.22476	
36	l	7	12		52 52	48	77541	90462					
37 38	1		56		53	56 4	77558 77575	90452					
38 39	l		00 48		53	12	77575	90443 90434					
	<u> </u>			<u> </u>									
40	7		40	4	53	20	9.77609	9.90424	1	10.1281		10.22391	
41	İ		32		53	28	77626						
42	1		24		53	36	77643						
43		_	16		53	44	77660						
44		6	8		53	52	77677	90386					
45	7	6	0	4	54	0	9.77694					10.22306	
46			5 2		54	8	77711	90368					
47			44		54	16	77728						
48	l		36		54	24	77744						
49		5	28		54	32	77761	90339	87425	12578	09661	22239	11
50	7	_	20	4	54	40	9.77778	9.90330		10.1255	10.09670	10.2222	10
51	١.		12		54	48	77795					2220	9 1
52	1	5	4		54	56	77812	90311					s -8
53	1	_	56		5 5	4	77829						
54	١.	4	48		55	12	77846	90292	87554	12444	09708	22154	
55	7	4	40	4	55	20	9.77862	9.90282	9.87580	10.12420	10.09718	10.22138	5
56	ľ		32	٠. ا	55	28	77879	90273					
57	ł	_	24		55	36	77896	90263					
58	ı	4	16		55	44	77913	90254	87659				
59	1	4	8		<i>5</i> 5	52	77930	90244					
60	L	4	0		56	0	77946	90235	87711	12289			
M	Hou	rp.		He	NF.		Co-sine.	Sine.	Co-tang	Tanges	Co-secant	Secant.	M
	Dec	_							- 50 0015	- august			valle.

126 Degs.

Degs. 53.

TABLE XXVII.

37 3	Dec	rs.				I	⊿og. Sme	s, renge	ents and	Secants.	•	Degs. 1	49.
			.M.	Н		- w	Sine.	Co-sine.	Tongont	Cotana	Second	Co-secant	M
										Co-tang.			
0	7	4	0	4	56		9.77946	9.90235	9.87711		10.09765		60
1		3	52	l	56		77963	90225	87738	12262		22037	59
2	ł	3	44		56		77980	90216	87764	12236	09784	22020	58
3		3	36	ł	56		77997	90206	87790	12210	09794	22003	57
4		3	23		56	32	7801 3	90197	87 817	12183	09803	21987	56
5	7	3	20	4	56	40	9.78930	9.96187	9.87843	10.12157	10.09813	10.21970	55
6		8	12		56	48	78047	90178	87869	12131	09822	21953	54
7	1	3	4	(<i>5</i> 6	56	78063	90168	87 895	12105	09832	21937	53
8		2	56	į	57		78 080	90159	87922	12078	09841	21920	52
9	١	2	48	1	57	12	78097	90149	87948	12052	09851	21903	51
10	7	2	40	4	57	20	9.78113	9.90139	9 87974	10.12026	10.09861	10.21887	50
lii	₹.	2		1	57		78130	90130	88000	12000	09870	21870	
12		2		!	57		78147	90120	88027	11973	09880	21853	48
13		2		1	57		78163	90111	88053		.09889	21837	47
14	l	. 2			57		78180	90101	88079	11921	09899	21820	
	7			<u>; — </u>									
15	١,	2	0		58	-	9.78197	9.90091		10.11895	10.09909	10.21803	
16	1	1	52		58		78213	90082	88131	11869	09918	21787	
17	1	1	44	1	58			90072	88158	11842	09928 09937	21770	43 42
18	i	1	36	ì	58				88184	11816		21754	
19	L.,	_1	28		58				88210	11790	09947	21737	41
20	7	1	20	4	58		9.78280		9.88236			10.21720	40
21		1	12	1	58		78296	90034	88262	11738	09966	21704	39
22	l	1	4	ļ l	58			90024	88289	11711	09976	21687	3 8
23	ì	0]	59		78329	90014	8 8315	11685	09986	21671	37
24	ŀ	0	48	1	59	12	78346	90005	88341	11659	09995	21654	36
25	7	0	40	4	59	20	9.78362	9.89995	9.83367	10.11633	10.10005	10.21638	35
26	l	0	32	}	59	28	78379	89985	88393		10015	21621	34
27	l	Ō	24		59	36	78395	89976	88420	11580	10024	21605	
28		0	16	١.	ð9	44	78412	89966	88446	11554	10034	21588	
29		Ō	-8	l	59	52	78428	89956	88472	11528	10044	21572	
30	7	0	-ò	5	0		9.78445	9.89947					30
31		59	52	١ '	ŏ		78461	89937	9.88498	11476	10063	21539	29
32	۳	59	44	Į.	ŏ		78478	89927	88524 88550	11450	10073	21522	28
33	1	59	36	İ	ŏ		78494	89918		11423	10073	21506	27
34		59	28	I	ŏ		78510	89908	88577 88603	11397	10092	21490	26
	<u>_</u>			-		_							
35	6	59	20	5	0	40	9.78527	9.89898		10.11371	10.10102	10.21473	25
36		59	12	1	0		78543	89888	88655	11345	10112	21457	24
37		59	4	ı	0		78560	89879	88681	11319	10121	21440	23
38	l	58	56		1	4 12	78576	89869	88707	11293	10131	21424	22
39	L	58	48				78592	89859	88733	11267	10141	21408	21
40	6	5 8	40	5	1	20	9.78609	9.89849	9.88759	10.11241		10.21391	20
41	l	58	32	1	1	28	78625	89840	88786	11214	10160	21375	19
42	1	58	24)	1	36	78642	89830	88812	11188	10170	21358	18
43	Ī	5 8	16	i	1	44	78658	89820	88838	11162	10180	21342	17
44	L	58	8	i_	1	52	78674	89810	88864	11136	10190	21326	16
45	6	58	0	5	2	0	9.78691	9.89801	9.88890	10.11110	10.10199	10.21309	15
46		57	52		2	8	78707	89791	88916	11084	10209	21293	14
47		57	44		. 2	16	78723	89781	88942	11058	10219	21277	13
48		57	36		2	24	78739	89771	88968	11032	10229	21261	12
49		57	28		2	32	78756	89761	88994	11006	10239	21244	11
50	6	57	20	5	4	40	9.78772	9.89752			10.10248	10.21228	10
51	3	57	12	•	2	48	78788	89742	89046	10.10980	10258	21212	9
52		57	4,		2	56	78805	89732	89073	10934	10268	21195	8
53		56	56		3	4	78821	89722	89099	10901	10208	21179	7
54		56	48		3	12	78837	89712	89125	10875	10288	21163	6
	_				_								
55	6	56	40	5	3	20	9.78853	9.89702				10.21147	5
56		56	32		3	28	78869	89693	89177	10823	10307	21131	4
		56	24		.3	36 44	78886	89683	89203	10797	10317	21114	3
57		*^						89673	89229	107711	10327	21098	2
58		56	16		3		78902						
		56 56 56	8		3	52	78918 78934	89663 89653	89255 89281	10745 10719	10337	21082 21066	ĩ

Log. Sines, Tangents and Secants.

38 D													. 141.
M	H	our	A.M	H	DUL	P.W.	Sine.	Co-sine.		Co-tang.		Co-secant	M
0	6	56		7 -	-	•	9.78934			10.10719			60
1 1	!	55			4								
3	1	55 55			4		78967 78983	89633 896 2 4			10367 10376		
1 3	1	55			. 4 4		78983 78999	89624 89614					
	+-	55		1_			9.79015	9.89604					
5	0	55			4		79031	89594				10.20985 20969	
7	ł	55		1	4		79047	89584		1	10416		
1 8	1	54		3	5		79063	89574	1				52
9	1	54	48		5	12	79079	89564	89515	10485	10436	20921	51
10	6	54	40	5	5	20	9.79095	9.89554	9.89541	10.10459	10.10446	10.20905	50
11	1	54	32	4	5		79111	89544	89567	10433	10456	20689	49
12	1	54			5		79128	89534	89595		10466	20872	48
13	1	54	16		5		79144	89524	89619		10476	30856	47
14	<u> </u>	54	8		_ 5	52	79160	89514			10486	20840	46
15	6		0		6	0	9.79176	9.89504		10.10329			45
16	1	53			6	.8	79192	89495		10303	10505	20808	44
17	1	53 53	44 36	•	6 6	16 24	79208 79224	89485 89475	89723		10515 10525	20792 20776	43
18	1	53	28		6	32	79224	89465	89749 89775	10251	10535	20760	41
	6	_			6	40	9.79256	9.89455					40
20	١٥	53 53	20 12		6	48	79272	9.89400 89445	9.89801 89827	10.10199 10173	10.10545	20728	39
22	1	53	12		6	56	79212 79288	89435	89853		10565	20712	38
23	1	52	56		7	4	79304	89425	89879	10121	10575	20696	37
24	1	52	48		7	12	79319	89415	89905	10095	10585	20681	36
25	6	52	40	5	7	žn	9.79335	9.89405	9 89931	10.10069	10.10595	10.20665	35
26	ľ	52	32		7	28	79351	89395	89957	10043	10605	20649	34
27	1	52	24		7	36	79367	89385	89983	10017	10615	20633	33
28	1	52	16	1	7	44	79383	89375	90009	09991	10625	20617	32
29	L	52	8		7	52	79399	89364	90035	09965	10636	20601	31
30	6	02	Ü	5	8	0	9.79415	9.89354	9.90061	10.09939	10.10646	10.20585	30
31	Ì	51	52	1	8	8	79431	89344	90086	09914	10656	20569	29
32	1	51	44	ł	8	16	79447	89334	90112	09888	10666	20553	28
33	ļ	51	36 28	l	8	24 32	79463 79478	89324 89314	90138	09862 09836	10676 10686	20537 20522	27 26
34	با	51	-	Ļ	8				90164				
35	6	51	20	5	8	40	9.79494	9.89304		10.09810		10.20506	25 24
36 37	1	51 51	12		8	48 56	79510 79526	89294 89284	90216 90242	09784 09758	10706 10716	20496 ₁ 20474	23
38	1	50	56		9	4	79542	89274	90268	09732	10726	20458	22
39		50	43	l	9	12	79558	89264	90294	09706	10736	20442	21
40	6	50	40	5	9	20	9.79573	9.89254		10.09680			90
41	ľ	50	32	ľ	9	28	79589	89244	90346	09654	10756	20411	19
42	l	50	24	٠	9	36	79605	89233	90371	09629	10767	20395	18
43	l	50	16	ŀ	9	44	79621	89223	90397	09603	10777	20379;	17
44	L	50	8	L	9	52	79636	89213	90423	09577	10787	20364	16
45	6	50	0	5	10	0	9.79652	9.89203	9.90449		10.10797		15
46		49	52	ł	10	8	79668	89193	90475	09525	10807	20332	14
47		49	44	l	10	16	79684	89183	90501	09499	10817	20316	13
48	l	49 49	36 28	l	10 10	24 32	79699	89173	90527	09473	10827	20301	12
49	<u> </u>			<u> </u>		_	79715	89162	90553	09447	10838	20285	
50	6	49	20	5	10 10	40	9.79731	9.89152		10.09422		10.20269	10
51 52		49 49	12	l	10	445 56	79746 79762	89142 89132	90604	09396 09370	10858 10868	20254 20238	8
53		48	56		11	7	79778	89132	90630 90656	09344	10878	20222	7
54		48	48		11	12	79793	89112	90682	09318	10888	20207	6
55	6	48	49	5	11	20	9.79809	9.89101		10.09292			5
56	۲,	48	32	"	ii	28	79825	89091	90734	09266	10909	20175	4
57		48	24		ii	36	79840	89081	90759	09241	10919	20160	3
58	ł	48	16		11	44	79856	89071	99785	09215	10929	20144	2
59		48	8		11	52	79872	89060	90811	09189	10940	20128	1
60	_	48	0		12	0	79887	89050	90837	09163	10950	20113	0
M	Ho	urp.	M.	Ho	UFA	.m.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secant	Secual.	MI

128 Degs.

Degs. 51.

	Degs.					. 0	,			•	Deg. 14	6 0.
M	Houra	H.	Ho	uri	. M	Sine.	Co-sine.	Tangent.			Co-secant	M
0	6 48	9	5	12		9.79887	9.89050		10.09163	10.10950	10.20113	60
1 2	47	52 44		12 12		79903			09137	10960	20097	
3	47	36	-	12		79918 79934		90889 90914	09111 09086	10970	20082	58
4	47	28		12		79950	89009	90940		10980 10991	20066 20050	57 56
5	6 47	20	5	12	40	9.79965			10.09034			
6	47	12	٦	12	48	79981	88989		09008	11011	10.20036 20019	55 54
7	47	4	ŀ	12	56	79996	88978	91018	08982	11022	20004	53
8	46	56	1	13	4	80012		91043	08957	11032	19988	52
9	46	48		13		80027	88958	91069	08931	11042	19973	51
10	6 46	40	5	13	20	9.80043	9.88948		10.08905			50
11	46 46	32 24		13 13	28 36	80058		91121	08879	11063		49
13		16	٠,	13		80074 80089	889 27 88917	91147 9117 2	0885 3 08828		19926	48
14	46	8		13	52	80105	88906	91198	08802	11003		46
15	6 46	ō	5	14	-0	9.80120			10.08776			45
16		52	Ĭ	14	8	80136	88886	91250				44
17	45	44		14		80151		91276				43
18	45	36		14	24	80166		91301	08699	11135	19834	42
19		28		14	_	80182	88855	91327	08673	11145		41
20	6 45	20	5	14	40	9.80197	9.88844		10.08647			40
21	45	12		14 14	48 56	80213	88834	91379		11166		39
23	44	56		15	4	80228 80244		91404 91430			19 772 19 7 56	
24	44	48		15	12	80259	88803	91456	,		19741	36
25	6 44	40	5	15	20	9.80274			10.08518			35
26	44	32		15	28	80290		91507				
27	44	24		15	36	80305		91533		11228	19695	33
28	44	16		15	44	80320		91559		11239	19680	32
29	44	8		15	52	80336	88751	91585	08415	11249	19664	31
30	6 44	0	5	16	0	9.80351	9.88741		10.08390			30
31 32	43	52 44		16 16	8 16	80366		91636	08364		19634	29
33	43	36		16	24	8038 2 80397	88720 88709	9166 2 91688	08338	11280	19618	28
34		28		16	32	80412	88699	91713	08312 08287	11291 11301	19603 19588	27 26
35	6 43	20	5	16	40	9.80428			10.08261			25
36	43	12		16	48	80443	88678	91765			19557	24
37	43	4		16	56	80458		91791			19542	23
38	42	56	1	17	4	80473		91816		11343	19527	22
39	42	48	_	17	12	80489		91842	08158	11353	19511	21
40	6 42	40	5	17	20	9.80504			10.08132			20
41	42	32 24	1	17 17	28 36	80519		91893				19
43	42	16	l	17	44	80534 80550		91919 919 4 5				18 17
44	42	8	1	17	52	80565		91971	08029	11406	19435	16
45	6 42	_ <u>o</u>	5	18	0	9.80580			10.08004			15
46	41	52	ľ	18		80595		92022			19405	14
47	41	44	1	18		80610	88563	92048	07952			13
48	41	36	l	18		80625		92073		11448		
	41	28	با	18		80641	88542	92099		11458		11
50 51	6 41	20 12	5	18		9.80656		9.92125		10.11469		10
52	41	1Z 4	Ī	18 18		80671 806 86	88521 88510	92150 92176				9
53	40		1	19		7						7
54	40				12							6
55	6 40	40	5	19	20						10.19269	5
56	40	32		19	28	80746	88468	92379				4
57	40				36			92304	07696	11543	19238	3
58 59	40				44	80777						2
60	40	8		19 20	5 2	8079 2 80807	88 43 6 88 42 5	92356 92381	07644 07619	11564 11575		1
			11.			Co-sine.						-
190	Degs.	-=-	440	ur.		CU-SINE.	oute.	Co-tang.	Tangent.	Co-secant		M
_~,	5 -									,	Degs.	gV.

Log. Sines, Tangents and Secants.

Degs.				,	og. om	es, rang	ciits aiiu			Degs. 1	39.
Houra	.M.	Ho	urr	.м.	Sine.	Co-sine.				Co-secant	M
6 40	0	5	20	0	9.80807	9.88425	9.92381	10.07619	10.11575	10.19193	60
	52		20	8	80822	88415	92407	07593	11585	19178	59
	44	1	20	16	80837	88404	92433	07567	11596 11606	19163	58 57
39	36		20	24	80852	88394 88383	92458 92484	07542 07516	11606	19148 1913 3	56
39	28		20	32	80867			10.07490			55
6 39	20	5	20	40	9.80882	9.88372	9.92510	07465	11638	19103	54
39	12		20 20	48 56	80897 80912	88362 88351	92553 92561	07439	11649	19088	53
39 38	56		20 21	<i>3</i> 0	80912	88340	92587	07413	11660	19073	52
38	48		21	12	80942	88330		07383	11670		51
6 38	40	5	21	20	9.80957			10.07362	10.11681	10.19043	50
38	32	0	21	28	80972	88308	92663	07337	11692	19028	49
38	24		21	36	80987	88298	92689	07311	11702	19013	48
38	16		21	44	81002	88287	92715	07285	11713	18998	47
38	8		21	5 2	81017	88276	92740	07260	11724	18983	46
6 3 8	0	5	22	0	9.81032	9.88266	9.92766		10.11734		45
37	5 2		22	8	81047	88255	92792	07208	11745		44 43
37	44		22	16	81061	88244	92817 92843	07183 071 <i>5</i> 7	11756 11766	18939 18924	42
37 37	36 28		22 22	24 32	81076 81091	88 234 88 223	92868	07132	11777	18909	41
		<u> </u>	-	_		9.88212		10.07106			40
6 37	20 12	5	22 22	40 48	9.81106 81121	9.88212 88201	92920	07080	11799	18879	39
37 37	12		22 22	56	81136	88191	92945	07055	11809	18864	38
36	56		23	4	81151	88180		07029	11820	18849	37
36	48		23	12	81166	88169	92996	07004	11831	18834	36
6 36	40	5	23	20	9.81180	9.88158	9.93022	10.06978	10.11842	10.18820	35
36	32	•	23	28	81195	88148	93048	06952	11852	18805	34
36	24		23	36	81210	88137	93073	06927	11863	18790	33
36	16		23	44	81225	88126	93099	06901	11874	18775	32 31
36	8		23	52	81240	88115	93124	06876	11885	18760	
6 36	0	5	24	0	9.81254	9.88105		10.06850		10.18746 18731	30 29
35	52		24	.8	81269	88094	93175 93201	068 2 5 06 79 9	11906 11917	18716	28
35	44		24 24	16	81284 81299	88083 88072	93227	06773	11928	18701	27
35 35	36 28		24	32	81314	88061	93252	06748	11939	18686	26
6 35	20	5	24	40	9.81328	9.88051		10.06722	10.11949	10.18672	25
35	12	_	24	48	81343	88040	93303	06697	11960	18657	24
35	4		24	56	81358	88029	93329	06671	11971	18642	23
34	56		25	4	81372	88018	93354	06646	11982	18628	22
34	48		25	12	81387	88007	93380	06620	11993	18613	21
6 34	40		25	20	9.81402	9.87996		10.06594		10.18598	20
34	32		25	28	81417	87985	93431	06569	12015 12025	185 83 18569	19 18
	24		25	36	81431	87975	93457	0654 3 06518	12036	18554	17
	16			44 52	81446 81461	87964 87953	9 3482 93508	06492	12030	18539	16
34	8			_				10.06467			15
6 34	0		26	0	9.81475	9.87942 87931	935559	06441	12069	18510	14
	52 44		26 26	16	81490 81505	87920	93584	06416	12080	18495	13
33 33	36		2 6	24	81519	87909	93610	06390	12091	18481	12
33	28		26	32	81534	87898	93636	06364	12102	18466	11
6 33	20	5	26	40	9.81549	9.87887	9.93661	10.06339	10.12113	10.18451	10
	12		26	48	81563	87877	93687	06313	12123	18437	9
33	4		26	56	81578	87866	93712	06288	12134	18422	8
32			27	4	81592	87855	93738	06262	12145	18408	7
32	48		27	12	81607	87844	93763	06237	12156	18393	
6 32	40	5	27	20	9.81622	9.87833		10.06211	10.12167	10.18378	5
	32		27	28	81636	87822	93814	06186	12178 12189	18364	4
32			27	36	81651	87811	93840	06160	12189	18349 18335	2
32	16			44	81665	87800	93865 ¹ 9 3 891	06135 06109	12211	18320	î
32 32	8		27 28	5 2	81680 81694	87789 87778	93916	06084	12222	18306	ō
							Co-tang.				M
ourp	M.	Ho	ura	·M·	Co-sine.	Sine.	Co-ung.	Tankenr.	CO-SECULIA	D	10

Degs.

Degs. 49.
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TABLE XXVII.

Log. Sines, Tangents and Secants.

						Log. Sin	es, Tan	gents and	d Secant	3.	D	100
	Degs.		77					7 0		10	Degs.	
M	Hour		-	_		Sine.		Tangent.			Co-secant	
0	6 32			28			9.87778				10.18306	
1	31			28	8		87767		06058			59
2	31	44		28			87756		06033			58
3	31	36		28	24	81738	87745	93993	06007			
4	31	28	1	28	32	81752	87734	94018	05982	12266	18248	56
5	6 31	20	5	28	40	9.81767	9.87723	9.94044	10.05956	10 12277	10.13233	55
6	31	12	"	28	48		87712	94069	05931	12288		
1 7	31	4	١.	28			87701	94095				
8	30	-	l	29	4	81810	87690	94120	05880	1		
9	30	48		29	12	81825	87679	94146	05854	12321	18175	
			 									
10	6 30		5	29	20	9.81839	9.87668	9.94171		10.12332		50
11	30		i	29	28	81854	87657	94197	05803			
12	30	24		2 9	36	31868		94222	05778	12354		
13	30	16	i	2 9	44	81882	87635	94248	05752	12365		
14	30	8	1	29	52	81897	87624	94273	05727	12376	18103	46
15	6 30	0	5	30	ō	9.81911	9.87613	9.94299	10.05701	10.12387	10.18089	45
16	29		1	30	8		87601	94324	05676	12399	18074	1
17	29	44	l	30	16	81940	87590	94350	05650			
18	29	36	ĺ	30	24	81955	87579	94375	05625	12421	18045	
19	29	28	l	30	32	81969	87568	94401	05599	12432	18031	41
I		i	-									
20	6 29	20	5	30	40	9.81983	9.87557	9.94426			10.18017	40
21	29	12	1	30	48	81998	87546	94452	05548	12454	18002	
22	29	4		30	56	82012	87535	94477	05523	12465	17988	
23	23	56	1	31	4	82026	87524	94503	05497	12476	17974	37
24	28	48	ı	31	12	82041	87513	94528	05472	12487	17959	36
25	6 28	40	5	31	20	9.82055	9.87501	9.94554	10.05446	10.12499	10.17945	35
26	28	32		31	28	82069	87490	94579	05421	12510	17931	34
27	28	24		31	36	82084	87479	94604	05396	12521	17916	
28	28	16	!	31	44	82098	87468	94630	05370	12532	17902	32
29	28	8		31	52	82112	87457	94655	05345	12543	17888	31
I				_								
30	6 28	0	5	32	0		9.87446		10.05319		10.17874	30
31	27	52		32	8	82141	87434	94706	05294	12566	17859	29
32	27	44		32	16	82155	87423	94732	05268	12577	17845	28
33	27	36	1	32	24	82169	87412	94757	05243		17831	27
34	27	28	Ĺ.,	32	32	82184	87401	94783	05217	12599	17816	26
35	6 27	20	5	32	40	9.82198	9.87390	9.94808	10.05192	10.12610	10.17802	25
36	27	12		32	48	82212	87378	94834	05166	12622	17788	24
37	27	4		32	56		87367	94859	05141	12633	17774	23
38	26	56		33	4	82240	87356	94884	05116	12644	17760	22
39	26	48		33	12	82255	87345	94910	05090	12655	17745	21
40	6 26	40	5	33	20		9.87334			10.12666	10.17731	20
41	26	32		33	28	82283	87322	94961	05039	12678	17717	19
42	26	24		33	36	82297	87311	94986	05014	12689	17703	18
43	26	16		33	44	82311	87300	95012	04988	12700	17689	17
44	26	8		33	52	82326	87288	95037	04963	12712	17674	16
45	6 26	0	5	34	0	9.82340	9.87277	9.95062	10.04938	10.12723	10.17660	15
46	25	52		34	8	82354	87266	95088	04912	12734	17646	14
47	25	41		34	16	82368	87255	95113	04887	12745	17632	13
48	25	36		34	24	82382	87243	95139	04861	12757	17618	12
49	25	28		34	32	82396	87232	95164	04836	12768	17604	11
50	6 25	20	_	34	40	9.82410	9.87221		10.04810	10.12779		10
51	25	12		34	48	82424	87209	95215	04785	12791	17576	9
52	25	4			56	82439	87198	95240	04760			8
53		56		35	4		87187	95266	04734			7
54	24	48		35	12	82467	87175	95291	04709	12825	17533	6
55	6 24	40	5	35	20	9.82481	9.87164	9.95317	10.04683	10.12836	10.17519	- 5
	24			35		82495	87153	95342	04658	12847	17505	
56		24		35		82509	87141	95368	04632			3
56 57							87130	95393	04607	12870	17477	
57		16		35	44	82523						
57 58	24			35 35		82523 8 2537						
57		16 8 0		35 35 36		82523 8 2537 82551	87119 87107	95418 95444	04582 04556		17463	0

42	Degs.	_		, ,				Degs. 1	
	Houra.M.	Hourp.m.	Sine.	Co-sine.	Tangent.	Co-tang.	Secant.	Co-secant	M
0	6 24 0	5 36 0	9.82551	9.87107		10.04556	10.12893	10.17449	60
ĭ	23 52	36 8		87096	95469	04531	12904		59
2	23 44	36 16		87085	954 95		12915		58
3	23 36	36 24	82593	87073	95520			17407	57
4	23 28	36 32	82607	87062	95545			17393	56
5	6 23 20	5 36 40	9.82621	9.87050	9.95571	10.04429	10.12950	10.17379	55
6	23 12	36 48	82635	8703 9	95596		12961	17365	54
7	23 4	36 56	82649	8702 8	95622	04378	12972		53
8	22 56	37 4	8 2 663	87 016	95647	04353	12984		52
9	22 4 8	37 12	82677	87005	95672	04328	12995	17323	51
10	6 22 40	5 37 20	9.82691	9.86993	9.95698		10.13007		50
11	22 32	37 28	82705	86982	95723	04277	13018	17295	49
12	22 24	37 3 6	82719	86970	95748	04252	13030	17281	48
13	22 16	37 44	82733	86959	95774	04226	13041	17267	47
14	22 8	37 52	82747	86947	95799	04201	13053	17253	46
15	6 22 0	5 38 0	9.82761	9.86936			10.13064	10.17239	45
16	21 52	3 8 8	82775	86924	95850	04150	13076	17225	44
17	21 44	38 16	82788	86913	95875	04125	13087 13098	17212	43
18	21 36	38 24	82802	86902	95901	04099 04074	13110	17198 17184	41
19	21 28	38 32	82816	86890	95926				
20	6 21 20	5 38 40	9.82830	9.86879		10.04048	10.13121		40
21	21 12	38 48	82844	86867	95977	04023 03998	13133 13145	17156 17142	39 38
22	21 4	38 56	82858	86855 86844	96002 96028	03998	13145 13156	17128	37
23	20 56	39 4 39 12	8 2872 8 28 85	86832	96053	03947	13168	17115	36
24	20 48						10.13179		35
25	6 20 40	5 39 20	9.82899	9.86821 86809	9.96078	03896	13191	17087	34
26	20 32	39 28 39 36	82913 82927	86798	96129	03871	13202	17073	33
27 28	20 24 20 16	39 36 39 44	8 294 1	86786	96155	03845	13214	17059	32
29	20 8	39 52	82955	86775	96180	03820	13225	17045	31
				9.86768		10 03795	10.13237	10.17039	30
30	6 20 0	5 40 0 40 8		86752	96231	03769	13248	17018	29
31 32	19 52 19 44	40 16	8 29 96	86740	96256	03744	13260	17004	28
33	19 36	40 24	83010	86728	96281	03719	13272	16990	27
34	19 28	40 32	83023	86717	96307	03693	13283	16977	26
35	6 19 20	5 40 40	9.83037	9.86705	9.96332	10.03668	10.13295	10.16963	25
36	19, 12	40 48	83051	86694	96357		13306	16949	24
37	19 4	40 56	83065	86682	96383	03617	13318	16935	23
38	18 56	41 4	83078	86 670	96408	03592	13330	16922	22
39	18 48	41 12	83092	86659	96433	03567	13341	16908	21
40	6 18 40	5 41 20	9.83106	9.86647	9.96459	10.03541		10.16894	20
41	18 32	41 28	83120	86635	96484	03516	13368	16880	19
42	18 24	41 36	83133	86624	96510		13376	16867	18
43	18-16	41 44	83147	86612	96535	03465	13388	16853	17
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45	6 18 0	5 42 0		9.86589			10.13411		1.5
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47	17 44	42 16		86565	96636	03364	13435	16798	13
48	17 36	42 24	83215	86554	96662	03338	13446	16785	13
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50	6 17 20	5 42 40		9.86530			10.13470		10
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M Houra.m., Houra.m. Sine. Co-sine. Tangent. Co-tang. Secant. Co-secant M			_						∠og. Sine	s, Teng	ents and	Secants.	•	_	
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41 10 32 49 28 83927 85924 98003 01997 14076 16073 19 42 10 24 49 36 83940 85912 98029 01971 14088 16060 18 43 10 16 49 44 83954 85900 98054 01946 14100 16046 17 44 10 8 49 52 83967 85888 98079 01921 14112 16033 16 45 6 10 0 5 50 0 9.83980 9.85876 9.98104 10.01896 10.14124 10.16020 15 46 9 52 50 8 88993 95864 98130 01870 14136 16007 14 47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85839 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.85815 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85803 98256 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.85815 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85803 98256 01794 14173 15967 11 55 9 4 50 56 84072 85791 98281 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 51 28 84125 85742 98383 01617 14258 15875 4 55 8 24 51 56 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 84151 85718 98433 01567 14222 15849 2 59 8 8 51 552 84164 85706 98458 01542 14294 15836 1 560 8 0 52 0 84177 85693 98484 01516 14307 15823 0		40	6	10	40	5	49	20	9.83914	9.85936	9.97978	10.02022	10.14064	10.16086	20
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43 10 16 49 44 83954 85960 98054 01946 14100 16046 17 44 10 8 49 52 83967 85868 98079 01921 14112 16033 16 45 6 10 0 5 50 0 9.83980 85876 98100 01870 14136 16047 14 47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85859 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 98451 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85803 98256 01744 14197 15941 9 52 9 4 50 56 84072 85791 98281 01719 14209 15928 8 53 8 56 51 4 84085 85799 98307 01693 14221 15915 7 54 8 48 51 12 84085 85769 98307 01693 14221 15915 7 55 6 8 40 5 51 28 84125 85764 98332 01668 14234 15902 6 55 6 8 40 5 51 28 84125 85742 98383 01617 14246 10.15888 5 56 8 48 51 12 84088 85769 98307 01693 14221 15915 7 55 8 24 51 56 84138 85730 98408 01592 14270 15862 3 58 8 16 51 44 84151 85718 98433 01567 14252 15849 2 59 8 8 51 52 84164 85706 98458 01542 14294 15835 1	ı		1		24	l				85912	98029				
44 10 8 49 52 83967 85888 98079 01921 14112 16033 16 46 6 10 5 60 9 83990 9.85876 9.98104 10.01896 10.14124 10.16020 15 47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85839 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.84815 9.98231 10.01769 10.14185 10.12954 10 51 9 12 50 48 840459 85815 9.98231 10.01769 10.14185 10.12954 10 52	l	43	l		16		49								
46 9 52 50 8 88993 85864 98130 01870 14136 16007 14 47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85839 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.85015 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85803 98256 01744 14197 15941 9 52 9 4 50 56 84072 85791 98221 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 51 20 9.84112 9.85754 9.98357 10.01643 10.14246 10.15888 5 56 . 8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01592 14270 15802 3 58 8 16 51 44 84151 85718 98433 01567 14224 15893 2 59 8 8 51 552 84164 85706 98458 01542 14294 15836 1	ll		1	10		1	49	52	83967	8568 8	98079	01921	14112	16033	16
46 9 52 50 8 88993 85864 98130 01870 14136 16007 14 47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85839 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.85015 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85803 98256 01744 14197 15941 9 52 9 4 50 56 84072 85791 98221 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 51 20 9.84112 9.85754 9.98357 10.01643 10.14246 10.15888 5 56 . 8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01592 14270 15802 3 58 8 16 51 44 84151 85718 98433 01567 14224 15893 2 59 8 8 51 552 84164 85706 98458 01542 14294 15836 1	l	46	6	10	- 0	5	50	0	9.83980	9.85876	9.98104	10.01896	10.14124	10.16020	15
47 9 44 50 16 84006 85851 98155 01845 14149 15994 13 48 9 36 50 24 84020 85859 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.84815 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85603 98266 01744 14197 15941 9 52 9 4 50 56 84072 86791 98261 01744 14197 15941 9 53 8 56 51 4 84085 85799 98307 01693 14221 15915 7 54 8 48 51 12 84088 85766 98332 01663 14231 15902 6 55 6 8 40 5 120	۱۱		۱					_							
48 9 36 50 24 84020 85839 98180 01820 14161 15980 12 49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.84915 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85603 98236 01744 14197 15941 9 52 9 4 50 56 84072 85791 98281 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 120 9.84112 9.85764 9.98357 10.01643 10.14246 10.15888 5 56 8 32 <td< th=""><th>ı</th><th></th><th>1</th><th></th><th></th><th>1</th><th></th><th></th><th></th><th></th><th></th><th></th><th>14149</th><th></th><th></th></td<>	ı		1			1							14149		
49 9 28 50 32 84033 85827 98206 01794 14173 15967 11 50 6 9 20 5 50 40 9.84046 9.84815 9.98231 10.01769 10.14185 10.15954 10 51 9 12 50 48 84059 85603 98256 01744 14197 15941 9 52 9 4 50 56 84072 85791 98281 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 51 20 9.84112 9.85754 9.98357 10.01643 10.14246 10.15888 5 56 8 32 51 28 84125 85742 98383 01617 14268 15875 4	1		ı				50								
51 9 12 50 48 84059 85603 98256 01744 14197 15941 9 52 9 4 50 56 84072 86791 98251 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14224 15902 6 55 6 8 40 5 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 84151 85718 98433 01667 14282 15849 2 59 8 8 51 52 84164 85706 98458 01642 14294 15836 1 60 8 0 52 0 84177 85693 98484	1	49	1	9	28	l	50	32	84033	85827	98206	01794	14173	15967	11
51 9 12 50 48 84059 85603 98256 01744 14197 15941 9 52 9 4 50 56 84072 86791 98251 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14224 15902 6 55 6 8 40 5 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 84151 85718 98433 01667 14282 15849 2 59 8 8 51 52 84164 85706 98458 01642 14294 15836 1 60 8 0 52 0 84177 85693 98484		50	6	9	20	5	50	40	9.84046	9.85815	9.98231	10.01769	10.14185	10.15954	10
52 9 4 50 56 84072 85791 98281 01719 14209 15928 8 53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 12 984125 85764 9.98357 10.01643 10.14246 10.15888 5 56 8 32 51 28 84125 85742 98383 01617 14245 10.15888 5 57 8 24 51 36 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 34151 85718 98433 01667 14282 15849 2 59 8 8 51	1		١			ľ		48							
53 8 56 51 4 84085 85779 98307 01693 14221 15915 7 54 8 48 51 12 84098 85766 98332 01668 14234 15902 6 55 6 8 40 5 51 28 84112 9.85754 9.98357 10.01643 10.14246 10.1588 5 56 8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01592 14270 15862 3 58 8 16 51 44 84151 85718 98433 01607 14282 15849 2 59 8 8 51 52 84164 85706 98458 91542 14294 15836 1 60 8 0 52 0 84177 85693 98484 01516 14307 15823 0	1		ł			1									8
55 6 8 40 5 5 20 9.84112 9.88754 9.98357 10.01643 10.14246 10.15888 5 56 . 8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01992 14270 15862 3 58 8 16 51 44 84151 85718 98433 01667 14282 15849 2 59 8 8 51 52 84164 85706 98458 01642 14294 15836 1 60 8 0 52 0 84177 85693 98484 01516 14307 15823 0	1			8	56	1	51	_		85779	98307	01693			
56 .8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 84151 85718 98433 01667 14282 15849 2 59 8 8 51 52 84164 85706 98458 01642 14294 15836 1 60 8 0 52 0 84177 85693 98484 01516 14307 15823 0	l	54		8	48	l	51	12	84098	85766	98332	01668	14234	15902	6
56 .8 32 51 28 84125 85742 98383 01617 14258 15875 4 57 8 24 51 36 84138 85730 98408 01692 14270 15862 3 58 8 16 51 44 84151 85718 98433 01667 14282 15849 2 59 8 8 51 52 84164 85706 98458 01642 14294 15836 1 60 8 0 52 0 84177 85693 98484 01516 14307 15823 0	l	55	6	8	40	ā	51	20	9.84112	9.85754	9.98357	10.01643	10.14246	10.15888	5
57 8 24 51 36 84138 85730 98408 01592 14270 15862 3 58 8 16 51 44 84151 85718 98433 01567 14282 15849 2 59 8 8 51 52 84164 85706 98458 91542 14294 15836 1 60 8 0 52 0 84177 85693 98484 91516 14307 15823 0	1		Ĭ			ľ								15875	
58 8 16 51 44 84151 85718 98433 01567 14282 15849 2 59 8 8 51 52 84164 85706 98458 91542 14294 15836 1 60 8 0 52 0 84177 85693 98484 91516 14307 15823 0	ll		ľ											15862	3
59 8 8 51 52 84164 85706 98458 91542 14294 15836 1 60 8 0 52 0 84177 85693 98484 91516 14307 15823 0	1		l		16									15849	
	, 1	59	Ī												
M Hours.m. Hours.m. Co-sine. Sine. Co-tang. Tangent Co-secant Secant. M	H	60		8	0		52	0	84177	85698	98484	01516	14307	15823	0
		M	Ho	urr	ж.	Ho	ULA	. M.	Co-sine.	Sine.	Co-tang.	Tangent	Co-secant	Secant.	M

133 Dega.

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TABLE XXVII.

	Dan	_					rog. Sir	ies, ran	genus am	u Secam	3.	Dam 1	95
	Deg			11.			Sine.	Conino	T	C- 4	Q	Degs. 1	
M	Ho			-					Tangent.			Co-secant	M
0	6	8	0	5		0	9.84177	9.85693		10.01516			60
1	ļ	7	52		52	8	84190 84203	85681	98509	01491	14319	15810	59
2	ł	7	44 36		52 52	16 24	84216	85669 85657	985 34 98 560	01466 01440	14331 14343	15797	58
3	l	7	28		52	32	84229	85645	98585	01415	14355	15784 15771	57 56
_	 -			-		40	9.84242	9.85632					
5	6	7	20 12	0	52 52	48	9.04242 84255	856 2 0	9.98610 98635	01365	10.14368 14380		55
6 7	1	7	4		52	56	84269	85608	98661	01339	14392	16745 16731	54 53
8	1	6	56		53	4	84282	85596	98686	01314	14404	15718	52
9	l	ő	48		53	12	84295	85583	98711	01289	14417	15705	51
10	6	6	40	-5	53	20	9.84308	9.85571		10.01263			50
ii	١ ٠	6	32	٠	53	28	84321	85559	98762	01238	14441	15679	49
12	1	6	24		53	36	84334	85547	98787	01213	14453	15666	48
13	l		16		53	44	84347	85534	98812	01188	14466	15653	47
14	1	6	8		53	52	84360	85522	98838	01162	14478	15640	46
15	6	6	7	5	54	0	9.84373	9.85510	9.98863	10.01137	10.14490	10.15627	45
16	١	5	52	-	54	8	84385	85497	98888	01112	14503	15615	44
17	1	5	44		54	16	84398	85485	98913	01087	14515	15602	43
18	I	5	36		54	24	84411	85473	98939	01061	14527	15589	42
19	}	5	28	_	54	32	8 442 4	85460	98964	01036	14540	15576	41
20	6	5	20	5	54	40	9.84437	9.85448	9.98989	10.01011	10.14552	10.15563	40
21	1	5	12		54	48	84450	85436	99015	00985	14564	15550	39
22	1	5	4		54	56	84463	85423	99040	00960	14577	15537	38
23	1	4	56		55	4	84476	85411	99065	00935	14589	15524	37
24		4	48		55	12	84489	85399	99090	00910	14601	15511	36
25	6	4	40	5	55	20	9.84502	9.85386	9.99116	10.00884		10.15498	35
26		4	32		55	28	84515	85374	99141	00859	14626	15485	34
27		4	24		55	36	84528	85 3 61	99166	00834	14639	15472	33
28	1	4	16		55	44	84540	85349	99191	00809	14651	15460	32
29		4	8		55	52	84553	85337	99217	00783	14663	15447	31
30	6	4	0	5	56	0	9.84566	9.85324		10.00758			30
31	1	3	52		56	8	84579	85312	99267	00733	14688	15421	29
32	1	3	44		56 56	16 24	84592	85 299 85287	99293 99318	00707	14701	15408	28
33 34		3 3	36 28		56	32	84605 84618	85274	99318	00682 00657	14713 147 2 6	1 539 5 1 53 82	27 26
	<u> -</u>			_									
35	6	3	20 12	5	56 56	40	9.84630 84643	9.85262			10.14738		25
36	1	3	4		56	48 56	84656	85250 85237	99394 99419	00606 00581	14750	15357 15344	24 23
37 38	1	2	56		57	4	84669	852 2 5	99444	00556	14763 14775	15331	23
39	1	2	48		57	12	84682	85212	99469	00531	14788	15318	21
40	6	- 2	40	5	57	20	9.84694	9.85200					
	ľ	2	32	9	57	20 28	84707	9.85200 85187	9.99496	00480	14813	15293	20 19
41 42	ļ	2	24		57	36	84720	85175	99545	00455	14825	15293	18
43	i	2	16		57	44	84733	85162	99570	00430	14838	15267	17
44	i	2	8	l	57	52	84745	85150	99596	00404	14850	15255	16
45	6	2	-0	5	58	 0	9.84758	9.85137	9.99621	10.00379	10.14863		15
46	1 "	1	52	_	58	8	84771	85125	99646	00354	14875	15229	14
47	ı	î	44		58	16	84784	85112	99672	00328	14888	15216	13
48	1	1	36		58	24	84796	85100	99697	00803	14900	15204	12
49	1	1	28		<i>5</i> 8	32	84809	85087	99722	00278	14913	15191	11
50	6	1	20	5	58	40	9.84822	9.85074	9.99747	10.00253	10.14926	10.15178	10
51	١	î	12	_	58	48	84835	85062	99773	00227	14938	15165	9
53	1	1	4		58	56	84847	85049	99798	00202	14951	15153	8
53	1	0	56	ļ	59	4	84860	85037	99823	00177	14963	15140	7
54	L	0	48		59	12	84873	85024	99848	00152	14976	15127	6
55	6	0	40	5	59	20	9.84885	9.85012	9.99874	10.00126	10.14988	10.15115	5
56	l	0	32		5 9	28	84898	84999	99899	00101	15001	15102	4
57	l	0	24		59	36	84911	84986	99924	00076	15014	15089	3
58	ı	0	16		59	44	84923	84974	99949	00051	15026	15077	2
59 50	ı	0	8	_	59	52	84936	84961	99975	00025	15039	15064	I
60	<u> </u>	0	0	6	0	0	84949		10.00000	00000	15051	15051	0
M	Ho	UPP	.M.	Ho	NIL	.м.	Co-sine.	Sine.	Co-tang.	Tangent.	Co-secant	Secant.	M

134 Degs.

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TABLE XXVIII. For reducing the time of the Moon's passage over the Meridian of Greenwich to the Time of its passage over any other Meridian.

of its passage over any other Meridian.

The numbers taken from this Table are to be added to the Time at Greenwich in West Longitude, but subtracted in East.

TABLE XXIX.

Correction of moon's altitude for Parallax and Refraction.

																	Corr		Con
	Ľ	aily	var	iatio	on of	the	Mo	on's	pas	sing	the	Me	ridia	m.	į	Deg.	Min.	Deg.	Min
Sh's	1	7	17	17	17	7	7	7	71	7	7	1	1	17	Sh's	10	51	51	35
Lon.	40	42	44	46	48	50	52	54	56	58	60	62	64	66	Lon.	11	52	52	35
0	 -	7	17	7	17	7	7	-	7	7	7	7	7	17	0	12	52	53	34
0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	13	52	54	33
5	ĭ	ĭ	ĭ	li	ľ	i	1	i	1	1	ĭ	1	1	ĭ	5	14	52	55	32
10	li	î	l i	li	١i.	l î	ī	i i	2	2	2	2	2	2	10	15 16	52 52	56 57	32 31
15	2	2	2	2	2	2	2	2	2	2	2	3	3	3	15	17	52	58	30
20	2	2	2	3	3	3	3	3	3	3	3	3	4	4	20	18	52	59	29
25	3	3	3	3	3	3	4	4	4	4	4	4	4	5	25	19	52	60	28
30	3	3	4	4	4	4	4	4	5	5	5	5	5	5	30	20	51	""	
35	4	4	4	4	5	5	5	5	5	6	6	6	6	6	35	21	51	61	27
40	4	5	5	5	5	6	6	6	6	6	7	7	7	7	40	22	51	62	26
45	5	5	5	6	6	6	6	7	7	7	7	8	3	8	45	23	51	63	26
50		6	6	6	7	7	7	7	8	8	8	9	9	9	50	24	50	64	25
55	6	6	7	7	7	8	8	8	9	9	9	9	10	10	55	25	50	65	24
60	7	7	7	8	8	8	9	9	9	10	10	10	11	11	60	26	50	66	23
65		8	8	8	9	9	9	10	10	10	lii	11	12	12	65	27	49	67	22
70		8	9	9	9	10	10	10	11	11	12	12	12	13	70	28	49	68	21
75	8	9	9	10	10	10	11	11	12	12	12	13	13	14	75	29	49	69	20
80	9	9	10	10	11	11	12	12	12	13	13	14	14	15	80	30	48	70	19
85	9	10	10	11	11	12	12	13	13	14	14	15	15	16	85	31	48	71	18
	10	10	11	11	12	12	13	13	14	14	15	15	116	16	90	32	47	72	17
95	111	11	12	12	13	13	14	14	15	15	16	-16	17	17	95	33	47	73	17
100	11	12	12	13	13	14	14	15	16	16	17	17	18	18	100	34	46	74	16
105	12	12	13	13	14	15	15	16	16	17	17	,18	19	19	105	35	46	75	18
110	12	13	13	14	15	15	16	16	17	18	18	19	20	20	110	36	45	76	14
115	13	13	14	15	15	16	17	17	18	19	19	20	20	21	115	37	45	77	13
120	13	14	15	15	16	17	17	18	19	19	20	21	21	22	120	38	44	78	19
125		15	15	16	17	17	18	19	19	20	21	22	22	23	125	39 40	44	79	11
130	14	15	16	17	17	18	19	19	20	21	22	22	23	24	130	11-	43	80	10
135	15	16	16	17	18	19	19	20	21	22	22	23	24	25	135	41	42	81	1 5
140		16	17	18	19	19	20	21	22	23	23	24	25	26	140	42	42	82	1 5
145		17			19	20	21	22	23	23	24	25	26	27	145	43	41	83	13
150		17			20	21	22	22	23	24	25	26	27		150	44	40	84	1 9
155	17	18	19	20	21	22	22	23	24	25	26	27	28	28	155	46	40 39	85	1 3
160	18	19	20	20	21	22	23	24	25	26	27	28	28	29	160	47	38	86	
165		19			22	23	24	25	26	27	27	28			165	48	38	88	}
170		20	21	22	23	24	25	25	26	27	28	29			170	49	37	89	
175		20		22			25	26	27	28	29				175	50	36	90	16
180	20	21	22	23	24	25	26	27	28	29	30	31	32	33	180	1		, 50	' '
	40	42	7 44	46	7 48	50	52	54	56	58	60	62	64	66	/	1			

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TABLE XXX.
or reducing the Moon's Declination, as given in the Nautical Almanac for Noon and
Midnight at Greenwich, to any other time under any other Meridian.

Time	1	Va	iatio	a of t	he M	oon's	Dec	inati	on in	twel	ve H	ours.	0110303	Time	ī
from Noon.	0 5		0 /	0 /	0 1	0 7	0 /	0 7		0 1		0 1	1 5	from Noon.	
Oh O'	O	0'	9	0	0/	0'	0'	0'	0'	9'	0'	0'	0° 0′ 0 1	12h 0	1
0 12 0 24	0	0	0	0	0	0	1	1	1	1 2	2	1 2	0 2	12 12 12 24	l
0 86	0	0	1	1	1	1	2	2	'2	2	3	3	0 3	12 36	l
0 48	0	1	1	1	2	2	2	3	3	3	5	4	0 4	12 48	-
1 0 1 12	0	1	1	2 2	2 2	.8	3	4	4	4 5	5	5	0 6	13 0 13 12	1
1 24	1	1	2	2	3	3	4	5	5	6	6	7	0 8	13 24	l
1 36 1 48		1	2 2	3	3 4	4	5	5 6	6	7	8	8	0 9	13 36 13 48	1
2 0	i	2	2	3	4	5	6	7	7	8	9	10	0 11	14 0	1
2 12 2 24	1	2	3	4	5	5 6	6	7 8	8	9	10 11	11 12	0 12 0 13	14 12 14 24	1
2 24 2 36	1	2 2	3	4	5 5	6	8	9	9 10	ii	12	13	0 14	14 36	l
2 48	1	2	3	5	6	7	8	9	10	12	13 14	14 15	0 15	14 48	١
3 0	$\frac{1}{1}$	3	4	5	7	8	9	10	112	12	15	16	0 16	15 0	-
3 24	1	3	4	6	7	8	10	11	13	14	16	17	0 18	15 24	
3 36 3 48	1 2	3	4 5	6	8	9	10	12 13	13 14	15 16	16 17	18 19	0 19	15 36 15 48	
4 0	2	3	5	7	8	10	12	13	15	17	18	20	0 22	16 0	1
4 12	2	3	5	7	9	10	12	14	16	17	19	21	0 23	16 12	1
4 24 4 36	2 2	4	6	8	9 10	11 11	13 13	15 15	16 17	18 19	20 21	23	0 24	16 24 16 36	Ī
4 48	2	4	6	8	10	12	14	16	18	20	22	24	0 26	16 48	ı
5 0	2	4	6	8	10	12	15	17	19	21	23	25 26	0 27	17 0	ł
5 12 5 24	2 2	4	6	9	11 11	13	15 16	17 18	19 20	22 22	24 25	27	0 28	17 12 17 24	1
5 36	2	5	7	9	12	14	16	19	21	23	26	28	0 30	17 36	ı
5 4 8 6 0	2 2	8	7	10 10	12 12	14	17	19 2 0	22	24 25	27	29	0 31	17 48 13 0	1
6 12	3	5	8	10	13	15	18	21	23	26	28	31	0 34	18 12	1
6 24	3	5	8	11	13	16	19	21 22	24	27	29 30	32 33	0 35	18 24	١
6 3 6 6 4 8	3 3	6	8	11	14 14	16 17	19 20	23	25 25	27 28	31	34	0 36	18 36 18 48	l
7 0	3	6	9	12	15	17	20	23	26	29	32	35	0 38	19 0	1
7 12 7 24	3	6	9	12 12	15 15	18 18	21 22	24 25	27 28	30 31	33 34	36 37	0 5 9	19 12 19 24	l
7 36	3	6	9	13	16	19	22	25	28	32	35	38	0 41	19 36	1
748 80	3	6	10 10	13 13	16 17	19 20	23 23	26 27	29 30	32 33	36 37	59 40	0 42	19 48 20 0	1
8 12	3	7	10	14	17	20	24	27	31	34	38	41	0 44	20 12	1
8 24	3	7	10	14	17	21	24	28	31	35	38	42	0 45	20 24	Ì
8 36 8 48	4	7	11 11	14 15	18 18	21	25 26	29 29	32 33	36 37	39 40	43	0 47	20 36 20 48	
9 0	4	7	11	15	19	22	26	30	34	57	41	45	0 49	21 0	
9 12 9 24	4	8	11 12	15 16	19 2 0	23 23	27	31 31	34 35	38 39	42	46 47	0 50	21 12 21 24	
9 36	4	8	12	16	20	24	28	32	36	40	44	48	0 52	21 36	1
9 4 8 10 0	4	8	12 12	16	20	24 25	29	33 33	37	41	45	49	0 53 0 54	21 48	
10 12	4	8	13	17	21	25	29 30	34	37 38	42	46	50 51	0 54	22 0	-
10 24	4.	9	13	17	22	26	30	35	39	43	48	52	0 56	22 24	1
10 36 10 48	4	9	13 13	18 18	22	26 27	31 31	35 36	40 40	44 45	49 49	53 54	0 57 0 58	22 36 22 48	
11 0	5	9	14	18	23	27	32	37	41	46	50	55	1 0	23 0	
11 12	5	9	14	19	23	28	33	37	42	47	51	56	1 1	23 12	
11 24 11 36	5	9 10	14	19 19	24 24	28 29	33 34	38 39	43 43	47	52 53	<i>5</i> 7	1 2	23 24 23 36	1
11 48	5	10	15	20	25	29	34	39	44	49	54	59	1 4	23 48	
12 0	5	10	15	20	25	30	35	40	45	50	55	60	1 5	24 0	e

TABLE XXX

For reducing the Moon's Declination, as given in the Nautical Almanac for Noon a Midnight at Greenwich, to any other time under any other Meridian.

	Mid	dnight	at Gre	enwic	b, to	any ot	her tin	e und	er any	other	Merid	ian.
	Time	V	ariatio	n of th	e Mo	on's D	eclinat	ion in	twelve	How	rs.	Time
	from	0 /	1	_		1		0 /	0 /	0 /	0 /	from
	Noon.	1 10	1 15	1 20	1 25		1 35	1 40	است	1 50	1 55	Noon.
	0h 0'	00 O	-	00 0	00 O		00 O	00 O	1 7 1	00 O	00 0	12h 0'
	0 12	0 1	0 1	0 1	0 1		0 2	0 2		0 2	0 2	12 12
	0 24	0 2	0 2	0 3	0 3		0 3	-		0 4	0 4	12 24 12 36
	0 36 0 48	0 3	0 5	0 5	0 6						0 8	12 48
	1 0	0 6	0 6	0 7	0 7		0 8	0 8	!		0 10	13 0
	1 12	0 7	0 7	0 8	0 8		0 9			0 11	0 11	13 12
1	1 24	0 8	0 9				0 11	0 12	0 12	0 13	0 13	13 24
	1 36			0 11	0 11						0 15	13 36
	1 48	0 10	0 11	0 12	0 13			0 15	0 16 0 17	0 16 0 18	0 17 0 19	13 48 14 0
	2 0	0 12		0 13			-	0 17			0 21	
	2 12 2 24	0 13	0 14	0 15 0 16	0 16 0 17		0 17 0 19	0 20		0 20 0 22	0 23	14 12 14 24
	2 36				0 18			0 22		0 24	0 25	14 36
	2 48	0 16	0 17	0 19	0 20			0 23		0 26	0 27	14 48
	3 0	0 17	0 19	0 20	0 21	0 22	0 24	0 25			0 29	15 0
	3 12	0 19		0 21	0 23			0 27			0 31	15 12
	3 24	0 20	0 21	0 23	0 24	0 25	0 27				0 33 0 34	15 24
	3 36		0 22	0 24 0 25	0 25 0 27		0 28 0 30	0 30		0 3 3 1 0 3 5	0 34 0 36	15 36 15 48
	3 48 4 0		0 25		0 28						0 38	16 0
	4 12	0 24	0 26	0 28	0 30			0 35		9 38	0 40	16 12
	4 24	0 26	0 27					0 37	0 38		0 42	16 24
	4 36			0 31	0 33	0 34	0 36	9 38		0 42	0 44	16 36
	4 48	0 28	0 30		0 34					0 44	0 46	16 48
	5 0	0 29	0 31					0 42		0 46	0 48	17 0
	5 12	0 30	0 32		0 37 0 38			0 43 0 45		0 48 0 49	0 50 0 52	17 12 17 24
	5 24 5 36	0 31 0 33					0 44	0 47			0 54	17 36
	5 48	0 34	0 36	1			0 46	0 48			0 56	17 48
	6 0	0 35	0 37	0 40	0 42	0 45	0 47	0 50			0 57	18 0
	6 12	0 36	0 39		0 44			0 52			0 59	18 12
	6 24	0 37	0 40			0 48				0 59 1 0	1 1 1 3	18 24 18 36
	6 36 6 48	0 38		0 44	0 47					1 2	1 3	18 48
	7 0				0 50			0 58		1 4	1 7	19 0
	7 12	0 42	0 45	0 48	0 61		0 57	1 0	1 3	1 6	1 9	19 12
	7 24	0 43			0 52		0 59			18	1 11	19 24
	7 36				0 54	0 57	1 0	1 3		1 10	1 13	19 36
	7 48	0 45	1	0 52	0 55	0 58	1 2	1 5		1 11 1 13	1 15 1 17	19 48 20 0
	8 0	0 47	0 50		0 57	1 0				1 15	1 19	20 12
l	8 12 8 24	0 48	0 51 0 52	0 55 0 56	0 58 0 59	1 1	1 5	1 8		1 17	1 20	. 20 24
Ī	8 36	0 60			1 1	1 4	1 8	1 12		1 19	1 22	20 36
l	8 48	0 51	0 55	0 59	1 2	1 6	1 10	1 13	1 1	1 21	1 24	20 48
	9 0	0 52	0 66	1 0	1 4		1 11	1 15		1 22	1 26	21 0
ŀ	9 12	0 54	0 57	1 1	1 5	1 9	1 13	1 17	1 20	1 24	1 28 1 30	21 12 21 24
ı	9 24	0 55	0 59	1 3	1 7	1 10 1 12	1 14	1 18		1 2 6 1 2 8	1 32	21 36
İ	9 56	0 56 0 57	1 0	1 5	1 8	1 13	1 18	1 22		1 30	1 34	21 48
	10 0	0 58	1 2		1		1 19	1 23		1 32	1 36	22 0
	10 12	0 59	1 4	1 8	1 12	1 16					1 38	22 12
l	10 24								1 31	1 35	1 40	22 24
	10 36								1 33		1 42	22 36 22 48
l	10 48 11 0	1 3		1 12			1 25	1 30		1 41	1 45	23 0
	11 12	1 5				1 24	1 29	1 33			1 47	23 12
1	11 24	1 6	iii	1 16	i 21	i 25	1 30	1 35	1 40	1 44	1 49	23 24
	11 36	1 8	1 12	1 17	1 22	1 27	1 32	1 57		1 46	1 51	23 36
	11 48	1 9		1 19	1 24		1 33			1 48	1 53	23 48
	12 0	1 10	1 15	1 20	1 25	1 30	1 35	1 40	1 45	1 50	1 55	24 0

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For reducing the Moon's Declination, as given in the Nautical Almanac for Noon and Midnight at Greenwich, to any other time under any other Meridian.

	Time	1	Vari	ation of th	Moon	's Declination	in twelve Hours.	Time
No. No.	_	0 7						
0 24 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 5 0 2 0 2 0 2 0 2 0 3 0 3 0 3 0 3 0 3 2 2 2 4 0 36 0 36 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 7 0 7	Noon.	2 0	2 5	2 10 2 1	2 20	2 25 2 30	2 35 2 40 2 45 2	
0 24 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 4 0 5 0 2 0 2 0 2 0 2 0 3 0 3 0 3 0 3 0 3 0 3	Oh O	00 0/	00 0	00 0/ 00	100 0	100 N 00 W	00 00 00 00 00 00	0 0 19h O
0 48 0 8 0 8 0 9 0 9 0 9 0 9 0 10 0 10 0 10		0 2	0 2	0 2 0			0 3 0 3 0 3 0	
0 48 0 8 0 8 0 9 0 9 0 9 0 10 0 10 0 10 0 11 0 11 0 11 0 12 12 0 12 </td <td>0 24</td> <td>0 4</td> <td>0 4</td> <td>0 4 0</td> <td>0 5</td> <td>0 5 0 5</td> <td>0 5 0 5 0 5 0</td> <td>6 12 24</td>	0 24	0 4	0 4	0 4 0	0 5	0 5 0 5	0 5 0 5 0 5 0	6 12 24
1 0 0 10 0 10 0 11 0 11 0 11 0 12 0 12	0 36		0 6	0 6 0 .	7 0 7	0 7 0 7		
1 2	0 48	0 8	0 8	0 9 0	0 9	0 10 0 10	0 10 0 11 0 11 0	11 12 48
1 12 0 12 0 12 0 13 0 13 0 14 0 14 0 15 0 15 0 16 0 16 0 17 1 3 12 1 24 0 14 0 15 0 15 0 16 0 16 0 17 0 18 0 19 0 19 0 20 0 20 1 3 24 1 3 6 0 16 0 17 0 17 0 18 0 19 0 19 0 20 0 20 1 3 24 0 25 0 25 1 3 3 6 1 48 0 18 0 19 0 19 0 20 0 21 0 22 0 22 0 23 0 24 0 25 0 26 0 27 0 27 0 28 1 3 36 1 48 0 18 0 19 0 19 0 20 0 21 0 22 0 22 0 23 0 24 0 25 0 26 0 27 0 27 0 28 1 4 0 2 2 0 22 1 2 2 2 2 2 2 2 2 2 3 3 2 4 0 25 0 25 1 3 48 1 2 2 2 4 0 24 0 25 0 26 0 27 0 27 0 28 1 4 0 0 25 0 2 2 2 2 2 2 2 2 2 3 3 0 2 4 0 25 0 26 0 27 0 27 0 28 1 4 0 0 2 2 2 2 2 2 2 2 2 2 3 3 2 4 0 2 5 0 2 6 0 27 0 27 0 28 1 4 0 0 2 2 2 2 2 2 2 2 2 2 2 2 3 3 2 2 3 2 2 4 0 2 5 0 2 6 0 27 0 27 0 28 1 4 0 0 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	1 0	0 10	0 10	0 11 0 1	0 12	0 12 0 12	0 13 0 13 0 14 0	14 13 0
1 1 2 2 2 2 2 2 2 2	1 12		0 12	0 13 0 1	0 14	0 14 0 15	0 15 0 16 0 16 0	
2 48 0 18 0 19 0 19 0 20 0 21 0 22 0 22 0 22 0 23 0 24 0 25 0 25 13 48 2 2 12 0 22 0 22 0 22 0 22 0 22 0 22 0								
2 0 0 20 0 21 0 22 0 22 0 25 0 25 0 24 0 25 0 26 0 27 0 27 0 28 14 0 0 2 12 2 0 22 0 23 0 24 0 25 0 26 0 27 0 27 0 28 0 29 0 30 0 31 14 12 2 2 36 0 26 0 27 0 28 0 29 0 30 0 31 0 32 0 33 0 33 1 32 0 33 0 34 14 24 2 36 0 26 0 27 0 28 0 29 0 30 0 31 0 32 0 34 0 35 0 36 0 37 1 38 0 40 0 41 36 2 2 48 0 28 0 29 0 30 0 31 0 33 0 34 10 35 0 36 0 37 0 38 0 40 14 36 2 3 3 12 0 32 0 33 0 35 0 36 0 37 0 38 0 0 31 0 35 0 36 0 37 0 38 0 40 14 48 3 3 0 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 44 0 0 41 0 42 15 0 0 3 12 0 32 0 33 0 35 0 36 0 37 0 38 0 40 0 41 0 43 0 44 0 45 15 0 0 3 12 0 33 0 33 0 35 0 36 0 37 0 38 0 40 0 41 0 43 0 44 0 45 0 47 0 48 15 24 3 36 0 36 0 37 0 38 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 3 36 0 36 0 37 0 38 0 40 0 41 0 43 0 44 0 45 0 47 0 48 15 24 4 4 0 44 0 45 0 47 0 48 15 24 4 4 0 44 0 45 0 47 0 48 15 24 4 2 0 44 0 44 0 45 0 47 0 48 15 24 4 2 0 44 0 44 0 45 0 47 0 48 15 24 4 2 0 44 0 44 0 45 0 47 0 48 15 24 4 2 0 44 0 44 0 45 0 47 0 48 0 50 0 52 0 53 0 55 0 57 16 0 0 4 1 0 43 0 44 0 45 0 47 0 48 15 24 4 2 0 44 0 44 0 45 0 47 0 48 0 50 0 52 0 53 0 55 0 57 16 0 0 4 1 0 43 0 44 0 45 0 47 0 48 0 50 0 52 0 54 0 56 0 58 0 59 0 57 16 0 0 4 1 0 43 0 44 0 45 0 47 0 48 0 50 0 52 0 54 0 56 0 58 0 59 0 57 16 0 0 54 0 54 0 54 0 54 0 55 0 57 0 59 1 0 51 0 52 0 54 0 54 0 54 0 54 0 54 0 55 0 57 0 59 1 0 1 2 1 2 1 2 1 2 1 2 1 2 1 2 1 2 1 2		1						
2 12 0 22 0 23 0 24 0 25 0 26 0 27 0 28 0 29 0 30 0 31 0 32 0 34 14 24 2 34 6 28 0 29 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 14 48 3 0 30 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 14 48 3 0 30 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 14 48 3 3 0 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 14 48 3 3 0 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 14 48 3 3 6 0 38 0 37 0 38 0 34 0 35 0 36 0 37 0 38 0 40 0 41 0 42 0 44 0 44 0 45 0 47 0 48 15 24 0 44 0 44 0 44 0 44 0 44 0 44 0 44								
2 24	2 0	0 20	0 21	0 22 0 2	0 23	0 24 0 25	0 26 0 27 0 27 0	28 14 0
2 36 0 26 0 27 0 28 0 29 0 30 0 31 0 32 0 34 0 35 0 36 0 37 14 36 2 48 0 28 0 29 0 30 0 31 0 33 0 34 0 35 0 36 0 37 0 38 0 41 0 42 13 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 38 0 40 0 41 0 42 13 0 32 0 34 0 35 0 36 0 37 0 38 0 40 0 41 0 42 13 0 32 0 34 0 35 0 36 0 37 0 38 0 40 0 41 0 42 13 0 34 0 35 0 36 0 37 0 39 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 0 36 0 37 0 39 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 0 44 0 45 0 47 0 49 0 51 0 52 0 54 0 56 0 58 0 37 0 39 0 40 0 41 0 42 0 44 0 45 0 48 0 49 0 51 5 36 3 48 0 38 0 40 0 41 0 43 0 44 0 46 0 47 0 49 0 51 0 52 0 54 15 48 4 0 40 0 0 42 0 43 0 44 0 46 0 47 0 49 0 51 0 52 0 54 15 48 4 0 40 0 0 42 0 43 0 44 0 46 0 47 0 49 0 51 0 52 0 54 15 48 4 50 0 46 0 48 0 60 0 52 0 54 0 56 0 57 0 59 1 0 1 2 16 2 16 2 4 24 0 44 0 46 0 48 0 60 0 52 0 54 0 56 0 58 1 0 0 1 2 1 4 1 6 6 1 8 16 36 4 48 0 48 0 50 0 52 0 54 0 56 0 58 1 1 1 3 1 5 1 8 1 10 1 12 1 15 1 17 1 19 1 11 1 17 0 1 1 2 1 5 1 6 36 6 48 0 58 1 0 1 3 1 5 1 8 1 10 1 12 1 15 1 17 1 19 1 11 1 17 0 1 1 2 1 5 1 6 1 7 1 10 1 1 2 1 1 4 1 1 1 1 1 1 1 1 1 1 1 1 1 1	2 12	0 22	0 23	0 24 0 2	0 26	0 27 0 27	0 28 0 29 0 30 0	31 14 12
2 48 0 28 0 29 0 30 0 31 0 32 0 34 0 35 0 36 0 37 0 39 0 40 0 41 14 48 3 3 0 0 38 0 31 0 32 0 33 0 35 0 36 0 37 0 39 0 40 0 41 0 42 15 0 3 2 3 2 4 0 34 0 35 0 37 0 38 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 3 3 6 0 36 0 37 0 39 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 3 3 6 0 36 0 37 0 39 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 4 0 44 0 45 0 47 0 49 0 51 0 52 0 54 0 56 0 58 0 51 15 36 48 4 0 0 40 0 42 0 43 0 44 0 46 0 48 0 49 0 51 0 52 0 55 0 55 15 5 48 4 0 0 40 0 42 0 44 0 45 0 47 0 49 0 51 0 52 0 55 0 55 16 5 0 57 16 0 0 4 12 0 42 0 44 0 45 0 47 0 49 0 51 0 52 0 55 0 55 16 5 0 57 16 0 0 4 12 0 42 0 44 0 45 0 47 0 49 0 51 0 52 0 55 0 55 10 57 16 0 0 4 12 0 42 0 44 0 45 0 48 0 49 0 51 0 53 0 55 0 57 0 59 1 0 1 2 16 24 4 3 45 0 48 0 48 0 60 0 52 0 54 0 56 0 58 0 57 0 59 1 0 1 2 16 24 4 48 0 48 0 60 0 52 0 54 0 56 0 58 1 0 1 2 1 5 1 7 1 9 1 11 1 17 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1	2 24	1		0 26 0 2	7 0 28	0 29 0 30		34 14 24
3 0 0 80 0 31 0 82 0 34 0 85 0 36 0 37 0 39 0 40 0 41 0 42 15 0 31 12 0 32 0 33 0 35 0 36 0 37 0 39 0 40 0 41 0 42 0 43 0 44 0 45 0 47 0 48 15 24 3 36 0 36 0 37 0 39 0 40 0 41 0 42 0 43 0 44 0 46 0 47 0 48 15 24 3 36 0 36 0 37 0 39 0 40 0 42 0 43 0 45 0 46 0 48 0 49 0 51 15 36 3 48 0 38 0 40 0 41 0 43 0 44 0 46 0 47 0 49 0 51 0 52 0 54 15 48 4 0 0 40 0 42 0 43 0 45 0 47 0 49 0 51 0 52 0 54 15 48 4 0 0 40 0 42 0 43 0 45 0 47 0 49 0 51 0 52 0 54 15 48 4 0 0 40 0 42 0 43 0 45 0 47 0 49 0 51 0 52 0 54 15 48 4 0 40 0 40 0 42 0 43 0 45 0 47 0 49 0 51 0 52 0 53 0 53 0 55 0 57 16 0 0 40 4 12 0 44 0 46 0 48 0 49 0 51 0 52 0 54 0 56 0 58 1 0 1 2 1 4 1 6 1 8 16 48 5 0 0 50 0 52 0 54 0 56 0 58 1 0 1 2 1 4 1 6 1 8 16 48 5 0 0 50 0 52 0 54 0 56 0 58 1 0 1 2 1 4 1 6 1 8 16 48 5 0 0 50 0 52 0 54 0 56 0 58 1 0 1 2 1 5 1 7 1 9 1 11 17 0 0 1 5 12 0 50 0 50 0 52 0 54 0 56 0 58 1 0 1 2 1 1 5 1 7 1 9 1 11 17 0 0 1 5 12 0 5 0 50 0 52 0 54 0 56 0 58 1 1 1 1 3 1 5 1 8 1 10 1 12 1 15 1 17 1 19 1 17 0 0 1 5 12 1 5 1 7 1 10 1 12 1 15 1 17 1 19 1 17 1 17 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1								
3 12 0 32 0 33 0 35 0 36 0 37 0 38 0 40 0 41 0 43 0 44 0 45 0 47 0 48 15 24 3 36 0 36 0 37 0 38 0 40 0 41 0 42 0 44 0 46 0 47 0 48 15 36 3 48 0 38 0 40 0 41 0 43 0 44 0 46 0 47 0 49 0 51 0 52 0 54 15 36 4 4 0 46 0 47 0 49 0 51 0 52 0 54 15 36 4 4 12 0 42 0 44 0 46 0 47 0 48 0 50 0 52 0 54 0 56 0 58 0 57 16 0 1 4 12 4 0 44 0 46 0 48 0 49 0 51 0 52 0 54 15 48 4 0 44 0 44 0 44 0 44 0 44 0 44 0 4				1				
3 24 0 34 0 35 0 36 0 37 0 38 0 40 0 41 0 42 0 44 0 45 0 47 0 48 15 24 3 3 6 0 36 0 37 0 39 0 40 0 42 0 43 0 44 0 46 0 47 0 48 0 45 0 0 51 15 36 3 48 0 38 0 40 0 41 0 43 0 44 0 46 0 47 0 48 0 50 0 52 0 53 0 55 0 67 16 0 0 40 0 42 0 43 0 45 0 47 0 48 0 50 0 52 0 54 0 56 0 58 0 59 16 12 4 24 0 44 0 46 0 48 0 49 0 51 0 53 0 55 0 57 0 59 1 0 1 2 16 24 4 36 0 46 0 48 0 49 0 51 0 53 0 55 0 57 0 59 1 0 1 2 16 24 4 36 0 46 0 48 0 49 0 51 0 53 0 55 0 57 0 59 1 0 1 2 16 24 4 36 0 46 0 48 0 50 0 52 0 54 0 56 0 58 1 0 1 2 1 4 1 6 1 8 16 48 5 0 0 50 0 52 0 54 0 56 0 58 1 1 1 3 1 5 1 7 1 10 1 12 1 15 1 7 1 19 1 11 1 77 0 5 12 0 52 0 54 0 56 0 58 1 1 1 3 1 5 1 5 1 7 1 10 1 12 1 15 1 17 1 19 1 11 1 70 0 5 12 0 50 0 50 0 52 0 54 0 56 0 58 1 1 1 3 3 1 5 1 6 36 0 48 0 58 1 0 1 2 1 1 5 1 7 1 10 1 12 1 15 1 17 1 19 1 17 36 0 1 1 2 1 1 5 1 17 1 19 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	3 0	0 30	0 31	0 32 0 3	0 35	0 36 0 37	0 39 0 40 0 41 0	42 15 0
S 36	3 12	0 32	0 33	0 35 0 3	0 37	0 39 0 40	0 41 0 43 0 44 0	45 15 12
S	3 24	0 34	0 35	0 37 0 3	0 40	0 41 0 42	0 44 0 45 0 47 0	
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TABLE XXX.

For reducing the Moon's Declination, as given in the Nautical Almanac for Noon a Midnight at Greenwich, to any other time under any other Meridian.

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1 86	0 23			0 25		0 27		0 28 0 29	0 29 0 30	13 36
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2 24	0 35			0 38	r	0 40	0 41	0 42 0 43	0 44 0 45	14 24
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3 12	0 47		0 49	0 51	0 52	0 53	0 55	0 56 0 57	0 59 1 0	15 12
3 24	0 50	1	0 52	0 84	0 55	0 57	0 58	0 59 1 1	1 2 1 4	15 24
3 36	0 52	0.54	0 55	0 57	0 58	1 0	1 1	1 3 1 4	1 6 1 7	15 36
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11 36	2 52	2 57	3 2	3 7	3 12	3 10 3 13 3 17	3 22	3 26 3 31	3 33 3 37 3 36 3 41	23 45
12 0	2 55	· · · · · · · · ·		3 10		3 20	3 25	3 30 3 35	3 40 3 45	24 (
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TABLE XXXI.

For reducing the Sun's Right Ascension in Time, as given in the Nautical Almanac for Noon at Greenwich, to any other time under any other Maridian

				t Gr												eridi	m.		
Time		1	Dail	y Var	riati	on of	the	Sun	's R	ight A	Asce	nsio	n in	Tim.	B.		,]	er:
from Noon.	3	30	3	32	3	34	3	36	3	38	3	40	3.	42	3	44	3		Ship's Long
Oh O'	0	0//	0	0//	O'	0//	0	Q/	07	0"	.oʻ	0"	0'	0"	O'	0"	0	0"	1-0-
0 12	0	2.	0	2	0	2	0	2	0	2	0	2	0.	2	0	2	0	2	3
0 24 0 36	0	. 3 5	0	5	0	4	0	4 5	0	4 5	0	14.1 5	0	6	0	6	0	4 6	6
0 48	0	7	0	7	0	7	0	7	0	7	0	7	0	7	0	. 7	0	8	12
1 0	0	- 9	0	9	0	.9	0	9	0	9	0	9	0	9	0	9	0	9	1.5
1 12 1 24	0	10 12	0.	11 12	0	11	0	11 13	0	11	0	11 13	0	13	0	11	0	11	13 21
1.36	ŏ	14	0	.14	ŏ	14	ŏ	14	o	15	ø	15	ŏ	15	·o	15	ő	15	24
1 48 2 0	0	16 17	0	16 18	0	16	0	16	Q	16	0	16	0	17	0.	17	0	17	27
2 0	0	19.	0	19	0	18	0	18	0	20	0	18 20	0	20	0	19	0	21	5.
2 24	ŏ	21.	ŏ	21	ő	21	ŏ	22	0	22	ŏ	22	0	. 22	0	22	ŏ	23	မှ 3€
2 36	0	23	0	23	0	23	0	23	0	24	0	24	0	24	0,	24	0	24	39
2 48	0	24 26	0	25 26	0	25 27	0	25 27	0	25 27	0.	26 27	0	26 28	0	26 28	0	26 28	42
3 12	0	28	0	28	U	29	0	29	0	29	Ð	29	0	30	0	30	0	30	48
3 24	0	30	0	.30	0	30	0		0	31	0	31	0	- 31	0	32	0	32	51
3 36 3 48	0	31 33	0	32 34	0	32 34	0	32 34	0	3 3 3 5	0.	33 35	0	33 35	0	34 35	0	34 36	54 57
4 0	ŏ	35	ŏ	35	ŏ.	36	o	36	ŏ	36	ŏ	37	o	37	o	37	ŏ	38	60
4 12	0	37	0	37	0	37	0	33	Ů.	3 8	0	38	·Q	39	0	39	0	40	63
4 24 4 36	0	3 8 4 0	0	39 41	0	39 41	0	40 41	0	40 42	0	40 42	O O	41	0	41 43	0	41	66
4 48	ŏ	42	ő	42	ŏ	43	ŏ	43	ŏ	44	ő	44	ŏ	41	ŏ	45	0	45	69 72
5 0	0	44	0	44	0	45	0	45	0	45	0	46	ø	46	0	47	0	47	75
5 12 5 24	0	45 47	0	46	0	46	0	47	0	47	0	48	0	48	0	49	0	49	78
5 36	Ö	49	0	48 49	0	48 50	0	49 50	0	49 51	0	49 51	0	50 52	0	50 52	0	51 53	81 84
5 4 8	0	51	0	51	0	52	Ó	52	0	53	0	53.	0	54	0	54	ō	55	87
6 0	0	52	0	53	0	53	0	54	0	54	0	55	0	55	, 0	56	0	56	90
6 12 6 24	0	54 56	0	55 57	0	55° 57	0	56 58	0	56 58	0	57 59	0	57 59	0	.58 0	0	58 0	93 96
6 36	ρ	58	0	58	0	59	0	59	1	0	1	ő	ĭ	1	ĵ.	2	i	2	99
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7 12	1	3	<u>-</u>	4	1		÷	-5	1	5	+	6	1	7	1	- <u>5</u>	1	<u>6</u>	105
7 24	1	5	1	5	1	6	1	7	li	7	i	8	i	8	i	ģ	ì	10	111
7 36 7 48	1	6 8	1	7 9	1	8 10	1	8	1	. 9	1	10	1	10.	1	11	1	12	114
8 0	i	10	î	11	1	11	i	10 12	1	11 13	1	11 13	1	12 14	1	13 15	1	13 15	117
8 12	1	12	1	12	1	13	1	14	1	14	1	1,5	ī	16	1	17	i	17	123
8 24	1	13 15	1	14	1	15	. 1	16	1	16	1	17	1	18	1	18	1	19	126
8 36 8 48	1	17	1	16 18	1 1	17 18	1	17 19	1	18 20	1	19 21	1 1	20 21	1	20 22	1	21 23	129 132
9 0	1	19	1	19	1	20	1	21	i	22	i	22	î	23	1	24	i.	25	135
9 12	1	20 22	1	21	1	22	1	23	1	24	1	24	1	25	1	26	1	27	138
9 24 9 3 6	1	22 24	1	23 25	1	24 26	1	25 26	1	25 27	1	26 28	1	27 29	.1	23 - 3 0	1	29 30	141
9 48	1	26	1	27	1	27	1	28	1	29	1	30	i	31	1	31	i	32	147
0 0	1	27	1	28	1	29	1	30	1	31	1	32	1	32	1	33	1	34	150
0 12	1	29 31	1	30 32	1	31 33	1 1	32 34	1	33 34	1	33 3 5	1	34	1	35	1	36	153
0 36	1	33	1	34	1	35	1	35	1	36	1	37	1	36 38	1	37 39	1	38 40	156 159
0 48	1	34 36	1	35	1	36	1	37	1	38	1	39	1	40	.1	41	1	42	162
1 12	$\frac{1}{1}$	38	+	37 39	1	38 40	1	39 41	1	40	!	41	1	42	1	43	1	44	165
1 24	i	40	i	41	1	42	1	43	1	42	1	43 44	1	44	1	45 46	1	45 47	168 171
1 36 1 48	1	41	1	42	1	43	1	44	1	45	1	46	1	47	1	48	1	49	174
2 0	i	43 45	1	44 46	1	45 47	1	46 48	1	47	1	48 50	1	49 51	1	50 52	1	51	177 LEO
	3/	30″	3/	32"		34//	3/	36"		38"	3/	40	÷	42"	3	44	N	53 464	2 13.2.3
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For reducing the Sun's Right Ascension in Time, as given in the Nantical Almanac for Noon at Greenwich, to any other time under any other Meridian.

1 000	101 140	Della	Vorie	ion of	he Sun	r ume	Ascens	ny otne		lan.	
from	7 "	/ //	/ //	1011 01	/ //	i / "	/ //	10 m m	Cime.	7 //	Shir
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0 36	0 6	0 6	0 6	0 6	lo 6	0 6	0 6	0 6	0 6	9 6	} ;
0 48	0 8	0 8	0 8	0 8	0 8	0 8	0 8	0 8	0 8	0 8	15
1 0	0 9	0 10	0 10	0 10	0 10	0 10	0 10	0 10	0. 10	0 10	1:
1 24	0 13	0 13	0 14	0 14	0 14	0 14	0 14	0 14	0 12	0 12	2
1 36	0 15	0 15	0 15	0 16	0 16	0 16	0 16	0 16	0 16	0 16	24
1 48	0 17	0 1.7 0 19	0 17	·0 18	0 18	0 18	0 18	0 18 0 20	0 18	0 18	2:
2 12	0 21	0 21	0.21	0 21	0 22	0 22	0. 22	0 22	0 22	0 23	3:
2 24	0 23	0 28	0 23	0 23	0 24	0 24	0 24	0 24	0 24	0 25	3€
2 36 2 48	0 25	0 25 0 27	0 25 0 27	0 25	0 26	0 26	0 26	0 26	0 26	0 27	39 49
3 0	0 28	0 29	0 29	0 29	0 29	0 30	0 30	0 30	0 30	0 29 9 31	4:
3 12	0 30	0 31	0 31	0 31	0 31	0 32	0 32	0 32	0 33	0 33	41
3 24	0 32	9 33	0 33	0 33	0 33	0 34	0 34	0 34	0 35	0 35	51
3 36 3 48	0 34	0 34 0 36	0 35	0 35	0 35 0 37	0 36 0 38	0 36 0 38	9 36 0 38	0 37	0 37 6 39	54 57
4 · 0	0 38	0 38	0 59	0 39	0 39	0 40	0 40	0 40	0 41	0 41	6(
4 12	0 40	0 40	0 41	0 41	0 41	0 42	0 42	0 42	0 43	0 43	65
4 24 4 36	0 42	0 42	0 43	0 43 0 45	0 43	0 44	0 44	0 44	0 45	0 45 9 47	66 65
4 48	0 46	0 46	0 46	0 47	0.47	0 48	0 48	0 48	0 49	0 49	79
5 0	0 47	0 48	0 48	0.49	0 49	0 50	0 50	0 50	0 51	Ò 51	7!
5 12	0 49	0 50	0 50	0 51	0 51	0 52	0 52	0 52	0 53	0 53	78
5 24 5 36	0 51 0 53	0 52	0 52	0 53 0 55	0 53 0 55	0 54	0 54	0 54 0 56	0 55 0 57	0 55 0 57	81 84
5 48	0 55	0 56	0 56	0 57	0 57	9 58	0 58	0 58	0 59	0 59	81
6 0	0 57	0 57	0 58	0 58	0 59	0 59	1 0	1 0	1 1	1 1	.9(
6 12	0 59	0 59	1 0.	1 0	1 1	1 1	1 2 1 4	1 3	1 3	1 4	9; 9(
6 36	î ŝ	1 3	i 4	1 4	1 5	1 5	1 6	1 7	1 7	1 8	99
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7 24	i 10	î ii l	1 i2	1 12	1 13	1.13	1 14	1 15	1 15	1 16	111
7 36	1 12	1 13	1 13	1 14	1 15	1 15	1 16	1 17	1 17	1 18	114
7 48 8 0	1 14 1 16	1 15	1 15	1 16	1 17 1 19	1 17 1 19	1 18 1 20	1 19 1 21	1 19 1 21	1 20 1 22	117
8 12	1 18	1 19	1 19	1 20	1 21	1 21	1 22	1 23	1 23	1 24	12:
8 21	1 20	1 20	1 21	1 22	1 23	1 23	1 24	1 25	1 25	1 26	126
8 36 8 48	1 22	1 23	1 23	1 24	1 · 25 1 · 27	1 25 1 27	1 26	1 27	1 27 1 29	1 28	129 139
9 0.	1 25	1 26	1 27	1 28	1 28	1 29	1 30	1 31	1 31	1 32	13:
9 12	1 27	1 28	1 29	1 39	1.30	1 31	1 32	1 33	1 34	1 34	131
9 24 9 36	1 29 1 31	1 30 1	1 31	1 32	1 32	1 33 1 35	1 34	1 35	1 36 1 38	1 36 1 33	14
9 48	1 33	1 34	1 35	1 36	1 36	1 37	1 38	1 39	1 40	1 40	14
10 0	1 35	1 36	1 37	1 57	1 38	1 39	1 40	1 41	1 42	1 42	150
	1 37	1 38	1 39	1 39	1 40	1 41	1 42	1 43	1 44	1 45	15:
10 24 10 36	1 39	1 40	1 41	I 41	1 42	1 45	1 44	1 45	1 46 1 48	1 49	15(15)
10 48	1 43	1 43	1 44	1 46	.1 46	1 47	1 48	1 49	1 50	1 51	16:
11 0	1 44	1 45	1 46	1 47	1 48	1 49	1 50	1 51	1 52	1 53	16
11 12 1	1 46	1 47	1 48 1 50	1 49 1 51	1 50 1 52	1 51 1 53	1 52	1 53 1 55	1 54 1 56	1 55 1 57	16
11 36	1 50	1 51	1 52	1 53	1 54	1 55	1 56	1 57	1 58	1 59	17
11 48	1 52	1 55	1 54	1 55	1 56	1 57		1 59	2 0	2 1 2 3	17' 18'
12 0	3/ 48"	2/ 50//	3/ 52/	1 57 3 54"	3/ 56"	1 59 3/58"	2 0 4/ 0"	4 2"	4 44	406#	-16
1	J 40"	J 301	J J2''	0 04"	J 30"	0.00	7 0	Diditized.	لكاتب ود	UYU	<u> </u>

TABLE XXXI.

For reducing the Sun's Right Ascension in Time, as given in the Nantical Almanac for Noon at Greenwich, to any other time under any other Meridian.

Noon		or Noo	Dail	v Vari	ation (of the S	Sun's 1	light	Ascens	ion in	Time.	
Noon.	Time	7-11										Shim's
0 12 0 2 0 2 0 2 0 2 0 2 0 2 0 2 0 2 0 2		4 8	4 10	4 12	4 14	4 16	4 18	4 20	4 22	4 24	4 26 4 28	Long.
0 12 0 2 0 2 0 2 0 2 0 2 0 2 0 2 0 2 0 2	<u> </u>		0' 0'	0 0	lo o	0' 0'	0' 0"	10'0	0 0	0' 0"	0' 0''0' 0	
0 356											0 2 0 2	
0				0 4			0 4	0 4				6
1 0 0 10 0 10 0 10 0 10 0 11 0 11 0 1	0 36								1			1 - 1
1 12 0 12 0 13 0 13 0 13 0 13 0 13 0 13					17 7							
1	-				.					احسنا		
1					1							
2 0 0 21 0 21 0 21 0 21 0 21 0 21 0 21												1 - 1
2 0 0 21 0 21 0 21 0 21 0 21 0 21 0 21				L								
2 12 0 23 0 23 0 23 0 23 0 23 0 23 0 25 0 24 0 24 0 24 0 24 0 24 0 25 25 25 25 0 25 0												
2 24 0 25 0 25 0 25 0 25 0 25 0 26 0 26 0 26					-					1		
2 36 0 27 0 27 0 27 0 28 0 28 0 28 0 28 0 28										L I		1
2 48 0 29 0 29 0 29 0 29 0 30 0 30 0 30 0 30												, ,
3				1 -	1							, 1
3 12 9 33 0 34 0 34 0 34 0 34 0 34 0 35 0 35 0 35 0 35 0 35 0 35 0 35 0 35 0 35 0 35 0 35 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 36 0 39 0 39 0 39 0 39 0 40 0 40 0 40 0 41 0 41 0 41 0 41 0 42 0 42 0 42 0 43 0 44 0 46 0 45 0 55 0 55 0 55 0 55 0 55 0 55 0 55 0 55 0 55 <td< td=""><td></td><td></td><td></td><td></td><td>0 32</td><td>0 32</td><td></td><td></td><td></td><td></td><td></td><td></td></td<>					0 32	0 32						
3 24 0 35 0 35 0 36 0 36 0 36 0 36 0 37 0 37 0 37 0 37	3 12	9 33	0 33	0 34	0 34	0 34	0 34	0 35	0 35	0 35	0 35 0 36	49
3 36 0 37 0 37 0 38 0 38 0 38 0 39 0 39 0 39 0 40 0 40 0 40 0 40 54 4 0 40 0 41 0 41				4								
4 0 0 41 0 42 0 42 0 42 0 43 0 43 0 43 0 44 0 44		0 37	0 37		0 38	0 38				0 40	0 40 0 40	54
4 12 0 43 0 44 0 44 0 44 0 45 0 45 0 46 0 46 0 47 0 47 0 48 0 48 0 48 0 49 0 49 66 4 36 0 48 0 48 0 48 0 49 0 49 0 49 0 49 0 49		, 1		,								4
4 24	4 0	0 41	0 42	0 42	0 42	0 43	0 43	0 43	0 44	0 44	0 44 0 45	60
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5 0 6 2 0 5 0												1 -7 1
5 12 0 54 0 54 0 54 0 55 0 55 0 55 0 56 0 56												
8 24 0 56 0 57 0 57 0 58 0 58 0 58 0 59 0 59 1 0 1 0 1 1 1 1 1 2 1 2 1 3 3 84 5 36 0 58 0 59 0 59 1 0 1 0 1 1 1 1 1 2 1 2 1 3 3 3 4 1 4 1 4 1 5 5 5 7 0 0 1 2 1 2 1 3 1 3 1 3 1 4 1 4 1 5 5 1 6 1 6 1 6 1 7 1 7 1 8 1 8 1 9 1 9 0 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1												
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6 36	5 270	1			42 2	1				1 10		
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7 12	1											
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9 36 1 39 1 40 1 41 1 42 1 42 1 43 1 44 1 45 1 46 1 46 1 47 1 44 9 48 1 41 1 42 1 43 1 44 1 45 1 46 1 47 1 48 1 49 1 49 1 47 10 0 1 43 1 44 1 45 1 46 1 47 1 48 1 49 1 50 1 50 1 50 1 51 1 52 1 50 1 10 12 1 45 1 46 1 47 1 48 1 49 1 50 1 51 1 52 1 50 1 50 1 51 1 52 1 50 1 50	9 12	1 35			1 37	1 38	1 39.	1 40	1 40	1 41	1 42 1 43	138
9 48 1 41 1 42 1 43 1 44 1 45 1 46 1 47 1 48 1 49 1 49 1 47 1 0 0	1 - 1											
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				1 61	1 50	1 53	1 54	1 00 1 66	1 56	57	57 T 50	
				1 53	1 54	1 55	1 56	1 57	I 58	1 59 1	0 2 1	
11 12 1 56 1 57 1 58 1 59 1 59 2 0 2 1 2 2 2 3 2 4 2 5 168 11 24 1 58 1 59 2 0 2 1 2 2 3 2 3 2 4 2 5 168 11 34 1 58 1 59 2 0 2 1 2 2 2 3 2 3 2 4 2 5 2 6 2 7 171 11 36 2 0 2 1 2 2 2 3 2 4 2 5 2 6 2 7 2 8 2 9 2 10 174 11 48 2 2 2 3 3 4 2 5 2 6 2 7 2 8 2 9 2 10 174				1 55	1 56	1 57	1 58	1 59	2 0	2 1 5	2 2 3	
11 24 1 58 1 59 2 0 2 1 2 2 2 3 2 3 2 4 2 5 2 6 2 7 171 11 36 2 0 2 1 2 2 2 3 2 4 2 5 2 6 2 7 2 8 2 9 2 10 174 11 48 2 2 2 3 2 4 2 5 2 6 2 7 2 8 2 9 2 10 2 11 2 12 177						1 59		2 1	2 2			
11 36 2 0 2 1 2 2 2 3 2 4 2 5 2 6 2 7 2 8 2 9 2 10 174 11 48 2 2 2 3 2 4 2 5 2 6 2 7 2 8 2 9 2 10 2 11 2 12 177	11 24 1	58 1	59	2 0		2 2	2 3	2 3	3 4	2 5 5	6 2 7	
1 11 48 12 2 12 3 12 4 12 5 12 6 12 7 12 8 12 9 12 16 12 11 12 12 1 177	11 36 2	0 2	i [2 2	2 3	2 4	5	2 6	2 7	2 8 2	9 2 10	174
	11 48 3		5	4	2 5	2 6	7	8 5	2 9	2 10 2	11 2 12	177
12 0 2 4 2 5 2 6 2 7 2 8 2 9 2 16 2 11 2 12 2 13 2 14 180								2 10	2 11			
4 8" 4 10" 4 12" 4 14 4 16" 4 18" 4 20" 4 22" 4 24" 4 26" 4 28" 0 0	1 4	8/4	10/	12/	¥ 14·	V 16"	V 18"	20″	4/ 22/4	V 241-1	26"14 28"	nogia

TABLE XXXII.

Variation of the Sun's Altitude in one minute from noon.

1	T		Decli	pation	of a d	ferent	Dame			itude.			
l i	00	10	20	30	40	50	60	70	80	90	100	110	
Lat.	"	P	"	"	"	"	"	"	"	"	H	"	Lat.
00	1	1	1	28.1	28.1	22.4 18.7	18.7 16.0	16.0 14.0	14.0	12.4	11.1	9.3	0°
1 2		ļ	28.1	22.4	18.7	16.0	14.0	12.5	11.2	10.2	9.3	8.6	2
3	28.1	28.1	22.4		16.0	14.0	12.5		10.2	9.3	8.6	8.0	3
5	22.4	18.7	18.7	16.0	14.0	12.5	10.2	9.3	9.3	8.6	7.4	7.4	5
6	18.7	16.0	14.0	12.5	11.2	10.2	9.3	8.6	8.0	7.5	7.0	6.6	6
7	16.0	14.0	12.4	11.2	10.2	9.3	8.6	8.0	7.5	7.0	6.6	6.2	7
8 9	14.0	12.4	11.2	9.3	9.3	8.6	8.0 7.5	7.5	7.0 6.6	6.6	6.2	5.9 5.6	8
10	11.1	10.1	9.8	8.6	8.0	7.4	7.0	6.6	6.2	5.9	5.6	5.3	10
11	10.1	9.3	8.6	8.0	7:4	7.0	6.6	6.3	5.9	5.6	5.3	5.1	11
12	8.5	8.5 7.9	7.9	7.4 6.9	7.0 6.5	6.5	6.2 5.8	5.9	5.6	5.0	5.Q 4.8	4.8	12
14	7.9	7.4	6.9	6.5	6.3	5.8	5.5	5.3	5.0	4.8	4.6	4.4	14
15	7.3	6.9	6.5	6.1	5.8	5.5	5.3	5.0	4.8	4.6	4.4	4.2	15
16	6.8	6.5	5.8	5.8	5.5	5.2	5.0	4.8	4.6	4.4	4.3	3.9	16
18	6.0	5.7	5.5	5.5	5.2	4.8	4.8	4.4	4.4	4.1	3.9	3.8	18
19	5.7	5.4	5.2	4.9	4.7	4.5	4.4	4.2	4.0	3.9	3.8	3.6	19
20	5.4	5.1 4.9	4.9	4.7	4.5	4.3	4.2	4.0	3.9	3.8	3.6	3.5	20 21
21	4.9	4.7	4.5	4.5	4.3	4.2	4.0 3.9	3.9	3.7	3.5	3.4	3.4	22
23	4.6	4.4	4.3	4.1	4.0	3.8	3.7	3.6.	3.5	3.4	3.3	3.2	23
24	4.4	4.2	4.1	3.9	3.8	3.7	3.6	3.5	3.4	3.3	3.2	3.1	24
25	1 22	4.1 3.9	3.9	3.8	3.7° 3.5	3.5	3.4	3.8	3.2 3.1	3.1	3.1	3.0	25
27	3.9	3.7	8.6	3.5	3.4	3.3	3.2	3.1	3.0	2.9	2.9	2.8	27
28	3.7	3.6 3.4	3.5	3.4	3.3	3.1	3.1	2.9	2.9 2.8	2.8 2.8	2.8	2.7	28 29
30	3.4	3.3	3.2	3.1	3.0	3.0	2.9	2.8	2.7	2.7	2.6	2.5	30
31	8.3	3.2	8.1	3.0	2.9	2.9	2.8	2.7	2.6	2.6	2.5	2.5	31
32	8.1 3.0	2.9	8.0 2.9	2.9	2.8	2.8	2.7	2.6	2.6	2.5	2.5	2.4	31 33
33	2.9	2.8	2.8	2.7	2.7	2.7	2.6 2.5	2.5	2.5	2.4	2.3	2.5	34
35	2.8	2.7	2.7	2.6	2.5	2.5	2.4	2.4	2.3	2.3	2.2	2.2	35
36	2.7	2.6	2.6	2.5	2.5	2.4	2.4	2.3	2.3	2.2	2.2	2.1	36 37
37	2.5	2.5	2.5	2.4	2.3	2.3	2.5	2.2	2.2	2.2	2.1	2.1	38
39	2.4	2.4	2.3	2.3	2.2	2.2	2.1	2.1	2.1	2.0	2.0	2.0	39
40	2.3	2.3	2.2	2.2	2.2	2.1	2.1	2.0	2.0	2.0	1.9	1.9	40
41	2.3	2.2	2.2	2.1	2.1	2.1 2.0	2.0	1.9	1.9	1.9 1.9	1.9	1.8	41 .
43	2.1	2.1	2.0	2.0	2.0	1.9	1.9	1.9	1.8	1.8	1.8	1.7	43
44	2.0	2.0	2.0	1.9	1.9	1.9	1.8	1.8	1.8	1.7	1.7	1.7	44
45 46	2.0	1.9	1.9	1.9	1.8	1.8 1.7	1.8	1.7	1.7	1.7	1.7	1.6	45 46
47	1.8	1.8	1.8	1.7	1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.6	47
48	1.8	1.7	1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.6	1.5	1.5	48 49
50	1.6	1.6	1.6	1.6	1.6	1.5	1.5	1.5	1.5	1.5	1.4	1.4	50
52	1.5	1.5	1.5	1.5	1.5	1.4	1.4	1.4	1.4	1.4	1.4	1.3	52
54	1.4	1.4	1.4	1.4	1.4	1.3	1.5	1.3	1.3	1.3	1.3.		54 56
56 58	1.3	1.3	1.3	1.2	1.3	1.3	1.2	1.2	1.2	1.2	1.2	1.2	58
60	1.1	1.1	1.1	1.1	1.1	1.1	1.1	1.1	1.0	1.0	1.0	1.0	60
62	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	0.9	62
64 66	1.0 0.9	0.9	0.9	0.9	0.9	0.9 0.8	0.9	0.9	0.9 0.8	0.9 0 .8	0.9	0.9	64 66
68	0.8	0.8	0.8	0.8	0.8	8.0	8.0	0.8	0.8	0,7	0.7	0.7	68
70	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	70
	0	1	2	3	4	5	. 6	7	8 Di	jit 9 ed b	10	ARI	

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			clinati										1 1	ı
123	130		150	160	170	180	190	200	210	220	230	240	1_ 1	ı
"	"	"	"	"	"	"	"	"	"	"	"	"	Lat.	l
9.2	8.5	7.9	7.3	6.8	6.4	6.0	5.7	5.4	5.1	4.9	4.6	4.4	0	ĺ
8.5	7.9		6.9	6.5	6.1	5.7	5.4	5.1	4.9	4.7	4.4	4.2	1	l
7.9	7.4		6.5	6.1	5.8	5.5	5.2	4.9	4.7	4.5	4.3	4.1	3	l
7.4	6.9	6.5	6.1	5.8	5.5	5.2	4.9	4.7	4.5	4.3	4.1	3.9 3.8	4	١
		_	_										5	ı
6.5	6.2 5.8	5.8	5.5	5.2	5.0 4.8	4.8	4.5	4.3	4.2	4.0 3.9	3.8	3.7	6	l
5.9	5.6	5.3	5.0	4.8	4.6	4.4	4.2	4.0	3.9	3.7	3.6	3.5	7	ı
5.6	5.3		4.8	4.6	4.4	4.2	4.0	3.9	3.7	3.6	3.5	3.4	8	l
5.3	5.0	4.8	4.6	4.4	4.2	4.1	3.9	3.8	3.6	3.5	3.4	3.3	9	l
5.0	4.8	4.6	-4.4	4.2	4.1	3.9	3.8	3.6	3.5	3.4	3.3	3.2	10	ı
4.8	4.6	4.4	4.2	4.1	5.9	3.8	3.6	3.5	3.4	3.3	3.2	5.1	11	ı
4.6	4.4	4.3	.4 1	3.9	3.8	3.7	8.5	3.4	3.3	3.2	3.1	3.0	12	ı
4.4	4.3	4.1	3.9	3.8	3.7	3.5	3.4	3.3	3.2	3.1	3.0	2.9	13	ı
4.2	4.1	3.9	3.8	3.7	3.5	3.4	3.3	3.2	3.1	3.0	2.9	2.8	14	ı
4.1	3.9	3.8	3.7	3.5	3.4	3.3	3.2	3,1	3.0	2.9	2.8	2.8	15	l
3.9	3.8	3.7	3.5	3.4	3.3	3.2	3.1	8.0	2.9	2.8	2.8	2.7	16	ı
3.8 3.7	3.7	3.5	3.4	3.3	3.2	3.1	3.0	2.9	2.8	2.8	2.7	2.6 2.5	18	l
3.5	3.4	3.3	3.2	3.1	3.0	2.9	2.9	2.8	2.7	2.6		2.5	19	ĺ
3.4	3.3	3.2	3.1	3.0	2.9	2.9	2.8	2.7	2.6	2.6	2.5	2.4	20	l
3.3	3.2	3.1	3.0	2.9	2.8		2.7	2.6	2.6	2.5	2.4	2.4	21	İ
3.2	3.1	3.0	2.9	2.8	2.8	2.7	2.6	2.6	2.5	2.4	2.4	2.3	22	l
3.1	3.0	2.9	2.8	2.8	2.7	2.6	2.6	2.5	2.4	2.4	2.3	2.3	23	ı
3.0	2.9	2.8	2.8	2.7	2.6	2.5	2.5	2.4	2.4	2.3	2.3	2.2	24	l
2.9	2.8	2.7	2.7	2.6	2.5	2.5	2.4	2.4	2.3	2.3	2.2	42.	25	l
2.8	2.7	2.7	2.6	2.5	2.5	2.4	2.4	2.3	2.3	2,2	2.1	7.1	26	ĺ
2.7	2.7	2.6	2.5	2.5	2.4	2.4	2.3	2.2	2.2	2.1	2.1	2.1	27	l
2.6	2.6	2.5	2.5	2.4	2.3	2.3	2.2	2.2	2.1	2.1	2.1	2.0	28	ı
2.6	2.5	2.4	2.4	2.3	2.3	2.2	2.2	2.1	2.1	2.0	2.0	2.0	29	l
2.5	2.4	2.4	2.3	2.3	2.2	2.2	2.1	2.1		2.0	2.0	1.9	30	ı
2.4	2.4	2.3	2.3	2.2	2.2	2.1	2.1	2.0	2.0	2.0	1.9	1.9	31 32	l
2.3 2.3	2.3	2.2	2.2	2.2	2.1	2.1	2.0	2.0	1.9	1.9	1.9	1.8	33	ı
2.2	2.2	2.1	2.1	2.0	2.a	2.0	1.9	1.9	1.9	1.8	1.8	1.8	34	į
2.2	2.1	2.1	2.0	2.0	2.0	1.9	1.9	1.8	1.8	1.8	1.7	1.7	35	l
2.1	2.1	2.0	2.0	1.9	1.9	1.9	1.8	1.8	1.8	1.7	1.7	1.7	36	Į
2.0	2.0	2.0	1.9	1.9	1.9	1.8	1.8	1.8	1.7	1.7	1.7	1 i.6	37	Į
2.0	1.9	1.9	1.9	1.8	1.8	1.8	1.8	1.7	1.7	1.7	1.6	1.6	38	۱
1.9	1.9	1.9	1.8	1.8	1.8	1.7	1.7	1.7	1.6	1.6	1.6	1.6	39	l
1.9	1.8	1.8	1.8	1.7	1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.5	40	I
1.8	1.8	1.8	1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.5	1.5	1.5	41	l
1.8	1.7	1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.5	1.5	1.5	1.5	42	l
1.7	1.7	1.7	1.6	1.6	1.6	1.6	1.5	1.5	1.5	1.5	1.4	1.6	43	I
1.7	1.6	1.6	1.6	1.6	1.5	1.5	1.8	1.5	1.5	1.4	1.4	1.4	44	ł
1.6	1.6	1.6	1.5	1.5	1.5	1.5	1.5	1.4	1.4	1.4	1.4	1.4	45	١
1.6 1.5	1.6	1.5	1.5	1.5	1.5	1.4	1.4	1.4	1.4	1.4	1.3	1.3	46	١
1.5	1.5	1.4		1.4	1.4	1.4	1.4		1.3	1.3	1.3	1.3	13	Ì
1.4	1.4	1.4	1.4	1.4	1.3	1.3	1.3	1.3	1.8	1.3	1.2	1.2		l
1.4	1.4	1.4	1.3	1.3	1.3	1.3	1.3	1.3	1.3	1.2	1.2	1.2	50	İ
1.3	1.3	1.5	1.3	1.3	1.3	1.2	1.2	1.2	1.2	1.2	1.1	1.1	51	ĺ
1.2	1.2	1.2	1.2	1.2	1.2	1.2	1.1	1.1	1.1	1.1	i.i	1.1	54	ĺ
1.2	1.1	1.1	1.1	1.1	1.1	1.1	1.1	1.1	1.1	1.0	1.0	1.0	56	ĺ
1.1	1.1	1.1	1.1	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	1.0	58	
1.0	1.0	1.0	1.0	1.0	1.0	1.0	0.9	0.9	0.9	0.9	0.9	0.9	60	l
0.9	0.9	0.9	0.9	0.9	0.9	0.9	0.9	0.9	0.9	0.9	Q.9	0.8	62	l
0.9	0.9	0.9	0.9	0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.8	64	l
0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.8	0.7	0.7	0.7		66	l
0.7	0.7 0.7	0.7 0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7			'	68 - 70	ĺ
12	13	14	0.6	0.6	0.6	18	0.6	90	21	99	93		۳	l
12 1	13	14.	ומו	16	17	ıx'	141	711	711	77	2.3	.74		4

							•						
					TAI	BLE	XXX	III.					23
		Var	istion	of the	Sun's	Altitu	de in (one mi	nute f	rom n	00B.		
					n of th	e sam	e name	70 .	e latiti	<u> 90</u>	100	110	
- .	00	10	20	30	11	"	//	"		111	"	"	Lat.
Lat.	ļ <i>"</i> —	-		-	28.1		18.7	16.0	14.0	12.4	11.1	10.1	02
1		ľ	1		۱. سر		22.4	18.6	16.0	13.9	12.4	11.1	1
2	ł	l	l	ł	1	1	28.0		18.6 22.3		13.9 15.8	12.3 13.8	3
3 4	28.1	1		١.		1		21.3	27.8	22.2	18.5	15.8	4
5	22.4	28.0	 	-			 			27.7	22.1	18.4	5
6	18.7	22.4	28.0		1	1	1		l	1	27.6	22.0	- ·6 7
7 8	16.0 14.0	18.6 16.0	22.3 18.6	27.9 22.3	27.8	1				1		****	8
9	12.4	13.9	15.9	18.5		27.7				İ			9
10	11.1	12.4	13.9	15.8	18.5	22.1	27.6				1		10
11	10.1	11.1	12.3	13.8	15.8	18.4	22.0 18.3	27.4 21.9	27.3	1	1)	11 12
12 13	9. 2 8.5	10.1 9.2	11.1	12.3	13.8 12.2	15.7 13.7			21.7	27.1	·		13
14	7.9	8.5	9.2	10.0	10.9	12.1	13.6	15.5	18.0		26.9		14
15	7.3	7.8	8.4	9.1	9.9	10.9	12.1	13.5	15.4	17.9 15.3	21 4 17.8	26.7 21.3	15 16
16 17	6.8	7.3 6.8	7.8	8.4 7.8	9.1	9.8	19.8	12.0 10.7	13.4	13.3		17.6	17
18	6.0	6.4	6.8	7.2	7.7	8.3	8.9	9.7	10.6	11.8		15.0	18
19	5.7	6.0	6.3	6.7	7.2	7.6	8.2	8.9	9.6	10.6	11.7	13.1	19
20	3.4.	5.7	5.6	6.3 5.9	6.7	7.1	7.6 7.0	8.1 7.5	8.8	9.5	10.5 9.5	11.6	21
21 22	5.1	5.1	5.3	5.6	5.9	6.2	6.6	7.0	7.5	8.0	8 6	9.4	22
23	4.6	4.8	5.0	5.3	5.5	5.8	6.1	6.5	6.9	6.8	7.9	8.5 7.8	23 24
24	4.4	4.6	4.8	5.0	5.2	5.5	5.4	5.7	6.4	6.4	6.8	7.2	25
25 26	4.2	4.4	4.6	4.7	5.0 4.7	4.9	5.1	5.4	5.7	6.0	6.3	6.7	26
27	3.9	4.0	4.1	4.3	4.5	4.7	4.9	5.1	5.3	5.6	5.9	6.2	27
28 29	3.7	3.8	3.8	4.1 3.9	4.3	4.4	4.6	4.8	5.0 4.7	5.3	5.5	5.8	28 29
30	3.4	3.5	3.6	3.7	3.9	4.0	4.2	4.3	4.5	4.7	4.9	5.1	30
31	3.3	3.4	3.5	3.6	3.7	3.8	4.0	4.1	4.3	4.4	4.6	4.8	31
32	3.1	3.2	3.3	3.4	3.5	3.7	3.8	3.9° 3.7	4.1 3.9	4.2	4.4	4.6	32 33
33 34	3.0 ⁷ 2.9	3.1 3.0	3.2	3.3	3.4	3.5	3.6	3.6	3.7	3.8	3.9	4.1	34
35	2.8	2.9	3.0	3.0	3.1	3.2	3.3	3.4	3.5	3,6	3.7	3.9	3.5
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37 38	2.6 2.5	2.7	2.7	2.8	2.9	2.9	2.9	3.1	3.2	3.2	3.2	3.3	38
39	2.4	2.5	2.5	2.6	2.7	2.7	2.8	2.9	2.9	3.0	3.1	3.2	39
40	2.3	2.4	2.4	2.5	2.6	2.6	2.7	2.7	2.8	2.9	3.0	3.0	40
41	2.3	2.3	2.4	2.4	2.5	2.5	2.6	2.6	2.7	2.8	2.8	2.9	41 42
42	2.2	2.1	2.2	2.2	2.3	2.3	2.4	2.4	2.5	2.5	2.6	2.7	43
44	2.0	2.1	2.1	2.1	2.2	2.2	2.3	2.3	2.4	2.4	2.5	2.5	44
45	2.0	2.0	2.0	2.1	2.1	2.2	2.2	2.2	2.3	2.3	2.4	2.4	45 46
46 47	1.9	1.9	2.0	2.0	2.0	2.1	2.1	2.2 2.1	2.2	2.1	2.2	2.2	47
48	1.8	1.8	1.8	1.9	1.9	1,9	2.0	2.0	2.0	2.1	2.1	2.1	48
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2.7 2 8 2.9 3.0 3.0 3.1 3.2 3.3 3.5 3.6 3.7 3.9 4.0 43 2.6 2.7 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 3.6 3.8 44 2.5 2.6 2.7 2.8 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 3.5 45 45 2.4 2.4 2.5 2.6 2.7 2.8 2.8 2.9 8.0 3.1 3.2 3.3 3.4 3.5 3.5 45 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 3.1 47 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.7 2.8 2.9 5.0 3.0 3.1 47 2.1 2.1 2.1 2.1 2.1 2.1 </td <td>3.0</td> <td>3.1</td> <td>3.2</td> <td>3.3</td> <td>3.4</td> <td>3.5</td> <td></td> <td></td> <td>3.9</td> <td></td> <td>4.2</td> <td></td> <td></td> <td>41</td>	3.0	3.1	3.2	3.3	3.4	3.5			3.9		4.2			41
2.6 2.7 2.7 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 3.6 3.8 44 2.5 2.6 2.6 2.7 2.8 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 3.6 3.8 44 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 4.5 4.6 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 3.1 47 2.2 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 5.0 48 2.1 2.1 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 49 2.0 2.0														
2.5 2.6 2.6 2.7 2.8 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 45 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.8 2.9 3.0 3.1 3.2 3.3 3.4 3.5 46 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 3.1 47 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 3.1 47 2.2 2.2 2.3 2.4 2.4 2.5 2.6 2.7 2.8 2.9 5.0 48 2.1 2.1 2.1 2.1 2.1 2.1 2.2 2.2 2.3 2.4 2.4 2.5 2.6 2.7 2.8 4.9 3.0 4.9 4.2 4.2 4					4 -								1	
2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.8 2.9 5.0 3.1 3.2 3.3 46 2.3 2.3 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 5.1 47 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.7 2.8 2.9 5.0 48 2.1 2.1 2.2 2.2 2.3 2.4 2.4 2.5 2.6 2.7 2.8 2.9 5.0 48 2.1 2.1 2.1 2.2 2.2 2.3 2.4 2.4 2.5 2.6 2.7 2.8 49 5.0 48 2.0 2.0 2.1 2.1 2.1 2.1 2.2 2.2 2.3 2.4 2.4 2.5 2.6 2.7 2.8 49 1.8 1.9 1.9 1.9 2.9 2.0 2.1 </td <td>-</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>٠</td> <td></td> <td></td> <td></td> <td></td>	-									٠				
2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 2.9 3.0 3.1 47 2.2 2.2 2.3 2.3 2.4 2.5 2.6 2.6 2.7 2.8 2.9 5.0 3.0 5.1 47 2.1 2.1 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 2.9 5.0 48 2.0 2.0 2.1 2.1 2.1 2.1 2.1 2.2 2.3 2.3 2.4 2.5 2.6 2.6 2.7 2.8 49 2.0 2.0 2.1 2.1 2.1 2.1 2.1 2.2 2.2 2.3 3.4 42 1.7 1.7 1.7 1.8 1.8 1.8 1.8 1.9 1.9 1.9 2.0 2.0 2.1 2.1 5.1 5.5 1.5 1.5 1.5 1.5 1														
2.1 2.1 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 2.7 2.8 49 2.0 2.0 2.1 2.1 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 50 1.8 1.9 1.9 2.0 2.0 2.1 2.1 2.1 2.2 2.2 2.3 2.4 52 2.2 2.3 2.4 52 2.2 2.3 2.4 52 2.5 2.6 50 50 2.8 50 2.0 2.1 2.1 2.1 2.1 2.2 2.2 2.3 2.4 52 2.7 2.8 49 2.0 2.1	2.3	2.3	2.4	2.4	2.5	2.6	2.6	2.7	2.8	2.9	2.9	3.0	3.1	57
2.0 2.0 2.1 2.1 2.2 2.2 2.3 2.3 2.4 2.4 2.5 2.6 2.6 50 1.8 1.9 1.9 2.0 2.0 2.1 2.1 2.1 2.2 2.2 2.3 2.4 52 1.7 1.7 1.8 1.8 1.8 1.9 1.9 1.9 2.0 2.0 2.1 2.1 54 1.5 1.6 1.6 1.6 1.7 1.7 1.7 1.8 1.8 1.8 1.9 1.9 2.0 2.0 2.1 2.1 54 1.5 1.6 1.6 1.6 1.7 1.7 1.7 1.8 1.8 1.8 1.9 1.9 2.6 1.4 1.4 1.4 1.4 1.4 1.4 1.6 1.6 1.7 1.7 58 1.3 1.3 1.3 1.4 1.4 1.4 1.4 1.4 1.5 1.6 1.5			1					1						
1.8 1.9 1.9 1.9 2.0 2.0 2.1 2.1 2.1 2.2 2.2 2.3 3.4 52 1.7 1.7 1.8 1.8 1.8 1.9 1.9 1.9 2.0 2.0 2.1 2.1 54 1.5 1.6 1.6 1.6 1.7 1.7 1.7 1.8 1.8 1.8 1.9 1.9 2.6 1.4 1.4 1.5 1.5 1.5 1.5 1.6 1.6 1.6 1.7 1.7 58 1.3 1.3 1.3 1.4 1.4 1.4 1.4 1.4 1.5 1.5 1.5 1.5 60 1.2 1.2 1.2 1.2 1.3 1.3 1.3 1.3 1.3 1.4 62 1.1 <td>-'</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td>	-'													
1.7 1.7 1.7 1.8 1.8 1.8 1.9 1.9 1.9 2.0 2.0 2.1 2.1 54 1.5 1.6 1.6 1.6 1.7 1.7 1.7 1.8 1.8 1.3 1.9 1.9 56 1.4 1.4 1.5 1.5 1.5 1.5 1.6 1.6 1.6 1.6 1.6 1.7 1.7 58 1.3 1.3 1.3 1.4 1.4 1.4 1.4 1.4 1.5 1.5 1.5 60 1.2 1.2 1.2 1.2 1.3 1.3 1.3 1.3 1.3 1.4 62 1.1<				1.9	2.0								2.4	
1.5 1.6 1.6 1.6 1.6 1.7 1.7 1.7 1.8 1.8 1.8 1.9 1.9 2.6 1.4 1.4 1.5 1.5 1.5 1.5 1.6 1.6 1.6 1.6 1.6 1.7 1.7 58 1.3 1.3 1.3 1.4 1.4 1.4 1.4 1.4 1.5 1.5 1.5 1.5 60 1.2 1.2 1.2 1.2 1.2 1.3 1.3 1.3 1.3 1.4 62 1.1 1.1 1.1 1.1 1.1 1.1 1.2 1.2 1.2 1.2 1.2 1.2 1.2 64 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.1				1.8	1.8							1	1.1	,
1.3 1.3 1.3 1.3 1.4 1.4 1.4 1.4 1.4 1.5 1.6 1.5 1.6 60 1.2 1.2 1.2 1.2 1.2 1.3 1.3 1.3 1.3 1.3 1.3 1.4 62 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.2	1.5	1.6		1.6	1.6	1.7	1.7	1.7	1.8	1.8		1.9	1.9	56
1.2 1.2 1.2 1.2 1.2 1.2 1.3 1.3 1.3 1.3 1.3 1.4 62 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.2 1.2 1.2 1.2 1.2 1.2 1.2 64 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.1 1.1 1.1 1.1 66 0.9 0				-						-}				
1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.2 1.2 1.2 1.2 1.2 1.2 1.2 64 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.1 1.1 1.1 66 0.9														
1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.0 1.1 1.1 1.1 1.1 66 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 1.0 1.0 68 0.														
0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 0.9 1.0 1.0 68 0.8														
							0.9	0.9	0.9	0.9	9.9	1.0	1.0	
1 120 130 140 150 160 170 180 190 200 210 220 230 240	d			1				-			_	1	. 1	70
	1 120	1 130	140	1 150	16°	1170	180	1 190	1 200	1 210	1 220	: 230	134°	ولهر

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To reduce the numbers of Table XXXII, to other given intervals of time from mon.

100	No.	- 1	-	174	67	Time	from 1	Nova.	3		4.0			
TS.	1 0'	1	147	1 8/	1 4	I M	6	71	1 '8'	9	10/	BIZ I	1190	8 7
0	0.0	1.0	4.0	9.0	16.0	25.0	36.0	49.0	64.0	01.0	100.0		141.0	-
	0.0	1.0	4.1	9.1	16.1	25.0 25.2	36.2	49:2	61.3·	91 3	100.3	121 (4	144.4	T
2 3	0.0	1-1	4.1	9.2	16.3	25.3	36.4	49.5	(FA 10	M1 6	300.7		141.4	
4	0.0	1.1	4.2	9.3	16.4	25.5	36.6	49.9	DA T	81.9		22.1	140 0	134
5	0.0	1.2	4.3	9.5	16.7	25.8	37.0	50.2	65.3	82.5	103.7	(00)	PHONE .	
6	0.0	1.2	4.4	9.6	16.0	26.0	37.2	50.4	65 6	88.8	102.0	128 :	(BG.4	
7	0.0	1.2	4.5	9.7	16.3	26.2	37.4	50.6	65.9	83.1	102 3	125.6	ILIO.	3
8	0.0	1.3	4-6	9.8	17-1	26.4	37.6	50.9	66.1	83.4	102.7	124-17	150	
10	0.0	1.4	4.6	10.0	17.4	26.5 26.7	37-8	51.4	66.4	84:0	405.4	No.	TAR P	10
fi	0.0	1.4	4.8	10.1	17.5	26.9	38.2	01.16	67.0	84.8	103.7	125.1	1	400
12	0.0	1.4	4.8	10.2	17.6		38.4		67.2	84.6	D-4 0	EU.	COX	
13	0.0	1.5	4.9	10.3	8.71	27.0 27.2	38.5	51.8 52.1	67.3	84.9	204.4	793 1	195 1	
14	0.1	1.5	5,11	10.5	17.9	27.4	38.9	52.3 52.6	67.8	85.3	104.7			10
15	0.1	1.6	5.1	10.6	18.1	27.6	39.1	52.6	68.I	85.6	105.1			18
17	0.1	1.6	5.2	10.8	18.3	27.9	39.5	53 0	68.6	86.2	105.7		150.9	15
18	0.1	1.7	5.3	10.9	18.5	28.1	39.7	53.3	68.8	86.5	106 L	_		
19	0.1	1.7	5.4	11:0	18.6	28.3	39.9	53.5	69.2	86.8	106 4			
20	0.1	1.8	5.4	11.1	18.8	28.4	40.1	53.8	69.4	87.1	106.8	128.4	152.1	20
21	0.1	1.8	5.6	11.2	18.9	28.6	40.5	54.0	69:7 70.0	37.4	107.1	199 0	152.9	田田田
23	0.1	1.9	5.7	11.4	19.2	29.0	40.7	54.5	70.3	88.0		129.6		8
24	0.2	2.0	5.8	11.6	19.4	29.2	41.0	54.8	70.6	88.4			153.8	24
25	0.2	2.0	5.8	11.7	19.5	29.3	41.2	55.0	70.8	88.7	103.5		154 4	3
26	0.2	2.1	5.9	11 8	19.7	29.5	41.4	55.3	71.1	39.0	108.5			26
27	0.2	2.1	6.0	12.0	19.8	29.7	41.8	55.5	71.4	89.6	109.2 109.6		155,0	27 28
29	0.2	2.2	6.2	12.1	20.1	30.1	42.0	56.0	72.0	89.9	1000	151 9	105.8	1
30	0.2	2.4	6.2	10.2	20.2	50.2	42.2	56.2	72.2	90.2			156 1	
31	0.3	2.3	6.5	12.4	20.4	30.4	42.5	56.5	72.5	90.6	110.5			31
32	0.3	2.4	6.4	12.5	20.6	30.6	42.7	56.8	72.8	90.9	111.0			
33	0.3	2.4	6.5	12.6	20.7	30.8	42.9	57.0	73.1	91.2	111.3			33
34	0.3	2.5	6.5	12.7	20.9	31.0	43.1	57.3 57.5	73.7	91.8	112.0	133.5		50
36	0.4	2.6	6.8	18.0	21.2	31.4	43.6	57.6	74.0	92.2			156.8	36
37	0.4	3.6	6.8	13.1	21.3	31.5	43.8	58.0	74.3	12.6		134.9		37
38	0.4	2.7	6.9	13.2	21.5	31.7	44.0	58 C	74. 5	92.8	113.1		159.6	38
39	0.4	2.8	7.0	13.3	21.6	31.9	44.4	58.5	75.1	98.1	115.4		160.0	39
41	0.4	2.8	7.2	13.6	21.9	35.3	44.7	59.0	75.4	93.3		136.0	liin.	100
42	0.5	2.9	7.3	15.7	28.1	32.5	44.9	59.3	75.7	94.1	F1 5 78			42
43	0.5	2.9	7.4	13.8	22.2	32,7	45.1	59 5	76.0	94,4	lis.s	137.3	161.7	43
44	0.5	3.0	7.5	15.9	22.4	32.9	45.3	59.8	76.3	DE-7	113.2	137.7	162.1	44
45	0.6	3.1	7.6	14.1	20.6	33.1	50-A	60.3	76.6	95.4	115.6		162 B	
47	0.6	3.1	7.7	14.2	22.9	33.4	46.0	60.6	77.1	11.7		138.8	163.3	46
43	0.6	3,2	7.8	14.4	23.0	38.6	46.2	60.8	77.4	96.0	116 0		163.8	THE PERSON NAMED IN
49	0.7	3.3	7.9	14.6	23.2	33.8	46.5	61 1	77-7	96.4	117.0	139 6	154.3	49
50	0.7	3.4	8.0	14.7	23.4	34.0	\$6.7	01.4	78.0	96 7	117.4	407.49	164.7	50
51	0.7	3.4	8.1	14.8	23.5	36.2	16.9	B1.9	78.3	97.4	117.7			51
52	0.8	3.5	8.2	15.1	23.8	34.6	17.2 57.4	62.1	78.6	97.7	118.4			
34	0.8	3.6	8.4	15.2	24.0	34.8	47.6	62.4	79.2	901 (1	-	41.6	166.	54
55	0.8	3.7	8.0	15.3	24.2	35.0	47.5	62.7	79.5	08	910 =	42 0	166.6	
56	0.9	3.7	8.6	15.5	24.3	35.2	48-1	62.9	79.8	90 T			167.3	56
57	0.9	3.8	8.7	15.6	24.5	35.4	18.3	65.2	80. I	29 0	120.3		162.7	57
58	1.0	3.9	8.8	15.7	24.7	35.6	48.8	63.7	80.4	89.3	120.6			59
-	0	10	97	3/	4/	NJ.	-	- 200	- 91	[MA.	100			
	- 0	-	- 20	1000	1	The Real Property lies				- Carrell I		-	The same of	

A INDUES AAA	IV, AAAY, AND AAAYI.											
TABLE XXXIV.	1			TAI	ILE I	XXX	V.					
Errors arising from a deviation of in the parallelium of the surfaces			i.	d	A	gle	of d	eria	tion			4
the central mirror.		107	15	120	20	130	135	40	40	100	IAB /	1.60
in d. Obs. to Obs. to Ohs. Fifth	D	11.	111	1 10	11	11	111	11	11	11	W	100
ngle, right. left. cross col.	0	0	0	0	0	0	0	0	0	10	-	0
D / " " " " " "	10	ő	0	1 1	l i	l i	2	2	3	10	3	5
0 0 0 0	20	0	i	l î	=	3	4	5	6	16	9	11
16 2 1 2 0 20 5 2 4 2	30	0	H	2	3	4	6	7	9	12	14	17
30 5 2 4 2 30 10 1 6 4	40	1		3	4	6	8		13	16	16	10
40 16 0 8 7	80	i	2	3	5	7	10	13	16	20	26	15
45 19 1 9 9	60	1	2	4	6	9		16	20	25	50	36
50 23 2 11 11	65	i	3	4		10	14	18	93	28	54	40
55 28 4 12 14	70	1	3	6		11	15	20		31	57	44
60 33 5 14 17	75		3	5	8	12	16	21	27	35		-
65 39 7 16 21	80	1	3	6		13	18	10		37	80	53
70 46 10 18 25	85	2	4	6		14		26	10	40	48	58
75 54 12 21 30	90	2	4	7	11	16	žI.	18	35	The same		63
80 1. 4 16 24 35	95	2	4	8		17	23	50	39	43	58	69
85 1.15 19 28 41	100	2	5	8	13	19	25	33	42	52	Ea l	85
90 1.27 23 32 48	105	2	5	9	14	20	28	36	145	57	53	1
95 1.43 28 37 56	110	2	6	10	16	22		40	50		E5	30
100 2:1 33 44 1.6	115	3	6.	11	17	25	34	44	55	68	55	55
105 2.23 39 52 1.16 110 2.49 45 1.1 1.29	120	3	7	12	19	27	37	48	61	76	31	109
110 2.49 45 1. 1 1.29 115 3.23 54 1.14 1.44		-										
20 4.05 1. 4 1.30 2. 3												
30 2.51												
40												
THE RESIDENCE OF STREET	Tanes \	FVV	VI		50.0	=						

TABLE XXXVI.												
Corrections of the mean Refraction for various heights of the Thermometer and Baromener. Th. 20° 24° 23° 32° 36° 40° 44° 48° 52° 56° 60° 64° 68° 72° 76°												
h Th. 200 210	230 3	20 360	400	440	480	520	560	60-	640	680	700	760
arum. 32.00 31 66	31 32 30	99 30.67	30.36	30.05	00.75	90.45	99 16	WILL LESS	28.60	10.11	20.00	_
	1+1/2											-
40 40			-		-							1
0 0 2 41 2 18	1 -			000		130	-	1	15 2			
0 30 2 181 28	1 3911	33 1 12 20 1 2		30 26	10	10	29	48	1 7	1 25	1 13	3.1
1 0 1 49 1 49		9 53	38	22	9	8 7	25	41	58	1 13	1	1 44
1 30 1 43 1 29	THE REAL PROPERTY.	0 46	33	19	6	6	19	36	50	1 3	1 3	1 30
2 0 1 30 1 18	1 50	53 40	29	17	6	6	16	27	43	55	. 3	1 18
2 30 1 20 1 8		46 36	25	1	_	-						
3 0 1 11 1		41 32	20	15	5	5	15	24	33	43	01	1 0
4 58 49	100	33 26	18	11	3	4	13	21	30	38	46	53
5 48 41		28 22	15	9	3	3	10	17	24	31	37	43
6 41 35		24 18	13	8	9	9	9	16	20	26	31	- 35
7 36 31	100	ACM DESCRIPTION OF	11		- 0	- 0		-		-	20	31
32 27		21 16 18 14	10	7	2	2	7	11	10	19	53	37
9 28 24		16 13	9	6	20	3	6	10	13	17	20	84
0 26 22		15 11	8	5	2	*	5	9	12	15	18	21
2 21 18		12 10	7	a a	3	1	2	6	30	14	16	120
4 18 16	1000	and producted t			_	-				-	-	16
6 16 14	11	8 11	6		- 5	- 3	3	6	8	10	12	14.
8 14 12	10	9 7	5	9	4	- 3	3	5	2	2	10	100
1 12 10	9	7 5	-	9		-	3	3	6	3	3	11
4 10 9	7	6 5	3	3	1	3	0	3	-0	-	0	
7 9 8	6	5 4	- 0	-	-	-	-	-	-	- 0		- 0
0 8 7		5 4	3	7	0	- 3	- 4	3	14	- 3	-5	1
7 6		4 3	9	1	0	10	1	3	0	4	9	
6 5	- 4	3 2	3	1	0	- 60	-	20	3	3	4	
5 5 4	3	3 2	î	1	0	0	-	7	3	3		
4 3	-	-	i	-	-	-	-	-	-	-	_	-
3 0	8	0 7	1	3	0	0	-	31	2	2	3	3
2 1	3	1 1	-	o o	0	0	2	1	-	1	3	1
7 1	in	0 0	0	- 6	0	00	0	4	-	-	21	1
0 0	0	0 0	0	6)	0	- 0	2	2	2	9	1	300

Longitudes and Latitudes of Stars, for Jan. 1, 1820.

Longitudes and Latit	ndes.	of Stars, for J	an. 1, 1520		
Names of Stars.	Mar.	Longitude.	Ann. Var.	Latitude.	Ann. var.
	200		aft. 1820,	0 / 1/	aft, 1820.
	2	0 6 00 00	50.09		-0.10
y PegasiAlgenib,	0	0. 6.38.48	19.98	12.35.42 N. 35.41. 7 N.	+ 0.12 + 0.16
Andromeda Stpheratz	4.3	0.11,48. 6	30,16	5.22. 2 N.	+0.25
ARIETIS	2.3	1. 5. 8.44	50.27	9.57.38 N.	+0.10
Ceti Menkar	2	1.11.48.18	50.27	12.35.44 S.	-0.57
Pleiadum Aleyone	3	1.27.28.35		4. 2. 3 N.	+0.43
Tauri	3	2, 3,16,52		5.45, 1 S.	-0.45
Tauri	2.4	2. 5.56.33	50.20	2.35. 6 S.	-0.46
Tauri ALDEBARAN	1	2. 7.16.23	50.21	5,28,44 S.	-0.33
3 Orionia Rigel		2,14,18.45	50.24	31. 8.44 8.	-0.47
Aurigio Capella	1	2.19.20.25		22.52.12 N.	+0.48
Orionis	1 3	2.19.50.55	50.20	23,34,34 S.	-0.48
Tauri	1 2	2.20. 3.36		5,22,26 N.	+0.48
Orionis	2	2.20.57. 1 2.22.10. 4	50,20	24,31,43 S. 25,18,5E S.	-0.08
Tauri	3	2.22.16.10		2.13. 0 S.	-0.48
Orionis Betelguese	1051	2.26.14.20		16. 3. 4 5.	-0.48
Ugminorum	3.4	3. 0.55.32	50.20	0.54.33 S.	-0.48
Gemingrum	3	3. 2.46.52	50.20	0.50, 4 8,	ALC: UNKNOWN
Geminorum	2.3	3, 6,35,13	50.18	6.45,41 5.	-0.47
Geminorum	3	3. 7.25.24	50 20	2. 2.55 N.	+0.46
Cante Majoris Sirius	1	3.11.36.34	50.07	39.22.31	-0.40
Geminorum	3.4	3,12,28,30		2. 3.36 S.	-0.45
Geminorum	3	3.16, 0.21	50.20	0.11.54 S.	-0.65
Geminorum Custor	1.2	3,17,44, 2	50.23	10. 5. 0 N.	+0.43
Geminorum Pollux		3.20.43.51	49.50	6.40.17 N.	+0.26
Canis Minoris Procyon	4.3	3,23.18.48	50.12 50.16	15.57.47 S. 5. 5.35 F.	-0.31
FiydraAlphard	2.0	4.11. 7 30 4.24.46.30	50.02	22.23.38 S.	-0.22
Leonis	3.4	4.25,23.18	50.23	4.51.19 N.	+0.32
Leonis Regulus		4.27.19.34	49.94	0,27,39 N	10.22
Leonis Denebola	1.2	5.19. 7.31	50.30	12 17 10 N.	+0.03
Virginis	3	5.24,35,54	50.20	0.41.32 N.	-0.02
Virginis	4.3	6, 2,19,20	50.21	1.22.23 N.	-0.08
Vivginis	3	6. 7.39.42		2.48.43 N.	-0.13
Virginis Spica		6,21,19,44	50,08	2. 2,20 S.	+0.17
Bootis Arcturus	1	6.21.43.30	50.45	80.54. 0 N.	-0.24
Coronæ BorAlphuoca	2,3	7, 9,45, 7	50.51	44.20.46 N.	-0.35 -0.37
Scrpentis	2.3	7.12.34.26	50.32	0.21.29 N. 25.31.31 N.	-0.40
			50.92	4.24.24 N.	-0.42
Seorpil	2.3	7.22.37. 5 7.28.36.38	50.18	5.27.45 S.	+0.44
Scorpii	3.2	8. 0. 3.22	200 -00	1.57.38 8.	-0.44
Scorpii	3	8, 0.25.27	50.18	5.26.59 8.	+0.45
Scornii	2	8. 0.40.27	50.20	1. 1.57 N.	-0.45
Scorpit ANTARES	- 1	8. 7.14.54	50.12	4.32.41 5,	+0.42
Onkinghi	3	8.18.52.48		1.49. 1 S.	+0.48
Ophiuchi Ras Alhague	2	8.19.55.23	50.21	35.52.26 N.	-0.48
Sagittarii	3	9, 9,52, 9	49.89	3.25.18 S.	+0.46 -0.45
Lyra Vega		9.12.47.19		61.44.26 N.	
Sagittarii	324	9.13.44.18	50.19	1.27,46 N. 31,15,43 N.	-0.45 -0.39
Aquile	1 3	9.28.25.48 9.29.14.10	50.79	29.18.45 N.	+0.08
Aquile	1 4	9.29.55.13	1000000	26.42.32 N	-0.35
2 Capricorni	3	10. 1.20.30		6.56.59 N.	-0.37
Capricorni	3	10, 1.31.52		4.36.32 N.	-0.37
Capricorni	4.3	10.19.16. 3	50.21	2.39.15 %	+0.96
Capricorni	3	10.21. 1. 7	50.21	2.33.49 8.	+0.25
Aquarii	3	11. 0.30.36	50.11	10,40,16 N.	-0,18 +0,31
Piece, Aust FOMALHAUT.		11. 1.19.32	50.59 49.42	31. 6:40 S. 59.54.57 N.	-0.16
Cygni Deneb	1.2	11, 2,51,16		19.24.44 N.	+0.10
Pogasi Manual	-	11.00.00.41	0011	The state of the s	

TABLES XXXVIII, XXXIX, XL. AND XLE

THE RESERVE AND ADDRESS OF THE PARTY OF THE	The state of the s													
TABLE XXXVIII.	TABLE XXXIX													
lust, or lat, and Hor. Par.			Aburr	ation					Commi	traile				
for Ellipticity 1		LIF.	- OCL	-1101	-				Tell I					
(Reduct, Red Dillor Par.						V	enu	5.		Mer	171			
	Elong.	Uran	Sat.	Jup.	Mars		-		-			-		
6) Lat. HorizontalPar.	1			1	ALC:	Ele	ng 1/	Ab.	Simo	ELAP	MALA	a Po		
53' KET 18656	-	-			1	200					-	-		
	D	-	-			100	D-	-	TOTAL ST		60 line	W.		
	Con. 6	259	27/	型9/4	36//	S.C	2.04	58	S.C.			N INCH		
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0.47.9 0.0 0.0 0.0		40	21	26	33		30 3			0 3				
1 35.0 0.1 0.1 0.1							45		10	71.00				
	4.5	19	41	-8-	23				_	_		_		
	60	Lin	16	10	23	CIL	Ell	200	- 4					
8, 9,2 0,2 0,2 0,2	7.5	10	KE.	110	13		40	10	100	5 4		1		
3.54.0 0.3 0.3 0.4	90	50	6	49	12		30	0	Gt.E	1 13	100	19		
4.39.3 0.5 0.5 0.5		1			100			10	0					
The state of the s	1 400		100	-	1			2	0					
5.22.4 0.6 0.7 0.7	100		1	3	100		10		1	7		100		
6. 3.9 0.8 0.9 0.0			1	+		Int	1	34	Carried Street	4 -1	10	III.		
0.43.7 1.0 1.1 1 1	120	5	-	1	13		1		. 1	0 3				
7.21.5 1.2 1.2 1.4	135	10	8	5	+		1		1	0 1	1	115		
	150	13	11	9	2	1	1			5 6	100	520		
7.57.2 1.5 1.6 1.7								-	2.00	-	111	1114		
8.30.7 1.8 1.9 2.0	165	15	13	IL	3	1	-1		Infa	2 20	100	111		
9, 1,6 2,0 2,3 2,3	ObTRO	15	13	II	4	1/20	-							
9.29.9 2.0 2.5 2.7	The	alloca	retirem.	100.0	No. S	TYPE .	in I	OWA	itork		NI-www.	y 30°		
	Line	Hoell	acton	OF	ne o	mit	HI.	OH	Hoon		100			
9.55.4 2.7 2.9 3:1	The op		u pin	ce is	given	1 In	ine	ES II	unites	100	TIME!			
10:18.1 3.0 3 2 5.4	by addi	mg 30	r' the	Sun	is ten	n lo	mgit	red	e, will	THE I	OCC IN			
10.37.8 3.5 3.6 3.3		40	BLE I	V E	-	T)			Was	ine 7	18.8			
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10.54.3 3.7 3.5 4.5	E	mar.	Equit	TOLES	in		Abe	tra	tion	III L.O	ME I			
11. 7.7 4.0 4.3 4.6		Lo	moilu	de		-	200		ration in Long and Let					
	Longitude.							1.0	ong=@long-#lun-					
11.17.8 4.4 4.7 5.0			-		-									
	1		-	Node								3410		
11.24.7 4.7 3.1 5.5		Long	-	Node										
11.21.7 4.7 3.1 5.5 11.28.2 0.1 5.5 5.9			-				Arg	lat						
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.5 5.9 6.3	D	Long	·) 'a	Node 2				lat	-A					
11.24.7 1.7 3.1 5.5 11.20.2 5.1 6.5 5.9 11.20.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7	D	Long	·) 'a	Node 2			Arg	lat	-0 +					
11.24.7 1.7 3.1 5.5 11.20.2 5.1 6.5 5.9 11.20.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7	D	Long	·) 'a	Node			Arg	lat	-A					
11.24.7 4.7 5.1 5.5 11.26.2 5.1 5.5 5.9 11.28.4 5.5 5.9 6.3 11.25.1 5.9 6.3 6.7 11.13.6 6.2 6.7 7.2	D	Long 10 +6	1 +7	Node 2 + 8	1000	The state of the s	Arg	lat	-0 +6	T T T T T T T T T T		3 de la		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.3 11.13.6 6.2 6.7 7.2 11.8.8 5.0 7.1 7.6	0.	10 + 6 0 0	1 + 7 8"9	2 + 8 10"8	30	Name of the last	Arg	lat	0 + 6	T + 7	+ 10-7	3 4 20		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.13.6 6.2 6.7 7.2 11.35.6 6.9 7.6 8.0	0.2	-0. +6 0.0 0.6	1 + 7 8//9 9.5	15.8	30		Arg	lat	- A	T 17"3	+ 10-0	3 de la 30 22 22 22 22 22 22 22 22 22 22 22 22 22		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.3 11.13.6 6.2 6.7 7.2 11.8.8 5.0 7.1 7.6	0.24	+ 6 0.6 0.6 1.2	1 + 7 8//9 9.5 10.0	15.8 16.1	30 28 26		Arg	lat of the same of	- 0 + 6 20 0 20 0 20 0	1. + 7 17/3 17.0 16.6	+ 1000	30 77 16		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.0 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 6.7 7.1 7.2 10.35.6 6.9 7.6 8.0 10.39.3 7.3 7.8 8.4	0. 2 4 6	0 + 6 0.6 0.6 1.2 1.9	1 + 7 8"9 9.5 10.0 10.5	15.8 16.1 16.4	30 28 26 24	THE REAL PROPERTY.	Arg	lat of the state o	- A O O O O O O O O O O O O O O O O O O	1. 17.0 17.0 16.6 16.5	+ B 10 10 10 10 10 10 10 10 10 10 10 10 10	30 77 16		
$\begin{array}{c} 11.24.7 & 4.7 & 5.3 & 5.5 \\ 11.26.2 & 9.1 & 5.5 & 5.9 \\ 11.28.4 & 5.5 & 5.9 & 6.3 \\ 11.25.1 & 5.9 & 6.3 & 6.7 \\ 11.13.6 & 6.2 & 6.7 & 7.2 \\ \hline 11.8.8 & 6.9 & 7.1 & 7.6 \\ 10.35.6 & 6.9 & 7.8 & 8.4 \\ 10.19.9 & 7.6 & 8.2 & 8.8 \\ \hline 10.19.9 & 7.6 & 8.2 & 8.8 \\ \end{array}$	0 24 45 00	0 + 6 0.6 0.6 1.2 1.9 2.5	1 + 7 8"9 9.5 10.0 10.5 11.0	15.8 16.1 16.4 16.6	30 28 26 24 29		Arg	lat of the state o	- 0 + 6 20 0 20 0 20 0	1. 17.0 17.0 16.6 16.5	+ 1000	30 77 16		
$\begin{array}{c} 11.24.7 & 4.7 & 5.3 & 5.5 \\ 11.26.2 & 9.1 & 5.5 & 5.9 \\ 11.28.4 & 5.0 & 5.9 & 6.3 \\ 11.25.1 & 5.0 & 6.3 & 6.7 \\ 11.18.6 & 6.2 & 6.7 & 7.2 \\ 11.8.8 & 6.9 & 7.1 & 7.6 \\ 10.55.6 & 6.9 & 7.8 & 8.0 \\ 10.39.5 & 7.8 & 7.8 & 8.4 \\ 10.19.9 & 7.6 & 8.2 & 8.8 \\ 2.77.4 & 7.9 & 8.5 & 9.1 \\ \end{array}$	0. 2 4 6	0 + 6 0.6 0.6 1.2 1.9 2.5	1 + 7 8"9 9.5 10.0 10.5 11.0	15.8 16.1 16.4 16.6	30 28 26 24 29		Arg	Dat S	- A O O O O O O O O O O O O O O O O O O	17/3 17/3 17.0 16.6 16.2 16.2	+ B 10 10 10 10 10 10 10 10 10 10 10 10 10	30 77 16		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.8 7.8 8.4 10.19.9 7.6 8.2 8.8 9.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.3	0 2 4 6 8 10	- 0 + 6 0.6 1.2 1.9 2.6 3.1	1 + 7 8"9 9.5 10.0 10.5 11.0	3 + 8 15.8 16.1 16.4 16.8	30 28 26 24 23 20		Arg	lat	0 + 6 20 0 20 0 19 9 19 8 19 7	1. 17/3 17/3 17.0 16.6 16.2 15.3	+ 10 0 9.4 8.5 8.7 7.8 6.8	30 20 20 20 20 20		
$\begin{array}{c} 11.24.7 & 4.7 & 5.3 & 5.5 \\ 11.26.2 & 9.1 & 5.5 & 5.9 \\ 11.28.4 & 5.0 & 5.9 & 6.3 \\ 11.25.1 & 5.0 & 6.3 & 6.7 \\ 11.18.6 & 6.2 & 6.7 & 7.2 \\ 11.8.8 & 6.9 & 7.1 & 7.6 \\ 10.55.6 & 6.9 & 7.8 & 8.0 \\ 10.39.5 & 7.8 & 7.8 & 8.4 \\ 10.19.9 & 7.6 & 8.2 & 8.8 \\ 2.77.4 & 7.9 & 8.5 & 9.1 \\ \end{array}$	0 2 4 6 8 10	+ 6 0 0 0 0 0 0 1.2 1.9 2.5 3.1	1 + 7 8"9 9.5 10.0 10.5 11.0 11.5 12.0	3 15.8 16.1 16.4 16.6 16.8 17.0	30 28 26 24 29 20		Arg	lat	0 + 6 20 0 20 0 19 9 19 8 19 7	1. + 7 17 17 16 .6 16 .2 16 .3 14 .9	+ 10-6 9.4 8.5 8.7 6.2 6.2	30 27 16 14 22 20		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.8 7.8 8.4 10.19.9 7.6 8.2 8.8 9.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.3	0 2 4 6 8 10	+ 6 0 0 0 0 0 0 1.2 1.9 2.5 3.1 3.7 4.3	1 + 7 8"9 9.5 10.0 10.5 11.5 12.0 12.4	15.8 16.1 16.8 17.0 17.2	30 28 26 24 29 20		Arg	lat	- A1 - 0 + 6 - 20 0 - 30 0	17/3 17/3 17/3 16.6 16.2 16.3 16.3	+ H 10-6 9.4 8.5 7.8 6.1 5.5	30 20 20 20 20 20 11		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.0 6.3 6.7 11.18.6 6.2 6.7 7.2 11.38 6.7 7.1 7.6 10.39.3 7.3 7.8 8.4 10.19.9 7.6 8.2 5.8 2.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 8.32.9 8.9 9.0 0.2	0 2 4 6 8 10 12 14 16	+ 6 0 0 0 0 0 0 1.2 1.9 2.5 3.1 3.7 4.3	1 + 7 8"9 9.5 10.0 10.5 11.5 12.0 12.4	3 15.8 16.1 16.4 16.6 16.8 17.0	30 28 26 24 29 20 18 16 14		Arg	lat	0 + 6 20 0 20 0 19 9 19 8 19 7	17/3 17/3 17/3 16.6 16.2 16.3 16.3	+ 10-6 9.4 8.5 8.7 6.2 6.2	30 27 16 14 22 20		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.6 8.2 8.8 2.77.4 7.9 8.5 9.1 9.32.0 8.8 8.9 9.5 9.32.0 8.8 8.9 9.5 9.32.9 8.8 9.0 9.0 0.2 7.59.6 9.1 9.6 0.5	0 2 4 6 8 10	+ 6 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	1 + 7 8/9 9.5 10.0 11.5 12.0 12.4 12.9	Node 	30 28 26 24 29 20 18 16		Arg	lat	- A1 - 0 + 6 - 20 0 - 30 0	17'3 17'3 17'3 17.0 16.6 16.2 16.3 16.3 16.3	+ H 10-6 9.4 8.5 7.8 6.1 5.5	30 20 20 20 20 20 11		
11.24.7 4.7 5.1 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.0 6.3 6.7 11.18.6 6.2 6.7 7.2 11.38 6.7 7.1 7.6 10.39.3 7.3 7.8 8.4 10.19.9 7.6 8.2 5.8 2.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 8.32.9 8.9 9.0 0.2	0 2 4 6 8 10 12 14 16 18	1.00g + 6 0.00 0.60 1.29 2.5 3.1 3.7 4.9 5.5	1 +7 8"9 9.5 10.0 10.5 11.0 11.3 12.0 12.4 12.9 13.3	3 15.8 16.1 16.8 17.0 17.2 17.4 17.6	30 28 26 24 23 20 18 16 14 12		Arg	1at 100 100 100 100 100 100 100 100 100 10		17.03 17.03 17.03 16.6 16.2 16.3 16.3 16.3 16.3 16.3 16.3	+ 10 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	30 25 25 25 20 21 10 14 11		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.6 8.2 8.8 2.77.4 7.9 8.5 9.1 9.32.0 8.8 8.9 9.5 9.32.0 8.8 8.9 9.5 9.32.9 8.8 9.0 9.0 0.2 7.59.6 9.1 9.6 0.5	0 2 4 6 8 10 12 14 16 18 20	1.2 1.9 2.5 3.7 4.9 5.5 6.1	1 + 7 8"9 9.5 10.0 11.5 12.0 12.4 12.9 13.3 13.7	Node 10"8 15.8 16.1 16.6 17.0 17.2 17.4 17.6	30 28 26 24 29 20 18 16 14 12 10		Arg	1at 100 100 100 100 100 100 100 100 100 10	0 + 6 20 0 20 0 19 9 19 8 19 7 19 8 19 7 19 8 19 7	1 7 17 17 17 18 18 18 18 18 18 18 18 18 18 18 18 18	+ 10 0 9.4 8.5 8.7 7.8 6.1 5.3 4.8	30 20 20 20 20 20 11 14		
$\begin{array}{c} 11.24.7 & 4.7 & 5.3 & 5.5 \\ 11.29.2 & 9.1 & 5.5 & 5.9 \\ 11.28.4 & 5.0 & 5.9 & 6.3 \\ 11.25.1 & 5.9 & 6.3 & 6.7 \\ 11.13.6 & 6.2 & 6.7 & 7.2 \\ 11.3.6 & 6.2 & 6.7 & 7.2 \\ 11.3.6 & 6.9 & 7.5 & 8.0 \\ 10.39.3 & 7.5 & 7.5 & 8.4 \\ 10.19.9 & 7.6 & 8.2 & 8.8 \\ 9.7.4 & 7.9 & 8.5 & 9.1 \\ 9.32.0 & 8.3 & 8.9 & 9.5 \\ 9.3.0 & 8.6 & 9.2 & 9.9 \\ 8.32.9 & 8.0 & 9.0 & 10.5 \\ 7.21.4 & 9.10.1 & 0.3 \\ 6.45.9 & 9.5 & 10.3 & 1.0 \\ \hline \end{array}$	0 2 4 6 8 10 12 14 16 18 20	1.00g + 6 0.00 0.60 1.29 2.5 3.1 3.7 4.9 5.5	1 + 7 8"9 9.5 10.0 11.5 12.0 12.4 12.9 13.3 13.7	3 15.8 16.1 16.8 17.0 17.2 17.4 17.6	30 28 26 24 23 20 18 16 14 12		Arg	lat		1 7 17 17 17 18 18 18 18 18 18 18 18 18 18 18 18 18	+ 10 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	30 25 25 25 20 21 10 14 11		
11.24.7 4.7 5.1 5.5 11.28.2 0.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11. 28.8 6.7 7.1 7.6 10.35.6 6.9 7.8 8.0 10.39.3 7.3 7.8 8.4 10.19.9 7.6 8.2 8.8 2.77.4 7.9 8.5 8.1 9.32 0 8.8 8.9 9.5 9.38 8.6 9.2 9.9 8.32.9 8.9 9.0 0.2 7.59.6 9.1 9.1 0.3 6.5.9 9.5 10.511.0 6.6.0 0.6 10.5 11.3	0 2 4 6 8 10 12 14 16 18 20	1.09 -0 -0 0.0 0.6 1.2 1.9 2.5 3.1 3.7 4.9 5.5 6.1 6.7	1 7 8"9 9.5 10.0 10.5 11.0 12.4 12.9 13.3 13.7	Node 10"8 15.8 16.1 16.6 17.0 17.2 17.4 17.6	30 28 26 24 29 20 18 16 14 12 10		Arg	lat	0 + 6 20 0 20 0 19 9 19 8 19 7 19 8 19 7 19 8 19 7	17 17 17 18 18 18 18 18 18 18 18 18 18 18 18 18	+ 10 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	30 25 25 25 20 21 10 14 11		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11. 8.8 6.7 7.1 7.6 10.35.6 6.9 7.8 8.0 10.39.3 7.3 7.8 8.4 10.19.9 7.6 8.2 8.8 2.77.4 7.9 8.5 9.1 9.32.0 6.3 8.9 9.9 8.32.9 6.3 8.9 9.9 8.32.9 6.3 8.9 9.9 8.32.9 6.3 8.9 9.3 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.6 10.5 7.28.6 9.1 9.7 10.5 7.28.6 9.1 9.6 10.5	0 2 4 6 8 10 12 14 16 18 20	1.00 -6 0.00 0.60 1.29 2.5 3.1 3.7 4.9 5.5 6.1 7.5	-1 +7 8"9 9.5 10.0 10.5 11.0 12.4 12.9 13.3 13.7 14.1	Node 	30 28 26 24 29 20 18 16 14 12 10		Arg	1at 15 15 15 15 15 15 15 15 15 15 15 15 15	0 + 6 6000 0 19.9 19.8 19.7 19.8 19.8 19.8 19.8 19.8 19.8 19.8	1 17 3 17 0 16 6 6 16 . 2 18 . 3 14 . 9 14 . 4 18 . 9 11 . 8 11 .	+ R 10-6 9.4 8.5 6.2 5.4 6.2 5.4 6.2 5.4 5.5 5.5	30 22 26 20 20 10 14 11 10		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.5 7.5 8.8 10.19.9 7.6 8.5 9.4 9.37.4 7.9 8.5 9.4 9.37.4 7.9 8.5 9.4 9.37.8 6.9 9.2 9.9 8.32.9 6.3 8.9 9.5 7.28.6 9.2 9.9 8.32.9 6.0 9.4 10.1 10.3 7.28.6 9.1 9.6 10.5 11.3 5.21.3 10.0 10.7 11.3 5.21.3 10.0 10.7 11.3 5.21.3 10.0 10.7 11.3	0 2 4 6 8 10 12 14 16 18 20 22 24 25	Long 	3/9 9.5 10.0 11.0 11.0 12.9 13.3 13.7 14.1	Node 3 10"8 15.8 16.1 16.6 16.6 17.0 17.2 17.4 17.5 17.6 17.7 17.8 17.9	30 28 26 24 29 20 18 16 14 12 10		Arg	lat	-A)	1 17 3 17 0 16 6 6 16 . 2 18 . 3 14 . 9 14 . 4 11 . 2 11 .	+ 10 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	30 22 26 20 20 10 14 11 10		
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11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.13.6 6.2 6.7 7.2 11.28.6 6.9 7.6 8.0 10.35.6 6.9 7.6 8.0 10.39.3 7.8 7.8 8.4 10.19.9 7.6 8.2 8.8 9.77.4 7.9 8.5 9.4 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 9.32.0 6.3 8.9 9.5 10.311.0 10.3 10.3 11.0 8.5 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3 9.32.0 6.0 0.6 10.5 11.3	0 2 4 6 8 10 12 14 16 18 20 22 24 25	Long 	3/9 9.5 10.0 10.5 11.0 12.4 12.4 13.3 13.7 14.1 14.5 14.8	Node 3 10"8 15.8 16.1 16.6 16.6 17.0 17.2 17.4 17.5 17.6 17.7 17.8 17.9	300 282 266 244 222 200 188 166 144 122 100 866 44 22		Arg	1at	-A)	1 17'3 17'3 17'3 17'3 17'3 17'3 17'3 17'	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.1 1.6 0.7	30 22 26 20 20 10 14 11 10		
11.24.7 4.7 5.3 5.5 11.28.4 5.0 5.9 6.3 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.13.6 6.2 6.7 7.2 11.3.6 6.2 6.7 7.2 11.3.6 6.9 7.6 8.0 10.39.3 7.8 7.8 8.4 10.19.9 7.6 8.2 8.8 9.37.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.5 9.3.8 8.6 9.2 9.9 8.32.9 8.3 8.9 9.0 0.2 7.58.6 9.1 9.10.5 11.3.6 6.0 0.6 10.5 11.3 5.24.3 10.0 10.7 11.6 4.11.0 10.1 10.1 10.3 3.19.4 10.4 11.2 12.0	0. 24 6. 8 10 12 14 16 18 20 27 24 26 28	Long 	3/9 9.5 10.0 10.5 11.0 12.4 12.4 13.3 13.7 14.1 14.5 14.8	Node 3 + 8 10"8 15.8 16.1 16.4 16.6 17.0 17.2 17.2 17.6 17.7 17.8 17.9 17.9	300 282 266 244 222 200 188 166 144 122 100 866 44 22		Arg	1at	-A 0 + 6 6 6 6 6 0 0 0 0 0 0 0 0 0 0 0 0 0	1 17'3 17'3 17'3 17'3 17'3 17'3 17'3 17'	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.1 1.6 0.7	30 22 26 20 10 14 11 10 G 4 11		
11.24.7 4.7 5.3 5.5 11.28.2 9.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.6 6.2 6.7 7.2 11.8.6 6.2 6.7 7.8 8.4 10.19.9 7.6 8.2 8.8 2.07.4 7.9 8.5 8.4 9.32 0 8.3 8.9 9.5 9.38 8.6 9.2 9.9 8.32.9 8.8 9.6 0.2 7.59.6 9.1 9.6 0.2 7.59.6 9.1 9.6 0.2 7.59.6 9.1 9.1 0.3 6.5.9 9.5 0.5 11.0 5.21.3 10.0 10.7 11.3 5.21.3 10.0 10.7 11.3 3.19.4 10.4 11.2 12.0 2.23.7 10.5 11.3 12.5	0. 24 6. 8 10 12 14 16 18 20 27 24 26 28		1 +7 8"9 9.5 10.0 10.5 11.0 11.3 12.0 11.3 13.3 14.1 14.1 14.8 15.2 15.8	Node 2 1-8 10"6 15.8 16.1 16.4 16.6 17.0 17.2 17.4 17.6 17.7 17.9 17.9 17.9	300 282 264 242 202 200 188 166 144 122 100 866 44 22 0		Arg	1at	-A 0 + 6 6 6 6 6 0 0 0 0 0 0 0 0 0 0 0 0 0	1 17'3 17'3 17'3 17'3 17'3 17'3 17'3 17'	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.1 1.6 0.7	30 25 26 20 20 20 10 10 10 10 10 10 10 10 10 10 10 10 10		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.6 8.2 8.8 10.19.9 7.6 8.2 8.8 2.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.3 9.32.0 8.3 8.9 9.3 9.32.0 8.3 8.9 9.3 10.32.0 8.6 9.2 9.9 8.52.9 8.0 9.0 0.2 7.59.6 9.1 9.6 10.5 7.23.4 9.4 10.4 10.3 6.45.9 9.5 10.5 11.0 6.6.0 0.8 10.5 11.3 5.24.3 10.0 10.7 11.5 5.24.3 10.0 10.7 11.8 3.10.4 10.4 11.2 12.0 2.23.7 10.5 11.3 12.1 1.36.2 10.5 11.3 12.1	0. 24 6. 8 10 12 14 16 18 20 27 24 26 28	Long 	1 + 7 8"9 9.5 10.0 10.5 11.0 11.5 12.9 13.3 13.7 14.5 14.5 15.2 15.8	Node 2 + 8 15"8 16.1 16.8 16.1 16.6 17.0 17.2 17.4 17.8 17.7 17.8 17.9 17.9	300 282 266 244 222 200 188 166 144 122 100 866 44 22		Arg	1at	-A 0 + 6 6 6 6 6 0 0 0 0 0 0 0 0 0 0 0 0 0	1 17'3 17'3 17'3 17'3 17'3 17'3 17'3 17'	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.1 1.6 0.7	30 22 26 20 10 14 11 10 G 4 11		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.13.6 6.2 6.7 7.2 11.3.6 6.2 6.7 7.2 11.3.6 6.9 7.6 3.0 10.39.3 7.8 7.8 8.4 10.19.9 7.6 8.2 8.8 9.37.4 7.9 8.5 9.4 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 9.32.0 8.3 8.9 9.5 7.23.6 9.1 9.6 10.5 7.23.6 9.1 9.6 10.5 7.23.6 9.1 9.6 10.5 7.23.6 9.1 9.6 10.5 7.23.6 9.1 9.6 10.5 11.3 5.24.3 10.0 10.7 11.3 5.24.3 10.0 10.7 11.3 3.19.4 10.4 11.2 12.0 2.23.7 10.8 11.3 12.1 1.36.2 10.5 11.3 12.1 1.36.2 10.5 11.3 12.1	0. 24 6. 8 10 12 14 16 18 20 27 24 26 28	Long - 0 + 6 0 0 0 0 6 1 2 2 5 3 1 1 3 .7 7 4 .3 4 .9 5 6 .1 6 .7 7 .5 7 .8 8 .4 8 .9 - 5 + 5	1 + 7 8"9 9.5 10.0 10.5 11.0 11.5 12.0 11.3 12.9 14.5 14.5 14.5 15.2 15.8	Node 2 + 8 15"8 16.1 16.8 16.1 16.6 17.0 17.2 17.4 17.8 17.7 17.8 17.9 17.9	300 282 264 242 202 200 188 166 144 122 100 866 44 22 0		Arg	1at	-A 0	1 1703 1770 1770 16.65 16.52 16.31 1	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.1 1.6 0.7	30 25 26 20 20 20 10 10 10 10 10 10 10 10 10 10 10 10 10		
11.24.7 4.7 5.3 5.5 11.28.2 5.1 5.5 5.9 11.28.4 5.0 5.9 6.3 11.25.1 5.9 6.3 6.7 11.18.6 6.2 6.7 7.2 11.8.8 5.6 7.1 7.6 10.35.6 6.9 7.6 8.0 10.39.3 7.6 8.2 8.8 10.19.9 7.6 8.2 8.8 2.77.4 7.9 8.5 9.1 9.32.0 8.3 8.9 9.3 9.32.0 8.3 8.9 9.3 9.32.0 8.3 8.9 9.3 10.32.0 8.6 9.2 9.9 8.52.9 8.0 9.0 0.2 7.59.6 9.1 9.6 10.5 7.23.4 9.4 10.4 10.3 6.45.9 9.5 10.5 11.0 6.6.0 0.8 10.5 11.3 5.24.3 10.0 10.7 11.5 5.24.3 10.0 10.7 11.8 3.10.4 10.4 11.2 12.0 2.23.7 10.5 11.3 12.1 1.36.2 10.5 11.3 12.1	0. 24 6. 8 10 12 14 16 18 20 27 24 26 28		1 +7 8"9 9.5 10.0 10.5 11.0 11.3 12.0 11.3 13.3 14.1 14.1 14.8 15.2 15.8	Node 2 1-8 10"6 15.8 16.1 16.4 16.6 17.0 17.2 17.4 17.6 17.7 17.9 17.9 17.9	300 282 264 242 202 200 188 166 144 122 100 866 44 22 0		Arg	1at	-A 0 + 6 6 6 6 6 0 0 0 0 0 0 0 0 0 0 0 0 0	1 17'3 17'3 17'3 17'3 17'3 17'3 17'3 17'	10°00 9.4 8.5 7.6 6.2 4.8 4.8 2.0 2.0 1.6 1.6 1.6 1.6 1.6 1.6 1.6 1.6 1.6 1.6	30 25 26 20 20 20 10 10 10 10 10 10 10 10 10 10 10 10 10		

able AL, routains the equation of the equinoses in longitude to be applied with its an interologistaties of all line betweenly bodies. Thus on July 16, 1820, when the longitude of the
is ascending node way 11s. 26.7 Of the equation of the equinoses was + 14.2,
the correction in Table XLL corresponding to the Argument of Longitude being found, with
rithin added to the log, secam (less radius) of the star's latitude will be the log, of the
rithin is founding that to be applied with its sign to the mean longitude. The logistical of the
continuity is founding to be applied with its sign to the description of the table of the logistic continuity of the logistic co

long 5, 2, 47 Tab, 41 + 10°, 3 log, L03342 Arg, lat. 1s, 2°, 47 Tab, 41 - 15°, 0 °, 120°, Latitude 19° 25° Sec. 0.02543 verr. long. + 11" .5 Log 1,05855 * Aber Latt - 5", 6

TABLE XLIL

Aberration	in Right Ascension and	Declination.
PART L	PART IL.	PART III.
Arg. R.A = * R.A O Long.	Ar.R.A.= * R.A.+ OLon.	Ar2dDec=Olon-j-kDec Adage
Arg. Dec = Arg.R. A.+3 signs.	Ar.Dec .= Arg. R.A+3signs.	Ar3dDec=Olon-*Dec \ itbeds
D 0. 6. 1. 7.12. 8.	D 0. 6. 1. 7. 2. 8.	D 0. 6. 1. 7. 2. 8.
1 - +-+	+-+-+-	0 -+-+-+
0 19", 17 16", 60 9", 59 30	0 07, 83 07, 72 07, 41 30	0 3/1.98 3/1.45 1/1.99 30
1 19 17 16 .43 9 .30 29	10 .83 0 .71 0 .40 29	1 3 98 3 41 1 93 20
2 19 16 16 ,26 9 .00 28	20 .830 .700 .39 28	2 3 98 3 38 1 87 28
319 -1516 -088 -70 27	30 ,830 ,690 ,38 97	3 3 .98 3 .34 1 .81 27
4 19 .13 15 .90 8 .40 26	40 -820 -690 -36 26	4 3 .97 3 .30 1 .75 26
5 19 .10 15 .71 8 .10 25	50 .820 .680 .35 25	5 3 .97 3 .26 1 .68 26
6 19 .07 10 .51 7 .80 24	60 .820 .670 .34 24	6 3 .96 3 .22 1 .62 24
7 19 .03 15 .31 7 .49 23	70 .820 .660 .32 23	7 3 95 3 18 1 56 23
8 18 .99 15 .11 7 .18 22	22 18, 0 65, 0 28, 0 3	8 3 .94 3 .74 1 .49 22
9 18 ,91 14 .90 5 .87 21	90 .820 .640 .30 21	9 3 .93 3 .09 1 .45 21
10 18 .88 14 .69 6 .56 20	100 .310 .630 .28 20	10 3 .92 3 05 1 .36 20
11118 82 14 .47 6 .24 19	110 .810 .620 .27 19	11 3 .91 3 .00 1 .30 19
12.18 .70 14 .25 5 .92 18	120 .810 .610 .26 18	12 3 .89 2 .96 1 .23 18
13.18 .68 14 .025 .61 17	130 .810 .600 .24 17	13 3 .88 2 .91 1 .16 17
14 18 .60 13 .79 5 .28 16	140 .800 .590 .23 16	14 3 .36 2 .86 1 .10 16
15 18 .02 13 .564 .96 15	15 0 .80 0 .58 0 .21 15	15 3 .85 2 .82 1 .03 16
1618 43 15 324 64 14	16 0 .79 0 .57 0 .20 14	16 3 .83 2 .77 0 .96 14
17 18 . 34 13 08 4 .31 13	170 .790 .560 .19 13	17 3 .81 2 .72 0 .90 13
18 10 .03 12 .53 3 .99 12	130 .790 .550 .17 12	18 3 .79 2 .66 0 .83 12
19 14 - 10 12 - 50 3 - 66 11	19 0 .78 0 .54 0 .16 11	19 3 .76 2 .61 0 .76 11
20 18 .02 12 .32 3 .33 10	20 0 .78 0 65 0 .14 10	20 3 .74 2 .56 0 .69 10
21 17 .90 12 .073 .00 9	210 .770 .600 .13 9	21 3 72 2 31 0 62 9
## 17 .78 11 .RO.2 .67 8	220 .770 .510 .12 8	22 3 69 2 40 0 55 8
23 17 .65 11 .64 2 .34 7	23 0 .76 0 .800 .10 7	23 3 .66 2 .40 0 .49 7
24 17 .52 11 .27 2 .00 6	24 0 .76 0 .49 0 .09 6	24 3 .64 2 .34 0 .42
25 17 .38 11 .00 1 .67 5	25 0 .75 0 .47 0 107 5	25 3 .61 = .28 0 .35 5
26 17 .23 10 .32 1 .34 4	260 .740 .460 .06 4	26 3 .58 2 .23 9 .28 4
27 17 .08 10 .441 .00 3	270 .740 .450 .04 3	27 3 35 2 17 0 21 3
28 16 .93 10 .16 0 .67 2	28 0 .73 0 .44 0 .03 2	28 3 .52 2 .11 0 14 2
29 16 .77 9 .87 0 .33 1	290 -720 -430 01 1	29 3 .48 2 .05 0 .07 1
30 16 .60 9 .59 0 .00 0	30 0 -72 0 -41 0 :00 0	30 3 .45 1 .99 0 .00 0
= += += + p	+-+-+-D	-+-+-+
11. 5. 10. 4. 9. 3. 1	11. 5. 10. 4. 9. 3.	11. 5 10. 4 9. 3.
	THE RESERVE OF THE PARTY OF	

To find the Aberration of a Star in Right Ascension.—Find the Equations in Part I: and I corresponding to the arguments of R. A. at the top of those tables, and connect them a cording to their signs, and to the log. of this sum or difference add the log. secant (less r lins) of the star's declination, the sum will be the log. of the aberration in Right Ascension seconds of a degree, which divided by 15 will be reduced to time, to be applied to the

To find the Aberration of a Star in Declination .- Increase the former arguments of R. y 3 signs, and connect together the corresponding equations of Part I, and II, to the log-faich add the log, sine of the ** seclination, the sum will be the log, of arch 1st. Wi-be arguments at the top of Part III, find in that Table arches 2d and 3d. These three arch connected with their signs will be the abstration in declination, to be applied to the meaelimtion.

EXAMPLE. Required the Aberration in R. A. and Dec. of a Pegasi, July 16, 1320 ?

table 8. * R.A = 12h. 55′ 48″ = 11s. 13°,57°, *Dec. 14° 14′ N. and by N.A. ⊙long 5a. 23° R.A. 11s. 13°,57′, bin. 3, 23° 46′ 7, 20, 11 Part I. +12,#27 3, 7, 43 Part II. - 0, 11 Diff. + 3s= 10s 200 H' Sum + 3 = 6, 7 43 * Decl. 14° 14° +12. 16 log. 1.08493 Arch 1st - 31.82 Aher R. A. +12". 5 log. 1.09847

Olong + * Dec 4s. 8°; 0' Arch 2d + 2, 45 } } If Decl. is S. auhl 6 Olong - * Dec 3, 9, 32 Arch 3d + 0, 66 } ; is Argum, Pari Ab inR. A. in time 0 .83

Nutation in Right Ascension and Declination to be applied to the mean values.

	The state of the s	
PART L Y	PART IL.	PART III. 1
IR. HA = * RA Lon. D node	Arg. R. A = *R.A.+lon. D node	Marie Company
4 6 signs if Dec. is S.	+6 signs if Doc. is S.	Equation Equators in R.A.
rg. Dre. = Arg. R. A. + 3 signs.		Arg. = Long. D andell
	ID 6 IF 7 10 0 1	0 6 / 7 7 2 2
D 0. 0.1. 7.2. 8.	D C C C C C C C C C C C C C C C C C C C	Dog at the same
_ + - + - +		Sealer - I
0 07 33 27 21 47 16 30	0 1", 22 1", 06 0", 61 30	0 0 0 5 4 14" 22
1 3 .33 7 .14 4 .04 29	1 1 221 050 59 29	1 0 3 0 414 329
2 8 .32 7 .06 3 .91 28	2 1 .22 1 .03 0 .57 28	2 0 6 8 714 520
3 8 .32 6 .99 3 .72 27	3 1 .22 1 .02 0 .55 27	3 0 .9 8 .5 14 .5 27
1 8 .31 6 .91 3 .65 26	4 1 .221 .010 .53 26	4 1 1 9 214 786
The state of the s		THE RESERVE AND THE PARTY OF TH
The Party of the P	5 1 .22 1 .00 0 .52 25	5 1 . 0 9 14 . 325
6 0 .28 6 .74 3 .39 24	6 1 .210 .990 .30 24	6 1 -7 9 -6 15 -6 24
7 8 .27 6 .65 3 .25 23	7 1 .210 .970 -48 23	7 2 .0 9 .975 125
8 8 .25 6 56 3 .19 00	3 1 .210 .960 .46 22	0 2 3 10 1 15 4 15
9 8 .23 6 .47 2 .99 21	9 1 .200 .950 .44 21	9 2 6 10 3 15 3 1
10 8 .20 6 .38 2 .85 20	10 1 .200 .930 .42 20	10 2 8 10 4 10 4 20
	THE RESERVE AND DESCRIPTION OF THE PERSON NAMED IN	The State of the S
11 11 . 18 6 . 29 2 . 71 19	11 1 .200 .920 .40 19	11 3 -1 10 700 -6 19
12 8 15 6 19 2 57 18	12 1 -190 .910 .38 10	12 3 -4 II -0 to 6 18
13 8 10 6 19 2 44 17	13 1 -190 .890 -36 17	13 3 7 11 2 15 7 17
14 8 .08 5 .99 2 .30 16	14 11 -180 -880 -34 16	14 4 .011 .410 716
15 3 .05 5 89 2 .16 15	15 1 180 860 32 15	16 4 211 Gla MA
	DOMESTIC OF THE PARTY NAMED IN COLUMN	
20 4 20 100 20 100 20 100 20	TO IN THE PROPERTY OF THE PARTY	16 4 311 -014 -414
COLUMN TO THE PARTY OF THE PART	17 1 -17 0 .83 0 .27 13	17 4 -812 .016 1133
10 7 .92 5 .57 1 .73 12	18 1 .16 0 .82 0 .25 12	18 5 .1 12 .2 16 1112
19 7 88 5 46 1 59 11	19 1 .150 .800 .23 11	19 5 .3 12 .6 16 111
20 7 .83 5 .35 1 .45 10	20 1 .150 .780 .21 10	20 5 5 12 0 15 1 10
21 7 .78 5 .24 1 .30 9	21 1 .140 .770 .19 9	21 0 912 216 2 9
25 7 72 5 13 1 16 8	22 1 130 750 17 8	22 6 712 576 2 3
The state of the s		Man the Man Man and Alexander
COLUMN TO THE REAL PROPERTY OF THE PARTY OF	24 1 -110 -720 -13 6	24 6 .713 -316 3
25 7 .55 4 .78 0 .73 5	25 1 110 .700 .11 5	28 6 913 - 16 3 4
26 7 .49 4 66 0 58 4	26 1 .100 .680 .09 4	16 7 2 10 .616 .5 1
27 7 .42 4 .54 0 .41 3	27 1 .090 .660 .06 3	27 7 413 716 4 3
28 7 .35 4 .41 0 .29 2	28 1 -080 650 04 2	28 7 713 916 4 2
29 7 .29 4 .29 0 15 1		
30 7 21 4 16 0 00 0		
	30 1 .06 0 .61 0 .00 0	30 8 .114 716 -4 "
后 大点 大	To the the plan	14-14-14-1
111 5 10 4.9 3.	11. 5. 10. 4.9. 3.	111. 5.10. 4.5. 2.

To find the Nutation of a Star in Right Ascension.—Find in Parts I. II. the Equations or recoponding to the arguments of R. A. at the top of the tables, connect them according the signs, and to the log, of the sum or difference add the log, tangent of the star administration, the sum will be the log, of an arch, to which apply the equation of the equation, the sum will be the log, of an arch, to which apply the equation of the equation, the sum will be the log, of an arch, to which apply the equation of the equation, the sum will be the Nutation in Right Ascension in seconds of a degree, which divides by will be reduced to seconds of time.

To find the Nutation of a Star in Declination.—Increase the arguments of R. A. Parts I. by S signs, and connect the corresponding equations of those tables, which will be the table of declination. Note. In patting the R. A. of the star equal to 3 sizes, the star in the connect the equation of the obliquity of the ecliptic.

EXAMPLE. Required the Nutation of a Pegasi, in R. A. and Doel. July 16, 1930 (R. A. Tab. 8 18-139-57)

Node N. A. 11. 26. 0

11. 17.57 Part 1. -2//15 11. 9.57 Part II. -1, 15

Diff. + 3s = 2s.17°57' Part I. - 10'74 Sum + 3s = 2 2.57 Part II. - 12

9. 30log 0.96843 E Declination 140 14/ tang. 9.40425

Nut. in Dec. -3.36

6s.— D Node=5s. 4d. 6s.+ D Node=5s. 26d.

Arch — 20.4 log 0.37273 os. + proceeding the Land of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the arguments Part I. U. of Right Assemblan and the Declination of the Star was south, the Assemblan and the Declination of the Star was south and the Declination of the Star was south as the Declination of the Star was south at the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Star was south as the Declination of the Declination of the Star was south as the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Declination of the Decl dination must be becaused 6 si,

To find the Augmentation of the Moon's Semidiameter, by the altitude of the Nonagesimal and the apparent distance of the Moon therefrom.

Part L. Part II.
Arg. = Alt, nona - ap, dist. # PART III.
Alt prince and die Arg. D's Argument ('s Parallag in Tat.
from nona, Som of Corr. true 0' 10' 30' 30' 40' 100' 100'
D + + + Part I. + South
0 0.000 4.000 7.009 30 10 0000 60 0 000000 30 00,60 00 92 10 23 10 52 10 52
1 U. 144 27 16 29 1 a 1 0 00 1 5 0 10 1000 250 500 701 0 10
4 4 0 02 3 30 0 000 170 360 500 750 961 17
6 0 86 4 817. 43 24 6 9 0 05 2 20 0 000 120 240 380 520 670 83
11 1. 56 5. 37 7. 74 19 9 8 0 10 0. 40 0 .000 040 .080 130 190 900 84
In L 845 587 23 17 10 3 0 .11 0. 20 0 000 020 050 090 130 180 24
1 38 0 69 7 87 16 11 2 0 13 North
15 9 90 6 90 10 11 6 0 14
17 2 395 997 97 13 12 4 0 16 0 20 0 01 0 020 01 + 0 020 05
12 66 6 18 8 03 11 13 1 0 18 0 40 0 000 000 000 000 000 000 000 000
11 5 5 3 5 5 10 13 5 0 49 1 0 0 000 000 000 000 000 000 000 000
5. 076. 458. 10 8 14 2 0 21 1 20 0 000 000 110 160 190 20 24
25 3 336 692 14 6 14 5 0 22 2 0 0 000 000 180 260 330 300 4
25 3 46 6, 70 8, 15 5 15 2 0 24 2, 40 0 000 130 240 350 460 550
20 3 396 798 16 4 15 5 0 25 3 0 0 000 110 000
3. 72 6 868. 17 S 15 8 0 26 3. 39 0 000 170 330 480 620 630 73 32 32 46 948. 18 2 16 1 0 27 4 0 0 000 170 330 480 620 750 88 3 977 028. 18 1 16 4 0 98 5 0 000 190 370 550 720 871 03
30 1. 09 7. 09 8. 18 0 16 7 0 99 8 0 0 000 290 470 700 311 121 32
11 to 1 D
Apr. Part IV. Arc. 4's Horis Sand Dis.
im of 10 15 16/ two equations correct
e. Eq. 407 500 07 107 207 307 407 507 07 107 207 307 407 507 ponding in the arguments at the top and
1 0.15 0.14 0.12 0.10 0.03 0.06 0.04 0.02 0.00 0.02 0.04 0.06 0.09 0.11 With this sum or dif-
3 0.48 0.42 0.36 0.30 0.24 0.18 0.12 0.06 0.00 0.05 0.13 0.17 0.21 ference take out the
2 0.20 0.20 0.24 0.20 0.16 0.12 0.08 0.04 0.02 0.00 0.02 0.04 0.06 0.09 0.11 With this sum or diff- 3 0.48 0.42 0.36 0.30 0.24 0.18 0.12 0.06 0.00 0.06 0.13 0.17 0.21 ference take out the corresponding correc- 4 0.54 0.56 0.43 0.41 0.33 0.25 0.16 0.08 0.00 0.08 0.17 0.25 0.34 0.43 0.00 0.00 0.00 0.00 0.10 0.21 0.32 0.44 0.05 0.00 0.00 0.00 0.00 0.00 0.00
0.96 0.94 0.73 0.61 0.49 0.37 0.35 0.12 0.00 0.13 0.25 0.32 0.31 0.64 III. is to be found with 1.12 0.98 0.85 0.71 0.57 0.43 0.29 0.15 0.00 0.15 0.29 0.44 0.50 0.75 the g Parinlat at the 1.381 1.12 0.97 0.31 0.65 0.49 0.33 0.17 0.00 0.17 0.34 0.51 0.68 0.06 to p and her true lat. at 1.41 1.26 1.09 0.91 0.73 0.55 0.37 0.19 0.00 0.19 0.38 0.57 0.77 0.76 0.96 the side, but in solar 1.60 1.41 1.21 1.010 32 0.62 0.41 0.21 0.00 0.19 0.38 0.57 0.77 0.77 0.96 the side, but in solar
8 1.28 1.120.970.31 0.65 0.49 0.33 0.17 0.000 1.70 0.34 0.51 0.68 0.86 top and her true lat. at
9 1.44 1.25 1.09 0.91 0.73 0.55 0.37 0.19 0.00 0.17 0.34 0.31 0.88 0.86 the side, but in solar 10 1.50 1.41 1.21 1.01 0.32 0.62 0.41 0.21 0.00 0.21 0.42 0.63 0.85 1.07 Eclipses this p. is no-
11 1.76 1.55 1.33 1.12 0.90 0.68 0.45 0.23 0.00 0.23 0.46 0.70 0.94 1.18 thing. Councet these
13 1.921.69 1.45 1.22 0.93 0.74 0.49 0.25 0.00 0.25 0.51 0.76 1.02 1.28 three parts and with 2.03 1.83 1.57 1.32 1.06 0.20 0.54 0.27 0.00 0.27 0.55 0.33 1.11 1.39 the same energies side
14 3.24L.97l.70.1.42[1.06.0.00.0.40.270.000.270.550.831.1111.39 the sam enter the side 4.24L.97l.70.1.42[1.140.360.580.290.000.290.590.891.191.50 column of P. IV. and 4.24L.97l.13.21.52[1.220.920.620.310.000.310.630.961.28] followed by the first Horizontal State of the first
bled, with its sign to the sum of the three first parts will give the Aug. of the (* S. D. That in Ex. I. Prob. 5. Appendix. The Alt. Nonng. is 2a.7° 59′ Dis. Nonng. (D.+P) 20° (S. D. by N. A. 16′ 27″, 7. Hence Arg. P. I. are 2s.7° 59′ +20° 46′ that is 2s. 28° 45′ 11s. 17° 13′ to which correspond + 8″ 18 + 6″ DI = 1. 17′ 13′ 13′ to which correspond + 8″ 18 + 6″ DI = 1. 17′ 13′ 13′ 13′ 13′ 13′ 13′ 13′ 13′ 13′ 13
14. 170 13' to which correspond + 2" 12 1, are 2s.70 59' + 200 46' that = 2s. 280 44'
1 ls. 17° 13' to which correspond + 8" 18 + 6" 01 = + 14" 19. This gives in P. II. + 0" 111. is 0". The sum of the three perts is + 14". 4 with which and the (S. D. 16" 27" 7. IV. is nearly + 0" 8, this connected with 14". 4 gives the Aug. of ('s S. D. 15" 2, as in Prob.
Appendix.
Tighting CaOOOC

TABLE XLV.

Equation of Second Differences to be applied to the mean length side on histories with a sign contents to that of the mean of the mean of the second differences.

-	_				-80	comit i	Per Marie			_		-	-11
up, time after noon	100	-	3/ 1	1 1	- 59	T 6/		8, 1	. 0	10%	- 10	1 12	
or midnight.	-	-	-		-	-	1	101					10
A m I h m	11	-01	"	100	-	0:0	0.0	0.0.	0.0	0.0	10.6	0.0	5 114
0. 0. 12. 0	0.0	0.0	0.0	0.0	0.0				3.7	4.1			978
PO 10 11-50	0.4	0.8	T.2	1.6	2.0	3.5		17.02					90
0.20 11.40	0.8	1.6	2.4	0.2	4.1	4.9	0.7	6.8	7.0	8.1			411
		2.4	3.6	4.0	6.11	7/3	8.2	9.6	10.8	12.0		au o	400
0.30 11.30	1 2 1		4.7	2 2	7.9	9.4	11.0	12.6	10:2	10.7	17.1		300
0.40 11.20	1.6	3.1		200	9.7	11.6		15.0	37-4	119.4	SER OF	100:0	211
0.30 11.10	1.9	3.9	9.3	7.5	9.1						_	100	31
1. 0 11: 0	0.7	4.6	6.9	9.2	11.5	13.7	16.0	13.3	20.5	22.5		3 100	
	2.5			10.5	13.2	15.8		21.1	23.7	20.3	10.0	121.1	386
1.10 10.50	2.0	5.3	7.9			17.8		23.7	26. 7		100	200.4	611
1.20 10.40	3.0	5.9	8.9	1149	14.8			26.2	10.0			300.	4
1.30 10.30	3.3	6.6	9.8	13.1	16.4	19.7			32.3			area.	ш
1.40 10.20	3.6	7.2	10.8	14.6	17.9	21.0	20.1	28.7 31.1					91
	3.9	7.8	11.6	15.5	19.4	23.3	27.2	31.1	54.9				a j
1.00 10.10	20.0					25.0	29.2	33.3	137.4	83.7		50.	0
2. 0 10. 0	4.2	11.3	123	16.7		20.0	10000		39.2	44.4			3
2,10 9.50	4.4	8.9	13.3	17.3	22.2	26.6	31.1	35.5					
	4.7	9.4	14.1	18.8	23.5	28.2	2 38.19	37.6	42.3	47.0			धा
				19.8	24.7	29.7	34.6	39.6	44.5	49.1	100	E 201	мп
≘.30 0.30	4.9	9.9	14.8		25.9	31.1		41.5	46.7	51.	157	0 62	
= 40 9.20	5.2	10.4	13.6	20.7				45.5	48.7	24	1 3/5	5 164	9
1.50 9.10	5.4	10.B	16.2	21.6	27.1	32.2							
		11 17	16.9	22.5	28.1	33.7	7 39.4	15.0	30.6		5 61	3 57.	-21
3. 0 9. 0	n.6	2.11			29.1	35.		46.6	52.4	68.	3 64	1 09	9
3.10 8.50	5.5	11.7	17.5	23.3			1000	48.1	04.3		I No.	2 72.	-
3.20 8.40	6.0	12.0	18.1	24-1	30.1			49.6				2 74	4
3.30 8.30	6.2	12.4	18.6	24.8	21.5		2 43.4		55.8	6.7	7 70		
3.40 8.20	6.4	12.7	19.1	25.5	31.8			150.9			2 (20)	2 40	
		13.0	19.6	26.1	32.6	39.	1 45.7	52 2	58.7	63	2 71	4 48-	.2
3.50 8.10	6.0	-					0 40 7	53.3	60.0	66.	T 35	3 E0	
4. 0 8. 0	6.7	13.3	20.0	26.7	33.3							HIRO	
4.20 7.40	6.9	13.8	20.8	27.7	34.6			55.4	62.3		5 E	HILL	ы
4.40 7.20	7-1	14.3	21.4	28.5	35.6	42.	8 49.9	57.0	64.2	71.	2 15	2 20	MI.
	12.5		21.9	29.2	36.8			58.3	65.6	100	9 100	(E) 117	
+ 5 0 7- 0	11.3	14.6		29.6	37.8	-		59.3	66.7	74	1 11	5 100	15
	77 A												_
5.20 6.40	7.4	14.8	24. 1					50.0	677.3	74	B 250	4 00	-
			22.4	29.9	37.4	44.	9 52.3	59.8	67.3	74	E 155	5 00	Н
5.40 6.20	7.5	15.0	22.4		37.4	45	9 52.3	60.0	67.5	70	0 112 0 112	5 50	a a
				29.9	37.4	45	9 52.3	60.0	67.3	74	0 12	5 90	10
6. 0 6. 0	7.5	15.0	22.4	30.0	37.4	44. 45. Secon	9 52.3	60.0	67.5	74	0 02	5 00	10.1
5. 40 6.20 6. 0 6. 0	7.5	15.0	22.4	29.9	37.4	\$44. Secon	9 52.3 0 52.5 ad Differ 2" 3"	ence.	67.3	74	0 M2	5 50	10.0
5.40 6.20 6. 0 6. 0 Ap.time after no or midnight.	7.5	15.0	22.4	30.0	37.4 37.5	Secon	9 52.3 0 52.5 nd Differ	160.0 rence.	67.5	6"	P I	5 80	100
5.40 6.20 6. 0 6. 0 Ap, time after no or midnight. R m h m	7.5	15.0	22.4	30.0	37.4 37.5	\$45. Secon	9 52.3 0 52.5 ad Differ 2" 3"	60.0 rence.	67.3 67.5	6"	77 0 0		10.0
5.40 6.20 6. 0 6. 0 Ap, time after no or midnight. 6 m h m 0. 0 12. 0	7.5	15.0 15.0 15.0	22.4 23.5 30" 0.0	40" 0.0	50° 0.0	144. Secon 1" " 0.0	9 52.3 0 52.5 nd Differ 2" 3" " " 2.0 0.0	60.0 rence.	67.5	6"	77 1	0.1 8	43
5.40 6.20 6. B 6. 0 Ap, time after no or midnight.	7.5 7.5 00 107 0.0 0.1	15.0 15.0 15.0 0.0 0.1	22.4 22.5 30" 0.0 0.2	40" 0.0 0.3	50" 0.0 0.3	Second 1" " 0.0 0	9 52.3 0 52.5 nd Differ 2" 3" " " 2.0 0.0	60.0 rence.	67.5 67.5 0.0 0.0	6"		0.1 8	
5.40 6.20 6. 0 6. 0 Ap, time after no or midnight. 6 m h m 0.0 12 0 0.10 11.50 0.29 11.40	7.5 7.5 00 0.0 0.1 0.1	20" 0.0 0.1 0.3	22.4 22.5 30" " 0.0 0.2 0.4	99.9 30.0 40" 0.0 0.3 0.5	50" 0.0 0.3 0.7	Secon 1" 0.0 0.0 0.0	9 52.3 0 52.5 nd Differ 2" 3" " 2.0 0.0 0.0 0.0	60.0 rence. 4" 0 0.0 0 0.0 0 0.1	67.3 67.5 0.0 0.0 0.1	6"		0.1 8	43
5.40 6.20 6. 0 6. 0 Ap, time after no or midnight. 6 m h m 0.0 12 0 0.10 11.50 0.29 11.40	7.5 7.5 00 107 0.0 0.1	20" 0.0 0.1 0.3 0.4	30" 0.0 0.2 0.4 0.6	40" 0.0 0.3 0.5 0.8	50° 0.0 0.0 0.3 0.7 1.0	Second 17 17 18 18 18 18 18 18 18 18 18 18 18 18 18	9 52.3 0 52.5 nd Differ 2" 3" " 2.0 0.0 0.0 0.0 0.0 0.0	60.0 rence. 4" 0.0.0 0.0 0.0 0.1 1 0.1	67.3 67.5 0.0 0.0 0.1 0.1	6" " 0.0 0.0 0.1 6.1	0.0 0.1 0.1	0.1 8	43
5.40 6.20 6. 0 6. 0 Ap. time after no or midnight.	7.5 7.5 00 0.0 0.1 0.1	20" 0.0 0.1 0.3 0.4	22.4 22.5 30" " 0.0 0.2 0.4	99.9 30.0 40" 0.0 0.3 0.5	50° 0.0 0.0 0.3 0.7 1.0 1.3	44. 5 45. Second 17 7 10.0 1	9 52.3 0 b2.5 nd Differ 2" 3" " 2.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0	60.0 rence. 0 0.0 0 0.0 0 0.1 1 0.1	67.3 67.5 0.0 0.0 0.1 0.1	6" 0.0 0.0 0.1 0.1	0.0	0.1 0 0.1 0 0.2 0	43
5,40 6.20 6, 0 6, 0 Ap, time after no or midnight. \$\hat{k} m \ h \ \pi 0.0 12, 0 0.10 11.50 0.29 11.40 0.30 1.40 0.40 11.20	7.5 7.5 0.0 0.1 0.1 0.2 0.3	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5	30" 0.0 0.2 0.4 0.6	40" 0.0 0.3 0.5 0.8	50° 0.0 0.0 0.3 0.7 1.0	44. 5 45. Second 10 10 10 10 10 10 10 1	9 52.3 0 52.5 nd Differ 2" 3" " 2.0 0.0 0.0 0.0 0.0 0.0	60.0 rence. 0 0.0 0 0.0 0 0.1 1 0.1	67.3 67.5 0.0 0.0 0.1 0.1	6" " 0.0 0.0 0.1 6.1	0.0 0.1 0.1	0.1 8	43
5, 40 6, 20 6, 0	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.3	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 10.6	30" 0.0 0.2 0.4 0.6 0.8 1.0	40" 0.0 0.3 0.5 0.8 1.0 1.3	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6	5 45. Secon 1" , 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0	9 52.3 0 b2.5 nd Differ 2" 3" " 2.0 0.0 0.0 0.0 0.0 0.0 0.1 0.1	60.0 (a) 0.0 (b) 0.0 (c) 0.1 (d) 0.1 (d) 0.1 (e) 0.	67.3 67.5 0.0 0.0 0.1 0.1	6" 0.0 0.0 0.1 0.1	0.0	0.1 0 0.1 0 0.2 0 0.3 0	43
5,40 6.20 6, 0 6, 0 Ap, time after no or midnight. \$\hat{k} m \ h \ \pi 0.0 12, 0 0.10 11.50 0.29 11.40 0.30 1.40 0.40 11.20	7.5 7.5 0.0 0.0 0.1 0.1 0.2 0.3 0.3	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0	40" 0.0 0.3 0.5 0.8 1.0 1.3	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6	Second 1" " 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0	9 52.3 0 52.5 ad Differ 2" 3" " 2.0 0.0 0.0 0.0 0.0 0.0 0.1 0. 0.1 0.	160.0 rence. 4/7 / / / / / / / / / / / / / / / / / /	67.3 67.5 0.0 0.0 0.1 0.1 0.1 0.2	6" 0.0 0.0 0.1 0.1	0.0 1.0 1.0 2.0 2.0 2.0	0.1 0 0.1 0 0.2 0 0.3 0	1日日 日本
5, 40 6, 20 6, 0	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.3 0.4 0.4	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 0.8 4 0.9	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0	29,9 30.0 40" 0.0 0.3 0.5 0.8 1.0 1.3	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.3	Second 17 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	9 52.3 0 52.5 ad Differ 2" 3" " " " 0.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0	160.0 rence. 4/7 // // // // // // // // // // // // /	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2	6" 0.0 0.0 0.1 0.1 0.2 0.2 0.3	0.0 1.0 2.0 2.0 2.0 2.0 2.0 3.0	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0	1. 日日日日日日 日本
5,40 6.20 6, 0 6, 0 Ap, time after no or midulght. 6 m h m 0.0 12, 0 0.10 11, 50 0.29 11, 40 0.30 11, 30 0.40 11, 20 0.50 11, 10 1, 0 11, 0	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.3 0.4 0.4	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 0.8 4 0.9	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0	29,9 30.0 40" 0.0 0.3 0.5 0.8 1.0 1.3 1.3 2.0	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6	144. Second 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0	9 52.3 0 52.5 ad Differ 2" 3" " " " 0.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0	160.0 rence. 4" 0.0.0 0.0.0 0.1 1.0.1 1.0.1 1.0.2 1.0.2 1.0.2	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2	6" 0.0 0.0 0.1 0.1 0.2 0.3 0.3	0.0 1.0 1.0 2.0 2.0 2.0	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0	1日日 日本
5,40 6,20 6, 0 Ap, time after no or midnight. \$\begin{array}{l} \be	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.3 0.4 0.4	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 0.8 4 0.9 5 1.0	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0	29,9 30.0 0.0 0.3 0.5 0.8 1.0 1.3 1.3 2.0 2.2	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.2 2.5 2.7	144. Second 177 7 8.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0	9 52 . 3 0 52 . 5 ad Differ 2" 3" " 0.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0	60.0 rence. 47 0 0.0 0 0.0 0 0.1 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2 0.3	6" 70,0 0.0 0.0 0.1 0.2 0.3 0.3 0.3	0.01122	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0	1. 日日日日日日 日本
5,40 6.20 6, 0 6, 0 Ap, time after no or midnight \$\bar{R}\$ m h m 0, 0 12, 0 0, 10 11, 40 0, 30 11, 30 0, 40 11, 20 0, 50 11, 10 1, 0 11, 0 1, 10 10, 50 1, 20 10, 40 1, 30 10, 30 1, 30 10, 30	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.3 0.4 0.4 0.4	15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 0.8 4 0.9 5 1.0	30" "0.0 0.2 0.4 0.6 0.8 1.0 1.3 1.5 1.6	29,9 30.0 0.0 0.3 0.5 0.8 1.0 1.3 1.3 2.0 2.2	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6	144. Second 177 7 8.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0	9 52.3 0 52.5 ad Differ 2" 3" " " " 0.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0	60.0 rence. 47 0 0.0 0 0.0 0 0.1 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2 2 0.2	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2	6" 0.0 0.0 0.1 0.1 0.2 0.3 0.3 0.3	0.0 1.0 2.0 2.0 2.0 2.0 2.0 3.0	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0	1. 日日日日日日 日本
5,40 6.20 6, 0 6, 0 Ap, time after no or midnight. \$\begin{align*} alig	7.5 7.5 0.0 0.1 0.1 0.3 0.3 0.4 0.4 0.5 0.4	15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.5 1 0.6 1 0.8 4 0.9 5 1.0 5 1.4 5 1.2	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0 1.1 1.3 1.5 1.6	29.9 30.0 0.0 0.3 0.5 0.8 1.0 1.3 1.3 2.0 2.2	50" 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.2 2.7 3.0	Second 1	9 52 . 3 0 52 . 5 ad Differ 2" 3" " 0.0 0.0 0.0 0.0 0.0 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0 0.1 0.0	60.0 rence. 47 0 0.0 0 0.0 0 0.1 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2 0.3	6" 70,0 0.0 0.0 0.1 0.2 0.3 0.3 0.3	0.01122	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0 0.4 0 0.4 0 0.4 0	1. 日日日日日日 日本
5,40 6.20 6, 0 6, 0 Ap, time after no or midnight \$\bar{R}\$ m h m 0, 0 12, 0 0, 10 11, 40 0, 30 11, 30 0, 40 11, 20 0, 50 11, 10 1, 0 11, 0 1, 10 10, 50 1, 20 10, 40 1, 30 10, 30 1, 30 10, 30	7.5 7.5 0.0 0.1 0.1 0.3 0.3 0.4 0.4 0.5 0.5 0.4 0.5 0.5	15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.5 1 0.6 1 0.8 4 0.9 5 1.0 5 1.4 5 1.2	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0 1.1 1.3 1.5 1.6 1.8	99.9 30.0 0.0 0.3 0.5 0.5 1.0 1.3 1.3 2.0 2.2 2.4 2.6	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.2 2.5 2.7 3.0 3.2	Secondary 1	9 52.3 0 52.5 and Difference 50 50 50 50 50 50 50 5	60.0 rence. 4" 0 0.0 0 0.0 0 0.1 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2 2 0.2 2 0.3	67.5 67.5 0.0 0.0 0.1 0.1 0.1 0.2 0.2 0.3 0.3	6" 0.0 0.0 0.1 0.1 0.2 0.3 0.3 0.3 0.4	0.0 0.1 0.2 0.2 0.3 0.5 0.4 0.4	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0 0.4 0 0.4 0 0.4 0	11日日日日 日本日の日本
Ap, time after no or midnight. \$\bar{R}\$, \$\text{im}\$ a fter no or midnight. \$\bar{R}\$ m & h m o. 0 12. 0 0. 10 11. 50 0. 20 11. 40 0. 30 11. 30 0. 40 11. 20 0. 50 11. 10 1. 0 11. 0 1. 10 10. 50 1. 20 10. 30 1. 30 10. 30 1. 40 10. 20 1. 50 10. 10	7.5 7.5 0.0 0.1 0.1 0.2 0.3 0.4 0.4 0.4 0.5 0.4 0.5	15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.5 1 0.6 1 0.8 4 0.9 5 1.0 5 1.4 5 1.2	22.4 23.5 0.0 0.2 0.4 0.6 0.8 1.0 1.1 1.3 1.5 1.6	29.9 30.0 0.0 0.3 0.5 0.8 1.0 1.3 1.3 2.0 2.2	50" 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.3 2.7 3.0 3.2	Second 100 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0	9 52 3 0 52 5 ad Drifte 2" 3" 2 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	160.0 rence. 4" " 0 0.0 0 0.0 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2 2 0.3 2 0.3	67.3 67.5 0.0 0.0 0.1 0.1 0.2 0.2 0.2 0.3 0.3	6" 0.0 0.0 0.1 0.1 0.2 0.3 0.3 0.3 0.4 0.4	0.0 0.1 0.2 0.2 0.3 0.5 0.4 0.4 0.5	0.1 0 0.1 0 0.2 0 0.3 0 0.3 0 0.4 0 0.4 0 0.4 0 0.5 0 0.5 0 0.6 0 0.6 0 0.7 0 0.8 0 0.	111111111111111111111111111111111111111
6, 40 6, 20 6, 0 Ap, time after no or midnight. 8 m h m 0. 0 12 0 0 0.29 11, 40 0, 30 11, 30 0, 40 11, 20 0, 50 11, 10 10, 50 1, 20 10, 40 1, 30 10, 30 1, 40 10, 20 1, 50 10, 10 10, 20 1, 50 10, 10 10, 20 1, 50 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20 10, 10 10, 20	7.5 7.5 000 1077 0.00 0.1 0.1 0.2 0.3 0.3 0.4 0.4 0.6 0.1 0.4 0.6 0.1 0.6 0.1	15.0 15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 1 0.8 4 0.9 5 1.0 5 1.2 8 1.3	22.4 23.5 0.0 0.2 0.4 0.6 1.0 1.1 1.3 1.6 1.8 1.9	99.9 30.0 0.0 0.3 0.5 0.5 1.0 1.3 1.3 2.0 2.2 2.4 2.6	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.2 2.5 2.7 3.0 3.2	Secondary 1	9 52 3 0 52 5 ad Differ 2" 3" " " " " 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	160.0 rence. 4" 0 0.0 0.0 0.0 0.1 1 0.1 1 0.2 1 0.2 2 0.2 2 0.3 2 0.3 2 0.3 2 0.3	67.3 67.5 0.0 0.0 0.1 0.1 0.2 0.2 0.2 0.3 0.3 0.3	6" 0.0 0.0 0.1 0.2 0.3 0.3 0.3 0.4 0.4	0.0 0.1 0.2 0.2 0.3 0.5 0.4 0.4 0.5	0.1 0 9.1 0 0.2 0 0.3 0 0.4 0 0.4 0 0.4 0 0.5 0 0.6 0 0.6 0	11日日日日 日本日の日本
5, 40 6, 20 6, 0	7.5 7.5 000 1077 0.00 0.1 0.1 0.2 0.3 0.3 0.4 0.4 0.2 0.2 0.4 0.5	15.0 15.0 0.0 0.1 0.3 0.4 0.5 1 0.6 0.8 4 0.9 5 1.0 5 1.4 6 1.3 7 1.4 7 1.5	22.4 23.5 0.0 0.2 0.4 0.6 1.0 1.1 1.3 1.6 1.8 1.9	40° 0.0 0.0 0.5 0.5 0.8 1.0 1.3 2.0 2.2 2.4 2.6	50° 0.0 0.0 0.3 0.7 1.0 1.3 1.6 1.9 2.3 2.7 3.0 3.2 3.5 3.7	Second 100 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0	9 52 3 0 52 5 ad Differ 2" 3" " " " " 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	160.0 rence. 4" " 0 0.0 0 0.0 1 0.1 1 0.1 1 0.2 1 0.2 2 0.2 2 0.3 2 0.3	67.3 67.5 0.0 0.0 0.1 0.1 0.2 0.2 0.2 0.3 0.3	6" 0.0 0.0 0.0 0.1 0.2 0.3 0.3 0.3 0.3 0.4 0.4 0.4	0.0 0.1 0.1 0.2 0.5 0.5 0.4 0.4 0.5 0.5 0.5	0.1 0 9.1 0 0.2 0 0.3 0 0.4 0 0.4 0 0.4 0 0.5 0 0.6 0 0.6 0 0.6 0	111555 444566 1077
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5, 40 6, 20 6, 0	7:5 7:5 0.0 0.0 0.1 0.1 0.2 0.3 0.3 0.4 0.4 0.4 0.6 0.1 0.6 0.1 0.6 0.1 0.6 0.1 0.6 0.1 0.6 0.6 0.1 0.6 0.6 0.6 0.6 0.6 0.6 0.6 0.6 0.6 0.6	15.0 15.0 10.8 0.8 0.1 0.3 10.4 10.6 10.	22.4 23.5 0.0 0.2 0.4 0.6 1.0 1.3 1.5 1.6 1.8 1.9 2.1 2.2 2.3 2.6 2.7	29,9 30.0 0,0 0,0 0,5 0,5 0,8 1,0 1,3 1,3 2,0 2,2 2,4 2,6 3,5 3,5 3,5 3,5	37.5 50°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°°	1 44. 5 45. Second 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.0 0 0.1 0 0	9 52 3 0 52 5 ad Differ 2" 3" " " " 0 0 0 0 0 0 0 0 0 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	160.0 rence. 4" 0 0.0 0 0.0 0 0.1 1 0.1 1 0.1 1 0.2 1 0.2 2 0.3 2 0.3 2 0.3 3 0.3 5 0.4	67.3 67.5 67.5 0.0 0.0 0.0 0.1 0.1 0.2 0.2 0.2 0.3 0.3 0.3 0.4 0.4 0.4 0.4 0.4	6" 0,0 0,0 0,0 0,1 0,2 0,3 0,3 0,3 0,4 0,4 0,4 0,5 0,5 0,5 0,5 0,6	0.0 0.1 0.2 0.3 0.4 0.4 0.5 0.5 0.5 0.6 0.6	0.1 0 0.1 0 0.2 0 0.3 0 0.4 0 0.4 0 0.4 0 0.6 0 0.6 0 0.6 0 0.7 0	111 EE & 44 A & 6 10 10 10 10 10 10 10 10 10 10 10 10 10
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This Table contains the Latitudes and Longitudes of the most remarkable Harbours, Islands, Shoals, Capes, &c. in the WORLD, founded on the latest and most accurate Astro-

nomical observations, surveys and charts. The Longitudes are reckoned from the meridian of Greenwich. I. COAST OF THE UNITED STATES OF Long. D. M. D. M. AMERICA. Light houses on Baker's Long. Island . . 42 34N 70 47W Lat. 42 34 D. M. D. Beverly 70 52 42 33 70 52 ENTRANCE of St. SALEM 42 32 70 50 45 8W 7N 67 Marblchead Croix River . Nahant Point (N. E Island of Campo-Bello Point of Boston har-(middle or West pas-42 28 70 54 bour) sage of Passamaquod-42 20 70 54 44 57 66 54 Boston light-house dy Bay) . 42 23 71 BOSTOÑ Wolves' Islands 45 66 41 CAMBRIDGE (Mass) 42 23 71 E. end of Grand-Manan 44 47 66 43 42 11 Grand-Manan N. head 44 53 70 41 66 45 Scituate light 41 59 70 34 do. West end . 44 40 66 55 Plymouth lights 42 6 70 14 Titmanan Light . 44 25 67 40 Race Point light 42 5 70 Entrance of Machias CAPE COD light . 44 44 67 20 41 43 69 56 Chatham light Sandy Point or Malabar 41 34 Gouldsboro' Harbour 44 34 69 59 67 52 43 52 68 9 Shoal of George, Mount Desert Rock 41 34 ह 67 40 Great Shoal S. É. P. Long-Island (south of W. P. 41 42 67 59 do. do. Mount Desert or en-41 48 67 47 N. E. P. trance of Blue-Hill do. do. 41 53 67 43 68 31 do. North Shoal Bay) 41 51 2 67 26 44 00 69 40 lale of Holt do. Third Shoal 41 47 5 67 19 Castine 44 24 68 46 do. East Shoal 43 50 Matinicus Island 68 55 NANTUCKET light-41 23 70 00 Wooden Bald Rock 43 45 68 54 house 43 44 69 15 Sancoty head on Nan-Island of Manheigin Penmaquid Point . 43 48 69 27 41 16 tucket Island 69 59 41 14 Bantum Ledges . 43 42 69 33 Tom Nevers head Kennebeck River en-Nantucket South Shoal 69 55 43 43 69 47 (true lat. by Blunt) . 41 4 Cape Poge, (Vineyard) 41 25 trance 70 25 Seguine Island light 43 41 69 46 43 40 69 52 Squibnocket-head Cape Small point Casho's Ledge (shoulest southwesterly part of 70 48 43 69 11 Martha's Vineyard) . 41 17 part) Alden's Ledge (off Cape Gay Head light-house, 43 28 41 21 70 50 Elizabeth) . . . 70 9 (Vineyard) 70 49 43 52 Noman's Land Island . 41 15 Brunswick 70 17 New-Bedford light hou. 41 35 70 53 PORTLAND light-hou. 43 39 43 33 Buzzard's Bay entrance 41 **|70 50** Cape Elizabeth 70 15 . . 41 29 71 18 43 28 Saco River entrance 70 26 NEWPORT Rhode Island light-hou. 41 28 71 23 70 23 Wood Island L. House 43 27 Point Judith L. house . 41 24 71 29 70 41 Agamenticus Hill 43 16 71 53 Watch Hill P. light . . 41 20 43 21 70 26 Cape Porpoise Watch Hill P. light . . . 41 14 72 10 Wells Harbour 43 19 70 33 41 10 71 39 Bald Head 43 13 70 35 Block Island Cape Neddock Nubble 43 10 70 36 43 Thames York River . . 70 39 trance of .41 21 72 11 York Ledge 43 6. 70 34 river) light-house 72 45 Boon Island light 43 6 70 31 Falkland Island L.house 41 15 Boon Island Ledge . NEW-HAVEN entr'ce. 41 17 72 58 43 70 27 PORTSMOUTH light-Montock Point (E. end 70 44 house of Long-Island) light 70 46 71 55 Portsmouth 43 house East Hampton in do. . 41 00 72 16 Isles of Shoals light 42 56 70 39 40 42 74 00 NEWBURYPORT NEW-YORK City lights on Plumb Island 42 48 70 51 New-York light-house Ipswich entrance 40 28 74 . 42 43 . 42 42 on Sandy Point 70 49 74 30 40 30 70 41 Perth Amboy Squam light 74 22 39 30 Sandy Cove (or Bay) . 42 41 70 39 Little Egg Harbour 39 18 74 34 CAPE ANN light hous-Great Fgg Harbour 38 57 74 58 Cape May es on Thatcher's Isl .|39

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	1	D.	M.		M.				M.		M.
4	Cape Charles	37			15W		Muskito or N. Smyrna	Γ.		-	
3	Cape Henry light	36	56	76	18	ž.	entrance	28	52N	80	56W
3	Norfolk (Vir.)	36	53	76	34	orida.	Cape Canaveral	28	19	80	33
7	Petersburgh (Vir.)	37	12	77	44	F_{bo}	Outer breakers off do.	28	20	80	13
۱. ا	York Town (Vir.)	37	12	76	39	g.	Las Tortolas or Hum-			l	
2	RICHMOND (Vir.) .	37	30	77			mocks	27	35	90	30
Æ	Amnapolis (Mar.)			76	37	Coast	Hillsborough Isl. N. P.	27	31	90	19
Ę	ALEXANDRIA (Vir.)			77	4	ŏ	— S. P	27		80	13
8	WASHINGTON City	38	5 3	77	3	782	Mount Pelado or Baid			l	
pres	Chincoteague Shoals					\mathbf{E}_{0}		27	1	80	11
	(on Maryland shore)			75				26		80	8
125		39		76			Cooper's Hill	26	42	180	3
50		35		75		ľ	Sand Hills	26		180	3
5		35		75				56		90	6
Г	Extreme shoal off do		3	75			Middle River entrance		7	80	7
	Deep soundings off do.			75	9	١.	CAPE FLORIDA			90	9
		35		75		1		25	20	90	20
		34		76		L	Cuyo Largo or Long	١.		00	, l
	Extreme shoal off do	34		76 76				24		190	
				76				24		80	
	Deep soundings off do.			76			Sombrero or Hat Key .	24	3Z	81	
		34 34			46		Loge Key	24	28	81 81	
ŀ		34		76		13		24	25	1 -	
ğ		34			14	Ę	Sand Key or C. Arens	124	21	81	37
North-Carolina.		34		77		K	C Florido	L.	90	82	31
ş	Bear Inlet	34	36	77		5	C. Florida Tortugas Islands and	34	20	02	31
P		34	34		30	3	Pople N W nest	04	7.4	83	2
\$	Stump Inlet	34	31	77		Sogs	Banks. N. W. part	6	37		45
δ	New Topsail Inlet	34	97		44	Ľ	— N. E. do	67	33	1	45
r	Sandy Inlet	34	19		55	1	— S. E. do	01	25 _.		00
	Deep Inlet	34	14		00	3	Key Marquis	24	30		13
	WILMINGTON	34	17		10		Boca Grande or Great		30	-	
	Brunswick	34	3		10	١.		24	39	88	11
	Smithville	33	54	78		ı		24		91	
ŀ	New Inlet	33	57	78	6	ı	Island of Pines	24	42	81	
I	CAPE FEAR	33	48	79	9	ı		24		81	
ı	Extreme shoal off do	33	36	77	47	ı.		24		81	17
	Deep soundings off do.	33	11	77	26	ľ		24		81	16
4	Lockwood's Folly Inlet	33	53	78	25	١.	Cana Sable on Tomaka	04		81	19
棒	Shallot	33	51	78		ğ	Cape Romano or P.			ł	
Ę	Little River Inlet	33	50		49	5	Larga	26	00	81	51
Ş	Brunswick Smithville New Inlet CAPE FEAR Extreme shoal off do. Deep soundings off do. Lockwood's Folly Inlet Shallot Little River Inlet GEORGETOWN Ditto light Shoals off do. Cape Roman CHARLESTON	33	25	79		K	Cape Romano or P. Larga Boca Grande ent. B.	1		1	
ş	Ditto light	33	13	73		ि	Carlos	26	41	1	10
ğ	Onosis off do	33	8	73			Boca Serraxota	27	16	82	
ľ	Ottant noman	33	2	79	6	Coast	Spirito Sante Bay ent.	27	38	82	47
		30	40		48			28		83	7
	Charleston light-house.				40	Ē	Keys of St. Martin	1 -	49	83	1
		32			59	z		59	8	83	5
	South Eddisto Inlet .	32	20	80	7	ı			23	83	5
	BEAUFORT (S. C.)	32		80		I	St. Marcos de Apalache		9		19
	Port Royal entrance . Tybee light .	32	8	80		Ĭ			48	84	29
eorgia.	SAVANNAH	72		90			St. George's Key, S. P.	29	30		18
ģ	St. Catherine's Sound .	32	2	31	3		Cape St. Blas	29	36	85	35
3		31		81			Bay St. Andres (E.point	L			
	St. Simon's Sound Brunswick (Geo.)	31	1	81	30			30	31	80	43
		31	10	Ì			Bay St. Rosa (W. point				
Ą	Amelia Sound (entrance of St. Mary's river).		44		42		of do.)		19	37	
Florida.	Cumberland Isl. (S. P.)	30	12	81			PENSACOLA		24		27
દ	Amelia Island (S. P.)	30	40	91			River Perdido		18		46
Č	River Nassau entrânce	30	20 00	18			Mobile Point		13		21
Ľ	River St. John entrance	20	91	81			MOBILE		40	98	
Đ,		29		81			Massacre Island		12	98	
Pi,	Island Anastasia, N. P.	90	51	81					13	88	
l '	— 8. P.	29		81 81			Candelarius, N. P.	39		88	
۱-		47	31	81	11	<u>. </u>	— S. P. '	(27)	28	189	12

TABLE XLVI. Latitudes and Longitudes.

9.

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1			at.		ong.				at.	Lo
ı	Van Bankar		M.	D.	M.	ľ		D.	M.	D.
1	Key Breton Entrance of MISSIS-	27	ZON	ניסן	18W	Ŀ	St. Christ'rs or St. Kitts - N. W. point	17	24N	ko i
1	SIPPI, N. E	29	12	99	9	Ē		17		63
ł	- La Balisa	29	8	89	6	H		17		63 1
1	S. E		59	89	13	F	Aves or Bird's I. about	15	40	63 4
1	— s. w.		56	89	29	ě		17		61 4
١.	NEW-ORLEANS .		57	90			St. Bartholomew, E. P.			62 4
į	Baton Rouge	30	36		13	į	St. Martin's, E. P.	19	.4	63
	Long-Island	29	10 52	90	14	۴	1 Tarrois_ "	18 18		63 62 :
1	I. Tonbalie, S. P. I. del Vino W. end .		56		24			18		63
Г	Bancos de Hostiones,		•	1	~-		Isle of Dogs, western .			63
1	- S. P.		50	91	44	l		18		63 :
1	₩. P.	29	26	93	4	ş	St.Croix orSt.Cruz E.P.			64 :
1	Iron Point or Point Fi-			_	_	lotands.	- W. P	17		64 :
•	erro	29	14	92	7	E	Anegado, S. P. of shoal			64
1	Deer Point	צצ	26	92		Æ	W. P	18 18		64 5
ł	Point del Pajaro River Lobos, ent	29	32	93	48 4	E	Virgin Gorda, E. P. Tortola, E. P.	18	90	64
1	Salt water Rev	199	96		28	٩	W. P.	19	25	64 4
ł	Constant Bay	29	27	93		1	St. John's	18	22	64
1	River Mermentao	29	38		11	ı	St. Thomas	18	22	64 !
1	Point ent. river Sabine			94	57	ı	Bird Kev	18	15	64 (
1	II felm Jo to M.	TW.	-1 T-	22		i	Serpent I. E. part .	18	19	65 1
ı	IL. Islands in the	_				ı	— Crab I. E. part .	19	10	65 1
ı	TO TAKE A D	١_	Lat.		ong.	ı	la a			
1	TRINIDAD,	D.			M. 30W	ł	Cape St. John or N. E.			65 : 66
1		10	4	61		l		18		67
1		10	9		55	Ŀ	Point Bruquen or N.W. Point St. Francisco	18		67 1
ł		10	51	60		Zice Sice	la n . a m n l			67
ı		11	29	60			TT - 3 # 100			67 1
1		11	5	60		Ĕ	Point Coama	17	55	66 5
15	Grenada, N. E. point .			61		۴	Point Coama C.MalaPasqua or SE.P. Shoal	17	59	65 4
1	- S. W. point		58	61		ı	Shoal	19	20	65 £
indenard lelands	Grenada Bank Middle			62		ļ	he			
12	Barbadoes, S. P.	13 13	1	59	36	ı	IT - 34 T	17		66 8 67 8
Š	E. do	13	8 5	59				18 18	9	67
	- N. W. point	13		59		ı	Zacheo or Dessecheo I.			67 \$
E	St. Vincents, N. point			61		ı			~-	
	-8. do	13	4	61		1	Cape Engano	18	35	68 5
ł	St. Lucia, S. point .	13		61	00	l		18		68 :
		13		1	56	ł	St. Catherine's I	18		68 :
1	Martinico, S. E. point .			1 -	56	ı	1	18		69 !
ı		14		61	6	Ę	La Catalina	18	8	70 1
ı		14		61	9 29	ş	Cape Beata	17 17	43	71 S
	1	14	14	I -	25	Ĩ	Altavela rock off do Cape Jacquemel	18	13	72
ı			39	61		2	Island Baca	18	4	73
1			52		37	Ľ		18		73 :
ł		16	4	61	14	6	Cape Tiberon	18		74 :
ı	- S. do	15	53	61		I§	Navaza Island	18		75
ı	Guadaloupe, S. W. P.	15	58		48			18	38	74 5
1	- N. W. do	16	20		56	B	Jeremy	18	38	74
f			30		32	17	Cavmito	18		73 4
Į		16			15	M		18		72 !
		17	21 5	61	8 44	1				78 : 72 :
1		17			00	i		18 18		72.1
I			42		17	l	- N. W. P.	18	56	73
	- N. P.		50		17	1	St. Mark	19	4	72
1	Redondo Island	16	56		22		St. Nichola Mole .	19	49	73 :
1	Nevis	17	9		33	1	Tortudas W. P	20	6	72 !
•	Ke Philippin and Bibbs	.1		1			led 🖳 l	jav.	٥	79

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			at.		ong.				at.		ong.
		D.	M.		M.		T - Colombia C DID		M.	D.	Èi.
		20			55W		Los Colorados, S. W.P.				44W
		19		71			Point Ivan and Yaunita	22		33	8
		19		71	10		Point Juan and Jaunito	22		84 83	
		19		69			Da. 11-1-	22		53	5
	O O		16	69	7		Port Cabanas	99		ŝż	
		iğ	3	68			MARIEL	23		82	
		-	-		•			23		82	
	Morant, E. P	17	58	76	9	1	HAVANNAH (the	l l	_		~ }
	KINGSTON	18	1	76	51	١.	Moro)	23	9	82	19
ġ	Port Royal	17	59	76	55	Ė	Point Escondido	23	8	81	47
Ē	Portland Point	17	42	77	14	Cuba.	Point Guanos	23		81	
Janaica	Pedro Bluffs	17		77		5	Pan of Matanzas			81	
٠,		18	.1	78	1			23		81	
			13	1 -	23	ğ	Point Ycacos	23	10	31	9
	Cape Negril, S. point . N. do.	10	24	78	37	orth	Stone Key off do Key Cruz del Padre .	23 23	13	81 80	.9
	Montego Bay	18	31	78		Ş	Las Cabezas	23	16	80	
,		18		77	49	ટ્ર		23		80	
	St. Ann's	18	31	77		ı	Key Carenero	22		79	
	Galina Point	18	29	76		I	Key Francis		40	1	17
	Arnatta Bay	18	21	76	51	ĺ	Key William (northern-			1	
	N. E. Point	18	13	76	20 ,	l	most)	22	36	78	
		l			·	ı	St. Juan	22	14	78	58
	Morant Keys or Las			L		•	Key Coco 9. side Baha-			L	
		17	25	76	00	•	ma channel		29	78	17
	Pedro Shoals	.~	00	~~			Key Point Paredon do.			79	5
	- Rattlesnake, N. W.	11	w	"	13			22 22		77 77	56 40
		17	5	79	13		Cayo or Key Verde		5	77	
	- South part	16	-		26	l	Guajava	21			25
	Formigas Shoal, N.E.P.				42		Point Maternillos	21	40	76	
			28	75		ı	Point de Mangle	21	13	76	
	Little Cayman, S. W. P.	19	36	30	5		Point de Mulas	21	7	75	34
			43	79	52		Tanamo	120	43	75	
	Grand Cayman, S.W.P.			81	5		Key Moa	20	44	74	
	— E. P	19	18	80			Point Guarico	20	40	74	
		17		84	4	i	Baracoa	30	22	74	25
		15 18	94	79 75	8	l	Shoal	19	e C	69	5
		1.3	~~	1'3	•	i	Nativity bank or E. reef		8 8	68	-
	Cape Mayze	20	14	74	4		Superb Shoal	20	-	68	
	C. Bueno or Guanos .	50	6	74		ļ.	Silver Key, S. E. end .			69	
	Point ent. Cumberland	ł		_			— N. E. do	20		69	
	Har	19	54	75	11		— W., do	20	29	69	59
	St. JAGO DE CUBA,	l. <u>.</u>			_	ı	Square Handkerchief,			l	
			57	76	5	ı	N. E. P	21			23
	Tarquin's Peak	13	54	76		l	S. E. P	50	-	1	28
	Dana dal anta	20 20	47 10	77 79	8	ŀ	- S. W. P.	20	53	1	56
Cuba.		21	6	79		ŧ	Turk's-Island, Grand'T Salt Key	21	90	71	3 58
3			44	80	5	ľ	- and Key	21	19		10
₽	Bay Xagua	21	53	80		ľ	- Endymion's Rocks .		7		15
side	Stone Keys	21	47	81	45		Great Cavcos, south	ŧ .	-	ļ -	
	Los Jardines	21	37	91	31	7	Part	21	31	71	27.
*	o. E. hour of the Drift	21	24	81	18	8	- N. E. P. or shoal				
Sout			94	31	50	Craycos	St. Philip	21			22
	Keys Jardines		24	92	4	ပ်	- N. W. part	21			47
	I. Pines. S. W. P.		22	82	-		Liver Cayoos, middle .				57
•	Indian Keys	21		82			Booby Recks off do	21	58	71	57
		21 21		83 83	49	ĺ	Providence Caycos, N. W. P.	21	50	72	9,
	Cape Corientes		43	84			Little Caycos, S. W. P.	91	36	72	
	Cape St. Antonio		54	34				21		72	7
	ancho Pedro Shoal .	23	4	85		ĺ		21		72	3
_	Shoal discovered in 1797.	22	6 -	85	2	L		21			430

-							•				
T	1 `		Let.		ong.	1			at.		ong,
ì	Great Inagua or Hene-	υ.	M.	D.	М.	l	Leeward Stocking I	D.	M. 50N	D.	M.
ł	aga, N. E. P	21	19N	73	1 W	1	THE HOLE IN THE	23	JU11	,,,	101
ı	- S. E. P	21	00	73	6	l	WALL	25		77	13
1			54		41	Jank.	Rocky point of Abaco .	56		77	5
ı		21	8	73		\mathbf{B}_{α}	Elbow Key	1	31 36	77	
1	Little Heneaga, E. P — W. P.	21		73	55 7	2	Man of War Key . Great Guano Key .		30 43	77	5 11
	Hogsties or Corrolaes .	21	39		00		I as Calanages N D	107	22	78	
ı		21	57	72		율	Lit. Bahama Bank, N.P.	27	50	79	
1		22	17 30		39	¥	Memory Kock	26 26	58 E.4	79 79	4
1			20 20	73 73	6	Little	Sand Key		34 46	79	1
1.	French Keys or 1. Pla-					1	Great Bahama I. W. P.			78	
ands.	nas .	22			34	ŀ	— S. P	26	21	78	
Ş	Miraporvos Keys .	22	7	74	32	Ŀ	- E. P	26		78	9
13	Castle Island or South	22	8	74	20	I		24 23	1 58	79 79	
ğ	Fortune Island, W. P.			74		ğ	Double-headed Shot	-	•	`	٠.
ä		22		74		Ž	Key, Western	23	52	90	14
Г	Crooked Island, W. P.				18	Salt	Salt Key	23	39	80	.8
1	- E. P	22	35	73	50	Ņ	Anguila, E. P	23	27	79.	14
1		23	5	73	35	ı	- GEORGETOWN,	32	22	64	33
ı	- W. P	23	3		49	å	- Wreck Hill, western-				
1	Rum Kev	23	34		57	ş	most-land	32	15	64	50
1	Watland's I. N. E. P.	24	6		26	Bet	Best Latitude to run for				
1	Conception, or Little I.	23 93			37 16	٣	Bermuda	32	8	'	
1	St. Salvador, or Guana-		-	١.٠	•	ı	III. East Coast of Amer	ica.	from	the	Gul
1	hari, S. P.	23	57	75	32	ı	of Mexico to Ca	pe Ť	forn.		•
1	— N. P.	24			49	l					
	Little St. Salvador, N.P. Eleuthera or Hetera I.	24	32	76	12	ı	1	D.	at. M	D.	ong. M.
ı	- Powers Point, S. P.	24	39	76	23	l	Point Gulebrao, E. part		IVE.	۳.	147.
1	- Point Palmeto	25	12		26	1	I. St. Louis	29	10N	96	5 V
	- James Point	25	24		36	1	Point St. Francisco, en-			l	
1			29		50	ı	trance of Bay St. Ber-		20	96	==
1	Egg Island, W. P New Providence	Zo	28	77	6	ı		28	53 8	97	
ı		25	5	77	22	ł	Point of the Coast .			97	
1	- E. P		59	77	9	ł	Bar de St. Jago	26	5	97	31
4	- W. P.		59	77	35	Ļ			55	97	
Benk	Andros Islands, S. P. — N. P.	24	4 24	78	45 3	Ę	River St. Fernando. ent. Inlets to Laguna Madre	25	22 2	97 97	
9	Berry Islands, Eastern	25			41	ş	Bar de la Marine, en-	23	•	7"	36
Ę	- Northern	25	49	78	1	2	Amana	23	45	97	58
88	- Great Harbour	25	49	78	5		Bar del Tordo	22		97	
18	Little Isaac, Eastern .	25 26		78	46	Coast	Mount Commandante .		48 20	97	
12	Berry Islands, Eastern — Northern — Great Harbour Little Isaac, Eastern Great Isaac Bemini Island, northern	120	1	79	2	ast	Bar de la Trinidad . Bar Ciega		39 34	97 97	
ľ	fresh water key	25	43	79	8	Eg	River Tampico		16	98	
1	Cat Key	25	23	79	10	ľ	Point de Xeres	21	55	97	45
1	Riding Rocks	25	17	79	4	1					
	Orange Keys, North — South	24 24	59 53	79 79	6	1		21		97	
1	Key Guinchos	22	44	78	1	1		21 26		97 97	
1	Key Lobos	22	25		33		River Cazones	20 4	44	97	
ı	Las Mucaras		10		12		Tenestequepe	20 4	40 .,	97	12
1	South edge of the Bank		5		22		Boca de Lima	20		97	7
1			45 56		45 19		River Tocoluta, ent Mount Gordo	20 20		9 7 96	1 57
1	Key Verde Island	22	1	75	5		River Nauta, ent.	20		96	
1	Key Sal	22	12	75	41		River Palina, ent.	20		96	
	Yuma or Long I. S. P.			74			Point Piedras	2 0 (DQ.	96	35
ı	Proma N. W. D.	23		75 75			River de Santa Nos!	19. (56	96	30
1	Exuma, N. W. P	23	<i>5</i> 0	75	DI.			οσ	10		

		at	Π.	ong.				at.	T.	ong.
	D.	M.	D.	M.	1			M.	D.	M.
t Delgada	1.0	52N				Bay Ascension, ent		26N	88	317
	liś		96		l	Island Cosumol, N. P.			86	
	19			21		- S. E. P		52	86	32
r St. John Angel .	19	32	96		ŀ	Rio Hondo, ent	19	4	88	
pa	19	32		50 .		I. Ubero, N. P.	19	20	98	3
de Orizaba	19	2	97	9		— S. P	13	22	87	53
t de Sampola .	19	30	96	16			19		87	52
St. Carlos	19	26	96	15		Key Jaicos	18	14	97	52
r Antigua	19	20		14	l	North Reef	18	2	87	50
Gorda	19	15	96	4		Chief Channel	17	54	87	55
A CRUZ	19	11	96	4	ı	Wallis's River, ent	17	52	88	19
ohn de Ulloa .	19	15	95	58		El Chinchorro I. N. P.	18	58	87	11
ana	119	4	96	6			18		87	6
r Medellin, ent	119	6	95	59		Misteriosa I	18	38	85	\$ 5
t Auton Lisardo .	119	4	95	45					84	44
le Alvarado	18	46	95	38			17		94	4
otalpan	18	35	95	29		South Keys, N. P.	17	30	87	12
	18	38	95	18			17	00	87	8
t Roca-Partida .	18	40		59	١.	Longcriffe, or Glover's	١.		١.	
t Morillos	18	41	94	51	ន់	Reef, S. P.	16		1 -	41
la *. •	18	18	95	5	Honduras.	Sapotillas Keys, S. E.P.	16	00		12
t Zapolitan	18	34	94	41	ŝ		16		86	
: Xicacal	18	27	94	37	윤		16			57
t St. John	18	19	94		Γ	Guanaja or Bonacca I.	16	32	86	7
illa	19	7	94	27		Point Manabique .	15	39	98	29
Guazacoalios .	18	8	94		l	Omoa	15	37	1	57
r Tonelado	18	8	93		1	Point Sal	15	47	37	29
rSt. Ann	18	8	93		ı	Triunfo de la Cruz	15	41	137	17
Cupilco	18	13	93	8	ı		16		87	2
Bocas	18	13	92		ı		15	53	86	6
r Chittepeque .	18	14	92		ı	Cape Delegado, or Hon-				
	18		98	7	l	duras	16			11
r St. Peter & Paul	18	27	91		ı	Cape Cameron	16	8		10
Llicalango	18	44	91			Cape False	15	14	93	3
	18		91		١.		14		83	
Escondido	19	-50	90			Caxones, W. P	16	3		11
r Chen	19	20	90				15			27
Morros	19	40	90				15		83	
PECHE	19	50	90		١.	Key John Thomas	15	23		49
Desconocida .	20	55	90			Alagarte Alla, N. W. P.			83	5
t Gorda	21	6	90				15	5 .	81	54
l Piedras	21	9	P:	13	١.		16	5	80	.2
	21	2U	89		1	Serrana or Pearl I. N.P.				47
lara	21	32	88	-			14		79	51
s de Silan	21	26	88		ŀ	Guana Recu, N. P.	14	47	1	44
	21	30	97		١.	S. P	13 13	27 20	B 0	41
d Jolvas, N. P.	21	30		11.	quitos.	Roncador	13	37		46
d Contoy, N. P.	ZI		96		Į.	Musketeers	13	21 07	29	46
Arcas Islands .	20	1b	91		18	Providence I. N. P	13	## #0	90	39
Obispo	20	32	92	5	Mos	Musquito Keys, N. P.	1.7	77	82	19
ngles Islands .	20	59	92	7	ľ	Ned Thomas' Keys,	۱, ۵	12	60	a, 1
Shoal	20	33	91		ı	S. P				21
	21		91			Bracman's Bluff	13		82 82	
	22	7	91	ZO	ł	Man of War Keys .	13	4	82	36
		96	۱.	01	ı		12		32	11
		36	91	21	•		13 11			
Sisal		27	90	2				33	83 81	
ran		29	צסן	26	I					
art of Bank off this	1	45		46	ĺ	E. S. E. Keys	-	22	80	31
	23			43	ı	S. S. W. Key, or Al-		6	61	8
. do	23	27	96	37	l		12	-6 -90	81	
Mugeres, or Wo-	1		0-	40	ĺ			20 41	32 32	3 4
ın's l		18		49	Ę	St. John's Point		41 90	53 53	7 5
inkun, S. P.		42		58	ĮŞ	Port Boco Toro		29 14		57
River	X)	26		15	16	I. Escudo, N. P.		20	80	3
Lacales	20	5	67	34	ď,	River Chagre, ent	1 7	-Thous	ĮΟV	است

-	18 19 1	Lat.	Long.		All and	Lat.	Long.
M	DODTO PELLO	D. M.	D. M. 79 35W	1	New Barcelona	D. M.	D. M.
	PORTO BELLO .	9 40	79 33 W	Cumana.		10 8N	64 46V
	Farallen I. N. P.	1 -		ma		10 20	64 48
	Point Manzanillo .	9 38	79 20	3		10 16	64 31
2	Point St. Blas	9 33	78 40	~	The state of the s	10 27	64 15
Darte	Point Conception .	9 19	77 53		484 36 7 36	10 35	64 20
4	Isle of Pines	8 55	77 39		the state of the s	10 42	63 54
	Cape Tiburon	8 40	77 29	М	Escondido or Hidden	10 41	CO 00
	River Suniquilla, ent	7 57	76 54	12	Port	10 41	63 27
1		8 37	76 57	П	Cape Malapasqua .	10 42	63 4
4	Point Arboletes	8 49	76 32	ш		10 46	62 44
2	Island Fuerte	9 20	76 13	ш		10 45	62 33
Carthag	I. St. Bernerd, N. W.P.	9 48	75 50	П	Point Pena or Salina .		61 53
	CARTHAGENA .	10 25	75.29	U		10 41	61 48
٥	Galera de Samba	10 48	75 20	н	River Gaurapiche, ent.	10 12	62 43
	West ent. River Mag-				Point Morro	9 54	61 58
	dalen	11 3	74 56	П	Oronoco River	8 23	60 26
à	St. Martha	11 15	74 11	L	Cape Barma	8 22	60 4
	Cape Aguia	11 21	74 12	1	Essequebo River	7 00	58 20
	Bank Navio quebrado .	11 36	73 11	П	DEMERARA river, ent.		1000
	Hacha	11 31	72 56	1	Corrobano Point	6 48	57 58
	Cape la vela	12 11	72 14	1.	River Berbice, Ent.	6 20	57 11
	Point Galinas .	. 12 27	71 41	Surinam	SURINAM River, ent.		55 15
00	Monges Islands, N. P.	. 12 31	70 59	1.5	Paramaribo .	5 49	55 15
83	Cape Chichibacoa	. 12 17	71 17	12	R. Marouri, entrance	5 50	53 52
JE .	Point Espada .	. 12 5	71 8	ľ	CAYENNE .	4 56	52 15
8	St. Carlos	. 11 3	71 12	1	Oyapock River, St Louis	3 51	51 40
S	MARACAYBO .	.10 43	71 17	ъ	Cape Orange .	4 12	51 20
	Coro	. 11 24	69 46	п	R. Cassipour, entrance	3 54	51 10
	D C	. 11 35	70 20	т	Cape North .	1 48	50 10
	Point Cardon . Point Macolla . Cape St. Roman	. 12 6	70 19	п	Mouth of River Amazon	0 18	50 00
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	Island Oruba, N. W. P	12 38	70 9	1	Point Tagiora	0 33	47 28
		. 12 25	69 58	L	Para	1 28	47 58
		. 11 57	69 53	1	Cape Magoany Point Tagioca Para Bay Maracuno	0 37	47 10
	Point Savannos .	.11 33	69 10	г	Caite Harb. Cape Gurapi Shoal	0 47	46 33
		11 14	68 35	п	Cone Gurani	0 42	45 22
	Point Soldado .	10 57	68 19	1	Shoal .	0 52	43 40
	Key Borrocho .	10 51	68 17	п	- marie		
	Tucacas	10 29		Т	Island of St. Joao	. 1 17	44 13
		10 19	68 4	Т	Bay of Mt. Luis	1 5	43 18
	Valencia	. 10 18	68 7	1	Bay de Cabalo de Velh		43 54
28.	Point St. John Andre	\$10 30	67 48	н	Point of B. Atius	. 2 3	43 44
20	Point Oricaro .	. 10 34	67 17	Т	Itaculumi	. 2 7	43 50
D.A.	Point Trinchera .	.10 38	67 4	h	OF THEIR COS .	. 2 27	43 40
100	LA GUIRA .	. 10 37	66 59	П	va. ue Macantin		43 47
-	CARRACCAS .	. 10 30	66 57	1	St. Luis de Maranhan		43 40
13	Centinela I. or Whit	e	Le 2	1	Coroa Grande, or Grea		Lan.
	Rock	. 10 50	66 6		Crown Banks, N. E		Jan 2
	Curacoa I. N. P. S. E. P.	. 10 36	66 3	1	Point	. 2 12	43 18
	Curacoa I. N. P.	. 12 24	69 13		Fin dosLancoesGrande		42 40
	- S. E. P.	. 12 2	68 46	1	I. St. Anna	. 2 18	43 5
H	- S. E. P. Little Curaco Buenayre, N. P S. P. Birds or Aves I. wester	.11 59	68 41	1	Bay of Rio Perguicas	. 2 23	42 4
	Buenayre, N. P.	. 12 21	68 26		Iquarasu ent. Parnhaib		41 20
1	- S. P	.12 2	68 18		Jericoacoara .	. 2 44	40.15
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N	- Eastern .	. 11 58	67 29	1	Mount Melancias Poir	t 3 7	39 7
9	Roca. W. P.	. 11 51	66 58	1	Searra	. 3 32	38 27
1	- E. P	. 11 51	66 32	1	Bay Iguape	. 3 40	38 14
	Orchilla I.	. 11 49	66 5		Roccas (dangerous)	. 3 59	33 26
1	Blanca I.	. 11 52	64 40	1	St. Lorenzo .	3 57	37 59
	Tortuga I.	. 10 57	65 19	1	Point Daniel .	4 42	37 24
	Seven Brothers mid.	. 11 46	64 27	1	Baxos de Salino .	4 40	37 00
ı	Margarita, W. P.	.11 2	64 28		Dollar De Lane	4 52	36 38
Į.	E. P.	.11 00	63 50		Cape St. Roque	5 8	35 38
		. 10 49	64 14	-	River Parahiba, ent.	6 48	35 10
	I. Cuagua or Pearl I.		63 48		I. Tamarica	7 46	
1	Friars I.	.11 14					34 5
-	I. Sola	.11 20	63 38		Pernambuco .	0829	35 5
1	Testigos I.	.11 24			Cape St. Augustine		34 51
	River Orquilla ent.	,10 8	65 32		Rio St. Francisco	. 10 57	36 4

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s Is. St. Ann's Bay	22	35	42	5		- False Cape Horn .		42	68	8
	22	44	41 42	50 6		— Yorkininster	55 54	27 7	70	4 35
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atherine's Island			48		1		41	49	C	53
St. Pedro	32	Š	52	8	ł	P. Quedal	41	5	74	9
St. Mary	34	39	53	58	l	P. de la Galera	39	54	73	46
	35	2		42	i	VALDIVIA, entrance .		51	73	33
	34		54	50	<u>.4</u>	P. Tirua	38	29 20	73	46
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gellan's Straits .		24	68	25	I	Point Tames	22	33	70	10
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uth point of en-		•			ı	Pavellon de Pica .	20	58	70	16
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del FuegoC.Penas	53	45	67 65	29 5			19 18	26 27		19 19
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rn	54	48		42	ž			17	73	
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ibelas Islands, E.P.				46	ĺ	P. Chilca	12		76	
E HORN, South	l		Ì			I. St. Lorenzo, W. P.	12	5	77	8
t of Hermit's Isl.		59	67	21	ĺ	LIMA	12	3	76	55
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1	Los Hormigas Islands .	11	56S.	77	48W	ł	Agualco	16	2N.	96	52
1	I Pelada	111	27	77	41	ł		16		100	
1	Island St. Martin .	111	3	77	30	ı				105	
1	Point Sentander	110	39		41	ŀ		21		104	
ź	Rock seen in 1792	110	48		48 ·	1	Tres Marias.	21		106	
1,5	Ferrol (entrance) .	19	7		30	ı	St. Joseph	23	4	109	42
r	l ` .	18		78	63	l	Cape St. Lucas	22	44	109	
•	I. Malabrigo (port)	7	48		21	ı	Morro Hermosa	27	46	114	41
1	isi. Lobos de Mer .		58		44	1	Redondo Island	29	49		
1	Isl. Lobos de Tierre	6	24		46	ł		. 30	23	115	
ş		6	56	79	49			31		116	- :
ł	Point de Ajuga	5	69	91	4	ł	Port Diego	. 32	89	116	
1	Point Payta	5	· 3	81	2	ł	Point Conception .	34	32	120	6
1	Cape Blanco		19	81	6	I	Monterey	36	38	121	
1	P. Malpelo	3	32		17	•	Port St. Francisco	37	49	122	
1	GUAYAQUIL City	. 2	12	1 -	42	1	Cape Mendocino .	40		124	7
1		3	4	130	8	ı		. 41		123	
1	Point St. Helena .	2	10	80	48	ł.	Cape Blanco or Orford	. 48	53	124	_
f	L Pelado		56	80	36	â	Cape Gregory	43	36	124	6
1	Point del Callo		23		34	2	Cape Foulweather	44	52	124	
1	I. de la Plata, W. P.	1	18	4. "	57		Cape Rond	45	43	123	
ł	Cape St. Lorenzo	1	4		43		Cape Disappointment .	46	13	123	
13	Manta		57		32	ο,	Cape Flattery	48	24	124	
13	Manta Cape Pasado Quito		27		20	3	Breaker's Point		24	126	
þõ	Quito	0	18		18	ಽ	1 '	49	36	126	
ł	Arbol		15N			'n	Woody Point	50	6	127	
1	Cape St. Francisco .		39		52	2	Bay St. Louis	50			14
1	P. de la Galera		48	79	51	7	Isles de Sartine (orBcott)	50	20	128	
1	R. Esmeraldas entrance	9 0	58		23	5	Cape Scott	100	48	128	
1	P. Mangles		37	P	52	څ	Cape Caution	51	120	127	
ŧ	I. Tumaco		47		38	١,	Cape Hector or James .	101	40	131	7
1	P. Guascama		29		23	ı	Bay de la Touche	52 52	42	132 132	
1	II. Gorgona middle .		53	78	7	ŀ		53	to.	133	
1	River Cajambrie, ent	1	19	77	3	ľ	Bay de Clouard	54	90	133	
1	I. de Malpelo		55	80	4	ı	Core St Portsland	55	19	133	
1	I. de Palmas		57	77	7	ŀ		56	19	134	
1	P. Chirambira		13 34	77	15	ŀ	Cape Ommaney	56	38	135	
1,	Cape Corientes	6	3	77	1				43	135	4
15	Limones		49	77		ľ	C. Engano or Edgecumb		2	135	
Iŝ	P. St. Francisco Solano	8	8		12	l	Port Gaudaloupe	57		135	
P.	P. Garachine PANAMA		57		22	ı	Port de los Remedios .	57	24	185	
1	P. Mala	7			53		Cape Cross	57	57	136	24
1	m m			80		ľ	Port des Français		37	137	
1	I. Quibo, N. P.	1 .	41	31		ļ	Cape Fairweather .		55	137	
1	Los Ladrones	7		82		Ī	Religions's Roy	59	18	139	00
ı	Point Burica	l s	3	82		ŀ	Point de la Boussole	59	50	140	5 5
1	Guife Duice, W. P.		23	83		l	Mount St. Elias	60	23	140	45
1	I. Cano ent. off English	1		-		ŀ	Cape Hinchingbroke .	60			16
1	harbour	8	48	83	51		Cape Elizabeth	59		151	
1	Cape Herradura	9	37	84		ĺ	Barren Isles			151	
1	Cape Blanco			84		ŀ	Point Banks	58	41	159	_6
ł	Nicoya		42	84		l	Cape Douglass	58	56	152	
1	Morro Hermoso	-		85	5	ı	Cane Whitemaday	158	15	151	46
1	P. St. Catharine	-	28	85		Ī	Cape Grenville	157	33	152	
ı		11	22	95		Ī	Trinity Islands	.156	36	153	
1	Point Desolado	11		86		I	Foggy Island	56	10	156	40
ı	Leon	12	22	86	46		Halibut Head Island .	54		162	
1	Realejo	12	27	87	5	ı	Ounalashka Island, N.P.	53	55	166	
1	Aserradores	12	35	97	20	Ĭ.	Bristol R. entrance	58	12	157	
1	Point Cosignina	12	53	87	37	þά		58	37	159	
1	Point Candadillo	13	7		57	lî.	Cape Newnham	58	34	161	
ı	Sacatecoluca	13			47	ĿĔ	Shoalness		00	161	
1	Point Remedios .	13			42	16	KLADE STEDDORS •		33	162 161	
ı	P. Gautimala .		54		53	Ø	1 T	104	317	166	
L	Puerto Ventosa	116	6	95		_	jours its day	64	34	100	-
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II TAB.

TABLE XLVI. Latitudes and Longitudes.

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dulgrave .67 46 165 12 APE .70 99 161 42 om the River St. Croix Cape Cape Egament .46 23 54 46 37 64 10 ce of St. Oroix Lat. Long. M. D. D. M. Lat. Long. M. D. <td>wines of Wales</td> <td>65</td> <td>45N.</td> <td>168</td> <td>17W</td> <td>Ŀ</td> <td>Richibucto Harbour</td> <td></td> <td></td> <td></td> <td></td>	wines of Wales	65	45N.	168	17W	Ŀ	Richibucto Harbour				
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Hope	POOT	144				1				63	54
tte Bay	Hone .	43	53			ı				61	58
tte Bay		44	13			١.	Entry Island	47	15		
## A series of the plane of the		-1-		63	53	1					
AX Harbour	rero light-house	44	30	63	32	1	Magdalen Isl. N. E. P.	47	41		
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ı	Mortier Harbour.	47	10	55	3		Rad Bay	151	46	56	
ı	Red Island, S. P.	47	94 .	54	8		York Point .	51	59	55	
1	Virgin Rocks	47	11	54	3	Ì	Cape Charles	. 59	13	55	
I	Point Brehin	47			12	l	Great Bay of Eskimau			57	
ł	Cape St. Mary	46			00	l	Cape Harison .	. 54		56	
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ŀ	Cape St. John				30	١.	Cape Walsingham	. 62	39	77	48
ł	Horse Islands			55			Cape Digges .	. 68	41	78	50
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15	II Angtoff lights	52	29		46	•		D. M.	D. M.
K		52			44		Pentland Skerries .	58 43N	3 3W
13	Winterten Nass II.	52 52				ľ			1 .
P	Consister Vacant				43			58 43	3 14
P	SOUTH & THOM	. 52			20	1	South Ronaldsha, S. P.	DO 44	3 5
	massorough Sand, S. P.	52	51	1	48			58 54	2 47
13	N. P.	53	9	1	35		Lam's Head on Stromsa	ł	1
ĸ	Sherringham Shoele	53	3		20		Island	59 4	2 40
12	Hasborough lights	.152	49		34	٠,	North Donaldaka M D		2 31
₹.	Cromer lights	152	56		26	Ą	Mould Head, on Papa Westra Island	1	1
r	Lemon and Ower, N. P.	52	14	_		Į S	Westra Island	5 9 2 1	2 0
I	S. P.	23			58	. 45			1 "
		53			00	reg	Noup Head, on Westra		
	Cromer Bank	53		1				5 9 18	3 14
	Dudgeon light	53		1	9	6	Marwick Head, on Po-	i	
		53	33	1	14		mona Island	59 6	3 28
	Inner Dowsing	43			44	H	Stromness	58 57	3 26
	Lynn Knock		3		29		How Head, on How		
	Spurn lights	53			20	1	Wells Island	58 55	3 31
	Plamborough Head					1	Sine Skerry	59 3	
		54	7	Ó	6		- New Process	PZ _P _	4 16
1		24	42 '	_			Pain Lefandunian Left	ے مم (میرا	4 477
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	XVII. The Shell	ınd İslan	is.	1		D. M.	D. M.	
	<u> </u>		 .	4	Solway Firth	54 47N.		
		Lat.	Long.	ł	CARLISLE	54 54	2 44	
		D. M.	D. M		Ga Books IVand Viela	54 30	3 30	
•	Sunbro Head, S. point	107 47IN	1 200	4		54 32	3 22	
3		.60 13	1 00	ı	Selker Rock	54 16	3 19	
Ì	Brassa Sound, Lerwick Out Skerries	: 60 37	0 8	1	Lancaster	54 2	2 41	
WHENDER	Whalsey Isle .	60 32	0 32	1	Formby Point	153 32	2 58	
ż	Ulst Island, N. E. P.	.61 7	0 15	ı	LIVERPOOL	53 22	2 52	
		.60 8	2 16	ı	Point of Air Light	53 20	3 11	
		.,,	1	-	Great Orms Head	.153 20	3 43	
	XVIII. Ferro	Islands .		1	Point Lines Light	153 24	4 11	
				:13	Skerries Light .	53 24	4 30	
		Lat	Lon	5 \$	Holyhead, W. P.	53 18	4 32	
	71 - 36 - 1 D - 1	D. M.	D. N	4 8	Branchy Pool Head	5% 47	4 37	
;	The Monk Rock appear	ISI 15N	6 478	ĸ	Bardsey Island	58 44	4 39	
	like a ship under sail	u lot 1914	0 4/1	٦.	Barmouth	52 43	3 52	
2 2	Fucloe I. (N. E. part of Ferro)	62 15	6 2	15	Aberiswith	59 23	3 53	
4	East point of Mygene		"	10	1614	52 6	4 38 5 10	
	Island, N. W. part o	2	1	13		52 1	5 20	
	Ferro Islands	.62 3	7 32	12	CON SOUTH OF STORES	51 55	5 23	
	1-Ct10 Islands .			٠ľ		51 52 51 45	5 36	
	XIX. From Duncan	sbay Hea	d to th	4	St. Ann's ditto, Milford		1000	
	Land's E	nd.		1	Haven	51 41	5 10	
	<u></u>	,		-1	St. Gowan's Head	51 38	4 58	
		Let.	Long.	1	Caldy Island	51 40	4 40	
		D. M.	D. M.	1	Worm's Head	51 34	4 17	
		58 40N		V	Mumble's Point & light		3 57	
		58 43	3 29	ı	Nash Point	151 26	3 33	
		58 38	5 00		BRISTOL	61 27	2 35	
	Cape Wrath, or Barre	1	۱		Flatholm light	51 23	3 6	
	Head .	58 36	5 20	ł	Lundy Island, entrance) [
1	A Rock seen at Ebb .	58 45	5 21	1	of Bristol Channel	51 10	4 38	
1	Rona Island	59 55 . 57 42	6 15	k	Mort P. entrance of		1 . [
	Rockal St. Kilda	57 51	14 6	ł	Bristol Channel .	51 11	4 19	
i	Butt of the Lewis .	58 29	8-56 6-34	ı	Hartland Point	51 1	4 30	
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;		58 13	7 30	ı	(Cow and Calf	150 33	5 5	
1		57 23	7 38	ļ	Towan Head	50 25	5 9	
1		57 8	7 10	ı	St. Ive's Bay	50 13	5 28	
i		56 48	7 33	I	Cape Cornwall	50 8	5 41	
1		57 50	5 44	ł	The Seven Stones .	50 2	6 6	
- 1		57 3	6 38	l	The Wolf Rock	49 63	5 51	
		56 56	6 38	ł	The Land's End	50 4	5 42	
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1		56 30	6 52	1	An. Fee			
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1	Skerryvore	56 17	7 8			D. M.	D. M.	
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ŀ		55 39	6 10	i,	Fastnet Rock	51 19	9 44	
١	Mull of Cantire light			L	Crook Haven	51 26	9 52	
١	house	55 20	5 37	Iŝ	Mizen Head	51 25	10 8	
	I. of Arran, S. E. part		4 57	13		51 33	10 5	
١	Cumray I. entrance of			۴	Bantry Bay		10 10	
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		55 59	4 16	Š	Dursey L. W. P Bull Rock	51 37	10 36	
		55 20	4 55	ķ	Bull Kock	51 38	10 49	
		55 39	4 30	Ę	Cow ditto		10 39	
		55 26	4 28		Cod's Head		10 97	
ľ	Loch Ryan	55 3	5 00	F	Kenmare Bay	51 44	10 30	
I		54 48	4 58	1	Lamb's Head		10 28	
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and.	Foze Rock	15%		11	6	-:	WICKLOW lights .	52 5	9	-	1 .
ij	Ferriter's Island			11	1	Γ.	Arklow	52 4	9	6	6
Ę	Tiraght Rocks	528		11	4	Ŧ	Glasscarrick	52 3	6	6 1	1
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Ę	Ennis Tuscan	52		11		l	WEXFORD	52.2	1	6 2	8
ğ	Ennis Tuscan Dunmore Head Dunorling Head Brandon Head	59	12 ·	10	53 .		Tusker Rock	52 1	8	6	8
ŭ	Dunorling Head	52	19	10	49	ŀ		52 1		6 1	
ij	Brandon Head	52	23	10	36	l	The Saltees Rocks		6	6 3	
10	The Seven Hogs Rocks	52	26	10		ł	Hook light, Waterford		_		•
r	Kerry Head, S. entrance	1				ı	harbour		5	6 5	6
ł	of Shannon River .	152	30	10	24	ų	Dungarven		3	7 3	
1	Loop Head, N. entrance		••		~		Helwick Head	52 0	Š	7 3	
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	Slime Head	53	35	10	32	Ľξ	Dundedy Head	151 3	2	8 5	i6
ı	Ennis Shark I.	53	46	10	36	8	The Stags, off Toe head	51 2	7	9 1	16
1	iEnnis Turk I	53	52	10	21	ŀ	BALTIMORE harbour	51 2	7	9 2	:6
ł	Clare Island	53	58	10	14	ľ		,			
1	Achil Head	54	7	10	30	ŀ	XXI. The Isla	of M	lant.		
ł	Black Rock	54	13	10		ŀ	2321 2 70 200	V			
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ı	Ciris rican	-	~	10	10	ľ	•			Loi	
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ł	Down Patrick Head .		27		36	76	Point of Air	54 2		4 1	
			19		27	9	Peel Hill	54 1	2	4 3	7
l-i	Sligo Bay	54	24		59	ľ	Castletown	54	3	4 3	5
Ħ	Ennis Murray Island .	54	32	8	57			<u> </u>			
Ę	Donnegal Bay	54	38	8	50	ŀ	XXII. From Cala	is to t	he S	caw.	
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1	Giant's Causeway .		16	6	26	ı	- Middleburgh	51 3	0	3 3	7
1	Racklin I. W. point	55	20		16	1	Goeree Island	51 4	6	3 5	
1	Fair Head	55	15	6	6		Schowen Island light .	51 3	9	3 4	
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Æ	Carrickfergus		43	5	46		Haerlem	52 2		4 3	
5	BELFAST		35	5	57	ı	ROTTERDAM	51 5		4 2	
	Belfast Loch		43	5	35	Ī	AMSTERDAM	52 2		4 6	
Coast	Mew Island lights .	54	41	5	24	1	Alkmaer	59 3	9	4 3	
ř	South Rock light	54	21	5	24	1	Texel, S. point	53	2	4 3	3
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TABLE XLVI. Latitudes and Longitudes.

D. M. D. M. Cast of 16N 10 -00E AMBURGH 53 38 9 56 ade 53 58 9 28 inhaven 53 59 9 28 inhaven 53 59 8 38 6				المعاقب فيان بالمعادر المراجع والمراجع والمراجع	برجم ومرخم خوب
D. M. D. M. D. M. AMBURGH 53 32N 8 32E. AMBURGH 53 33 9 9 56 ade 53 36 9 82 akstadt 53 48 9 28 krhaven 53 56 8 39 be River, entrance 54 00 8 20 anningen 57 15 N 10 30E, homout 57 25 9 34 be River, entrance 54 00 8 20 anningen 57 25 9 34 be River, entrance 57 25 9 34 Amburgh 57 43 10 37 XXIII. Cattagat and Seand. Lat. Long. D. M. D.		Lat.	Long.	1	Lat. Long.
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こ	— Bórgholms Slott . — S. end light	56	52	16	37		Askwold	61	22	5	12
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l l		55			47	B		12.	10	Į	
ı k	S. E end	54	58	15	14	R	Wardhuus Island . River Kola . Nagel Island .	. 70	22]31	6
. ∤.	— Svanike	55	8	15	16	14	River Kola	.69	15	33	94
ı ı		54		13		r	Nagel Island .	.68			55
1						•					
ŀ	— BERGEN	54	25	13	32	l	Cape Sweetnose .	67		37	
, l	- S. E. end New	1		1		l	Cape Orlogenose	. 67	00	39	21
: 1		54	16	12	52	ı	Cross Island .	. 66		38	49
(. j	Deep	104	10	113	JZ	1					
, 1·						1	Onega	-1	36	37	
i I	XXV. Gulfs of Finlan	nd o	nd B	oth:	is.	14	Cape Donega .	164	45	35	46
į t						ĸ	Cape Donega . ARCHANGEL .	. 64		38	59
				4							
. 1			Lat.	L	ong.	闰		65			10
. 1		D.	M.	D.	M.	E	Cape Good Fortune	. 66	24	40	
, I.	Odensholm light				18E.	厗	Morshom Island .	66	40	40	35
					TOD.	1	Cana Candinose	. 68		41	
		59		24	3	ĺ	Cape Camuniose .			12-	
	Surep Head light .	59	28	24	23	ı	Welgate's Straits .	. 70		62	2
	Nargen I. N. point	59	36		34	•	Nova Zembia .	.[78	00	170	00
	prvri	59	26			•		•			
					<i>4</i> 8 .	•	XXVIII. From Usha		On.	14	
l l		59		25	6	•	AATIII. ETUM USING	***	GW	COST	w.
1	Chalk Ground	. 59	41	26	15			_			
		59		26		i	1		Lat.	11	ong
		122	40		- 1	ı	1	D.	M.	ÌΩ	. M
	Little Titters Island .	59	47	27	3	l	USHANT	1	29N		3W
5	Great Titters Island .	59	50	27	20		1	1			
	Lavanscar, N. end	59	50	27			BREST	48	23	14	29
						l	Saint Matthew's light	48	19	14	47
		59		28			Point Raz	48	ĩ		47
_ 11	Narva	59	20	28	24			120			
1		59		29	6	ŀ	Saints Rocks	48	4	5	5
							Point L'Abbe	47	49	14	12
		60		29			Quimper	47	58	14	8
		59		29			Isles de Glenan	47	44	14	00
, I	PETERSBURG	159	56	30	19		isies de Gienan .	1.	77		
		60	8	29		ě	Quimperlay		5 3	13	34
, [cojis cuucii , ,						L'ORIENT .	47	45	13	21
. 1		60		29			Port Louis	47	42	1 -	21
. 1	Fredericksham	60	30	27	25		TI VICE LIVES	F.			
	Aspo	60	14		22	15	Isle de Gross	47	36	3	
	Uppland Island Habes	150	- 4			Š	Quiberon, S. point	. 47	26. 1	13	4
	Hogland Island lights .		3	27	7	10	Belle Isle, N. end	47	23	3	14
, þ	Orrengrund's Beacon .	60	14	126	39	ž	0 1	140		1,	
		60			28		S. end	47	17	13	4
						×	Vannes	. [47	39 ^	-12	46
		60			50	ı	Houat Isle	47		12	56
	Helsinfors	60	10	25	7	1		47	22	12	36
i I	Hango Beacon	59	45	22	57	I	Dumet Isle	۲.	4.5		
įŤ		,-,-		<u> </u>		ı	NANTES	47	13	1	3 3
	WWI Cult	í D	· • • • • • • • • • • • • • • • • • • •			ı	Croisic	147	18	72	31
()	XXVI. Gulf of	ъ	umia	•		ı	St. Gildas point .	47			16
ı l		-				i					
1			Lat.	L	ong.	1	Noirmoustier I. S. W.	47			15
1			M.	n	M.	1	Dieu Isle	46	42	13	27
j _ k	lia linka					ı	St. Gillies	46	40		51
.₫ (Uto light	لاترا	7/N		25E.	1	or Guises	146	15		24
囊	Abo	60	27	38	15	ı	Roche Bons	46	10		
5 l	Wasa	63	13		55	ı	Rec Isle Light	46	15	11	34
Both	TORNEA	120	E 1	5		ı	ROCHELLE .	46	9		9
71	LUMBA	65	91	24	9	i			5 6		59
. 1		_				ľ	ROCHEFORT .				
, į	XXVII. From the No	ize i	lo An	cha	ngel.	İ	Oleron Isle light	46	- 5	11	94
, I.						ı	leland Aix	46	1	11	10
ľ		, 1	Lat.	I	AAC	ı			3ŝ		10
					ong.	Ī	CORDOVAN light H.				
١,		D.			M.	ı	Medoc	45		10	45
١		IPO	1N.	17	14E.	ĺ	BORDEAUX	144	50	10	34
ż	The NAZE	158				•					
	The NAZE						Cana Panet	laa.	41	MT.	174
S	Lister Land	58	10	6	38		Cape Feret . Digitized	44	43		14
S		58	10	6			Cape Feret Digitized & BAYONNE	44	43) (99		14 99

-		, Lat.	Long.		Let.	1100
		D. M.	D. M.		D. M.	Long.
1	St. John de Luz	.43 23N		Bay of Roses		D. M.
1	0. 0.1	.43 21	1 57	10 O	43 15N. 43 20	
1	O	.43 28	2 40			3 17
1	Cape Machicao .		2 44		49 49	8 54
1	BILBOA	.43 15			43 19	3 28
Ŀ	SAINT ANDERO	. 43 28	3 40		43 15	3 29
Spate	Saint Vincent .	.43 30	4 16		43 24	3 42
Į,	Villa Viciosa .	. 43 28	5 18	Narbonne	43 11	3 00
	Cape Penas .	. 43 42	5 46	Montpeller	43 37	3 52
9	Cape Burola .	. 43 42	7 17	MARSEILLES	43 18	5 22
18	Cape Vanas . Cape Ortegal .	43 47	7 35	h⇒lCiotta	43 10	5 36
K	Cape Ortegal .	. 43 47	7 48	TOULON	43 7	5 55
ls.	Cape Prior	. 43 34	8 14	TOULON Cape Taillar St. Tropez	43 7	6 44
Lξ	FERROL	. 43 30	8 6	St. Tropez	43 16	6 40
13	CORUNNA .	. 43 23	8 20	Frejus	43 96	6 44
1	Cape Belem .	. 43 10	9 10		43 99	7 00
1	Cape Turiana .	. 43 3	9 17	Antibes	43 35	7 7
ł	Cape Finistere .	. 42 54	9 17	St. Marguerite Island	43 31	7 3
1	View Bow	.42 14	8 37	Cape de Oropes .	43 28	7 10
	Cone Forelle	.41 59	8 45			
1	lo nonmo	. 121 07	8 39	I	43 49	7 17
1		.41 11	1 :		43 40	7 20
ì	Averies	. 40 39	8 41		43 56	8 8
H	Coimbre	. 40 14	8 24		44 17	8 30
19	Cape Mondego . Cape Fiseraon . The Burlings .	.40 12	8 54		44 25	8 56
Ιŧ	Cape Fiseraon .	. 39 24	9 18	Rapallo	44 23	9 16
P.	The Burlings .	. 39 25	9 31	Point Venere	44 3	9 45
1	The Rock of Lisbon	. 38 46	9 29		43 33	10 22
1	LISBON	. 38 42	9 9	l 1= = = :	43 33	10 17
1	Cape Epichel .	. 38 25			43 25	10 20
	St. Ubes	. 38 32	8 50		43 20	10 41
1	Sines .	. 37 50	8 54	Piombino	42 56	
١.	Care St Vincent		0 04	Being Hannels		10 35
ŀ₿	Cape St. Vancent	. 37 3	9 2		42 23	11 10
Ĭğ	Lagos	.37 8	8 39			11 46
1	Cape St. Vincent Lagos Cape Carbonera Cape St. Mary Point Arenilla	. 37 7	8 19			12 28
15	Cape St. Mary	. 36 57	7 52	Cape Dazia	41 26	12 40
13	Point Arenilla . St. Lucar .	. 37 8	6 50	Cercello Point		13 00
ಟಿ	St. Lucar · .	. 36 46	6 20		41 15	13 38
Je.	SEVILLE	. 36 59	5 58	NAPLES	40 51	14 11
13	CADIZ	. 36 32	6 18		40 42	14 46
×	Cape Trafalgar .	. 36 10	6 00		40 18	14 57
Ł	Tarifa Island .	. 36 1	5 35		40 7	15 43
	P. Carnero	. 36 5	5 23		39 00	16 42
1	Algesiras	. 36 7	5 24		58 47	
1	GIBRALTAR .	.36 6	5 20	S Cape Bancan		16 25
ı	GIBRALI AR .	.130 0	(0 %)		38 20	16 10
1	WWW (7) 10 at 4		34	Se prime	37 56	16 15
1	XXIX. The North	oast of the	Ja editer-	S. Point of Italy	37 53	
ı	ranean Sea, from (sibrallar lo	Constan-		37 57	16 23
1	tinople.			S cape stillo	38 26	16 55
1				Catanzaro	39 1	16 55
1	1	Lat.	Long.	Cape Rizuta	39 54	17 31
		D. M.	D. M.	Cape Callone light .	39 4	17 36
	GIBRALTAR .	. 36 6N	. 5 20W		40 28	17 35
1	MALAGA .	. 36 43	4 24		40 00	18 20
1	Modril	. 36 45	3 33	Cape St. Mary, the en-	1	
ı	Almeria	. 36 51	2 31	trance to the Gulf of	4	J
3	Cape de Gatt .	. 36 44	2 13	Venice	39 50	18 50
ÌĠ	Point Calla .	. 36 47	2 00	Otranto	46 19	18 55
S	CARTHAGENA	. 37 36	1 1		1	,
	Cape Pallos .	37 37	1	J	40 38	18 12
12	ALICANT	. 38 18	0 41		41 9	17 00
ğ	Cape St. Martin		0 29		41 40	16 8
ľ		. 38 47	0 10E.		49 20	14 30
S	VALENCIA .	. 39 96	0 20W	Ancona	43 38	13 39
South	Cape Gropeso	.40 6	0 SE.	Ramino	44 4	12 33
K	tiver Ebro .	. 40 42	0 27	VENICE		12 21
1	erragona	. 41 9	1 19	TRIESTE		13 47
1	BARCELONA .	. 41 22	2 12			13 48
١.	Cape St. Sebastian	- 41 53	3 9	St. Marie		14 15
					,	44 (

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ſ		Lat.	Long.	1	ì		Lat.	Lon
1_		D. M.	D. M.		Como Porreit		D. M.	D. N
.92	Pescera	44 45	114 58E.	1	Cape Ferrat .	•	36 2N 36 38	
Ę		14 17	15 35	i	Cape Tennis .	•	36 49	1 10
1	Cape Sesto	43 49	16 10	į	Cape Matifor .	:	36 55	2 4
S	Rosaro	43 🙎	17 10	1	Cape Beringu .		36 53	3 30
6	Raguea	49 40	18 11	1	Cape Tenels .		36 59	4 10
1	Cataro	49 23	19 29	١.	Cape Bugaroni .		36 50	5 47
ł	Durazo	41 30	20 00	i	Cape Ferro	•	36 52	6 54
1	Cape Patti La Valona	41 20 40 45	19 46 20 16	1	Bona Tabarca	•	36 32 36 48	7 30
1	Cape Linguette .	40 39	19 55	Bar	Cape Serra	•	37 10	9 24
ł	Butrinto	39 47	20 33	9	Cape Blanco .		37 20	9 5
1		39 94	20 36.	3	TUNIS		36 32	10 34
		. 38 16	22 4	පි	Cape Bon	•	37 5	11 9
ı		36 46	21 50		Susa	•	35 45	10 50
1		36 23	22 30		Cape Paul	•	35 14	11 1
1		36 27 37 53	23 12 23 2	1	Cape de Zoara . Tehy	•	33 53 33 12	11 10
اي.		37 58	23 2 23 46	1	TRIPOLI	•	32 54	13 1
Š		38 27	23 44		Lebida .	:	32 8	14 5
1	Cape Doro .	38 9	24 40	18	Cape Mensurato		32 12	16 1
F		39 25	23 17	1	Cape Lorat .		30 50	17 29
ıy fa	SALONICA, or Salo		00	12	Cape Linconta .	•	30 29	19 20
P.		40 38	22 56 23 40		Cape Serabion .	•	31 21	20 10
Ľ	Cape Ballouri . Mount Athos .	40 00	23 40 24 15	l	Zoara	•	30 44	20 4
	Contorn	40 39	23 58	1	Bengaza	•	32 16 32 55	21 20
;		40 40	94 50	1	DERNE .		32 48	22 1
		40 30	25 38	1	Cape Razatin .		32 30	23
1		. 40 3	26 6	1	Cape Luco .		31 50	25 00
		41 3	27 6	18	Cape Soliman . Point Ramadan .		31 44	25 18
	Galipoli CONSTANTINOPLE	40 26 41 1	26 37	15	Come Kamadan .		31 32	26 00
	CONSTANTINOPLE	41 1	128 55	S.	Cape Lagosego . Cape Capopera .		31 23 31 13	27 20 28 44
	XXX. The Black Sea	and Sea	of Azof.	1	Cape Rose		31 8	29 30
					ALEXANDRIA		31 13	30 16
•		Lat.	Long.	1	Aboukir		31 18	30 38
	T	D. M.	D. M.	1	Rosetta	•	31 24	30 58
Sea	Ismayl Akerman		. 28 50E	·	Cape Bourles		31 33	31 30
ä		46 11 46 28	30 16 30 37	1	Demietta		31 24 31 20	32 7 33 30
嚣		46 38	32 56	1	El Arish		31 13	33 50
	Koslof	45 15	33 25	1	Jaffa		32 5	35 3
•		44 40	33 36	1	M. Carmel .		32 50	35 16
1		45 22	36 27	Ţ.	Acre		33 00	35 26
1		45 12	36 30	3			33 10	35 19
	Teganroy . Trebizonde .	47 18	38 39		Cape Serpente .		33 28 34 22	35 35
1	Cape Yassoun	41 10	39 37 37 48		Cape Vardo . Tripoly	•	34 46	35 50 36 2
	las "	42 3	35 9		Tortosa	:	35 22	36 8
•					Cape Zaret		36 14	35 53
,	XXXI. The South Co.	est of the	Mediter-	·	ALEXANDRETTA	l, or		l
	, ranean S				Scanderoon .	•	36 35	36 20
ż		Lat.	Long.	1	Yasso	•	36 58	36 15
Į.	·	D. M.	D. M.	1	Cape Urco .		36 11 36 40	37 10 34 12
E	Cape Spartel	35 49N		1	Point Calvero	:	36 34	33 25
2	TANGIER .	35 40	5 49	1	Cape Draumonte		36 30	32 20
12	Ceuta	35 54	5 16	ż	Satalia	•	37 3	30 55
1	Tetuan	35 19	5 27	E	Satalia Cape Chelidoni . Rosso Island .	•	36 18	30 40
10	Cape Negril	35 14	4 24	K	Kosso Island .	•	36 8	29 59
1	Cape Baalai	35 10 35 98	3 44 2 57	1	Cape Seven capes	•	36 26 36 4 0	29 00
	Cape Three Forces	34 57	2 8	١.	Macri Cape Baibe		36 38	29 30 27 45
1	Cape Fegalle	35 47	1 9	ľ		Οc	36 45	27 15
1		35 59	0 48	1	Cape St. Mary		37, 38	127 7

	- -	at.	4	ong.	_	the second second	1	~	at.	Lon	2
	D.	M.		M.	.		- 1		M.		M.
Comp Displa				25E.		Capra	١ ١			.14 18	
	38		27					1			
	39	~	27	5		MESSINA .		38	14	16 49	9
	39			10	ı	Cape Orlando .	•	38	20	14 40	-
	40	3		24		Cape Cefala .		38		14 6	5
	10			50		Cape Cafrana		38	18	13 36	
	10		29	4	1	PALERMO .		38	7	13 20	
					l	Cape Alos	•	38	18	13 23	
XXXII. Islands with	hin i	the St	rail	3.	1	Cape St. Visto	•	38		12 50	
					1	Tripano	•	38	.9	12 36	
		at.		ong.		Cape Ruxe .			17	13 20	
	D.	М.		M.	Ē	Cape Alicate	•	37	3	13 50	
Alboran	36	1N.		00W		Cape Secha	•	136	49	14.26	
Fromentera, W. point	35	39		57E.	1	Cape Passari .	•	36		15 38	
	38			24 55	ſ	Saragossa .	•		5	15 30	
	38			5 5	ł	Cape Carmale	. •			15 39 15 43	
	39		1 -	25 50	1	Cape Moline .	•	37	3/	19 40	•
Salina	38		1 -	50 27	ĺ	Camer bell		20	56 ·	15 44	
Cabrera S. point .	39	12	Z	37	1	Stromboli	•		37	15	
MA IODGA S	20	ο0	ا	40	1	Lipari, S. point	•		31 44	14 5	
MAJORCA, S. point	37	20		49 00	l	Salina			32	15	
- N. ditto	40 39	7 45	2	7	1	Volcano , .			35	14 2	
				17	İ	Felicudi	•		36	14 19	
- E. ditto		42 34		39	ı	Ustica	•		5 <u>1</u>	13 20	
- MAJORCA TOWN	137	J-1	1	.,	1	Levaci		38		12 9	
Dragon Jelend	30	49		59	l	Maretimo .	•	38	ī	12	-
Dragon Island	39			30	1	Favognana .	•	37	5 6	12 2	
Minorca, S. point	39	43		42	1	Quill Rocks	•	37	35	ii i	
- PORT MAHON	30	52		22		Pantellaria .	•	36	45	12 3	
- N. point	40			35		Linosa		35	52	12 5	
ATT PUBLIC .	1~		١			La Pidossa	•		31	12 4	
Cape Corse	43	2	9	19	1	Lampien	•		30	12 30	
Saint Florenzo	142	35		16	l			ı			
Calvi	142	30		40	1	Gozo, N. point .	_	36	3	14 8	5
Aiscio	41	52	1 -	44	١.	C. Comoneto		35	54	14 11	
South point	41	21		21	R	La Valette .	:	35	54	14 29	
Cape Signo	140	14		37	Malta.	Cape Nicholas .		35	47	14 39	
BASTIA .		27	9		r	-	·			[·	
			1		l	Fano, entrance to	the			l	
N. P. Lagosardo	41	14	9	2	1	gulf of Venice .		40	5	19 2	2
Cape Asinara	40	53	8	6	çe.	Pelagosa .		43	23	16 9	0
Cape Caccia	40	31	8	7	Ę	Plana		42	20		3
Cape Otano	39	9	8	14	E	Tremiti		42	19	15 4	0
Cape Malfetena	. 138	50		54	F	Lissa, S. point .		42	57	16 1	5
CAGLIARI .	. 139	15	9	30	10	Pomo		43	13	15 40	6
Cape Carbonera	. 38	57	-	48	E	Longa, S. E. poin		44	1	15 4	
Cape Frances	, 39	39		50	۲	Cortu, S. E. point		39	47	17.	1
Olaster	40		-	34	1	Paxu, S. point		39	24	20 2	
Cape Cavallo .	. 40	48	9	47	1	St. Maure, W. poir	nt.	38	54	20 4	
					1	Cefalonia S. point		38	7	20 5	
Asinara Isl. N. point	41	7		14	1	— Cape Viscardo		38	30	20 4	
Antioch Island	. 38	5 ō		15	ا	Zante, S. point .		37	50	20 49	
Toro	. 38	47		18	ış	Cerigo, S. Point		36	20	23 1	
Galita Island	. 37	38	9		Į	Cerrigotto .	•	35	54	23 9	
Gorgona	. 43			51	臣	Cerrigotto Milo, Town Scio, Town Mytelene, Town		36		24 50	
	. 43			50	12	Scio, Town	•		30	26	
	42			12	12		•		13	26 27	
	42			7	ĺ	Tenedos	•		50	26	
	. 42			5	1	Lemnos	•	39	54	26 25	•
Monte Christo		17		26	Ŀ	la		L-	10		, I
	42			00	H	C. Crio	•		12	23 39	
Ganute	42			10	Įŝ	Cape Spada •	•	30	47	23 57	
Palmaria	40			00	p	Suda	.•	35	30	24 84	
Posca.		42	13		1	Cape Sassouso .			35	25	
Ischia, S. point	. 40	38	113	56	1	CANDIA .	nigi.	35	19	125 18	51

ı			al.	T	ong.	Т		TL	at.	7.0	ong.
ł			M.	D	M.	ı			M.	D.	
1	Cape Sidera				28E.		Madeira, Tristam poi				
1.5	Cape Salamone	35	00	26	20	Ę	-FUNCHAL .	. 32		16	54
13		Į.		l		ij	S. Deserms, S. point			16	25
Ę	Goza	34	50	24	1	Ę	Isl. Salvages, middle	. 30	13	15	42
1	Goxa, S. point	35	24	27	8	F	Piton	. 30	5	15	54
ł	Seapanto	35		27	7	ŀ	XXXVI. Can	fa	1 2-		
1	Rhodes, Town	86			30,	l	AAAVI. Can	ary [s	LUTLUS		
ł	Rhodes, CapeSt. Gioane	35	57	38	21				at.		ong.
I		l		١.	-	I		D.		D.	M
1		35			32	ı	Palma, Town .		39N.		
13		35			47	ľ	- N. point	. 28		18	
15		. 35	3		41	ſ	— S. point .	. 28		17	
15		34		33	8	l	Ferro, Valverde	. 27		17	-
#	Cape Grego	35	7	34	5	Į.	Gomero, St. Sebastia		6	17	8
1	XXXIII. The Coast of	AF	rica fi	rom	Cape	نو	Teneriffe, Hidalgo poi			16	
1	XXXIII. The Coast of Spartel to Cup	e V	rd.			P	- Orotava	. 28		16	
1						Ē	Tena Point .	. 28		17	1 .
•	ł	1_	at.	Γ_{Γ}	ong.	F.	Peak	- 28		16	
1.		D.			М.	•	— Port Christianos	. 27		16	
1.			49N.		54W	l	SANTA CRUZ	. 28		16	
1.g		35			13	ı	Canary, N. E. point	. 28		15	
	New Sallee	34	5		43	ł	— Palmas	.128		15	40 3
		33			25	ł	— S. W. point .	. 27	40	16	9
1	1 4	33	.9		40	•	Fuerteventura . Point Gorda .	: 28	46	13	50
ł	Cape Cantin		35	9	5	ł	G 187	. 28	4	14	
1	Saffia Bay	33	90		46 81	ı	Lanzarote, S. point	. 28		13	-
Ìģ		31 30	350			ı	— Puerto de Naos	. 28		13	
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Benegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayal, S. E. point	22 20 20 19 18 15 14 ster D. 39 38	19 15 56 33 10 7 58 47 11 Isla Lat. M. 44N. 96	16 17 16 16 16 16 17 nuls	40 10 38 20 3 31 33 ong. M.	St. Ant	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock Isle of May St. Jago	La D. 17 17 16 16 16 16 16 16	11. M. 12N. 8 2 59 46 46 28 45 49 6	L D 25 25 25 25 24 24 24 22 23 23	. M. 19W 8 15 6 55 37 12 56 45 14 5
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vo.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The Wa Corvo Flores Fayad, S. E. point Pico, Point de Repertal Summit of Peak St. George, S. E. point Graciosa, Villa da Praya	26 22 20 20 19 18 15 14 ster D. 39 38 38 38 38	18 15 55 33 10 7 58 47 10 10 10 10 10 10 10 10 10 10 10 10 10	16 16 16 16 16 16 16 17 nuls 28 28 28 27 27	40 10 38 20 3 31 33 ong. M. 7W 74 42 35 28 50 40	St. Ant	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock Isle of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle	Ls D. 17 17 16 16 16 16 16 15 15	14. M. 12N. 8 2 59 46 46 28 45 4 49 6 54 57 52	Lo D 25 25 25 25 24 24 22 23 23 24 24 24	. M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 22 23
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Pico. Senegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayal, S. E. point Pico, Point de Espertal — Summit of Peak St. George, S. E. point Graciosa, Villa da Praya Terceira, Angra St. Michael, P. Delegada	26 22 20 20 19 18 15 14 39 39 38 38 38 38 38 39	18 15 55 33 10 7 58 47 47 44N 44N 96 30 96 89 45	16 16 16 16 16 17 18 11 28 28 27 27 27 25 25	40 10 38 20 3 31 33 33 00 7 49 49 49 40 40 40 40 40 40 40 40 40 40 40 40 40	St. Ant	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock lale of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point	Ls D. P. 17 - 17 - 16 - 16 - 16 - 16 - 16 - 16 - 16 - 15 - 14 - 15 - 14 - 14 - 14	at. M. 112N. 8 2 559 446 446 228 4 4 9 6 557 550 550	L D 25 25 25 25 24 24 24 22 23 23 23 23 24 24 24 24 24 24 24 24 24 24 24 24 24	. M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 22 23 43
Pico.	Cape Bajador Cape Das Bartas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayal, S. E. point Pico, Point de Repertal — Summit of Peak St. George, S. E. point Graciosa, Villa da Praya Terceira, Angra St. Michael, P. Delegada — Point Ferraria — N. E. Point	26 20 20 19 18 15 14 ster D. 39 38 38 38 38 37 37 87	18 15 55 33 10 7 58 47 10 10 10 10 10 10 10 10 10 10 10 10 10	16 17 16 16 16 17 18 11 19 18 18 18 18 18 18 18 18 18 18 18 18 18	40 10 38 20 3 3 31 33 33 M. 7W 74 49 35 50 40 13 39 55 55 15	St. Ant	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Bonavista Leton Rock Isle of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point	Ls D. P. 17 - 17 - 16 - 16 - 16 - 16 - 16 - 16 - 16 - 15 - 14 - 15 - 14 - 14 - 14	at. M. 112N. 8 2 559 446 446 228 4 4 9 6 557 550 550	L D 25 25 25 25 24 24 24 22 23 23 23 23 24 24 24 24 24 24 24 24 24 24 24 24 24	. M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 22 23 43
Pico.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayad, S. E. point Pico, Point de Espertal — Summit of Peak St. George, S. E. point Terceira, Angra St. Michael, P. Delegada — Point Ferraria — N. E. Point Formigas or Ants	22 20 20 19 18 15 14 ster D. 39 38 38 39 38 37 37 37	18 15 55 33 10 7 58 47 11 Isla 44N 26 30 96 92 87 81 8 84 49 17	16 17 16 16 16 17 18 1 10 81 31 98 98 97 27 27 25 25 24	40 10 38 20 3 3 3 3 3 3 3 3 7 4 4 2 8 5 5 0 40 13 13 13 5 5 5 6 40 11 12 13 14 15 15 15 15 15 15 15 15 15 15 15 15 15	St. Ant	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock lale of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point	Ls D. P. 17 17 16 16 16 16 16 15 15 14 11 14 14	at. M. 112N. 8 2 559 446 446 228 4 4 9 6 557 550 550	L D 25 25 25 25 24 24 24 22 23 24 24 24 24 24	. M. 19W 8 15 6 6 55 37 12 56 45 14 5 30 48 82 23 43 Cap
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Pico. Senegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayal, S. E. point Pico, Point de Espertal — Summit of Peak St. George, S. E. point Graciosa, Villa da Praya Tercelra, Angra St. Michael, P. Delegada — Point Ferraria — N. E. Point Formigas or Ants St. Mary, Town	22 20 20 19 18 15 14 ster D. 39 38 38 39 38 37 37 37	18 15 55 33 10 7 58 47 11 Isla 44N 26 30 96 30 98 45 54 49 17 59	16 17 16 16 16 16 16 17 nuls 18 1 28 28 27 27 27 25 25 24 26	40 10 38 20 3 3 3 3 3 3 3 3 7 4 4 2 8 5 5 0 40 13 13 13 5 5 5 6 40 11 12 13 14 15 15 15 15 15 15 15 15 15 15 15 15 15	54 55.	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock lale of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point	Ls D. D. 17 16 16 16 16 16 15 15 14 15 14 14 14 14	at. M. 19N. 8 9 46 446 445 4 45 6 54 45 57 57 57 57 57 46.	L D 25 25 25 25 24 24 22 23 23 24 24 24 24 24 24	M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 22 23 43 Cap M. M.
Pico. Senegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayat, S. E. point Pico, Point de Espertal — Summit of Peak St. George, S. E. point Terceira, Angra St. Michael, P. Delegada — Point Ferraria — N. E. Point Formigas or Ants St. Mary, Town — West Point	26 22 20 90 19 18 15 14 ster D. 39 38 38 38 39 38 37 37 37 37 37 36 36	18 15 55 53 10 7 58 47 10 10 10 10 10 10 10 10 10 10 10 10 10	16 17 16 16 16 16 16 17 18 18 18 18 18 18 18 18 18 18 18 18 18	40 10 38 20 3 31 33 33 7W 7 48 35 40 11 39 55 15 54 10	54 55.	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock Isle of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point XXXVIII. From Co of Good CAPE VERD Goree Isle	Ls D. D. 17 17 16 16 16 16 16 15 15 14 14 14 14 14 14 14 14 14 14 14 14 14	at. M. 12N. 8 2 546 446 228 44 49 6 54 49 6 54 48 M. 48N. 440 At. M. 440 At.	L D 25 25 25 25 24 24 22 23 23 24 24 24 24 24 24	. M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 29 23 43 Cap M 33W
Pico, Senegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayad, S. E. point Pico, Point de Espertal — Summit of Peak St. George, S. E. point Terceira, Angra St. Michael, P. Delegada — Point Ferraria — N. E. Point Formigas or Ants	26 22 20 20 19 18 15 14 ster D. 39 38 38 38 37 37 37 37 37 37 36 36	18 15 55 53 10 7 58 47 10 10 10 10 10 10 10 10 10 10 10 10 10	16 17 16 16 16 16 17 18 1 19 18 19 18 19 18 19 19 19 19 19 19 19 19 19 19 19 19 19	40 10 38 38 38 31 33 33 33 31 33 35 40 12 35 50 40 12 39 11 11 11 11	14° . 20	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Nicholas, East point Salt Island Bonavista Leton Rock lele of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Brava, S. point XXXVIII. From Ca of Good CAPE VERD Goree Isle Cape Naze	Ls D. D.	at. M. 12N. 8 2 546 446 228 44 49 6 54 49 6 54 48 M. 48N. 440 At. M. 440 At.	25 25 25 24 24 22 23 23 23 23 24 24 24 24 21 21 21 21 21 21 21 21 21 21 21 21 21	. M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 29 23 43 Cap M 33W
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Pico, Senegal.	Cape Bajador Cape Das Barbas Cape Blanco Cape St. Ann Cape Myric Portendic SENEGAL, R. Breberic CAPE VERD XXXIV. The We Corvo Flores Fayal, S. E. point Pico, Point de Espertal Summit of Peak St. George, S. E. point Graciosa, Villa da Praya Terceira, Angra St. Michael, P. Delegada Point Ferraria N. E. Point Formigas or Ants St. Mary, Town West Point XXXV. Madeir	26 22 20 20 18 15 14 15 14 20 39 39 39 38 38 38 37 37 37 37 37 36 36 36 36 37	18 15 55 33 10 7 58 47 11 Isla 12 M. 44N. 96 96 927 81 89 45 54 49 17 59 69	16 17 16 16 16 16 16 17 18 1 18 1 28 28 27 27 27 25 25 24 25 25	40 10 38 38 38 31 33 31 33 33 7W 7 42 28 55 40 11 12 55 15 54 10 14	agumbia.	St. Anthony, N. W. N. E. point SANTA CRUZ St. Vincent St. Lucia St. Nicholas, N. poi East point Salt Island Bonavista Leton Rock Isle of May St. Jago PORTO PRAYA N. W. point Fogo, N. point Middle Brava, S. point XXXVIII. From Co of Good CAPE VERD Goree Isle Cape St. Mary, ent. the River Gambia	Ls D. D. 17 17 16 16 16 16 16 15 15 14 14 14 14 14 14 14 11 11 11 11 11 11	at. M. 12N. 8 2 559 446 445 4 449 6 54 426 557 550 rd to 48N. 49N. 40 29 19	L D 25 25 25 25 24 24 24 22 23 23 24 24 24 24 24 24 24 24 24 24 24 24 24	M. 19W 8 15 6 55 37 12 56 45 14 5 30 48 92 3 43 Cap M 33W 25 12 40
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Femba of Reddree,S. P. 6 30 40 19	ŭ	- N. P	5		-	-	1	Cape Israel	15	15		
Mombus Harbour, ent. Chence River	1	remon or Leadree, S. P.	6				ı	Gebel Zabayr	15	3		
Chence River 3 37 40 7 Gebel Zeghir 14 2 42 52 Quilise River 3 25 40 19 Gebel Zeghir 14 2 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 42 52 Great Arr	1	— N. P.	4				ĺ	Hodeida	14			
Chence River 3 37 40 7 Gebel Zeghir 14 2 42 52 Quilise River 3 25 40 19 Gebel Zeghir 14 2 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 41 42 52 Great Arroe 18 42 52 Great Arr	1	Mombas Harbour, ent.						Shoal off Ras Magame!	14		42	56
Quilise River	1	Chence River	3	37			ı	Gebel Zeghir	14	2	49	52
Cape Bassas 4 50 48 49 Cape Bassas 4 50 48 49 Cape Bassas 4 50 48 49 Cape Guardafui 11 50 51 20 Cape Guardafui 11 50 51 30 56 30 Cape Chanseley 18 Cape Chan	1	Quilise River	3	25			1	Great Arroe	18	41	42	52
Formosa Bay, S. point 3 00 41 2 2 39 41 91	1	Leopard's Reef	3	16	41	2		MOCHA	13	20	43	20
— N. point	1	Formosa Bay, S. point	3	00			1		l	7	1	
Daedalus Shoal 0 23 43 4 Cape Hargiah 13 30 47 2 Juba 1 8N 44 10 Magadoxa 1 8N 44 10 Cape Bassas 4 50 48 49 Moro Cobir Point 8 30 50 45 Cape Orfui 10 00 51 17 Cape Orfui 10 00 51 17 Cape Guardafui 11 50 51 32 Cape Guardafui 11 50 51 32 Cape Guardafui 11 50 51 32 Cape Curia Muria, I Western 17 33 55 40 Cape Chanseley 18 2 56 30 Cape Chanseley	1	- N. point	2	39			į	Cape St. Anthony .	12	39	44	14
Daedalus Shoal 0 33 43 4	ı				41	18	H	Cape Aden	12	44		
Juba 0 19 43 2 Macula Bay 13 57 47 58	1		0	23	48	4	Æ	Cape Hargiah	13	30		
Name	1		0	12	43	9	١,					
Magadoxa 2 5 45 49 Kisseen Point 15 19 51 20 Cape Bassas 4 50 48 49 Cape Fartash 15 34 51 56 Moro Cobir Point 8 30 50 45 Cape Delgado, north 10 00 51 17 Cape Oriui 10 92 51 39 Cape Guardafui 11 50 51 32 Cape Guardafui 11 50 51 32 Socetra Island, E. P. 12 30 54 52 Cape Chanseley 18 39 56 30 Cape Chanseley 18 39 56 30 Cape Chanseley 18 30 56 30 Cape Changeley 18 30 56 30 Cape Changeley 18 30 56 30 Cape Changeley 18 30 56 30 Cape Changeley 18 30 56 30 Cape C	ľ			8N.	-144	10	1					
Cape Bassas 4 50 48 49 Cape Fartash 15 34 51 56 Moro Cobir Point 8 30 50 45 Cape Delgado, north 10 00 51 17 Cape Orfui 10 92 51 39 Cape Guardafui 11 50 51 32 Socetra Island, E. P. 12 30 54 52 Cape Chanseley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Chanseley 18 32 56 30 Cape Chanseley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changeley 18 32 56 30 Cape Changele	1		ه ا	5	45	49	ı					
Moro Cobir Point 8 30 60 45 Dofar 17 3 54 10 Cape Delgado, north 10 92 51 37 Cape Morebat 17 00 54 32 Cape Guardafui 11 50 51 32 Cape Guardafui 11 50 51 32 Cape Chanseley 18 2 56 30		Cape Bassas					1					
Cape Delgado, north 10 00 51 17 Cape Morebat	1	Moro Cobir Point					ı					
Cape Orfui	1	Cape Delgado, north										
Cape Guardafui . 11 50 51 32 Curia Muria, I. Western 17 33 55 40 Cape Chanseley . 18 2 56 30	1	Cape Orfui					1					
Socotra Island, E. P. 19 30 54 52 Cape Changeley	1						ı					
	1							Cape Changeley by	18	916		
		- W. P.										
	ضة				Ť		_					

TABLE XLVI. Latitudes and Longitudes.

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		al.	Ĺ	ong.	1	1		at.	17	ong.
		. M.	D		ı	G		M.		. M.
azeira Islands, S.W.P.					ı			50N.		
N. E. P.	20	35	22	56	•	Cape Dajera	26	370	54	
salhad, (Cape Rasal-	1			1	ı		26		54	
gat)	22	22		58		Linga	100	33	55	1
pe Kuriat	23	20	58		S.	Kishm Island,	L.		l	
uscat	23	3 8	58		3		26	32	55	22
ırka	23	41	57	56	Ö	Point Bassadore, or			l	
int Deba	25	34	56	20	.₹	N. W. P	26	38	55	
	1		1		Persian	— LuՈւ	26	55	55	55
pe Mussendorn .	26	21	56	38	Pe	- N. Point	27	2	56	20
ab, or Gap Island .	26	22	56	42		- Kishm Town	26	57	56	24
	26		56	44	The	Angaum I. mid	26	39	55	57
pe Jedee	26	8	56	12	ŀ	Lanek I. (conical hill)	26	52	56	28
mps	25	54	56	3		Ormus I. N. end, fort			56	37
	25		56				27	13	56	22
nrah, or Red Island			55			Res Koli, or Cape Hill	26	20	67	30
arga	25	22	55			Kohumbarek, or Bom-				
	24		53		i	barack Rock	25	52	57	46
Taland	25		53			Cons Vahumbash	los			46
	25		52			Kohumbarek Shoal .	95		57	
rnain Island	25		52			! <u></u>	i		٢	
zenie Island .	24		52			Cape Jask	25	38	68	10
	24		52		l		25	16	61	
					l			4	63	
	24		52		l	Cape Gwadur or Guadel		7	65	
	25		52		ď		25	•	۳	~
	25	15	52	19	ersla.	Cape Monoze or Mewa-			Lee	£A.
wlooi Island, (pro-			_		5	ree	34	51	66	
	25		52		Г	Korauchee or Crotchey		40	67	7
s Reccan	26		51		l	River Scind, ent. grand		_	L _	
or Hussan	26	3	51		ŀ	branch	24	8	67	
tif Bay, N. P.	26	36	50		ł	Tattah	34	44	68	
ree low sandy isls.	27	42	49		1	Bate Castle	22	385		20
	27		49		l	Point Gigat	22	99		16
	. 98	5 3	48	16	ı	Diu Head	90	42	71	6
.sulurhud, (S. E. par	t]				l	Jaffrabad	20	55		36
of Graen Haven)	29	20	47	57		Searbett Island	20	55	71	
	28		48	38	ı	Gospust Point	91	12	72	
aen	29		47	48	ı	Cambay	32	94	73	39
erleeche Island .	29	3 0	48	25	ı	Swallow Point (Vanx	1		1	
sra river bar	29	57	48	42	l	Tomb)	21	4	73	51
SRA or BUSSORAH	30	29	47	40	ı	Surat Castle	21	11	73	5
lam	29	55	50	18	1	Demaun	90	22	73	9
	. 29	46	50	27	i		90	3	73	6
	29	16	50	27	ļ	Terrapore Point	19	50		49
sheer	29		50	56	i	Basseen River	19	18		12
dilah Hills, N. P.	29	19	51	26	ı	Bombay, (flagstaff) .	18	56	72	58
	28	29	51		ŀ	Bombay light house .	18	54	72	
zarine	28	2	49		i	Henery and Kenery Isl-	1		1	
yn	27			4	i		18	49	72	67
pe Berdistan	27	58	51		ı	Coulisba Island	18	37		59
nklun	27			6	i	Chaoul	18	39	73	
pe Nabon	27	94	52		1		18	16	73	
	26		54		ı	Radjapour harbour . Bancoot River	17	57	73	
	26		53		ı	Same Jane	17	47		13
	26		50		Ī	Sevendroog	17	46		
rable Shoal	26				ı	Dabul	17	20	73	
escent Shoal	1		50		1	Argenwell Fort	17	9E		16
	26	77	51	43	I		1:5	20 16	1,3	17
ditto Favourite	مما	-0			ı	Zughur Point		16	123	2/
soundings	26		51		ı	Rettna-Geriah	17	2		23
	26		53		i	Radjapour Fort	16	47		\$ 5
	26		54		I	Geriah point & flagstaff	116	31	73	
	26		54		ı	Angrias Bank, N. P	116	38	73	8
eat Tumb I	26		55			— S. P	116	18	73	8
tle Tumb I	26		55		ľ	Dewghur harbour .	16	23	73	39
bfleur I.	26		54	34		Atchera River	116	11	73	35
mesa Island	125		55	9		Melundy (fortified Ial.)	16	3	73	364

LITTE

_		Lat.	1.	ong.			Lat	Long.
I	•	D. M.		M.	l		D. M.	D. M.
		18 7N	93	56E.		P. Brala, or Capas de		
	Church (or St. John's			~		Mer .		103 37E.
	Rocks)	17 28 16 52	94 94	7	Į.	P. Capas de Terra	5 15 5 21	103 8 103 4
		16 21	94			Tringany River, ent Great Redang Island .		103 00
	Cape Negrais	16 2	94	13		Pulo Printian		102 40
		15 52	94	19		Calantan Road		102 17
	Sunken I. (or La Guar-	15 41	94	15		Cape Patani	7 4 7 19	101 51 102 43
	da) Rang oon, or Pegu rive r		73	10		Pulo Lozin	8 29	100 58
	entrance	16 29	96	25	١.	Siam River, entrance .	13 30	101 15
	PEGU	18 00	96		Ę	JUTHIA or SIAM . Cape Liant	14 18	101 13
	Tanan Daint	16 28 13 33	98	30 6	Š	Cape Liant	12 27 9 55	101 37 103 40
Ę.	Tavay Island	13 13	98	9	ı	Pulo Ohy Folse	1840	104 34
7	Cabossa Island	12 46	97	29		Pulo Oby	8 25	104 54
	West Canister Island .	12 40	97		ı	Cambodia Point	8 35	104 56
		12 36 12 12	97	30 24	1	Cambodia River, W. en.		106 30
	Mergui Tores Islands, western	11 50	97	3	1	Cape St. James (E. ent. Saigon R.)	110 17	107 4
	Small Rock	11 21	97	15	1	Cape Trivoane	10 21	107 16
		11 10		57	1	Point Babeck	10 30	107 33
	St. Matthew's I. Sever's Islands, N. P.	9 55 8 43	98	4 48	i	Brittos Bank, N. E. P. Cow Island	10 32	107 56 107 52
	_ S.P.	8 28	1	48		Point Kega	10 41	106 4
	Junkseylon I. N. P.	8 9		20	ı	Point Kega Point Vinay	10 54	168 19
	— S. P.	7 46		20	l	Mui-guio or Little Cape		108 31
	Parlis River	6 21	100 100		l		11 9	108 40 108 48
	Queda	6 6	100		I	Pulo Ceicer de Terre . Cape Padaran .	11 21	109 00
	Prince of Wales' I. For				1	Padaran Bay	11 35	109 4
	Cornwallis .	5 24	100		1	Cape Varela False	11 44	109 12
	Cape Caran	3 32	101	.8	1	Carmaigne Harbour,en.	11 49	109 12 109 19
	Salangore Hill and For	3 20	101	10	ď	Water Islands Tre Island		1109 19
	S. P.	2 56	101	16	謹		12 21	109 19
	Parcelar Hill	2 52	101		ပ္	Nhiatrang	12 26	109 10
	Parcelar Point	2 42	101	32 ·	Ę	Three Kings Rocks	12 43	109 23 109 12
	Tanjong Tuan, (C. Ra- chado)	2 29	101	52	ట్	Cape Varela, or C. Pagoda	12 45	103 12
	Tanjong Clin or Peer				l	goda	12 55	109 25
ay.		. 2 17	102		ı	Perforated Rock	12 59	109 25
Malay	Fisher's Island	. 2 13 . 2 12	102 102	_	1	Phuyen Harbour, ent.	13 23	109 14
•	Malacca Fort Water Islands, southern		102	-	1	Coumong Harbour, ent	13 33	109 13 109 21
	Mount Mora or Moar	. 1 59	102		1		13 44	109 14
	Mount Formosa	. 1 49	102		ł		13 50	109 14
	Mount Battoo Ballo	1 39	103 103		i		14 11	109 14 109 7
	Fulo Pisang .	1 28	103		1		14 19 14 39	109 7 108 56
	anjong Boero	1 15	103		1		15 23	109 6
	Little Hill, or False Jo		١.,		ı		15 23	108 44
	hore Hill	. 1 26	104		I		15 54	109 33
	Johore Hill . Barbucet Hill .	. 1 23 . 1 25	104		1	Cape Turon or Tienchu Callaohanne I. (N. ent.		108 15
	l'oint ROMANIA	1 23	1	18		Turon)	16 11	108 7
	False Barbucet Hill	. 1 30	104	16		Cape Chouvay	16 21	107 51
	Romania Reef	. 1 25	104	25	ı	Hue or Huesso River		107 26
	Eastern Bank, (oute part)	1 39	100	35	I	Tiger Island Hainan I. and adjacent	16 55	107 93
	Pulo Tingy	2 17		: 33 : 8	1	Islands.	7	
	Blair's Harbour	. 2 43	103	40		4	18 94	108 59
	Pulo Varola	. 3 16		47	ı	- Yulenken Bay. Zenby	18 11	109 35
1	Pahan Road Fingoram	. 3 31	102	18	ı	- South Pt. of Hainer	112 10 112 10	109 34
	Howard's Shoal	4 14		31	1	Galong Bay . Digit Brother's I's. Eastern	18 11	109 41

ſ		Lat.	Long.				æt.		ong.
1		D. M.	D. M.	ĺ		D.	M.	D.	M.
1	— Luengsoy Point S. P.	15 22N	110 OOE.	l	Prince Edward's islands	۱			400
		19 26	110 8				538.		46E.
ı	— Saddle Island	18 35 18 40	110 11 110 24	l		46	40	38	8
1	— Point of land	18 38	110 24 110 21		Kerguellan's Land, or	1		1	
1	Tinhosa Island .	18 40	110 21	ĺ	Isle of Desolation, — Bligh's Cap, N. P.	48	29	68	44
Į		18 49	110 34	ı	- Christmas Harbour	48	_	1 ::	4
1	- Toongean Mount.pt.		111 2	ı	— Port Paliser	49	3	69	
ı	— Hainan Head N.E.P.	19 59	110 54		— Cape Digby, or E. P.	49	23	70	
1.	-South Taya Island .	19 49	111 12		— Cape George, or S,P.	49	54		10
Ę	- North Taya Island .	19 59	111 17		— Island Solitaire	49	49	68	
Ę	Nowchou Ty-foong-kyoh I. (Tienpak harb.)	20 58	110 26			49	3	68	23
Œ	Ty-foong-kyoh I. (Tien-				St. Paul's (or Amster-			77	Ea
r	pak harb.)	21 22 21 26	111 13 111 2 5		dam Island) . Amsterdam (orSt.Paul's		52	77	026
1		21 31	111 40		* * * * * * *	38	47	77	59
1	Mamee-Chow, or the	r. "	*** 40		Danish Rock, doubtful			98	
	Twius, near S. W. P.				Cloate's I. (lon. uncer.)		7	112	1
1	of Hai-ling-shan .	21 34	111 50		Tryal Rocks	20	40	105	
1	In 10 10 10 10 10 10 10 10 10 10 10 10 10	21 43	112 15		Rosemary Islands) very	20	23	115	55
1	Nampang I	21 34	112 12		A reef 10 miles near			1	
ĺ		21 28	112 22		N.W. of Rose New	20	18	1	
1		21 39	112 29		mary Island Hol-	00	90		3E '
	Haw-cheun, S. W. end	21 30	112.31		Abrohlos Shoals J land			113 105	
1	Passage I. (near S.W.P.	21 35	112 34			10	33	103	~
1	Wy-Causs I. (near S.	~: 35	-10 34		Cow Isles, — Northern	11	50	97	4
•		21 34	112 47	Н		12		97	_
I		21 36	112 52		Clarke's Reef and Im-				
1		21 47	113 1			17	32	119	14
ſ	Ty-kam I	21 52	113 1		Dampier's or Scott's			l	
1.	Cou-cock I.	21 60	113 7		Reef, N. W. end .	13		121	
Ę	Tyloo I. S. P	21 52	113 14	ı	— N. É. end	14	1	122	
Ë			113 44	Н	Coral Bank	13	32	124	
ł	TOTO OF T ASSESSO 1.	22 2	113 39		Coral Bank, 9 fathoms	13	25	124	
1		22 8	113 49	П		12		194	
ł			113 33 113 32		Cartier's Sandy I. or Bank		zo	193	90
ł	Macao, City Lantoa or Tyho I.S.W.P.	22 10 22 19	113 50		Red Island, (very near	15	9	124	99
1		22 24	113 48		New Holland) Coral Bank. 10 f. or less	• -		124	
1	Asses Ears	21 54	114 1			iĩ		128	
1	Great Lema I. N. E. P.	22 5	114 18	ŀ	Sahul Shoal, S. W. P.	1		1	
ł	Nine Pin Rock	22 16	114 22	l	12 fath.	11	35	124	14
I	Whampo anchorage .	23 6	113 92		Echo's Soundings, 14	l			
	CANTON	23 7	113 14		fath.	I	16	125	
1	XLI. Islands and Shoa	la in the	INDIAN	1	Coral 7 fath. Bank .	ļŸ	56	129	25
1	XLI. Islands and Shoa OCEAN, between ti			١.	Partura Charl	33	8	43	5
1	Cape of Good Hope as				Fortune Shoal	35 35		41	
1	ding those West and J			ŀ	Union Shoal Dutch Bank		44	44	
1	Holland.			1	Otter's Shoal, doubtful			36	
1		Lat.	Long.	l	Princess Augusta's Shoal			1	
1	1		D. M.	ı	doubtful		44	36	16
1	Dutch Bank, Stot Van				Union Rocks, doubtful			41	20
1	Capelle,	ł	1		Swallow Rockett Break.	j i		1	
ŧ	Various situations from	40 008.	38 50E.	١.	ers, doubtful	28	20	42	10
1	├─_ to . <u>.</u>	36 00	43 30		Belliquese Shoal, doubt.	28	43	42	83
1	Telamaque Shoel,	ł	1	ľ	Star Bank	25	15	44	10
ł	(doubtful)	40 0		Ŀ	Madagascar Island,	۱, ۵		40	25
1	Various situations from	39 9	21 57	Ŀ	- Cape Amber, N. P.	12	14		30
1	— to Brunswick Bank.doubt.	38 90	23 24	ğ	Cape East	146	27		23
ł	French Shoal, doubtful	39 9	36 19 43 6	ŝ	— Bay Antongil — St. Mary's Island .	17	00	50	25
1	Atlanta's Rock, doubtful		52 00	Madagascar	— Foul Point	117	41	49	36
•	Wellington Shoal, very		l ~~ ~~	¥	— Fort Dauphin Digitized	25	J 5)(P46	35 1
ŀ		39 53	71 43	١,	_ Cape St. Mary .	25	40	95	13
-		سننسند				_	_	_	

_		-				-		-7		7	
1 1	1		at.	L	ong.	1			at.		ong.
1		D.	М.	D.	М.				M.		M.
1	- St. Augustine Bay			۱.,		ı		19	52S.	55	23E.
1			39 S		00E.	ı	Galega or S. Roquepiz,		95		20
1	- Cape St. Vincent	21		43		ı	mid.		25		39
	— Cape St. Andrew .	16	2	15	16		Saya de Malha Bank				20
1	Madagascar Island .	١.		Ι.	_		limits)		18		58
	- Table Cape	15		46	6	ı	Fortune Bank		12		40
	— Bembatooka Bay 🛭 .	15	43	46	28		John de Nova		15		12
•	- Majambo Bay .	15		47	6		Providence Island .	9	9		00
1	- Narrenda Bay .	14	31	47	45		Coetivy Island	7	14	56	32
1	— Dalrymple Bay .	13	31	49	9		Chagos Archipelago .	۱.		1	
1	- Passandava Bay .	13	45	48	23		- Diego Garcia	7		79	22
1	- Cape St. Sebastian	12	28	48	54		- Diego Garcia }		14	۳.	
I8.	(25	7	100	16		— Pitts' Bank		29	71	95
12	Star Bank }	25	25	44	10	9	- rice bank	-	<i>5</i> 0		
Š	Bassas de India	22	28	40	37	rpelago.	- Centurion's Bank .			70	53
٤	Europa Rocks	21	28	40	8	lä	- Ganges Bank	7	26	70	50
じ	Sussex Rocks	21	25	42	36	Ę	— Owen's Bank	6	46	70	12
1	Bazaruto Islands .		20		12	Arch	- Egmont's, or Six Isl-	1		1	
1	Barren Islands, western				15	Ι,	ands	6	37	71	24
1	English Bank	17	40		15	ŝ	- Danger Island .	6	21	71	18
	Juan de Nova or St.	ľ		1		å,	— Eagle Island		10	71	23
ı	Christopher's I.	17	3	43	7	Ü	- Three Brothers	6	9		35
1		17		44		ı	- Peros, Banhos Isl'ds.	1 -	92		53
ŧ	Chesterfield Shoal		20	14	_	ł	- Saloman's Islands .		23		20
ł			54	45			- Sandy Islands .		17		37
1	Mayotta Island		20		50	ı	- Speaker's Bank	4 -	ô		26
1	Mohilla Island .		15		34	l	- Speaker's Deale	1 -	••	١:-	
1	Johanna Island, Peake		32		25	ı	Pona Molubque Atoli,	1		1	
1	Comoro :				_	ı	S. P.	10	41	72	20
ł		12	30	40	50	ı	_ N. W. P.		34	1	12
1	John Martin's Island		1.5			ı	N. E. P.		83		25
1	(doubtful)	10	15		50	l			21		35
•	Firebrass Bank		16		20	ı	Addon Island (mid.)		21	13	~
1	Aldabra Islands, N.W.P.	9	23		46	ı	Suadiva, southern group		037	-	12
ı	Assumption Island .		46		16	l	- South Reef	0	9N		
1	Cosmoledo Island .	9	46	48	20	1	South Island		11		19
1	Marquis of Huntley's	١.		l		1	- S. W. Island		18	73	4
1	Bank		57		20	ı	N. W. Island		28	73	3
1	St. Peter's Island .		28		42	1	- N. Island		34	13	8
1	Natal Island (doubtful)		26		12	ı	- Northern group, S.			L	
1	Sandy Island		10	48	10	l	W. Island		48		19
	ot. Lawrence Island .	9	13		58	1	N. W. Island .		51		20
	Zanzibar Island, S. P.	6	28	39	46	1	- N. E. Island .	10	58	73	33
1	— N. P.	5	40	39	46	1	Adoumatis Atoli	1		1	
1	Amirante Island, N. W.P	. 5	10	53	45	1	S. W. extremity	1	50		27
1.	- S. E. P	. 6	20		30	ago.		1	49		33
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8	- S. E. P.	. 5	30		59	Archipel	N. W. Island	2	7	73	35
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1	- South Islet	. 16	47	59	34	1	Cordivia Island		00		36
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Pulo Pisang		104 6		Three Sisters		105 48
little Fortune Island				North Island	5 41	105 49
Engano or Deceit I.N.P.	5 15	102 25	, S	Two Brothers, northern	5 9	106 5
- E. point	5 22		الق	Lynn Shoal	5 12	106 13
— S. E. point	5 30	102 38	3 8	46 · · · · · · · ·		105 58
		102 29				106 14
S. point W. point		102 19	4	ucepera Sent St. Banca	3 19	106 10
•	ļ		- Ji			105 48
Java Head	6 48	105 11	- 1		3	
First Point		105 19	- {ı	Banea Island	. 1	
second Point	6 36	105 21	Ľ		ا ۽ و	106 40
Chird Point		105 40			36	100 40
	6 3	105 54	1	-Tanjong Panjong or	ا م. ـ	٠,,,
Bantam or St. Nicholas P.	5 52	106 2	1		1	106 4
		06 10	J,		1	105 14
RATAVIA Observat	6 5		Į.		1 43	105 20
BATAVIA Observatory	6 9	106 52	-	- Tanjong Muncoods,	j	
Carawang Point		107 12	1	N. of Banca .		105 51
edary Point			ŀ	— Tanjong Tuan .	1 39	106 6
		107 49	-	- Songy Leat Bay .	1 50	106 9
Voerden Castle Rock		107 52	-			106 14
'rincessCharlotte Shoal		107 54	. -	- Goonong Marass		
		"	۱:		1 53	105 53
		108 22	ĕ -			06 52
Jumkin's Island or out-		, [Banca			06 54
	5 47	108 23	7		- 50	.w. 43
heribon Mountain .		108 26	٦	- Entrance Point, or	. ا م	INC 54
	6 50	104 14	ſ	S. E. P	3 2	106 54
***	, ,,	M 2 172		^	igitizad by	

1-	*************************************	_	et.	1	ng.	-	· · · · · · · · · · · · · · · · · · ·	_		. Yan	_
ı			M.	Ď.	.M.	ı			Lat. M.	Lon D. N	
1	Essex Shoal or Fairlie					ı	Bintang I. N. W. P.	i.	10N.	104 19	9 E
ł	Rock	3	27S.	107	2E.		Johore Shoal	ī		104	
1	Vansittart's Shoels	3	11	107	2	1	Shoal ent. Rhio Straits	1	8	104 1	
1	(107	8	I	Sincapour I. E. P.	1	22	104 00	
1	Pulo Leat or Middle I.			107	5	ı	Pulo Battain, N. E. P. St. John's I. S. P. Rocky Reefs	1	10	104 4	
1	Alceste Shoal	2		107		ł	St. John's I. S. P.	1	14	103 51	
1				107		ı	Rocky Reefs Middle Island Coney Island Buffalo Rock Rocks Red Island	1	9	103 55	5
1				107			Middle Island	1	13	103 46	5
1	North island			107		l	Coney Island	1	9	103 41	L
1	General Hewitt's Rock			107		ı	Buffalo Rock	1 1	9	103 48	3
•	Discovery Rock			106			Rocks	1	6	103 45	
ł	Pulo Glassa or Gasper I.			107	-	l	Red Island	1	6	103 38	
Ł.	Tree Island	2	28	107		ı	Tree Island	11	8	103 30	
	Warren Hasting's Shoal	Z	23	106		ı	Alligator Island	1	10	103 40	
ŧ	Belvidere's Shoal Vansittart's Shoal Hillsborough Shoal Magdalen's Shoal Severn's Shoal	2		107			Rocks]]	12	103 36	
15	Willshammer Shoel	Z	3	106 106		Ŀ	Little Carimon Great Carimon, N. P	1 !	8	103 24	
ŏ	Mondolonie Shool	3	56	107		Ø	Great Carimon, N. P.	!!	.7	103 21	
1	Severn's Shoal .	1		106		ř	The Brothers Pulo Cocob	1:		103 21	
ı	Billiton Island, S. E. P.	2	99	108		ğ	Pulo Cocob	1:		103 25	
1	S W Point	3	15	107		ş	Water Islands on Perm	1 *	20	103 16	9
1	—S. W. Point	2	33	107		2	Pulo Pisang Water Islands, or Four Brothers, S. P.		4	102 20	n .
	N. W. Island off Billiton				00	ı	Fisher's Island	6	_	102 12	
1	Shoe Island (formerly		-			l	Bambeck Shoal			101 4	
1	Bird I. and White R.)	3	47	108	2	1	Pulo Callam or Colong,	7	J	101 31	•
1	Fox Shoal	3		110	4	ı	S. P.	9	56	101 16	6
1	Pulo Mancap	3	5	110		l	Two and half fathoms		•	102 10	
1	P. Mancap Shoal, S. P.	3	22	110	11		Bank		54	hoı 4	1
1	Discovery's West. Bank	3	39	108	43		Round Arron	1 .	49	100 49	
1				109	10		Blenheim's Shoal .		3	101 2	
1	— Reef	3		108		ı	Long or Great Arroa .		52	100 44	4
	Osterly's North Shoal .	3	19	108	40	•	Two Brothers, Pulo	ļ			
1	Cirencester's Sand Bank			109		1	Pandan	3	24	99 54	1
1			54	108	58	ı	Pulo Salanama .		21	99 52	2
ı	Montaran Islands south					•	Pulo Varela	3		99 36	5
1	Eastern		31	108	52	ı	Pulo Jarra			100 14	ŧ
ı	— Tockockemou (high-		••	-00			Sambilang I. Southern		3	100 35	
•	est Isl.)	7	31	108			Dinding I. W. P.		16	100 35	5
1	Minto Rocks	7	14	109		1	Prince of Wales' I. Fort			l	_
I	Ontario's Shoal	3		108		يدا	Cornwallis		24	100 21	
1	Rendezvous I	25	44 42	110		Š	Pulo Pera		42	99 1	
Ŀ	Souroutou, W. P Carimata I. Peak .	1		108		Š	Boonting I. Southern .	5	45	100 18	
	Pulo Donen	1		108 109		Ę	Pulo Bonton (Dome)			99 20	
15	Pulo Panumbangan	1	12	109		3	Pulo Ladda, S. P	6	8	99 42 99 39	3
Bulliton	Massa Teega Isles	ĥ		109		ı	Trotto I. N. P	7	47 10	98 50	
F	Greig's Shoal			108		•	Sangald or Guilder Rock	1 4	14		
		ľ	70	1.00	3,	I	Pulo Telibon The Brothers	7	31	99 29	
1	TheSeven Islands, N.W.	1	8	105	94	1	Pulo Rajah or P. Taya		36	98 20	
I	Pulo Varela or Barallah			104			Junkandon C D	7		98 20	
1	Pulo Taya		45	104		ı					
H			35	103		ı	XLIII. Islands and Sho		in th	chij	N.
1	Ilchister Shoal	0	26			I	SEA.				
	Lingin, Tanjong Eang,	-			•	ı		L	at.	Long	
	N. K. ovi		20	105	4	ı				D. M	
1	East Domino Island .	0		105	4		St. Barbe Island .			107 15	
1	Geldrias Bank, same as	ľ		1		ı	Direction Island .		15	109 5	
1	Dogger Bank	0	48N.	104	58	Ľ	Pulo Datoo	0	7	108 36	
1	Rhio		57	104	30	ľĔ	Welstead's Rock .	-	32	107 5	
1	Eastern Island off Pulo			1		k	St. Esprit Islands, E			107 13	
Ŀ	Panjang	0	54	104		Lŝ	Green Island:	-		107 30	
	Island		48	104		Tambelaris	St. Julian Island	0	54	106 48	5
I	Three Brothers, south .		31	103		Γ	Tambelan Islands, East	١.		-0= -	_
IĐ	Pedro Branco	_	20	104		ı	or Great I.	~ ~	00 ∏	107 35	
Lŝ	lalands off P. Romaine		23	104		ł	Gap Rock Digitized by			107 3	
種	Bintang L (the hill)	1	2	104	30	1	Europe Shoal	1	13	107 9	_
											7

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			Lat.	Lor		1	1		Lat.		ng.
	<u> </u>		М.				ا ہے ا		М.	D.	M.
•	Rocky Island	1			14E.	1	Round Island or Great		المدر	١	
ı	Camel's Hump	1		106			Catwick				52 L
1	Saddle Island	1		107				10		108	5 3
1	French White Rock	i.		106				10		110	_
1	Victory Island	i		106		1	· '	1	j	1	
	Acastee Rock	i		106		1	Investigator's Coral	1	ì	1	
	White Rock	2		105		()		14	19	112	59
1		2		105		1	Triton's I. or Bank S.		1	1	-
ſ	Macedonian Reef					1 1		15	AR I	1111	11
ı	South Anambas, li- §	2			8	()				111	
1	mits · · (·	2		106		1	Passoo Keah (Sandy I.)	10	3	111	10
1	Pulo Domar		45	105	27	()	Bombay Merchant's	۱	ا م	I	•
ŀ	Middle or G. Anambas,		ا . ـ	١	ا	1	1 '	16		112	
ı	W. limit			105			—S. P	15		112	
1	North Anambas		27	106		1	Discovery Shoal, W. P.	16	11	111	
3	Pulo Tingy		17	104	8	(- E. P	16	16	111	
3	Ex. Islet off P. Tingy			104	-		Jehangire's Coral Bank	16	18	112	35
5	Pulo AOR or Wawoor	2		104		1	Vulador's Shoal, E. P.	16	19		7
Æ	Pulo Pisang or Pambee-		- I	(~	- N	1	- W. P.	16	18	112	
1	lan	9	37	104	21	1	Crescent Chain	۱	- 1	1	-
		1	44	104		1	- Money's Island	16	28	111	30
ı	Pulo Timoan, S. P.	-	54 54		- 1	1		16		111	
1	- N. P	1 -		104	70	1		16	32		
ı	- Bay on S. W. side		48	ļ		1				111	
ı	- N.Islet off N.W.side		56		,~	acels	- Drummonds's Island		47	111	22
ļ	Pulo Varela		16	103	47	ğ	- Governor Duncan's		ا ہے	l	أحم
1	Pulo Brala or Capas de	1		1.		Par		16		111	
ı	Terre	4		103	_	~	and the property of	16		111	
1	Pulo Capas de Terre .	5	15	103	8	1	Observation Bank N. P.	16	37	111	
1.	St. Pierre Islands .			108		1	Pyramid Rock	16	35	112	37
1-3	- Ledge of Rocks			108		1	Lincoln Island	16	41	112	
Š	Larkin's Reef		-	109		1	Rocky Island	16	52	112	
	South Haycock Island.	-		109		1	Woody Island	16	50	112	
		, #		1.03	J1	1	Amphiteiteleland TET To	lic	50	113	
Naturas	South Natura's Islands	1	. 1	l	4	1	Amphitritelslands, W.P.	16	54	112	
ž	South Island or Sa-		00		۱ م	1		10	-		
Š	pata			109		1		17		111	
r	— East Island			109			— E. P	17	6	111	32
1	- West Island	-		108					1		, 1
1	- North or Flat Island		3	108	54		Macclesfield Bank, li- 5	15	17	113	
	Low Island	3	00	107	49		mits)	16	21	114	
1	Hutton's Shoal	3		107			Scarborough or Mar-	15	4	117	
ſ	Diana Shoal	3		107		()	singola Shoal, limits	15		117	
j	North Haycock Island		-	107			St. Esprit Shoal (by Lt.	1		1	- 1
1	Grand or Great Natu-		1	108		1	Ross)	19	30	113	6
1				108			— (by Asseveido)	19	6	113	5
ı						()			-	[~# *	-
1	Peaked Island			108		ار ا	Pratas or Prater's Shoal	90	47		ا رو
1	Pyramidal Rocks .			107	Z4	Pratas.	N. E. P.	20	AE I	116	
1	N. W. Island	4		107	0Z	ş	- N. W. P.	20	40	116	
I	— Coral Reef	4		107		1	- Anchorage	20	43	116	
1	— Coral Reef	_		107	47		- Island	20	43	116	
ı	North Natuna Is. S. P.			107		1		21	57	113	44
ı	— N. P	4		108			[The Islands near Can-		١	۱	ļ
1	- Rock above Water .			107			ton are given in No.		ŀ	١	Ì
1	— Saddle Island .			107			XL.and in No. XLVI.	}	ł	1	١
1		4		107		1	Pedro Branco	22	19	115	9
1	Pulo Oby			104	_		Lamock Islands, outer-	1	-	1	٠ ١
1							Lamock Islands, outer-	23	34	117	10
1	The Brothers (eastern)			106			most	20	- 10	1***	47
1	Pulo CONDORE .			106		1.1	A-3-3	1	ì	ŧ	ì
1	Charlotte's Bank .			107	59	8	Andrade Rock (very doubtful)	-		1	_
1	Phæton Bank		00	107	29	8	dowbtful)	9	56	111	4
	Royal Bishop's Bank .	9	44	108	21	(8)	Luconias Shoals .	1	<u> </u>	Ì	_ 1
1	Britto's Bank, N. E. P.	10	32	107	56 ·	١	Luconias Shoals — Hard Rocks — Two Fathern Shoal		94	113	
1	Hollond's Bank, S.W.P.	10	36	108	32	٦	- Two Fathom Shoal		5	113	
1	⊢ N. E. P.	iŏ		108	47	Part	- Dry Sand		57	119	
ı	Pulo SAPATA	10		109	9	ø,	Sea Horse Reef		35	113	
1	Pyramid Rock or Little	٦٠٠	• 1	1	~	E.	— Half Moon Breakers			116	
1		10	2	109	00	S	- Bank Digitized by	10	57	117	
-		140		LAUT	w	9	■ Digitzed by 4	44	كلتات	CIA	

TABLE XLVI. Latitudes and Longitudes.

Lat. D. M.	Long. D. M.	Lat. Lo D. M. D.
	D. M.	Breakers*
Paraquas, 5 or 6 leagues from Palawan 9 10N	117 28E.	Falmouth's (or Essex)
Euphrates Shoal . 5 38	113 24	low Island*) .10 58 112
T		
ditto 5 49	113 2	— Bank, or Gossard's B. 11 25 114 g Essex (or Falmouth) 2 low Island* . 11 2 112
Louisa's Breakers 6 20	113 18	low Island* .11 2 112
Mantannane Isles 6 39	116 7	Gossard's Reef (or Mid-
Barton's Shoals . 6 55	116 6	Gossard's Reef (or Middleburg R.) . 8 58 111 - Small Island 10 42 113
Royal Charlotte's Rocks 6 57	113 38	Small Island 10 42 113
— Sands	114 29	Cornwallis Breakers . 10 00 114
Swallow or Investiga-	*** ~/	6 — ditto
tor's Rocks 7 23	113 44	Sabut Jung low Island 11 32 113
Viper's Bank 7 30	115 7	☑ — Bank
- Breakers 8 00	115 25	Gaspar Shoals 11 36 113
Ardasier's large coral	110 20	South Sea Castle's San-
flats and gaps	1	dy Islands and dan-
- W. P. (Walpole,	1	gers, limits (by Lieut.
Cornwallis and A.) 7 56	113 12	Ross)
- N. E. P. (Walpole		-do
and A.) 7 54	114 24	Two lelands
E. P. (Ardasier) . 7 40	114 47	An Island (Investigator) 11 8 114
- S. P. (Pennsylvania		An Island ditto . 10 44 114
and A.) 7 30	114 84	A Reef 10 15 113
Gloucester Shoal . 7 47	114 50	(10 00
Stag's Shoal 8 24	112 57	Discovery's Reef . \[\begin{pmatrix} 10 & 00 & 113 \\ 10 & 8 & 13 \end{pmatrix} \]
Prince of Wales Bank (8 3	110 24	York Breakers 9 50 117
limits 8 13	110 34	Pennsylvania Breakers 8 17 114
London Breakers . 9 36	112 26	ditto 8 48 15
Reef, western . 8 55	112 00	- ditto (Viper's) . 8 58 115
— do. eastern 8 48	112 24	- ditto 9 4 115
Reef, western 6 55	113 14	1 L. ditto . (10 00)(15
— aitto 7 92	113 8	- ditto (Fanny) 9 45 114
Ganges Breakers 9 22	114 12	1 — ditto
— ditto	115 10	- ditto 9 47 116
Investigator's Shoal W.P 8 5	114 35	ditto 9 5% 110
- E. P 8 10	114 51	- ditto 10 23 116
Shoal 9 12	116 32	- ditto
- Shoal	114 34	XLIV. Islands and Shoals between
Coral Rocks . \$ 9 40 ditto . \$ 9 42	113 4 113 15	via and New Guinea, South of th
Cavallo Marino's Shoal 5 54	114 18	lebes.
1 ##c Q 21	114 21	Lat. Lo
Black Rocks 9 39	114 58	D. M. D.
	115 7	Carimon Java . 5 50S. 110
White Sand 10 48 Low Black Island 11 1	115 13	Lubeck or Babian Island 5 49 112
Low Black Island .11 1	115 17	Arrogant's Shoal . 5 12 113
1 (15.50)	112 34	Madura I. N. W. P 6 53 112 113
I of Friendship's Shoa!	112 49	Pondy Island 6 53 113 Greet Solombo I. (Hill
Hardwick's Reef* (or		Great Solombo I / Hill
Dolphin's) 9 54	112 17	Great Solombo I. (Hill on S. E. P.) 5 38 114
Breaker's* (ditto) 10 2	112 12	Little Solombo I. 5 24 114
log Royal Captain's Shoal 9 4	116 43	Arentes Island 5 10 114
Bombay's Shoal 9 27	116 55	Little Pulo Laut (mid.) 4 51 115
# Dolphin's Reef* (or		Four Brothers, sunken
Hardwicke's) 9 59	112 17	islands 7 00 114
- Breakers* 9 45	112 30	Tirk Island 7 15 115
Breakers* (ditto) . 10 8	112 15	Kangelang or Cangay-
Great Reef, N. P. * 10 7	112 9	ang I. N. P 6 53 115
- Long Island*	112 35	S.P 7 19 115
- Breakers* . 10 22	112 31	- S. E. I. or Hasting's
First Island* 10 35	119 38	Island 6 56 116
Ledge*	112 47	Kalkoon Islands, north-
Breakers* 10 46	1112 47	ern, about 6 10 115
* The longitudes of these p	aces ought	
probably to be increased.		Digitized by GOOGLE
		- 10 and

M m TAB.

Careal Patermoster's Lat. Long.	-					_		_	_		
Great Paternoster's Is. W. P			Lat.			ı	(Lo	ng.
S. W. I. 734 117 16		G4 T 4 4 4	D. M.	υ.	MI.	l.		D.	M.		М.
S. W. I. 734 117 16			~	l	007	8	Bally Island,	_		1	
Section Colorado						К		8	508.	115	2E.
Eastern Island		- 8. W. L				5		8	34	116	24
Eastern Island	E	S. I				ä		8	18	115	43
Eastern Island	ğ	Two low islands .				S	Bally Straits, S. ent.	8	50	114	40
Eastern Island	ž	- E. P	0 42	118	40		A shool man the anches			ı	
Eastern Island	를	Postillion's Jalands, N.				ğ	age at Balambuang,	ŀ		ł	
Content Cont	3	W. P.				ż	bears S. W. W from	ł			
Noesa Sera Islands						1	the flagstaff, distant			1	
Noesa Comba, shout							mile from shore	ŀ		١	
S.Bank off NoesaCombals 46 117 00 Calooship of Rotterdam 15 15 117 36 Hen & Chickens, S. P. 5 28 117 54 Zalinaff or Safanaff or Laer's I.		Noesa Sera Islands .		1	-		Mynder's Rocks	7	41	114	22
S.Bank off NoesaCombals 46 117 00 Calocohig or Rotterdam 5 15 16 17 06 Hen & Chickens, S. P. 5 28 117 54 Lombock I. S. P. about 8 60 116 26 Lombock I. S. P. about 8 60 117						i	Banditti Island	8	46	115	15
Calcochi orRotterdam 5 15 117 36 Hen & Chickens, S. P. 5 28 Zalinaff or Saflanaff or Laer's I 5 31 118 25 25 25 25 25 25 25 2								8	50	116	00
Hen & Chickens, S. P. 5 28 Zalinaff or Saflanaff or Laer's I.				117	36	l					
Laer's I. 5 51 118 25				117	54						
Coral Bank off ditto, S. P		Zalinaff or Saffanaff or	1			ŀ	Lombock I.				
- Coral Bank off ditto, S. P		Lacr's I	5 31	118	25	1		8	13	115	59
S. P		- Coral Bank off ditto,	l	l		l				\	
		S. P	5 54					l		١	
- ditto, W. P			1				Town	8	49	116	33
Five Fathoms Bank 5 52 118 20 Tonyn Islands, S. W. I. 6 31 118 36 Shoal			ŀ				Selonda Island	8			
Tonyn Islands, S. W. I. 6 31 118 36 118 36 118 46 118 36 119 37 119 37 119 37 119 38 117 25 117 25 118 36 119 36 119 36 119 36 119 36 119 36 119 37 117 18 117 18 117 18 118 36		- Five Fathoms Bank	5 52	118	20		Pulo Majo orMayo N.P.	8			
E. I		Tonyn Islands, S. W. I.	5 31	118	36		Flat Island .				
Sample S		— E. I	5 31	118	46	١.	Sandbuy's 4 Shoals. (
Fanakeka or Tunikik I 5 34 119 24 119 24 119 24 119 24 119 24 119 24 119 24 119 26 119 00 119 26 119 00 119 28 110 00				119	5	ĺ	limits				
Brill Shoal, N. P.		Tanakeka or Tunikik I.	5 34	119	24						
S. P. 6 5 119 00 Mansfield Shoal 5 40 120 13 Salayer I. N. P. 6 49 120 28 Tumbora Mountain Biena Bay, rugged point Biena Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 118 36 Stay Bay, anchorage Bill 119 4 Stay Bay, anchorage Bill 119 4 Stay Bay, anchorage Bill 119 4 Stay Bay, anchorage Bill 119 4 Stay Bay, anchor		Brill Shoal, N. P	6 00	119	2		- Timor Yung I (off	1	-	1	
Mansfield Shoal		├─ S. P. '	6 5	119	00		N. W. P.)	8	91	1116	57
Middle Island 5 40 120 28 Salayer I. N. P. 5 49 120 28 Cambyna I. S. P. 5 40 122 30 South Island 5 40 122 30 Hegadis Island 6 13 122 40 Sapy Bay, anchorage 8 30 119 3 118 51 118 5		Mansfield Shoal . •	5 45	120	13		- Sumbayo Roy		·		
Salayer I. N. P.		Middle Island	5 40	120	28		- Tumbora Mountain			1	
Cambyna I. S. P. 5 30 121 1 1		Salayer I. N. P.	5 49				- Biema Raw wasad	ľ	,	1	73
Peak		Cambyna I. S. P	5 30				point	R	11		51
South Island		Peak	5 21	121	1						
Hegadis Island		South Island	5 40			ı	Serv Roy onehomor				
Bouton Island, S. P. 5 42 122 44 — Town . 5 27 122 48 — N. E. Point . 4 23 123 4 — Calansoese Harbour 4 55 123 11 — East Point . 5 15 123 15 — Resey's Islands, . William 1 124 1 — Velthoens or Koko I. 5 58 124 48 — ditto		Hegadis Island	6 13				S F Point	9			
- Town		Bouton Island, S. P.	5 49			į					
- N. E. Point		- Town	5 27			ı	Comodo Island				
Calansoese Harbour 4 55 123 11 S. W. P. about 8 60 121 30 East Point 5 15 123 15 S. P. about 9 00 121 30 Token Bessy's Islands, 124 1 124 1 125 125 125 125 125 125 125 125 Wangiwangi, N.WI. 5 15 123 33 124 48 124 1 125 125 125 125 125 125 125 125 Wanalakjee I. (N. W. Tonin I.) 6 41 120 14 120 14 120 14 120 15 120 28 120 56 125 1							Flores or Manager	0	223	117	37
- East Point							S W D	۵	1		1
Token Bessy's Islands, — Wangiwangi, N.WI. 5 15 — Pinnunko S. P. 6 14 124 1 — Velthoens or Koko I. 5 58 124 48 — ditto		- East Point				ŀ	S P shout	1 -			<u></u>
- Wangiwangi, N.WI. 5 15 123 33 124 124 12		Token Bessy's Islands.	10		10	l	J. F. about			131	3U
- Pinnunko. S. P. 6 14 124 1		- Wangiwangi N.WI	5 15	193	23		N D Di	0	38	1	
- Velthoens or Koko I. 5 58 124 48 - ditto . 6 10 . 7 12		- Pinnunko S. P.	6 14				M. F. Flores Head,				
ditto		- Velthoens or Koko I	5 59				or aron cape	•	Đ	123	3
St. Matthew's Islands, (mid.)		- ditto	6 10		-10		Semile of Fi	_	40		ا ۔
Mamalakjee I. (N. W. 5		St. Matthew's Islands		l			Sandel West 1			123	3
Mamalakjee I. (N. W. Tonin I.)		(mid.)	5 18	124	16		Ring on Tay A. P.			1	
Tonin I.)					.0		Bidli of West P.				
Schiedam Islands, N.W. 7 1 120 28 — S. E		Tonin L)	6 41	120	14		B and				
- S. E		Schiedam Islanda N W	7 1			•	Peden.	10	w	120	30
Shoal		- 8. E.	7 12			١,	ingle Bar	_			
New Island 7 12 121 40 121 36 121 36 122 36 36 36 36 36 36 36		- Shoat	7 27			1	Samu John J	.,	37		
Alfred's Shoal		Kalatoa Island	7 19			i	Non Island				
Pulo Comba or Cambay 123 41 124 6 124		Alfred's Shoel		(. <u> </u>		ı	TAGM TRIBUG	10	4 b	121	3
Core Shoal, about		lagger's Reafor Range-	•		30	•	Puls Carl	_]
- ditto, another estimate		lore Shoal short	7 40	191	12	ł	Tombie	7	49	123	41
mate		- ditto, another coll	- 40	***	10	ŀ	Position I. Peak (on	_		1	(
Angelica's Shoal				101	AC		T. W. P.)				
- ditto, another estimate			7 25			i	E. Point			194	00
mate		- ditto enother	, 22	121	95		rantar I. N. E. P.		10	194	5
Rusa Rafi es Lusardy I. 8 17 121 38 181 18		mote amount esti-	7 40	100	10	ı	East Island, Str. of Aloo		20	124	00
Rusa Linguette or Ro- sagalet I 8 6 122 00 N. W. P 8 9 124 27 - E. P 8 17 126 15		Russ Reff on T	9 17				middle Island, ditto .	8	93	193	55
88galet I. 8 6 122 00 -E. P. 8 17 126 15		Russ Linguists on To	° 1/	121	38	l i	Umbay or Mallao I	_	_ i	1	
1 ho m			ا ہے ہ	102	<u></u> l			8	9	194 9	27
Dastarus 40 14 1122 41 Kotto orRottel. S.W.P. 11 2 1122 65		the Three Partaria				ll		8	17		
	-	. appraised some	0 14	133	41		Motto orRottel. S.W.P.	Li b	,2 _	122)	550

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	- Booca Bay on S. side	10	46 S.	193	20E.		Cape William		34 S.	118	58E.
l	Timor Island, S. W. P.	10	23	123	30		Cape Temoel or Samsa,			1	
ï	- Copang, Fort Con-			l			S. P	0	8 •N		۵£
Temor	cordia	10	9	123			— N. W. P	0		119 119	
E	- Peak		41 94	124 123		ن ا	Cape Donda Cape Rivers	-	15	120	
[]	- N. W. Point	-	12	124			Manado	_	28	125	
	— Tulycaon Bay — Batto-gady		57	124			Cape Coffin	-	42	125	
	- Point nearest Ombay	8	39	125	13		I. Banca		59	125	
	- Dilly or Diely .	8	35	125			Kema Village	-	22	125	
	East end	-	21	127			Castican Bay		48 28	125 123	
	Pulo Batto		16	124	5		Goonong Tella River . Cape Talabo		48 S.		
	Pulo Cambing or Pas- sage I. S. P.	9	21	125	39	ŀ	Weywongy Isl. about .	4	3		
	- N. P.		11	125		ŀ	Waxway leland, mid		34	123	14
	Wetter Island, E. P.		46	126		ĺ	Cambyna I. Peak .		21	122	1
١,	- Pulo Baby, near	l	_				Middle Island		40	120 120	28 9
	8. W, P		5	l			Buele-comba Hill .	9	33	1.20	7
ľ	Geonong apy or Burn-	a	35	126	40		Waller's Shoals and	4	30	117	7
•	ing l		33 41	126	3		Laurel Rocks, limits		37	117	
ı	Kisser Island		00	127	7	ŀ	Noesa Sera Islands .	5		117	9
ĺ	Pulo Jackee or Noosa					ł	Noesa Comba		15	117	9
l	Nessing		19	127		à	Shoal off Noesa Comba			117 115	
ı	Lettee I. W. P		16	127			Little Pulo Laut I. mid. Moresses or Manevessa,	3	51	113	03
	Roma Island	1 -	39	127	30	ខ្ល	Island	4	25	116	3
	Lucapin-hay or Luce- pera I.		40	127	21	2	Dwaalder Island		12	116	
	Turtle Islands, castern	1 -	25	127		9.	Royal George Shoal .		17	116	
l	Cerowa Island, about		10	129	53	ig .	Two Brothers Great Pulo Laut, N.E.P N. P.	4	26	116	
ı	Babber Island, about	1 -	25	130	_	5	Great Pulo Laut, N.E.P	3	21 11	116	41
ľ	Timor Laut, S. P.		15	131		``	— N. P	-	7		
l	Arroe Island, S. ext	1 9	00	135	w	ı	The Three Alike Islands		39	116	54
ı		<u> </u>		<u>' </u>		l	Dry Sand Bank		37	117	48
	XLV. Borneo, Celebes	, L	uconic	ı, wı	in the	ı	Triangle Islands, mid.		3	117	53
	adjacent Islands and as New Guinea.	SM	oaus, c	is ja	r eust	•	Little Paternosters, S.P.		50		20
•	as Jew Guntea.						⊢ E. P ⊢ N. W. P	2	10 8	117 117	
ł	1		Lat. . M.		ong. M.		Pamaroong or Dondre-		0		7.4
1	Tanjong Sambar, SW.P				14E.		kin I. S. P		54	117	36
ı	Succedana .	1	16	110		ł	Seven Islands	0	32	119	43
ı	Tanjong Factie		16	109		1		_			_
ĺ	Pontiana or Lewa, R.en.	. 0	2N.				Banguey Peak		19N.		
ė	Point Mampava		17	168	58		Balambang I. N. Harb. Balabac I. (Hill)		16 59	116 117	
Borneo.	Piner Sambas entrance		4 13	109	3		Mangsee Islands		32	117	
å	River Sambas, entrance Tanjong Apoc		5 5	109		ł	St. Michael's Islands.				
ı	Tanjong Dateo		00	110			(Bangcawang) .	7	48	118	
1	BORNEO Road .		00	114			Toob-Bataba Shoal .		00	119	
ı	Pulo Teega		38	115	7		Palawan, W. end .	.8		117	
I	Abai Harbour		21	116	15		— N. P	11		119 119	
ı	Keeney Balloo Moun-	6	ŏ	116	40		Ragged Island		00	118	
ı	tain Tanjong Sampanman-	۱,	J	***	-20		Soolo Island, Town	-	1	121	
1	gio, N. P.	. 7	1	116	46		Takoot Paboonoowan				•
1	Point Unasang	. 5	17	119			Shoal			121	
1	Point Kanneecongan	. 1	5	119			Pangootaran I.		15	120 122	40 4
1	River Passier, entrance		54 S.				Belawn I. E. P Tapeantana Island, E.P.		00 14	192	
ı	Ragged Peint . Shoal Point .		10 35	116 116			Tamook Island			121	
1	Point Salatan, S. P.		10	114		ä	Mataha Island, S. P.			121	50
1		1 -			-:	ΔŽ	Mataha Island, S. P Pecias I. N. P		41	121	
•	Point Layk, S. W. P.		37	119		ş	Ballook Ballook			121	
1	MACASSAR Town		9	119		ķ	Basilan, I. E. P		30 560	22 Q .:	3U 1
١,	Cape Mandhar .	.13	35	1119	9	<u>_</u>	Santa Cruz I. Digitized by	-th	ear O	71	

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		D,	M.	D.	M.			M	i. <u>D</u>	ID.	λ. Ι
	Sangboys or Hare's Lips	6	48N. 52	121 121	41E		Group of Islands, S. P.	12	8N.	120	23E.
	Teynga Island	•	52	121	43	١.	N. P. Turret Island	12 12	17	١	
i .	Catanduanes I. S. P.	13	38	124	16	ŀ	North Rock	12	97	120 120	
	C. del Espiritu Santo.					6	Mindoro I. S. P.	12	11	121	4
	N. E. P. Samur I.	12	40	125	38	Mindoro	- Point Dongan or	ı		- -	~~
	St. Bernardino Island .			124	38	ŭ	Pandan .	12	48	190	58
	Ticao I. Port St. Jacinto	12	34	123	34	₹	Point Calavite .	113	27	120	
						ŀ	Luban	13	44	120	
	Manilla	14	36	121	2	ļ	Goat Island	13	51	120	7
,	Cavite	14	29	120	55		Babuyan Islands	l		1	Ì
nţa	Ent. Manilla Bay Point Capones Two Sisters Islands Point Boliano	14	20				- Lapurip or Dalu-	١			
8	Total Capones	15	5 2	120	3		peril	19	15	121	34
10	Point Boliano	16	97	119 120		ŀ	- Fuga or New Babu-	١.۵	. 1		40
		18	49	121			yan I — Camiguin I	19 19		121 122	
	Point Cavnaien	18	48	121		١.	Babuyan Islands,	7	7 (122	1.7
ľ	Cape Enganno	18	39	122			- Guinapac Rocks .	19	5	122	25
	Mauban	14	8	121			- Didicas Rocks			122	
	Cape St Ildefonso .	15	25	121	46	ŀ	- Claro (or Old) Ba-				
		١.		l			buyan	119	37	122	17
ö			5 3	122			— Calayan I	19	25	121	46
na m	Point Balagonan	7	51	122	24		Bushee Islands	ı		1	
à	Point Balagonan . Suriago village, near N. Point . Cape St. Augustine, S. E. P. South or Serangi Point	۱.	411				- Balintang or Rich-	١. ـ			
ij	Point	7	47	125	25	ŀ	mond isles	19	58	122	
٠.	r b	6	4	126	40	ė	- Sabtang I	20	14 14	155	
	South or Serangi Point	5	39	125	20	Bashee Islands.	- Goat I.	ZU	15	122 122	9
H	Mindanao	7	10	124	35	751	Batan or Monmouth	ZU	13	122	•
		1	••		•	2	I. S. P.		17	122	15
1	Negros, South Point .	9	6	123	3	48	- ditto Mount, N. P.	20	23	192	
	Point Sojoton .	9	50	122	32	æ	- Graston or High	Γ			
	Cagayanes Islands, mid.	9	34	121	23		Round 1	20	34	122	13
ے ا	Panay I. Point Nasog,				_		- Bayat or Orange I.	20	37	192	7
9	S. P		25	122	6		North Bashee, High I.	21	3	122	8
Ponay	- Asloman village .	10		122			- Northernmost I	21		122	8
	- I OHICE OLOD OF IVEL		48 24	122			Gadd's Reef	31	43	121	43
		10		121 121			Cumbrian's Reef, doubt-	1			
	Sombrero Rock	10	28	121			ful, probably the same as Gadd's Reef	١.	9.	L_	4.
	Cuyos Islands			.~.	21		Little Botel Tobago	21	33	131	40
	- Quiniluban (North-			1				91	56	121	45
	ern Island)	11	28	121	11			21		121	
	- Grand Cuvo	10	52	121			Vele-rete Rocks	21	42	120	
	— Southern I	10	40	121			Formosa I. South Cape	21	54	191	
				121				1			
	Betsey's Bank, 5 fathoms	11	42	120	57	l	Gomano Island		56 S.		
	Ylin Islands, S. P. off S. P. Mindoro	12	٥	10.			Lissamatula I. S. E. P.		46	1::6	
1	Coral Shoal, West of		9	121	19		Lulie Bessey, S. E. P.		29	125	58
ı			11	120	57	2	— N. E. P		58		40
ı			36	120		Ş.	Xulla Mangola, W. end	1	58	125	
نيا	— E. P	12	40	120		šes	1		43 40	125	
Ba. F.	- N. P	12		120		a B	Greyhound Straits }		56	124	30
۳	S W D (Tales)	110	An .	120	29	Xulla	Haycock I. off S. W. P.	-	30	l	
12	- West, or Great Islet		39	120		×		1	58	124	36
1,2			40	120			Skelton's Island, on N.]			
1			46	120		ĺ	W. P. ditto		45	194	36
1			3	119			Middle Island			194	96
			9	119		1	Albion's Island			124	
ľ			18	119		•	Bouro Island, N. W. P.		6	125	57
1	Sail Rock		23 22	119		•	- N. ext		. 2		_
1	n			119 119			- N. E. P.	3	15	127	
1	Calavite or High I.	12	21	119			— Cajeli or Bouro Bay — South Point			197	•
L		1		١.,		ı	Ambiaw Island		49	107	1.4
		_					. DUBIGL WAILURE	0	58	197	14

Manipa Island	-		_		-		_			,	
Manipa Island	1						l	!	Lat.		
Bonos L about	1	34					,	m: 1 3.64 *-			
Section Sect	I						ł i				
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Loeuwarden Island				53	129	49	ŀ			198	95
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Matabella Islands	2	Ob I					ŀ				
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Noesa Laut I. 3 40 128 52				40	128	15	ı		-	"	
Lookisong or Landscape I. S. P. 1 46 198 10 Pulo Gasses, S. P. 1 41 198 20 Kekik 1 33 198 37 Pulo Pissang 1 23 198 53 Horsburg's Rocks 1 8 198 20 Boo Islands 1 12 199 18 Weeda Islands 0 40 128 25 Kawary Islands, Grand Cl 1 44 199 42 Pulo Popo, S. E. P. 1 12 199 52 Battanta I. Cape Cambo, W. P. 0 56 130 23 Waygecoo I. S. E. P. 0 7 Point Pigot 0 56 130 23 Waygecoo I. S. E. P. 0 7 Point Pigot 0 131 12 199 13 Kools. 0 50 131 12 131 18 Mir's Bay 122 27 114 30 131 14 Mir's Bay 122 27 114 30 131 14 Mir's Bay 122 27 114 30 131 14 Mir's Bay 122 27 114 30 131 14 Mir's Bay 132 27 114 30 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 132 14 Mir's Bay 133 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay 134 14 Mir's Bay	1	Noesa Laut I	3	40	128	52	ı			125	44
I. S. P	1	Banda Island, anchorage	4	31	130	00	ł	— N. end	3 46	125	38
Pulo Gasses, S. P. 1 41 128 20	1						l	Glatton's Rock	3 50	125	56
Kekik	1		1	46	128	10		Sallibebo or Toulor Isls.	.[ı	
Pulo Pisang	1						1			126	55
Pulo Pisang					128	37	1	— Tulour or Karkalang			
Serangi islands, S.P. 5 20 125 31 125 32										126	44
Weeda islands							1		5 00	127	17
Kanary Islands, Grand C	1						ı		5 20		
Pulo Popo, S. E. P. 1 12 129 52 129 52 130 23 141 14 141 115	1		•				ł		i	125	32
Pulo Popo, S. E. P. 0 56 130 25 131 18 0 56 130 25 130 25 131 18 1					129	42	ľ	— N. P	5 31	125	43
Battanta I. Cape Cambo, W. P. 0 56 130 25							ŀ	XI.VI From CAN	TON H	V	a MT
Do. N. P. 0 56 130 25 Fisher's Island 0 56 130 23 Waygecooe I. S. E. P. or Point Pigot 0 21 131 18 Offak Harbour 0 00 130 50 Mir's Bay 22 27 114 30 Mir's Bay 22 31 115 Mir's Bay 22 31 115 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 116 30 Mir's Bay 22 31 Mir's Bay 22 31 Mir's Bay 22 31 Mir's Bay 22 31 Mir's Bay 22 31 M	1		1	12	129	52	Ł	SKATKA with the a	diacent l	olana olana	204 I
Pisher's Island 0 0 130 23 23 23 23 23 23 23	1		٦		İ		ì	Shoals.		********	
Waygecooe I. S. E. P. or Point Pigot . 0 21 131 18	1		-				ì				
or Point Pigot	1		v	56	130	23	ĺ	ļ ,			ong.
- Offak Harbour	1		٦	٥.			ı	CANDON			
Amsterdam I.	i i						l		23 IN		
Amsterdam I.	1	D 10 1					ı	Single Island on Channel	22 27	114	30
Fow or Faux Island	1						l	Single raising of Cuttens	00 05		40
Geby I. N. W. end . 0 4N 139 19	1			_			ł	Mendozale Island	20 21		
Geby I. N. W. end O 4N 139 19 Syang I. O 22 129 55 Eye Island O 24 129 55 Islet E. of Pulo Moar O 39 128 58 Catharine's Ialands O 39 129 11 Canton Packet Shoal O 35 128 55 O 46 Ditto soundings 15 fath Yowl or Aiou Islands O 35 131 00 End of the Island O 36 Island O 38 Island O 38 Island O 38 Island O 38 Island O 38 Island O 36 Island O 36 Island O 36 Island O 38 Island O	1						1	Fokoi Point	20 33		
Syang I.	1							Pedro Branco	99 19		
Canton Packet Shoal . 0 35 128 55 Breaker Point	1	Swann 1						Point Shalong-Tow	22 39		
Canton Packet Shoal . 0 35 128 55 Breaker Point	1						4	Point Tengmee	22 45		
Canton Packet Shoal . 0 35 128 55 Breaker Point	1		1 -				.5	Point Cun-chee-san	22 49		
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Ditto soundings 15 fath O 42 130 3 Yowl or Aiou Islands O 35 131 00 Lamo or Name I. W. P. 23 28 117 16 66 N. E. Island O 36 131 15 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 24 11 119 20 Lamock Island, S. west. 25 117 46 Lamock Island, S. west. 26 117 40 Lamock Island, S. west. 27 117 21 Lamock Island, S. west. 28 17 117 21 Lamock Island, S. west. 24 11 16 60 Lamo or Name I. W. P. 23 28 117 10 Lamock Island, S. west. 25 117 10 Lamock Island, S. west. 26 117 40 Lamock Island, S. west. 27 117 21 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 117 11 118 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. west. 28 117 10 Lamock Island, S. w	1	Canton Packet Shoal .	0	35			ı	Breaker Point	22 57		
Ditto soundings 15 fath 0 42 130 3 Yowl or Aiou Islands 0 25 131 00	1				1		1	Cape of Good Hone	23 14		
Yowl or Aiou Islands	1				130	3	ı	Small round Island	23 26	1	
Aiou, the largest Isle	1				1		1				
- N. W. Island . 0 38 131 8 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 117 21 Lamock Island, S. west. 23 17 Lamock Island, S. west. 24 11 Lamock Island, S. west. 24 11 Lamock Island, S. west. 24 11 Lamock Island, S. west. 24 11 Lamock Island, S. west. 24 11 Lamock Island, S. west. 24 11 Lamock Island, S. Lamock Island, S. Lamock Island, S. Lamock Island, S. Lamock Island, S. Lamock Island, S. Lamock Island, S. Lamock Island, S.	1	- Aiou, the largest Isle	0	25	131	00	ı				
- N. E. Island	1	— N. W. Island .			131	8	1				
	1		0	36	131	15	ı	The Brothers, southern	23 32		
Asia's Islands, S.W. Isle N. E. Island 1 04 131 23 Chin-Chew Bay Chin-	}	- Reef North part .			l		1	Chapel Island	24 11		
Chin-Chew Bay 24 54 119 40	ŧ	Asia's Islands, S.W. Isle	1	00	131	17	1	Amoy Harbour	24 20		
Gillolo I. N. end 2 23 0 45 128 22 Maba Village 0 53 Islet near Pulo Moar 0 9 129 58 Point ent. Straits Patientia 0 13S. 137 45 Coccoanut Pt. or S. P. 0 45 Batchian I. S. E. P 0 48 128 3 Amsterdam Island 0 20 127 53 N. P 0 7N. Negory Kalam, N. P 0 28 127 67 N. P 37 25 122 45 Wolf Rock 0 30 127 67 Tehoo-san Island 37 36 121 28 129 29 129 29 12	Ī		1	4	131	23	1	Chin-Chew Bay .	24 54		
- Ossa Village . 0 45 128 22 Ting-hoy harbour . 26 10 119 57 128 128 128 128 2	•	Gillolo I. N. end .	12	23	l		ł	Lamyet Islands, south .	24 59		
- Maba Village . 0 53	1	- Ossa Village	0	45	128	22	1	Ting-hoy harbour .	26 10	119	57
Point ent. Straits Patientia	1	Maba Village .	10	53	ĺ		•	Heysan or Black Islands	28 53	1	
tientia 0 13S. 127 45 — Coccount Pt. or S. P. 0 45 Batchian I. S. E. P 0 48 128 3 Amsterdam Island . 0 20 127 53 Kayo or Cayo I. S. P 0 1 127 23 — N. P	ł			9	125	58	ł			121	52
tientia	ł						•			121	42
N. P 0 7N. Negory Kalam, N. P 0 28 127 37 Ten-choo-Foo City	1.	tientia.			137	45	1			119	00
N. P 0 7N. Negory Kalam, N. P 0 28 127 37 Ten-choo-Foo City	13	- Cocoanut Pt. or S. P.				_	ı		30 20	122	36
N. P 0 7N. Negory Kalam, N. P 0 28 127 37 Ten-choo-Foo City	ž	Batchian I. S. E. P.					1		L	1	
N. P 0 7N. Negory Kalam, N. P 0 28 127 37 Ten-choo-Foo City	Z	Amsterdam Island .	4 :				ı				
Negory Kalam, N.P. 0 28 127 37 Ten-choo-Foo City . 37 48 120 32 Wolf Rock 9 30 127 6 Tehoo-san Island 38 1 121 1						23	ı				
Wolf Rock 0 30 127 6 Tchoo-san Island 39 71 2121 1	1	- N. P					ı				
	1						ı				32
I lidore island, 3. ext1 U 34 1127 24 I UKeusen Is. northern .138 8 1190 44	ı						1				1
130 12	Ł	indore island, 3. ext	Įυ	34	1127	24	<u> </u>	Reusen is. northern .	38 8	120	44

_		_	o.t	-	ope (-			at.		
1	۱ ،		at. M.		ong. M.	١.			M.	D.	obg M.
		υ.	MI.	ש.	IVI.						00E.
	Pekin River, anchorage	70	ron.	. 10	00E.			25		121	6
	at Peiho	30	9214.	110	WE.		NT Dains		18	121	-
1	High Peaked Island, S.	34	5	125	15		N. E. Point	25	11	121	
	W. ext. Corea	3 6		129		l	Lamay Island	22	10	120	
	la *	37		128			Pehoe or Pescadore Is.	~	17	1.20	•
ı	l	45	12	137		i		23	8		
. 1	Ternai Bay	47	51				- High Island, S. W.	~	٠		
Ę	Suffren Bay	47	DI To	139			fligh thiand, co. vv.	23	14	119	96
ž	Cape Lessops	49	3U	141				23 23	19	119	
F	Castrie's Bay	51		141			- Largest Island .	30	32 54	1177	40
F	Vanjuas Point	52	7	142		l		23			92
	Bay de Langle	17	49	141				22		119	
	Bay d'Estaing	48	59	141		ŀ	- ditto, seven fathoms	22	DI	119	ı
1	Monneron Island	46	20	141			Pat-chow or Madjicose-	l	- 1		
	La'Dangereuse Rock .	45	47	142	9.		mah Islands .	١	_ {		
		l		l	1			24	6	123	53
	Cape Crillon, (ent. Pe-	l					- Bluff Point. West	١	1		
	rouse's straits) .	45	54	141	59		ext. Great I	24	17	123	
l	Cane Aniwa	46	2	143	30	ľ	Kumi I	24	28	193	5
	Cape Lowenorn	46	23	143	40	ŀ	— Eastern I.Ty-pin-san	24		125	36
	Bay Mordwinoff .	46	48	143		ŀ		25	6	125	11
1	Cane Tonyn	46	50	143			Lieu-Chew Islands .			l	
	Point Siniavin .	47	16	143		ŀ	-Great Lieu-Chew, S.e.	26	3	128	18
	Mount Spenhers or Rev.	-			ا ٽ	l	- ditto, adjacent I.N.P.	27	34	-	
1	Mount Spenberg or Ber-	47	33	142	90		Western laland	196	20	127	17
ł į	nizet	47	50	142		ŀ	Hoapinsu I.	25	44	123	
		41	50	140	44		Two or which I	25	55	123	
	Cape Alexander Dal	1.5	91	142	50		Ty-ao-yu-au I Sulphur Isis. southern			141	
		48					Sulpant raise southern	24	48	141	
		48		143	2			24 25	1.4	141	
Ĭ	River Neva, entrance .			143	2		— northern	20	3U 1.2	123	
	Gulf Patience, North P.	49	19	l			Group of four Islands, }	27	30 40		
	Robber Island Reef N.					ŀ	limits)	27	40	128	30
	E. P	48		144			L	_	40		_
	_ S. W. P	48	28	144				29		130	5
	Cane Patience	48	52	144			Ormsbee's Peak	29	40	140	
١,١	Cape Billinghausen .	49	35	144	26	l	A rock	30	45	123	
Ę	Mount Tiers	50	3	143	37	Į	South Island	31	30	140	
ğ	Cape Ratmanoff	50	48	143	53		Gotto I. S. end	32	38	128	
g	Cane Crovere	51	00	143			Asses Ears	32	3	198	
Š	iiiowns roint	51	53	143			Quelpaert I. S. P	33	8	126	19
	Shoal .	52	30	143		Ì	Kinsin Island	1		1	
	Wurst Point	52	57	143		Ĺ	- Cape Tschirikoff .	32	14	131	41
1	Cape Klokatschef	53	40	143		ŀ	- Cone Danville	31	27	131	27
ı	Cape Klokutscher	20	3	143		ŀ	- Cape Nagaeff .	31	15	131	
1	Cape Lowenstern .		5 <i>4</i>	142			- Mount Schubert .	31	41	131	
1	NT 41 D	54	24 1 C	1			- Mount Horner, Peak		3	130	
1	North Bay	54	10	143			- Cape Tschitschagoff				
		54	17	142			S. P	30	57	130	36
1		54		142		١.	— Cape Tachesma W.P.	21	94	130	
1	Cape Golowgtscheff .	53	30	141	55	Islands.	- Cape I schesma vv.s.	31	372 40	130	7
1				١	1	Ę		31			-
1	Cape Romberg	53	26	141		75]	- Mount Unga, volcano	21	4 3	130	14
	Cape Chavaroff	53	39	141	26		— Nangasky harbour	_			16
	Jonas Island	56	25	143	16	Japan	ent		44	199	40
	Ochotsk	59	20	143	19	Ę	— Cape Nomo S. P. of	1			
l	Vamale	60	46	154		7		32	3 5	199	
1	Bolcheretsk			156		ŀ	- Cape Seurote .	32	58	129	
ł	Cape Lopatka, Kamt-	-					Sanao-sima Island N.P.	30	49	131	00
		51	2	156	46	l	- S. P.	30	24	l	
13	skatka .			158		ı	Tenegasima I. (middle)			130	30
Russia.		53				l	Volcano I	30	43	130	
r		53		159		İ		30		130	
ı		56		162						130	
1	Cape Tschulkolskoi .	64			24W	l		30			
	East Cape	66	6	169				30		130	
	Cape Serdze Kamen .	67	3	171	49	ŀ	St. Claire I.	30		129	
1		68	56	179	9	ŀ	Symplegados Islands N.		1 Ο0	129	ξ
=					•	•	E. P	31	30 1	127	18

7	•		at.	T.	ong.	_		, ,	at.	7	011.0
			M.		M.		l *	D.			ong.
	Symplegados Islands, S.		***	۳.	414.		Raakok Island		M.	142	M.
	V. P.	21	26N.	190	37 P			48	ON.		15E.
	Meac-Sima I. S. W. P.	21	25					48		153	
		31		129		ľ		48		153	
				199			Charamukatan I. Peak			154	
		31		129				50		155	
		34		129			Peak Fuss (S. W.P.)			155	10
	— Cape Fida-Buengono			129			— East Point	150	28	156	9
	- North Point	34		129			XLVII. NEW HOLL	AN	תי	42	
1	Coinett's Island	84		129	56	l	jacent Islands a				e aa-
ŀ	Dagelet I. N. E. P	37	25	131	22	į.	jeceni Isianas a	na i	Snoais	·.	
	Niphon I			1		į		١Ī	aL	L	ong.
Islands.	- South Point	38	5	135	55	i		D.	M.	D.	M.
Ĭ.	- Cape Noto	37	36	137	54	ı	Pedro Branco (Rubrick)	44	008.	147	45E
3	A Rock	37	36	137	34	ı	South West Cape .	43	34	146	6
		37	51	137		ł	Mew Stone	43	42	146	
Japan	— Jedo			140			South Cape	43	38	146	
3	- Cape Kennis	37		141	- : :	•		43		147	8
7	— Zach's Mountain .			132		Ь		43		147	9
1	- Russian's Promonto-				20	ŀ		43		147	-
li	ry, S. P.	39	46	139	44	Ŀ				147	-
ı	NI IS D	40		1.09	77	ş	D'Entrecasteaux's Port Adventure Bay	70	01		6
1		40		140	e		Adventure Bay	43	21	147	
				140		E	Frederick Henry Bay .	42	25	147	
	Cape Gamally .	40	35	139		82	Cape Pillar	43	12	148	9
		40		140			Oyster Bay	42	42	148	8
H		41	9	140	8	å	St. Patrick's Head .	41	48	148	22
li	Cape Sangar (ent.						Cape Portland	40	49	148	15
1	Straits of Sungar)	41		140		£	Port Dalrymple	41	3	147	11
ł	Osima Island	41	31	139	19	4	Circular Head	40	43	146	15
	Kosima Island	41	21	139	46		Cape Grim, N.W.Prom.	40	41	144	46
	Okosir Island (middle)	42	9	139	30		West Cape or Sandy Pt.			144	
	Jesso Island	1					Macquarie's Harbour			145	
	Cape Nadeshda (ent. of			ŀ				43		145	
		41	25	140	9	ı	Port Davey	43	98	146	
		41		133			Total Davidy	P-0	20	730	~
ı	— Matzumay town .	41	39	140			Ent. to Bank's Straits .	مدا	20	148	an i
	— Cape Oota-Nisawu	40	18	139		L		Γ-ν	•	140	20
1	Cope Vutusoff	42	20	140		į.	Furneaux Islands, .	40	05		00
1		42			.1	ľ	- Barren I. S. E. ext.			148	20
ll				140			Clarke's I. S. ext.	40	34	1	
. 1	Cape Okamay, S. P.			140			- N. Sister, near N. P.				
	— Cape Taka-sima .			140		l		39	38	147	
		42		141		i	Endeavour Rock .	39	39	147	
ı		43		141		ŀ	Kent's Group .	39	29	147	17
		44		141		ł	The Pyramid	39	43	147	13
		44	00	141	54			40		147	32
	— Cape Romanzoff (N.			I			Hunter's Islands	ı		l	
	Point)		26	141	84		- Black Pyramid, W.	40	33	144	22
	— Cape Soya	45	31	141	51	ł	- Albatross I. N. West			144	
		45	21	142			King's Island N. E. P.			143	
	Peak de Langle, Rio-	1					West ent. Bass's Straits			143	
	schery I	45	11	141	12			`	-	. =0	-
	Cape Guibert, Reifun-			**	-~		Cane Albany Oteran	38	52	143	20
	schery N. E. P.	45	98	141	4		Cape Albany Otway .				
	Jeurire Island	44	28			ı	Port Philip, ent	38	1	144	
		44		141			Western Port	38	aı	145	10
	Janiecocki iskud .	[**	47	1-21	22	l	Wilson's Promontory,	-			ا بـ
	States Island S IF and	اء ا	o٤	1.4-		I	S. ext.	39	11	146	
i	Staten Island, S.W. end	44	20	147	28		Ram Head	37	89	149	45
	- Cape Vries, (Vries	١					Cape Howe	37	30	150	7
	Straits)	45	26	149	43	H	Cape Dromedary .	36	18	150	
ĺ	Company's Island .			ŀ				35		150	
ŀ	- Cape Shouten .	46	18	150	58			34		151	
ĺ	- North end	46	28 '	151			Botany Bay. entr. (C.	- '	- 1		. 1
ĺ	Marikan I. N. end .		10	163		å	Banks)	34	00 l	15 T	23
l	South end (Bousole	Ī				0		33		151	
l	Straits)	46	46	152	39	Ę	Port Jackson, ent. Broken Bay	33		151	
ı		48		153		<u>م</u>	Post Stanbane	32			
'	·······] tocaton 1: 1 CGA ·	-		-53	12	_	Port Stephens		PY()!	159	10
									-	5	

_						_
		Lat.	Long.	1	Lat	Long.
. 1		D. M.	D. M.		D. M.	D. M.
	Cape Hawke		152 80E.			160 39E
	Smoky Cape	30 51	158 7			147 OU
	5 PA 7 January 5	30 9	153 21		31 30	140 00
ľ	Solitary Islands . }	29 56	1 1	Week's Reef 36' N. E.		1 1
l	Cape Byron	28 7	153 30	and S. W	81 15	153 9
	Point Danger	28 7	153 30	listand	31 00	147 6
	Shoals off ditto	28 7	153 39	Ganges Island		154 25
		27 1	158 23	Bank of Soundings	80 30	177 30
3	Cape Morton Shoal	26 58	153 28	Island	30 00	137 (10
4	Sandy Point		153 17	Island	80 00	139 00
	Sandy Point Cape Capricorn Keppel Bay Regries Reef S. P.		151 00	Island	30 nn	141 30
7	Karnel Ray		150 36	Island	80 nn	143 00
3	Barrier Reef, S. P.		152 36	Island	30 00	144 94
_	Defiter recoil at a .		150 11	Rica de Oro		157 3
Į	Cape Townsend		149 00	Island		143 00
•	Cape Palmerston	21 00	148 33	Island	29 83	187 00
ı		20 32	148 30			143 00
	Cape Conway					175 45
	Cape Gloucester .	19 38	148 6	0.1		
	Cape Cleveland		146 40			158 00
			146 8	ditto(another account)	20 53	168 00
ĺ	Cape Grafton		145 54			176 50
١.	Cape Flattery	14 52	145 18			175 44
	Cape York	10 38	149 33	Disappointment Island		139 25
ŀ	New Year's Island .		133 18	1 1		142 48
ŀ	Vandieman's Cape .	11 12	129 54	Bassiosos I		178 27
	Red Island, off P. Vul-			Island	3 6 6	164 36
	can	15 9	124 22	Reef	26 6	160 00
		17 14	118 57	Copper Island	26 00	131 48
1			114 40	Tree Island	26 00	145 44
	Clarke's Reef N. of			Lasker's Island	26 00	173 94
	Rosemary I	20 17	1	Island	25 58	181 17
	Eastern Rosemary I.	120,22	1	Island	25 42	131 18
	N P D	20 26	1	Reef		152 50
	N. E. P	20 35	115 40			139 00
	Part 46-1 Charles	21 37	112 25	North Island	25 14	141 14
				Televil	25 12	131 86
			114 56	Island		146 00
	Shoel (land of N. Hol-		!	Grampus Island .	25 10	141 29
	hand in sight from the		<u> </u>	Sulphur Island		
	mast head)	20 15		Kendrick's Rock	24 30	183 36
		21 50	114 25	Marcus Island	24 18	153 42
	Dirk Hartog's Road,	1		Weeks Island	84 00	154 00
	ent. to Shark's Bay .	25 6	113 15	Dexter's Island	23 24	163 5
ì	Houtman's or Abrohlos			Island	23 3	162 57
	Shoals		113 85	Reef	28 6	142 28
	Rottenest Island	31 58	114 24	Jardines	21 35	151 30
	Cape Leuwen or S. W.			Parel or Pern I.	21 10	141 40
	Cane	34 22	115 6	Abregoes Shoal	21 1	136 43
	Cape Chatham		116 22	Reef	20 42	158 00
	Cape Howe	35 9	117 88	Douglas Reef	20 32	136 13
	K. George III. Harbour		118 1	Lamire I	20.30	166 42
	Point Hood	34 23	119 36	Island		152 50
			121 58	Bishop's Rock	90 16	186 53
		36 27	127 2	Week's or Wilson's I.	19 21	166 55
ŀ			135 45			165 49
l	Port Lincoln		137 56	Welcons T		163 33
•			10/ 30	Haleyon I	10 00	
	Endeavour Shoal, off		.00 01	Folger's I	19 22	155 15
			139 31			156 13
ŀ	XLVIII. Islands, Rock	e and She	als in the			160 00
	NORTH PACIFI	C OCF	A.N		16 36	169 43
				Island	16 00	171 49
		Lat.	Long.			
	47	D. M.	D. M.	Pajares Islet, northern.		145 48
l	Alcootskia I.	1	1	Urracas, about	20 20	146 15
١.			170 42E.	Assumption Island .	19.45	145 35
	- Consission		166 22W	Almagan Islandigetzed be		146 21
_	Bank (64 fathoms)		178 30E.			0
_						

1			⊿at.	1.0	200			-	- 4		
1			M.	D.	M.				Lat.	FOI	ng. M.
1			47N	146	171.	•	lohanna	υ.	M.	υ.	M.
ı			00			1	Johannes	Ď	55N.		
l				145		1	Lion's Island	5	16	132	
1	Guam, Umatac Bay .	13	21	144	20	ı	St. Andrew's Island	5	20	132	
1	Della della della AVIII I			i		l	Pulo Anna	4	38	132	
ı	Radack chain of Islands,	1				•	Pulo Mariere			132	28
1	viz.					ł	Lord North's L		3	131	
•	Aour, circular group of		٠,			l	Ganges Shoal, S. W. P.			131	7
1	32 islands extending			ŀ		l	← N. E. P	3	6	131	23
1	13 miles N. W., and		_	ŀ			Helen's Shoal	3	50	131	41
1	S. E. anchorage .		19	171	12		Freewill orSt. David's } Islands, limits . }	0	49	134	17
ı	Kaven group 33 miles			İ			Islands, limits .	1	2	134	30
1	N. W. and S. E.	1				ı	, ' '	1			
1	— Araksheef Island,			Ì		l	 ,			l	
	(largest I.)	8	54	170	49	1	Pelew Islands,	l		l	
•	 Southern Island . 	8	29	171	11	1	- Baubelthouap, E. P.	1 7	41	134	55
ł,	Tchitchagoff, circular					1	- Northernmost, Ky-				-
	group of Islands N.W.					ı	angle	8	8	134	50
	& S.E. 24 miles, mid-					ı	- Large Reef, part dry		18	134	
•	dle	9	6	170	4	8	- Southernmost, An-			1.0.	
1	Romanzoff, circular	ł			_	Islands.	gour		53	134	91
1	group of 65 islands,					150	Matelotes, N. E. I.		34	137	
1	E. &W. 30 miles. & 10	i		1			- Southernmost		19	137	
1	miles wide inclosing a					Caroline	Yap or Hunter's I.N.P.	۱۵	40	101	40
ł	sca 12 miles wide &					Ē	- S. P.	13	20		
1	27 miles long					Š	Philip Islands		30	138	
ı	- Odia I. eastern; an-						Thinks There	9		140	3
	chorage	a	28	170	16	9	Thirteen Islands .		18	14:	
١.	Legiep or Hayden group			169			Haweis' Island	1 '	30	1 16	
13	Ailou group 15 miles	1	31	107	13	1,0	Strong's Island	1 2	12	162	
Ē	Ailou group, 15 miles					7	Islands	5	23	153	
3	long, 5 miles wide					d	Islands	. 15	47	157	
×	- Krusenstern-Capeni-	10	07		00	æ	islands	6	9	160	51
Ę	Legiep or Hayden group Ailou group, 15 miles long, 5 miles wide — Krusenstern-Capeni- us I. (northern) I. Du Nouvel An	10	2/	170		Š	Islands		17	159	
ΙŽ	I. Du Nouvel An	10	8	170	55	•	Hope's Islands		15	165	12
 	wareson of Calrick group		i	1			Baring's Islands		35	168	13
1	separated by a chan-					ŀ	Palmyra Island		49	162	29
	nel from a southern						Cluster of Islands .		38	161	26
	group called Souvor-						Ditto . • .	9	55	ŀ	
1	off or Tagay, extend-						Brown's Range				
1	ing N. & S. 25 miles		. 1				- Arthur's Island, N	11	43	162	42
		11	11	169	50				19	162	
1	Group north of Kutosoff			ı					52	166	
	Mille		16				Lydia's Island	9		165	
	Medjuro , . .		15						14	166	
1	Arno		25				Arrecife's Island		36	161	8
1	Bigar, south of Kutosoff	11	40	l		•	Muskitto Group, low)	1 7	20	168	
ł				1			Muskitto Group, low } and dangerous . }	7	47	1-03	~~
1	Pescadores Is. southern	11	00	167	30			8	54	166	35
ı	northern	11	20	167	2		Peterson's Island	a	20	171	
1							Reef			179	
1	Ralick chain of Islands,	l		1		ı	Calvert's Islands	٥	48	172	
	extend nearly N. &	1		1		l	Ibbetson's Islands .	8	6		
	S. about one degree	1		l		1				172	8
1	west of the Radack	l		ĺ			Mulgrave's Islands			168	
1	chain, viz.	l		1		l	Ronhom's Island	3	54	172	
1:		5	50	167	15	ĺ			50	169	
Ę	- Nosmureck I	5	30	1.07	10		Cook's Island Hall's Island		18	171	
13	Kuli group		40			1			54	173	
			30	l		1	Reef	1		179	
18	Odia group		30 15	1		l	Pitt's Island	2		174	
F	Odia group Namou group		00			l	Matthew's Island .		50	175	
æ	T ital Talama			1		•	Simpson's Island .		26	175	
1	— Littei island		55	I		1	Macasgill's Islands .		12	160	
1	— Tebot Island		30	l		ı	St. Bartholomew .		10	163	48
1	Quadelon group		20	ļ			Cornwallis or Smyth's			1	
1	Oudia-Milai group .		45	1			Isles		46	169	29
1	Radogala group		00			ı			00	166	
Ļ	Bigini (northern) .	11	20	167	15				(24)	166	

Nn Tab.

~			- t	110	ne.	_				_	
1			Lat. M.	D.	ng. M.	l	}	La D.		1 10	ng. M.
1	Gasper Island	15	3N.	177	00E.		Massachusetts Island .	22 2		177	5W
	Gasper Rico I	14	48	169		ı	Island	24	4	168	00
ı	Wake's Rocks	17	48	173		1	Henderson Island .	24	6	128	30
ı	St. Peter	11 8	3	179 178	55W	1	- another account .	24 9			
1		22		175		1	Gardner's Reef Pollard's Island	24 1 24 4		168 168	9
l		23		164		i	Allen's Reef	25 0	90	167	
ł	French Frigate's Shoal	23	45	165	50	ı	Cooper's Island	25	4	131	
ı		26		173	40	ı	Maro's Reef	25 2		170	
	O. J. D	20	4.77			1	Island	25 2	2	131	
ı		20 19	3/	155 154		ı	A Rock Laysan's Island	25 3 25 5	50	174 171	3
ı		18		155		ı	Liscanskey's Island .	25 5	12	173	
ŀ	- Karakakoa Bav .	19	28	155		ı	Neva Island	26	5	172	
1	Mowee, E. point	20	50	155		ı	Maro'sReef (dangerous)	26	6	170	24
•	- S. Point	20	34	156			Island & Rock	26 Q	4	170	
1	- W. point · ·	20	25	156 156			Pearl& Kermes group or	27 4	6	176	15
ż	Ranai, S. point	20	46	156		ĺ	Clarke's Reef 60 miles N. W. and S. E.	27 4	18	176	6
Ĕ	Morotoi, W. point	21	10	157			Bunker's Island	128 O	XO 1	173	-
E	Woshoo	21	43	157	58	ĺ	Island	28 9	25	178	
Ę	— W. point Tahoorowa Ranai, S. point Morotoi, W. point Woahoo Attoi, Whymoa Bay Tahoora Oneeheow Bird's Island	21	57	159			Island	28 5	i4	178	
3	l'ahoora	121	40	160 160		ŀ	Swift's Island	32 5	3	119	6
ě	Oreehous	23	2	160							
2	Bird's Island	23	8	161			Culpepper's Island .	1 4	n	92	00
	Gardner's Island, disco-					ľ	Wenman's Island	1 2		91	
		25		167			Redondo Rock	0 1	5	91	34
	Maro's Reef, ditto	25	28	170	20	ŝ	Abington Island, C. Ib-		_		
	Gallego Island	1	42	104	5	Islam	etson Albemarie I. C Berke-	0 2	9	90	43
	Christmas or Noel I.			157		7	lev	0	2	91	21
	Sidney or Fanning's I.			159		g Se	- Christopher's Point		õs.		
		4		126		epa	James I. Harbour .	0 1		90	
	New-York Island		44	160	6		Charles I. S. P.	1 3		90	
	Cocoss Islands, or Chat- ham Bay	5	97	87	15	9	Chatham I. N. E. P.	0 4		89	.9
	Palmyra I.	5 5	49	162				0 5		89	
	Island	6	36 I	1 6 6	50	1	XLIX. Islands, Rocks	and	Shoa	ls, i	n the
ı	Barber's Island	8	50	178			SOUTH PACIFI	C O	CEA		
	Reef	10	UU I	179				Le		Lo	ng.
H		11		109 154		1	New Guinea	D . 1	м.	D.	M.
		11		164			- Middleburg Island	n e	ne l	1 20	16E.
	Island	13	9	168			- Cape of Good Hope	02	:O 1	132	
ľ		13		170			- Flat Point	0 4	6	134	
	Shoal	14	3U	170		١.	Cape Valshe!	S 2	6	137	23
	Cluster of Islands \cdot .	17	00	133 136		l '	Cape Rodney . King William's Cape	10	2	147	
	Island	16	30	163		ı	Torres or Endeavour	0 4	"	148	31
	Passion Rock	16	56	109			Straits		1		
	Cornwallis I	16	54	169			EasternFields or Reefs,		- 1		
l		18		114			N. E. end	10	2	145	43
l	Clarion Island Island	18 18	21	114 155			— N. W. part Murray's Islands	9 5		145	
ı	Shoal	is		170		ľ	Murray's Islands . Wamvax or Darnley I.	99	ရှိ	144 143	3 40
	Socora Island	18	48	110			Pandora's Shoals, N. P.	9 5	5	143	
		19		166	32	8	Wreck Reef, S. P.	11 ž		144	
1		19		109	53	Straits.	Portlock's Reef .	9 4	s i	144	45
1		19 19		115 111	10	orres	- ent. Torres Straits .	9 5		144	
1		19		165		Ę	Boot Reef	9 5	7	144 144	
1	Cloud's Island	19		115		-	Halfway Island			143	
1	Copper Island	20	6	131	54	l	Booby Ísle	10 2	7	141	
1		21		176			York I. (Mt. Adolphus)	10 3	7	143	
١.	oualer's island	22	6	112	14		- "			I	

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						_			_	_	
F			at.	Γo	ng. M.			La		Lo	mg. M.
ł		D.	М.	Į D.	М.				М.		
1	Prince of Wales' group,			l		•	Onascuse or Hunter's I.				
ı	N. P	1 -	00S.		12E.		De Peyster's Islands .			178	
ı	Kangaroo coral reef	13	22	143	47		Ocean's High Island .	0 4		170	
1	Providence Islands						Pleasant Island	0 2	0	167	10
ł	- Little Providence or			1			Gardner's Island	10		169	
1	Danger I	0	11	135	12			10 0	0	166`	50
ı	- N. W. ext. of Shoal	1		1			Ganges' Island	9 4	4	166	43
	off ditto	0	1	135	8		Stewart's Island	8 2	4	163	00
ı	Louisiade Isles					1	Egmont or SantaCruzI.		- 1		
1	- Cape Deliverance .	11	42	154	30		- Cape Boscawen .	10 5	5	166	10
ł	Stephen's Island	0	21	137			Pitt's or Alderney I	11 5	o	166	46
ı	Durour's Island	1	17	143	30			11 3		170	24
ł	Admiralty Islands, (1		146				10 3		166	12
ł	limits	3	10	148	6	ı		11 4	9	170	42
1	Sydney Shoal		20	146	50			12 1		169	
ı	Active's first Reef (dis-	-						12 1		172	
1	covered 1811)	3	40	146	53			11 4		174	
1	- second reef (ditto)		41	146				-			
ſ	New Ireland				3.		Sir J. Banks's Island .	13 2	7	167	24
ı	- Cape St. George .	4	54	152	59	1	Espiritu Santo, C. Lis-	•			
ı	— Carteret's Harb.	1 .	48	152	-		burne	15 4	1	166	57
•	New Hanover W. end		25	149	6			14 3		166	
ı	New Britain	ľ		17	٠		- Bay St. Philip and				
1	- Cape Palliser	4	18	152	10		St. James	15 1	o l	167	5
1	— Cape Orford		40	152				14 5		167	
ı	- Port Montague .		12	151	2			15 2		167	
ı	- Cape Ann		27	149				16 3		167	
ı	Cocos Islands		30	156			Mallicolo, C. Sandwich			167	
ł	Shoals W. of Bougan	_						16 2		167	
1	ville's Strait .	6	11	154	22			15 4		167	
ı	Bouganville Strait .		00	155		1	Aurora Island	15		168	17
ı	Laughian's Islands, S. E.					1	Table Island	15 3		167	7
	ext.	9	20	153	42	1		15 4		168	20
ł	Bridgewater Shoal .	8	54	156	49					168	13
8	Cape Deception		30	156		ı	Paoom Island	16 3	i O	168	
ł	Cape Nepean	8	51	157	32	١.	Three Hills	16 5	9	168	22
	Cape Marsh . ,	9	7	159	46	Hebrides.		16 4	6	168	22
ķ	Deliverance small Isld's	10	51	162	27	- 2	Sheppard's Islands .	16 5		168	42
H	Indispensable Strait, S.	ŀ		1		8	Monument	17 1		168	38
۲	ent	10	15	161	15	_ `	Montague Island .	17 2	6	168	32
Į	Bellona Island	11	6	159			Hinchinbroke Island .	17 2	5	168	38
Į		12	5	159	48			17 4		168	33
S	Pandora and Indispen-	l					Erromango, Traitor's	ı	ı		
1	sable Shoal, N. P.	12	9	160			Head	18 4		169	
1	S. P		46	160				19 1		169	
ı		12		157			l'anna, Port Resolution			169	
	Port Prasiin	7		157			Erronam	19 3	9	170	
l	Stewart's Island	-	24	163			Enatum	20 1	νį	170	4
1	Bradley's Shoal		45	161				22	. 1	169	2
t	Lord Howe's Group .		24	159				22 3	9	169	16
ſ	Hunter's Islands		48	157			Matthew's or Hunter's		. 1		
	Shank's Island		28	163			Island . , .	22 2	4	172	15
ı	Blaney's Island		39	174		ı	701	۱	. 1		
ı	Dundas Island		15	173			Diana's Bank, about .	15 4		150	
ı	Drummond's Island .	1		176		ı	Bougainville's Reefs	15 3		148	
ı	Byron's Island	1		177		ı		15 1		147	
ı	Hope Island		47	176		ı	Alert's Reef	17	2	151	49
	St. Augustine Island .		36	176			Mellishe's Keys and		ا ۔		
	herson's Island		56	176			Reefs	17 1		156	
	cillice's Group		29	179	6	ı		18 4		158	2
1	litchell's Group	9	6	179				19 3		158	
ł	Plaskett's I.	P - "	18	179				19 3		158	
ı	Independence I	10	25	179		ı	Chesterfield Bank .	19 5 90 5		158	
-	Littahall I		45.1	. s 7U	114	- 1	INDIAND SAMBLE				40

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	_				_					
•		at.	Loi	ng. M.	1	*		at.	Lu	
1	D.	M. 50S.			H	Museuspiele Island		M.	D.	M.
	21		159							45E.
1	50		158			The Judge and his clerk			160	7
Baring's Shoals .	20		159			The Bishop & his clerk			159	
_ (21	24	158				٥Û ده		166	
	21	9	155		1	Campbell's Island .	52		169	
			154				17		179	2
Mid-day Reef	21	58	134	20			¥9	33	179	2
G . Il land		•	162	51		Chatham Island, C.	١.,	40	. ~~	-011
Small low woody Island	10	3	162		ı		43			59W
	18		162				14	_	175	
	19						30	8	179	
The second secon	19	58	163				29		178	
	20	7	164				25		174	
Pudyona, N. W. P.	20	6	1	7			23		173	
Cape Colnet	20	30	164		ł		23		177	
	22	5	167	8		(20	6	168	36
Queen Charlotte's Fore-			164	••		D		- 1	l	
land	22	10	167			Rotumah or Grenville's				
Isle of Pines	22	42	167	1			12			57E.
	22	Z/	167	11			10			00W
Prince of Wales Fore-	60	20	1,66	E0.	1		9	.9	171	
land, S. P	22	3 U	166				8		172	4
	22		165				10		170	
Loyalty Island	טש	54	166	30			6		166	
1 2 2 2 2 2	1						11	_	162	
Wreck Reef and Sand	-			10			13	.6	163	
Bank	22	11	155				13		163	
Cato's Bank	23	6	155					22	176	
Reef	23	40	160		1			53	175	
	23	48	164		1		17		175	-
	25	00	166		1		15		174	
	26	4	160	UU	l		15	50	174	8
(20	12	1.00	9.	1	Navigator's Islands,	١	_		
Sir C. Middleton's Is.	28	13	160		•		14	9	169	3
		14	158		ı	- Leone, S. P	14	8	169	
	30	5	159		ı		14	5	169	18
Island	31	14	160		i			17	170	3
Lord Howe's I	31		159		i		14	3	171	7
Norfolk I. (Mt. Pitt) .	27	2	168		ı		14		170	
Rosavetta Reef	30	30	173	23			13		171	
No. of Co.		07	. ~~				13		171	
	34		173	4	يا	Amargura	18	00	174	
	35		175		Ę	Vavaoo (Howe's) Is.			174	00
	36		175		į,	Lati or Bickerton I			174	
Mercury Bay	36	48	176	6	۴		19	2	169	
Cape East	37	44	178		Ę,			46	175	
	38	22	178	-	ž		19		174	
	39	6	178	2	Įž		19		174	-
		43	177		۳		20		174	
	40		176		1		20	36	175	17
Banks' I. E. end .		43	173			l'ongutaboo,	1	_	1	
Cape Saunders	45	37	170		l	- Van Dieman's Road		6	175	
	46	8	169		1		21		174	45
	48	6.	166		1	Pylstaart's Island .	22	22	175	41
	48	15	166		1		t		Ì	
Cape South		17	167		ı		27	46	176	00
	16		167		ı	King George's Reef .	19	56	167	30
1		28	166		Ι,	Palmerston Island .		00		57
		56	166	6	1	Whytootaeke		56		45
Dusky Bay	45	40	166		١.	Hervey's Island		17		48
Open Bay	13	51	168		ş	Wateoo Island	20		159	
Cape Foulweather .	41		171	30		Maria I		45	155	
Cape 'arewell	40	40	173	18	3	Mangeca Island .		57	158	
cueenCharlotte'sSound		5	174	40	2	Roxburgh Islands .		36	159	
Cape Campbell	41	34	174		Society		1			σle
Cape Paliser	41	24	175	41	Š	Scilly Island Digit	16	30		ao
epe Egmont		23	174	12		Lord Howe's I	16	-16	154	
ret Island	39	5	175	5			16		152	
			-		_		,			

Г		Lat	Long. D, M.		long. O. M.
ı	Bolabola Island .	D. M.	151 521	Onateays Island . 9 58S. 13	
ı	Clietea	16 45	151 31		8 49
ı		16 45	151 35		,
l	Huaheine, OwharreBay		151 8	Bunker's Shoal 0 17 16	60 40
1		17 25	150 58		9 50
I	Eimeo (Taloo harbour)		150 00	Island	88 54
3	Tethuroa	17 1	149 36	Brock's Island 1 13 13	59 3 0
œu;	Otaheite, Point Venus		149 36		73 45
Isla	— Cuitipeha Bay .	17 46	149 14	Hero Island 5 40	55 55
Į.	Osnaburg or Miatea .	17 52	148 6	Island 6 39 10	66 18
2			1		39 54
ģ		14 58	147 50	Pennryhn's Island . 9 1 1	57 35
۳.		15 38	146 30		56 57
ı		17 25	145 30		56 50
ı		20 31	145 54	10.00	60 49
ı		22 27	150 49	10.46	60 55
l		22 43	152 00	60.00	66 6
		23 25	149 20	1 2 2 2 2 2	59 25 56 7
ì		23 42	148 3	1 1 2 2 2	43 00
Ī	Byron's Islands	14 20	145 0		52 6
	— Taoukaa Island Disappointment Islands	14 30	145 9 141 22		52 0 55 12
l		17 7	144 22		44 59
l		17 11	143 7		46 10
ŀ		17 23	141 45	Isle de Chiens . 14 50	38 47
l	l	16 00	139 00		44 28
۱.	l ·	17 00	138 00	Isles de Kru-)	11 20
8		17 49	142 43	senstern exten-	
H	Bow Island	18 17	140 43		48 41
3	Pr. Henry's Island	19 00	141 22	and S. S. W.	
Archipe	Cumberland I.	19 18	140 52	15 miles	
5	Gloucester Island	19 11	140 20	Chaine du Rurick, N.	,
[§	Queen Charlotte's I.		138 20		
15	Whitsunday Island		138 12		46 30
12	Lagoon Island .	18 48	138 33	_ W. P 15 20	
احّا	Queen Charlotte's I. Whitsunday Island Lagoon Island	1		1 2 3 3 45 45 1	46 56
`	Osnaburg Island .	22 8	140 37	Dean, or Prince) of Wales, or W.P. 15 00	
ı		21 43	140 30		48 22
1	Carysfoot Island	20 49	138 33		47 12
1	Lord Hood's Island	21 31	135 32		39 00
1	Gambier's Island	22 55	135 00		38 0 0
ŧ		23 12	134 32		67 50
1		24 26	135 6		78 36
L		25 4	130 25		78 47
ı	Opero Island	27 36	144 11		61 4
ł		J		Anderson's Island (or	
1	Nukahiwa I. (Federal		100 40		28 11
1	Port Tochitschagoff		139 42		24 40
1	- Port AnnaMaria, ent		139 40		30 28
1:	Cape Martin S. E.P		139 32	1 1	79 47
Tolomile	South Point	. 8 59	139 44		80 4
		. 8 53	139 49		92 24
3	Oanuga I. (wasning-	. 8 58	139 13		05 34 109 40
	tou I.) W. P. Uapoa I. (Adams)	. 9 21	139 13		95 12
		9 29	107 99	Group of Islands 31. 3	129 24
Г	Level I. (Lincoln) Mottauity Islands	. 7 23		1 100 45	80 38
B	(Franklin) .	. 8 37	140 20	Juan Fernandez S.W.P. 33 45	79 6
ľ	Hiau I. (Knox, Roberts		140 13	E. P	78 53
1	Small Sandy Island	7 57	140 3		
ı	Fattuuhu I. (Hancock		140 6	NEW SOUTH SHET-	
1	.)	7.50	1	DE P	
1	Hood's Island .	. 9 26	138 52	Clarence Island, Floyd's	
1	Hood's Island Ohevahoa	. 9 41	139 2	Promontory 60 57	54 6
1	Ohitahoo, Resolutio				54 8
	Bar .	9 55	139 9	Cornwallis Island 61 00	54 28
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	T	Lat.	L	ong.		at.	L	ong.
	D.	M.	D.	M.	D.	M.	D.	M.
Seal Islands .	. 61	008.	55	32W				
Cape Valentine .	. 61	3	54	40	count)	428.	62	90 W
Sarah Island	. 61	22	55	30	Ditto the harbour	- 1		
Obrien's Islands	. 61	28	56	35	(by another person) 62	55	63	5
Bridgeman's Islands	. 62	00	57	12	New Plymouth . 62		61	37
Cape Melville .	. 62	00	57	46	Monroe's Island, Presi-			_
Sheriff Cape .	. 62	28	60	57	dent's Bay62	46	62	20
Ditto (another ac	e-l				Castle Rock (west of			
count)	. 62	21	61	47	Monroe's l.) .62	50	62	30
Yankee Straits .	. 62	30	60	22	Mount Pisgah 63		63	00
Ragged Island .	. 62	40	62	10	Ditto (another ac'nt. 62		63	

Note to page 279, Table XLVI. line 1.

I have considered the Essex Shoal to be the same as the Fairlie Rock, and have given its latitude and longitude as in Horsburgh's Directory, namely 3° 27′ S. 107° 2′ E. The place assigned by Captain Orne, of the Essex, is 3° 36′ S. 107° 00′ E. differing nime miles in latitude; and as it is possible that the rocks may not be the same, I have now given Captain Orne's estimate, made from a meridian observation two hours after striking on the shoal, June 26, 1804. He described it as "a small rock or coral patch, seen by the man at the mast head an instant before she struck; but there was no appearance of a breaker, though the breeze was fresh, and a short sea running. In the act of wearing ship, she struck rather on the side of the rock, which reduced her velocity from 8 to 2½ knots; after rubbing a few seconds, she fell off into deeper water (8 fathoms) without any material damage."

D0 4 0				with the vertical rise			_
PLACES.	SITUATION.	TIME.	R.	PLACES.	SITUATION.	TIM	E.
	-	H. M.	FŦ			H. 1	
	France	10.30		Bolt Head	England	5.5	
	Scotland	12.45		Bombay	India	11.1	
	Wales	7.30	13	Bombay Offing		12.	
	Ireland Isle of Man	5.30		Borkum Island	Holland	11.	
		10.30 3. 0		Boston	England	6.4	
	France	7.30		Boston	America	11.3	
lban's Head (St.)		11		Botany Bay	New Holland	8.	
	America France	6. 0 11. 0		Boulogne	France	10.4	
		10.30		Bourdeaux	France	3.	
	North Sea	1		Brassa Sound	Shetland ;	10.	
	America	8.30		Bray Head	Ireland	3.3	
	Anglesea.			Bremen	Cormany	6.	
	Holland	3. 0		Brest	France	3.4	
msterdam Island		8.30		Bridgewater	England	6.4	
ndrew's Bay, St.		2.15		Bridport	England	6.4	40
Ingra Bay	Terciera.	2.20	d	Brighton	England	10.	
	Cattegat	12. 0		Bristol	England	6.4	
	America	11.30	11	Broad Bay	America	10.4	
	America	11. 0		Broad Haven	Ireland	6.	
nticosta I. W. end		3.30		Burnt Island	Scotland	2.	
ntwerp	France	6. 0		Button's Islands	Hudson's Bay	6.4	5U
	Pacific Ocean	6. 0		C		_	_
	Russia	6. 0	Ì	Cadiz Caen	Spain	3.	
	Ireland	8.15	_	Caen	France	9.	
	Scotland	11.15	9	Caernarvon	Wales	9.	
	England			Calais	France	11.	30
lugustine, (St.)	America	7.30		Caldy Island	Wales	6.	
ugustine's Bay St.		2.15	ı	Calf of Man	St.George's Ch.	10.3	30
vranches	France	6.00	ı	Campbell (Port)	America	9.	
₽		1 1	1	Canary Island	Atlantic Ocean	3.	0
	Red Sea	12. 0		Canso (Cape) Cantire (Mull of)	America	8.3	30
alisore	India	9.45	12	Cantire (Muli of)	Scotland	10.3	30
allingskellings B.	Ireland	3.15	ı	Capricorn (Cape)	New Holland	8.	
	Ireland	9. 0		Cardiff	Wales	6.	0
	Ireland	6.45		Cardigan Bar	Wales	7.	0
	Ireland	4. 0		Carlingford	Ireland	10.	0
	Scotland	11.30		Carlisle	England	12.	
	Ireland	3.30	1	Carmarthen	Wales	6.	30
	Wales	8.15		Carrickfergus	freland	10.	
	France	7.30		Caskets	Eng. Channel	8.	
armouth	Wales	8. 0	13	Catherine's Pt. St.	Isle of Wight	9.	
arnstable Bay	England	5.50	26	Catness	White Sea	5.	
	England	10.30	~~	Cayenne	S. America	6.	
	France	3.30		Charente Riv. ent.		4.	ř
eachy (on shore)			90	Charles (Cape)	America	7.	4 4
	England	11. 0	-	Charleston Bar	America	7.	
ear Island	Hudson's Bay	12. 0		Chatham	England	1.	
eaumaris	Wales	10.15	94	Chepstow		7.	
	England	11.15	~	Cherbourg	England France	9.	
	Ireland	10.30		Chester Bar		11.	
	Bay of Biscay	3. 0		Chichester Heat	England		
	Isle of Wight	11.40		Chichester Harb. Christmas Sound	England	11.	ou or
	Norway	1.30		Chunchill (Corr.)	8. America	2.	
	Atlantic Ocean	7 7		Churchill (Cape)	Hudson's Bay	7.5	
erwick		7. 0			Ireland	3.	5 U
	England	2.19	10	Cod (Cape) Condore Pulo	America	11.5	
	Spain Sacia	3.40		Condore Pulo	China Sea	4.1	
		3.45		Conway	Wales	10.	15
Bilboa	Spain Basis			A . *			
liboa lakeney	England	6. 0		Copeland Island	Ireland	10.	
lakeney lanco (Cape)	England Africa	6. 0 9.45		Copeland Island Coringa Bay	Ireland India	10.: 9.:	15
Silboa Slakeney Slanco (Cape) Slaskets	England	6. 0		Copeland Island	Ireland	10.	15 45

TIMES OF HIGH WATER.

	TIMES	UF	-11	IGH WATER.			_
PLACES.	SITUATION.	TIME	R.	PLACES.	SITUATION.	TIME.	R.
		II. M.	FT			н. м.	FT
Cork Harbour, en.	Ireland			Fly or Vlie Gatway	Holland	6.45	
Corunna	Spain	3. 0		Fly or Vlie Road	Holland	7.30	
Coutance	France	6. 0		Foreland (North)	England	11.15	16
Cowes	Isle of Wight	10.15	15	Foreland (South)	England	11. 6	15
Crocotoa Island	Str. of Sunda	7. 0	3	Formby Point	England	11. 0	
Cromartie	Scotland			Fort St. John Fox Island	Newfoundland America	9. 0	1
Cromer Crookhaven	England Ireland	3.30		Fowey	England	5.30	
Cross Island	White Sea	4:15		Funchal	Madeira	11.30	7
Cuxhaven	Germany	1. 0		G			11
D				Gallicia (Coast of)	Spain	3. 0	1
Dartmouth	England	6.10	20	Galloper	Thames River	12.45	11
David's Head, (St.)		6. 0		Gallway Bay	Ireland	4.30	1
Deadman's Pt.	England	5.30		Galloway, Mull of	Scotland	11.15	
Deal	England			Gambia R. ent.		10.15	
Dee (River)	Scotland	11. 0		Gay Head	America	7.37	
Delaware R. ent.	America	9. 0 2.15		George's River	America America	10.45 7. 0	
	India E	10.30		Georgetown Bar	India	4.80	
Dieppe Dieple Pow	France Ireland	3.30		Good Hope, (Cape)		3. 0	
Dingle Bay Donnegal	Ireland Ireland	6.30		Good Hope, (Cape)	Africa	2.30	
Dover	England		14	Gorce Gatway	North Sea	1.30	
Douglas	Isle of Man			Gouldsborough	America.	11. 0	
Downs	England		15	Granville	France	7.30	N
Drogheda	Ireland	10.45		Gravelines	France	11.45	
Drontheim	Norway	2.15		Gravesend	England	1.30	
Dublin	Ireland			Grizness, (Cape)	France	11. 0)
Dudgeon Lights	North Sea	6. 0		. н		١, ,	J
	Scotland	2. 0		Haerlem	Holland	9. (
Duncansbay Head		10. 0 10.45		Hague La, (Cape) Halifax	Nova Scotia	7.30	ه اه
Dundalk Bay	Ireland Ireland			Hamburgh	Germany	6. 0	13
Dundedy Head Dundee	Scotland	2.15		Hartland Point	England	6. 0	
Dungarvon	Ireland	4.30		Hartlepool	England	3.45	
Dungeness	England			Harwich	England	11.30	14
Dunkirk	France			Hasborough Gatt	England	6.30	
Dunnose	Isle of Wight	9.15		Hastings	England	10.36	
Dursey Island	Ireland	3.30		Hatteras, (Cape)	America	9. 0	
E .				Havre de Grace	France	9. 0	
Eastern Brace	Bay of Bengal	9.45		Helena, St.	Atlantic Ocean	2.1	
Eddystone	Eng. Channel	2 30		Helens, St.	Isle of Wight Holland	11.43	
Edinburgh Elbe R. (red buoy)	Scotland North See	12. 0		Helvoetsluys Henlopen, (Cape)	America	8.4	
Elizabeth Town Pt.		8.54		Henry, (Cape)	America	7.40	
Embden	Germany	12. 0		Holybead Bay	Wales	10. 0	
Exmouth Bar	England			Holy I. Harbour	England	2.30	
Exuma Bar	Bahamas	6.35		Honfleur	France	9.30	
Eyder River	Germany	12. 0		I-Tull	England	6. (119
Eyemouth Harb.	Scotland	2.15		Humber R. ent.	England	5.13	
F				Hurst Castle	England	9.34	4
Fair Head	Ireland	9. 0		l		l	J
Falmouth	England			Ice Cove	Hudson's Bay	10.	
Fayal Road	Azores	8.20	43	Ipswich	England	12. (
Fear (Cape)	America	8. 0		Ireland, W. Coast Ireland, S. Coast	Atl Ocean	3. 0	
Fecamp	rrance Spain	10.30 3.0		Isle de Dieu	France	3.	
Ferrol Ferriters	Spain Ireland	3.30		Isle of Man, S. side	St. George's St.	10.9	1
Fifeness	Scotland	2. 0		Ives, (St.)	England	5.1	24
Filey	England	4.30		Jackson, Port	New Holland	8.1	
Finsterre (Cape)	Spain	3. 0		Janeiro Rio	S. America	4.30	
Finmark	Lapland	2.15		Johns, (St.)	Newfoundland	6. 0	İ
Fisguard Bay	Wales	6.30		Jutland Coast	Denmark	12. 0	
Flamboro' Head	England	4.30		K	L		۱ '
Florida Keys	America	8.50		Kedgeree	India	111 30	1
Fluehing	Holland	1 1. 0	•	Kenmare River	Ireland GOC	7319 0	<u> </u>

TIMES OF HIGH WATER.

	TIMES	UF	Π.	IGH WATER.			_
PLACES.	SITUATION.	TIME.	R.	PLACES.	SITUATION.	TIME.	ŀ
		н. м.	FT		1	н. м.	ŀ
Kennebeck	America	10.45	9	North Cape	Lapland	3. 0	1
Kentish Knock	R. Thames Ireland	11.45		0		}	
Killibegs King's Channel	River Thames	6.45 12.50		Olonne Oporto	France Portugal	3.30	
King's Road	Bristol Chan.	6.45		Orfordness	England	10.30	
Kinsale	Ireland	4. 0		Orkney Islands	North Sea	10.30	
Kinnaird's Head	Scotland	12. 0	1	Orm's Head	Wales	10.15	
L L	l		_	Ortegal (Cape)	Spain	3 0	
Lambaness Lancaster	Shetland	9.30		Ostend Owers	France	12.30	
Land's end	England England	11.15		Owers P	Eng. Channel	9.36	Ή
Leith Pier	Scotland		15	Padstow	England	5.45	1/2
Lemon and Ower	North Sea	7. 0		Passamaquoddy R.		11.30	
Lerwick	Shetland	9.45		Passier Roads	Bornes	5. 0	
Lewis Islands	Scotland	6. 0		Penmarks	France	3.30	
Lewis (Butt of) Limerick	Scotland Ireland	6. 0	٠,	Penobscot River Pentland Frith	America	10.45	
Lisbon	Portugal	3. 0	• 4	Penzance	Scotland England	5. 0	
Liverpool	England		27	Peter Head	Scotland	12. 0	ı
Lizard	England	5. 0		Plymouth Sound	England	6. 5	h
Loch Swilly	Ireland	6.30		Plymouth	America	11.30	1
Loire River	France	3. 0		Pol de Leon (St.)	France	5.15	
London Londonderry	England Ireland	6. 0		Poole	England Scotland	9. 0 11.45	
Long Sand Head	River Thames	11.30	1	Port Glasgow Port Hood	Cape Breton	7.30	
Longships	England	4.30		Port Howe	Nova Scotia	8.30	
Lookout (Cape)	America	9. 0	7	Port Jackson	Nova Scotia	8. 0	
Loop Head	Ireland	4.15		Portland Bill	England	7.15	
L'Orient	France	4.0		Portland Race	England	9.15	
Lundy Island Lyme Regis	Bristol Chan.	5.45 7.5	30 <u>i</u>	Portland	America France	10.45 4. 0	
Lynn Deeps	England England	6.30		Port Louis Porto Praya		11. 0	
F, 200M	Sug-ene	0.00		Port Roseway	Nova Scotia	8.15	
Machine	America	11. 0	19	Port Royal Island		8.15	l
Madeira	Atl. Ocean	11.30	7	Portsmouth Harb.	England	11.36	
Malacca Roads		10.30		Portsmouth	America	11.15	
Malo, (St.) Marbiehead	France America	6.30		Portugal (Coast of) Pulo Pinang	Europe India	3. 0 1.30	
Margate Road	River Thames	11.45	:4	T (110 I III) MING	Inum	1.30	ľ
Martin Vas	Atl. Ocean	3.45	٠٦	Quebec	Canada	8. 0	ı
Marys (St.)	Scilly Islands	4.40	- 1	•	India	10. 0	١
May (Cape)	America	8.45		R		م ما	
Milford Haven	England	6. 0 3.15		Raehlin's Island Ram Head	Ireland	9. 0 5.45	
Mizen-head Montrose	Ireland Scotland	1.30		Ramsey	England Isle of Man	10.30	
Morocco Coast	Africa	2.15		Ramegate	England	11. 0	
Mount's Bay	England	5. 0	10	Rhe Island	Bay of Biscay	3. 0	
Mount Desert	America	11. 0	12	Rhode-Island	America	6.45	
N		1 1		Rio Janeiro	S. America	4.30	
Nangasaki	Japan	7.53		Robin Hood's Bay		3.45	
Nantz Nantz River ent.	France France	4. 0 3. 0		Rochefort Rochelle	France France	3. 0 3.45	
Nassau	N. Providence	7.30		Rochester	England	1. 0	
Natal Rivier	Africa		12	Rodrigues Island	Indian Ocean	12.45	
Needles	Isle of Wight	l 9. O		Roman (Cape)	America	8.0	
Newcastle	England	4. 0		Roseness		10.30	
New-Bedford	America	7.57		Rotterdam	Holland	3.30 10.51	
Newburyport New-Haven	America America	10.16	14	Rye Harbour S	England	10.51	*
New-London	America	8.54		Sable (Cape)	Nova Scotia	8. 0	
Newport	Wales	6.45		Sable Island	America	8.30	ŀ
New-York	America.	8.54	5	Salem	America	11.30	1
Noatka Sound	North America			d	S. America	3.45	
Nore Light North Berwick					England OG	11. 0 9. 0	l,
MOTER DEFWICE	Scotland	2. 0		Sandwich Bay	Nova Scotia	7. 0	_

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TIMES OF HIGH WATER.

PLACES.	SITUATION.	TIME	R.	PLACES.	SITUATION.	TIME.	L
		н. м.	FT			H. M.	F
	New-Jersey			Telling (Cape)	Ireland	6. 0	
Scarborough	England	4.30	13	Terciera	Azores	11.45	ł
Scaw	Denmark	12. 0	16	Texel (ent. of)	Holland	6.45	
Scilly Islands	Eng. Channel	4.40	18	lexel Road	Holiand	7.45	1
Seal Islands	Bay of Fundy	8.45		Thames, R. mouth	England	12. 0	
Seine River	France	9. 0		Tinmouth	England	3. 0	1
Selsea Bill	England			Todhead	Scotland	18.45	
Selsea Harbour	England	11.15	15	Torbay	England	6.10	20
Senegal R. ent.	Africa	11. 3		Tory Island	Ireland	6. 0	1
	Lapland	9. 0	15	Townsend	America	10.45	1
	Ireland			Tuskar Rock	Ireland	7. 0	
	England	12. 0	15	Typa Roads .	River Canton	10. 0	
beepscut	America	10.45		Ü		I 1	
	North Sea	10.30	6	Ushant (within)	France	3.45	20
	England	3. 0	13	Ushant (without in			
	England	9.21			France	4.30	١
	Guinea	8.15		v v		1	١
	America	7.30		Vannes	France	4.30	1
Skerries	Wales	10. 0	1	Vincent (Cape St.)		2.30	
Skerries	Scotland	5.30		w w	opun.	1	1
Sky Island	Scotland	6. 0		Wardhuys	Lapland	4. 0	ı
	Ireland	6.45		Watchet	Bristol Chan.	6.45	
	Ireland	5.15		Waterford Harb.	Ireland	5.30	
Smalls	Walce	5.50		Weser, River ent.	Germany	12. 0	
	France	10.30		Vestern Brace	Bay of Bengal	9.36	1
	England			Wexford Harbour	Ireland	7.30	
	England	9. 0		" eymouth	England	6.15	
Spain (N. coast of)		3. 0	1	Whitby	England	3.45	
Spurn Point	England			Whitebayen	England	11.15	
	England			Wicklow	Ireland	9. 0	
	England	4.30		Winterton	England	8.15	
	Scotland	1. 0	•	Woolwich	England	2.15	ıi
	Orkneys	10.30			Scotland	7. 0	
	N. America	9.30		Wrath (Cape)	OCULANA	1 . 4	•
Sunderland	England			Yarmouth Roads	P-44	8.45	
Swansey	Wales	6. 0	معا	Yarmouth	England	9.30	
	Lapland			Yorkshire Coast	Isle of Wight	4.30	١-'
T	~ehrana	1	۱,۰	Youghall	England Ireland	4.30	

APPENDIX,

CONTAINING METHODS OF DETERMINING THE LONGITUDE BY OBSERVATIONS OF ECLIPSES, OCCULTATIONS, &c.

THE longitude of a place may be determined in a very accurate manner, by observing the beginning or end of a solar ealipse, or occultation of a fixed star by the moon, or the difference between the times that the moon and a known fixed star pass the meridian. These observations, when made on land with a good telescope and well regulated time-keeper, furnish, by far the most accurate method of determining the longitude, and when made on board a ship without a telescope, will in general give it to a greater degree of accuracy than any other method: For this reason, it was thought proper to insert in this Appendix the usual rules of calculating such observations, by means of the Nautical Almanac. The first thing to be taken notice of, is the method of determining the longitude, latitude, &c. of the moon or other object, having regard to the unequal velocity between the times for which these quantities are given in the Nautical Almanac. This calculation is rendered much more simple by making use of the signs + and -, and performing addition and subtraction as in the introductory rules of Algebra; and as it is possible that these rules may not be familiar to some readers of this work, it was thought proper to prefix an explanation, as far as will be necessary, in the present problems.

Quantities without a sign, or with the sign + prefixed, are called positive or affirmative, as 7 or + 7; and those to which the sign — is prefixed, are called negative, as — 7. Addition of quantities having the same sign, that is, all affirmative or all negative, is performed by adding them as in common arithmetic, and prefixing the common sign. Thus the sum of + 4 and + 3 is + 7. The sum of — 4 — 3 and — 5 is — 12. When the quantities have not the same sign, the positive quantities must be added into one sum and the respective into suches, as above the difference of these two parts. sum, and the negative into enother, as above; the difference of these two sums, with the sign of the greater sum prefixed, will be the sum of the proposed quantities. Thus the sum of +14, -7, +5, and -2, is found by adding +14+5, whose sum is +19; and then -7 and -2, whose sum is -9; the difference of 19 and 9 is 10, to which must be prefixed the sign of the greater number 19, which is +, so that the sought sum

is + 10. The following examples will illustrate these rules.

Subtraction is performed by changing the sign of the number to be subtracted from + 10in from - to +; and then adding the numbers by the preceding rule. Thus to subtract + 3 from + 7 the sign of + 3 must be changed, and the numbers - 3 and + 7 added together as in algebra, which by the preceding rule gives + 4; and if it were required to subtract - 3 from 7, the sign of - 3 must be changed, and + 3, + 7 added together. The sum + 10 represents the sought difference. It is not usual to make an actual change of the sign in any proposed question, it being sufficient to suppose the number to be subtracted to have a different sign from that prefixed to it, and to perform the operation accordingly. To illustrate this, the following examples are added.

Observing that when no sign is annexed to a quantity, the sign + is always understood

PROBLEM I.

To find the longitude, latitude, &c. of the moon at any given time at Greenwich, having. regard to the unequal velocity between the times marked in the Nautical Almanac, The intervals of these times being 12 hours.

RULE.

Take from the Nautical Almanac the two longitudes, latitudes, &c. next preceding the given time at Greenwich, and the two immediately following it, and set them down in succession below each other, prefixing the sign + to the southern latitudes or declinations, and the sign — to the northern. Subtract each of these quantities from the following for the first differences, and call the middle term arch A; subtract
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each first difference from the following for the second differences, and take the half sum or mean of them, which call the arch B, noting the signs of the quantities as in

algebra.

Find the difference between the given time and the second time taken from the Nautical Almanac, which call T, then to its logarithm add the log. of A and the constant logarithm 5.36452, the sum rejecting 10 in the index, will be the logarithm of the preportional part,* to which prefix the sign of the arch A; observing to express all these quantities in seconds.

Enter Table XLV. with the arch B at the top and the time T at the side, opposite to this will be the correction of second differences, to which prefix a different sign from that of the arch B, and place it under the proportional part found above, and the second quantity taken from the Nautical Almanac, and connect these three quantities tagether as in addition in Algebra; the sum will be the sought longitude, latitude, &c. The latitude or declination being south if it has the sign +; north if it has the sign -

EXAMPLE 1.

Required the longitudes and latitudes of the moon, December 12, 1808, at 15h. 49'-29" and 17h. 1m. 29" app, time by astronomical computation at Greenwich, which correspond to the immersion and emersion of Spica, calculated in Problem VII.

1808. Dec.	D long. N. A. 1st. diff.	1 0111	1st. diff. 21 diff.
12 noon	6 10 45 20 , 7 6 16	+5 2 +2 6 37	-94 21
12 midn.	6 17 51 36 A 7 11 18		A-96 45 -2 24
13 noon	6 25 2 54 7 16 5	B=+4.54.5 +0.51 18	_31 34 49
13 midu.	7 2 18 59		B=2 65

IMMERSION.

Constant 5. T= 3h. 48' 29'= 13709'' log. 4. A= 7 11 48 = 25878 log. 4	.13701 4.13701
+ 2 16 52.2= 8212.2 log. 3	.91446 11 89.7=_ 699.7 . log. 2.84494
+6 17 51 36 Second longitude. — 51.9 Tab. XLV. B = 4'	54".5 + 2. 6 37 Second Intitude. 13.7 Tab. XLV. B=-2' 5".5
6 20 7 56.3 D's longitude.	+ 1. 55 11.0 D's latitude south.

EMERSION.

Constant 5.36452 T= 5h. 1' 29"= 18089" log. 4.25742 A 7 11 18 = 25878 log. 4.41238	
+ 3 0 36.0= 10836 log. 4.03487	- 15' 23" 3 =-923.3 log. 296535
+6 17 51 S6 Second longitude, - S5.9 Table XLV. B=4' 54''.5	+2 6 37 Second latitude. + 15.4 Tab. XLV. B==-2' 6".5
6 20 51 36.1) y longitude.	+1 · 51 29.1 D's latitude south.

These quantities are made use of in Problem VII.

EXAMPLE II.

Required the longitudes and latitudes of the moon, June 16, 1806, at 2h. 49' 58." and 5h. 34' 6'.6 app. time, astronomical account at Greenwich, which correspond nearly to the beginning and end of the total eclipse of the sun as observed at Salam.

1806.	June.	1	long	. N. A	1st diff	,	2a. diff.	D	kt. Ņ.	4:	1st. diff.	1 1	M dic
15d. 16 16 17	midn. neon.	2 2	14 22 29	48 58 6 19 27 12 50 47	7 17 2 A 7 20 3 7 23 3	53 15	+3 52 +2 42 B=+3 7		-1 14 -0 84 +0 6 +0 47	6 13 33 28	1-90 53 A-1-40 46 1-49 55		+68 0 31

^a This may be found to minutes by Table XXX. by entering it at the top with half the arch A (because the table extends only to 3° 45') and at the side with the time T; the result doubled will be nearly the sought proportional part; but the table not being calculated to seconds, it is hardly accurate enough to be used in calculating eclipses. This correction may also be found by proportion; by saying as 12 hours are to the time T, so is the arch A to the sought proportional part, and this method is the shortest when T is an allquot part of 12 hours. Thus if T be 3, 6 or 9 hours, the proportional part will be \(\frac{1}{2}\), or \(\frac{1}{2}\) of the arch A respectively. This method is made use of in Problem XVII. in interpolating the distances of the moon and sun.

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[†] If the arch B consists of minutes and seconds, the correction for minutes, tens of seconds and units of seconds must be found separately, the sum of these three parts will be the sought correction. Preparational parts for the minutes of the time T may be taken in finding the correction of this table when secondary. In this rule, part of the correction of third difference is neglected. This part never excess with of the third differences, and rarely amounts to a small fraction of a second.

	BEGINNING AT 2h. 49' 50', 1=T.	
Second longitude A 7° 20′ 58′ Prop. part B 8 7 Tab. XLV.	2s. 22° 6′ 19″ Second latitude N. + 1 43 59.8 A 40′ 46″ Prop. part - 16.8 B 31 Tal. XLV.	- 0° 34′ 18″ + 9 37.0 - 2.8
D's longitude	2 23 50 2.0 D's latitude N.	- 0 24 38.8
	END AT 5h. 34' 6".6=T.	,
Second lengitude A 1° 20° 88° Prop. part B 8 7 Tab. XLV.	2s. 22° 6′ 19″ Second latitude N. + S 24 35.3 A 40′ 46″ Prop. part - 23.2 B=31″ Table XLV.	- 0° 54′ 13″ + 18 55.0 - 3.8
)'s longitude	2 25 30 31.1 D's latitude N.	- 0 15 21.8

The proportional parts of the arch A were calculated in this example by arithmetic without logarithms. By observations of the eclipse on that day, it was found that the moon's longitude was too great by 58".5 and her latitude too great by 11".4. These corrections are applied to the above longitudes and latitudes, in calculating the eclipse.

Remark 1. It will not be necessary to take notice of the second differences in calculating the parallax or semi-diameter of the moon, or any of the solar elements useful in calculating an eclipse or occultation. In this case the quantity immediately preceding and following the proposed time at Greenwich, must be taken from the Nautical Almawas, and their difference, will be the arch A, also the difference between the proposed time and that taken first from the Nautical Almanac is to be called the time T. Then by proportion, as the interval between the times taken from the Nautical Almanac is to the time T, so is the arch A to the correction to be applied to the first quantity taken from the Nautical Almanac, additive if increasing, subtractive if decreasing. rection may also be found by logarithms as above, using the constant logarithms 5.36452 if the interval of the times in the Nautical Almanac is 12 hours, and 5.06349 if the interval is 24 hours. The proportional part of the moon's parallax and semi-diameter may also be found by Table XI. and that of the solar elements by Tables XXX. XXXI. as taught in the explanation of these tables. To exemplify this, the rest of the quantities requisite in calculating the eclipse and occultation (Prob. VI. VII.) are here found.

	EXAMPLE I.													
1808.	B.	S.D.	D.	H.P.	1909.		10	ng.		R.	Α.			
Dec. 12 midn.	18	17"	59	40"	Dec. 12 noon.		შ 0°2			. 18		4		
Dec. 15 noon.	16	23	7 60	8	13 ricon.	8 2	1 2	3 10	177	22	29	5		
Difference A.	1	В	1	20	Difference A.		1	6	1	4	25	1		
Pro. T=\$h. 48' 29"	l	1.9	!	6.3	Pro. T=15h. 48' 29"	1	4	15	1	2	54	ť.		
Correspond. values	16	18.9	59	523	Correspond. val.	8 5	11	2 19	17	20	59	0		
Pro. T .== 5h. 1' 29"	l	2.5	•	8.4	Pro. T=17h. 1' 29"		4	3 21	1	3	. 8	1		
Correspond, values	16	19.5	59	54.4	Correspond, val.	8 2			17	21	12	ā		
EXAMPLE IL														
4000	_		_											
1906.	D:	8.D.	D:			l	0 k	ng.	ı	♠ R	. A.			
June 16 noon.	16		60'	21"	June 16 noon.	E∰	34	18"	5b.		20'			
16 mldn.	16	30	60	34	17 ndon.	85	31	35	5	40	80	0		
Differences A.	1 +	3	1+	13	Differences A.	١.	87	17	1	4	9	4		
Pr. part T. 2h. 49' 50".1		0.7	1 🛨	3.1	Pr. part T. 2h. 49' 50".1	1+	8	45.4	1	+	29	4		
Correspond. values	16	27.7	60	24.1	Correspond. val.	84	41	8.4	5	36	50	0		
Pr. part T. 5h. 34' 6".6	1 +	1.4	比土	6.0	Pr. part T. 5h. 34' 6".6	1+	13	17.5	1	+	57	9		
Correspond. val	16	28.4	60	27.0	Correspond. val.	84	47	35.6	5	57	18	5		

The semi-diameters thus found must be decreased 2" for inflexion, and augmented by the correction Table XLIV. in calculating an eclipse or occultation by Problem XIII. or in deducing the longitude from observations by Problems VI. VII. VIII. er !X.

The sun's semi-diameter by the Nautical Almanac, June 13, 1806, was 15' 46".3 and June 19, 1806, was 15' 45".9. Hence at the above time it was 15' 46".1. This in eclipses of the sun must be decreased 34" for irradiation.

Remark 2. The above rule for calculating the second differences of the lunar motions where the intervals in the Nautical Almanac are 12 hours, may be made use of when the intervals are three, six, &c. days, (as is the case with the elements of the motions of the planets) by taking two longitudes, latitudes, &c. before, and two after, the given time at Greenwich, and thence deducing the arches A, B, and the longitudes, latitudes, &c. and then making use instead of T, of the quattent of the difference between the given time and that marked in the Nautical Almanac against the second longitude, &c. divided by the number of half days in the given interval. Thus, if the interval is one day, the divisor is two: if the interval is 3 days, the divisor is 6; and if the interval is 6 days, the divisor is 12. Thus if it were required to find the geocentric longitude of Jupiter July 14d. 13h. 30, 1808, astron. acc. at Greenwich, the work would be as follows.

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	•			•				
	24 long.		1	Second longitude .				11s. 17° 58' 0
July 7				A=-11' Prop. part B=- 7' Tab. XLV				- 2 52
18	11 17 59	— 3'	8	B=- 7' Tab. XLV	•		•	+ 40
19	11 17 48	A11	6	2f longitude				11 12- 58 48
95	11 17 31	-17	B= 7	Il set minteriore	•	•	٠	

In this example the time T, is 3h. 7' 30", found by dividing by 12, the interval between July 13 and July 14d. 13h. 30". In general the correction of Table XLV. may be neglected in calculating the places of the planets. In the above rule the intervals of time at which the longitudes, &c. are marked in the N. A. are supposed equal. If that should not be the case, the correction Table XLV. may be neglected, on account of the trouble of calculating it.

PROBLEM II.

To find the horary motion of the moon in long, lat. &c. at any given time at Greenwick. RULE.

Take from the Nautical Almanac the four longitudes, latitudes, &c. two immediately preceding the given time at Greenwich, and two immediately following. Prefix the sign + to the southern latitudes or declinations, and the sign - to the northern. Then find the first and second differences, the Arch B, and the time T, as in Problem I. The mean of the two first differences, noticing the signs as in algebra, will be the approximate motion in 12 hours.

To the proportional logarithm of one fourth part of the time T add the proportional logarithm of the arch B, the sum will be the proportional logarithm of the correction of the approximate motion, to be applied to it with the same sign as the erch B, and the corrected motion of the moon in 12 hours will be obtained,* which divided by 12 will give the horary motion.

EXAMPLE I.

Required the horary motions of the moon in longitude Dec. 12, 1808, at 15h. 48' 29" and 17h. 1' 29" app. time at Greenwich.

This corresponds to Example I. preceding, in which T=3h. 48' 29" and 5h. 1' 29'. The two first differences in longitude 70 6' 16" and 70 11' 18"; their mean 70 8' 47" is the approximate motion in 12 hours, and the arch B is 4'54'.5. The rest of the calculation is as follows.

At 16h, 48' 29" T=3h, 48' 29" Arch B 4' 54" 5 P. L. 1,5644 † T 57 7 P. L. 4985	At 17k. 1' 29" T=8k. 1' 29". 1.5644 † T 1k. 15' 22" P. L. 5781
Corf. + 1 332.0629	Corr. + 2 3 P. L. 1.9425
Approx. mbt. 7 8, 47	Apprex. mot. 7 8 47
Mot. 12 hours 7 10 20	Mo. 12 hours 7 10 50 .
In 1 hour 35 51.7	In 1 hour 85 54.2

In a similar manner, if the horary motion in latitude was required at 12d. 17h. 33, the two first differences in latitude are — 34' 21", and — 36' 45", their mean — 35' 33" is the approximate motion in 12 hours. The correction found by the above rule with the time T. 5h. 33' and the arch B=—2'.6".5 is — 59", whence the true motion in 12 hours is — 36' 33" which divided by 12 gives the horary motion — 3' 2".7. The negative sign — indicates that the north polar distance is decreasing, the positive sign + that it is increasing. In the present example the north polar distance was decreasing, and as the latitude was south, it was also decreasing, as is evident.

EXAMPLE II.

Required the horary motions of the moon in longitude June 16, 1806, at 2h. 49' 50'. and 5h. 34' 6".6 app. time by astronomical computation at Greenwich.

This corresponds to Example II. preceding, in which T=2h. 49 50".1 and 5h. 34'6".6; the two first differences, are 7° 17'21" and 7° 20'53", the mean of which 7° 19'7" is the approximate motion in 12 hours, the arch B is + 3' 7".

^{*}The motion in 12 hours thus obtained, which for distinction will be called the arch M, is not per accurate, since the field and higher orders of differences are neglected; but the horary motion des therefrom, is abundantly sufficient for the purpose of projecting an eclipse or occultation. When go accuracy is required, site third differences may be taken into account in the following masser. By found the second differences may be taken into account in the following masser. By found the second differences may be taken into account in the following masser. By found the second and take of the the corresponding reportection which is to be increased by one shath of the arch h, without not ng the signs. To the quantity thus found is to be prefixed a sign differented of the arch h, such then it is to be applied to the arch M with its sign to obtain the true masser in 12 hours. Thus in the above example the second differences of longitude are + 5°2" + 4°4". Tracting the former from the latter, leaves the third difference or arch b = -15". Corresponding this amount of the time T Sh. 48°29" in Table KLV. is 1".6 which increased by one first of b = 2°. 5 give sought correction. 4".1 or 4", to which must be prefixed the sign + (because the sign of b = 2°. 5 give sought correction. 4".1 or 4", to which must be prefixed the sign + (because the sign of b is segar and in the above example for finding the horary motion in intuitede, the two accounts of the motion in 12 hours = 36° \$2" in + 10° making 1" or 3' 3' 5 per hour.

At 2b. 49°50" 1=T Arch B= + 3' 7" † Time T= 42 27	P. L. P. L.	1,7616 6274	Åt 5h. 34′ 6″.&=T ¼ T=1h. 23′ 32″ P. L.	1.7610 3334
Correction + 041	P. L.	2.3890	Correction + 1 27 P. L.	2.0950
Approx. mot. 7 19 7 Mot. in 12 hours 7 19 51	• •		Mot. in 12 Isours 7 20 34	:
Mot. in 1 hour 36 80.2	·.		Mot in 1 hour 36 42.8	

EXAMPLE III.

Fraction of the moon in right ascension in 12 hours, supposing it to increase uniformly with the velocity it had July 4th, 1908, at 9h. 20' app. time at Greenwich, by astronomical computa-

July 3, mid.	D R. A 225°. 57'	70	29	B= 4° 50° ‡ T=2h. 21 50	P. L. P. L.	1,6021 1045
4, noon	233 28 . 241 01	A 7 7	35 38	Correction + 3.52 Approx. mot. 7 52 0	P. L.	1.7066

True mot. in 12h. 7 55 32 or 90'.22".1.

In this example T=9h. 26' and the approx. motion is the ball sum of 7° 29' and 7° 35'.

REMARKS.

1. When it is required to find the motion of the moon in any given interval of time. the motion in 19 hours must be found for the middle of that interval.

In calculating an occultation of a star by the moon, the relative horary motion in longitude is the same as the horary motion of the moon, because the star is at rest; but in calculating a solar eclipse, the sun's horary motion must be found in page III, of the Nautical Almanac, and subtracted from the moon's horary motion in longitude, the remainder will be the horsey motion of the moon from the sun in longitude. Thus on the 16th of June, 1806, the sun's horary motion was 2'23".1, which subtracted from the horary motions found in Example II. 36'. 39".2 and 36'. 42".8 leaves the corresponding horary motions of the moon from the sun in longitude 34' 16".1 and 34' 19".7.

As the sun has no sensible motion in latitude, the relative horary motion of the moon from the sun in latitude, is the same as the true horary motion of the moon in

The horary motion of a planet may be found in a sintilar manner, making use of the arches A. B. T. found as in Remark 2, Problem I. Thus if the horary motion of Jupiter was required July 14, 1808, at 13h. 30', the arch B= -7' T=3h. 7' 30", and the approximate motion in the interval 6 days is the half sum of the two first differences—3' and—11', namely—7' 0". The certain found as in the adjoining calculation is—1' 1 T=46 52 P. L. 5844 49", hence the motion in 6 days is 8' 49", whence the horary motion is 3",67. The negative sign in-Corr. 1.9946 dicates that the motion is retrograde, or contrary to the order of the signs: in this case the relative mo- Ap. mo. - 7 tion of the moon from the planet in longitude would be found by adding their horary motions, because the Motion — 8 49 in six days, motion of the moon is always direct. Similar remarks may be made in finding the ho-

rary motion of the moon from the planet in latitude.

PROBLEM III.

To find the time of the ecliptic conjunction or opposition of the moon with the ear, a planet, or a fixed star.

The time of the ecliptic conjunction of the sun and moon is the same as the time of new moon given for the meridian of Greenwich in the first page of the month of the Nautical Almanac. Thus in January 1803, the ecliptic conjunction is on the 27th day at 4h. 9 apparent time at Greenwich. The times of the ecliptic conjunction of the moon and those fixed stars with which there may be an occultation, are also given in the same page, being marked with Bayer's characters of reference. The time of conjunction is placed first, then the characters of the moon and star, or moon and planet. Thus in 1808, December 12d. 17h. 33') a M, signifies that on the 12th day of December at 17h. 33' apparent time at Greenwich, the moon was in ecliptic conjunction with the star Spica, whose character is a Π , and that there might be an occultation of that star. Also, December 15, 1808, 5h. 53' D is signifies that at that moment apparent time at Greenwich, the moon and Saturn were in ecliptic conjunction, and there might be an occultation of that planet. These times being reckoned according to astronomical computation, and in calculating them, no attention is paid to the parallaxes. The time of the ecliptic opposition of the sun and moon is the same as at the time of full moon given in the same page of the Nautical Almanac. Thus the full moon or ecliptic opposition in May, 1808, was 9d. 19h. 39 at Greenwich. Digitized by GOOGIC

The time of the ecliptic conjunction, as given in the Nautical Almanac, is easily computed from the geocentric longitudes of the objects; and as it may sometimes be required to seconds, the rule is here inserted, adapted to the calculation of the conjunction of the sun and moon, which, with a slight modification, will answer for any of the planets.

RULE.

Take from the Nautical Almanac the two longitudes of the sun and moon at the noon and midnight* preceding the time of the conjunction, and the two immediately following. Subtract the longitudes of the sun from those of the moon, noting the signs as in algebra, the remainders will represent the distances of the sun from the moon on the celiptic. Subtract each of these from the following to obtain the first differences, and call the middle term the arch A, subtract each of these differences from the following for the second differences, and take their half sum or mean for the arch B, noting the signs as in algebra.

To the constant logarithm 4.63548 add the arithmetical complement of the log, of the arch A in seconds, and the log, of the second of the above found distances in seconds, the sum, rejecting 10 in the index, will be the logarithm of the approximate value

of Tin seconds.

With this time T at the side of Table XLV. and the arch B at the top, find the equation of second differences, the logarithm of which added to the two first logarithms used in finding T, will, in rejecting 10 in the index, give the logarithm of the correction of the approximate time T in seconds, to be applied to it with the seme sign as the arch B, and the apparent time of the conjunction at Greenwich, counted from the second neon or midnight, taken from the Nautical Almanac, will be obtained. From which the time of conjunction under any other meridian may be easily obtained, by adding to it the longitude in time when east, or subtracting when west.

REMARK I. When the time of the ecliptic conjunction of the moon and a planet is required, the longitudes of the planet must be found by Problem I. for the noon and midnight immediately preceding, and those immediately following the time of the conjunction, and these are to be used in the above note instead of the sun's longitudes. If the ecliptic conjunction of the moon with a fixed star is required, its longitude must be found in Table XXXVII. and corrected for the equation of the equinoxes and aberration by Tables XL. XII. as shown in the explanation of those Tables. This lengitude

is to be used instead of the sun's, in the above rule.

Remark II. By the same rule, the time, when the moon is at any distance from the sun, may be found, by increasing the sun's longitudes given in the N. A. by the quantity the moon is supposed to be distant from the sun, counted according to the order of the signs. Then supposing a fictious sun to move so as to have these increased lengitudes at the corresponding times, and finding by the above rule the time of conjunction of the moon with this fictious sun, which will be the sought time when the moon is at the proposed distance from the sun. Thus, to find the time of the first, second, or third quantiter of the moon, the sun's longitudes must be increased 3, 6, or 9 signs respectively (rejecting as usual '12 signs when the sun exceeds that quantity.) Thus, if the first quarter of the moon which happened after midnight, July 29, 1808, was required. The sun's longitudes increased by 3 signs give the longitudes of the fictious sun July 29d. th; 29d. 12h; 30d. 0h; and 30d. 12h. respectively, 7s. 6° 5′ 44″, 7s. 6° 34′ 26″, 7a. 7° 3′ 9″, and 7s. 7° 31′ 51″. The longitudes of the moon corresponding are 6s. 23° 49′ 19″, 7s. 0° 53′ 34′, 7s. 7° 57′ 30″, and 7s. 15° 0′ 57″. Hence the time of the conjunction of the moon with the fictious sun found by the above rule, was July 29d. 22h. 21′ at Gracawich, which is the time of the first quarter required. In a similar manner, by increasing the longitudes of a planet or a star, the time may be found when the moon is at any proposed distance from it.

EXAMPLE.

Requ	ired the	time of	the eclip	tic conj	unctio	n of the	e aus e	nd mo	on, in Ja	n. 180	16 ?
. 18	08, Jan.) Lon	g; OLo	ng. Di	stances	; 1	1st di	e.	24.	di st .	
26d.	12h.	9 27 58	46-10 5	57 26 -8	0.4	0	•	, ,,	1	•	••
	0	10 4 25	2-10 6	27 56 -2	2 5	4 .	+ 5	57 46	1	3	15
	12	10 10 50	3-10 6 41-10 7	58 26 + 3	bt 9	77 A=	= - 1-5 ∤	54 S 1		3 ·	
28	0	10 17 11	41-10 7	28 551 🕂 9	42 4	16.	+5	51 00	1 19-	 3	18
				Ce	mistant	4.63548					4/8648
λ		1"= 212		log. co		5.67221			4		5.67521
2 dis.	2 2 5	4 = 78	74	log.		3.86170	Tab. X	LV. Co	r. 22 '.4 lo	g.	1.35025
T	14976"	= 4b.	9' 86"	-		4.17538	Corre	tion 45	" lo	f .	1.65794
Correction	100		-45								

Conjunction 4 8 51 past noon Jan. 27, at Greenwich, apparent time, which agrees nearly with 4h. 9' marked in the Nautical Almanac. The time of conjunc
"The sun's longitude at midnight is the mean of the longitudes on the preceding and following socra-

tion under any other meridian, as for example 30° W. is found by subtracting the longitude Sh. from 4b. 8' 51", which leaves 2h. 8' 51". If the longitude had been 30° E.

the time of conjunction would have been 6h. 8' 51".

The usual method of calculating the parallaxee in eclipses of the sun or occultations, is that by the nonagesimal or minetieth degree of the ecliptic above the horizon. Seveis that by the nonagramus or minetized degree of the exciptic and very left horizon. Several methods have been proposed for calculating the altitudes and longitudes of his point, which are required at each of the phases. The following, which is an improvement on that given in La Lande's Astronomy, seems well adapted to the purpose, since several of the logarithms are the same at each of the phases, which much abridges the calculation, and on this account it admits of considerable simplification, by a table like that on page 578. The method of making these calculations will first be given at full length, and then in the abridged form, by means of the proposed table.

PROBLEM IV.

Given the apparent time at the place of observation counted from noon to noon, according to the manner of astronomers, the sun's right ascension, and the latitude of the place, reduced on account of the spheroidal figure of the earth, by subtracting the reduction of latitude, Table XXXVIII. To find the ultitude and longitude of the nonagesimal degree of the ediptic.

RULE NOT ABRIDGED.

Add 6 hours to the sum of the sun's right ascension and the apparent time of observation, and call the sum the time T, rejecting 34 hours when it exceeds that quantity. Seek for this time in the column of hours of Table XXVII. supposing that marked A. M. to be increased by 12 hours, as in the astronomical computation. The corresponding log. co-tangent being found, is to be marked in the first and second columns, as in the following examples.

If the reduced latitude is north, subtract it from 90°; if south, add it to 90°, the sum or difference will be the polar distance. Take half of this, and half the obliquity of the ecliptic, and find their difference and sum. Place the log. co-sine of the difference is the first column, its log. sine in the second column: The log. secant of the sem in the first column, its log. co-secant in the second column, and its log. tangent

in the third.

The sum of the logarithms in the first column, rejecting 20 in the index, will be the log. tangent of the arch G. The sum of these in the second column, rejecting 20 in the index, will be the log. tangent of the arch F. These arches being less than 90° when the time T is found in the column A. M. otherwise greater. This rule is general except in places situated within the polar circles. Within the north polar circle the supplement of F to 360° instead of F, must be taken; within the south polar circle the supplement of G to 180° must be taken instead of G, the other terms remaining un-F, G thus found, and 90°. Rejecting 360° when the sum exceeds that quantity.

Place in the third column the log. co-sine of G, and the log. secant of F, the sum of the three logarithms of this column, rejecting 20 in the index, will be the log. tangent of

half the altitude of the Nonagesimal.

EXAMPLE.

Required the altitude and longitude of the Nonagesimal at Salem, in the reduced latitude 42° 22' 4" N. June 15, 1806, at 22h. 6' 18". I apparent time by astronomical computation, when by the Nautical Almanac the sun's right ascension was 5h. 36' 50". and the obliquity of the ecliptic 23° 27' 48".

The sum of the apparent time, sun's right ascension and 6 hours, rejecting 24 hours, is 9h. 43' 8".1=T. The polar distance is 47° 37' 56", its half is 23° 48' 58", and the half obliquity 11° 43' 54", hence their difference is 12° 5' 4", their sum 35° 32' 52". The rest of the calculation is as follows:

Col.		Col.			Ćol. 3.	
Diff. 12° 5′ 4″ Sum 36 52 52 T. Sh. 48′ 8′.1 P. E.		Sine Co-secant	9.\$2068 10.36564 9.48826	G. F.	Tangent Co-sine Secant	9.85403 9.97215 10.00265
G. 159° 42′ 0′′ F. 173 40 51 90	Tan. 9.56810	F. Tan.	9.04468	ł	25''=Tan.	9.82883
Sum 63 22 81 reject	ing 360°, is longitude	Nonages.	,	67 58	50 = Air. 1	Nonang.

The two upper logarithms of the first and second columns, and the upper logarithm of the third column, vary but little in several centuries; and as these numbers occur

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TO FIND THE ALTITUDE AND LONGITUDE OF THE NONAGESIMAL.

in calculating a partial eclipse or occultation, and four times in a total, or annular se or transit, it will tend considerably to abridge the calculations, to have a table the following, containing their values for various places, for the obliquity 23° 27' 40", the variations for an increase of 100" in the latitude or obliquity. The logarithms . C, of the table, were calculated in the following manner.

north latitudes subtract the reduced latitude from 90°, in south latitudes add the red latitude to 90°, the sum or difference will be the polar distance; take half is and half of the obliquity of the ecliptic 11° 43′ 50″, and find the sum and

g. A is equal to the log. co-sine of the sifference added to the log. secant of the sum, rejecting 20 in the index.

g. C is equal to the log. tangent of the sum.

g. B is equal to the log. tangent of the difference, increasing the index by 10, less

us for Salem in the reduced latitude 42° 99' 4'', the half polar distance is 23° 48' 58'', alf obliquity 11° 43' 50'', the difference 12° 5' 8'', the sum 35° 32' 48''.

ence12°			Tang.+10=19.33065 Tang.=C = 9.85402
Sum. A	ļ	 0.07993	DiffB 9.47663

this way the logarithms may be found for places not included in the table. The es for an increase of 100" in the latitude or obliquity, are found by repeating the tion with these increased values, and ascertaining the corresponding changes in alues of A, B, C. These logarithms are given to six places of figures, though in al five will be quite sufficient, since the latitude and longitude of the nonagesimal rely required to a greater degree of accuracy than 10".

Places.	Reduced Latitude North.			Α.	Var A. +100". Lat. Obl.		В	Var. B. +100". Lat. Obl.		O:	Var. C. +100". Lat. Obl.	
	0	1	11			+	14.4	4	+	115,200		4
Albany	12	27	13	0.079670	53	97	9,475733	298		9,858328	228	220
Berlin	52	20	24	0.061608	49	75	9,324135	618	1099	9.771197	240	240
Cambridge, E.	52	1	28	0.062166	49	- 76	9.331054	600	1080	9.778925	240	240
Cambridge, A.	42	-12	2	0.080150	52	97	9.478363	288	733	9,855355	222	222
Dublin obs.	53	12	7	0.060090	48	73	9.304166	670	1155	9.763705	242	242
Edinburgh	55	46	2	0.055618	47		9,285401	878	1376	9.741011	249	249
Greenwich obs.	51	17	28	0.063466			9.346396	562	1038	9.780232	238	23
lavanna	23	3	34	0.120000				95		10.003045	210	210
Linderhook -	42	11	37	0.080163		98		289			222	222
Lancaster	39	51	18	0.084648	54			249			219	219
Leon I. obs.	36	16	62	0.091680				202		9.902005	216	216
London	51	19	29				9.345714	564		9,779944	238	235
Natchez	31	17	36	0.101899				152	577	9,940447	212	217
Oxford obs.	51	34	28	0.062963			9.340586		1054	9.777800	239	25
Paris	48	38	51	0.068207	50	83	9.394413	452		9.802627	233	255
Philadelphia	39	45	44	0.084828	53		9 501872	248		9.874738	219	218
Richmond obs.	51	16	56	0.063482	49	78	9,346576	562	1038	9,780308	238	231
Rutland	13	24	52	0.077866	52	95	9.465330	312		9,845648	224	224
Salem	42	22	4	0.079832	52	-98	9 476637	291	731	1.854016	222	222
Place Prob. VII.	19	52	38	0.127485	66	157	9.607602	78	500	10.027183	211	211

se logarithms are calculated for the obliquity 23° 27′ 40″. The columns marked epresent the variations of A, B, C, for an increase of 100″ in the reduced lat. llumn Obl. represents the variations of A, B, C, for an increase of 100″ in the ty of the ecliptic. The signs must be changed if the latitude or obliquity is less hat used in calculating the table.

EXAMPLE.

uired the values	of A, B, C, for Salem	, when the obliquity	7 is 23°	27' 48" 1
ır numbers	0.079832	9.476637		9.854016
>r + 8″ Obl.	+8	58		+ 18
t values	A=0.079840	B=9.476579	•	C=4.5¥034

Abridged method of calculating the altitude and longitude of the Nonagesimal, by the preceding Table.

Add together the sun's right ascension, the apparent time at the place of observation, (counted from noon to noon) and 6 hours, the sum, rejecting 24 or 48 hours if greater than those quantities, is to be called the time T; this is to be sought for in the column of hours of Table XXVII. supposing the column marked A. M. to be increased 12 hours, as in the astronomical computation.* The corresponding log. co-tangent, added to the log. A of the Table, gives the log. tangent of the arch G; this added to the log. B of the Table, rejecting 10 in the index, will be the log-tangent of the arch F; these arches being less than 90° when T is found in the column A. M. otherwise greater. † [This rule is general, except in places situated within the polar circles, which is a case that very rarely occurs. Within the north polar circle, the supplement of F to 360°, is to be used instead of F; within the south polar circle, the supplement of G to 180°, is to be taken instead of G, the other terms remaining unaltered.] Then the longitude of the Nonage-simal is equal to the sum of the arches F, G, and 90°, neglecting as usual 360° when the sum exceeds that quantity.

To the tabular log. C, add the log. co-sine of the arch G, and the log. secant of the arch F, the sum, rejecting 20 in the index, will be the log-tangent of half the altitude of the Nonagesimal.

EXAMPLE I.

Required the altitudes and longitudes of the Nonagesimal at Salem, June 16, 1806, at the times of the beginning and end of the eclipse, calculated in Problem VI.?

Beginning	of the Eck	pse.	. [_	,	,, 1	End of	the Ecliq	15G.	
h. , , , , , , , , , , , , , , , , , , ,					ь. 5 0 6	87	18.5	⊚ R. As Apperen A	scension of time 6.07984		
T 9 43 8.1 Co-tan.	9.48826			T	12	27	53.1	Co-tan.	8.78470		
G 159° 42′ 0″ Tan. 99 B	9.56810 9.47658		9.9 7216 9.854 0 8		4° 90	11'	13"	Tap. B	8.86454 9.47658	0-a	8.854 05
F 175 40 31 Tan.	9 04468	Sec 1	0.00265	F	1	15	23	Tan.	8.84112	Sec.	10.00819
63 22 31=lon, N.	53 59 25	Tan.	9.82883		95	26	36=	= lon. N.	55 28 53	Tan.	9.85297
Altitude Nonages	67 58 50	-	V A M		Mititu		Non	rges	70 57 46		

EXAMPLE II. Required the altitudes and longitudes of the Nonagesimal at the times and places mentioned in the Example of Problem VII.?

	•		,,		ersion.	1.00.		İ		,		E	nersion.		
	h 17 16 6	20 57	59	⊙ R. Asc Apparent A			•		h 17 18 6	21	12.5		Ascension nt time 0.12748	,	
Ŧ	16	18	28	Co-tan.	9.80098			г	17	31	41.5	Co-ta	n. 9.946 22		
G	40 90		7'	Tan. B	9.92846 9.60761	Co-6. C-	9 88233 10.02718		49° 90	50	18"	Tan. B	10 07370 9.60761	Co s.	9.89563 10.02718
F	18	57	48	Tan.	9,53607	Sec.	10.02423	F	25	38	40	Tan.	9.68131	Sec.	10.04504
	149	15	55:	=lon. N	. 40 S8 46	Tan.	9.93374		165	28	58=	ion. N.	37 17 39	Tan.	9.88175
1	Alcit	ude	No	nages	81 17 32				Altin	iq.	None	grs	74 35 18		

In these calculations it is usual to take the sun's right ascension, and the apparent times, to tenths of a second, and to take proportional parts for the seconds and tenths in the nearest logarithms. Thus in Example I. in finding the log. co. tang. of 9h. 43' 8".1; the nearest logarithms are 9.48849, 9.48904, corresponding to the 9h. 43' 4", and 9h. 43' 12". These logarithms differ 45, the times 8", and the difference between 9h. 43' 4", and 9h. 43' 8".1, is 4".1. Hence 8": 45:: 4".1: 23 the correction to be subtracted from the first log. 9.48849, (because it is decreasing) to obtain the sought log. co-tang. 9.48826.

Thus if the time T is 5 hours, it must be called 5h. P. M. If T is 14 hours, it must be called 2h. A. M. In making use of a common table of logarithms, you must turn the time T into degrees, and make use of the log. co-tangent of its half.

the log-co-tanguart of its nam.

† The arches F, G, are acute when the time T is found in the column A M. otherwise obtase. This is easily remembered from the circumstance that a is the first letter of acute and A. M. Some writegs have not taken notice of the cases of the values of F, G, within the polar circles.

† Strictly speaking, the quantity thus obtained is the distance between the north pole of the cellptic and the zenith of the place, which, in southern latitudes, and between the tropics, is frequently the supplement of the attitude of the Nonagesimal. The a ove form is made use of the collection of the strictles of the collection of the attitude of the Nonagesimal, or its supplement, is made use of parallexes. It is in Table XLIV.

PROBLEM V.

Given the altitude and longitude of the Nonagesimal; the longitude, latitude and horizontal parallax of the moon, and the latitude of the place of observation, to find the moen's parallax in latitude and longitude.

RULE BY COMMON LOGARITHMS.

From the horizontal parallax of the moon, subtract its correction from Table XXXVIII. corresponding to the latitude of the place, the remainder, in occultations of a fixed star, will be the reduced parallax; but in solar eclipses this quantity is to be diminished by the sun's horizontal parallax, 8". 8* to obtain the reduced parallax.

To the log:rithm of the reduced parallax in seconds, add the log. sine of the altitude of the Nonagesimal, and the log secant of the moon's true latitude, the sum, rejecting 20 in the index, will be a constant log. From the moon's true longitude,† increased by 360° if necessary, subtract the longitude of the Nonagesimal, the remainder will be the moon's distance from the Nonagesimal, which if less than 180° is to be called the arch D, otherwise its supplement to 360° is to be called the arch D. To the constant logarithm, add the log. sine of D, the sum, rejecting 10 in the index, will be the logarithm of the approximate parallax in longitude in seconds, which add to the arch D, then take the logarithm of the sum, and add it to the constant logarithm, rejecting 10 in the index, and the logarithm of the corrected parallax will be obtained. This will in general be sufficiently exact, but when great accuracy is required, the operation may be again repeated, by adding the arch D to the corrected parallax; then to the log. sine of the sum add the constant logarithm, rejecting 10 in the index, and the logarithm of the parallax in longitude P will be obtained. This is to be added to the true longitude of the moon when her distance from the Nonagesimal is less than 180°, otherwise subtracted to obtain her apparent longitude.

If the true latitude of the moon is south, prefix the sign + to it, if north the sign -Then to the logarithm of the reduced parallax in seconds, add the log. co-sine of the altitude of the Nonagesimal, and the log, co-sine of the moon's apparent latitude, the sulm, rejecting 20 in the index, will be the logarithm of the first part of the parallax in latitude in seconds, to which prefix the sign + when the altitude of the Nongesimal is less than 90°, otherwise the sign —, this added to the true latitude of the moon, due regard being had to the signs, will give her approximate latitude.

To the logarithm of the reduced parallax in seconds add the log. sine of the altitude of the Nonagesimal, the log. sine of the moon's approximate latitude, and the log. cosine of the sum of the arches D and & P, the sum, rejecting 30 in the index, will be the logarithm of the second part of the parallax in latitude in seconds, to which prefix the sign — when the arches $D + \frac{1}{2}P$, and the approximate polar distance? are both greater or both less than 90°, otherwise the sign +, this term connected with the approximate latitude will give the apparent latitude of the moon, ** which will be south if + north if -The moon's true latitude subtracted from her apparent latitude, noticing the signs, will give the parallax in latitude.

BY PROPORTIONAL LOGARITHMS.

The above rule will answer in calculating by proportional logarithms, with the following alterations. When the log. sine occurs, read log. co-secant; for log. co sine read log. secant; for log. secant read log. co-sine; and for log. co-secant read

*** This rule gives the apparent latitude in all cases, but it may not be amiss to observe, that in several late publications the cases where the moon is between the zenith and the elevated pole are by mistake neglected.

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^{*} This is the mean value of the sun's parallax, and may be used instead of the true parallax, which varies from 8".7 to 8".9. The true solar parallax at any time may be found by subtracting the logarithm of the sun's distance, given in the Nautical Almanac, from the log. of 8".8, increasing the index by 10 when necessary, the remainder will be the logarithm of the sought parallax in seconds.

§ Corrected for the errors of the tables, when known.

§ This sum D + cor. phr. is nearly equal to D + P the apparent distance of the moon's B. D.

§ In solar eclipses the apparent fatitude is so small that its log, sine may be put equal to 10.00000. In occulations you must calculate the first part of the parallax in alittude by approximation, making use of the true latitude instead of the apparent in the above rule, and deducing the apportance value of the first part of the parallax: this applied to the true latitude will give the approximate value of the first part of the parallax: this applied to the true latitude will give the approximate to a sufficient elegree of exactness.

§ The apparent polar distance is found by adding + 90° to the approximate latitude, due regard being had to the signs. To be perfectly accurate, the opportunit instead of the approximate latitude ought to be made use of in this part of the calculation, and the logarithms of this term ought to be increased by the log, secant less radius of \(\frac{1}{2} \) Fy to these corrections are too small to affect the result. In calculating the second part of the parallax in latitude, it will be sufficient to take the logarithms to three or four pieces of the decimals.

**This rule gives the apparent latitude in all cases, but it may not be amise to observe, that in several

log. sine. The parallaxes may be calculated to the nearest second by proportional logarithms. When greater accuracy is required, common logarithms must be made use of.

To illustrate this rule, the following examples, corresponding to the times of the begin-

To illustrate this rule, the following examples, corresponding to the times of the beginning and end of the total eclipse of the sun, of June 16, 1806, as observed at Salem, are given. The elements necessary for this purpose have already been calculated in Problems I. and IV. For greater accuracy the longitudes and latitudes of the moon are corrected for the errors—58".5 in longitude, and — 11".4 in latitude, which were found by comparing several observations of the eclipse made at different places.

EXAMPLE I.

Given the altitude of the Nonagosimal 67° 58′ 50″, its longitude 63° 22′ 31″, the longitude of the moon 83° 49′ 3″.5 her latitude 24′ 27″.4 N. her horizontal parallax 60′ 24″.1, the latitude of the place of observation 42° 33′ 30″: required the parallaxes in longitude and latitude?

The correction in Table XXXVIII. corresponding to the latitude 42° 33′ 30″, and parallax 60′ 24″.1 is 5″.6, this and the sun's horizontal parallax 8″.8 subtracted from the D's horizontal parallax 60′ 94″.1 leaves the reduced parallax 60′ 9″.7=3609″.7. The longitude of the Nonagesimal 63° 28′ 31″ subtracted from the moon's longitude 83° 49′ 3″, leaves the moon's distance from the Nonagesimal, 20° 26′ 32″ equal to the arch D, because less than 180°.

Calculation by common Logarithms.

Reduced par. Alt. Noneg. D's true lat.	3609" 7 67 58 50 24 27.4	Log. Sine Sec.	3.55747 9.98710 10.00001	Reduced par. Alt. Nonag.) 's app. lat.		809 7 5 8 50	Log. Co-si. Co-si.	3.55747 9.57394 10.90000
Constant log.	20 28 52	Sine	3.52458 9.54315	i prt. p. 1353".3 =)'s true latitude		22° 33′′.3 24 27.4		3.15141
Appr. par. 110	89''=19 2 9	Log.	3.06773	D's approx. lat.	_	1 54.1	Sine	F4 6.743
D + Appr. par. Constant log.	20 46 1	Sine	9.54970 3.52458	Reduced parallax Ait. Nonag. D + 1 P		20 \$6 25	Log. Sine Co-si.	3.557 9.967 9.971
Cor. par. =11	87"=19 47	Log.	3.07428	• •			٠.	
D + cor. par. Constant log.	20 46 19.	Sine	9.54980 3.52458	2 part par. Approx. lat.	_	1".7 1 54 1	Log.	0.238
Par. long. P 1186'	′.8==19′ 46.8 _,	Log.	3.07438	D's app. lat.	-	1 55 8	or 1'	55".8 N.
'D's true longitude	83 49 8.5							
)'s app. long.	84 8 50.5		;	I				

EXAMPLE II.

Given the altitude of the Nonagesimal 70° 57′ 46″, its longitude 95° 26′ 36″, the longitude of the moon 85° 29′ 33″.6, her latitude 15′ 10″.4 N. her horizontal parallax 60′ 27″.0, the latitude of the place of observation 42° 33′ 30″. Required the parallaxes in longitude and latitude?

The correction in Table XXXVIII. corresponding to the latitude 42° 33′ 30″ and parallax 60′ 27″, is 5″.6, this and the sun's horizontal parallax 8″.5, subtracted from the moon's horizontal parallax 60′ 27″.0 leaves the reduced parallax 60′ 12″.6. The longitude of the Nonagesimal 95° 26′ 36″, subtracted from the moon's longitude increased by 360°, viz. 445° 29′ 33″, leaves the moon's diplanee from the Nonagesimal 350° 2′ 57″, the supplement of which to 360° is 9° 57′ 3″, equal the arch D.

By proportional Logarithms.

		- j p p						
Reduced par. Alt. Nonag. D's true lat.	60' 12''.8 70 57 46 15 10.4	P. L. 0.4758 Co-se. 10.0244 Co-si. 10.0000	Reduced par. Alt. Nonag. B's app. lat.		70 57	12".6 1 6	P. L. Sec. Sec.	0.4756 10.4865 10.0000
Constant log.	9 57 8	0.5000 Co-se, 10.7624	1 part par. lat.	+	19 38	5	P. L.	0.9821
Approx. per.	9 50	P. L. 1.2624	D's true lat.	_	15 10	4		
The section			D's appr. lat.	+	4 28	1	Co-se.	12.886
D + appr. par. Constant log.	10 6 53	Co-se. 10.7554 0.5000	Reduced par- Alt. Nonag.	•		-	P. L. Co-se.	0.4756 10.0244
Corrected par.	10 0	P. L. 1.2554	D+PP 2 part par. lat.	_	10 2	5 4.4	Sec. P. L.	10.0067 8.8927
D + cor. par.	10 7 3	Co-se. 10.7553	- ber ber me	-				
Constant log.		0.5000	Approx. lat.	+	4 :	28.1		٠, ٩
Par. long. P.	10 0.0	P. L. 1.2553	Apparent lat.	+	-	32.5 o		T
D's true long.	85 2 9 32. 6				Digiti:	zed by (Goo	gle
) 's app. long.	85 19 32.6	i						

EXAMPLE III.

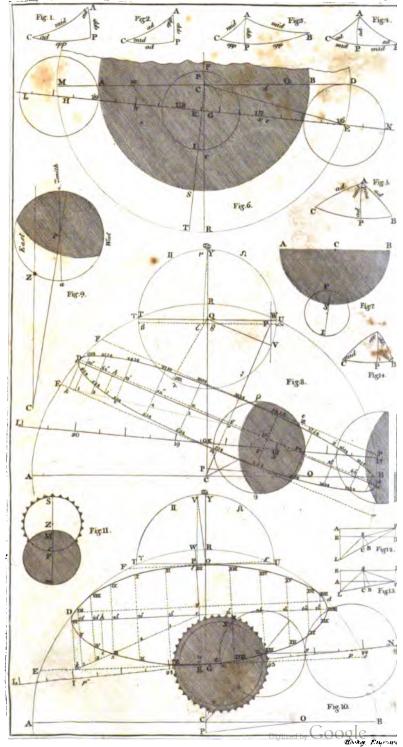
Required the parallaxes in longitude and latitude at the time of the occultation of Spica, Dec. 12, 1808, at the times and place mentioned in the Example of Problem FH.:

Immersion.

Reduced par-	59′ 50′′ A. P. L.	0.4782			Bec. 0.6199
Alt. Non.)'s true lat.	81 17 32 Co-se. 1 55 11 Cn-si	9.9998) app. lat.*		Sec. 10.0003
Constant D		4830 10.1108	1 part par. lat		P. L. 12984
Appr. par.	45 55 P. L.	5983	•	+ 1 55 11.0 + 2 4 14.8	Go-se, 11.4421
D + appr. par. Constant	51 37 56 Co-se.	10.1057 4880	Reduced par.		P. I. 0.4782
Corrected par-	46 25 P. L.	5887	Alt. Nonag. D + ½ P	51 15 13	Co-te. 10.0050 Sec. 10.3035
D + corr. par.	51 38 26 Co-se.	10 1056 4830	2 part par. lat		P. L. 21283
Par. long. P.	+ 46 25 P. L.	5886	D approx lat	+ 2 4 14.5 + 2 5 34.6 Se	omeths.
true long.	200 7 56.3			10 23.6	ora que
D app. long.	200 51 21.3		•		
_		T			
•		Emer	eun.		
Refluend par. Alt. Non.) true ist.	59' 53'' 0 P. L. 74 35 18 Co-sec. 1 51 29.1 Go-si.	0.4780 10.0159 9.5998	D app. lat.		0.4786 Sec. 10.5755 Sec. 10.0003
Alt. Non.	74 35 18 Co-sec. 1 51 29.1 Go-si.	0.4780	D app. lat. 1 part par. lat.	÷ 15' 54".2	Sec. 10.5755
Alt. Non.) true lat. Constant	74 35 18 Co-sec. 1 51 29.1 Go-si.	0.4780 10.0159 9.5998 0.4937	D app. lat. 1 part par. lat. D true lat.	+ 1 51 29.1	Sec. 10.5755 Sec. 10.0003 P. L. 1.0553
Alt. Non.) true lat. Genstant D	74 35 18 Co-sec. 1 51 29.1 Go-si. 35 22 38 Co-sec. 53 26 P. L.	0.4780 10.0159 9.5498 0.4937 10.2374	p app. lat. 1 part par. lat. 5 true lat. D approx. lat. Reduced par.	•	Sec. 10.3755 Sec. 10.0003 P. L. 1.0558 Co-sec. 11.4813 P. L. 0.4790
Alt. Non. b true lat. Constant D Appr. par. Dat. appr. par.	74 35 18 Co-sec. 1 51 29.1 Go-si. 35 22 38 Co-sec. 53 26 P. L.	0.4780 10.0159 9.5498 0.4987 10.2374 7\$11°	D app. lat. 1 part par. lat. D true lat. D approx. lat.	+ 1 51 29.1	Sec. 10.5755 Sec. 10.0003 P. L. 1.0558 Cosec. 11.4813
Alt. Non.) true lat. Genstant D Appr. par. Dat. appr. par. Communi	74 35 18 Co-sec. 1 51 29.1 Go-si. 35 22 38 Co-sec. 33 20 P. L. 35 56 4 Co-sec. 33 54 P. L.	0.4780 10.0159 9.5498 0.4937 10.2374 7511 10.2315 4937) app. lat. 1 part par. lat.) true lat.) approx. lat. Reduced par. Alt. Nonag. D + 1 P 2 part par lat.	1 51 29.1 + 2 7 25.3 35 39 35 + 1 44.2	Sec. 10.5755 Sec. 10.0003 P. L. 1.0838 Co-sec. 11.4813 P. L. 0.4730 Co-sec. 10.015.1
Alt. Non.) true lat. Genstant D Appr. par. Dal. appr. par. Genrected par. D + corr. par. Genstant Par. long. P.	74 35 18 Co-sec. 1 51 29.1 Go-si. S5 22 38 Co-sec. 33 20 P. L. 35 56 4 Co-sec. 35 54 P. L. 35 56 32 Co-sec. + 53 54 P. L.	0.4780 10.0159 9.5498 0.4987 10.2374 7S11* 19.2315 4937 7252	D app. lat. 1 part par. lat. D true lat. D approx. lat. Reduced par. Alt. Nonag. D + ½ P 2 part par lat. D approx. lat.	35 39 35 1 44.2 2 7 23.3	Sec. 10.3755 Sec. 10.0003 P. L. 1.0558 Co-sec. 11.4513 P. L. 0.4730 Co-sec. 10.0153 Sec. 10.0892 P. L. 20154
Alt. Non.) true lat. Constant D Appr. par. Par. appr. par. Constant Carrected par. D + corr. par. Constant	74 35 18 Co-sec. 1 51 29.1 Go-si. S5 22 38 Co-sec. 33 26 P. L. 35 56 4 Co-sec. 33 54 P. L. 35 56 32 Co-sec.	0.4780 10.0159 9.5498 0.4937 10.2374 7511* 19.2315 4937 7252 10.2314 4937) app. lat. 1 part par. lat.) true lat.) approx. lat. Reduced par. Alt. Nonag. D + 1 P 2 part par lat.	35 39 35 1 44.2 2 7 23.3 4 2 7 23.3 4 2 9 7.5 8	Sec. 10.3755 Sec. 10.0003 P. L. 1.0558 Co-sec. 11.4513 P. L. 0.4730 Co-sec. 10.0153 Sec. 10.0892 P. L. 20154

Having thus explained the method of calculating the parallaxes of the moon, it now remains to give the rules for finding the longitude by eclipses and occultations. The main object in these calculations is to determine from the observed beginning or end of the eclipse or occultation, the precise time of the ecliptic conjunction of the sun, or star and moon, free from the effects of parallax, counted on the meridian of the place of observation, since the difference of the times of conjunction, obtained in this mammer at two places, will be their difference of longitude. If the lunar and solar tables were perfeetly correct, the longitude might be determined by taking the difference between the time of conjunction, given in the Nautical Almanac, and that deduced from the observations of the eclipse or occultation; but it is much more accurate to compare the times deduced from observations actually made at the places for which the difference of longitude is sought. There are two different methods of finding the ecliptic conjunction, according as the latitude of the moon is supposed to be accurately known or not. If the latitude was given correctly by the lunar tables, or was accurately known by other observations, the ecliptic conjunction, and the longitude of the place, might be determined by each of the phases of the eclipse or occultation, by the method given in Problems VIII. and IX. But the moon's latitude not being generally given to a sufficient degree of accuracy, it is usual to combine together the observations of the beginning and end of the eclipse or occultation, or the beginning and end of total darkness in a total eclipse, or the two internal contacts of an annular eclipse, to ascertain the error of the amoon's latitude, by the method given in Problems VI. and VII. In making the calculations in these Problems, it will be necessary to know nearly the longitude of the place. In order to find the supposed time at Greenwich, so as to take out the elements from the Nautical Almanac: and if the longitude, deduced from the observation, should differ considerably, the operation must be repeated with the longitude obtained by this peration.

The moon's true latitude 1° 55′ 11″ must first be used, its log, secant bring 10,000c, which give the lat part par. 9′ 3′, which, added to the true latitude of the moon, gives the app. lat. nearly 2° 4′ 14′ the log, secant of which is 10,000s, as above. The calculation for the emersion in made in a similar manner.



PROBLEM VI.

Given the latitude of the place, and the apparent times of the beginning and end of a sofar eclipse, counted from noon to noon, according to the method of Astronomers, to find the

longitude of the place of observation.

In the rule for solving this Problem, references will be made to fig. 12, Plate XII. in which DSE represents a small arch of the ecliptic; S, the place of the centre of the sun, supposed at rest; F, L, the apparent places of the centre of the moon at the beginning and end of the eclipse respectively; FD, SC, and AEL, are perpendicular to DE, FA parallel to DE, and SB perpendicular to FL. Then it is evident that FD, LE represent the apparent latitudes of the moon, which fall below DL if south, above if north; and Si', SL represent the sums of the corrected semi-diameters of the sun and moon, at the beginning and end of the eclipse respectively.

RULE.

To the apparent times of the beginning and end of the eclipse add the estimated longitude of the place in time if it is west, but subtract if east; the sum or difference will be the supposed time at Greenwich, corresponding to which, in the Nautical Almanac, find, by Problem I. the moon's semi-diameter, horizontal parallax, longitude and latitude,* and the sun's semi-diameter, longitude and right ascension, also the moon's horary motion from the sun, by Problem II. Decrease the sun's semi-diameter 3d' for irradiation, and the remainder will be his corrected semi-diameter. Decrease the moon's semi-diameter 2" for inflexion, and to the remainder add the correction in Table XLIV. the sum will be the moon's corrected semi-diameter. Find also, in the Nautical Almanac, the obliquity of the ecliptic.

With these elements and the apparent time at the place of observation, calculate the altitudes and longitudes of the Nonagesimal, by Problem Iv. the parallaxes in longitude and latitude, and the moon's apparent longitudes and letitudes, by Problem V.

Take the difference between the apparent longit des of the moon at the beginning and end of the eclipse, and subtract therefrom the difference of the sun's langitudes at the same times, the remainder will be the relative motion in longitude DE or FA. The relative motion in latitude AL, is found by taking the difference of the moon's apparent latitudes at the beginning and end of the eclipse, if they are both north, or both south, but their sum, if one be north, the other south. From the logarithm of FA, increasing the index by 10, subtract the logarithm of M., the remainder will be the log-tangent of the angle of inclination DSB; this angle is to be taken greater than 90° when the moon's apparent latitude FD, at the reginning of the eclipse, is greater than at the end FL, otherwise less.; Then to the log co-secant of the angle of inclination, add the logarithm of the relative motion in longitude FA, the sum, rejecting 10 in the index, will be the logarithm of the apparent motion of the moon FL on her relative orbit. Then in the triangle SFL, the sides SF, SL, represent the sums of the corrected semi-diameters of the sun and moon at the beginning and end of the eclipse, and these with the relative motion FL, are given to find the angle FSB, (by Case VI. Obl. Trig.) Thus: to the log. ar. co. of FI., add the log. of the sum of SF and SL, and the log. of their difference, the sum, rejecting 10 in the index, will be the logarithm of the difference of the segments FB, BL; half of which, added to, and subtracted from half of FL, will give the two segments FB, BL, the greater segment being contiguous to the greater side, whether SF or SL. Then from the logarithm of the segment FB, increasing the index by 10, subtract the log. of SF, the remainder will be the log. sine of the angle FSB, which is always less than 90°:

nearly equal, they may be neglected, (as in the above rule) except in a case which very raselfloc-curs, namely, when the difference of SL. SF is greater than the difference of the two app. latitudes EL, FD, in which case the rule in this note must be made use of. Observing that the fractions EL FD

Corrected for the errors of the tables in longitude and latitude, when known.

^{**}Corrected for the errors of the tables in longitude and latitude, when known. It this correction must be found after the abitude and longitude of the Non-gesimal are calculated. It his rule is equally true whether the latitude be of the same or different names. If the latitudes are equal and of the same name, the angle DSB will be 90°. If they are equal, but of different names, the angle DSB may be taken acute or obtase, since in that case the angle PSB is 90°. Siricity speaking, when the points F, L. fall on different sides of the line DE, the angle DSB is greater or less than DE. FD EL

^{80°:} according as the expression — is greater or less than -- but as the divisors SL and AF are 8L

represent the quotients of the moon's apparent latitudes divided by the sum of the semi-dia-

meters of the san and moon.

I When SF, SL are equal, or their difference is so small that it may be neglected, the log. sine of the angle FSB may be obtained much more expeditiously by subtracting the log. of the sum of SF and SI, from the log. of FL increasing the index by 10. This melbox may almost always be made use of without much error. It is the rule adopted by Doctor Mackay in his treatise on longitude.

the difference between this and the angle of inclination DSB, will be the central angle

To the log. co-sine of the central angle, add the log. of the sum of the corrected semidiameters at the beginning of the eclipse SF, rejecting 10 in the index, the sum will be the logarithm of SD, the apparent difference of longitude of the sun and moon at that This is to be subtracted from the longitude of the sun at the beginning of the eclipse, if the central angle is less than 90°, but added if greater than 90°, the sum or difference will be the moon's apparent longitude: to this must be added the moon's parallax in longitude, when her distance from the Nonagesimal (found as in Problem V. by subtracting the longitude of the Nonagesimal from the moon's longitude, borrowing 360° when necessary,) is greater than 180°, otherwise the parallax must be subtracted; the sum or difference will be the moon's true longitude at the beginning of the eclipse.

Take the difference in seconds, between the sun's and moon's true longitudes at the

beginning of the eclipse, to the logarithm of which add the arith. comp. log. of the moon's horary motion from the sun* in seconds, and the constant logarithm 3.55630: the sum, rejecting 10 in the index, will be the logarithm of the time from the conjunction in seconds, which is to be added to the observed apparent time of the beginning of the eclipse, when the sun's longitude at that time is greater than the moon's true longitude, otherwise subtracted; the sum or difference will be the apparent time of the true ecliptic conjunction of the sun and moon at the place of observation. The difference between this and the time of conjunction at Greenwich, inferred from the Nantical Almanac by Problem III. will be the longitude of the place of observation. But if corresponding observations have been made at different places, it will be much more accurate to find the times of the conjunction at each place by the above rule, and the difference of those times will be the difference of meridians, if it does not differ much from the supposed difference of longitude. If there is considerable difference, the operation must be repeated, making use of the longitude found by this operation; and thus by successive operations, the true longitude may be obtained.

The long, of the place of observation being accurately known, the errors of the lunar tables in long. and latitude may be easily found. For the difference between the moon's true longitude deduced by the above method from the observations, and the longitude found from the Nautical Almanac, will be the error of the tables in longitude. To find the error in latitude, add the log. sine of the central angle DSF to the log. of the sum of the corrected semi-diameters at the beginning of the eclipse SF, the sum, rejecting 10 in the index, will be the logarithm of the moon's apparent latitude FD at that time, which will be south, if the point F falls below D, otherwise north. Take the difference between this and the moon's apparent latitude found by Problem V. if they are both north, or both south; but their sum, if one be north and the other south, and the error of the tables in latitude will be obtained.

REMARK.

The above rule will answer for deducing the longitude from the observed beginning and end of the internal contacts of a total or annular eclipse. The differences consist in reading the rule, beginning and end of the internal contacts, instead of beginning and end of the eclipse, and taking SF, SL equal to the differences of the corresponding semidiameters, instead of their sums.

EXAMPLE.

At Salem, in the latitude of 42° 33' 30" N. longitude by estimation 4h. 43' 32" W. from Greenwich, the beginning of the total eclipse of June, 1806, was observed at 15d. 22h. 6' 18".1, and the end at the 16d. 0h. 50' 34".6, apparent time, by astronomical computation. Required the longitude of the place of observation? Most of the following elements are calculated in Problems I. II. IV. V.

^{*} With the horary motion varies, it must be taken to correspond to the middle time between the beginning of the eclipse, and the conjunction or new moon.

† When the eclipse or occultation is nearly central or (in other words) when FD, EL are very small in comparison of SF, the latitude thus found cannot be depended on, as a small error in the times of observation, would produce a considerable error in the latitude. Indeed the case may occur when FD, EL are less than 30", that it may be uncertain whether the points F, L, fall above or below the line DE because the error of the lunar tables in latitude may sometimes be equal to 30". In this case the correct latitude of the moon may be found (1) By observations made at another place where the eclipse or occultation was not so central. (2) By the number of digits eclipsed, if it was a solar eclipse. (3) By the difference of decliantions of the moon und star observed before and after the immersion or emersion. (4) By the meridian attitude of the moon observed the same day, whence it may be found whether the moon was north or south of her place given by the tables. was north or south of her place given by the tables.

ELEMENTS OF THE ECLIPSE.	BEGINNING.	END.
	d. h. ' "	d. b. / "
Apparent times of observation	15 22 6 18.1	
Estimated longitude W. from Greenwich	4 46 32	
Supposed time at Greenwich	16 2 49 501	16 5 34 6.
O's witht assention	5 88 \$0.0	5 37 18.
Cat. of place 42° 83′ 30"—Reduc. Tab. XXXVIII. 11′ 26″	42 22 4	
UDBIQUITY of the ecliptic	23 27 48	9 1 11
D's long, by N. A. Err. Tab. 58' 5 True long. D Prob. I.	88 49 3.5	85 29 32.
Longitude of the Nonagetimal, by Prob. IV.	. 63 22 31	95 26 36
D's true long.—Long. Ronages.— D's dist. from Nonages	20 26 32	350 2 57
This distance or supp. if greater than 180° is such D.	D 202632	D 9 57 3
Altitude of Nonagesimal Prob. IV.	67 58 10	70 57 46
D's horizontal parallax, by Prob. L.	60 24.1	80 27.
O's hor, par. 8".8-Correction, Table XXXVIII. 5".6	- 14,4	14
Reduced parallex	60 9.7	60 12
D's S. Diam. by N. A.—Infexion 2'	16 26.7	16 26,
Add correction Table XLIV.	15.2	18.
D's corrected semi-diameter	16 40.9	16 42
o's semi-diameter by N. A. 15' 46".1-Irradiation 3".5	15 42.6	15 42
tun of the corrected semi-diameters	ST= 32 23.5	SL= 32 25
D's horary motion in longitude by Problem II. Ex. II.	36 39.2	. 86 42
o's horary motion	2 23.1	2 28
h's horary motion from the sun*	1 34 16.1	34 10.
D's parallex in longitude P	19 46.8	10 0
)'s apparent longitude—Error Tab. 58', 5, by Prob. V.	84 8 50.3	
has lengitude by Problem L.	84 41 8.4	84 47 35
Diff. D's app. lengitudes D app. mot.		1 10 42
Diff. Ors longitudes= @ app. mot.	1	6 52
Difference of motions of (1)	1	FA 64 10
b's true lat. by N. A. Prob. I.—Error Tab. 11".4	- 24 27.4	- 15 10
h's app. lat. cor. for error Tab. 11".4 by Prob. V.	FD=- 1 55.8	
b's lat, at end—Lat, at beginning	FD=- 1 00.8	AL=+ 6 28
D. s. tar. or cuto-row or orking	.1	AL

As the apparent latitude at the beginning of the eclipse is north, and at the end south, the point F corsesponding to this example falls above DE, the point L below it. The rest of the calculation is as follows.

FA 64' 10".2=3850".2 log. AL 6 28.3 = 388.3 log.		©'s longitude 24° 41′ 3″.4 - 32 23 .5	•.
Inclination 84° 14′ tan.	10.99631 co-s. 10.00220	B's app. long. 84 8 39 .9 by obs. D's par. long. — 19 46 .8	
Apparent motion FL Its arith. co	3869. 7 log. 3.58768 omp. 6.41232	D's true long. 83 48 58 .1 ©'s long. 84 41 8 A const.	3.55630
SF + SL = 64' 48".9 Diff. SF, SL	3888 9 log: 3.58988 1 9 log: 0.27875		8.49552 6.6867 <i>5</i>
Diff. segments	1 91 log. 0.28090	Time fr. conj. 1h. 31'18".1=5478".i log.	8.73168
Its half Half of FL Sum is great segment	0 95 1934 85 1935 8	App. ti. obs. 15 22 6 18.1 An. tl. conj. 15 23 37 36.2 at Salem	
Diff. is lesser seg, FB SF 32' 23".5=	1988 9 log. 13.28644 1943 5 log. 3.28858	Conjunct. 16 4 19 at Greenwich Diff. Merid. 4 41 25.8	
Angle FSB Inclination	84° 19′ sine 9.99786 84 14		
Diff. cent. ang. DSF SF	0 5 co-sin. 10.00000 lug. 3.28858	sine 7.16270	
6D=32 23"5 =	1943 5 log. 8.28858	App. lat. FD=2'.8 log. 0.45128	

In finding the time of conjunction or new moon, at Greenwich, 4h. 19', in the Nautical Almanac, the longitude of the moon was supposed to be given correctly by the tables. If the calculation be made by Problem III. after allowing for the error—58".5, the result will be 4h. 20' 47", whence the difference of Meridians—4h. 43' 10'.8, which differs so little from the assumed longitude 4h. 43' 32", that it will not be necessary to repeat the operation. If the eclipse was observed at Greenwich, the time of conjunction ought to be determined thereby, in a similar manner to the above calculation; or by those of Problem VIII. if only one of the phases is observed: by this means the errors of the tables will be wholly avoided. If the eclipse was not observed at Green-

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^a This borary motion iscreases from 34' 16".1 to 34' 19".7 or 3".8, during the eclipse 2h. 44' 18".5, which is 1".32 per hour. Now the ecliptic conjunction, or time of new moon at Greenwich by the N. A. was 4h. 19' or rather 4h. 20' 47', corresponding to 23h. 37' 15' at 8alem, which is 1h. 30' 87' after the heginaing of the eclipse: and the increase of the horary motion in half that time is 1", which added to 34' 16".1, gives the horary motion of 34' 17".1, corresponding to the middle time between the beginning of the eclipse and conjunction. This is used in calculating the correct time of conjunction.

wich, the observations at any other place whose longitude is known might be made use of, and thus the difference of meridians accurately obtained.

The moon's true longitude deduced from the above observation, is 83° 48′ 53".1; by the N. A. it is 830 50 2'.0, the difference-65".9 would be the error of the tables by this observation, if the assumed longitude 4h. 43' 32' and the solar tables were correct. By repeating the operation with the assumed longitude 4h. 43' 10".8 the error 68'.9 would be reduced to nearly the estimated value 55".5.

The eclipse was so nearly central at Salem, that a variation of a minute in the moon's latitude, would hardly alter the times, or duration of the eclipse, so that the latitude could not be determined by the above observations to any considerable degree of accuracy. From this cause it happens that the app. latitude at the beginning of the eclipse is by the above calculation 2".8 instead of 1'55".8, as found by allowing the error 11'.4, deduced from other observations made where the eclipse was not so nearly contral, and by the limits of the shadow of total darkness.

PROBLEM VII.

Given the latitude of the place, and the apparent times of the beginning and end of an occultation of a fixed star by the moon, to find the longitude of the place of observation.

In the following rule reference will be made to fig. 13, Plate XII. in which DSE represents a parallel to the ecliptic passing through the place of the star S; SF, SL the corrected semi-diameters of the moon at the beginning and end of the occultation; DF, EL the differences between the apparent latitudes of the moon and the star, when of the same name, or their sums when of different names; either of these lines falling below DE if the moon's apparent latitude is more southerly than that of the star, otherwise · above.

RULE.

To the apparent times of the beginning and end of the occultation add the estimated longitude of the place in time if it is west, but subtract if east; the sum or difference will be the supposed time at Greenwich, corresponding to which, in the Nautical Almanac, and by Problem I. the moon's semi-diameter, horizontal parallax, longitude and latitude, " and the sun's right ascension; also the moon's horary motion by Problem II. and the true longitude and latitude of the fixed star, by Table XXXVII. corrected for aberration and equation of equinoxes by Tables XL. XLI. Find also in the N. A. the obliquity of the ecliptic. To the moon's semi-diameter, add the correction in Table XLIV. and from the sum subtract the inflection 2', the remainder will be her corrected semi-diameter. With these elements and the apparent times at the place of observation, calculate the altitudes and longitudes of the Nonagesimal, by Problem IV. and the parallaxes in longitude and latitude, and the moon's apparent longitudes and latitudes by Problem V.

Take the difference between the apparent longitudes of the moon at the beginning and end of the occultation, which will be the moon's apparent motion in longitude, the logarithm of which, in seconds, being added to the log. co-sine of the meant of the apparent latitudes of the moon at the beginning and end of the occultation, rejecting 10 in the index, will be the logarithm of the motion of the moon on the parallel FA. The relative motion in latitude AL, is found by taking the difference of the moon's apparent latitudes at the beginning and end of the eclipse if they are both north or both south; but their sum if one be north and the other south. From the logarithm of FA, increasing the index by 10, subtract the logarithm of AL, the remainder will be the log. tangent of the angle of inclination DSB; this angle is to be taken greater than 90° when the difference of the moon's and star's apparent latitudes at the beginning of the occultation FD is greater than at the end EL, otherwise less. Then to the log: co-secant of the angle of inclination, add the logarithm of the relative motion FA, the sum, rejecting 10

in the index, will be the logarithm of the apparent motion of the moon in her orbit FL.

Then in the triangle SFL, the sides SF, FL (representing the corrected semi-diameters of the moon at the immersion and emersion,) and the relative motion FL are given to find the angle FSB (by Case VI. Oblique Trig.) Thus: to the log. ar. ce. of FL, add the log. of the sum of SF and SL, and the log. of their difference, the

Corrected for the errors of the tables in longitude and latitude, when known.
 This correction must be found after the altitude and longitude of the nonagesimal are calculated.

[†] This correction must be found after the altitude and longitude of the nongestimal arc calculated.

† The mean latitude is half the sum of the two latitudes, if they are of the same mame, but their half difference, if of different names. In solar eclipses, the correction for the mean latitude of the moon is neglected as too small to be taken notice of, the distance FA being taken equal to the difference of longitude DE (fig. 12, P. XII.)

§ This rule is equally true whether the points F, L fall on the same or on different sides of the line DE. If DF, EL are equal, and the points F, L fall on the same side of DE, the angle DSB will be 30. If they are equal, a d those points fall on different sides of the line DE, the angle DSB may be taken ague or obtained. In strictness when the points F, L fall on different sides of DE, the angle DSB is ET.

sum, rejecting 10 in the index, will be the log. of the difference of the segments FB, BL; half of which added to, and subtracted from half of FL, will give the two segments FB, BL, the greater segment being contiguous to the greater side, whether SF or SL. Then from the logarithm of the segment FB, increasing its index by 10, subtract the logarithm of SF, the remainder will be the log. sine of the angle FSB, which is always less than 90°. The difference between this and the angle of inclination DSB, will be

the central angle DSF.

To the log. co-sine of the central angle, add the logarithm of the moon's corrected semi-diameter at the immersion SF, and the log. secant of the star's latitude, the sum, rejecting 20 in the index, will be the logarithm of the apparent difference of longitude of the moon and star at that time. This is to be subtracted from the true longitude of the star, if the central angle is less than 90°, but added, if greater than 90°, the sum or difference will be the moon's apparent longitude; to this must be added the moon's paralax in longitude, when her distance from the Nonagesimal (found as in Problem V. by subtracting the longitude of the Nonagesimal from the moon's longitude, borrowing 360° when necessary) is greater than 180°, otherwise the parallax must be subtracted; the sum or difference will be the moon's true longitude at the beginning of the occultation.

Take the difference in seconds between the true longitudes of the star and moon at the beginning of the occultation, to the logarithm of which, add the arithmetical comp. log. of the moon's horary motion in seconds, and the constant logarithm 3.55630, the sum, rejecting 10 in the index, will be the logarithm of the time from the conjunction in seconds, which is to be added to the observed apparent time of the beginning of the occultation, when the star's longitude is greater than the moon's true longitude at that time, otherwise subtracted; the sum, or difference, will be the apparent time of the true ecliptic conjunction of the star and moon at the place of observation. The difference between this and the time of conjunction, inferred from the Nautical Almanac by Problem III. for the meridian of Greenwich, will be the longitude of the place. If corresponding observations be made at different places, it will be much more accurate to deduce from them the time of conjunction at each place, and take the difference of those times for the difference of meridians, if it does not differ much from the suppose difference of longitude. 'If there is considerable difference, the operation must be repeated, making use of the longitude found by this operation; and thus by successive operations the true longitude may be obtained.

The long, of the place of observation being accurately known, the errors of the lunar tables in lat. and long, may be easily found. For the difference between the moon's true longitude, deduced from the observations by the above method, and the longitude found from the Nautical Almanac, will be the error of the tables in longitude. the error in latitude proceed thus: To the log. sine of the central angle DSF, add the logarithm of the corrected semi-diameter of the moon at the immersion SF, the sum, rejecting 10 in the index, will be the logarithm of the apparent difference of latitude of the moon and star, which added to the true latitude of the star, with the sign + if the point F falls below the line DE, but with the sign — if above, will give the apparent latitude of the moon at that time, the difference between this and the apparent latitude found by Problem V. will be the error of the tables, always supposing the sign + to be

prefixed to southern latitudes, the sign — to northern, and noting the signs as in algebra.*

REMARK. In the two preceding Problems the time of the true conjunction is calculated by means of the triangle SFD, but it will be useful for the purpose of verification, to go over the The process is nearly the same in both calculation by means of the triangle SLE. methods. The differences consist in finding the angle LSB by subtracting the logarithm of SL from the logarithm of LB, increasing its index by 10, the remainder will be the log. sine of the acute angle LSB, which, added to the angle of inclination, (found as before) will give the central angle DSL, with which and the distance SL, corresponding to the end of the eclipse or occultation, may be found the apparent diff. of longitude between the sun and moon, and moon and star; this is to be added to the longitude of the sun or star at that time, if the central angle exceed 90°, otherwise subtracted, the sum or difference will be the apparent longitude of the moon corresponding, from which the time of the ecliptic conjunction may be obtained as before. If the central angle exceed 180° the sine and co-sine of the excess of that angle above 180° must be found instead

of the sine and co-sine of the central angle.

The apparent latitude of the moon is found as in the preceding rules, by making use of the central angle. The apparent latitude and the moon is found as in the preceding rules, by making use of the central angle DSL and the value SL, corresponding to the end of the eclipse or occutation; whence may be deduced the apparent latitude and the error of the tables in latitude.

It is evident that both these methods ought to give the same results and thus furnish a proof of the correctness of the calculations. All these calculations may be made by proportional logarithms, by creating in the rule, log. co-tung. for log. tang. log. co-secant for log. sine, &c. as was mentioned at the end of the rule in Problem V. and by using the constant log. 0.4771 instead of 3 55530.

^{*} I When SF=SL the angle may be found as in the note, with this mark in page 583.

† When this varies, it must be taken to correspond to the middle time between the immersion and true conjunction.

* See note; with this mark in page 584 guized by

TAE.

EXAMPLE.

EXAMPLE.

Suppose in a place in the latitude of 20° 0′ N. longitude 1b. 9m. 0s. east of Greenwich by estimation, the occultation of Spica by the Moon on Dec. 12, 1808, was observed; the immersion at 16b. 57' 23', expersion at 18b. 10' 23', apparent time by astronomical computation. Required the longitude of the place of observation?

Most of the elements in the following Table are calculated by Problems I. II. and VI.

ELEMENTS OF THE OCCULTATION.	IMMERSION.	
	d. h. ' "	d. b. "
Apparent times of observation	12 16 57 2 9	12 18 10 29
Estimated longitude E. from Greenwich	1 .9 0	1 9 0
Supposed time at Greenwich	12 15 48 29	12 17 1 29
We right accension	17 20 59.0	17 21 12,3
Lat. of place 20° 0'-Reduc. Tab. XXXVIII. 7' 22"	19°52′ 38′′	. •
Obliquity of the ecliptic	23 27 39	9 ' "
h's long, by N. A.—Prob. I.	200 7 56.3	
Congitude of the Nonagesium, by Prob. IV.	149 15 55	
D's long.—Long. Nonages— D's dist from Nonages	50 512 1	35 22 38
his distance or its supp. to \$60° is arch D	D 50 52 1	D 35 22 38
Altitude of Nonagesimal Prob. 1V.	81 17 32	
N'e havisantel merallay	59 52.S	59 54.4
Reduction, Table XXXVIII.	1.4	1.4
Reduced parallax	59 50.9	
D's S. Diam. by N. A.—Inflexion 2"	16 16,9	16 17.5
Add correction Table XLIV.	10.4	13.3
D's corrected semi-diameter	SF 16 27.3	SL 16 30.8
I's horary motion in longitude by Problem II. Ex. I*.	35 51.7	35 54.2
)'s parailax in longitude	46 25	33 54
's apparent longitude	200 54 21.3	201 25 30.1
Difference of D's app. longitudes		31 KJ
o's true lat. by N. A. Prob. I. South	1 55 11.0	1 51 29.1
's parallax in latitude	10 23.6	
)'s apparent latitude South	2 5 34.6	
k's tr. lat = lat. T. XXXVII. 2º 2' 13".9. ST. XLI. 0".6		
Difference D * app. lat.	FD= 3 21.3	
Difference of D's app. latitudes		AL 3 32 9
*'s tr. long. Long. Tab. XXXVII. 2010 10' 29'.3-Tab.		
XL. 11".5-Tab, XLI. 10".1.	201 10 36.7	}
	a at the beginning	

The difference of the apparent latitudes of the Moon and Star at the beginning of the occaitation 3'21" 3 being less than at the end 6' 54".2 the angle of inclination is less than 30°. In this example the moon's latitude is more southerly than the star's, bence the points F, L, fall below the line DE. Diff. app. long. D 31' 8' 8=1688".8 loc. 3.7156 1

Diff. of meridians 1 9 2		Digitized by GOOGLE.
Conjunction 17 53 0	place of observa- tion at Greenwich.	
Immersion 16 57 29	,	
Time 6278 1 44 33	log. 3.79748	Error Tab. + 1.9 in long.
D's true long. 200 7 58.2 Diff. true long. 3752.5 1 2 32.5 D's hor. mot. 2153.5 85 53.5	const. S.55630 log. S.57432 log.co-ar. 6.66686	Error Tab. — 1 .4 in lat. D's tr. lon. 200 7 56 .2 by obs. D's tr. lon. 200 7 56 .3 by N. A.
)'s par. long. — 46 25		D's ap. lat. 2 5 34 .6 by N. A.
) app. long. 200 54 23.2	by obs.	D's sp. lat. 2, 5 33 .2 by obs.
Diff. app. long. D * 967".5 = 16 7.5 * longitude 201.10 30.7	log. 2.98563	* lat. 2 2 13 .3
Star's lat. SF 2° 2′ 13′′	log. 2.99445 sec. 10.00027	FD 190".9= 3' 19".9 lon. 2.50053
Diff. is cent. angle 11 41	co-sine 9.99091	. sine 9.30ats
Inclination 83 30		·
FSB 71° 49'	sine 9.97775	İ
FB 938.0 SF 987.3	log. 2.97220 log. 2.99445	
Its half 1 .8 Half FL 939 .8		
Diff. segments 3 .7	log. 0.56628	
Diff. SF, SL 3.5	log. 0.54407	
Apparent motion FL 1879 .6 Its arith. cemp. BF + 8L = 52 58.1 = 1978 .1	6.72594 log. 3,29625	3.27406
Inclination 83° 30'	trn. 10.94308	co-sectint 10.00240
Distance FA D's diff. lat. AL=3 52.9=212 .9	log. 13.27126 log. 2.82818	5.27126
D's mean app. lat. 2 7 21	log. 3.27156 cos. 9.99970	

The moon's horary motion varies from 35' 51".7 to 35' 54".2 during the occultation, hence at the middle time 17h. 49' 45" between the immersion 16h. 57' 29" and the conjunction 18h. 42' (deduced from the Nautical Almanac) the horary motion was 35' 55".5 as is easily found by a calculation similar to that in the Example of Problem VI.

The difference of meridians deduced from the observation 1h. 9'2" differs but 2" from the assumed quantity 1h. 9'0". If the difference had been considerable, it would have been necessary to repeat the operation with the difference of meridians thus calculated, and so on till the assumed and calculated longitudes agree. of the tables above found were deduced upon the supposition that the observations were actually made at the place mentioned in this example, and that the true longitude of the place of observation was 1h. 90". For it must be observed, that the errors of the tables in longitude cannot be found by an observation of an eclipse or occultation without knowing by other observations the precise longitude of the place of observation. This is evident by observing, that by repeating the operation till the assumed and calculated longitude of the place of observation agree with each other, the long. of the moon, deduced from the calculation, will agree also with the longitude by the tables. The time of conjunction at Greenwich 17h. 33' 0" taken from the Nautical Almanac, is liable to a small error from the incorrectness of the tables. To obviate this error it will be necessary to deduce (by the above method or by Problem IX. when only the beginning or end is observed) the time of conjunction from observations actually made at two places, the difference of these times will be the difference of meridians free from the errors of the tables.

PROBLEM VIII.

To find the longitude of a place by an eclipse of the sun when the beginning or end only is observed, the apparent time being estimated from noon to noon, according to the method of astronomers; the latitude of the place being also known.

RULE.

RULE.

To the apparent time apply the estimated longitude of the place in time, by adding if west, substract ling if esst, the sum or difference will be the supposed time at Greenwich. Corresponding to this time in the Nautical Almanac, find by Problem II. the moon's semi-diameter, horizontal parallax, longitude, and latitude; and the sun's semi-diameter, longitude, and right ascension; also the moon's hornry motion from the sun by Problem II. Decrease the sun's semi-diameter 34" for inflation. Decrease the moon's semi-diameter? To inflation. Decrease the moon's semi-diameter? To inflation. Decrease the sun's semi-diameter 34" for inflation. Decrease the sun will be the moon's corrected semi-diameter. Find, also, in the Nautical Almanac, the obliquity of the cellptic.

With these elements and the apparent time at the place of observation, calculate the altitude and longitude of the nongesimal by Problem IV. and the parallaxes in longitude and latitude and the moon's apparent latitude, and find the logarithms of the sun and moon add and subtract the moon's apparent latitude, and find the logarithms of the sun and moon add and subtract the moon's apparent latitude, and find the logarithms of the sun and moon add and subtract the moon's longitude, and find the logarithm of an arch in seconds, to be added to the sun's longitude if the phase is after the apparent conjunction, but subtracted if before; § the sum or difference will be the apparent longitude of the moon. To this add the moon's parallax in longitude when the moon's distance from the monagesimal (found as in Froblem VI. by subfacting the longitude of the nonagesimal from the moon's longitude, borrowing 350° when necessary) is greater than 180°, otherwise subtracted, the sum or difference will be the orne longitude of the moon's horary motion from the sun in seconds, and the constant logarithm 3.55630, the sum, rejecting 10 in the index, will be the longarithm of the correction of the given time expressed in seconds. This is to be added to the apparent ti

EXAMPLE.

At Salem, in the intitude of 42° 33′ 30′ N. longitude by estimation 4h. 43′ 32′ W. from Greenwich, the legisming of the total eclipse of June, 1806, was observed at 15d. 22h. 6 18′.1 apparent time by astronomical computation. Required the longitude of the place from this observation? The elements must be calculated as in the example of Problem VI. for the beginning of the eclipse, except those marked in Italics. The rest of the calculation may be made by proportional logarithms

as follows :

operation, or by other observations.

† This correction must be found after the altitude and longitude of the nonagesimal are calculated.

† These calculations may be made in the same manner by using proportional legarithms, the only difference consists in using the constant logarithm 0.4:71 instead of 3.56:30 in finding the time of conjunction.

^{*} The longitude and latitude must be corrected for the errors of the tables, when known, by a previous

conjunction.

§ In general, the beginning of an eclipse or occultation precedes the apparent conjunction, and the end is after the apparent conjunction, but there is a case (which very rarely occurs) where the convary may take place: namely, where the point F or L, (P. XII. fig. 12, 13, [alis between C and B, which can happen only when the lines FI, EL are nearly equal to SF or SL. In this case it may be ascertained whether the phase precedes or follows the conjunction by making the calculation as in Prob. The or VIL with the times of beginning and end, calculated by Problem XIII. and as the central angle is greater or less than 90°, the phase will, follow or precede the apparent conjunction. The latitudes given for the fathly being numposed correct. by the tables being supposed correct Digitized by GOOGL

Kum semi-diam. (9) D App. lat.	D	32' 23''.5 1 55 .8		
Sum Diff.	,	54 19 .S 50 27 .7	P. L. P. L.	0.7197 0.7715
			Sum	1.4912
Half sum ⊙ Longitude		52 20 4 41 3 .4	P. L	7456
D App. long. D Par. long.		8 45 .4 - 19 46 .8	•	
D True long. True long.		3 48 56 .6 4 41 3 .4	Cons. log.	0.1721
Difference D Hor. mot. fr.		52 6 .8 31 17 .1	P. L. A Co. P. L.	0.5383 9,2798
Time from covj. App. time obs.		1. 31' 15'' 1. 6 18	P. L	0.2952
App. conj. Salem App. conj. Green.		37 SI 19 by N. A.		
Diff. Merid.	4	1 41 29		

It we suppose the time of conjunction at Greenwich to be 4h, 20' 47', as calculated in the example Problem VI, the difference of meridians would be 4h, 43' 16", agreeing nearly with the assumed longitude, so that it will not be necessary to repeat the operation. The remarks at the end of that example, respecting the errors of the lunar tables, and the comparing of actual observations at different places, are equally applicable to the present Problem.

PROBLEM IX.

To find the longitude of a place by an occultation of a fixed star by the moon, when the immersion or emersion only is observed, the apparent time being estimated from noon to noon, according to the method of astronomers, and the latitude of the place being known.

RITLE.

To the apparent time apply the estimated longitude of the place turned into time, by adding if west, subtracting if east, the sum or difference will be the supposed time at Greenwich. At this time find in the Nautical Almanac the sun's right ascension, the moon's semi-diameter, horizontal parallax, longitude and latitude * by Problem I. and the moon's horary motion by Problem II. also the latitude and longitude of the fixed star by Table XXXVII. and correct it for aberration and equation of equinoxes by Tables XL. XLI. Decrease the moon's semi-diameter 2" for inflexion, and to the remainder add the augmentation from Table XLIV. † the sem will be the corrected semi-diameter. Find also, in the Nautical Almanac, the obliquity of the ecliptic. With these elements and the apparent time of observation, calculate the altitude and longitude of the Nonagesimal by Problem IV. also the parallaxes in longitude and latitude of the moon's apparent latitude by Problem V.

Take the difference between the latitude of the star and the app. lat. of the moon. which add to, and subtract from the moon's corrected semi-diameter (these quantities being expressed in seconds) half the sum of the logarithms of these quantities increased by the log. secant of the star's latitude, rejecting 10 in the index, will be the logarithm; of an arch in seconds to be added to the star's longitude if the moon has passed the apparent conjunction, but subtracted if before, § the sum or difference will be the apparent longitude of the moon. To this add the moon's parallax in longitude when the moon's distance from the nonagesimal (found as in Problem VII. by subtracting the longitude of the nonagesimal from the moon's longitude, borrowing 360° when necessary) is greater than 180°, otherwise subtract it, the sum or difference will be the true longitude of the moon. Take the difference in seconds between the moon and star's true longitudes, and to its logarithm add the arithmetical comp. log. of the moon's horary motion and the constant logarithm 3.55630, the sum, rejecting 10 in the index, will be the logarithm t of a correction in seconds to be applied to the given time of observation by adding when the moon's true longitude is less than the star's, otherwise subtracting, the sum or difference will be the time of the true conjunction at the place of observation. The difference between this and the time of conjunction inferred from the Nautical Almanac by Problem III. for the meridian of Greenwich, will be the longitude of the place of observation, if the tables are correct; but it is much more accurate to compare the times of conjunction deduced

Corrected for the errors of the tables in longitude or latitude when known.

This correction must be found after the altitude and longitude of the nonagesimal are calculated.

See note with this mark in page 589.

Proportional logarithms may be used instead of common logarithms, the constant logarithms being 4271 instead of 3,55630, and the log. co-side being used instead of log. scent.

from actual observations at the different places in the manner mentioned at the end of the rule given in Problem VII.

EXAMPLE. Suppose in a place in the latitude of 20° 0' N. longitude by estimation 1h. 9' 0" east from Greenwich, the emersion of the star Spica was observed on December 12, 1608, at 18h. 10' 29", apparent time by astronomical computation. Required the longitude of the place of observation?

The elements must be calculated as in the example of Problem VII. for the emersion of Spica. The rest of the calculation, made by common logarithms, is as follows.

) Semi-diameter 16' 30''.8--990''.8 Biff. app. lat.) × 6 54 .2 414 .2 1405 .0 log. 3.14768 log. 2.76087 Difference 576 .6 A.90855 Sum * lat. 2° 2' 13". 10.00027 Arch 15' 0''.6 2,35154 2001.6 log. * longitude 201 10 30 .7 app. long. 201 25 31 .3 33 64 ك 37 37 200) true long.
Diff. true long. D Constant 18 53 .4 = 1153.4 35 54 .7 = 2154.7 log. 5.05438 6.66661 = 1894 Time of obs. 18 10 29 Conj. at place obs. Conj. at Greenwich 18 42

The difference of meridians by calculation 1h. 9' 3" differs but 3" from the assumed longitude, so that it will not be necessary to repeat the operation. All the remarks made at the end of the example in Problem VII. are applicable to this problem. also be further observed, that the emersion or immersion which happens on the dark limb of the moon can be observed with much more accuracy than on the enlightened limb; because the light from this limb prevents the observer from perceiving the star's immersion or emersion so instantaneously as on the dark side of the moon. PROBLEM X.

by obs.

by N. A.

17 33

To calculate an eclipse of the moon.

The time of beginning or end of a lunar eclipse at any place may be found by subtracting or adding the longitude to the times given in the Nautical Almanac for the meridian of Greenwich, according as the longitude is west or east. But as some readers may wish to know the method of deducing these times from the longitudes, latitudes, &c. of the moon and sun, given by the Nautical Almanac or by other tables, it was thought proper to insert the rule for these calculations.

An eclipse of the moon can only happen at the time of the full moon. If her longitude at that time is not distant from either node* of the moon's orbit more than about 12°, there may be an eclipse. To find whether there will be one, and to calculate the

times and phases, proceed as follows.

Difference merid.

RULE.

Find the time of full moon at Greenwich by the Nautical Almanac or Problem III. to which add the longitude of the place turned into time if east, but subtract if west, the sum or difference will be the time of the ecliptic opposition at the proposed place.

For the time at Greenwich find by Problem I. the moon's latitude, horizontal parallax and semi-diameter (which requires no augmentation) also the sun's semi-diameter. Then by Problem II. the horary motion of the moon from the sun in longitude, and the moon's horary motion in latitude.

Draw the line ACB (Plate XII. fig. 6) and perpendicularly thereto the line PCR. Select a scale of equal parts to measure the lines of projection, and from it take CG equal to the moon's latitude, and set it on CR from C to G, above the line AB if the latitude of the moon is north, below if south. Take CO equal to the horary motion of the moon from the sun in longitude, and set it on the line CB to the right of

^{*} The longitude of the moon's ascending node is given in the third page of the Nautical Almanac.
The longitude of the other node is found by adding or subtracting 6 signs.

† The northern latitudes found by Prob. I. have the sign —, the southern —. In the figure the latitude south.

If it had been north the point G must have been placed on the continuation of RC above C.

C. from C to O. Take CP equal to the moon's horary motion in latitude, as found with its sign by Problem II, and set it on the line CR from C to P, above the line AB if its sign is -, below* if +. Join OP which is equal to the horary motion of the moon from the sun, and parallel thereto through G draw the relative orbit of the moon from the sun NGL, on which are to be marked the places of the moon before and after the full, by means of the horary motion OP, so that the moment of full moon, or ecliptic opposition at the proposed place, may fall exactly on the point G. This may be done by making the extent OP equal to the transverse distance of 60, 60, on the line of lines of the sector, then measuring from the same lines the transverse distance corresponding to the minutes and parts of a minute in the time of full moon at the place of observation, and setting it on the line GN from G towards the right to the point x, where the whole hour preceding the full moon is to be marked. Then the distance OP set from x to the right hand on the line LGN reaches to the hours preceding the full moon, and set to the left hand reaches successively to the following hours. These intervals are to be divided into sixty equal parts representing minutes, if the size of the scale will admit of it.

Add 50" to the moon's horizontal parallax; and from the sum subtract the sun's semidiameter, the remainder will be the semi-diameter of the shadow CB, with which describe the circle ASB about the centre C. Add the moon's semi-diameter to the radius CB, and with that radius describe, about the centre C, the circle DRM, which if there be an eclipse will cut NL in the points E, H, representing respectively the places of the moon at the beginning and end of it. If there is no intersection there will be no eclipse. Draw the line CKST perpendicular to LN, cutting it in K and meeting the circles ASB, DRH in S, and T. With a radius equal to the moon's semi-diameter describe about the centres E, H, K, the small circles represented in the figure; of which that drawn round K cuts the line CKS in the points I, F, and if the eclipse is total the whole of this circle will fall within ASB, as in fig. 6, but if part of the circle falls without ASB, as in fig. 7, P. XII. the eclipse will be partial. In either case the number of digits eclipsed may be obtained by saying as the diameter of the moon FI is to the obscured part FS so are 12 digits to the number of digits eclipsed. When the eclipse is total, the beginning and end of total darkness may be found by taking a radius equal to CB decreased by the moon's semi-diameter, and sweeping with it round the centre C, a circle d e h m, cutting J.N in the points e, h, representing respectively the points of beginning and end of total Then the hours and minutes marked in the line NL, at the points E, e, K, darkness. h, H, will represent respectively the times of the beginning of the eclipse, beginning of total darkness, middle of the eclipse, end of total darkness, and end of the eclipse. this rule no allowance is made for the oblate figure of the earth, the correction from this source being much less than the errors of observation.

EXAMPLE.

Required the times of beginning, end, &c. of the eclipse of the moon of May 9th; 1808, at a place in the longitude of 30° W. from Greenwich?

By the Nautical Almanac the time of full moon at Greenwich was May 9th, at 19h. 39'. From this subtracting the longitude of the place of observation 30° W. or 2h. the remainder 17h. 39 was the time of full moon at the place of observation. Corresponding to the time at Greenwich, 19h. 39', the elements in the adjoined table were calculated by Prob. I. and II. and the values CB, CD, Cd, found by the above rule. Upon the centre C with the radii CB, CD, Cd, taken from a scale of equal parts, describe the circles ASB, MRD, mrd.

me of full	Elements of the Ectipse May	9, 1	9h. 3	19"	٠.
h, at 19h.	App. time conj. Greenwich, May 9		196.	39	1
longitude	Long. place 30° W.		2	0	ł
W on Oh	App. time conj. at place obs. D's lat. by Prob. I. S. decreasing		17	39	ł
v. or zu.	D's lat. by Prob. I. S. decreasing	CG	+	10 44".E	ı
me of full	D's Horiz. Paral.		·	61 13 -9	
Corres-	D's Semi-diameter	BD	1	16 49 .7	
10h 30'	O's Semi-diameter D's Hor. Mot. in long. Prob. IL.		1	15 51 .3	
, 1 , 14. 5 ,	D's Hor. Mot. in long. Prob. IL.			37 37 .8	
were cal-	O's Hor. Mot. in long.			2 24 .8	
he values	D's Hor Mot from @ in long.	CO	ł	3 5 13 4	
ile Unon	D's Hor. Mot. in lat. Prob. II. D's Hor. Par. +50' - O's S. D. =	CP	-	3 28 .2	
or open	D's Hor. Par +50 - 6's S. D. =	CB	1	46 12 J	
$\cup D, \cup a,$	ICB+ D's S. D.	CP		62 52 .5	
, describe	CB-D's S. D. =	Cq	!	29 31 .	7
Draw th	e line ACB representing th	e e	clip	tic, an	ď

^{*} In other words the point P will fall above C if the moon is approaching to the north pole of the ecliptic, otherwise below: That is, the point P must fall above C if the moon's latitude is south decreasing or north increasing, otherwise below. When no great accuracy is required the horary motion in initiation need not be found by Prob. II. Instead of which the angle COP may be taken east to 5° 40°, in eclipses of the moon or sun, and the line OP equal to CO increased by 9° or 10°: but this method will not answer in occultations, in which the angle COP varies above 5 degrees.

1. The detance Ga way also hardward by company activities to expine as 60 minuted are to the minutes.

in occumations, in which the angre COT varies above a degrees.

† The distance Gr way also begious by common arithmetic by saying as 60 minutes are to the minutes and seconds in the time of full moon (which in the present example is 39) so is OP to Gr. After marking the hours on the line LGN it is usual to divide them successively into halves and quarters of an hour, then into five minutes and one minute.

the semi-diameter of the shadow is increased by the earth's atmosphere from 20" to 60", according to the estimates of different astronomers. Mayer supposes this correction to be one 60th part of the shadow, varying from 37" to 46". The mean of Mayer's correction added to the sur's paraliar is nearly equal to 50" assumed as above.

make CG perpendicular thereto equal to the moon's latitude 10' 44":8 S. the point G being taken below C because that latitude is south. Make CO equal to the horary motion of the moon from the sun in longitude 35' 13".0, and CP perpendicular thereto equal to the horary motion in latitude-3' 28".2, the point P being placed above C, because the moon's horary motion in latitude has the sign - prefixed, or in other words, the latitude was south decreasing. Join OP, and parallel thereto draw through G the line NGL, and on it let fall the perpendicular CK. Make the distance OP a transverse distance of 60, 60, on the line of lines of the sector, and measure from the same lines the transverse distance 39,39 (corresponding to the minutes in the time of full moon at the blace of observation) this distance, set on the line GN, to the right of G, reaches to the point x, where the hour 17h, preceding the full moon is to be marked. Take the extent OP and lay it from 17h. to the right hand to 16h. and successively to the left to 18h. 19h. &c. Subdivide these lines into 60 equal parts, representing minutes, if the scale will permit, and the times corresponding to the points E, e, K, h, H, will represent respectively the beginning of the eclipse 15h. 56m. the beginning of total darkness 16h. 54m. the middle of the eclipse 17h. 41m. the end of total darkness 18h. 28m. and the end of the eclipse 19h. 26m. which times agree nearly with those in the Nautical Almanac, allowing for the difference of meridians 2 hours.

Calculation by Logarithms.

The phases of the eclipse may also be calculated by logarithms in a very simple manner. Thus suppose it was required to find the time of the beginning of the eclipse in the above example. In this case in the right-angled triangle OCP, there would be given CO=2113".0 and CP=208".2, to find OP=2123".2 and the angle OPC=84°22. This angle is equal to RGE, because GE, OP are parallel, and its supplement gives the angle CGE=95°38'. Then in the triangle CGE there are given the angle CGE=95°38' the moon's latitude CG=644".8, and the line CE (=CD) =3773".9 to find CEG=9°48', GCE=74°34' and GE=3654".5. Then say as OP (2123".2) is to 1 hour (3600".) so is GE (3654".5) to the time (6196"=) 1h. 43' 16" between the beginning of the eclipse and the full moon at the place of observation 17h. 39', and as the point E falls to the right hand of G, that time must be subtracted from 17h. 39', to obtain the time of the beginning of the eclipse 15h. 55' 44", which agrees nearly with the projection. As these calculations are very simple, it will be unnecessary to take notice of the different cases, or to give the galculations at full length, the whole being sufficiently evident from the figure. The middle of the eclipse is found by means of the triangle GKC similar to OCP, in which the angles and hypotenuse CG are given to find CK, KG. The time of describing KG being added to, or subtracted from the time of full moon at the place of observation, according as the point K falls to the left or right of G, will give the time of the middle of the eclipse. The distance CK 10' 41".7 subtracted from the radius CD or CT=63' 52".9 will leave a remainder equal to the eclipsed part F8 (=KT) 52' 11".2 as 12 digits to the left or proportional logarithms may be made use of.

PROBLEM XI.

' To project an eclipse of the sun for any given place.

An eclipse of the sun can happen only at the time of new moon. If the moon's longitude at that time is not distant from either node of the moon's* orbit more than 173° there may be an eclipse. To find whether there will be one, and to calculate the times and phases, proceed by the following.

RULE.

To the time of the new moon given in the Nautical Almanac (or calculated by Problem III.) add the longitude of the proposed place, turned into time, if east, but subtract if west, the sum or difference will be the time of conjunction at the proposed place. Corresponding to the time of new moon at Greenwich, find by Problem I. the moon's latitude, horizontal parallax, and semi-diameter, also the sun's longitude, semi-diameter and declination. Then by Problem II. find the horary motion of the moon in latitude, and the horary motion of the moon from the sun in longitude.

Draw the line ACB (Plate XII. fig. 10.) representing the ecliptic and perpendicularly thereto, the line PCR. Take a scale of equal parts to measure the lines of the projection; measure from it an interval equal to the moon's latitude, and apply it on CR from C to G, above the line ACB if the moon's latitude is north, below, if south.† Take CO equal to the horary motion of the moon from the sun in longitude and set it on the line .CB to the right hand of C to O; take CP equal to the moon's horary

† In the figure the latitude is supposed north. If it had been as much south, the point & weight have been as much below C as it is now above it.

^{*} See note with this mark in page 591. All the eclipses that can happen in any part of the earth are indicated in the Namical Almanac.

motion in latitude, found by Prob. II. and set it on the line CR, from C to P, above . the line ACB if the sign is -, below if +. Join OP which represents the horary motion of the moon from the sun on the relative orbit, and parallel to that line draw the relative orbit of the moon NGL, on which are to be marked the places of the moon before and after the conjunction, by means of the horary motion OP, so that the moment of the new moon, or ecliptic conjunction, at the proposed place may fall exactly on the point G, as in the figure where the new moon is at 23h. 35dm. This may be done by taking the extent OP equal to the transverse distance of 60, 60, on the line of lines of the sector, then measuring from the same lines, the transverse distance corresponding to the minutes and parts of a minute of the time of new moon at the place of observation and setting it on the line GN from G towards the right hand to the point x, † the place of the moon at the first whole hour preceding the conjunction (which in the present figure is 23h.) Then the distance OP being taken in the compasses and set from æ to the right hand, gives successively the hours preceding the new moon, and the same distance set to the left gives the following hours, as in the figure, where they are marked in succession 22h. 23h. 24h. 1h. These hours are to be divided into 60 equal parts representing minutes, the scale being taken sufficiently large for that purpose. I In the

present figure the subdivisions are carried only to five minutes.

From the moon's horizontal parallax subtract the sun's 8".8, the remainder is to be faken from the scale of equal parts for the radius CB, with which, on the centre C, describe the circle BRA, cutting CR in R. Open the sector till the transverse distance of 60°, 60°, on the line of chords, is equal to the radius CB, and measure from the same lines the transverse distance 23° 28' (equal to the obliquity of the ecliptic) which set on the circle ARB on each side of R to T and U. Join TU cutting CR in Q. On Q as a centre, with the radius QT, describe the circle TVU, on which set off the arch TV equal the sun's longitude. Through V draw the line VP parallel to CR to cut TU in P, the place of the pole of the earth. § Draw CP and continue it on either aide so as to cut the circle ARB in the point W, situated above AB if the latitude of the proposed place is north, below, if south. In the present figure the latitude is north. If it had been south the lower part of the circle ARB ought to have been made use of. Open the sector so as to make the transverse distance 60°, 60°, on the chords, equal to CB, and measure off the transverse distance equal to the chord of the com-plement of the latitude of the .place, which set from W on each side to D and & With the same opening of the sector measure the chord of the sun's declination, and With the same opening of the sector measure the chord of the sun's decunation, and set it on the same circle from D on each side to E and F, and from d on each side to e and f. Draw the dotted lines Ff, Dd, Ee, cutting CW in l, q, n. Bisset is in r, and erect the line VI. r, XVIII. perpendicular to CW, and make r, VI. and r, XVIII, each equal to qD. Open the sector to make the transverse distance 90° , 90° , on the sines, equal to qD, and measure off the transverse distance corresponding to 15° , 30° , 45° , 60° , 75° (or 1, 2, 3, 4, 5 hours,) which set on each side of the point r, on r, VI. and r, XVIII. to the points marked with those numbers 15° , 30° , &c. Through these points draw the lines I. XI, II. X; III. IX; &c. as in the figure parallel to CW. Quen the sector so as to make r, event to the transverse distance. figure parallel to CW. Open the sector so as to make m equal to the transverse distance of 90°, 90°, on the sines, and measure the complements of the former degrees

† See note with this mark in page 592.

† The scale I generally make use of is one inch to 10 minutes, reducing the seconds to decimals of a minute. Thus 30' 85" in decimals is 50'.8, which by this scale would be 5.06 inches, obtained by placing

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^{*} See note with this mark in page 592.

The scale I generally make use of is one inch to 10 minutes, reducing the seconds to decimals of a minute. Thus 30' 68'' in decimals is 50'.8, which by this scale would be 5.06 inches, obtained by placing the decimal point one figure to the left.

§ This may also be found as follows. After drawing TQU, as above, open the sector till the issueverse distance 90°, 90° on the sines is equal to QT, then measure from that line the extent QP' as a transverse distance corresponding to the sine of the difference between the sun's longitude and 90° or 270°. When the sun's longitude exceeds 6 agms, the point V will fall in the semi-circle below TU. This is not drawn in the figure for want of room. When the longitude exceeds 2, 4,6, &c. sigm, it will be convenient to mark on the circle TYU the points corresponding to those sigms, by setting off the radius QT as a chord from T to II, from II to \$\int_{\infty}\$, &c. and then taking from the sector the chord corresponding to the exceeded the given longitude above that of the point II, \$\int_{\infty}\$, &c. immediately preceding. Thus if the sun's longitude was 84° 44' it would be convenient to set of 80° from T to II, and 24° 44' from II, to the sought point V. In case of not having a sector, an arch as RT may be marked off by a plane scle even when the such its CR differs from that of the scale, by drawing by Prob. VI. of Geometrical Problems, the line CT making an angie with CR equal to the proposed arch 23° 28'. The intersection of that line with the circle ARB will give the sought point T. In a similar manner the point 10° 10°, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$ can be found by drawing a line QV making the angle TQV equal to the proposed arch TV. The points 15°, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\int_{\infty}\$, \$\infty\$. Each intersection of that line with the circle ARB will give the sought point T. In a similar manner the point 15°, \$\infty\$, \$\int_{\infty}\$, \$\infty\$, \$\infty\$. Acc. on t

as transverse distances on the sines, viz. 75°, 60°, 45°, 30°, 15°, and set them on the above lines I. XI; II. X; &c. from the points of intersection with the line VI. 7, XVIII. above and below that line. The points I. II. III. &c. obtained in this manner, will represent the situation of the spectator at the proposed place at those hours, and a regular curve drawn through these points will represent his path. In marking the hours it must be observed, that the place of noon will be at the lower point n, if the sun's declination is north; but at the upper point l, if the declination is south; the hours must be marked from noon towards the left in numerical succession completely round the curve ending at 24h. according to the method of astronomers. In the present figure the declination is north, and the point n the place of noon or 0 hours. If it had been south the point t would have been marked 0h. and the points marked XI. X. &c. would be I. II. &c. re-The path touches the circle ARB in two points representing the points of sun rising and setting, which in the present figure are respectively 16h. 26 and 7h. 34.

These points divide the path into two parts, of which one represents the path by day, the other by night, as is evident from the hours marked on the curve. Half hours or any other intermediate time, may be marked in a similar manner. Thus, for the time 3h. 30'=52° 30'. Set the sine of 52° to the radius r, VI. from r to h en the line r, VI. and erect the perpendicular hi equal to the sine of 37% (which is the complement of 5201) to the radius m, and the point i will be the place of the spectator at the proposed time. In this way the halves and quarters of hours may be marked on those parts of the path where necessary. The smaller subdivisions may generally be obtained to a sufficient degree of accuracy by dividing the quarters of hours into equal parts.

Take from the scale of equal parts an extent equal to the sum of the semi-diameters of the sun and moon, and beginning near N, find by trials the point p' of the moon's path and the point Z' of the path of the spectator, marked with the same time and at that distance apart. That time will be the beginning of the eclipse. If no such points can be found, there will be no eclipse at the proposed place. Proceed in the same way towards the point L and find the points p'', Z'', at the same distance apart, the corresponding time will be the end of the eclipse. Find by trials the point p of the moon's path and the point Z of the path of the spectator marked with the same times at the nearest distance from each other (which will in general be nearly the middle time between the beginning and end of the eclipse) that time will be the middle of the eclipse. On Z as centre with a radius equal to the sun's semi-diameter, describe the circle whose diameter is Se, representing the sun's disc, and on the centre p, with a radius equal to the moon's semi-diameter, describe the circle whose diameter is Mm, representing the moon's disc. The part of the sun's disc that is cut off by this circle will represent the part of the sun that is eclipsed. In the example of fig. 10 the centre p of the moon's disc is so near that of the sun Z, that the eclipse is nearly central, and as the moon's semi-diameter is greater than the sun's, the eclipse must be total. Under similar circumstances if the moon's semi-diameter had been least, the eclipse could have been annular. In case of a partial eclipse the sun's disc will not be wholly covered by the moon, as in fig. 11, Plate XII. where the circles representing the discs of the sun and moon are marked with the same letters as in fig. 10, but the objects are placed in a different situation. In this case the number of digits eclipsed may be obtained by drawing a line through the centres p, Z, to meet the discs in the points S, M, s, m, and by saying as the distance Ss (representing the whole disc) is to the obscured point M_s so are 12 digits to the number of digits eclipsed. The beginning and end of total darkness in α total eclipse are found like the beginning and end of the eclipse, except in taking in the compasses the difference between the semi-diameters of the sun and moon, instead of their sum. For the points of the path of the spectator and of the moon's orbit, marked with the same time, and at that distance from each other, will represent the situations and times of the beginning and end of total darkness. The beginning and end of the internal contacts of an annular eclipse are found in the same manner, the only difference is that in a total eclipse, the moon's semi-diameter is greatest, but in an annular eclipse the least.

In observing the beginning of a solar eclipse, it is of some importance for the accuracy of the observation, to know on what part of the sun's limb the eclipse will begin. This is easily found by means of the projection. Thus at the beginning of the eclipse, which corresponds to the point p' of the moon's path and the point Z' of the path of the spectator, the first point of contact g may be obtained by drawing about the centre p' with a radius equal to the moon's semi-diameter, a circle representing the moon's disc*; about Z' as a centre with a radius equal to the sun's semi-diameter another circle representing the sun's disc, touching the former in the point g.

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^{*} Instead of this virule, the line p' Z' may be drawn cutting the sun's disc in the sought point of contact x.

Draw the line CZ' meeting the sun's disc in the points a, c, the point c being the most distant from the centre C. Then the circle g a c being held between the cyc of the observer and the sun, the engraved or marked side of the figure towards the cyc, and the line c a in a vertical direction with the point c uppermost, will represent the appearance of the sun as viewed by the naked eye at that time, c will represent the upper part of the sun, a the lower, and g the point of contact. If the eclipse be observed with an inverting telescope, the contrary will be observed; that is, the part a must be uppermost, c the lowest, and g the point of contact will appear to the left hand of ca. In a similar manner the appearance of the objects may be obtained at any other part of the eclipse, but it is not necessary except at the beginning of it, where there is nothing to direct the eye of the observer.

EXAMPLE.

Bequired the times and phases of the total eclipse of the sun, June 16, 1806, at Salem, in the latitude of 42° 33′ 30″ N. and the longitude of 4h. 43m. 32s. west from Greenwich?

By the Nautical Almanac the time of new moon at Greenwich was June 16d. 4h. 19', corresponding to June 15, 23h. 35' 28", at Salem. At the time at Greenwich, 4h. 19' the elements of the eclipse were as in the adjoined table calculated by the above rule.

Draw ACB (Plate XII. fig. 10,) and perpendicular thereto the line CGR. Make CG equal to the moon's latitude 19'37" N. taken from a scale of equal parts, the point G being above C because the latitude is north. Make CO equal to the moon's horary motion from the sun 34' 18".1, to the right hand of the point C; and CP equal to the moon's horary motion in

ELEMENTS.			1
Conjunction at Greenwich, June 16, Salem W. from Greenwich Ecliptic conjune. at Salem, June 15, Latitude of Salem)*a Horizontal Parallax ()*a Horizontal Parallax ()*a Horizontal Parallax ()*s Reduced Horizontal Parallax ()*s Semi-diameter (44	95 33 60 16 15 32	25.7 8.6 16.9 25.1 46.1 14.2 42.0
D's Horary motion in Long. Prob. II. (a)'s Horary motion N. A.		2	41.2 23.1
D's Horary motion from O's Horary motion in lattinde	7+	- 3	18.1 22.5
b's Latitude by Prob. L Co	7 84	ı 44	
O's Declination D	2	322	N.

latitude + 3' 22'.5, the point P being below C because this horary motion has the sign Draw NGL parallel to OP. Make OP a transverse distance of 60, 60, + prefixed. on the line of lines of the sector, and measure from the same lines the transverse distance 354, 354 (corresponding nearly to the minutes in the time of new moon) this distance set on the line GN to the right of G reaches to the point x where the hour preceding the new moon is to be marked, viz. 23h. Take OP in the compasses and preceding the new moon is to be marked, viz. 23h. Take OP in the compasses and mark it successively on the line NL from x or 23h. to the right to 22h. and to the left to 24h. or 0h. 1h. &c. These are subdivided into five minutes, the sale not admitting smaller divisions. Take the moon's reduced horizontal parallax 60' 16".9 from the scale of equal parts, and with that radius describe about the centre C the circle ARB. Set off (by means of the sector) the arches RT, RU, each equal to 23° 28'. Join TQU and about that diameter describe the circle TYU. Make the arch TV equal to the sun's longitude 84° 44′ 36″, which is done by setting the radius QT as a chord from T to II, and then the arch \(\preceq V = 24\) 44′ 36″ by means of the sector. Draw P'V parallel to CR to meet TU in the point P'. Join CP' and continue it to meet the circle ARB in W. Make (by the sector) the arches WD, Wd, equal the complement of the latitude of the place 47° 26′ nearly, the radius being CB. In a similar manner make the arches DP, DE, df, de, &c. each equal to the sun's declination 23° 22'. Draw the lines Flf, Dad, Ene, cutting CW in l, q, n. Bisect ln in r. Draw the line VI, r, XVIII parallel to Dqd and make r, VI, r, XVIII, each equal to qD. Through the points l, VI, n, XVIII, l; draw the path of the spectator as taught in the above rule, and mark the hour of noon 6h. at the point a because the sun's declination is north. Mark the following hours in succession to the left I, II, III, &c. as in the figure. Take an extent in the compasses equal to the sum of the semi-diameters of the sun and moon 32 14".2 and beginning towards N, find, as above directed, the points p'Z' at that distance apart and marked with the same time 22b. 7' nearly, which is the time of the beginning of the eclipse. Proceed in the same way for the end of the eclipse corresponding to the points p'', Z'', and to the time 0h. 53' which is the time of the end of the cclipse. Take the difference of the semi-diameters of the sun and moon 42" in the compasses, and proceed in the same way to find the beginning and end of total darkness 23h. 27m. and 23h. 31m. The paints corresponding could not be drawn in the figure as they are so near to p and Z, and the scale small. Find by trials the points p Z marked with the same time and at the least distance apart, this will be the time of the middle of the eclipse 23h. 29'. With an extent equal to the moon's semi-diameter 16' 28'.1 as a radius, describe about p the circle whose diameter is Min

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representing the moon's disc, and with the sun's semi-diameter 15' 46".1, describe about Z the circle whose diameter is S s representing the sun's disc at the middle of the selipse. The sun's disc being wholly covered by the moon, indicates that the eclipse was total. Describe in the same way about p' and Z' the discs of the sun and moon, at the beginning of the eclipse, touching each other in g. Draw CZ' cutting the moon's disc in c and a. Then the arch c g will be the distance of the first point of contact of the sun and moon from the sun's zenith towards the western part of the limb.

REMARKS. 1. The correction for the spheroidal form of the earth, the augmentation of the moon's semi-diameter, inflexion and irradiation, are neglected in the above rule as not sensibly affecting the result of the projection, though these points might be attended to

by the following precepts.

2. From the latitude of the place subtract the correction of latitude of Tab. XXXVIII. and from the moon's horizontal parallax decreased by 8".8 subtract the correction of parallax in the same table; the remainders will be the corrected latitude and parallax to be made use of in the above rule to correct for the spheroidal form of the earth.

3. Decrease the moon's semi-diameter given by the N. A. by 2' for inflexion.
4. Decrease the sun's semi-diameter 3'4 for irradiation, and from the remainder subtract a correction equal to the augmentation (Tab. XV.) that the moon's semi-diameter would have when at the same altitude as the sun, the remainder will be the corrected semi-diameter of the sun, to be used in the above rule in finding all the times and phases of the eclipse. This method of decreasing the sun's semi-diameter produces nearly the same result as that by augmenting the moon's semi-diameter, horary motion and horizontal parallax, and taking the sun's semiliameter as given in the Nantical Almanac.

5. Besides these corrections, there are others depending on the change of the moon's semi-diameter, horizontal parallax and horary motion during the eclipse, but all these

corrections are usually neglected in projecting an eclipse or occultation.

6. The altitude of the sun, which is nearly the same as that of the moon during the eclipse, may easily be found by means of the projection. Thus if it were required at the beginning of the eclipse when the spectator is at Z': Take the distance CB and apply it as a transverse distance 90°, 90°, to the sines of the sector; then the distance CZ' applied in the same manner to those lines, will give the zenith distance of the sun, about 310, corresponding to the altitude 590. The correction (Table XV.) corresponding to this altitude is 14", which is nearly the correction to be subtracted from the sun's semi-diameter 15', 42".6 (corrected for irradiation) to obtain the corrected semi-diameter 15' 29".6, as taught in §4. Table XV. was calculated for the mean semi-diameter 15' 37" and the correction of the Table 14" ought to be increased in ratio of the sun's semi-diameter 15' 46".1 to 15' 37" when very great accuracy is required. The difference of the corrected semi-diameters of the sun and moon 15' 28".6 and 16' 26".1 is 57" 24, which is to be used instead of 42" in finding the beginning and end total rkness. The duration of the total darkness found by the corrected values 57" is 42 minutes, but with the uncorrected value 42" is only 34 minutes. It was probably owing to the neglect of this correction that some of the Almanacs published in this country, for 1806, menzioned the duration as 3 minutes.

7. The path of the spectator I, II, III, IV, &c. calculated for the proposed latitude 43° 33' 30' may be made to answer for any other latitude by altering the centre of projection and the scale of equal parts. By this means the trouble of repeatedly describing that path, when the eclipse is to be calculated for several places, may be avoided. To do this add the Prop. Lög. of the reduced parallax to the log. secant of the latitude of the place, the sum, rejecting 10 in the index, will be the Prop. Log. of an arch A. To this Prop. Log. add the log. secant of the sun's declination (or star's in an occultation) and the log. co-tangent of the latitude of the place, the sum, rejecting 20 in the index, will be the Prop. Log. of the arch B. Take the radius r, VI (or qD) in the compasses, and make it a transverse distance on the line of lines of the sector corresponding to the arch A, and with that opening of the sector measure the transverse distance corresponding to the arch B which set from r towards C on the line rC (continued if necessary) will reach to the centre of the projection corresponding to the proposed latitude; the transverse distance corresponding to the reduced parallax measured from the line of lines, with the same opening, will be the radius of the projection, and the transverse distance corresponding to the horary motion of the moon from the sun or star in an occultation, will be the horary distance to be made use of in marking the hours on the lunar orbit LN; lastly, the latitude of the moon at the conjunction is to be measured as a transverse distance, and set from the new centre of projection on a line drawn through it parallel to CR, and the point where it reaches will be the new point G corresponding to the place of the moon at the ecliptic conjunction. Through this point the line of the moon's path is to be drawn parallel to the line LN of the figure, and the hours are to be marked

on it as before. Whence the times of beginning and end of the eclipse may be found as in the above rule. An example of this method is not given, as it would render the scheme too confused.

PROBLEM XII.

To project an occultation of a fixed star by the moon, at any given place. The method of projecting an occultation is nearly the same as that of an eclipse of the sun, but to save the trouble of reference it was thought expedient to give the rule without abridgment.

RULE.

To the time of the ecliptic conjunction of the moon and star, given in the first page of the Nautical Almanac (or calculated by Prob. III.) add the longitude of the propose place turned into time, if east, but subtract if west, the sum or difference will be the time of conjunction at the proposed place. Corresponding to the time of conjunction at Grasnwich, find by Problem I. the moon's latitude, horizontal parallax and semi-diameter, also the sun's right ascension. Then by Problem II. find the horary motion of the most in longitude and latitude, and by Tables VIII. and XXXVII. the star's Right Ascension, Declination, Longitude and Latitude.

Draw the line ACB (Plate XII. fig. 8.) representing a parallel of the ecliptic passing through the star, and perpendicular thereto the line CPR. Take a scale of equal parts to measure the lines of the projection, and from it take an interval equal to the difference of the latitudes of the moon and star, and apply it to the line CR from C to Gabove the line ACB if the moon's latitude is north of the star's, otherwise below; Take CO equal to the horary motion of the moon in longitude, and set it on the line CB to the right hand of C to O; take CP equal to the moon's horary motion in latitude found with its sign by Problem II. and set it on the line CR from C to P, above* the hine ACB, if its sign is -, below if +. Join OP which represents the horary motion of the moon on her orbit, and parallel to that line draw the orbit of the moon NGL, on which are to be marked the places of the moon before and after the conjunction by means of the horary motion OP, so that the moment of the ecliptic conjunction at the proposed place may fall exactly at the point G, as in the figure where the conjunction is at 18h. 42°. This may be done by making OP equal to the transverse distance 60, 60, on the line of lines of the sector, then measuring from the same lines the transverse distance corresponding to the minutes and parts of a minute in the time of the ecliptic conjunction at the place of observation, and setting it on the line GN from G towards the right to the point x, the place of the moon at the first whole hourt preceding the conjunction (which in the present figure is 18h.) Then the distance OP being taken in the compasses, and set from x to the right hand, gives successively the preceding hours, and the same distance set to the left gives the following hours, as in the figure, where they are marked 17h. 18h. 19h. 20h. These hours are to be divided the figure, where they are marked 17h. 18h. 19h. 20h. into 60 equal parts representing minutes, the scale being taken sufficiently large for that purpose. In the present figure the subdivisions are carried only to five minutes. Take the moon's horizontal parallax from the scale of equal parts for the radius CB; with which on the centre C, describe the circle BRA cutting CR in R. Open the sector till the transverse distance 60°, 60°, on the line of chords, is equal to the radius CB, and measure from that line the transverse distance 23° 28' (equal to the obliquity of the ecliptic) which set on the circle ARB on each side of R to T and U. Join TU cutting CR in Q. On Q as a centre, with the radius QT, describe a circle TYUV, on which set off the arch TYV, equal to the star's longitude. Through V draw the line VP parallel to CR. Open the sector till the transverse distance 90°, 90°, on the sines, is equal to the radius CB, then take in the compasses from the same lines an extent equal to the transverse distance corresponding to the complement of the declination of the star, and with one foot in C sweep a small arch to cut the line VP' in P' the place of the pole of the earth.*† Draw CP', and continue it on either side so as to cut the circle ARB in the point W situated above AB, if the latitude of the proposed place is north, but below if wouth. In the proposed figure the latitude is north. (If it had been south the lower part of the circle ARB ought to have been made use of.) Open the sector

[¶] In strictness these quantities ought to be corrected for Aberration and Nutation by Tables XXXIX.—XLIII. but the correction is so small that it may always be neglected. If the Right Ascension and Declination only are given, the latitude and longitude may be found by Problem XIX. and if the latter are given, the former may be calculated by Problem XX.

†* In the figure the point G is placed above ACB, because the moon is in a less southern latitude than the star. This part of the rule may also be thus expressed. Find the moon's latitude with its sign as in Problem II. Prefix the sign + to the star's latitude in north, the sign — if south. Add the latitudes, noticing the signs as in algebra, and the distance CG will be obtained. If its sign is—the point G is to be placed above C, but below C if the sign is +.

See note with this mark in page 502.

† See note with this mark in page 592.

† See note with this mark in page 594.

The distance of the line WV from the line CR, the situation of the point I and the path of the spectator, may be found as in the note § page 594.

as before so as to make the transverse distance of 60° , 60° , on the chords, equal to CB, and take the chord of the complement of the latitude of the place, which set from W on each side, to D and d. With the same opening of the sector measure the chord of the star's declination, which set on the circle ARB from the point D on each side, to E and F. and fr m d on each side to e and f. Draw the dotted lines Ff, Dd, Ee, cutting CW in l, q, n. Bisect l n in r, and erect the line t r u perpendicular to CW and make rt, ru each equal to qD. Open the sector to make the transverse distance 90°, 90°, on the sines equal to r 1, and on each side of r mark on the line t r u the sines of 15°, 30°, 45°, 60°, 75°, (equal to 1h. 2h. 3h. 4h. 5h. respectively) to that radius, and mark the points with those degrees as in the figure; through these points draw the dotted lines parallel to ln as in the figure. Open the sector so that the radius r l may correspond to the transverse distance 90° , 90° , on the sines, and measure the complements of the former degrees as transverse distances on the sines, viz. 75° , 60° , 45° , 30° , 15° , and set them on the above dotted lines, on each side of the points 150, 300, &c. respectively. above and below the line tru. A regular curve it lun drawn through the extremities of these dotted lines, will represent the path of the spectator in the given latitude. Subtract the sun's right ascension from the star's (increasing the latter by 24 hours when necessary) the remainder will be the hour of the star's passing the meridian,* which is to be marked at the upper point l of the path if the star's declination is south, but at the lower point n if the declination is north. The other hours are to be marked from this point towards the left, by marking successively, at the points where the dotted lines meet the path, the hour of the star's passing the meridian, increased by 1h. 2h. 3h. &c. completely round the curve, observing to reject 24 hours when the sum exceeds 24h. In the present example the star's declination is south, consequently the upper point l of the path is taken for the hour of passing the meridian 19h. 54'. The extremities of the dotted lines to the left being marked successively 20h. 54', 21h. 54', 22h. 54', 23h. 54', Oh. 54', &c. The path touches the circle ARB in two points, representing the points of rising and setting of the star, which in the present figure are 14h. 9' and 1h. 39'. These points divide the path into two parts, of which one represents the path while the star is above the horizon, the other when below, as is evident from the hours marked on the curve. The half hours or any other intermediate time may be marked in a similar manner. Thus, for the time 4h. 24', which is 3h. 30' or 52° 30' from the time 7h. 54', marked at the point n; set the sine of 52° to the radius r t from r to h on the line rt, and crect the perpendicular hi, equal to the sine of 37% (which is the complement of 52%) to the radius m, and the point i will represent the place of the spectator at the proposed In this way the halves and quarters of hours may be marked on those parts of th where necessary. The smaller subdivisions may generally be obtained to a the path where necessary. sufficient degree of exactness by dividing the quarters of hours into equal parts.

Take from the scale of equal parts an extent equal to the semi-diameter of the moon, and beginning at the line NL towards N, find by trials the point p' of the moon's path. and the point Z of the path of the spectator, marked with the same time and at that distance apart. That time will be the beginning of the occultation or immersion at the proposed place. Proceed in the same way towards the point L, and find the points p, Z, at the same distance apart, the corresponding time will be the end of the occultation er emersion. About the points p', p, as centres, with a radius equal to the moon's semidiameter, describe the small circles meeting the paths of the spectator in the points Z', Z. These circles will represent the moon's disc; the points Z', Z, the places of the star, and the lines CZ', CZ, the vertical circles passing through the star at the times of immersion and emersion respectively. To render this part of the scheme more distinct to the eye, it is drawn separately in Fig. 9, Plate XII. in which the point C, p'Z', are similarly situated to the corresponding points of Fig. 8, marked with the same letters. Through p' draw the line a' p c' parallel to CZ', to meet the moon's disc in a', c'. Then the circle a Z' c' being held between the eye of the observer and the sun, the engraved or marked side of the figure towards the eye, and the line CZ' (or a' p' c') in a vertical position with the point Z' above C, will represent the appearance of the moon and star as viewed by the naked eye, c' will represent the upper part of the moon, a' the lower part, and Z' the point of contact. The contrary will be observed if the object be viewed by an inverting telescope. It will generally be conducive to the accuracy of an observation to estimate in this manner the point of emersion, so as to keep that point of the moon's limb in the field of view of the telescope, and the eye directed towards that point of the limb, so as to perceive the star at the first instant of its appearance.—The situation of the point of emersion with respect to the horns ρ , θ , of the moon may also be made use of for this purpose. The line op? connecting the moon's horns, is nearly parallel to the line CR, except very near the new or full moon, so that in general it will be sufficiently

[•] Or rather the horary distance of the ② and ** at the time of the coliptic conjunction of the saxon and dar.

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correct to draw through p the line p p θ parallel to CR. If greater accuracy is required; the following construction may be made use of. Subtract the sun's longitude from the moon's, * make the arch TYU λ equal to the remainder, and join $\mathcal{Q}\lambda$. Set on the same circle the arch $T\beta$ equal to the moon's latitude; below the point T if that latitude is south, above if north. Through β oraw the line β ε deparallel to TQ to cut $Q\lambda$ in ε and εR in δ . Take the extent QT and set it on the line δ Y above δ to μ . Join μ ε and parallel thereto through p draw the line p p θ cutting the moon's disc in the points p θ representing the horns, the figure being viewed as above directed. The enlightened part of the moon is that nearest to the sun, the dark part is the most distant from it.

BXAMPLE.

Required the times of immersion and emersion of Spica, Dec. 12, 1808, at a place in the latitude of 20° N. and in the longitude of 1h. 9m. east from Green-rich?

By the first page of the Nautical

By the first page of the Nautical	LEMENT	
Almanac for the month of December,	Conjunction at Greenwich Dec. 12	17h. 33
IMIM the time of the ecliptic con-	Il operius de la fementa	1 9
Junction of the moon and Spica (marked) a 117) was, December 12, 17h.	Conjunction at place of observation	18 42
ed D a III) was December 12, 17h.	X 8 R. ascen. Tab. VIII.	13 15 06
		17 21 18 19 63 50
18h. 42' at the proposed place. At	Latitude of the place	200 0 0
		59 55.2
elements of the seculation was the	D's semi-diameter by N. A.	16 19.8
		35 55.2
in the adjoined Table calculated by	D's horary mot. in lat. Prob. IL. CP X's longitude, Tab. XXXVII. TYV	- 3 02.7 201 10 51
the above rule.	N's letitude by N. A.	49 55 8.
Draw ACB, and perpendicular	* latitude, fab. XXXVII.	2 2 13 8.
thereto the line CGY. Make CG	Diff of la itudes D. N. of * CG	12 20N.
equal to the difference between the	X' decimation	10 10 8.
latitudes of the moon and star 12 20" to	aken from a scale of equal parts, th	e point G be-
ing above C, because the moon is north		
horary motion in longitude 35' 55'.2 to		
tion in latitude 3'2".7, the point P bein		
is south decreasing.) Draw NGL part	allel to OP Make OP a transver	e distance of
60, 60, on the line of lines of the sector,	and maggire from the same lines t	he transmess
distance 42, 42 (corresponding to the	-inutes in the time of the conjunct	tion) this dis
the second of the line CN from C to the i	minutes in the time of the conjunc	non, can us-
tance set on the line GN, from G towar		
path where the hour preceding the conj		
the compasses, and mark it on the line l		
left to 19h. 20h. &c. These are subdiv	ided into five minutes, the scale n	ot admitting
of smaller divisions. Take the moo		
parts, and with that radius describe ab		
means of the sector) the arches RT, RI	J, each equal to 23° 28′. Join TQ'	U, and about
that diameter describe the circle TYUV?	r. Make the arch TYV equal to tl	re star's lon-
gitude 201° 10' 31", which is done by	making the arch UV=21° 10′ 31″.	. Draw PV
parallel to CR, and with an extent equ		
79° 50', taken as a transverse distance		
one foot in C, sweep an arch cutting I	P'V in P'. Join CP' and continue i	t to meet the
circle ARB in W. Set on each side	of W the arches WD Wd same	to the com-
plement of the latitude of the place 7		
equal to the star's declination 10° 10',		
ting OW in l , q , n , Bisect l n in r ,		
equal to q D. Through the points l	, i, n, u, i, draw the path of the	spectator as
taught in the above rule, and mark th		
53' 50" or 19h. 54', at the upper po		
Mark the following hours in succession		
	asses equal to the moon's semi-	
19".8 and beginning towards N, find	as above directed the points p' , Z' .	. at that dis-
tance apart, and marked with the same	time 16h. 57', which is the time of	the immer-
	e emersion corresponding to the	
at the same distance apart, and the tim		
With the same extent describe about	and p' the small circles represent	ing the disc

^{*} In strictness the longitude and latitude of the moon at the time of immersion or emersion ought to be made use of, but it will be sufficiently exact to use the star's longitude instead of the moon's (increasing it by 360° when less than the sun's longitude) and the moon's latitude at the conjunction. Quantities of the same order as the moon's parallax are neglected in the value of the arch TYU).

of the moon at these times, and cutting the path of the spectator in the point Z, Z'. Join CZ', Cp', and parallel to CZ', draw c', p' a' cutting the moon's disc in c', a' (as in Fig. 9, P. XII.) and the arch a Z will represent the distance of the point of immersion from the lower part a' of the moon. The line CZ runs nearly through the point p, so that the top part of the moon c and the point Z nearly coincide, consequently the emersion happened near the moon's zenith. By subtracting the sun's longitude 261° 7' from the moon's or star's 201° 10', (increased by 360°) the remainder is 300° 3', which is to be marked on the circle TYUV to the point λ . Make the arch $T\beta$ equal to the moon's latitude 1° 49' 53'', taking the point β below T, because the latitude is south. Draw the lines λ Q β_L δ , μ , ρ , ρ θ as in the rule, and the points ρ , θ , will represent the Φ places of the moon's horns. The point of emersion Z will be to the westward of the upper horn ρ , about 60° measured on the moon's limb.

REMARKS

1. When it is thought necessary to take notice of the spheroidal form of the earli, the corrections of latitude and parallax of Table XXXVIII. must be subtracted from the latitude of the place and the moon's horizontal parallax respectively, to obtain the latitude and parallax to be made use of in the above rule.

2. Subtract 2' from the moon's semi-diameter given by the N. A. the remainder is

to be made use of without augmentation, on account of the altitude of the moon.

3. The corrections for the change of the moon's semi-diameter, horizontal parallax, and horary motion during the occultation, are neglected in the above rule, as not materially affecting the result.

4. The line C Z', measured on the sines as a transverse distance to the radius C B, will be the star's zenith distance at the immersion. In a similar manner it may be

found at the emersion at Z, or at any other point.

5. The curve l t n u may be made to answer for any latitude, as in Problem XI. Remark 7.

Calculation of an Occultation of a Planet by the Moon.

By a similar process the times of immersion and emersion of a planet may be calculated by finding the planet's right ascension and declination, geocentric longitude and latitude from the Nautical Almanac, and using them instead of the star's. Also by Prob. II. the horary motion of the moon from the planet in longitude and latitude, which are to be used instead of the horary motion of the moon. In this projection it will not be necessary to take notice of the parallax of the planet, but it may be easily allowed for, by taking the radius C B equal to the difference of the horizontal parallaxes of the moon and planet. The apparent diameter of the planet may also be neglected, making the distances pZ, p'Z' equal to the moon's semi-diameter. When great accuracy is required, the sum of the semi-diameters of the moon and planet must be made use of for finding the external contacts, and their difference for the internal centacts.

PROBLEM XIII. To calculate the beginning or end of a solar Eclipse.

RULE.

This must be done by approximation, by assuming a time for the beginning or end of the eclipse, as for example the time obtained by projection by Problem XI. the time of new moon at the place of observation, or an hour before or after, according as it is the heginning or end of the eclipse that is sought. With this time calculate the elements of the eclipse and the parallaxes, as taught in the first part of Problem VIII. The parallaxes applied to the longitude and latitude of the moon by the N. A. will give the apparent longitude and latitude. Find the difference of the apparent longitudes of the moon and sun, and from its prop. log. increasing the index by 10 subtract the prop. log. of the moon's apparent latitude, the remainder will be the log. tangent of an angle, whose corresponding log. co-sine is to be added to the prop. log. of the diff. of longitudes, the sum, rejecting 10 in the index, will be the prop. log. of the apparent distance of the centres of the sun and moon, which ought to be equal to the sum of the corrected semi-diameters, if the assumed time was correct. If this is not the case, the operation must be repeated with an assumed time differing a few minutes from the former, and the apparent distance of the centres of the sun and moon must be calculated in this new supposition. Then add together the arith. comp. of the prop.log. of the difference of the apparent distances thus calculated, the prop. log. of the difference between the first calculated distance and the sum of the semi-diameters, and the prop. log. of the interval of time between the two suppositions, the sum, rejecting 10 in the index, will be the prop. log. of the correction to be applied to the first assumed time, which at the beginning of an eclipse is to be added to the first assumed time, if the distance be greater than the sum of the semi-diameters, but subtracted if less; and the contrary in calculating the end of an eclipse; the sum or difference will be the approximate time of the beginning or end

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of the eclipse. If great accuracy is required, the operation may be repeated with this approximate time, combining this result with one of the former suppositions, and thus the operation may be repeated till the apparent distance of the centres at the assumed

time is found to be exactly equal to the sum of the corrected semi-diameters.

Remark. This rule, with some modification, will answer for calculating the time of an occultation of a fixed star or planet by the moon. In this case the star's longitude is to be found in Table XXXVII. and corrected for the equation Tables XL. XLI. (or the planet's longitude is to be taken from the Nautical Almanac) the difference between this and the moon's apparent longitude corresponding to the assumed time being found, its prop. log. is to be added to the log. secant of the moon's apparent latitude, and the sum is to be used in finding the distance of the centres instead of the prop. log. of the diff. long. of the sun and moon, with the index increased by 10. The latitude of the star is to be found by Tables XXXVII. and XLl. or the planet's latitude by the Nautical Almanar, and added to the latitude of the moon, if of a different name, otherwise their difference is to be taken and made use of, instead of the moon's latitude in the above rule. Lastly, instead of the sum of the semi-diameters, the semi-diameter of the moon is to be made use of. When very great accuracy is required, in calculating an occultation of a planet by the moon, the difference of the parallaxes of the moon and planet decreased by the correction of Parallax l'able XXVIII. is to be made use of as the reduced parallax, in finding the parallaxes in longitude and latitude. When the apparent distance of the centres of the moon and planet is equal to the sum of their semi-diameters, their limbs will just appear to touch each other, and when that distance is equal to the difference of the semi-diameters, the planet will be wholly covered by the moon.

EXAMPLE.

Required the time of the beginning of the solar eclipse of June, 1806, at Salem, supposing the errors of the moon's longitude and latitude in the Nautical Almanac to be unknown?

To abridge the present calculation, suppose the beginning of the eclipse to be June 15d. 22h. 6' 18".1 app. time the elements corresponding to which have been calculated in Problem VI.; namely, I's apparent longitude 84° 8′ 50″.3, D's apparent latitude, 1' 55".8 N. these being corrected for the errors of the tables, 56".5 and 11".4, hence the uncorrected values are 84° 9'.48".8, and 2' 7".2 N. The difference between this app. long. of the moon and the sun's longitude 840 41' 3".4, is 31' 14".6.

Diff. long. App. Lat.	31' 14' .6 P. L. 2 7 .2 P. L.	10.7605 1.9 2 89		0.7808
		Tan. 8.8316—0	Corresponding co	-sine 9.9890
		App. Dis. B	81'. 19"A	P. L7596

This apparent distance differs 1' 4".5 from the sum of the semi-diameters 32' 23'.5. It is therefore necessary to make a second supposition, as for example ten minutes later, or at 22h. 16' 19".1, with this time the elements are to be again calculated as in Problem VI. namely, D's app. long. uncorrected 84° 14' 7".1, \odot 's long. 84° 41' 27".2, their difference 27' 10".1, D's app. lat. uncorrected for error of tables 1' 55".8 N.

Diff. long. B App. Lat.		P. L. P. L.		.8212 .9586		0.8212
	Tang.		8	.8626	Corresp. co-sine	9.2988
. į s	Second App. Dist.	⊙)	27'.	14".7	P.	L. 8200
I	ïrst App. Dist.	⊙ D	31.	19 0		
1	iff. 1st. dist. & Se	ference mi-di <u>am</u> . erval	4. 1. 10.	4. S 4. 5 0.	P. L. Ar. eo. P. L. P. L.	8.3545 2.2233 1.2533
	Cor	rection	2.	38.	P. L.	1.8836
F	irst supposed tim	e 15d.	22h.	6. 18.1	`	
A	poroximate time	15.	92.	s. 40.1		

If this approximate time had differed very much from the assumed times, it would be necessary to repeat the operation till the last assumed and calculated times agree.

PROBLEM XIV.

Given the moon's true longitude to find the apparent time at Greenwich. RULE.

1. Take from the Nautical Almanac the two longitudes immediately preceding

the given longitude and the two following, and find the first and second differences as in Problem I. Call the middle term of the first differences the arch A, and the half sum of the second differences (noting the signs) the arch B.

2. To the constant logarithm 4.63548 add the arithmetical comp. log. of A in seconds, and the logarithm of the difference in seconds between the given longitude and the second longitude taken from the Nautical Almanac, the sum, rejecting 10 in ex,

will be the logarithm of the approximate time T in seconds.

3. Enter Table XLV. with the arch B at the top, and this time T at the side, and find the corresponding correction: to the logarithm of which add the two first logarithms above found, the sum, rejecting 10 in the index, will be the correction of the approximate time to be applied with the same sign as the arch B, and the correct apparent time, counted on from the second noon or midnight, will be obtained.

EXAMPLE

Suppose the moon's longitude Dec. 12, 1808, was 6s. 19° 38' 58". Required the apparent time? As in Example I. Problem I. $A=7^{\circ}$ 11' 18''.=25878". B=+4' 54".5 and the difference between the given longitude and the second longitude, taken from the Nautical Almanac, 6s. 17° 51' 36" is 1° 47',22",=6442". Hence (as in the following calculation) the approximate time past midnight is 2h. 59' 14'', this and the arch B gives, in Table XLV, the equation 27".5 whence the correction is +46', and the sought time 3h, past midnight or Dec. 12d. 15h.

A Diff. long.	Constant 25878" Ar. co. 6442"	log.	4.63548 5.58707 3.80302	Eq. Table XLV. 27"	.5 Log.	4.63548 5.58707 1.48935
Approx. time 2h. 59'	14''=10754''			Correction+46".	_	1,66188

PROBLEM XV.

Given the distance of the moon from a fixed star not marked in the Nautical Almanac, together with the altitudes of the objects, the apparent time of observation, and the estimated longitude, to find the longitude of the place of observation.

RULE.

To the apparent time of observation, by astronomical computation, add the estimated longitude in time, if west; subtract, if east, the sum or difference will be the supposed time at Greenwich, corresponding to which find the moon's latitude by Problem Le also the longitude and latitude of the Star by Table XXXVII. and correct them for absertion and nutation, by Tables XL. XLI.

With the apparent altitudes and distance of the objects, find the correct distance by

the usual rules of working a lunar observation.

To the correct distance add the latitudes of the moon and star, and find the difference between the half sum and the distance. Then to the log secants of the latitudes of the moon and star, rejecting 10 in each index, add the log co-sines of the half sum and difference if the latitudes are of the same name; or the log sines if of a contrary name; half the sum of these four logarithms will be the log co-sine of half the difference of longitude if the latitudes are of the same name, or its log sine if of a different name.

The difference of longitude is to be added to the apparent longitude of the star if the moon is east of the star, otherwise subtracted (borrowing or rejecting 360° when necessary;) the sum or difference will be the true longitude of the moon, whence the time at Greenwich may be found by Problem XIV. The difference between this and the apparent time at the ship will be the longitude, which will be west if the apparent time at Greenwich be greater than the time at the ship, otherwise east.

REMARK.

This method, with a slight modification, will answer for finding the longitude from the observed distance of the moon from a planet, as Jupiter, Venus, Mars, or Saturn. The only difference consists in finding from the Nautical Almanac by Problem I. the Geocentric long, and lat. of the planet, which are to be used instead of the longitude and latitude of the star in the above rule. For the daily variation of the longitude and latitude of a planet is so small, that no error of moment can arise from calculating those quantities for the supposed instead of the true time, at Greenwich, and the parallax and semi-diameter of the planet are so small as not to affect the calculation materially.

The latitudes of the moon and the fixed star or planet made use of in these observations, ought not to differ very much, on account of the decrease of the relative motion arising from this source. If the latitudes are of a different name, their sum; othcrivise their difference ought to be found, and if it does not exceed one third part of the

difference of longitude of the two objects, they may in general be made use of.

EXAMPLE.

Suppose that on the 7th. January, 1808, sea account, at 6h. 57m. P. M. in the longitude of 120° W. by account, the observed distance of the farthest limb of the moon from the star Aldebaran was 39° 7′ 4″, the observed altitude of the star 43° 18′, and the observed altitude of the moon's lower limb 52° 52′. Required the true longitude, without using the distances marked in the Nautical Almanac, upon the supposition that they were not given in it?

In this case the supposed time at Greenwich was Jan. 6d. 14h. 37m. D's horiz. par. 54' 35", D's S. D. 15' 5". Apparent distance of centres D * 38° 51' 59", whence (by the rule page 167) the correct distance is 38° 47' 26". The Moon's latitude deduced from the Nautical Almanac by Problem I. is 2° 37' 36" N. The Star's longitude and latitude is found by Tables XXXVII. XL. XLI. making use of the longitude of the Moon's node 7s. 28° 15', and the Sun's longitude 9s. 15° 43', as given in the Nautical Almanac.

Table XXXVII. * Long Table XLL * Aberration		67°.6′.11′ Å + 15 .3		5°.28'.45'.6 S. + 1.2
Table XL. Equal. Equino	x	+ 15 .2	* App. Lat.	5 28.51 S.
* Apparent longitude	_	67. 6. 52	-	,
Gorrect Distance D Latitude ** Latitude		5.28. 51 8	N, Sec. 0.00046 S. Sec. 0.00159	
•	-	n 46.53, 58	•	
, Diff. 🛔 su	Half sun m and dist.	23.26, 56 15 20, 30	Sine* 9.500% Sine* 9.42255	•
•			19.02481	. •
h Diff. of Long. Diff. of Long. **s Longitude .	•	18.59. 21 37.58. 42 67. 6. 52	Sine* 9.51249 D West of 未	
Longitude Lorg. Jan. 6d.[12h]		29. 8. 10. 27.49. 6		
Difference		1.19. 4=	=4744'' Diff.	
\	6. 12 27. 7. 0 33. 7. 12 39. ht log. 4.635 g. co. 5.668	53. 10 49. 6 46. 32 46. 6	5.55.58 5.57.26 5.59.34 Mean B.	4.63548 4.66865
Approx. time 2b. 39' 16"=3 Correction + 19	556" Log. S.980	27 Co	rrection + 19" L	og. 1.27726
	Hence time at App. time at sh		14h. 39' 35'' 6. 37. 0	
	Longitude	IEM VX		120°.55° į W.

PROBLEM XVI.

Given the intervals of time between the passages of the moon's limb and a fixed star over two different meridians, to find the difference of longitude of the two meridians.

In making these observations it is usual to note the times of transit by a clock regulated to siderial time, being the most convenient for calculation. If the intervals are given in mean solar time, they may be reduced to siderial, by adding a proportional part of the daily difference 3' 56".6. Thus if the interval was 6 hours mean time, the correction would be found by saying as 24h.: 6h.::3' 56".6: 59".1, which added to 6h. gives the interval in siderial time 6h. 0' 59".1. In the following rule it is supposed that the intervals are given in siderial time. The constant logarithm 4.63667 made use of in the rule, is the logarithm of 43318 seconds, the number of seconds siderial time in half a mean solar day. In strictness this quantity ought to be equal to the logarithm of the number of seconds siderial time in 12 hours apparent time, which may differ 15 seconds from 43318" on account of the daily variation of the equation of time. The correction arising from this source is very small, and may in general be neglected, though it can be allowed for in a very simple manner, since the logarithm varies an unit in the fifth decimal place for 1" of time. Hence the correction of the logarithm is equal to half the daily variation of the equation of time in

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Use cosme if the intitudes are of the same name.

seconds, given in the Nautical Almanac, to be added to 4.63667 when the equation of time is marked add and is increasing, or sub. and decreasing; otherwise subtracted. Thus if the observation was made July 4, 1808, the equation of time is marked add, and is increasing daily 10".5. half of which or 5 is the correction to be added to 4.63667 to obtain the logarithm 4.63672, to be made use of July 4, 1808.

RULE.

If the moon be observed at both places on the same side of the star, take the difference of the observed intervals, otherwise the sum, which reduce to seconds of siderial time, and find the corresponding logarithm, to which add the arith. comp. log. of the variation of the moon's right ascension* in 12 hours in seconds, and the log. 4.63667 (corrected for the variation of the equation of time, as directed above, when very reat accuracy is necessary.) The sum, rejecting 10 in the index, will be the log. of a number of seconds, from which subtract the above difference of intervals, the remainder will be the longitude in time.

The western place of observation corresponds to the greater interval if the star is west of the moon, the less if east. If the moon be observed on opposite sides of the

star, the western place will be where the star is to the westward of the moon.

EXAMPLE.

Suppose that on the 4th. of July, 1808, the interval in siderial time between the transit of the moon's western limb and Antares, observed at Greenwich, was 22 6"; and the interval at a second place was 20' 3"; the increase of the moon's right ascension in 12 hours (corresponding to the middle time of the moon's transit by the meridians of the two places reduced to Greenwich time, 9h. 26') being by Prob. II. Ex. III. 30' 22".1 the star being to the eastward of the moon. Required the longitude of the second place of observation?

Interval at Greenwich 22'. 6" **Bid.** time at second place 20. 3

2. 3 == 123" 2.08991 Diff. of intervals Var. D R. A. in 12h. 30' 22".1. in time=1822".1 Constant log. Corrected as above. log. log. co. 6.73943 4.63872 2925"

Subtract diff. intervals Remains long. in time 2802-46' 42" W. from Greenwich.

This method of determining the longitude admits of a very great degree of accuracy on account of the frequent opportunities of observation. Other methods of finding the longitude depending on the same principles have been proposed. One consists in observing the apparent time of the moon's passing the meridian, and comparing it with the time of passing observed at Greenwich, or deduced from the Nautical Almanac, and taking the difference of these times, and saying, as the daily difference of the moon's passing the meridian (deduced from the Nautical Almanac for the time of observation) is to 360°, so is the above difference to the longitude of the place. Another method consists in deducing the longitude from the change of declination of the moon, obtained from her observed altitude when on the meridian, and the known latitude of the place of observation, by a method somewhat similar to the preceding; but neither of these methods is susceptible of the same degree of accuracy as that in the above Problem.

It is not absolutely necessary that the same star should be made use of at both places; for if two stars be observed, whose difference of right ascension is accurately known, that difference will be equal to the interval of passing of the two stars to the meridian in the siderial time, and by applying this to one of the intervals, the observations may be reduced to be the same as if one star only had been used.

PROBLEM XVII.

Given the longitudes of the sun and moon, and the moon's latitude, to find their distance. RULE.—Find the difference of the two longitudes, and to its log. co-sine add the log. co-sine of the moon's latitude, the sum, rejecting 10 in the index, will be the log. co-sine of the sought distance, of the same affection as the difference of longitude. EXAMPLE.

July 16, 1808, at neon at Greenwich, by the Nautical Almanac, the sun's longitude was 3s. 23° 40′ 24″, the moon's longitude 1s. 3° 14′ 1″, and her latitude 1° 24′ 28″ N. Required their distance?

† Two arches or angles are said to be of the same affection when they are both greater or both less than 90°, but of different affection when the one is greater and the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°, which is a state of the other less than 90°.

^{*} In general it will be exact enough to take the difference between the moon's R. A marked in the Nantical Almanac for the nearest noon and midnight, but when very great accuracy is required, it may be found as in Prob. II. Ex. III. for the middle time between the two transits of the moon by the meridians of the two places, reduced to Greenwith time by adding the longitude if west, subtracting if east. If absolute accuracy is required, it would be proper to notice the variation of the moon's semi-diamster and declination between the observations, also to compute the effect of the variations of the horary motion, noticing the higher order of differences, but these circumstances would affect the result but very little.

O Longitude D Longitude		14	21" 1		
Diff. long. D Latitude	03	26 24	23 28		9.22033 9.9998 7
Distance	80	26	33	co-sine	9.22020

The distances being calculated from noon The same as in the Nautical Almanac. and midnight by this (or the following) Problem, they may be interpolated for every 3 hours by Problem 1. An example will sufficiently illustrate this.

EXAMPLE.

Given the distances of the sun and moon in July, 1808, at 15d. 12h. 16d. 6h. 16d. 12h. and 17d. 0h.; respectively 85° 52' 13" | 80° 26' 33" | 75° 0' 44" | and 69° 34' 9". Required the distances July 16d. at 3h. 6h. and 9h.?

1808, July 15d. midnight	Dist. © D 1st 85°. 52'. 13	diff. ,, 2d. df.,,		
16 noon	80. 26. 33 -5.	25. 40 - 9	•	•
16 midnight				
17 доол	69. 34. 09 -5. At Sh.	26. S5 B=-27] At 6h.	At:	9 L .
Second longitude	+ 80. 26. 33	+80. 26. 33		+80. 26. 33
Prop. part A Table XLV. T 3h.	- 1. 21. 27.2 1. 2.6 T=6h.	A — 2. 42. 54.5 + 8.4	T=Sb.	- 4. 4. 21.7 十 26
Table Yr 4 . 1 2m	T . 20 1-04	T	1	
Distance at 3h.	79. 5. 8 Dist. at	6h. 77. 43. 42	Dist. at 9h.	76. 52. 14

These distances agree with the Nautical Almanac.

PROBLEM XVIII.

Given the longitudes and latitudes of the moon and a star, to find their distance. RULE.

To the log. secant of the difference of longitude of the moon and star, rejecting 10 in the index, add the log. tangent of the greater latitude, the sum will be the log. tangent of the arch A, of the same affection as the difference of longitude. Take the sum of the arch A, and the less latitude, if the latitudes are of a different name, but their difference if of the same name, and call it the arch B. Then add together the log. secant of the difference of longitude, the log. secant of the greater latitude, the log. co-sine of the arch A, and the log. secant of the arch B, the sum, rejecting 30 in the index, will be the log. secant of the distance of the moon and star of the same affection as B.

EXAMPLE.

Required the distance of the moon and the star a Pegasi at noon at Greenwich, July 16, 1808, when by the Nautical Almanac the moon's longitude was 33° 14' 1", latitude rected for aberration 19° 24' 41" N.?

long.	350	49						
Diff. long. Greater lat.		24 24			secant tang.		secant	10.15173 10.025 42
Arch A Lesser lat.	25 1	30 24	45 28	N.	tang.	9.67374	co-sine	9.93541
Arch B	24	6	17				secant	10.03962

Distance D * 45° 15' 23" secant 10.15221

It may be observed that the log, secant of the distance is also equal to the sum of the log. co-secant of the greater latitude, the log. sine of arch A, and the log. secant of the arch B, rejecting 20 in the sum of the indices; but the above rule is in general the most convenient on account of the smallness of the greater latitude, except when the difference of longitude is nearly equal to 90°.

PROBLEM XIX.

Given the right ascension and declination of a celestial object, with the mean obliquity of the ecliptic E, to find its longitude and latitude.

RULE.

To the log. tangent of the declination add the log. co-secant of the right ascension of the object, the sum, rejecting 10 in the index, will be the log. tangent of the arch A, to be taken out less than 90°, and called north or south as the declination is. If the right ascension is less than 180°, call the obliquity of the ecilptic south, if above 180°, merch. If A and E are of the same name, take their sum, otherwise their difference, which call B, and mark it with the same name as the greater number, whether N. or S. Then call B, and mark it with the same name as the greater number, whether N. or S. add together the log. secant of A, the log. co-sine of B, and the log. tangent of the right ascension, the sum, rejecting 20 in the index, will be the log tangent of the lon-tude in the same quadrant as the right ascension, unless B be greater than 90°, in

ich case the quantity found in the same quadrant as the right ascension, subtracted

1 3600 will be the longitude.

To the log, sine of the longitude add the log, tangent of B, the sum, rejecting 10 in the index, will be the log, tangent of the latitude of the same name as B.

Remark. As the Tables of this collection are not marked above 180°, you must subtract 180° from the right ascension when it exceeds that quantity, and find the log. tangent and log. co-secant of the remainder; and then the arch, corresponding to the log. tangent of the longitude, is to be taken of the same affection as this remainder. and 180°, added thereto, the sum will be the longitude, unless B is greater than 90°, in which case the supplement of that sum to 360° is to be taken as observed above.

EXAMPLE. By Table VIII. the right ascension of a Pegasi, July 16, 1808, was 22h. 55' 14"=

343° 48' 30", and its declination 14° 11' N. the mean obliquity of the ecliptic 23° 27' Required its longitude and latitude?

tang.

co-sine

tang.

Declin. tang. 9.40266 co-sec. 10.55462 343 48 30 R. A.

11 12 N. tang. 47 N.

65 38 59 N. B

Longitude 350° 49' 19"

10.15020 sec.

9.46295

tang. 10.54431 9.61522 9.20837

Lat. 19° 24' 51" N. tang. 9.54708

PROBLEM XX.

The longitude and latitude of a celestial object being given, with the mean obliquity of the ecliptic E, to find the right ascension and declination.

RULE.

To the log. tangent of the latitude add the log. co-secant of the longitude, the sum, rejecting 10 in the index, will be the log. tangent of the arch A, which is to be called north or south as the latitude is. If the longitude is less than 1800, call the obliquity E north; if above 1800, south. If A and E are of the same name, take their sum, otherwise their difference, which call B, marking it with the same name as the greater number. Then add together the log. secant of A, the log. co-sine of B, and the log. tangent of the longitude, the sum, rejecting 20 in the index, will be the log. tangent of the right ascension in the same quadrant as the longitude, unless B be greater than 90°, in which case the quantity found in the same quadrant as the longitude, subtracted from 360°, will be the right ascension.

To the log. sine of the right ascension add the log. tangent of B, the sum, rejecting

10 in the index, will be the log. tangent of the declination of the same name as B.

Remark. If the longitude exceeds 180° you must subtract 180° from it, and find the log. tangent and log. co-secant of the remainder. The arch corresponding to the log. tangent of the right ascension is to be taken of the same affection as this remainder, and 1800 added thereto will be the right ascension, unless B is greater than 900, in which. case the supplement of that sum to 360° is to be taken as was observed above. EXAMPLE.

By Table XXXVII. the mean longitude of a Pegasi, July 16, 1808, was 350° 49' 11", its latitude 190 24' 47" N. and the mean obliquity of the ecliptic 230 27' 47". quired its right ascension and declination?

Lat. 19° 24′ 47′ N tang. 9.547′

350 49 11 65 38 34 N. E 47 S. 27

Long.

B

9.51705 co-sec. 10.79712 tang. 10.34417 tang.

10.38466

9.20817

co-sine 9.86931

9.95713

tang.

Right ascension 343° 48' 28"

10 47 N.

9.44559

Declination 14° 10' 50" N. If the given longitude, latitude, and obliquity are the mean values, the resulting right ascension and declination will be the mean values, but if the proposed quantities are corrected for aberration and nutation, the resulting quantities will also be corrected. This

remark is equally applicable to the preceding Problem. SPHERIC TRIGONOMETRY.

Most of the rules given in the preceding Problems may be easily demonstrated by Spheric Trigonometry. As for example that of Problem XVII. may be investigated as follows. In Plate XII. Fig. 1, let A be the place of the moon, C that of the sun, CP an arch of the ecliptic, and AP a circle of latitude passing through the moon and cutting the ecliptic at right angles at P. Then the difference of longitude of the sun and moon is equal to the arch CP, and the moon's latitude is AP, whence the distance AC may be found by the rule of Napier, radius \times co-s. AC = co-s, $AP \times \text{co.s}$, CP. This in logarithms gives log. co-s. AC = log, eo-s. AP + log, co-s. CP - log, radius, which is the formula made use of. Want of room prevents the insertion of the demonstrations of the methods of calculating the other Problems.

The celebrated rules given by Lord Napier for solving the problems of Right-Angled Spheric Trigonometry being very easily remembered, are much made use of by mathe-In a paper communicated by the author of this work to the American Academy of Arts and Sciences, and published in the third volume of the memoirs of that society, a method was given for the more easy application of those rules to oblique Spheric Trigonometry, and as the tables of this collection may sometimes be made use of in solving various problems of Spherics besides those given in the former part of this work, is was thought proper to insert this improved method, with the formulas most frequently made use of, to enable any person acquainted with Spheric Trigonometry to make use of the tables, without the trouble of referring to another work, for the rules.

In every Right-angled Spheric triangle there are five circular parts; namely, the two legs, the complement of the hypotenuse, and the complements of the two oblique angles, which are named adjacent or opposite, according to their positions, with respect to each The right-angle is not included as one of the circular parts, neither is it supposed to separate the legs. In all cases of right-angled Spheric Trigonometry, two of these parts are given to find the third. If the three parts join, that which is in the middle is called the middle part; if they do not join, two of them must, and the other part which is separate, is called the middle part, and the other two opposite parts, as in Plate XII. fig. 1, 2. Then putting the radius equal to unity, the equations given by Napier will become

Sine of middle part = Rectangle of the tangents of the adjacent parts. = Rectangle of the co-sines of the opposite parts.

The method of applying these solutions to the various cases of Right-angled Sphetic Trigonometry is very simple, and is explained in several treatises. To apply the method to Oblique-angled Spheric Trigonometry, it is necessary to divide the triangle into two right-angled spheric triangles by means of a perpendicular AP (Plate XII. fig. 3, 4, 5, 14.) let fall from the point A upon the opposite side BC: the perpendicular being so chosen as to make two of the given things fall in one of the right-angled triangles, or In other words the perpendicular ought to be let fall from the end of a given side and opposite to a given angle.* Each triangle thus found, contains, as above, five circular parts, the perpendicular being counted and bearing the same name in each of them : consequently the parts of each triangle similarly situated with respect to the perpendicular, must have the same name. In every case of Oblique-angled Spheric Trigonometry, there are three parts given to find a fourth, and in making use of the method of a solution by means of the perpendicular, there will in general be two of these parts in each of the triangles ACP, ABP, similarly situated with respect to each other. To each of these must be joined the perpendicular AP, and there will be three parts in each triangle, which are to be named middle, adjacent or opposite, according to the above directions. Then the equations for solving all the cases of Right-angled, and all except two cases of Oblique-angled Spheric Trigonometry are,

 $\{=\}$ Tangents of the adjacent parts.† X = X = X Co-sines of the opposite parts. 1. Sine middle part

"These equations, when applied to right-angled spheric triangles, signify as before, that the sine of the middle part is equal to the rectangle of the tangents of the adjacent parts, or to the rectangle of the co-sines of the opposite parts : but when applied to an oblique-angled triangle, they signify, that the sines of the middle parts are proportional to the tangents of the adjacent parts; or that the sines of the middle parts are proportional to the co-sines of the opposite parts of the same triangle; observing that the perpendicular being common to both triangles APB, APC, and bearing the same name in each of them, must not be made use of in the analogies, nor counted as a middle part. This can produce no embarrassment, because the cases of Oblique Spheric. Trigonometry may in general be solved in the shortest manner without calculating the pe<mark>rpendicular.</mark>

The first case not included in the above rules, is where the question is between two sides and the opposite angles, which may be solved by the noted theorem, that the sines of the sides are proportional to the sines of the opposite angles, or as it may be expressed in an abridged form or more easy reference.

2. Sine side & sine opp. angle.

This, combined with the above improved formula, furnish a complete solution of the various cases of Spheric Trigonometry, except where three sides are given to find an angle, or (which is nearly the same thing, by taking the supplementary triangle) The above rules marked (1,) (2,) are simple in their three angles to find a side.

When this can be done in two different ways (as in Cases H. IV.) it will generally produce the shortest solution to make use of that perpendicular which does not divide the required angle or size.

[†] It will be of considerable assistance in remembering these rules to note that the second letters of the words tangent and co-sine are the same as the first letters of adjacent and opposite.

form, and the first varies but little from that made use of by Napier, so that it is extremely easy to remember them. The case not included in these rules may be solved by one of the formulas of case V. or VI. which may be committed to memory with little trouble. To illustrate these rules, the following examples are given, which include all the cases of Oblique Spheric Trigonometry.

CASE I. PLATE XII. Fig. 3, 4, 5, 14.

Given AB, AC, and the opposite angle C, to find BC and the angles A, B.

In the right-angled spheric triangle APC are given AC and C, and by marking it as in fig. 2, CP may be found by the rules $sine\ mid.=tang.\ adj.\$ which gives $sine\ (co.\ C)=$ tang. CP \times tang. (co. AC,) or tang. CP=co-s. C \times tang. AC.* Then in the triangles ABP, ACP are given AB, AC and CP to find BP. If to these is joined the perpendicular AP it will be found that in the triangle ACP the complement of AC is the middle part (as in Fig. 3), and CP an opposite part. The triangle ABP is to be marked in a similar manner. Then the rule $sine\ mid.\ \infty\ co-s.\ opp.$ gives sine (co. AC): co-s. CP:: sine (co. AB): co-s. BP, and BC=BP+CP. By marking the segments as in Fig. 4,

the rule sine mid. \(\infty\) tang. adj. gives sine CP: tang. (co. C.):: sine BP: tang. (co. B.)

Having found BC, the angle A may be found by the rule sine side. \(\infty\) sine opp. angle

which gives sine AB: sine C:: sine BC: sine A.

Otherwise—If the side BC is not required, the angles A, B, may be found in the following manner. The rule $sine\ mid.=tang.\ adj.$ gives by marking as in Fig. 1. sine (co. AC)=tang. (co. CAP) or cot. CAP.=co-s. AC×tang. C, and by marking as in Fig. 5, the rule (sine mid. \propto tang. adj. or) $tang.\ adj.$ \propto sine mid. gives tang. (co. AC): sine (co. CAP): tang. (co. AB): sine (co. BAP,) then A=BAP+CAP. By marking the segments as in Fig. 14, the rule (sine mid. \propto co-s.

opp. or) co-s. opp. ∝ sin. mid. gives co-s. (co. CAP): sine (co. C)::eo-s. (eo. BAP): sine (co. B) or sine CAP: co-s. C:: sine BAP: co-s. B. Having A, C, and AB, BC may be found by the rule sine side ∝ sine opp. angle, which gives sine C: sine AB:: sine A: sine BC.

CASE II. Fig. 3, 4. Plate XII.

Given AC, BC and the included angle C, to find AB, and the angles A, B.

The rule sine mid.=tang. adj. gives as in Case I. tang. CP=co-s. Cxtang. AC, then

BP=BC+CP and the rule co-s. opp. \(\precedex \) sine mid. gives by marking, as in Fig. 3. co-s.

CP: sine (co. AC):: co-s. BP: sine (co. AB,) and by marking as in Fig. 4, the rule sine mid. ∞ tang. adj. gives sine CP: tang. (co. C):: sine BP: tang. (co. B.) Having found AB we may find A, by the rule sine side ∞ sine opp. angle, which gives sine AB: sine C:: sine BC: sine A.

If the angle A had been required and not B, it would have been shorter to let the perpendicular fall upon the point B, by which means the required angle A would not be divided into segments. In this case the side AB and the angle A might be found in a similar manner to that by which AB and B are found above.

. CASE III. Fig. 3, 4, 5, 14. Plate XII.

Given the angles B, C, and the opposite side AC to find BC, AB, and the angle A. The rule sine mid. ∞ tang. adj. gives as in Case I. tang. CP=co-s. C \times tang. AC. Then the rule tang. adj. ∞ sine mid. gives, by marking as in Fig. 4, tang. (co. C.): sine CP:: tang. (co. B): sine BP, then BC=CP+BP. Again, the rule co-s. opp. ∞

sine mid. gives by marking as in Fig. 3, co-s. C P: sine (co. AC):: co-s. B P: sine (co. AB.) Having found BC, the rule sine side ∞ sine opp. angle, gives sine AC: sine B:: sine BC: sine A.

Otherwise—The rule sine mid. = tang. adj. gives as in Case I. cot. CAP=co-s. AC \times tang. C, and the rule sine mid. \propto co-s. opp. gives by marking as in Fig. 14, sine (co. C.): co-s. (co. CAP):: sine (co. B): co-s. (co. BAP) or co-s. C.: sine CAP.:: co-s. B: sine BAP, and A=CAP+BAP. Then the rule sine mid. \propto tang. adj. gives by

marking as in Fig. 5, sine (co. CAP): tang. (co. AC):: sine (co. BAP): tang. (co. AB.) Having found A the rule, sine side ∞ sine opp. angle gives sine B: sine AC:: sine A: sine BC.

^{*} In putting this or any similar expression in logarithms, the radius must be neglected in the sum of the two logarithms of the second number.



allo sine mid-stong, self, gives us in Cost I, ver, CAP-see-s, ACX (sec. BAP=A+CAP). The rule size with a long, etc. (long, etc. by marking so in

co. sho mids given by marking on in Pay, 14, 1500 (no. CAP) , thus (no. CAP) , thus (no. CAP) ; thus (no. CAP) is thus (no. CAP) and the Pay of the pay of the radio sine side on sine app, on the given size B : sine AC : the AC : the AC : the BC. with BC built been renoted and not AB, it would be shorter to be the perpendill from the point C, by which means the required side BC would not be divided meats. In this case the side BC and the angle B maghs be found in a similar to that by which AB and B are found above.

CASE V. Fig. 3.

Given AB; AC, and BC, to find either of the ougles made

-1 (AR+AC+RC,) then the angle A may be found by either of the following

ner & A = Si (S-An) + sinc (S-Ac) + - An + co - Ac - m.

Co-a. A A=

CASE VL Vig. 3.

Given the mights of, R. C. to find either of the rides on BC.

in [A+B+C.) Then the side Wil may be found by either of the f Towing as, adapted to logarithms as in the tast example.

Co-rise b + section (5-A) + messe, B + co-res, C-10,

Co-sine & BO-

show includes all the cases of Oblique Trigonometry. The Mr. and 40% cases solved in a different manar by the following theorems, which so some ac-

me A (AC+BC); sine A (BC or AC) : root, A C : tang. A (A-B), conian A (AC+BC) : coming A (BC or AC) : root, A C : tang. A (A+B), -B), is been than 90° and A (A+B) is of the same affection as A (AC+B) and an and difference of the tarms A (A-B) and A (A+B) will give A and R.

Both the sides in Case IV, may be found these

inc \S (A+C) : sinc \S (A = C = larg. \S AC : tags. \S (BC = AB). Coulon \S (A+C) : re-sinc \S (A \cong C) : strag. \S AC -tags. \S (BC+AB.) \S AB) is less than \cong \S and \S (BC+AB) is of the size effection \cong \S (A+C.) to saim and difference of \$ (BC or AB) and \$ (BC+AB) give AB and BC.

no aim and differences of a City of Albert Spheric Tragonometry by the accomproved rule for solving the cases of Oblique Spheric Tragonometry by the accompany be easily deduced from those given by Levi Napier. For if we pot M middle part, a far the adjacent part, and B for the opposite part of the transition, 3, 4, 5, 13, Plate MIL) =, a, b, for the corresponding part of the transition of P for the perpendicular AP. Then if P is an adjacent part, the raise of sine M sine of

will give they I'- and tare. Pthur, A.

y time M - Long. Accommon langue. If P is an opposite mult, the same rolls

APPENDIX TO THE SIXTH EDITION.

ON FINDING THE LATITUDE BY TWO ALTITUDES.

SINCE the part of this work for the finding the Latitude by two altitudes was in the press, the following Table XLVIII. has been computed, by means of which the correction of either one of the observed altitudes can be computed for the change of declination of the observed object during the elaped time between the observations, and thus the Problems of double altitudes of the sun, moon, planet, or fixed star, can be reduced to the case of the declination, being invariably the same as at the time of the observation of the altitudes which is not corrected, and then the Problem comes under the first (or second) method of solution, which is much more simple and free from cases than the general solution by the third method. This process of correcting the altitude is somewhat similar to that before taught, for making allowance for the run of a ship during the time elapsed between the observations; and the same altitude, which is corrected for the run of the ship, can also be corrected for the change of declination. This method of correcting one of the altitudes is particularly applicable to the case where both observations are made on the same heavenly body, and the declination does not vary but few minutes, or in extreme cases more than one or two degrees; but the same process may be used when two different objects are observed, provided their declinations are nearly equal, or do not differ more than one or two degrees.

As either one of the altitudes may be corrected, the Problem admits of two different ways of solution. For the sake of precision, the altitude which is selected to be corrected, will be called the first altitude; and the corresponding declination, the first declination; the other altitude, which is not corrected, will be called the second altitude, and the corresponding declination, the second declination. These terms, first and second, having no reference to the order in which these observations are taken, since the altitude here defined as the first attitude, may be actually observed either before or after the other obser-

vation

The proposed table gives for various declinations, altitudes, and latitudes, the change of the first altitude, corresponding to a variation of 100" in the first declination. Thus, with the latitude 50° N. the sun's altitude 30°, and the declination 14° N. the Table gives 77" for the variation of that altitude arising from a change of 100" in the declination. If the actual change of declination is greater, or less than 100" the tabular number 77" must be increased or decreased in the same proportion. Thus, if the change of declination be 200", the change of altitude will be $200^{\circ} \times 7.7_{0.5} = 154$ ". If the change of declination be 60", the change of altitude will be $60^{\circ} \times 7.7_{0.5} = 46$. The correction of this first altitude having been found, it is to be applied to the first alfitude, corrected as usual, for dip, refraction, semi-diameter and parallax, and the corrected first altitude will be obtained, such as it would have been, if the declination at the time of observing that altitude had been equal to the second declination. With this corrected first altitude, the second altitude and second declination without correction, and the observed elapsed time, or hour angle, the computation of the latitude may be made by the First Method, explained in page 133

This Table is calculated for every 2° of declination, from 0° to 26°. If the change of declination is not very great during the elapsed time, it will in general be sufficiently exact to enter the table with the nearest declination, and take proportional parts for the degrees of altitude and latitude. The latitude by account is to be used in finding the numbers from this table, it being sufficiently accurate, since an error of 1° of latitude rarely produces more than 2' change in the numbers of the Table. Suppose now, that the tabular number was required, when the latitude was 37° N. the first altitude 28°, the tfirst declination 6° 25' S. In this case, using the declination 6°, and the altitude 20°, he tabular numbers corresponding to the latitude 30° S. and 40° S. are, respectively, 57" and 73", whose difference 16" corresponds to a change of 16° of latitude, and by T T Tab.

proportion, the change corresponding to 7° of latitude is $16'' \times 7 = 11''.2$, this added to 57", gives the correction corresponding to the altitude 20° and the latitude 37° S. to 57', gives the correction corresponding to the united 30', and 30', the two tabular numbers are 64'' and 81'', whose difference 17" multiplied by $\frac{7}{10}$ gives 11".9 to be added to 64" to get 75".9, the correction corresponding to the altitude 30° and the latitude 37° S. Hence it appears by changing the altitude from 20° to 30°, the correction changes from 68".2 to 75°.9, increasing 7".7, by an increase of 10° in the altitude, the corresponding increase for a change of 8° in the altitude is equal to 7'.7 \times 10 = 6'.2 nearly. This added to 68".2 gives 74".4, for the tabular number corresponding to the declination 6°, the altitude 28°, and the latitude 37° S. If the same calculation be repeated, using the declination 8°, the tabular number will be 76".2 instead of 74".4, increasing only 1".8 for an increase of 2°=120' in the declination, and the corresponding correction for the 25' of the first declination is $1'.9 \times \frac{1}{120} = 0'.4$, nearly. This added to 76''.2 gives the correct tabular number 76''.6, or 77'' nearly, corresponding to the proposed latitude, 37° S. altitude 28°, or declination 6° 25' S. The correction for the minute of declination is in this case small, and in general it will be so, and when the change of declination during the clapsed time is only a few minutes, it will be sufficiently e mact to take out, as was directed above, the numbers corresponding to the nearest de-clination in the table. As there is nothing peculiar in this method of finding the cor-rections for the intermediate degrees of altitude and latitude (several tables in the work having been arranged upon a somewhat similar plan) it will not be necessary to go into any further detail relative to the manner of finding the number from the table corresponding to any proposed declination, altitude or latitude. The use of these numbers in finding the correction of the first altitude, is, for the sake of easy reference, drawn up in the following rules.

RULE.

1. If the two declinations are of the same name, take their difference; if they are of different names, take their sum, and this difference, or sum, will be the change of declination corresponding to the two observations, or two objects.

2. Find in Table XLVIII. the number corresponding to the first declination, the first altitude, and the latitude by account. Multiply this by the change of declination, in seconds, between the two observations; the product, rejecting the two right hand figures, will be the number of seconds to be applied to the first altitude, with the same sign as in the table, * if at the second observation, the object is nearer to the elevated pole than at the first observation; but with a different sign from the Table, if at the second observation, the object is further from the elevated pole than at the first observation.

Thus, in the above example, where the tabular correction was 77'', if the second altitude was 48° and the second declination 6° 15' S. which is 10' or 600'' less than the first declination 6° 25' S. the product of 600'' by 77 (rejecting the two right hand figures) is 462'' = 7' 42'', being the correction to be added to the first altitude 28° , making it 28° 7' 42'', because the second declination is nearest to the elevated pole. If the second declination had been 6° 35' S. instead of 6° 15' S. the correction 7' 42'' would be subtractive, making it 27° 52' 18''.

It may be observed, that the method of correcting one of the altitudes does not alter the horary angles in any way whatever, and the regulation of the watch used in the observation is calculated in exactly the same manner as if the correction had not been made, and whichever altitude is corrected, the result will be very nearly the same; a difference of a few seconds will sometimes be found, owing to the small quantities neglected.

To illustrate this, the following examples are given.

EXAMPLE I.

The sun's correct central altitude was 32° 25', his declination 17° N. Eight hours afterwards, by a watch, his correct central altitude was 30° 8' and declination 16° 55' N. Required the latitude, supposing the latitude by account 53° 20' N?

The tabular correction corresponding to the first altitude 32° 25°, declination 17° N. and latitude by account 53° 20′ N. is 80″. Multiplying this by the difference of the de-

^a The signs in the Table are positive except in a few places between the tropics. In all cases without the tropics, when the distance from the elevated pole decreases, the altitude is to be increased, and when the polar distance increases, the altitude is to be decreased. The contrary takes place in these latitudes between the tropics where the tabular numbers have the sign—prefixed. It may also be observed, that the tabular number, corresponding to any possible situation of the object, cannot exceed 500°; it was however fund convenient to insert a few numbers exceeding 100°; for the purpose of finding more accurately the proportional parts for the intermediate degrees of altitude or latitude corresponding to profibble cases.

clination $17^{\circ}-16^{\circ}$ 55'=5'=300', the product (rejecting the two right hand figures) is 240'. 00=4', the correction of altitude. This is to be subtracted from 32° 25' because the sun recodes from the elevated pole, while the declination changes from 17° N. to 16° 55' N. therefore the corrected first altitude is 32° 21'. Using this with the second altitude 30° 5' the second declination 16° 55', and the elapsed time 8 hours, the calculation may be thus made by the first method, as follows—

Elapsed time		Cor		10.082471	Cor	2.				Co	L.	3.			
Declination	18	65 N.	sec.	10.01921			۱.						co-sec.	10.53614	ı
A			co-sec.	10.08168	co-sine	9.74812	1						co-sine	9.74812	ż
3 Sum Alt's.	31	141	co-sine	9.93196	00-sec.	10.28512	В	310	18	ď N.			co-sec.	10.28426	3
3 Diff. Alt's.	1	6}	sine	8.28650	soc.	10.00008			[B	less	tha	n 90	o named	as decl.]	i
C.	1	9	sine	8.30014	co-sine	9.99991							∞-sine	9.99991	i
[Z less than 9	(O)	amed as be	earing of	renith.]	Z sec.	10.03325	z	22	8	N.					
							E	53	26	N.			sine	9.90490)
		•				Latit	ude	53	25	N.			sine	9,90471	i

As it is entirely arbitrary which altitude is considered as the first, or the one to be corrected, it may not be amiss to repeat the operation, considering 30° 8' as the first altitude, and 16° 55' as the first declination. The tabular number corresponding to these quantities, and the latitude by account is 79'' which multiplied by the change of declination 300'' (rejecting the two right hand figures) is 237'' = 3' 57'' or 4' nearly. This is to be added to 30° 8' to give the corrected first altitude 30° 12', because the Sun approaches the elevated pole, while his declination changes from 16° 55' to 17° . Assuming, therefore, the corrected first altitude as 30° 12', the second altitude 32° 25', the second declination corresponding thereto 17° N. and the elapsed time, as before, 8 hours, the calculation may be then made as follows—

Thomas d simu	•••	Cor		10.06247	Cor	 2.				Cor.	3.		
Elapsed time Declination	170	N.		10.01940		٠.						co-sec.	10.53406
A			co-sec.	10.08187	co-sine	9.74350						co-sine	9.748 <i>5</i> 0
🛔 Sum Alt's.	810	184	co-sine	8.93165	co-sec.	10.28429	В	810	27'	N.		00-sec.	10.28256
d Diff. Alt's.	1	6 <u>}</u>	sine	8.28650	sec.	10.00008		4	[B l	ess th	an 90°	named	as decl.]
0	1	9	sine	8.30002	co-sine	9.99991						co-sine	9.99991
[Z less than i	100 III	ke bearing	of zeni	th.]	sec.	10.03278	Z	21	5 9	N.			
						Latitude	E	53	26	N.		sine	9.90480
								55	25	N.		sine	9.90471

So that the latitude is exactly the same by both methods.

If the middle time between the two observations was required, it would be obtained by adding the log. tangent of C 8.30263 to the log. secant of E 10.22493, whose sum, rejecting 10 in the index, is 8.52762 which sought for in the log. tangents correspond in the Col. P. M. to Oh. 15m. 26s. whose half Oh. 7m. 43s. is the middle time between the two observations. Taking the sum and difference of this and half the elapsed time, 4h, gives the times from noon when the observations were made, 4h. 7m. 43s. and 3h. 52m. 17s. the one being before noon, the other afternoon. The same result is obtained which ever altitude is corrected.

EXAMPLE II. [Same as Example XIII. page 145.]

Given the moon's correct central altitude 55° 20', the moon's declination 0° 36' N. The sun's correct central altitude at the same time 37° 40', his declination 0° 17' S. The hour angle, or difference of the right ascensions of the sun and moon 5 hours. Required the true latitude, the latitude by account being 23° 20' N.

The tabular correction corresponding to the latitude by account 23° 20' N. the sun's altitude 37° 40', considered as the first altitude, and declination 0° 17' S. is 50", and the change of the two declinations from 0° 17' S. to 0° 36' N. is (53' =) 3180', this multiplied by 50, and the two right hand figures rejected, gives the correction of altitude 1590"=26' 30", this is to be added to the altitude 37° 40' because the change from 0° 17'S. to 0° 36' N. approaches the sun to the elevated pole, therefore the sun's corrected

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altitude is 38° 6' 30" or simply 38° 6'. Using this with the moon's altitude 55° 20', the moon's declination 0° 36' N. and the hour angle 5 hours, the latitude may be found by the first method, in the following manner:

Col. 1.		Co	L. 2.			Col. 2	3.	
Etapsed time 5h. co- Declination 0° 36' N. sec.	ec. 10.21855 10.00002						. CO-SEC.	11.97998
A co-6	ec. 10.21557	co-sine	9.89947				co- sine	9.29947
1 Sum Alt's. 46 45 co-	ine 9.83608	co-sec.	10.15789	B	00 4	54 N.	co-sec.	11.87947
½ Diff. Alt's. 8 37 sine	. 9.17658	sec.	10.00495					
` O sine	9.22723	co-sine	9.99372				. co-sine	9.39372
	:	Z sec.	10.03601	Z	23 (ų n.		
				E	23 4	6 N-	sine	9.60532
			Latitu	de	230 2	4' N.	sine	9,59804

This agrees with the calculation by the third method.

If the moon's altitude 55° 20' had been considered as the first altitude and corrected. the tabular number corresponding to this altitude, the moon's declination 0° 36' N. and the latitude by account 23° 20' N. will be 70". Multiplying this by the change of declination 3180", and neglecting the two right hand figures, gives the correction of affitude 2226 = 37' 6" or simply 37', which is to be subtracted from the moon's altitude 55° 20' to obtain the corrected altitude 54° 43′, because the change from 0° 36′ N. to 0° 17′8, makes the moon recede from the elevated pole. Using the corrected altitude 54° 43′ the sun's declination 0° 17′ S. and the sun's altitude 37° 40′ with the hour angle 5h, the latitude may be found by the first method, in the following manner :

Col. 1. Sh.		10.21555	Cor	. 2.	Col. 3.					
Declination 00 17' S.	sec.	10.00001							co-sec.	12.30585
`_ A	co-sec.	10.21856	co-sine	9.89947					co-sine	9.89947
3 Sum Alt's, 46 11 3*	co-sine	9.84028	cn-sec.	10.14167	В	6 0	21'	j 8.	CO-60C.	12.20639
1 Diff. Alt's. 8 311	sine	9.17097	sec.	10.00482						
. σ	sine	9.22679	co-sine	9.99374			·		co-sine	9.98374
4			Z sec.	10.05970	z	24	7 }	N.	`	
-					F	23	46	N.	sine	9.60532
				Latitude	•	239	24'	N.	sine	9.59966

Which agrees with the preceding calculations.

EXAMPLES FOR EXERCISE.

1. The sun's correct central altitude was 41° 33′ 12′, his declination 14° N. After an interval of 1h. 30m. his correct central altitude was 50° 1' 12" and declination 13° 58' 38". Latitude by account 52° 5' N. Required the true latitude ?

The tabular number corresponding to the altitude 41° 33′ 12″ is 87″ and this being taken for the first altitude, is also corrected 41° 32′0″, the second altitude 50° 1′ 12″, elapsed time 1h. 30m. and declination 13° 58′ 38″ N. These make the latitude 52° 5″ N. Or, by taking 50° 1′ 12″ for the first altitude, and using the corresponding declination, the tabular number is 95′, the corrected first altitude becomes 50° 2′ 30″, using this with the second altitude 41° 33′ 12″ the declination 14° N. and the elapsed time 1h. 30. The latitude becomes as before 50° 5′ N. latitude becomes as before 52° 5' N.

2. Given the correct central altitude of the moon 53° 43', her declination 14° 16' N. After an interval in which the hour angle was 1h. 44m. 15s. her correct central altiande was 42° 29' and declination 13° 52' N. The latitude by account 48° 54' N. Required the true latitude?

With the first altitude and first declination the tabular number is 98", and the corrected first altitude 53° 19' 28", the second altitude 42° 29' with which and the declination 13° 5%' N. and the corrected elapsed time or hour angle 1h. 44m. 15s. the latitude will be found 48° 55' N.

^{* 4} in taking the half sum and half difference of the altitudes, it will be convenient to prove the accuracy of the calculation by adding this half sum to the half difference, for the sum will be the greater altitude.

The difference of the same numbers will be the least altitude. Thus in the present example 460 11' \(\frac{1}{2} + 8^\ \) 31' \(\frac{1}{2} - 340' \) 43' the greater altitude, and 460 11' \(\frac{1}{2} - 8^\ \) 31' \(\frac{1}{2} - 370' \) 40' the least altitude.

Or, by taking 42° 29' for the first altitude, and 13° 52' N. for the first declination, the tabular correction will be 83", the corrected first altitude 42° 49', using this and the second altitude 53° 43', the corresponding second declination 14° 16' N. and the hour angle 1h. 44m. 15s. the latitude will be found 48° 54' N. nearly agreeing with the former calculation.

3. Given the correct central altitude of the moon 55° 38', her declination 0° 20' S. After an interval in which the hour angle was 5h. 30m. 49s. her correct central altitude was 99° 57', and her declination 1° 10' N. The latitude by account 23° 25' S. Re-

quired the true latitude?

With the first altitude 55° 38' and the first declination 0° 20' S. the tabular correction is 71" and the first corrected altitude 54° 34' 6". Using this with the second altitude 29° 57', the second declination 1° 10' N. and the hour angle 5h. 30m. 49s. the true

atitude will be found \$30 23' 8.

Or, by taking 29° 57' for the first altitude, and 1° 10' N. for the first declination, the tabular correction will be 45' and the first corrected altitude 30° 37'. Using this with the second altitude 55° 35' the second declination 0° 20' S. and the hour angle 5h. 30m. 49s. the true latitude will be found to be 23° 24' S. nearly agreeing with the preceding calculations.

In making the calculations of these three examples the seconds were noticed, which is always best to be done, particularly when the altitudes are nearly equal; some difference might be found in the above results if the nearest minutes only were taken. Thus, Example XII. page 144, calculating to the nearest minute, only gives the latitude 55° 28′. If the calculation be made as in Ex. I. of this appendix, it becomes 53° 25′, differing 3′. This would be avoided by taking the angles to seconds, and in some extreme case it would require the use of 6 or 7 places of decimals.

TABLE XLVIII.

ewing the variation of the altitude of an object arising from a change of 100 seconds in the declination: moves the body towards the elevated pole, apply the correction to the altitude with the signs in the rwise, change the signs.

$\overline{\Gamma}$	1		ation.	E declin	TUD	LAT!	fores	 L	LA CITUDE Of same name as declination.										
Dec	Ail	70~	60°	50 0	40°	30 °	\$0 °	100	00	100	50 0	300		50°	,	·0°			
o	0° 10 20 30 40 50 66 70	94 93 100	87″ 88 92 100	76" 78 82 88 100	64'/ 65 68 74 84 100	50" 51 53 57 66 78	34" 35 36 39 45 53 68 100		0" 0" 0" 0" 0"	17" 18 18 20 22 27 35 51	34" 35 36 39 45 53 68 100	50" 51 53 57 65 78 100	64" 65 68 74 84 100	76" 78 92 98 100	87" 88 92 100	94" 95 00			
20	0 10 30 30 40 50 60 70	94 96 101	87 88 93 102	77 78 93 90 103	64 66 69 76 86 103	50 51 54 59 68 81 105	34 35 37 41 47 57 73 108	17 18 19 22 25 30 40 59	0 -1 -1 -2 -2 -3 -5 -8	17 17 17 18 20 24 30 43	34 34 35 38 42 50 64 92	50 50 52 56 63 74 95 139	64 65 67 73 82 97 124	77 77 81 87 98 116	87 87 91 98 111	94 95 99 07			
40	0 10 90 30 40 50 60	94 97 103	87 89 95 104	77 79 94 93 106	64 67 71 78 90 109	50 52 56 68 71 86 113	34 36 39 44 51 62 81 119	17 19 21 24 29 35 47 70	0 -1 -3 -4 -6 -8 -19	17 16 16 16 17 19 23 32	34 34 34 36 39 45 56 91	50 50 51 54 59 70 88 127	64 64 66 70 78 92 117	77 77 79 85 94 111	87 97 90 96 107	94 94 98 105			
60	0 10 20 30 40 50 60	94 98 104	87 90 96 107	77 80 86 95 109	65 67 73 81 93 113	50 53 57 64 74 91 119	34 37 40 46 54 66 87 129	17 20 22 26 32 40 53 80	-0 -2 -4 -6 -9 -13 -18 -29	17 16 15 14 14 15 17 22	34 33 33 34 36 41 51 72	50 49 50 52 57 66 82 118	65 64 65 69 76 88 111	77 76 78 93 92 107	87 87 89 94 105	94 94 97 103			
80	0 10 20 30 40 50 60	95 99 106	87 91 98 109	77 81 87 97 113	65 68 74 83 97 118	50 54 59 66 78 95 125	35 38 40 48 57 70 93 140	18 20 24, 29 35 44 59 90	-0 -3 -5 -8 -12 -17 -24 -39	18 15 14 19 11 11 11	35 33 32 32 33 37 45 62	50 49 49 50 54 62 77 109	65 63 64 67 73 84 105	77 76 77 81 89 104	87 86 88 93 102	95 94 96 101			
100	0 10 20 30 40 50 60 70	95 100	88 92 100	78 82 89 100	65 69 76 86 100	51 55 60 69 81 100	35 38 43 50 60 75 100	18 91 95 30 38 48 66 100	-0 -3 -6 -10 -15 -21 -31 -48	18 15 12 10 8 6 5	35 32 31 30 31 33 39 53	51 48 48 49 51 58 71 100	65 63 63 65 70 81 100	78 75 76 80 87 100	88 86 87 91 100	95 94 95 100			
190	0 10 20 30 40 50 60 70	96 101	89 94 109	78 83 91 103	66 70 78 88 104	51 56 69 71 85 105	35 39 45 53 63 80 107	18 29 27 33 41 53 72 110	-0 -4 -8 -12 -18 -25 -37 -58	18 14 11 8 5 2 -1 -6	35 32 29 28 28 29 33 44	51 48 47 47 49 54 65 91	66 63 62 64 68 77 95 134	78 76 76 78 84 97 120	89 86 86 90 98 112	96 94 94 99 1 0 8			
Dec gl	ui. OC	by G	gitized	50°	TITUL	LAT			0 2	100	20 °	30°		500	60°	70°			

TABLE showing the variation of the altitude of an object arising from a change of 100 seconds in the change moves the body towards the elevated pole, apply the correction to the altitude with the si otherwise, change the signs.

$\lceil \cdot \rceil$		1	Of	Patrone 1	LATI	TUDE	ation.		1	Of d	lifferen	LAT	TTUD	E declin	ation.
Dec	Alt.	70	600	50°	400	30°	20°	10°	00	100	200	30°			60°
14	10 20 30 40 50 60	97' 94 94 97 106	89" 86 86 89 96 109	79' 76 75 77 92 93 115	66" 63 61 62 66 73 99 125	52' 48 46 45 46 50 60 82	35" 31 27 26 25 25 27 35	18" 14 10 6 2 -2 -7 -16	0 4 9 14 91 30 43 69	23 28 35 44 59	35' 40 45 55 67 85 114	52° 57 64 74 89 110	66' 72 80 91 107	79' 85 93 106	89″ 95 104
169	0 10 20 30 40 50 60	98 94 94 96 104	90 86 85 87 94 106	80 76 74 75 80 90 110	67 63 61 61 63 79 84 117	58 48 45 44 44 47 54 73	36 31 27 25 22 21 21 25	18 13 9 4 0 6 14 26	-0 -5 -10 -17 -24 -34 -50 -79	23 30 37 48 62 86	36 41 48 58 70 90 121	52 58 66 77 99 115	67 73 82 94 111	90 96 95 109	90 97 106
180	0 10 20 30 40 50 60 70	99 95 93 95 102	91 87 85 86 92 103	81 76 74 74 78 97 105	68 63 60 59 61 66 79 108	53 48 44 49 41 43 49 64	36 31 26 23 20 17 16 16	18 13 8 2 -3 -10 -20 -36	-0 -6 -12 -19 -27 -39 -56 -89	18 24 31 40 51 67 93	36 42 50 60 74 95 128	53 59 68 79 96 121	68 74 94 97 116	81 88 98 112	91 98 109
2 0°	0 10 20 30 40 50 60	100 95 93 94 100	92 87 85 85 90 100	82 76 74 73 76 83 100	68 63 60 58 59 63 74 100	53 48 43 40 39 39 43 56	36 31 25 21 17 13 10	18 12 6 0 6 15 26 46	-0 -6 -13 -21 -31 -43	18 95 33 42 55 79 100	36 43 52 63 78 100	53 60 70 82 100	68 76 86 100	89 100	92 100
22 °	0 10 20 30 40 50 60	96 93 94 98 110	93 88 85 85 88 97 117	83 77 73 72 74 80 95	69 63 59 57 57 60 68 92	54 48 43 39 36 36 38 47	37 30 25 19 14 9 4	19 12 5 -2 -9 -19 -33 -56	-0 -7 -15 -93 -34 -48 -70	19 26 35 45 58 77 107	37 45 54 66 82 106	54 62 72 86 104	69 78 88 103	93 91 103-	93 102
94 0	0 10 20 30 40 50 60 70	97 93 93 97 107	95 88 85 84 86 93 112	84 77 73 71 72 77 91 123	70 64 59 56 54 56 64 83	56 48 42 38 34 32 32 38	37 30 94 18 18 5 -2 -13	19 11 4 -4 -12 -23 -39 -67	0 8 16 26 37 53	19 27 36 48 62 83 115	37 46 56 69 86 111	55 63 74 89 109	70 79 91 107	84 93 105	95 104
26 c	0 10 20 30 40 50 60 70	98 95 93 96 105	96 89 85 83 85 92 108	85 78 73 70 70 74 86 115	79 64 59 54 52 53 58 75	56 48 41 36 32 28 27 29	38 30 23 16 9 1 —8 —23	19 11 3 -6 -16 -28 -46 -78	19 18 18 28 41 -58	19 28 38 50 66 88 123	38 47 58 79 91 117	56 65 77 92 114	72 81 94 111	85 95 108	96 106
Dec.	A15.	700	60°	50°	40°	30°	20°	10°	00	10 ⁰	20 0 itized	30°	40°	50°	60°
		l 				PUDE as dec	ination	ı.		0	ુ યું કુર		TTUD		in aria

New Form of Table XX. for working a Lumar Observation.

SINCE the former part of this work was printed, a new form has been given to Table XX. by which the last corrections of a Lunar Observation may be found very expeditionally. The correction in this new Table will frequently exceed those deduced from the former one by two or three seconds, on account of having introduced the fourth and fifth corrections, which were formerly omitted.

In using this new Table, it will in general be sufficiently accurate to find the nearest altitudes and distance, and take out the corresponding correction, without the trouble of making a proportion for the neglected degrees and minutes. Thus in Example I. enter the new table with the X's altitude 50°, p's altitude 70°, distance 50°, which are the nearest numbers in the Table, the corresponding correction is 20°, as found before. In the second example the X's altitude 10°, p's 20°, distance 10°, give the correction 27°, being 4° more than the former method. In the third example the correction is 24°. In the fourth 22°. In the fifth 17°. In the sixth 19°. These corrections are always additive.

In using the second and third methods of working a lunar observation, this correction is to be found in the same manner, and added to the corrected distance, subtracting 18° from the sum. Thus, in Example L the correction 20° is to be added, and the constant quantity 18° subtracted, leaving 3° to be added instead of 1°, found by the former method.

TABLE VY / Aran E.

TABLE XX. (New Form.) Correction in Seconds, additive. True distance of the Moon from the Sun or a Star.

				l'rue	dista	ance	of th	e Mo	on fr	om t	he S	un or	a St	ar.			
ف¥ ۱۱۱۸	B Alt.	-0°	25°	30	50ن	40°	454	300	600	700	800	900	100c	1100	120°) Alt	
<u>(Jo</u>	10 ³ 20 30 40 50 60 70	96" 96 18	79' 71 39	67 64 46 18	58 46 28	53" 52 44 32 18	48' 48 42 53 23	41" 43 39 38 25 18	37" 37 34 30 26 22 18	31' 31 30 27 24 22 19 18	27" 27 15 24 22 21 19 18	28" 22 22 21 20 19 18 18	19" 18 18 18 18 18	15' 15 15 16 17 18	11" 11 13 15 18	10 ² 20 30 40 50 60 70	10°
<u>300</u>	10° 20 30 40 50 60 70	60 89 67 18	64 73 62 35	56 63 56 39 18	51 56 50 39 25	46 49 45 38 28 18	42 44 41 56 29 21	39 40 38 39 28 23 18	83 35 32 29 26 23 20 18	28 28 27 25 23 21 20 18	24 24 23 22 21 20 19	20 20 20 19 19 18 18	17 17 17 17 17 18 18	18 14 15 16 18	10 11 14 18	10° 20 30 40 50 60 70	200
Ю	10° 20 80 40 50 60 70	18 68 78 57 18	35 60 64 53 31	40 53 55 48 34 18	40 49 48 43 34 23	38 44 43 39 32 25 18	36 40 39 38 30 25 20	34 36 38 33 29 24 21 18	30 31 30 28 25 22 20 19	26 26 25 24 22 20 19	22 22 21 20 19 18	19 19 19 19 18 18	16 16 17 17 18	14 15 16 18	12 14 18	10° 20 30 40 50 60 70	900
i0c	10° 20 30 40 80 60 70	18 58 64 47 18	34 52 54 43 27	18 38 47 46 39 28 18	26 58 42 41 35 28 21	28 96 96 37 32 27 22 18	25) 34 35 33 30 26 21 19	28 32 32 30 27 24 21 19	27 28 27 26 24 21 19	24 24 24 22 21 19 18	21 21 21 21 20 19 18	19 19 18 18 18	17 17 17 18	15 16, 18	15 18	10° 20 30 40 80 60 70	40°
go.	10° 20 39 40 50 60 70	18 48 50 36 18	39 43 42 54 23	18 33 39 37 31 23 18	25 33 35 35 33 28 23 19	18 27 32 32 30 28 24 19	2½ 27 30 30 27 24 21 19	28 27 28 27 25 22 20 18	23 25 25 24 22 20 18	22 22 22 21 19 18	20 20 20 19 18	18 18 18 18	17 18 18	17 18	18	10° 20 80 46 80 60 70	500
00	10° 20 30 40 50 60 70	18 38 36 27 18	26 33 31 25 20	18 28 31 28 23 19	23 28 28 25 25 22 19	18 1 24 27 26 23 20 18	21 24 26 24 22 19	18 22 24 24 24 22 20 18	20 22 22 21 20 18	20 21 20 19 18	19 19 19 18	18 18 18	18 18	18		10° 30 30 40 50 60 70	80°
00	10° 20 30 40 40 60 70	18 27 25 20	25 25 23 19	18 23 23 21 18	21 26 22 19	18 22 22 22 20 18	20 21 21 10	18 20 21 20 18	18 90 20 19 18	19 19 19 18	19 18 18	18	18			10° 20 30 40 50 60 70	900
800	10 ^a 20 30 40 60 70 80	18 20 18	20	18 19 18	19 19	18 19 18	19 19	18 19 18	18 19 18	18 18 18	18	18	Digiliz	ed b	Gb	10° 8 90 50 40 50 80 70 0	e

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- DO. of the West-Indies, on four sheets, which may be had separate.
- DO. of the Coast of Guyana.
- DO. of the Coast of Brazil.
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Musical Instrument Makers, and Importers of ALL KINDS OF MUSICAL MERCHANDISE,

No. 358 Pearl-street, near Franklin-square, NEW-YORK,

Have constantly for sale an assortment of Piano Fortes, and Musical Instruments of the most celebrated Makers and of their own manufacture; together with every other asticle of Musical Merchandise, wholesale and retail, on the most liberal terms. BANDS supplied with Instruments, &c. Musical Instruments of all kinds tuned and repaired in the neatest manner. All orders thankfully received and attended to with care and despatch.

April, 1826.

N. KNIGHT.

BOOK, STATIONARY, AND VARIETY STORE,

Thames-street, late Fell-street, Fell's Point, BALTIMORE.

April, 1826.

CABINET WAREHOUSE.

BOSS & FAIRCHILD

Beg leave to inform their friends and the public that they have commenced the CABI-NET MAKING business at their Warehouse, No. 607 Broadway, where all orders will be thankfully received and punctually attended to: they will also find a large assortment of imitations of wood of all descriptions, painted by one of the first painters in the country. April, 1826.



A GENERAL ASSORTMENT OF

TACKLE. SPORTING

WHOLESALE AND RETAIL.

AT ISRAEL HORSFIELD'S HARDWARE STORE.

No. 230 Water-street, New-York.

Fishing Tackle.—Fish rods, swivels, shrimp nets, India grass, hair, gut and flax lines. brass and wood reals, floats, sinkers, gentlemen's fish cars, silkworm gut, single and twisted, Kirby and taper point fish-hooks fastened on gut, grass, hair, and flax snells. Patent spring pickerel, R. Hemming & Son's royal improved cast, and best steel Kirby, taper point mussel, mackerel, whiting, cod, and, sea fish-hooks, stamped I. P. Shooting Tackle.—Dartford canister, English and American gunpowder, patent shot.

shot bags, powder flasks, horns, game bags, gun worms, locks, &c. &c.

ALSO, A GENERAL ASSORTMENT OF HARDWARE.

May, 1826.

FLOUR STORE.

The subscribers have constantly on hand and for sale at No. 223 Front-street New-York, a large and general assortment of New-York Canal, I hiladelphia, Baltimore, and Richmond Superfine and Fine FLOUR-Also, Rye, Indian. and Buckwheat MEAL, which they sell at the lowest market prices, and warrant all l'lour sold by them to be equal to the representation.

Merchants, Masters of vessels, and private families, will always be supplied with a NEWELL & PARSONS, superior article.

New-York, April, 1826.

No. 223 Front-st. near Peck-slip.

SELDEN BR YNARD'S

Lottery and Exchange Office, No. 16 State-street, Boston.

FIVE DOORS ABOVE THE BRANCH BANK,

Where have recently been sold the following Capital Prizes, viz :-

1	Prize of	\$ 50,000	97	Prizes of	\$10,000
2	do.	15,000		do.	500
4	do.	10,000	218	do.	100
6	do.	5.000	318	do.	50

The Capital Prize of \$50,000 is the highest ever sold in Boston; and the cash was advanced on presentment of the ticket at the above office, on the day the drawing was received. Tickets may be had in all the Lotterics drawn in the United States, and the cash paid for prizes soon as drawn.

Orders enclosing cash (post paid) will meet with immediate attention.

May, 1826.

CARPETING.

J. & J. H. SACKETT.

Offer for sale at their Store, 96 Division-st., nearly opposite Market-st.

An extensive assortment of VENETIAN, BRUSSELS, and ENGLISH INGRAIN-ED CARPETING, large and elegant patterns, of all qualities; Nankin and Canton Matting; Brussels, Wilton and Imperial Rugs; able and Piano Covers, &c. &c.

Merchants and Ship-Masters, in furnishing ship's cabins, will find it greatly for their

interest to apply as above.

N. B. Also—a splendid assortment of light Carpeting, patterns and quality equal to any ever imported, at prices which will be an object to those who are in want of the above articles.

New-York, April, 1836.

YORK HOUSE,

NOS. 5 AND 7, COURTLANDT-STREET,

NEW-YORK,

BY A. YOUNG.

May, 1826.

GILBERT & SONS,

EXCHANGE-STREET, BOSTON.

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Bought and Sold.

Notes of Hand, &c., Bank Notes, J
LOTTERY TICKETS in all legally authorized Lotteries, constantly for sale.—Attend
generally to all business relating to Stock Exchange, Money and Lottery Brokers.
April. 1826.

J. HORSPOOL, CABINET, CHAIR, AND SOFA MANUFACTURER.

16 WHITE-ST.

J. H. assures the public, that all orders in his line will be attended to with punctuality and despatch.

Orders from the Southward will meet with immediate attention.

New-York, May, 1826.

WILLIAM BIGELOW,

BOOKBINDER.

No. 50 Fulton-street, Brooklyn.

BOOKBINDING, in all its various branches, executed with neatness and despatch.

Merchants' 'ccount Books, Writing Books, &c. ruled to patterns at the shortest notice.

Backgammon Tables, Chess Boards, Battledores, Dice Boxes, &c. wholesale and retail.—Maps, Pictures, &c. varnished in the neatest manner.

Brooklyn, 1826.

G. & R. WAITE'S LOTTERY AND EXCHANGE OFFICES,

Corner of Broadway and Maiden-lane, and the corner of Broadway and Fulton-street, New-York—south-west corner of Third and Chesnut-streets, Philadelphia—and corner of Charles and Market-streets, Baltimore. At all the above offices, Bank Notes are discounted at the lowest rates, and the highest premium given for Gold. Tickets and shares in all the Lotteries for sale. Cash advanced for prizes as soon as drawn. At Waites' offices have been sold and paid Prizes amounting to EIGHT MILLIONS OF DOLLARS.

NATIONAL HOTEL.

The public is respectfully informed, that the splendid edifice recently erected in the city of New-York, by Joseph Delacroix, Esq. is opened for their accommodation. - This establishment is situated in one of the most airy and agreeable parts of Broadway, nearly opposite the City Hotel, was built expressly for a house of public accommodation. and is believed to possess advantages for rendering a residence to strangers of business

or pleasure, pleasant and comfortable, not surpassed in this country.

The Leasers, anxious to render the National Hotel alike creditable to themselves and to this commercial metropolis, have spared no exertion or expense, to finish the establishment in a style they are fully confident will secure public approbation and support. Its immediate superintendance is confided to Mr. E. BOARDMAN, a gentleman whose experience and capacity of conducting a concern of this kind are well known to the public. and renders any pledge on the part of the proprietors for its proper management, unne-

Extensive arrangements have been made for the reception of permanent BOARDERS. and the accommodation of gentlemen with their families, with private suits of apart-May, 1826.

ments.

GEDNET KING.

MATHEMATICAL INSTRUMENT MAKER.

MO. 118

State-Street, opposite Broad-Street, BOSTON.

Ilas constantly for sale, wholesale and retail, a general assortment of Mathematical and Philosophical Instruments, of the best quality, (warranted,) comprising articles of almost every description in the mathematical line, viz :—

Sextants of ebony and metal, with silver, brass and ivory arches-Quadrants, with and without tangent and vertical screws-Day and Night Telescopes, with and without brass shades, and Telescopes of every description—azimuth, amplitude, storm, brass, and wood binnacle, hanging, and pocket Compasses—Binnacle Lamps, Time Glasses of every quality, Thermometers, Marine Barometers, Scales and Dividers, Parallel Rules. Protracters, cases of Instruments, &c. &c.

Bowditch's Practical Navigator, Blunt's American Coast Pilot, do. Seamanship and Naval Tactics, do. Nautical Almanacs, Shipmaster's Assistant, Merchant and Seaman's EXPEDITIOUS MEASURER, consisting of a set of Tables, which show at one view the Solid Contents of all kinds of packages and casks, according to their several Lengths, Breadths, and Depths: also, Rules for determining the contents of all sorts of casks in wine and beer measure—Stereotype Edition; Corrected by EDMUND M. BLUNT, Author of the American Coast Pilot, &c.—WARD'S Lunar Tables, together with every Nautical publication of merit, including an assortment of the most useful

Sextants, Quadrants, Compasses, Time Glasses, and other Instruments, cleaned and repaired at the shortest notice, and on the most reasonable terms.

April, 1826.

J. M. ELFORD'S

CHART AND MATHEMATICAL STORE.

No. 119 East Bay, sign of the Quadrant, CHARLESTON, S. C. OLD ESTABLISHED STAND.

FOR SALE-Charts, Nautical Books, and Mathematical Instruments of every description. Compasses, Quadrants, Spy-Glasses, &c. repaired and for sale. Chronometers rated

Published and for sale, J. M. Elford's LONGITUDE TABLES, being the shortest and most simple method of working Lunar Observations of any in practice. Elford's Circular POLAR TABLES, for finding the Latitude at any time of night by an Altitude of the Polar Star. Elford's Universal and Perpetual Circular TIDE TABLES, for finding the time of High Water every day in the year, at all the principal places in the world, by inspection or at sight. Also—The UNIVERSAL SIGNAL BOOK, with improvements, by J. M. Elford.

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N. B. An EVENING SCHOOL from 6 till 9—and private lessons given upon Lunar Observations at intervals. April, 1996.

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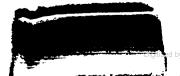
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